



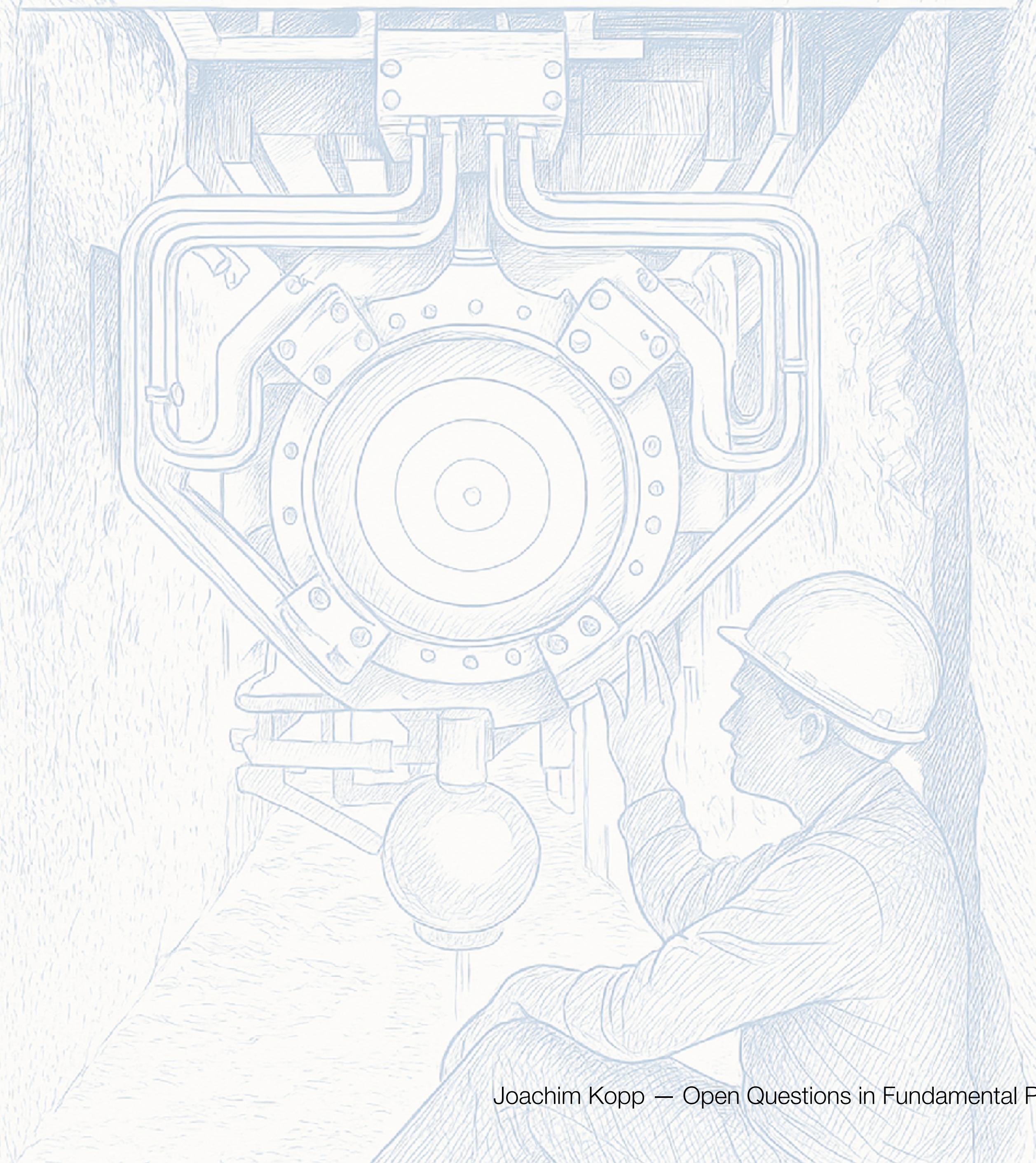
PARALLEL 3 / NEUTRINOS AND COSMIC MESSENGERS

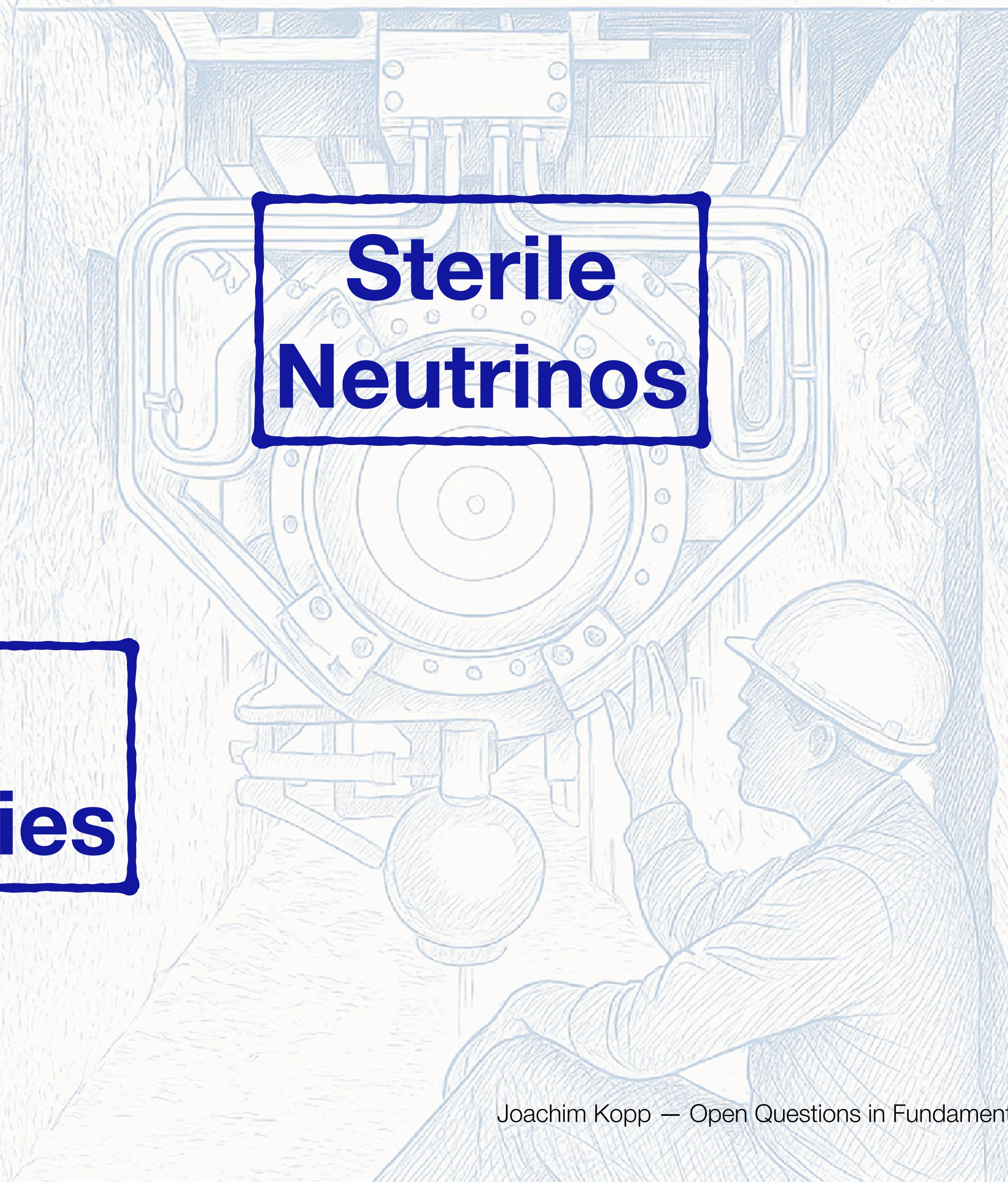
BSM Searches in Neutrino Experiments

Joachim Kopp (CERN and JGU Mainz)

23-27 JUNE 2025 Lido di Venezia







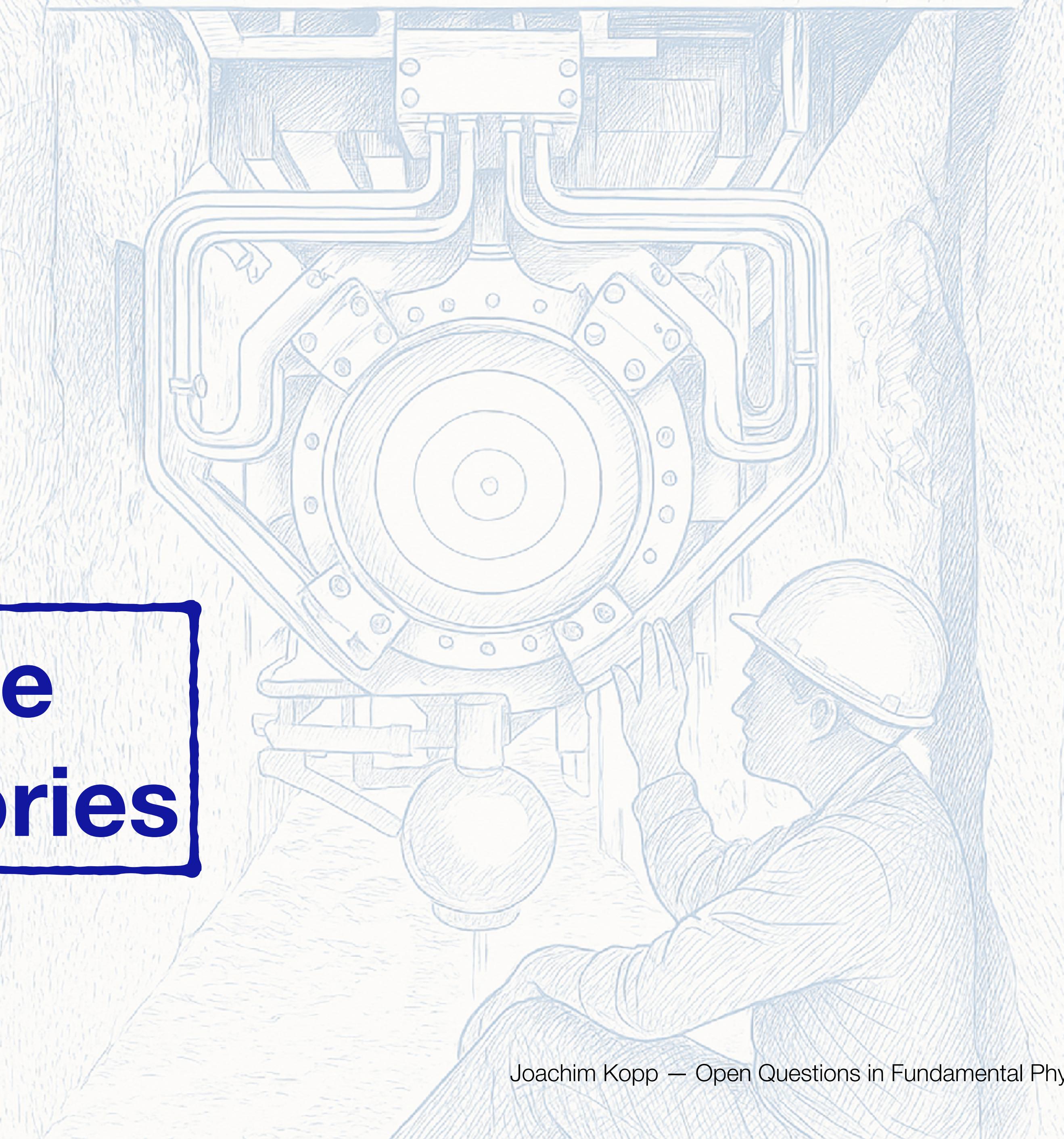
Sterile Neutrinos

Effective Field Theories

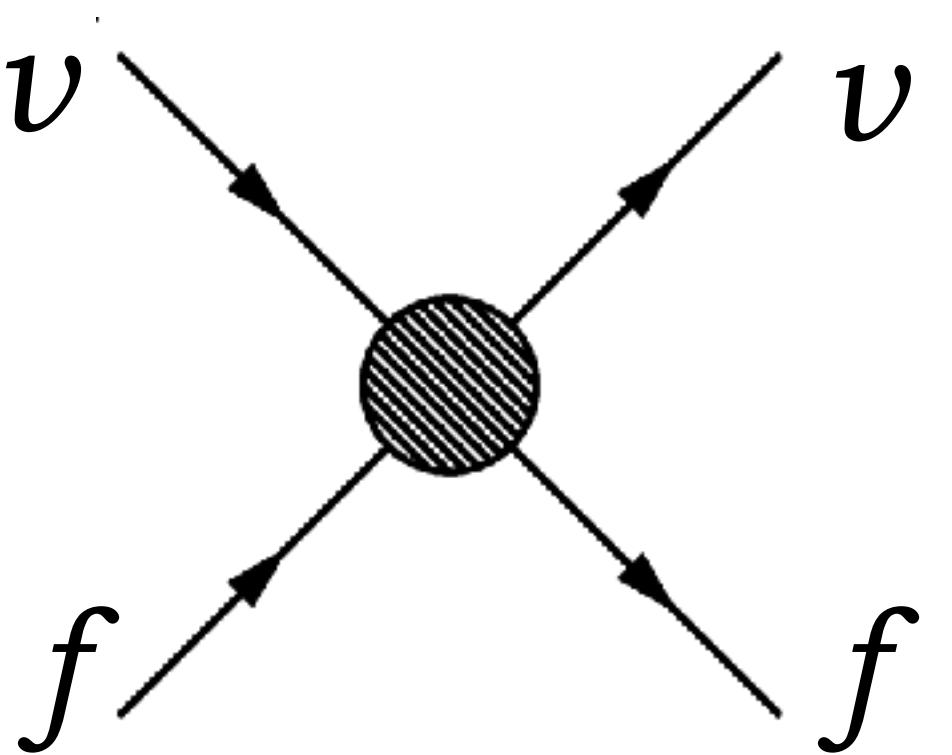
Dark Sectors



Effective Field Theories



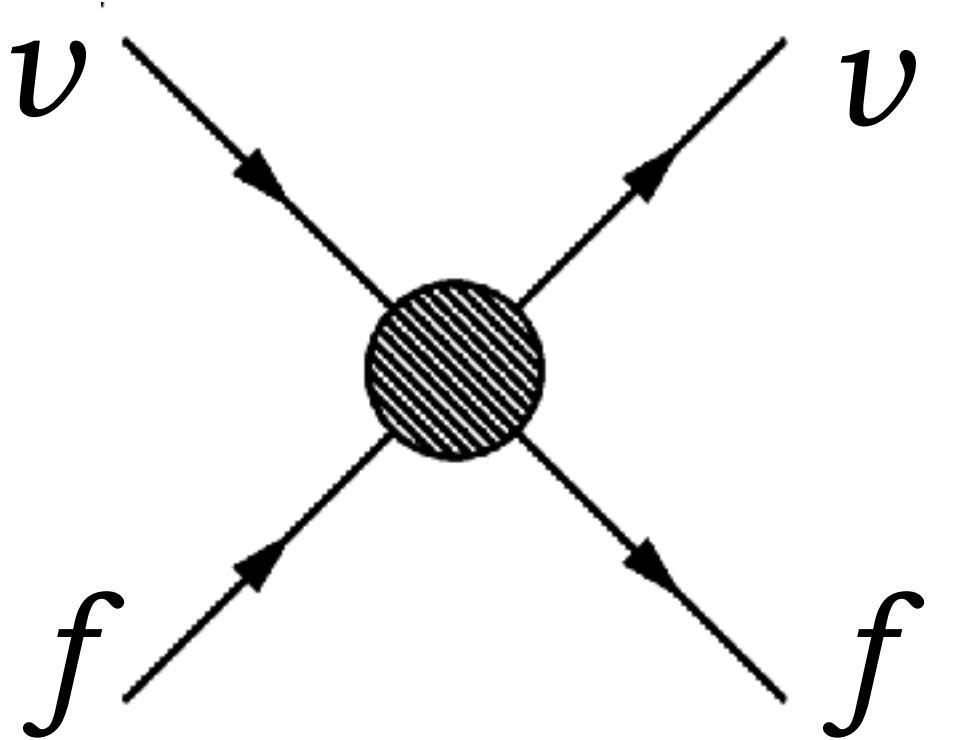
Effective Field Theories



Effective Field Theories

- Standard Model Effective Field Theory (SMEFT)
parameterises new physics above the electroweak scale

$$\mathcal{L}_{\text{SMEFT}} = \sum_i \frac{c_i^{(5)}}{\Lambda} \mathcal{O}_i^{(5)} + \sum_i \frac{c_i^{(6)}}{\Lambda^2} \mathcal{O}_i^{(6)} + \dots$$



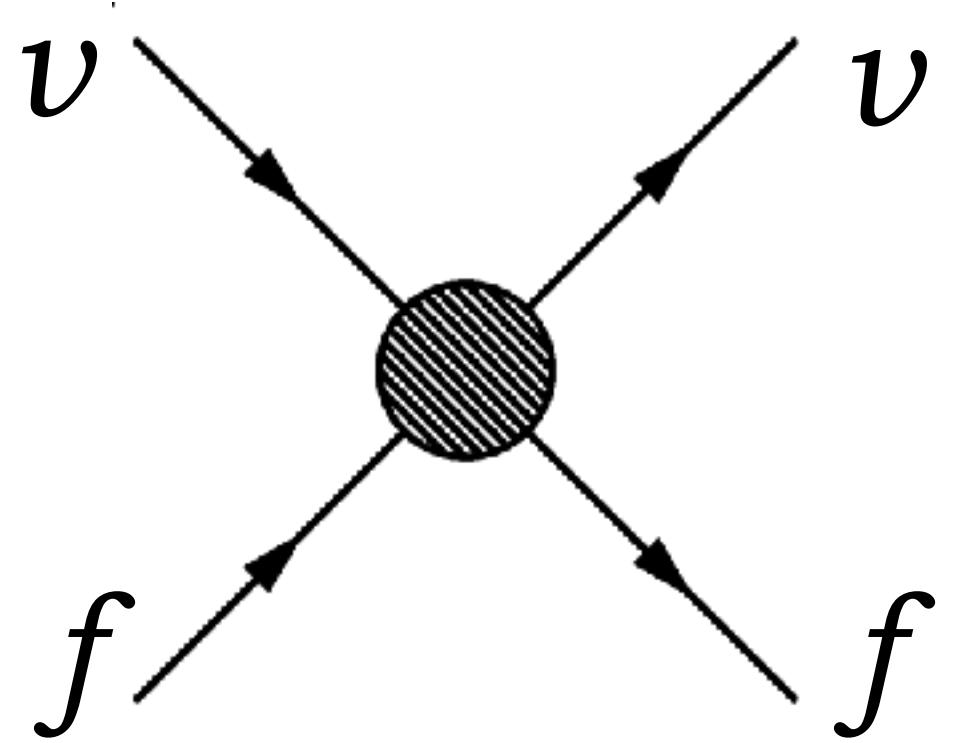
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↑
1 operator
(Majorana neutrino mass)

↑
2499 operators

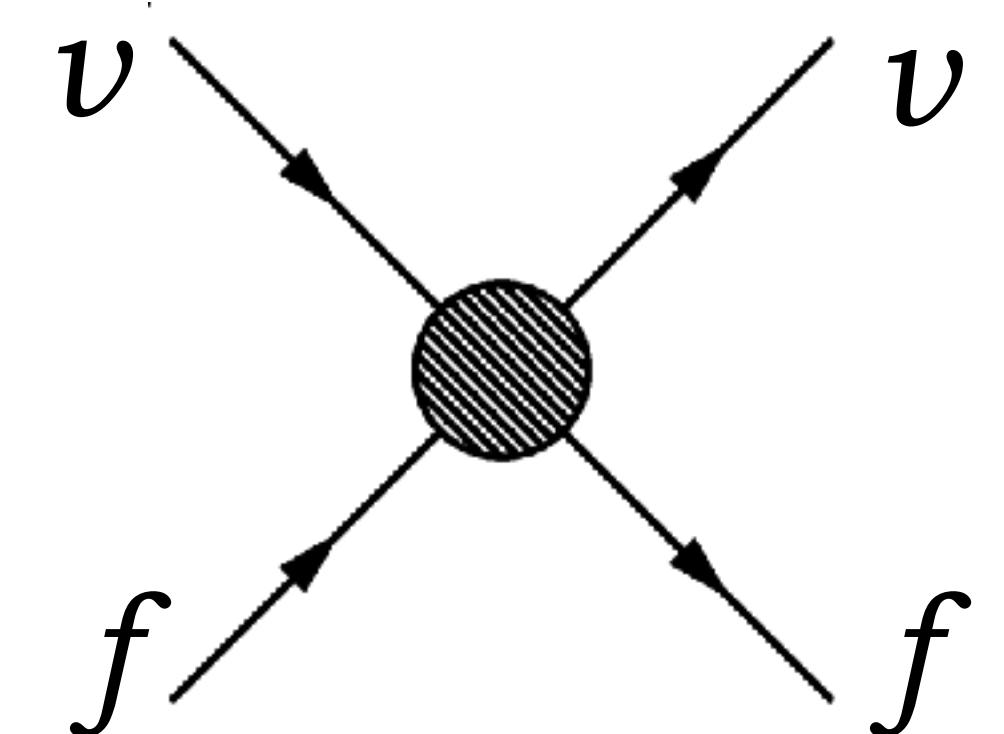


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- For **low-energy experiments** (like neutrino experiments):

map onto **Weak Effective Field Theory (WEFT)**: “Non-Standard Interactions” (NSI)
(EFT valid below the electroweak scale)

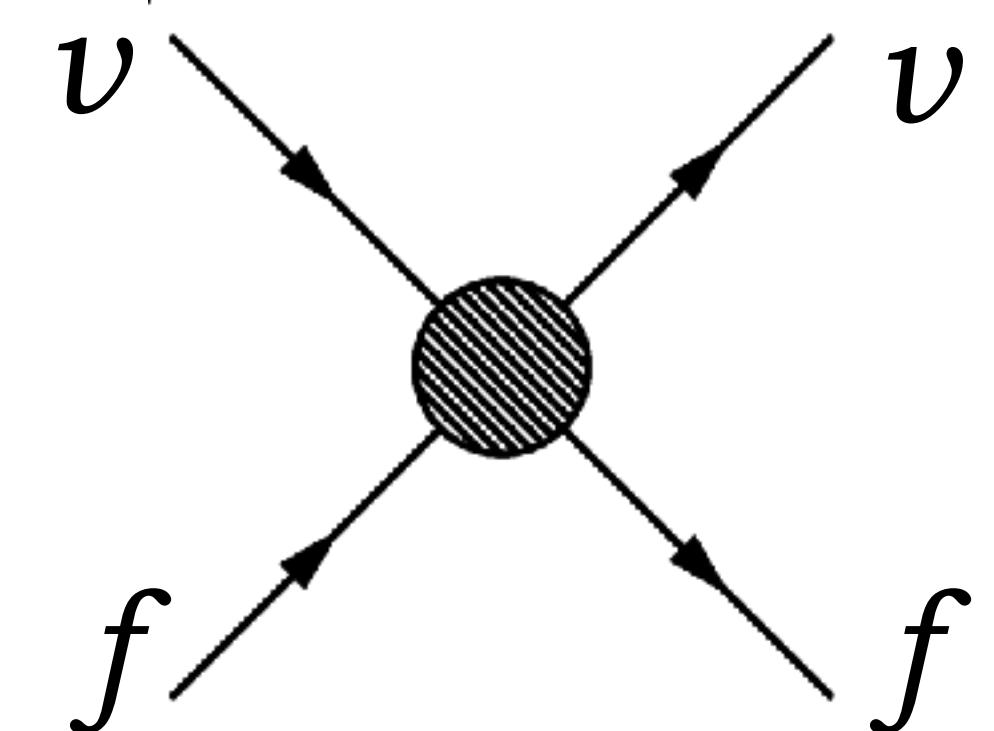
$$\mathcal{L}_{\text{NC}} = \sum 2\sqrt{2}G_F \epsilon_{\alpha\beta}^{q,XY} (\bar{\nu}_\alpha \Gamma_X P_L \nu_\beta) (\bar{q} \Gamma_Y q)$$

$$\mathcal{L}_{\text{CC}} = \sum 2\sqrt{2}G_F \epsilon_{\alpha\beta}^{qq',XY} (\bar{\ell}_\alpha \Gamma_X P_L \nu_\beta) (\bar{q}' \Gamma_Y q)$$

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dim-6 operators
with different Lorentz structures

dimensionless coefficients

(interaction strength relative to SM weak interactions)

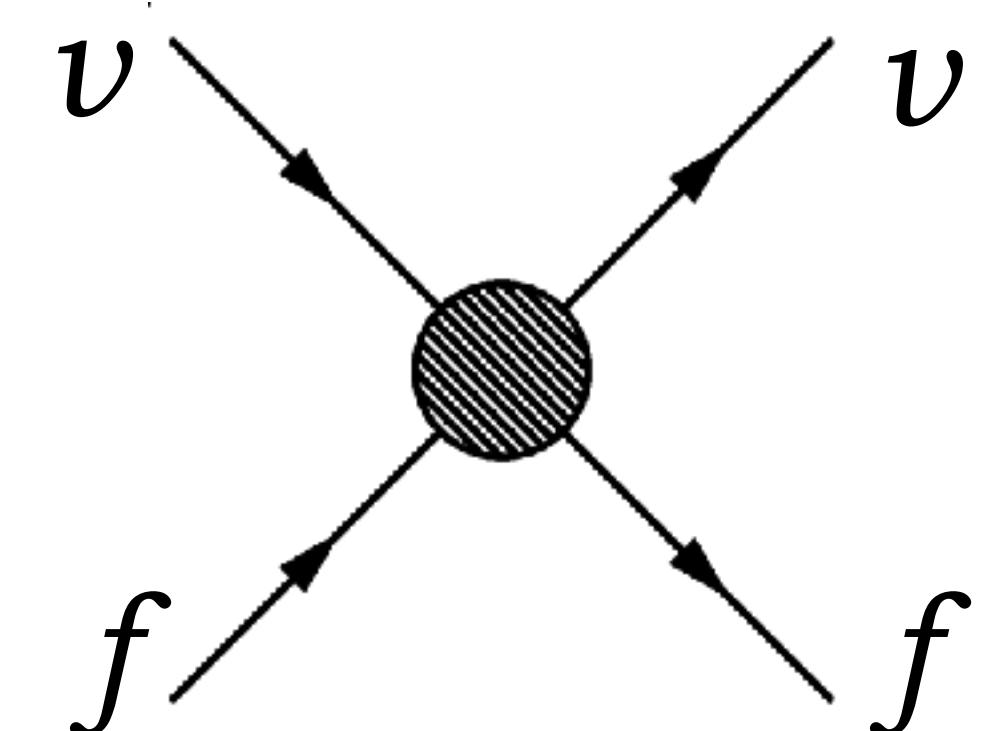


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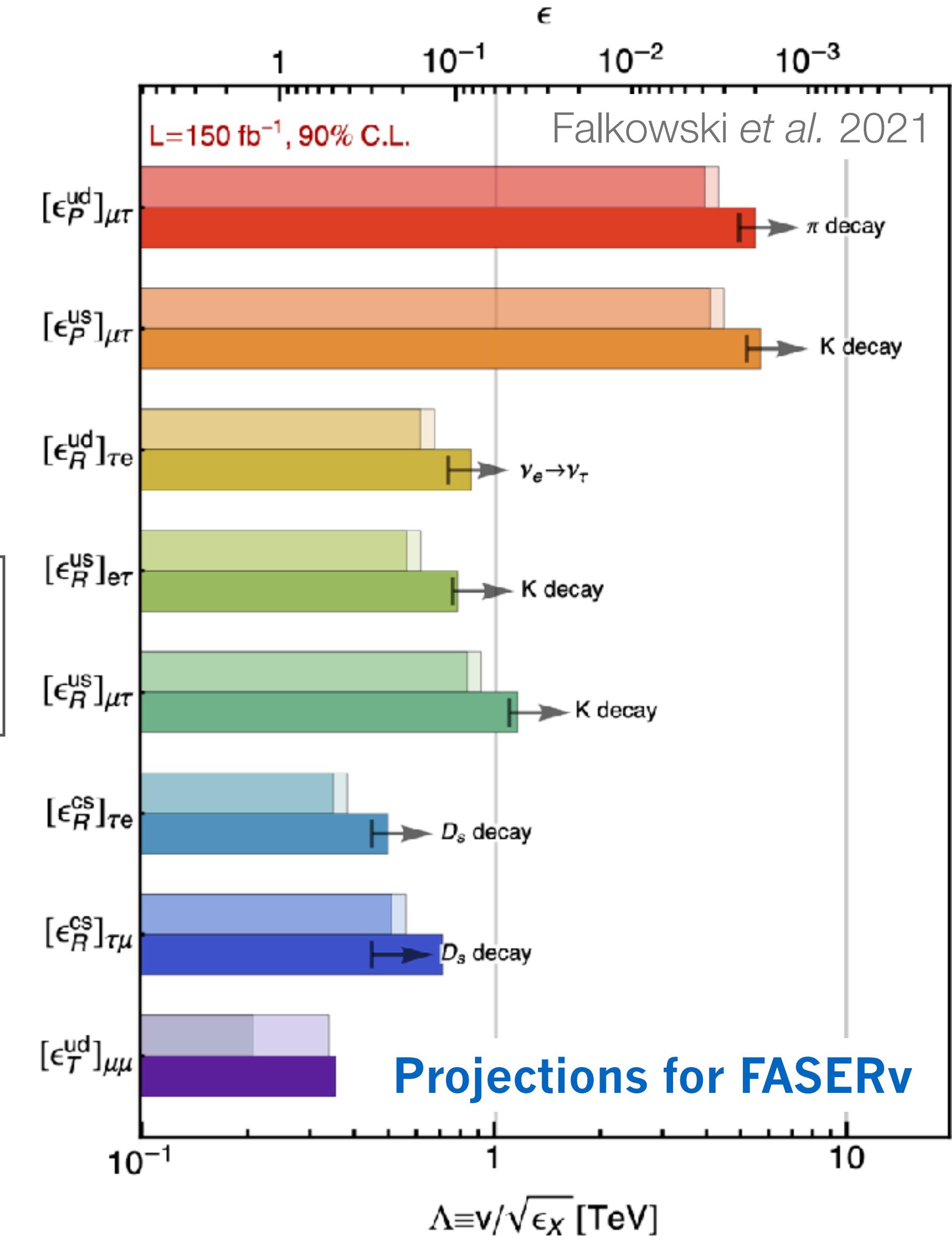
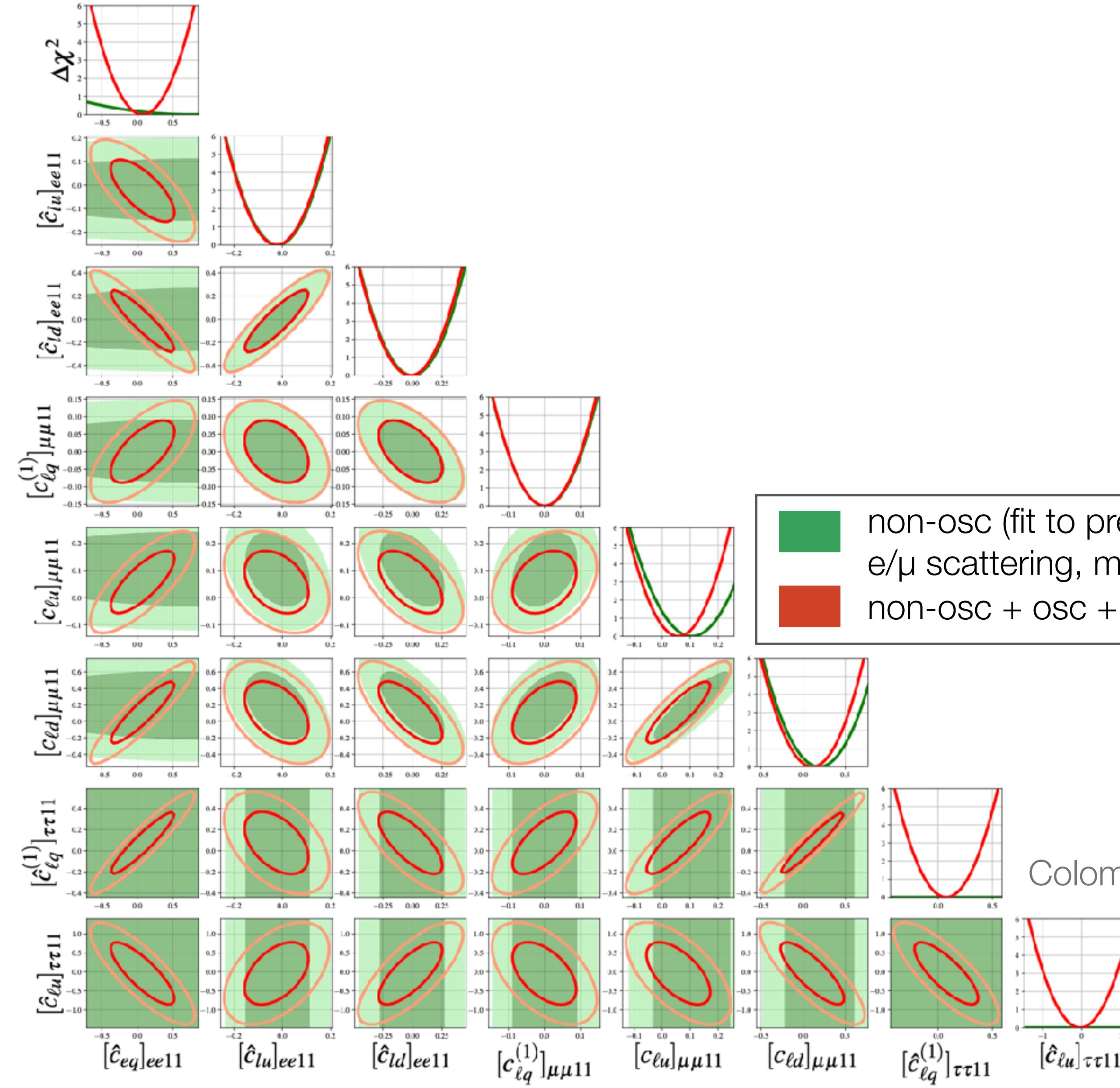
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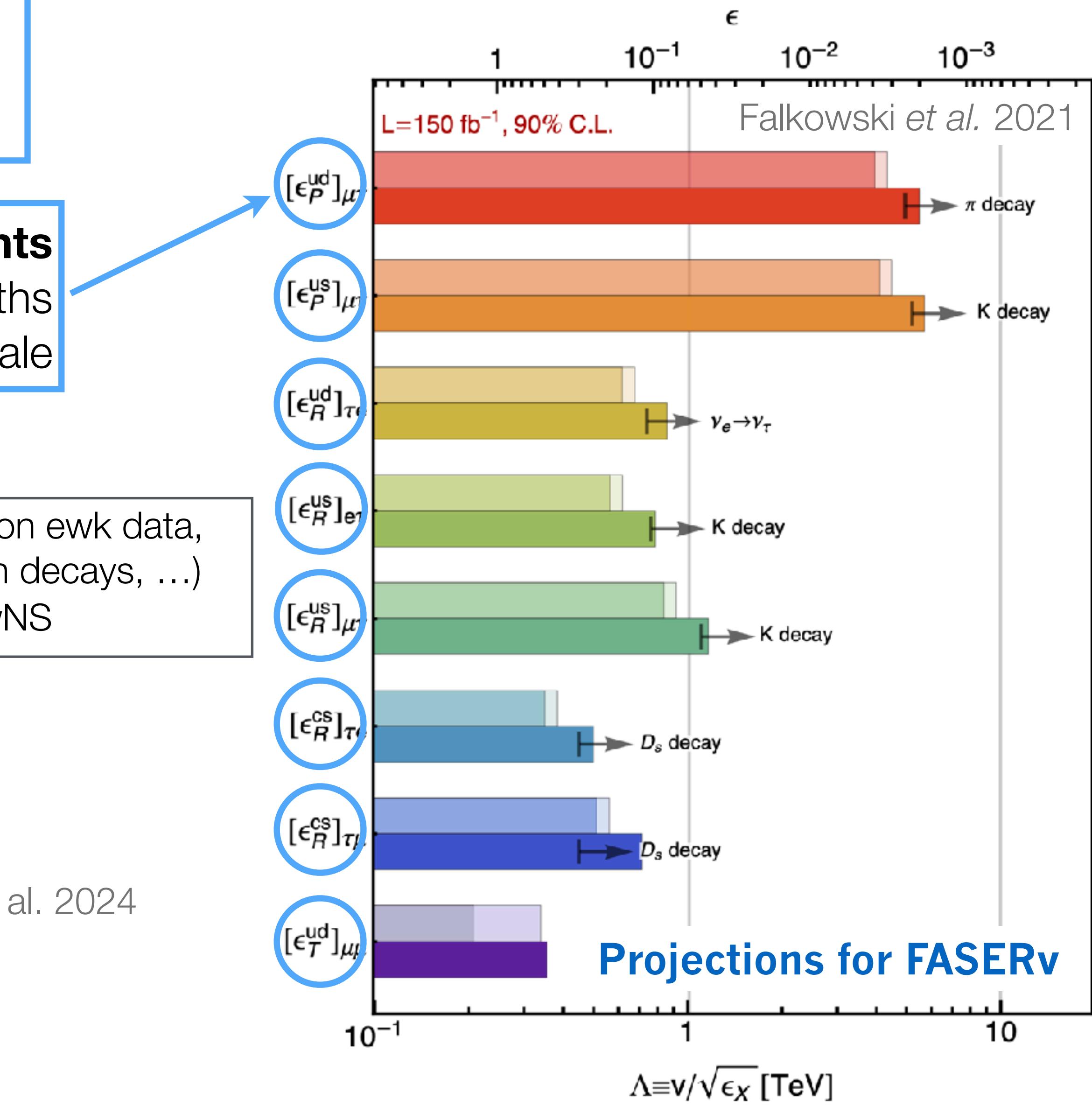
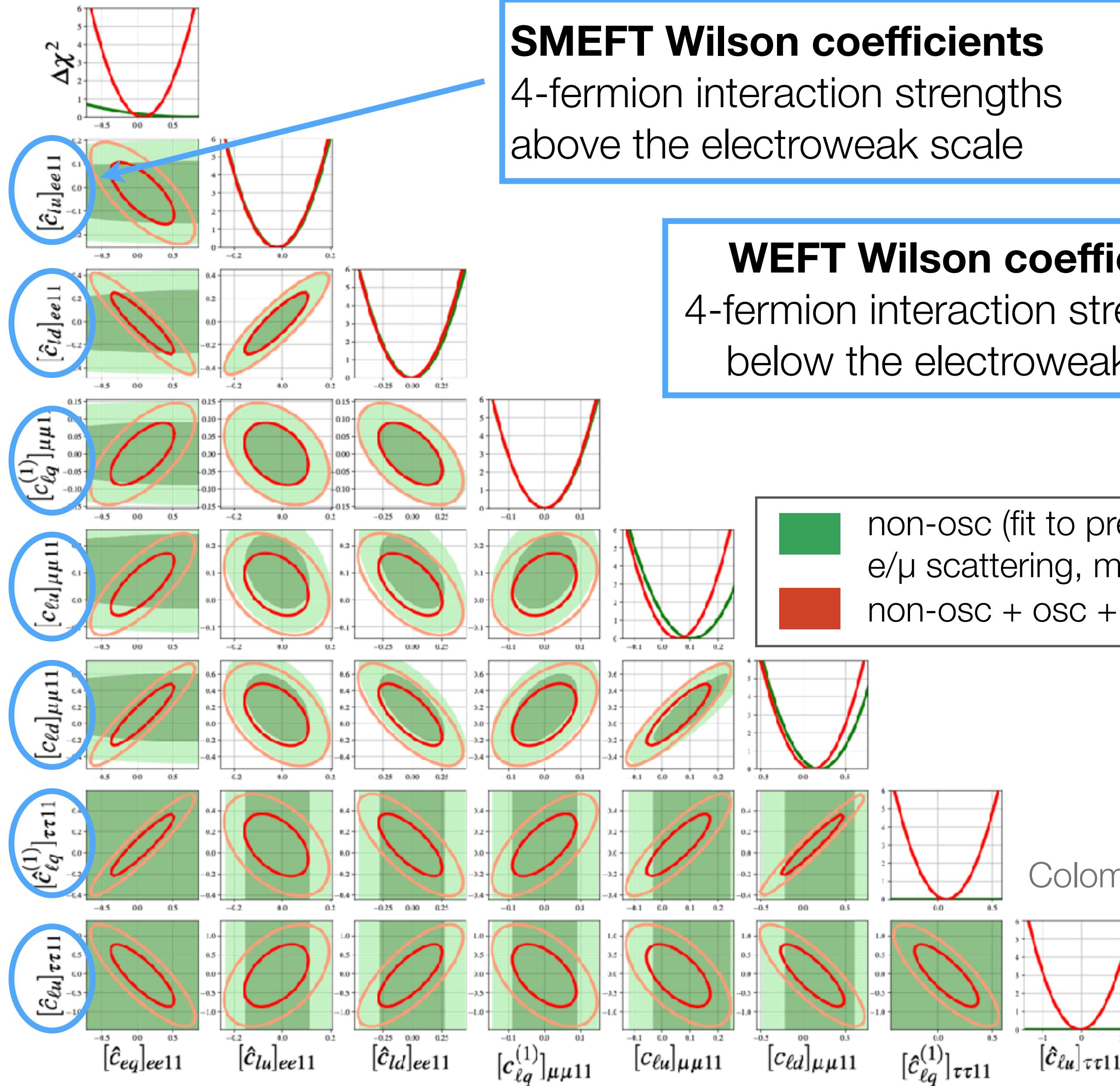
$$\mathcal{L}_{\text{NC}} = \sum 2\sqrt{2}G_F \epsilon_{\alpha\beta}^{q,XY} (\bar{\nu}_\alpha \Gamma_X P_L \nu_\beta) (\bar{q} \Gamma_Y q) \quad \text{(new matter effects in oscillations)}$$

$$\mathcal{L}_{\text{CC}} = \sum 2\sqrt{2}G_F \epsilon_{\alpha\beta}^{qq',XY} (\bar{\ell}_\alpha \Gamma_X P_L \nu_\beta) (\bar{q}' \Gamma_Y q) \quad \text{(new nu production/detection processes)}$$

EFTs: Importance of Neutrino Constraints

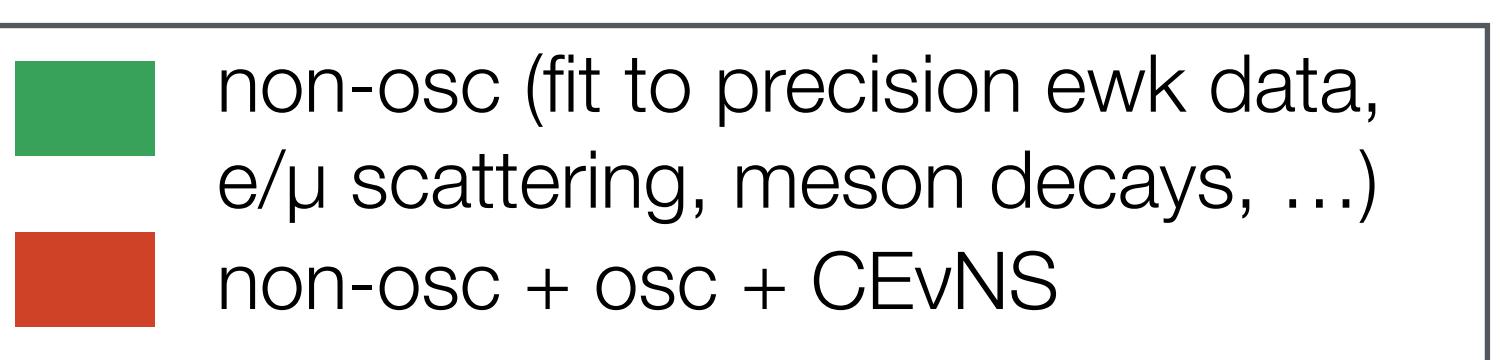


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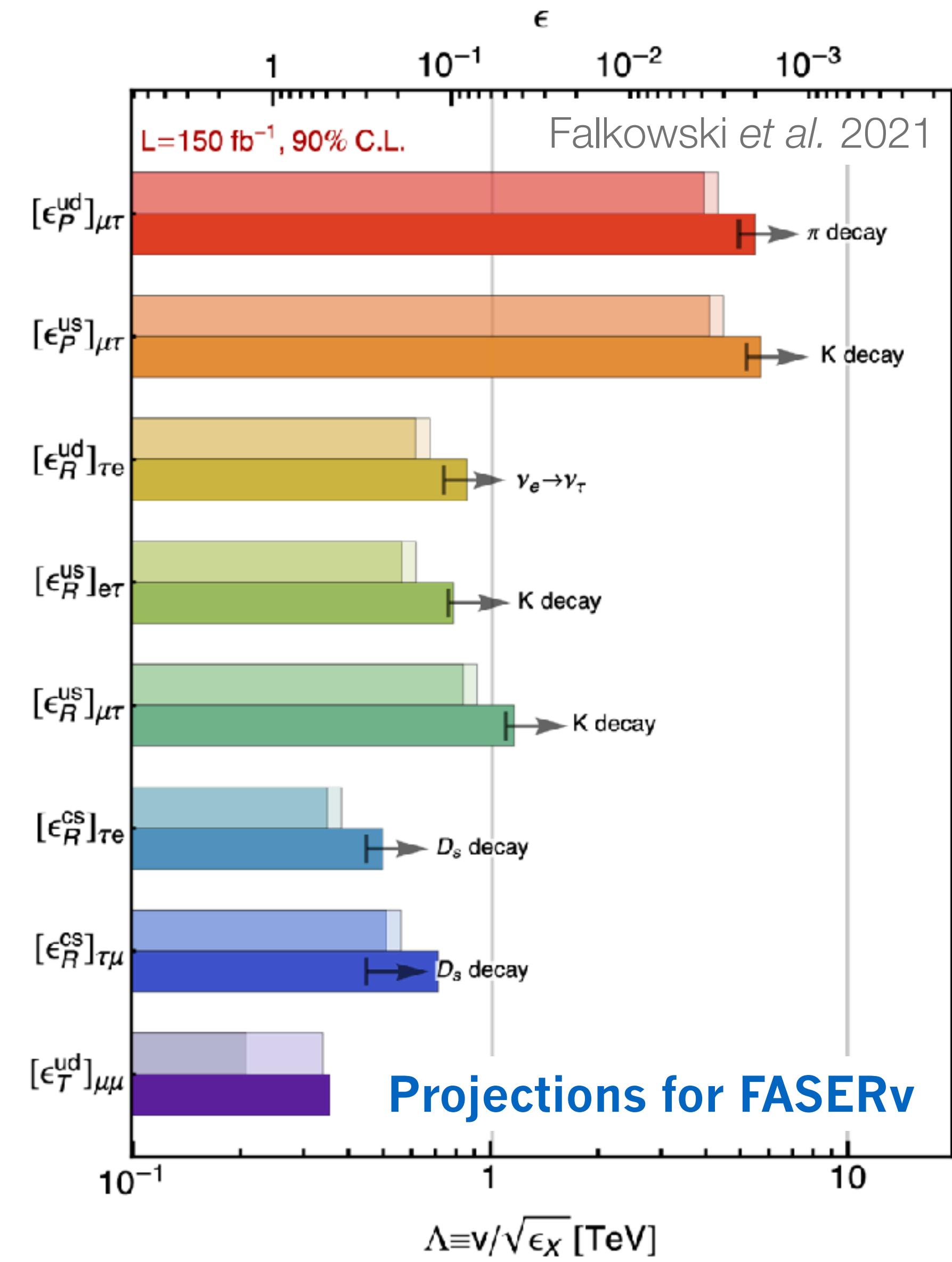
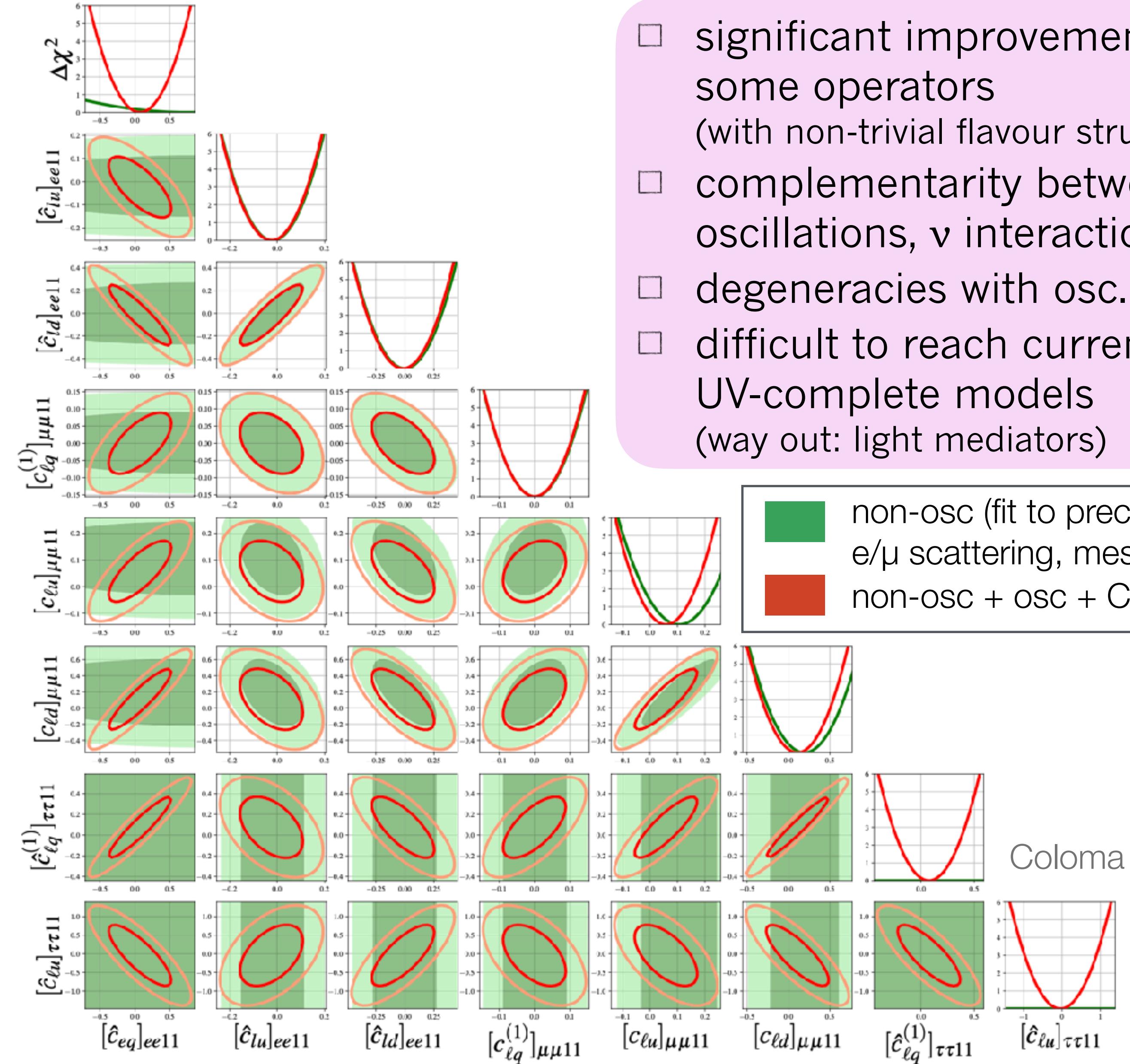


EFTs: Importance of Neutrino Constraints

- significant improvement for some operators
(with non-trivial flavour structures)
- complementarity between oscillations, ν interactions, CEvNS
- degeneracies with osc. params
- difficult to reach current bounds in UV-complete models
(way out: light mediators)

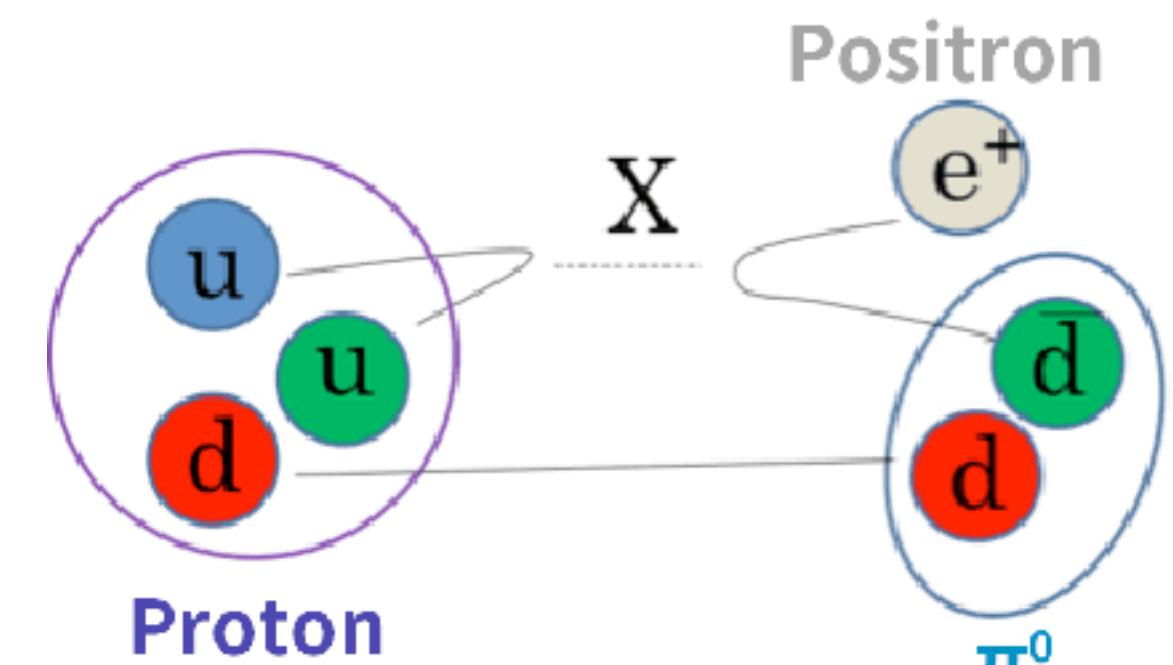


Coloma et al. 2024

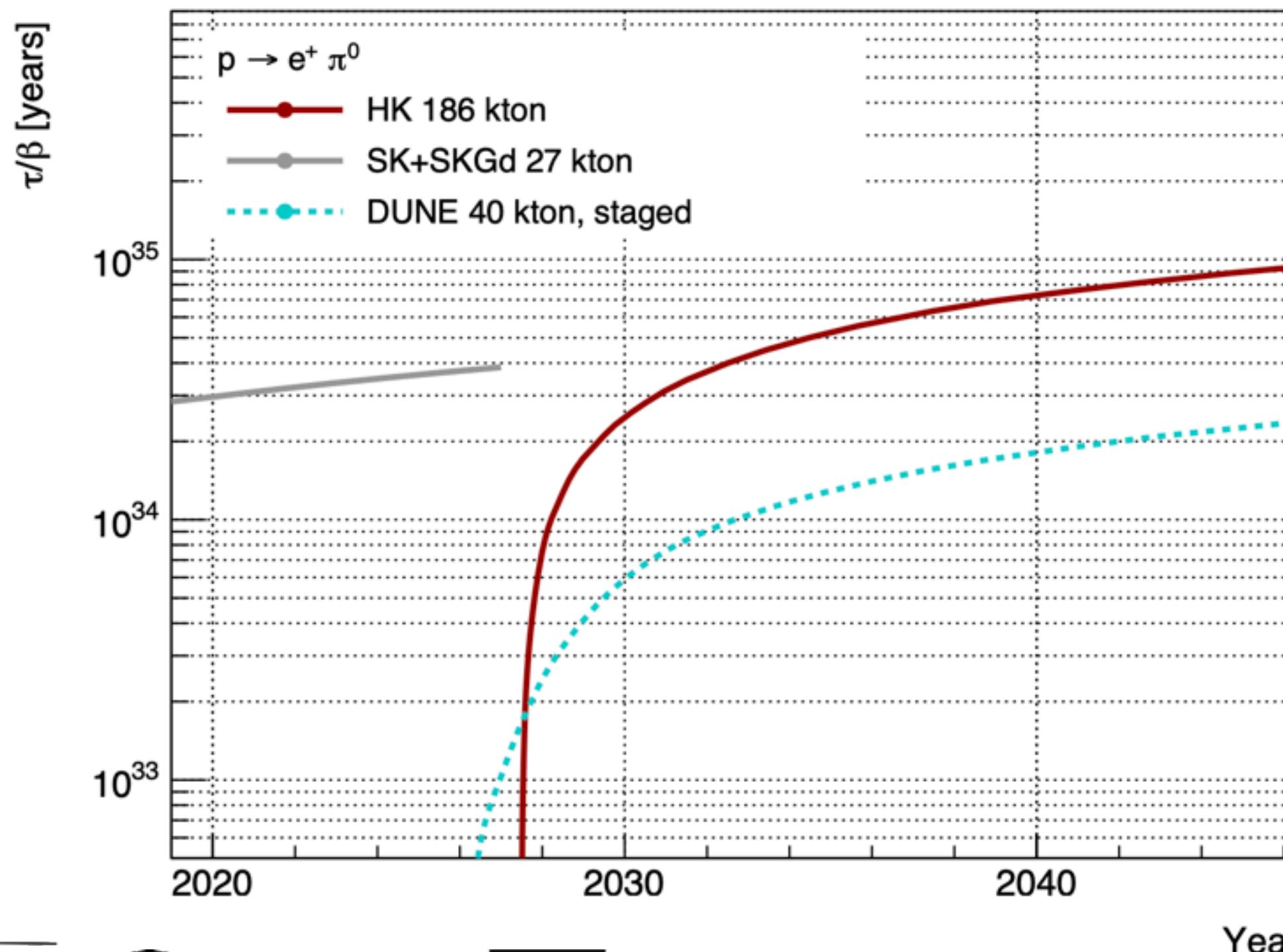


Nucleon Decay

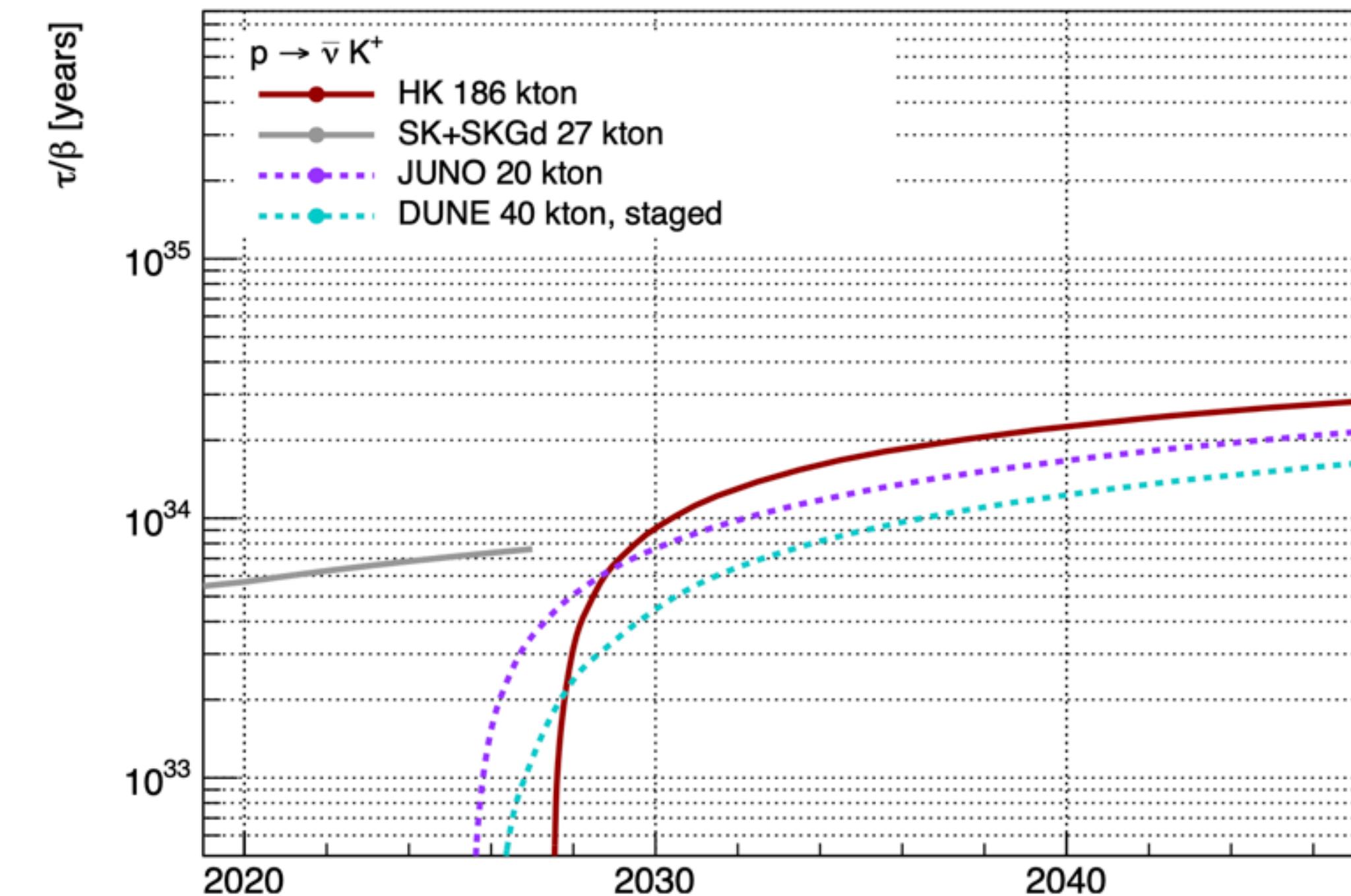
- Hallmark signature of **Grand Unification**
- **HyperK**: lots of protons, kaon ID only via decay products
DUNE: excellent particle ID, lots of *neutrons*



$$p \rightarrow e^+ \pi^0$$



$$p \rightarrow \bar{\nu} K^+$$



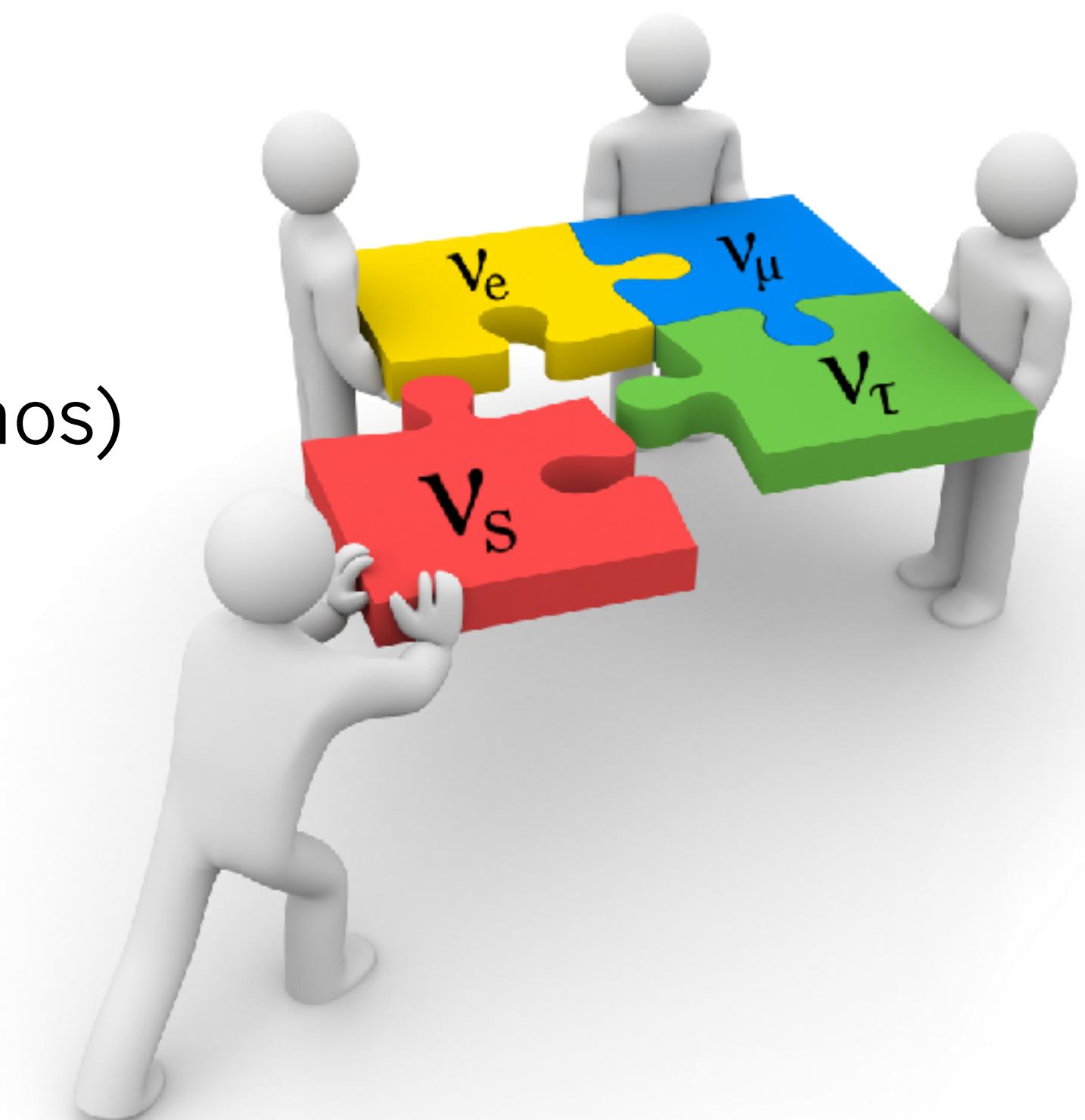
Sterile Neutrinos

Sterile Neutrinos

- **Neutrino Portal:** SM ν can mix with singlet fermions
(sterile neutrinos = heavy neutral leptons = right-handed neutrinos)
⇒ the only renormalisable SM coupling to singlet fermions
(dark matter?)

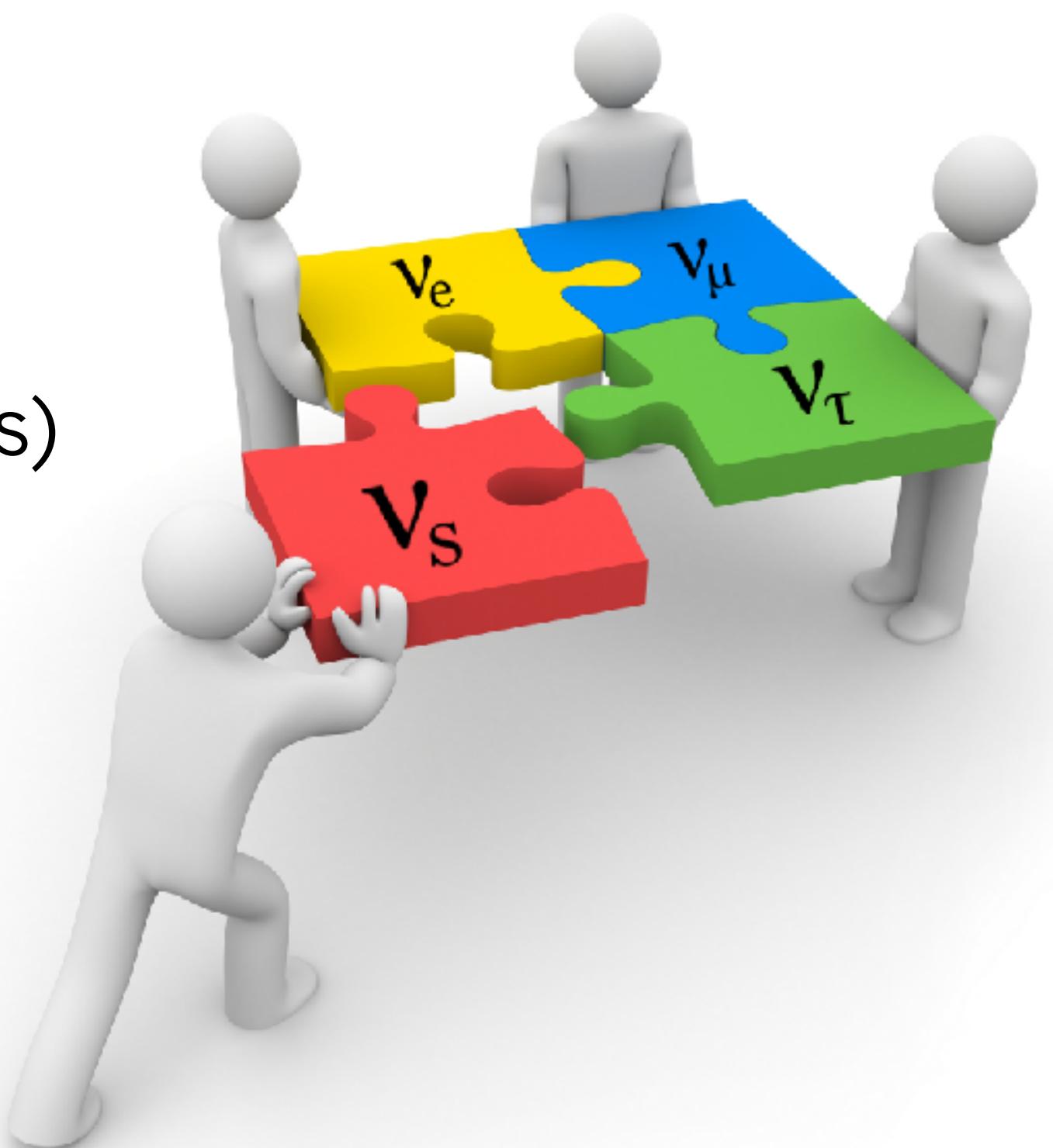
$$\mathcal{L} \supset y \bar{L} (i\sigma^2 H^*) N$$

↑ ↑ ↗
SM lepton SM Higgs sterile
doublet doublet neutrino

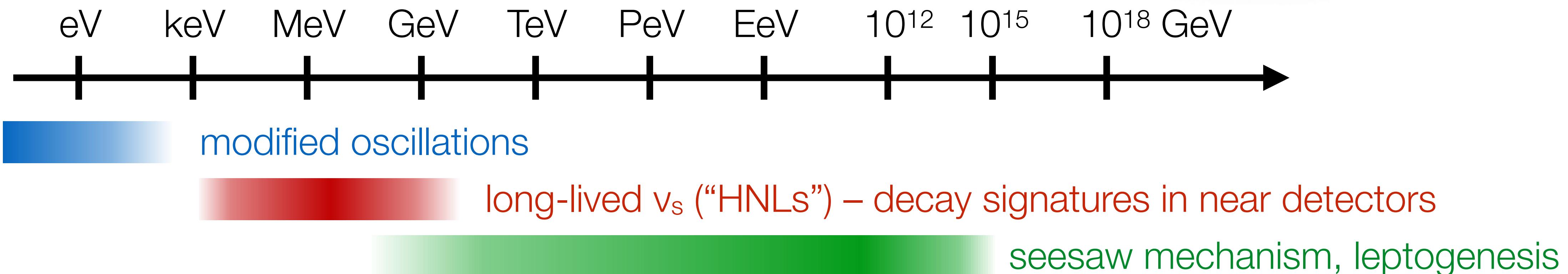


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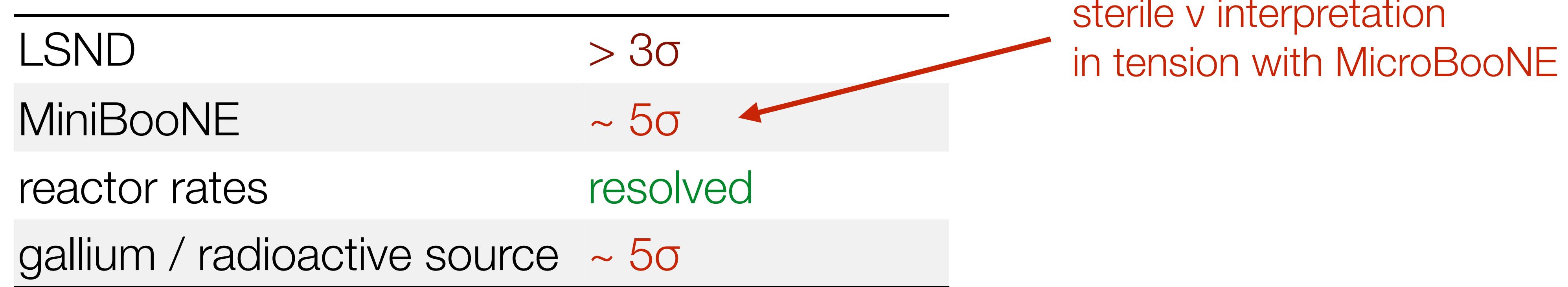


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eV-Scale Sterile Neutrinos

- long-standing anomalies, no consistent explanation

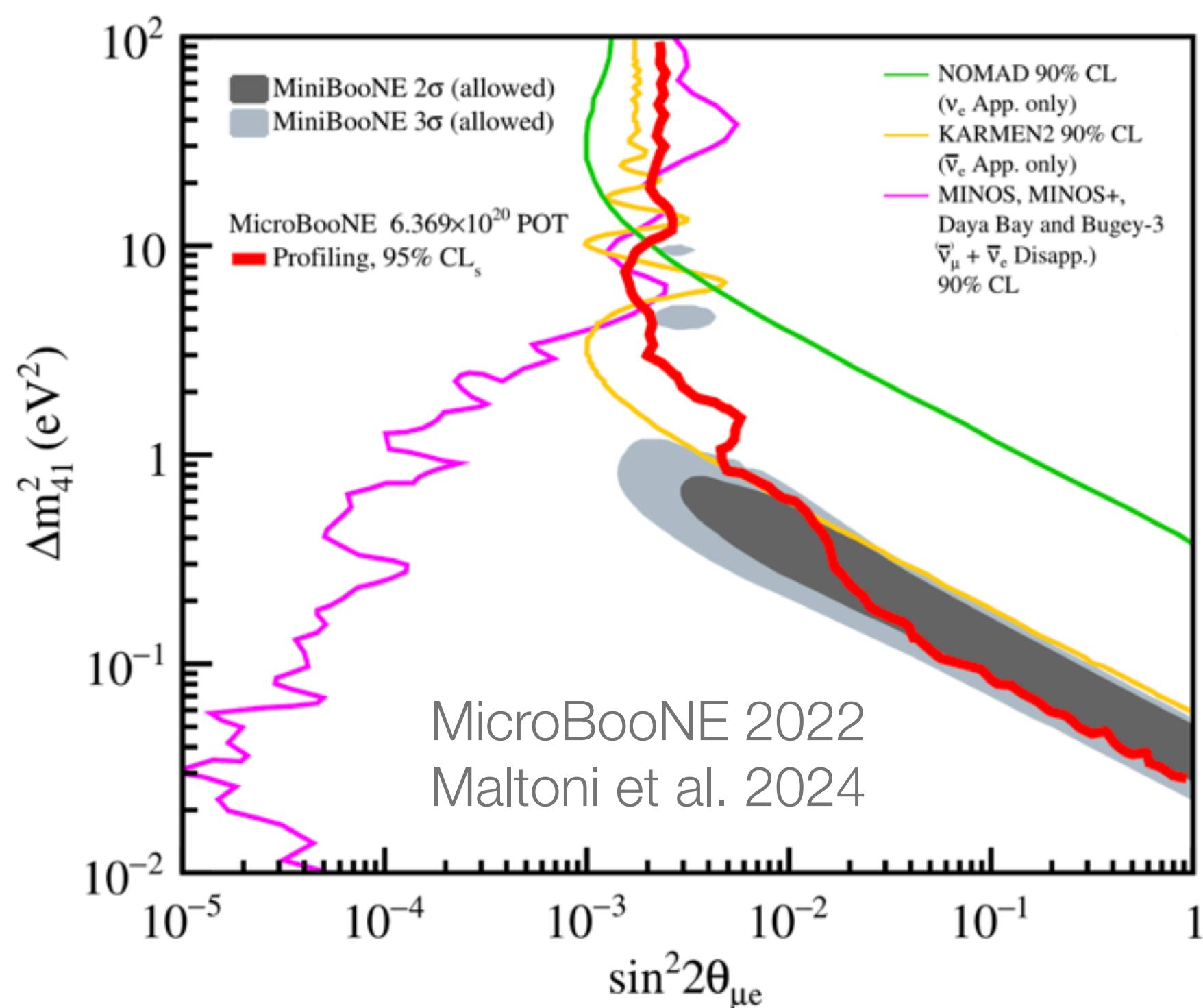


- vigorous experimental programme underway to resolve anomalies (SBN at FNAL)
 - expect important insights into ν interaction physics (or discover new physics)
- independent of anomalies sterile neutrinos remain an interesting possibility
 - integral part of BSM programme at upcoming experiments (DUNE, HyperK, JUNO, ...)
 - target windows in specific models

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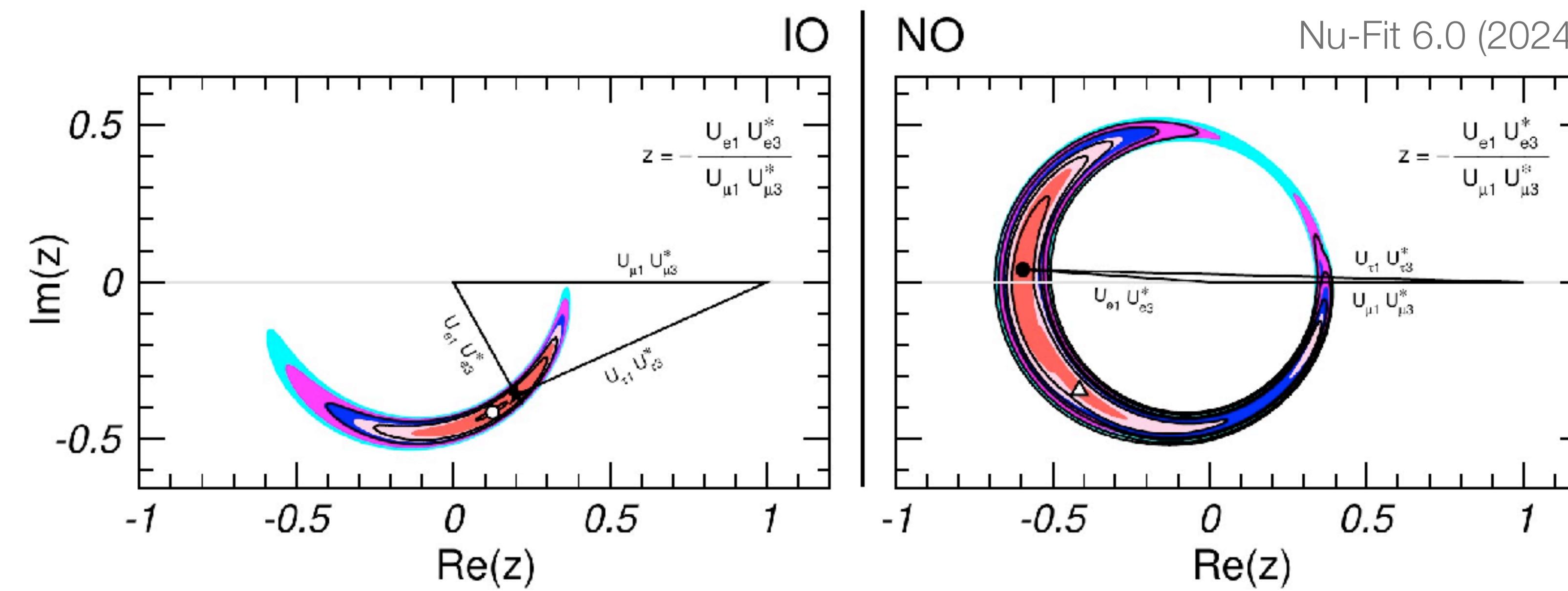
LSND	> 3 σ
MiniBooNE	~ 5 σ
reactor rates	resolved
gallium / radioactive source	~ 5 σ



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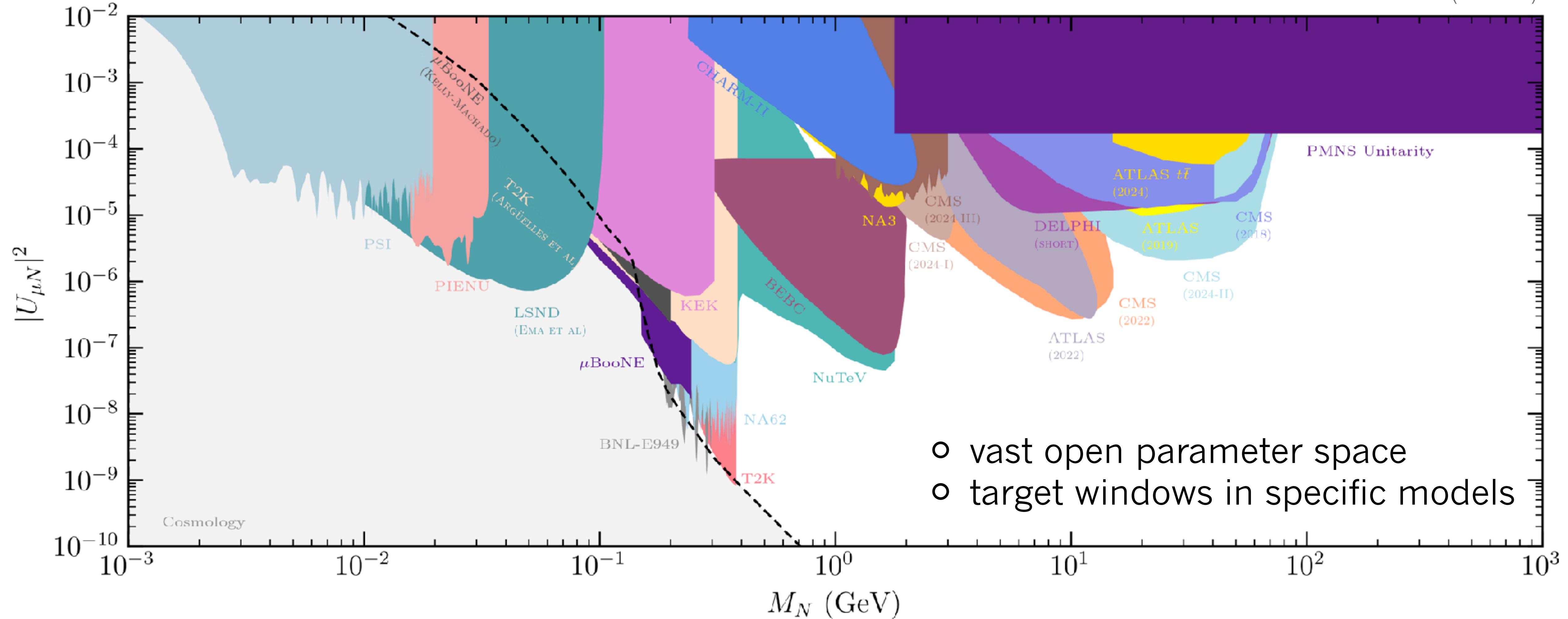
Heavy Neutral Leptons

- sterile neutrinos with $m \gtrsim \text{MeV}$
- can be produced in meson decays
- long-lived \rightarrow detection via downstream **decay**
- inaccessible in oscillation experiments \rightarrow effective PMNS matrix **non-unitary**

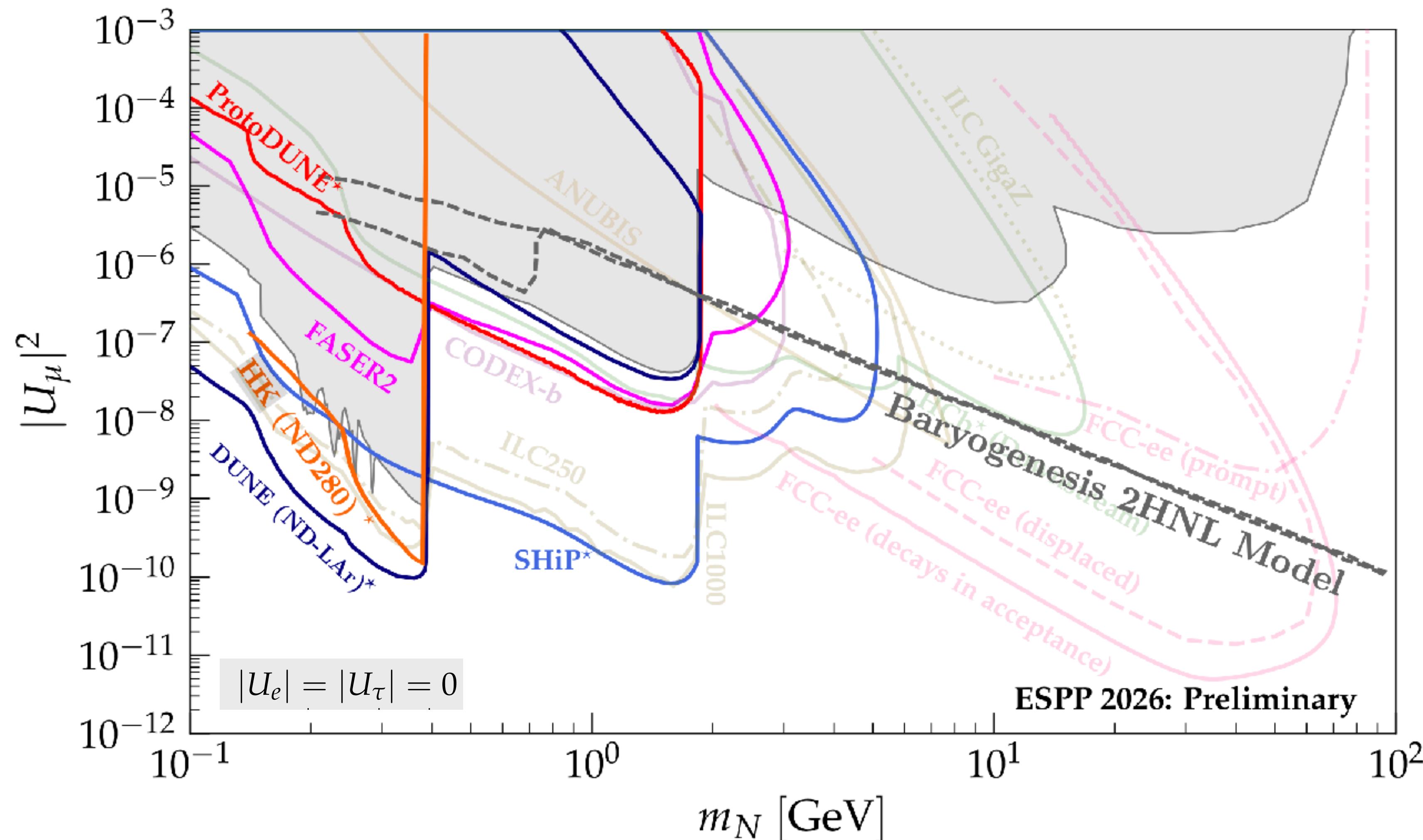


Heavy Neutral Leptons

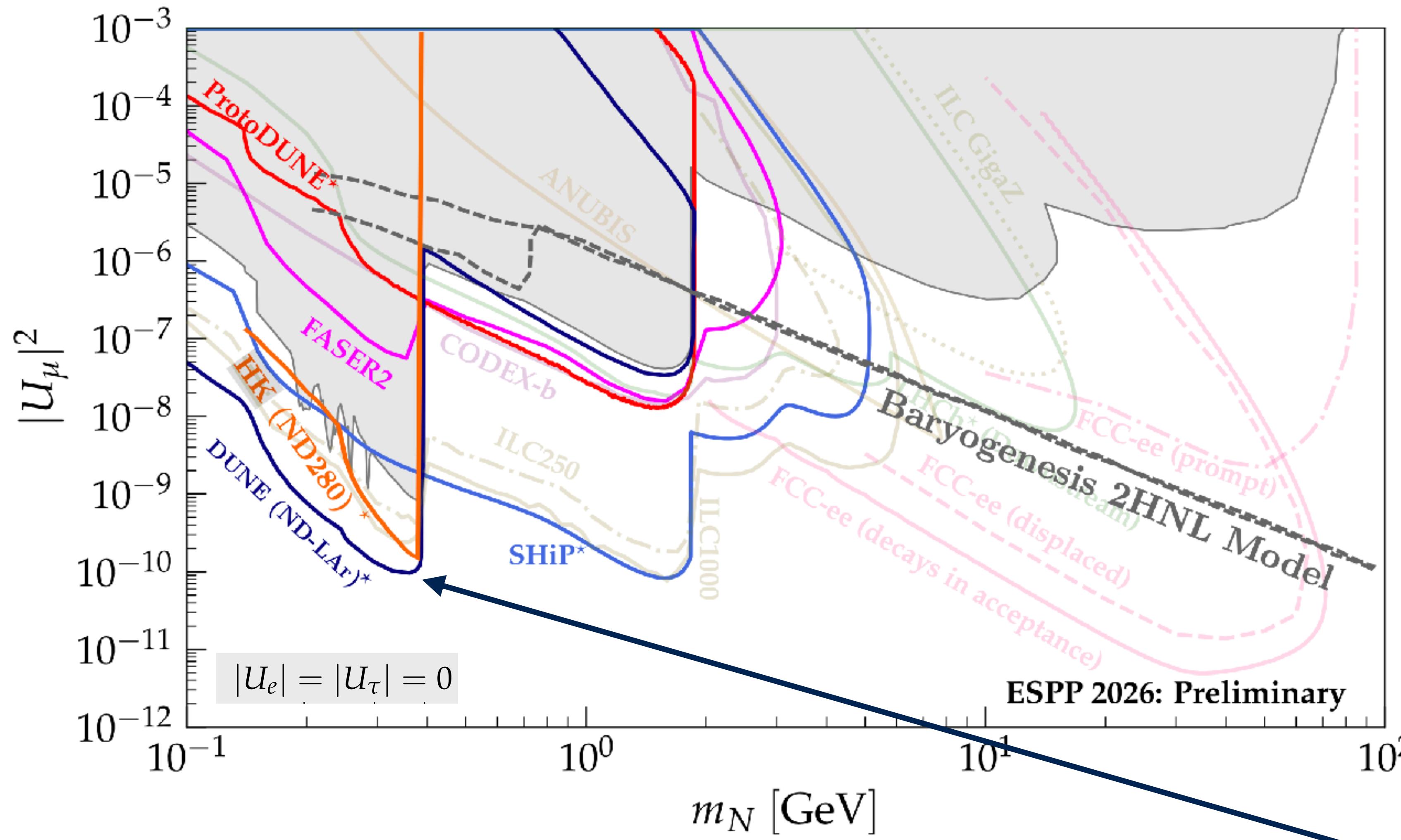
Fernandez-Martinez et al. 2023 / Hostert 2025 (GitHub)



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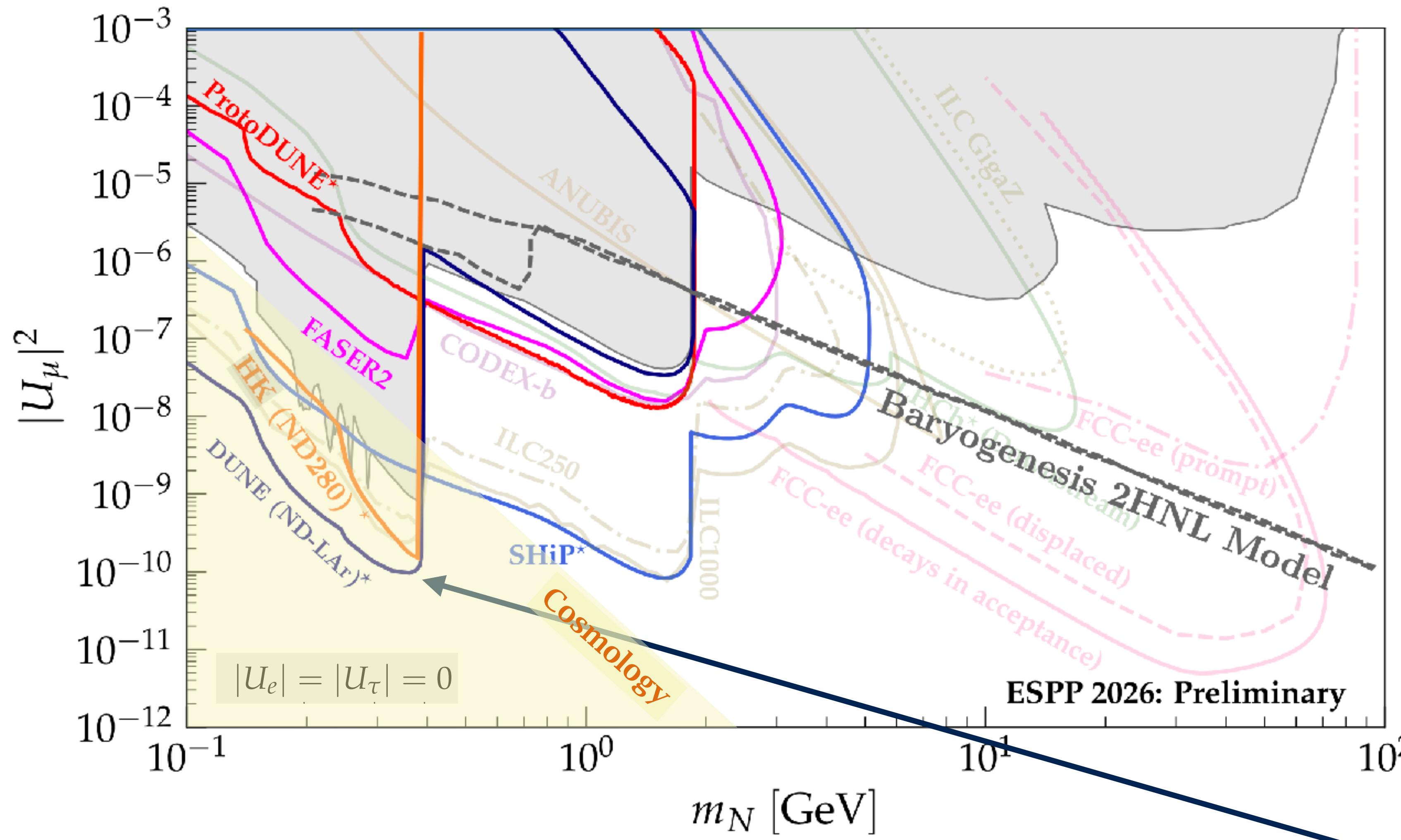


Heavy Neutral Leptons



below kaon threshold:
neutrino beams dominate
(but cosmologically disfavoured)

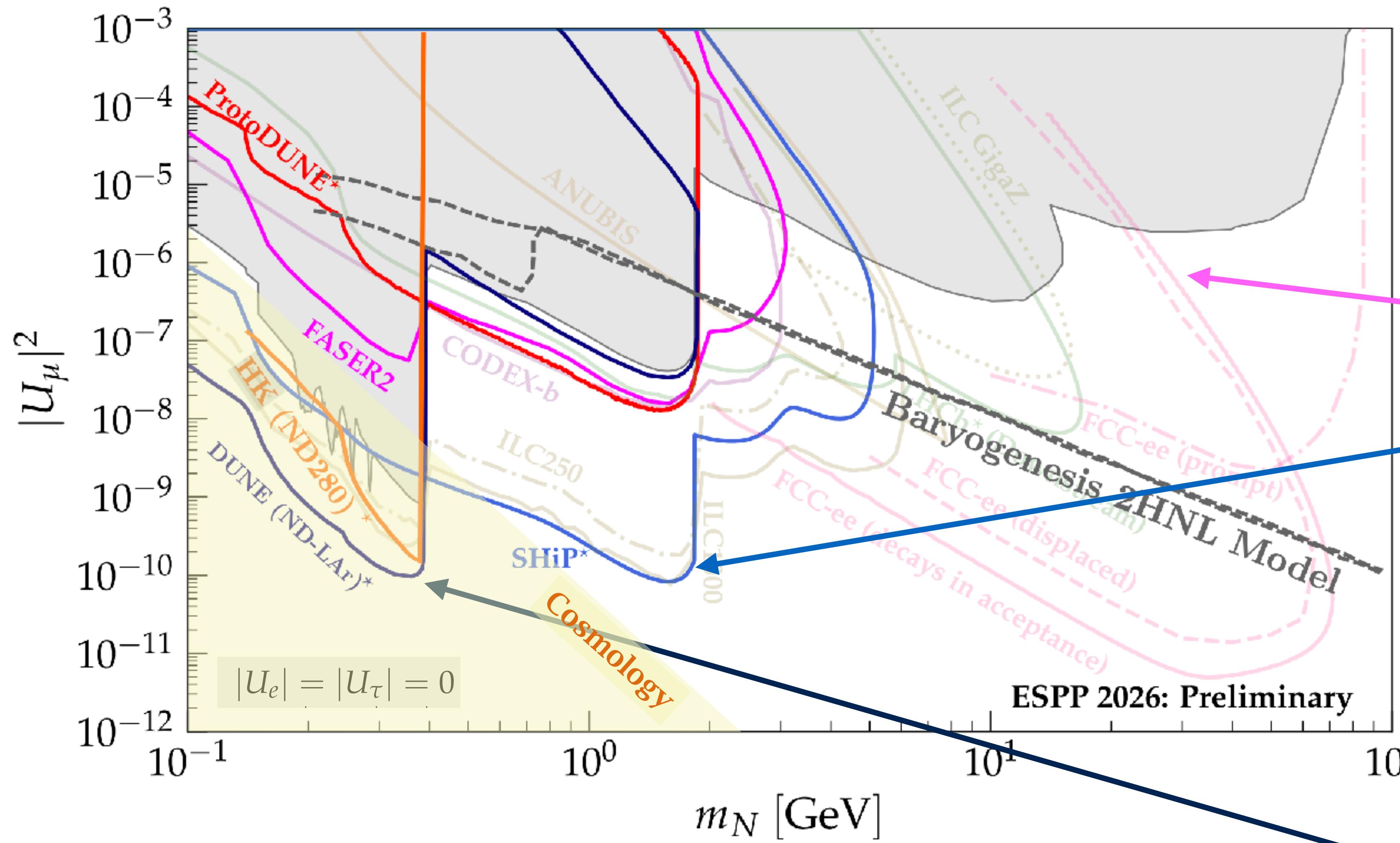
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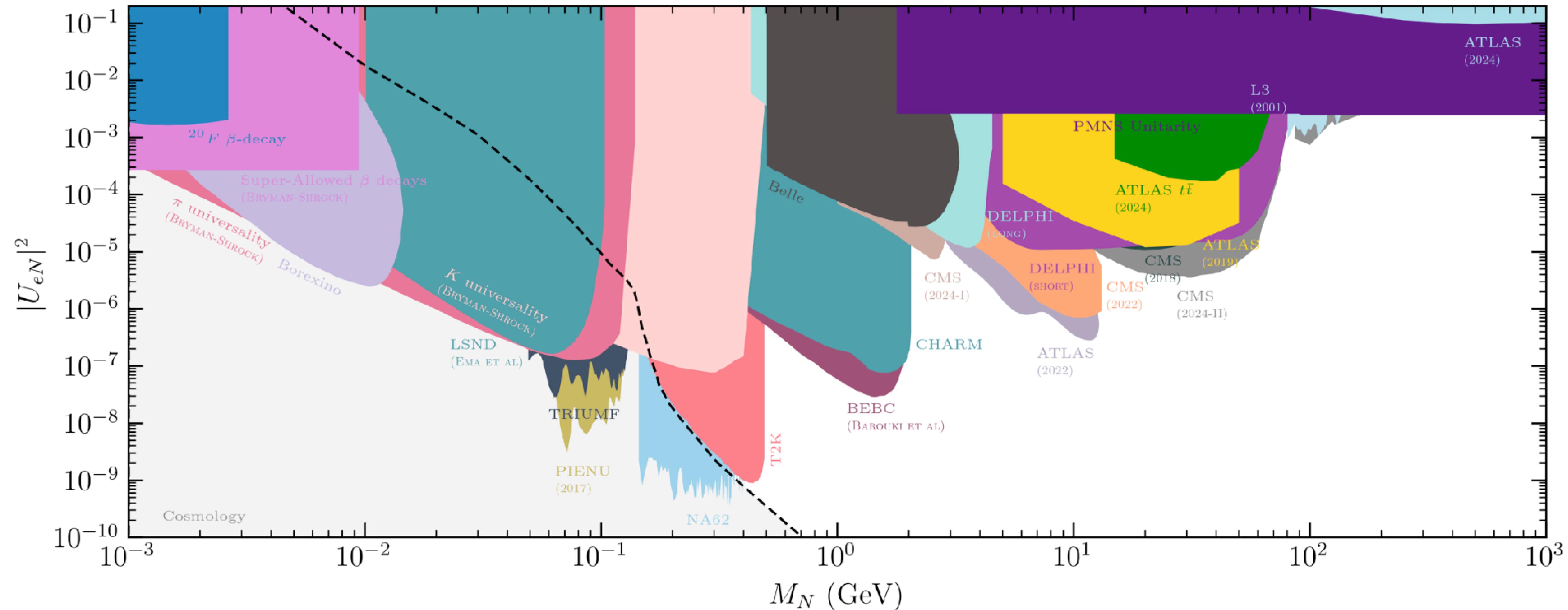


at the highest masses
colliders
production in charm/bottom decay:
SHiP / SPS dominate

below kaon threshold:
neutrino beams dominate
(but cosmologically disfavoured)

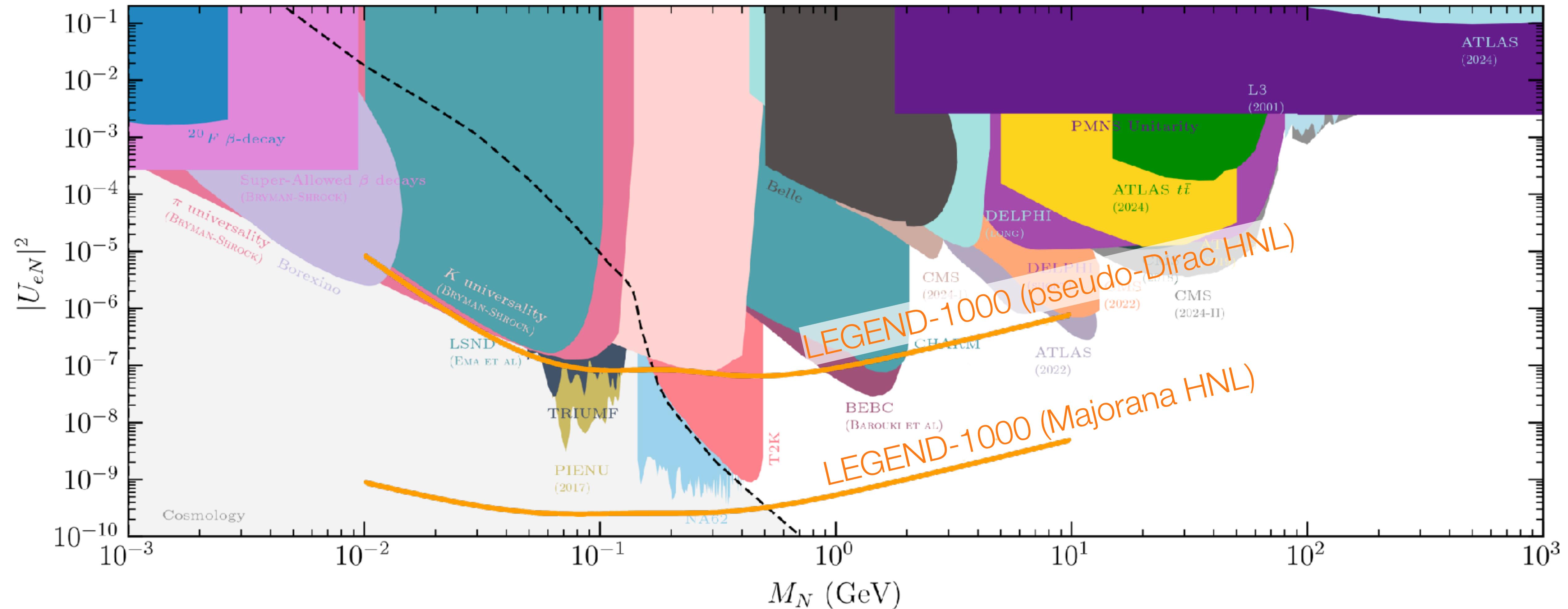
Heavy Neutral Leptons (coupling to electrons)

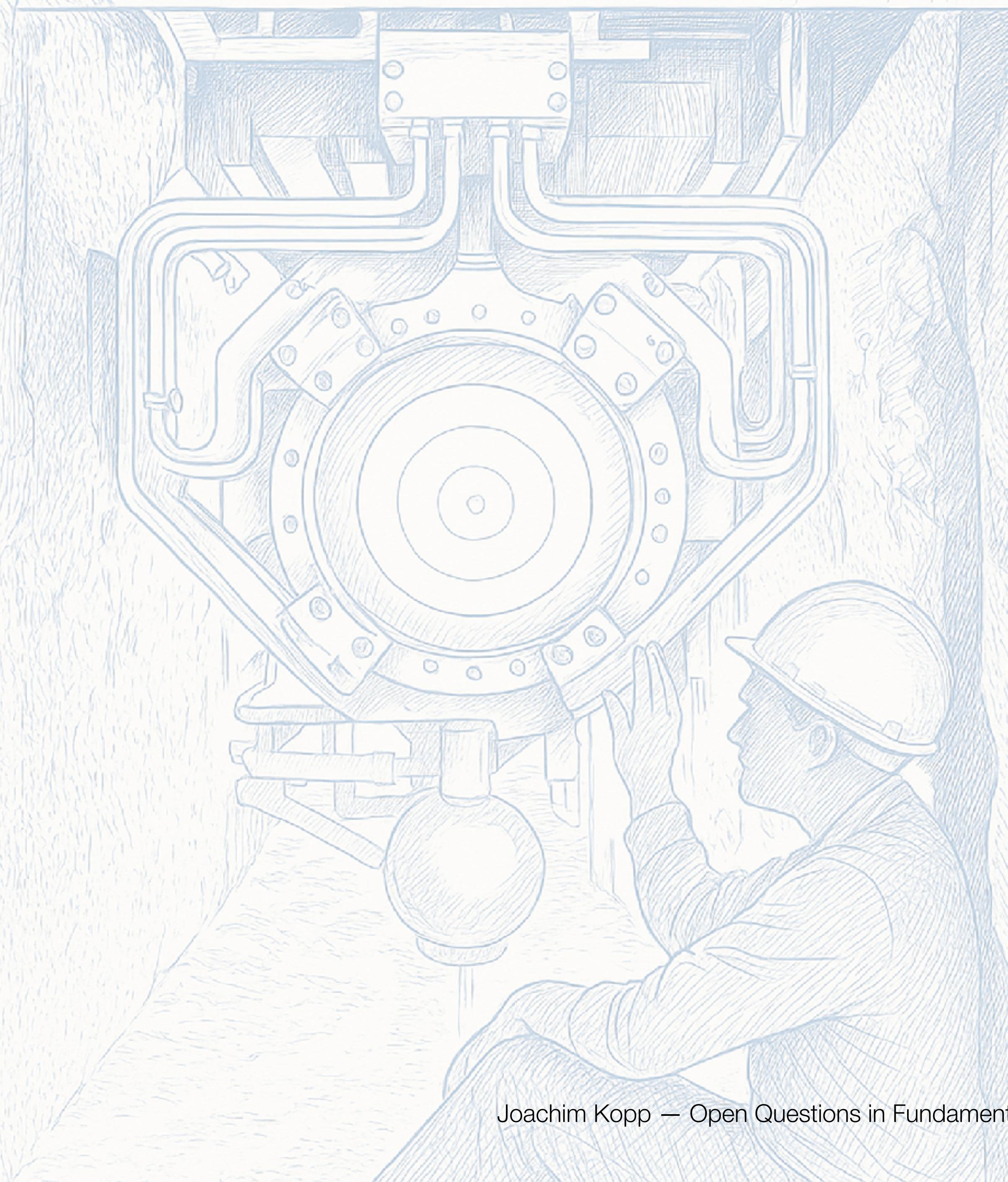
Fernandez-Martinez et al. 2023 / Hostert 2025 (GitHub) / Bolton et al. 2022



Heavy Neutral Leptons (coupling to electrons)

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Dark Sectors



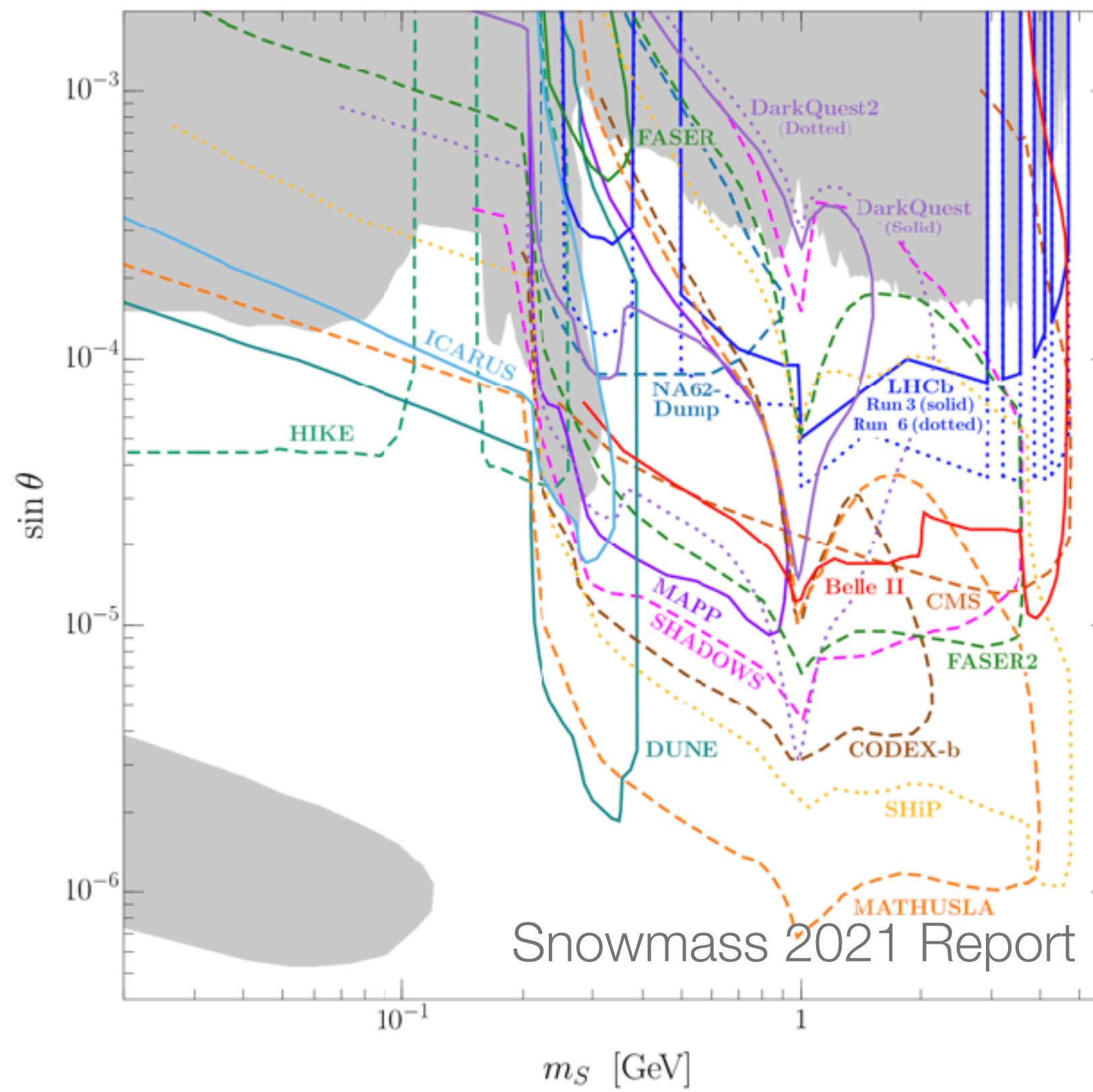
Long-Lived Particles at Neutrino Beams

- **huge luminosities** ➔ great for producing feebly interacting particles
- **large detectors** which excel in
 - event reconstruction
 - timing
 - spatial resolution
 - background suppression
- **main players**
 - DUNE (1.2–2.3 MW beam, 60–120 GeV)
 - T2K / SuperK / HyperK (0.8–1 MW beam, 30 GeV)
 - Fermilab short-baseline programme (0.05 MW, 8 GeV, currently taking data)
 - SPS @ CERN (0.35 MW, 400 GeV)
 - FASER / SND@LHC / FPF (\sim TeV neutrinos)
 - CEvNS

Decay Signatures: Examples

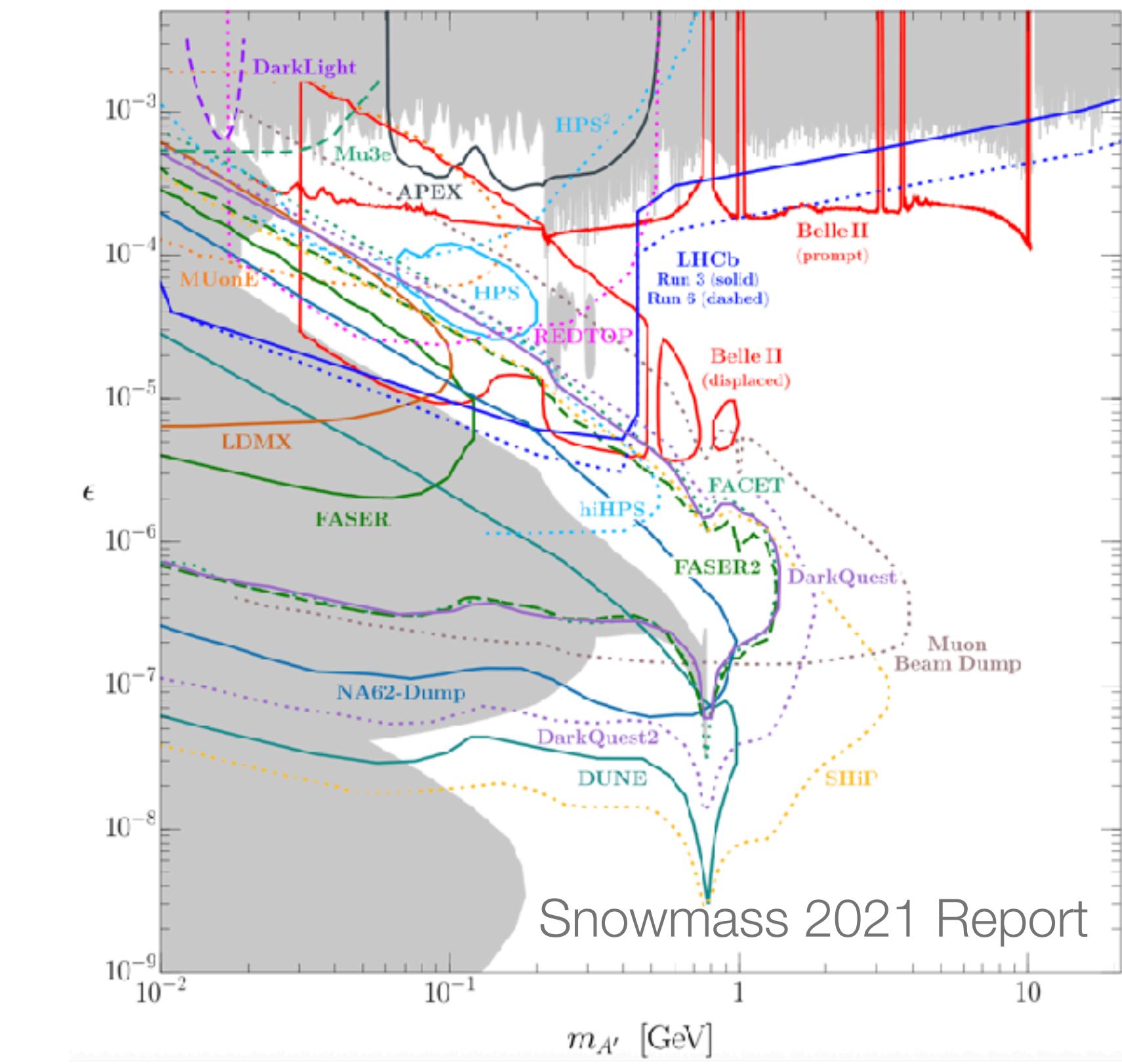
Higgs Portal Scalar

(production via meson decay, bremsstrahlung, Drell–Yann;
decay to SM pairs)



Dark Photons

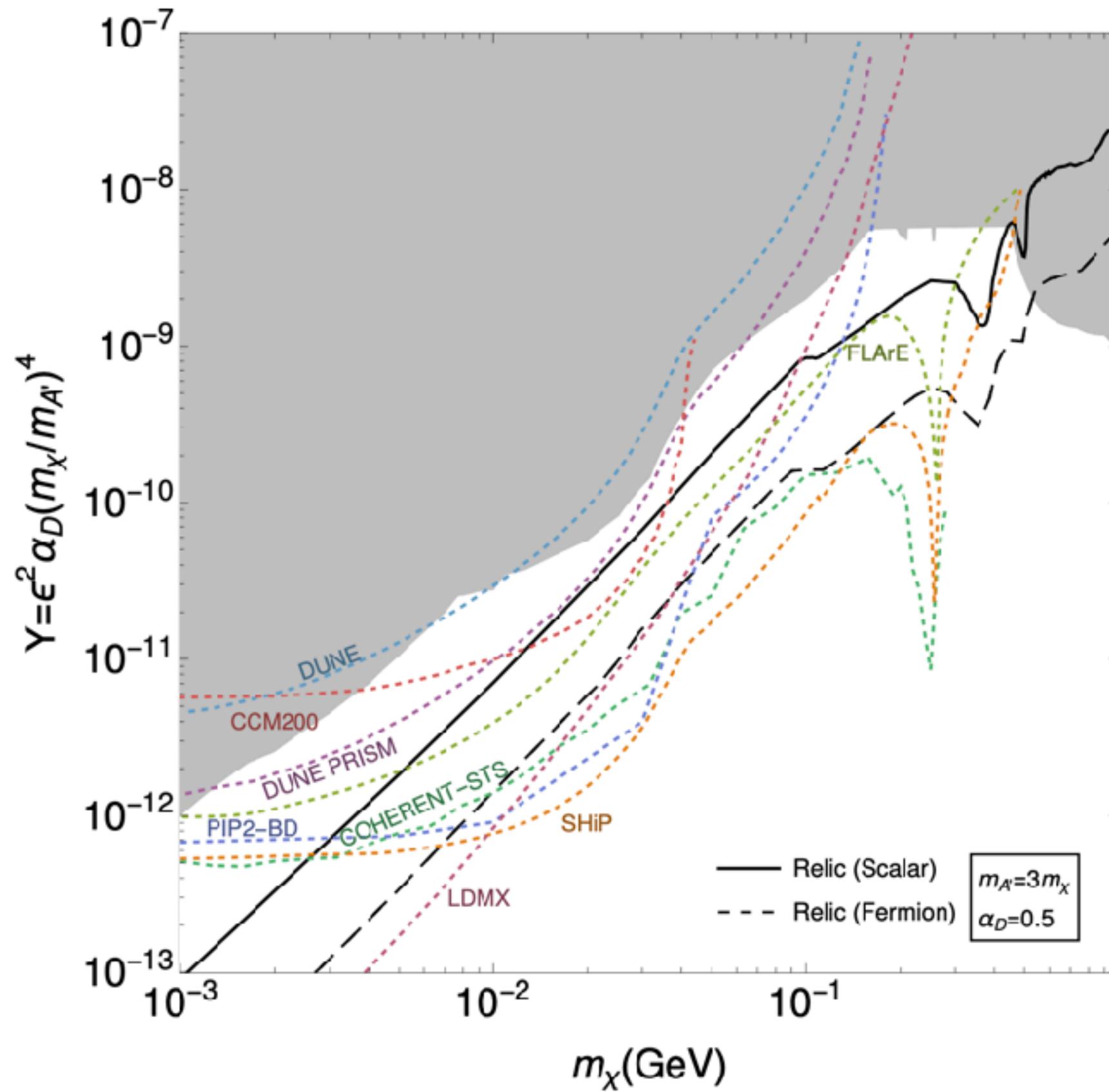
(production via meson decay, bremsstrahlung, Drell–Yann;
detection via decay, tridents)



Scattering Signatures: Examples

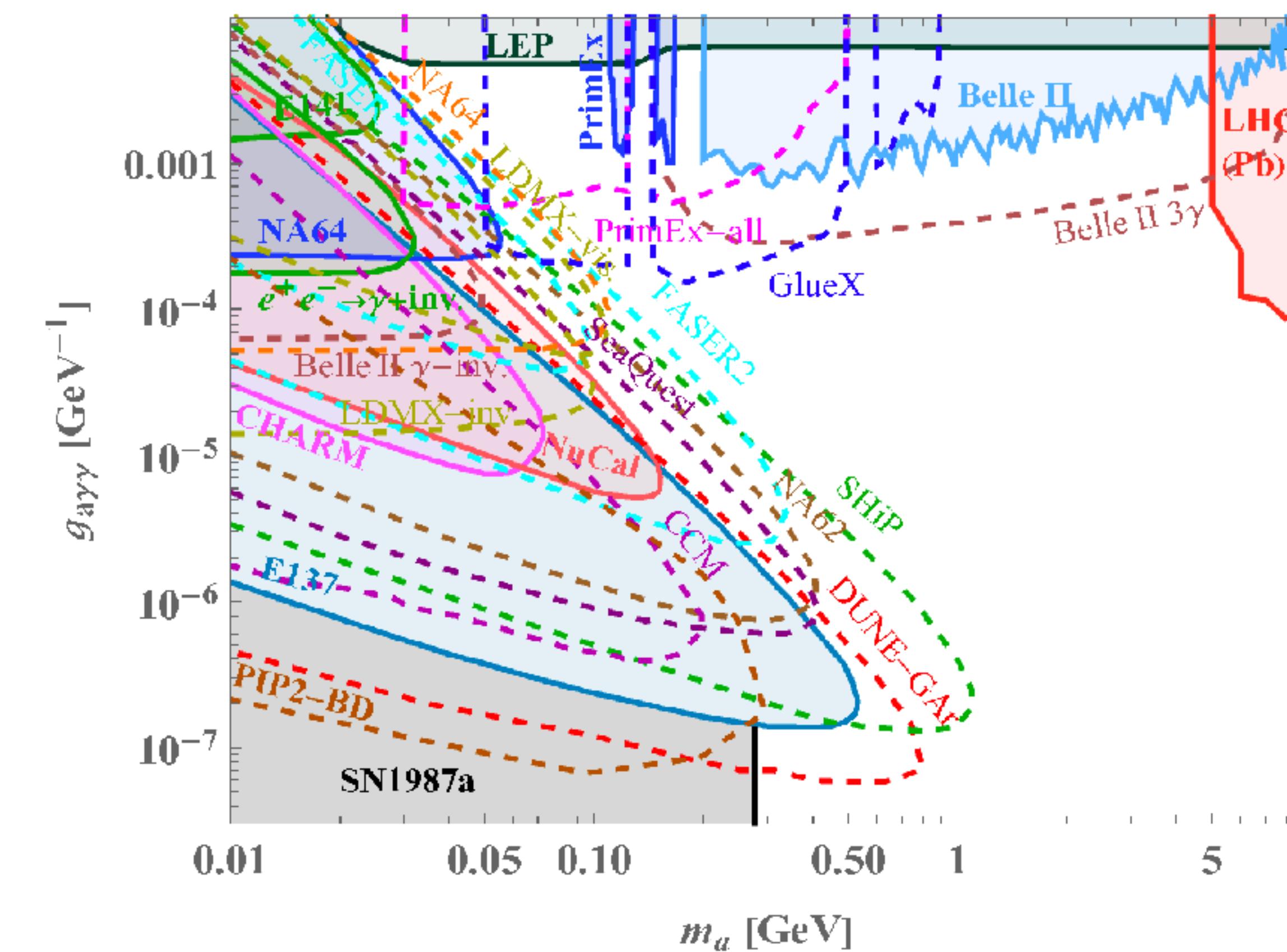
Vector Portal DM

(production via A' decay in the target
detection via scattering)



Axion-Like Particles

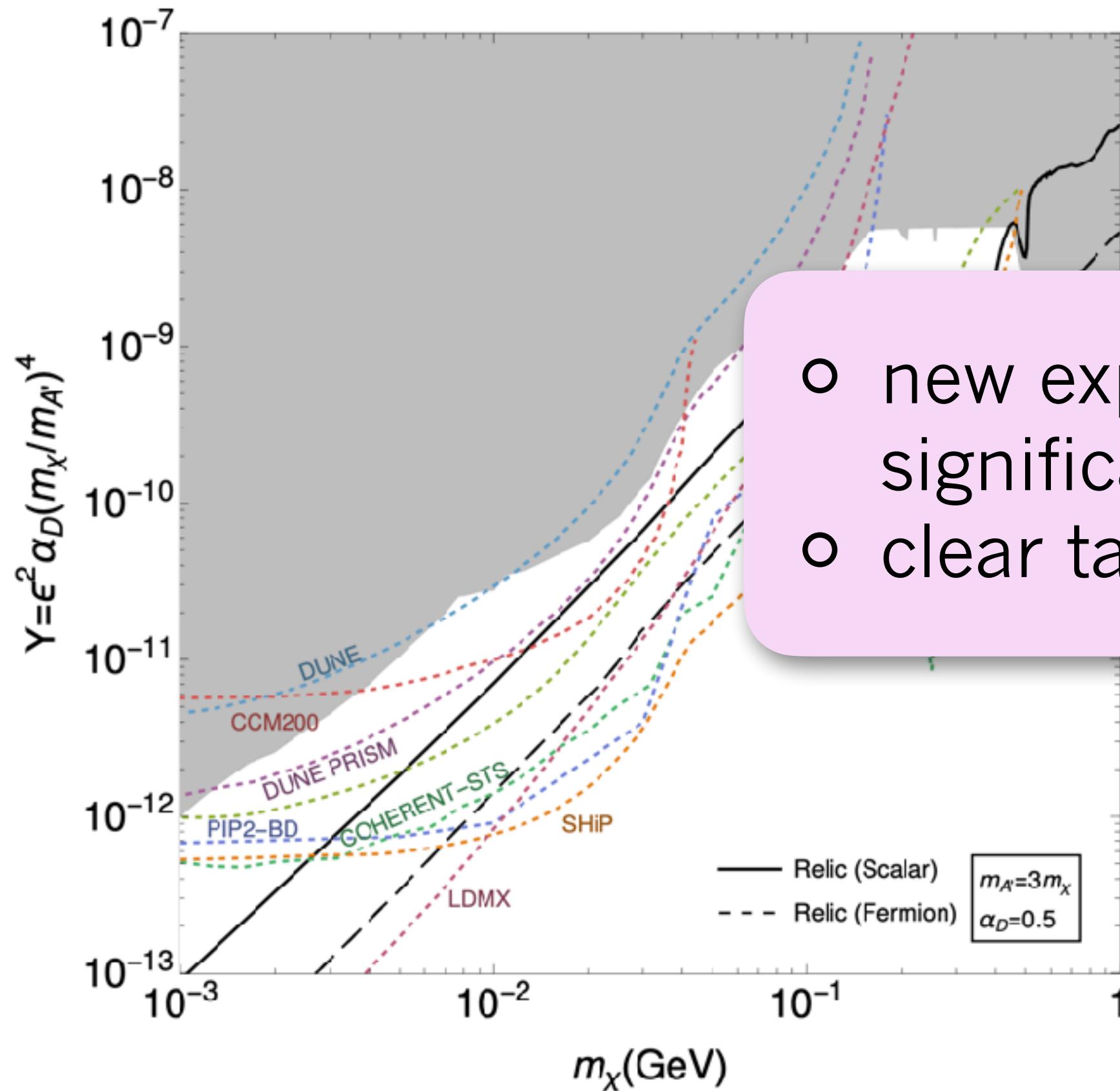
(production via Primakoff, Compton, nuclear reactions;
detection via decay, scattering)



Scattering Signatures: Examples

Vector Portal DM

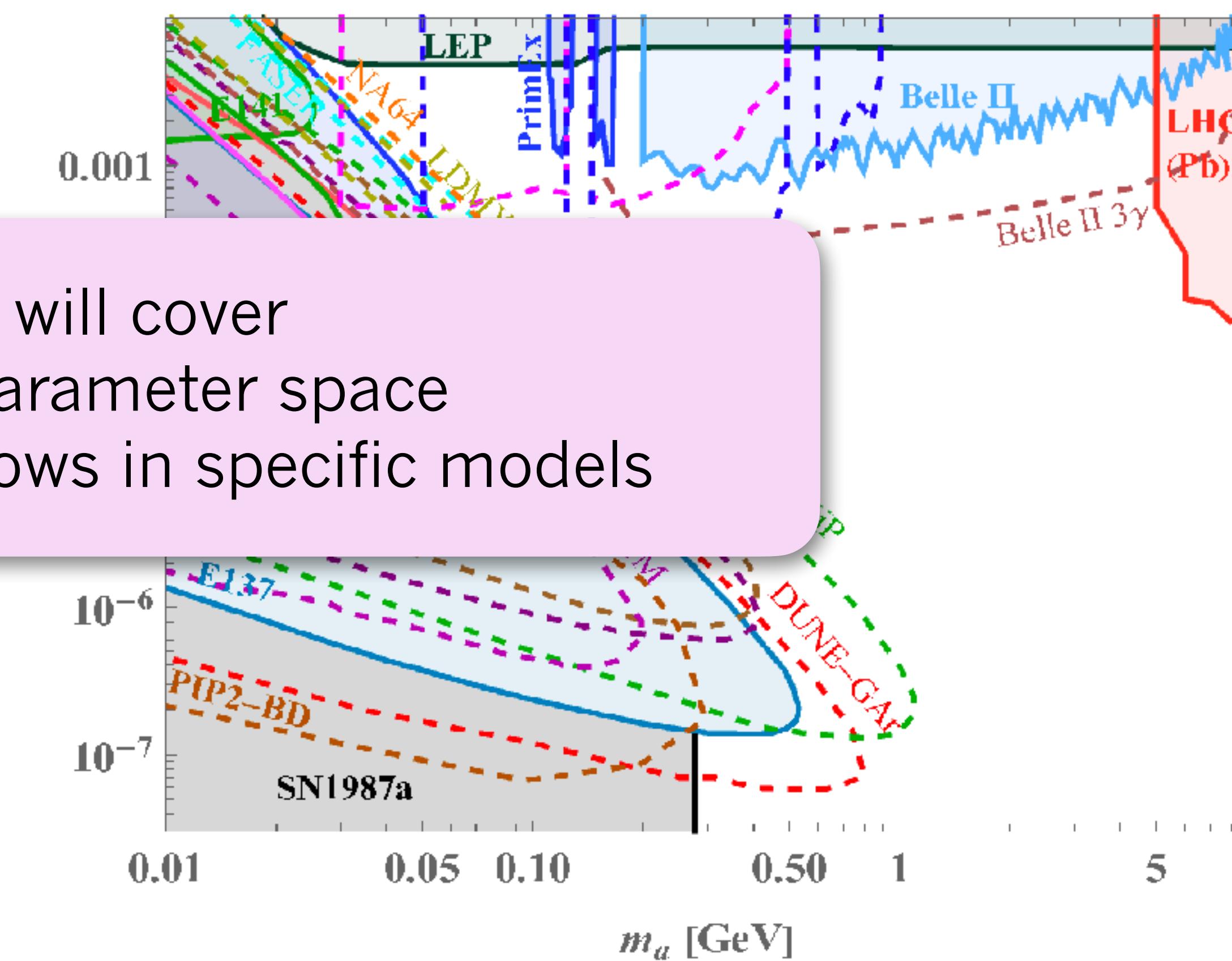
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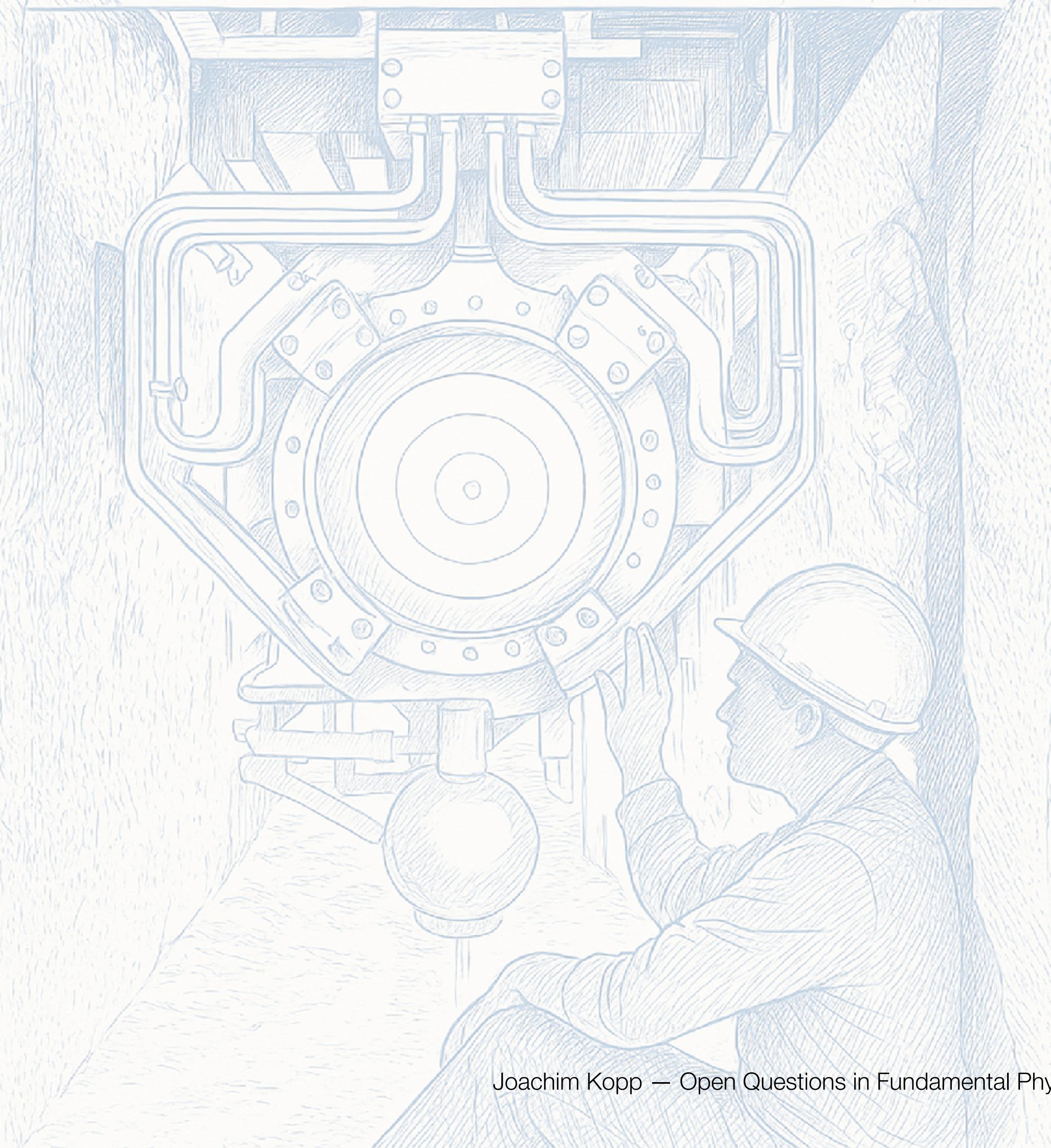
- new experiments will cover significant new parameter space
- clear target windows in specific models

Axion-Like Particles

(production via Primakoff, Compton, nuclear reactions;
detection via decay, scattering)



Summary



Summary

Sterile Neutrinos

- very generic (\rightarrow part of dark sector)
- target windows in specific models
- only detectable in neutrino experiments via oscillation or decay (“HNLs”)

Effective Field Theories

- highly competitive constraints on some (flavourful) operators
- model-dependent constraints more challenging

Dark Sectors

- MW-scale neutrino beams uniquely positioned
- cover broad range of motivated models

