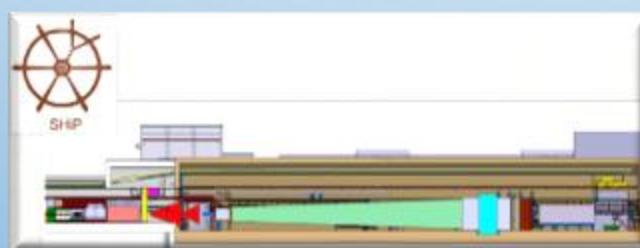
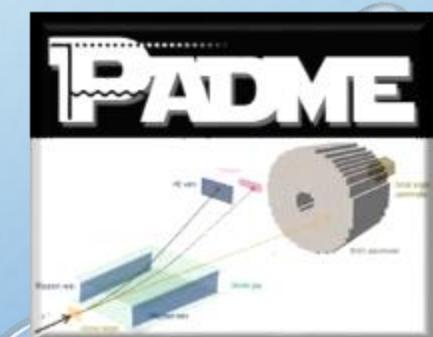
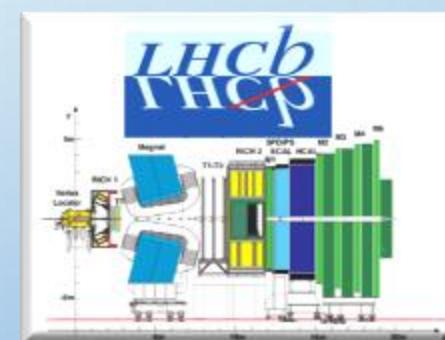
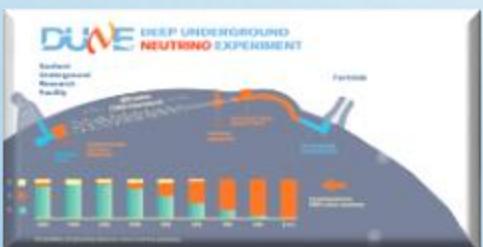


RIUNIONE DI GRUPPO 1

SEZIONE DI NAPOLI

16/01/2024



Riunione Gruppo 1 - Napoli

Thursday 16 Jan 2025, 09:00 → 19:00 Europe/Rome

2026

Francesco Alessandro Conventi (Istituto Nazionale di Fisica Nucleare)

Description La riunione di Gruppo 1 si terrà in aula 2026.

Link zoom: <https://cern.zoom.us/j/68194851871?pwd=AbJcvBqjYDRig6QH0HaSp2o7FFa95c1>

Indico page

09:00 → 09:30	DUNE	⌚ 30m	[X]
	Speakers: Francesco Di Capua (Università Federico II di Napoli), Francesco Di Capua (Istituto Nazionale di Fisica Nucleare)		
	DUNE_Gr1-Napoli...		
09:30 → 10:00	HYPER_K	⌚ 30m	[X]
	Speaker: Gianfranca De Rosa (Istituto Nazionale di Fisica Nucleare)		
	HyperK-Gr1-160120...		
10:00 → 10:30	SND@LHC/SHIP	⌚ 30m	[X]
	Speaker: Antonia Di Crescenzo (Istituto Nazionale di Fisica Nucleare)		
	2025_01_15_Neutri...		
10:30 → 11:20	Lightning talks (I)		[X]
10:30	Search for neutrino interactions in the emulsion target	⌚ 8m	[X]
	Speaker: Fabio Alicante		
	20250116_FA.pdf		
10:40	Design of the magnet for the detector upgrade	⌚ 8m	[X]
	Speaker: Daniele Centanni (Istituto Nazionale di Fisica Nucleare)		
	2025_01_16_DCent...		
10:50	Neutrino astronomy for multi-messenger studies in Super-Kamiokande	⌚ 8m	[X]
	Speaker: Aurora Langella (Istituto Nazionale di Fisica Nucleare)		
	Presentazione_Gr1...		
11:00	HK mPMT electronics overlook	⌚ 8m	[X]
	Speaker: Alessandro Di Nola (Istituto Nazionale di Fisica Nucleare)		
	alessandro_electro...		
11:20 → 11:40	Coffe break	⌚ 20m	
11:40 → 12:10	Comunicazioni del coordinatore	⌚ 30m	[X]
	Speaker: Francesco Alessandro Conventi (Istituto Nazionale di Fisica Nucleare)		
	Scheda_riunione_CS...		
12:10 → 12:40	CMS	⌚ 30m	[X]
	Speakers: Alberto Orso Maria Iorio (Istituto Nazionale di Fisica Nucleare), Orso Maria Iorio		
	CMS_Status_2024-5...		
12:40 → 14:00	Lunch Break	⌚ 1h 20m	

14:00 → 14:30	ATLAS	⌚ 30m	[X]
	Speaker: Francesco Cirotto (Istituto Nazionale di Fisica Nucleare)		
14:30 → 15:00	FCC	⌚ 30m	[X]
	Speaker: Marcello Campajola (Istituto Nazionale di Fisica Nucleare)		
15:00 → 15:30	Lightning talks (II)	⌚ 8m	[X]
15:00	Muon system: Run 2 studies and GEM MEO Upgrade at CMS	⌚ 8m	[X]
	Speaker: Leonardo Favilla		
	LFavilla_GR1_1601...		
15:00	Small Pad MicroMegas R&D for Future Colliders	⌚ 7m	[X]
	Speaker: Simone Perna		
	Lightning talk_Simo...		
15:17	Highlights dal Test Beam per il calorimetro elettromagnetico di IDEA al CERN	⌚ 7m	[X]
	Speaker: Lucrezia Borriello (Istituto Nazionale di Fisica Nucleare sezione Napoli)		
	Lightning talk_Lucr...		
15:24	Attività di simulazione per FCC	⌚ 6m	[X]
	Speaker: Antonio D'Avanzo (Istituto Nazionale di Fisica Nucleare)		
	FCC_simulation_Gr...		
15:30 → 16:00	NA62	⌚ 30m	[X]
	Speaker: Marco Mirra (Istituto Nazionale di Fisica Nucleare)		
16:00 → 16:30	BELLE2	⌚ 30m	[X]
	Speaker: Guglielmo De Nardo (Istituto Nazionale di Fisica Nucleare)		
16:30 → 16:50	GMINUS2	⌚ 20m	[X]
	Speaker: Michele Iacobacci (Istituto Nazionale di Fisica Nucleare)		
	Mi Stato di g-2 al Fe...		
16:50 → 17:05	PADME	⌚ 15m	[X]
	Speaker: Paolo Iengo (Istituto Nazionale di Fisica Nucleare)		
17:05 → 18:20	Lightning talks (III)		[X]
17:05	Validation of DUNE Photon Detection System with ProtoDUNE data	⌚ 8m	[X]
	Speaker: Gabriel Botagoske		
	250116-rv-Botagoske...		
17:15	Misura di $B \rightarrow \tau \nu \mu \bar{\nu}$ a Belle II	⌚ 10m	[X]
	Speaker: Giovanni Gaudino (Istituto Nazionale di Fisica Nucleare)		
	250116-rv-Gaudino...		
17:25	Search for $K^+ \rightarrow \mu^+ \nu A'$, $A' \rightarrow e^+ e^-$	⌚ 10m	[X]
	Speaker: Ilaria Rosa (Istituto Nazionale di Fisica Nucleare)		
	KmarusA_Gruppo1...		
17:35	Misura BR $K \rightarrow \pi^+ \pi^- \pi^+ \pi^-$ in NA62	⌚ 10m	[X]
	Speaker: Renato Fiorenza (INFN Napoli)		
	Fiorenza_2025-01-1...		

















COMUNICAZIONI DEL COORDINATORE

Pao lengo Adelina D'Onofrio Leonardo Merola vincenzo izzo Bernardino Spisso Raffaele

CSN1-INFN 2025

- **Riunioni del 2025:**
 - 19-20 Febbraio Presidenza INFN
 - 7-9 Maggio Pisa
 - 14-16 Luglio *da definire*
 - 8-12 Settembre Otranto (LE)
 - 24-25 Novembre *da definire*

 Istituto Nazionale di Fisica Nucleare

Search
AD Adelina D'Onofrio
AV Anna Vanacore
BS Bernardino Spisso
FC Francesco Cirotto
g giulio
IR Ilaria Rosa
LM Leonardo Merola
mm marco mirra
Ol Orso Iorio
Pl Paolo lengo
R Raffaele
SL Salvatore Loffredo
SM Stefano Mastroianni
vi vincenzo izzo

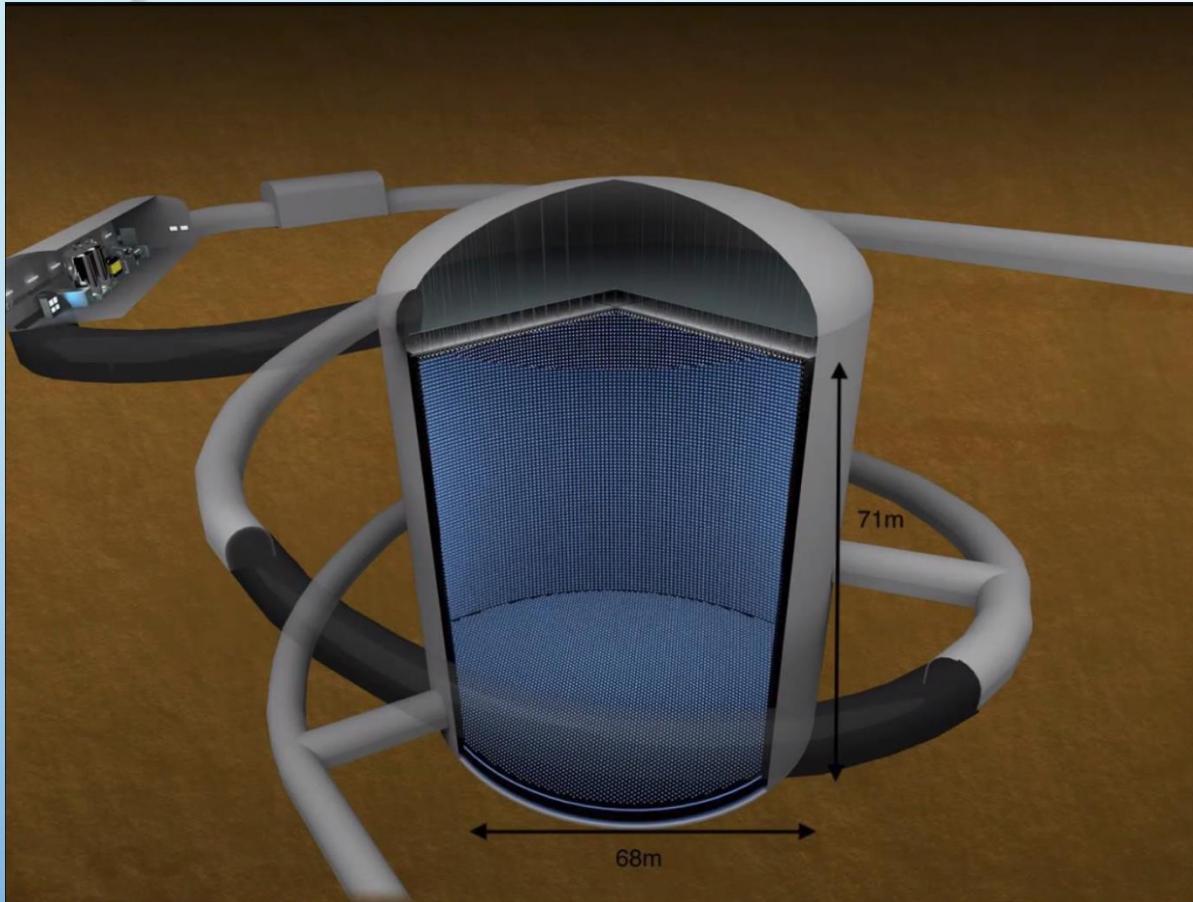
DUNE activities in Naples

Main involvement in the Photon Detection System

- Characterization of the PDS module (Megacell)
- PDE of SiPMs at cryogenic temperatures
- European site for evaporation of the wavelength shifter
- Active participation to ProtoDUNE (analysis and simulation)
- Proposal for a new design for Far Detector 3
 - Simulation and performance study
 - In situ R&D small prototypes

ER	Elvira Rossi	+	mic	video
A	Adele	mic	video	
AL	Aurora Langella	mic	video	
CB	Cristiano Bozza	mic	video	
FC	Francesco Cirotto	mic	video	
IR	Ilaria Rosa	mic	video	
mm	marco mirra	mic	video	
PI	Paolo Iengo	mic	video	
SM	Stefano Mastroianni	mic	video	
VT	Valeri Tioukov	mic	video	

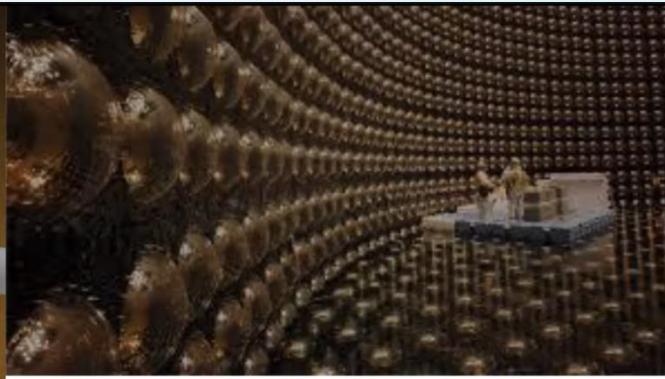
HYPER-K T2K SUPER-K



 **Hyper-Kamiokande**

 **T2K**

 **SUPER-K**



Hyper-K
T2K
Super-K

Gianfranca De Rosa
Per il gruppo Hyper-K di Napoli

ER Elvira Rossi



A Adele



AV Anna Vanacore



AL Aurora Langella



CB Cristiano Bozza



FC Francesco Cirotto



g giulio



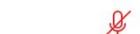
IR Ilaria Rosa



mm marco mirra



PI Paolo Iengo



SM Stefano Mastroianni

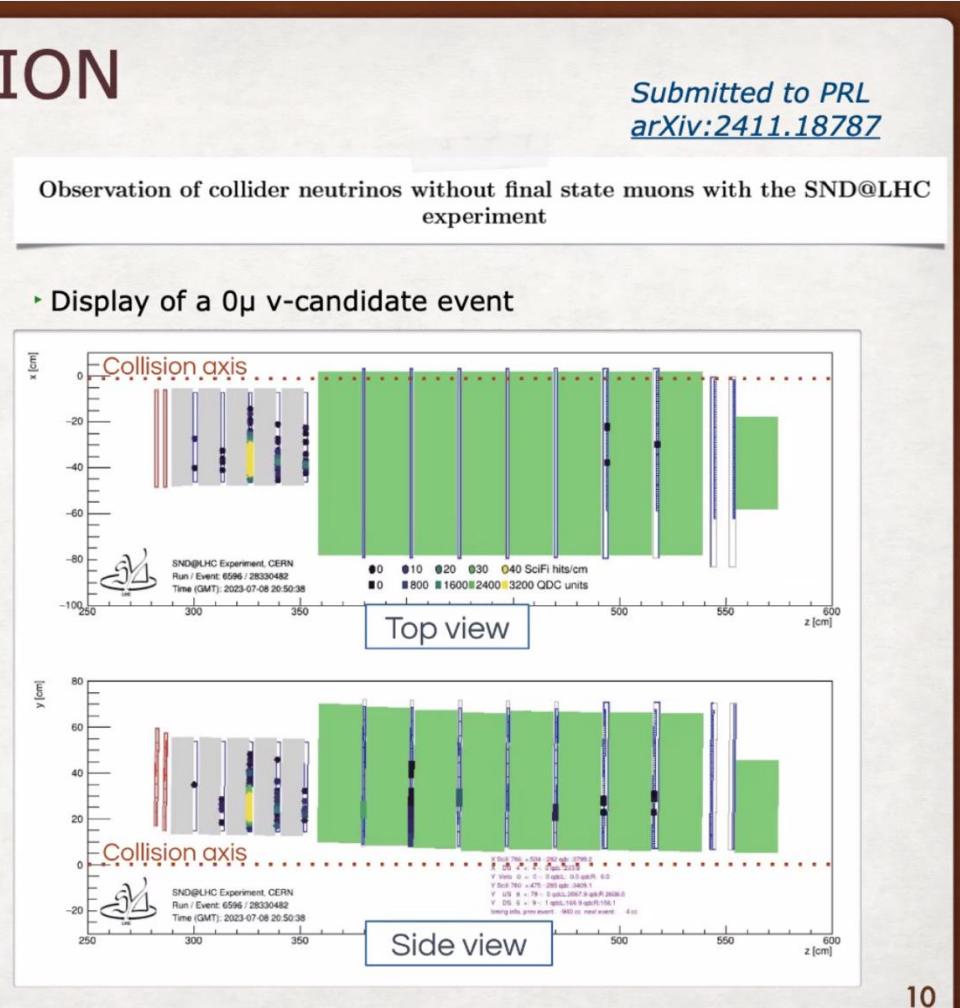
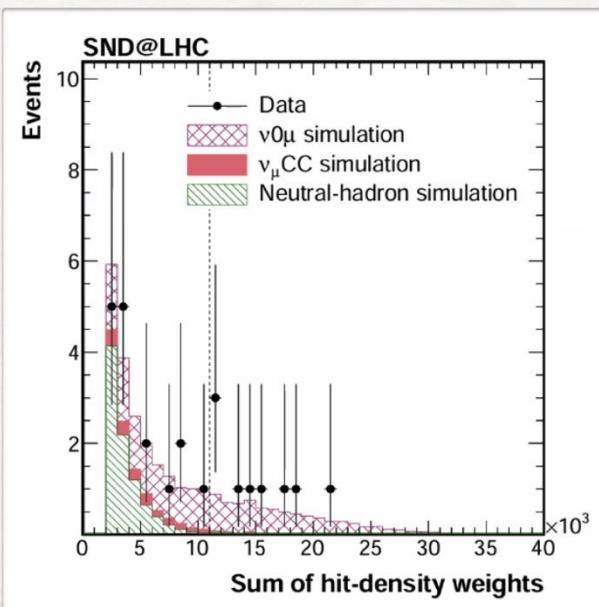


VT Valeri Tioukov



0 μ NEUTRINO OBSERVATION

- First observation of neutrino interactions without a muon in the final state based on **2022-2023** data
- 9** observed **0 μ v-candidates**
- Observation significance **6.4 σ**
- Evidence for **v_e interactions** at 6.4 σ



10

ER	Elvira Rossi		
AD	Adelina D'Onofrio		
AL	Aurora Langella		
CB	Cristiano Bozza		
FC	Francesco Cirotto		
g	giulio		
IR	Ilaria Rosa		
mm	marco mirra		
Ol	Orso Iorio		
Pl	Paolo Iengo		
SM	Stefano Mastroianni		
VT	Valeri Tioukov		
vi	vincenzo izzo		

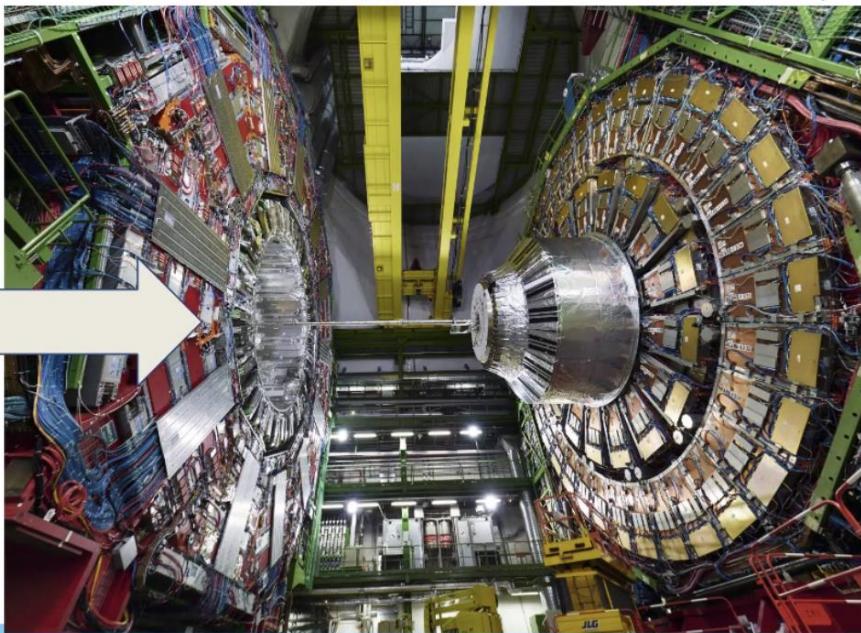
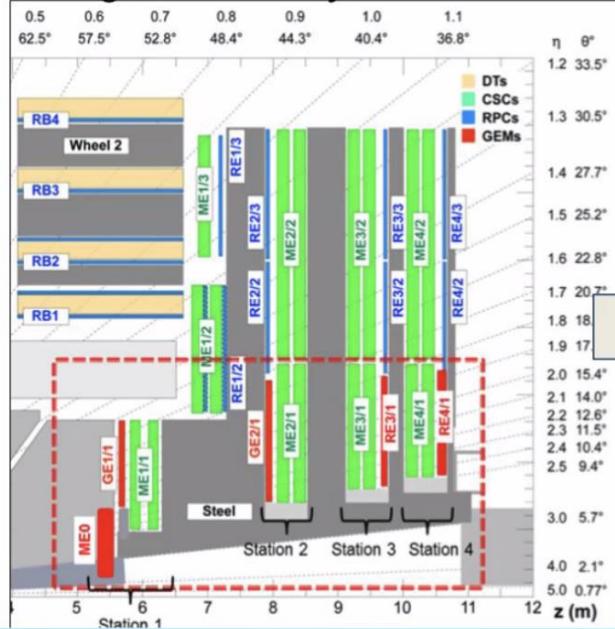
Looking forward

Phase 2 upgrade:

- additional low angle, high-rate muon detectors
→ GEM, iRPC

And also:

- Additional silicon tracker layers, trigger input
- High-Granularity Calorimeter



Alberto Orso Maria Iorio

ER Elvira Rossi



vi vincenzo izzo



AD Adelina D'Onofrio



BS Bernardino Spisso



FC Francesco Cirotto



g giulio



LM Leonardo Merola



mm marco mirra



OI Orso Iorio



PI Paolo Iengo



SL Salvatore Loffredo



VD Vincenzo Del Piano

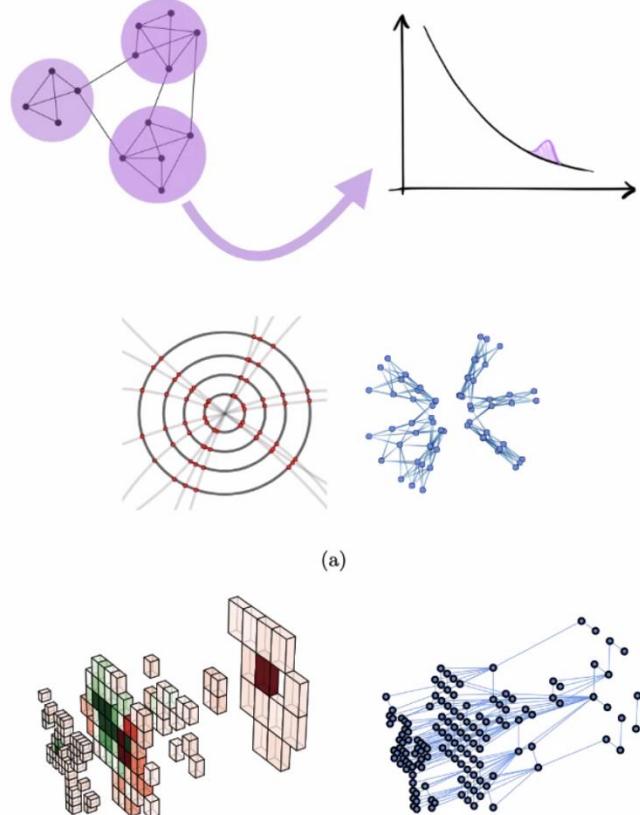


R Raffaele



GRAPH ANOMALY DETECTION FOR NEW PHYSICS SEARCHES

- Graph-structured data are ubiquitous across science, engineering, and many other domains
 - ↳ Used to describe and analyze relations and interactions
 - ↳ Can encapsulate object or event information
 - ↳ Can be employed in particle physics!
- Our strategy: **to represent jets as graphs and then apply machine learning to build an anomaly detection algorithm**
 - ↳ Targeting heavy resonance searches with hadronic final states in Run-3
 - ↳ Exploit event-based graphs to detect anomalies
- Jet information can be used as input features for neural network architectures.
 - ↳ A significant improvement in performances can be achieved by employing a set of features with basic information (**low-level**) such as information coming directly from the detectors.
 - ↳ Jet constituents represent challenging input features to achieve this goal

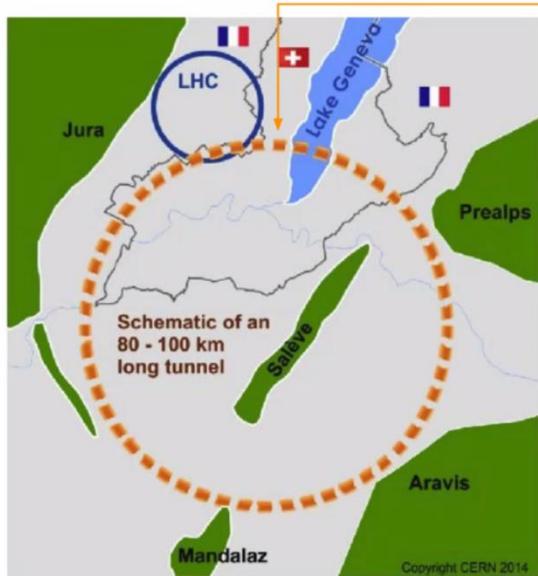


	Elvira Rossi		
	Marco Napolitano		
	Adelina D'Onofrio		
	Anna Vanacore		
	Francesco Di Capua		
	giulio		
	Ilaria Rosa		
	Leonardo Merola		
	marco mirra		
	Paolo Iengo		
	Raffaele		
	Stefano Mastroianni		
	vincenzo izzo		

The FCC integrated program

The 2020 European Strategy concluded that an **e+e− Higgs factory is the highest priority next collider.**

→ In 2021 CERN has launched the international **Future Circular Collider (FCC)** feasibility Study



FCC is a comprehensive long-term programme maximising physics opportunities:

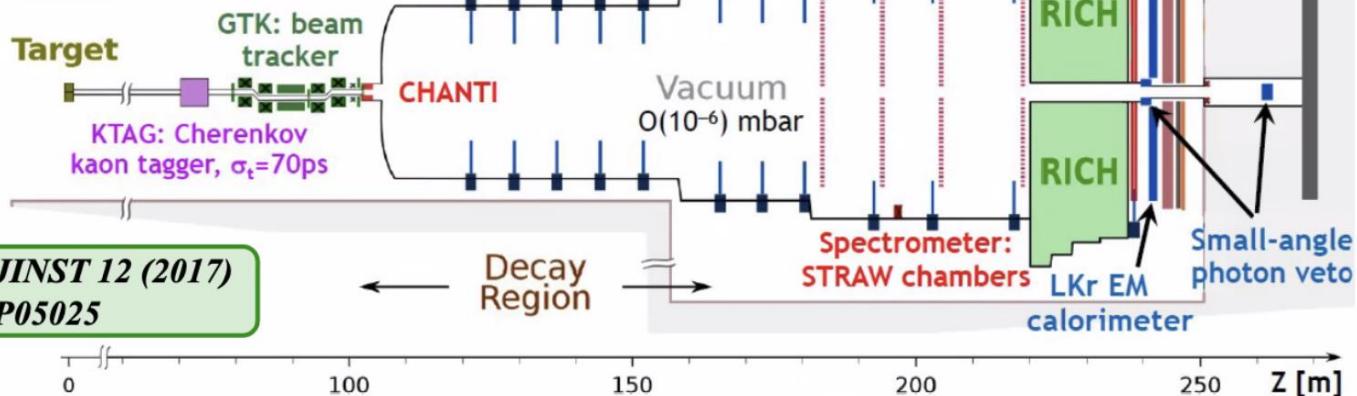
- **stage 1 (FCC-ee): an Higgs factory, electroweak and top factory at highest luminosities**
- **stage 2 (FCC-hh):** as natural continuation at energy frontier (~ 100 TeV)

Both sharing same technical infrastructures



Beamline and detector

SPS protons: 400 GeV, 3×10^{12} /spill.
 Unseparated secondary ($p/\pi^+/K^+$) beam:
 75 GeV/c ($\pm 1\%$), divergence $< 100 \mu\text{rad}$.
 Beam rate: 600 MHz; K^+ rate: 35 MHz;
 K^+ decays in fiducial volume: 4 MHz

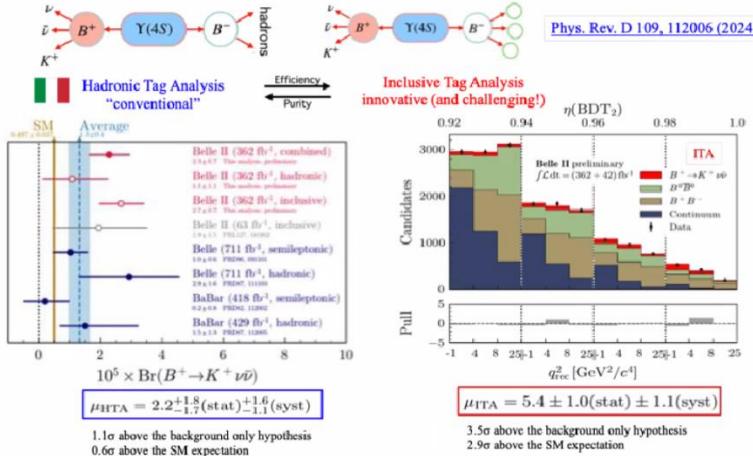


- ✓ Excellent time resolution: $\mathcal{O}(100 \text{ ps})$ to match beam and daughter particle information
- ✓ Kinematics: rejection of main K modes 10^{-4} via kinematics reconstruction
- ✓ PID capability (RICH+LKr+MUV): $\mathcal{O}(10^{-7})$ muon suppression for $15 \text{ GeV} < p < 35 \text{ GeV}$
- ✓ High-efficiency photon veto: 10^{-7} rejection of π^0 for $E(\pi^0) > 40 \text{ GeV}$

ER	Elvira Rossi	<input checked="" type="checkbox"/>	<input type="checkbox"/>
MN	Marco Napolitano	<input type="checkbox"/>	<input checked="" type="checkbox"/>
BS	Bernardino Spisso	<input checked="" type="checkbox"/>	<input type="checkbox"/>
FD	Francesco Di Capua	<input checked="" type="checkbox"/>	<input type="checkbox"/>
g	giulio	<input checked="" type="checkbox"/>	<input type="checkbox"/>
IR	Ilaria Rosa	<input checked="" type="checkbox"/>	<input type="checkbox"/>
PI	Paolo Iengo	<input type="checkbox"/>	<input checked="" type="checkbox"/>
SL	Salvatore Loffredo	<input checked="" type="checkbox"/>	<input type="checkbox"/>
SM	Stefano Mastroianni	<input checked="" type="checkbox"/>	<input type="checkbox"/>
vi	Vincenzo Izzo	<input checked="" type="checkbox"/>	<input type="checkbox"/>

BELLE2

Evidence for $B \rightarrow K \nu \bar{\nu}$

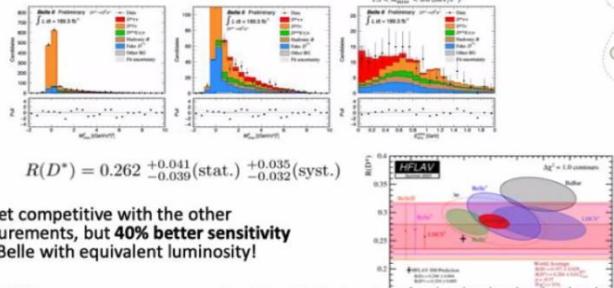


R(D^*) measurement at Belle II

- Long-standing $\sim 3\sigma$ anomaly in Flavor Physics:

$$R(D^*) = \frac{B(\overline{B} \rightarrow D^{(*)} \tau^- \bar{\nu}_\tau)}{B(\overline{B} \rightarrow D^{(*)} \ell^- \bar{\nu}_\ell)}$$

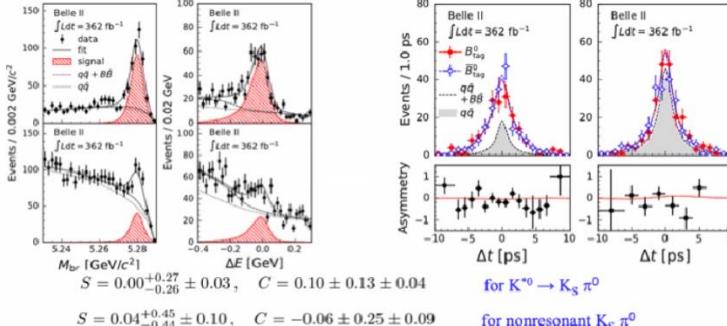
- First Belle II measurement (using only \sim half of the Run1 data):



Time dependent $B^0 \rightarrow K_S \pi^0 \gamma$

- The SM predict that the photon is $\sim 100\%$ right(left)-polarized for B^0 (B^0) decays;

- Any significant CP violation would be a sign of new physics;



World leading precision!

Search for LFV in $\tau \rightarrow \mu \mu \mu$ decays

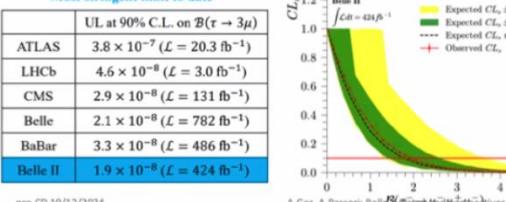
- Belle II will dominate the sensitivity for Lepton Flavor Violation searches in the coming years;

- Competition with hadron machines only on few channels, most notably $\tau \rightarrow 3\mu$;

- Compared to Belle/BaBar: more inclusive analysis and better control of the backgrounds;

- World leading upper limit!

Most stringent limit to date



30 7

ER Elvira Rossi

MN Marco Napolitano

BS Bernardino Spisso

FD Francesco Di Capua

g giulio

IR Ilaria Rosa

PI Paolo Iengo

SL Salvatore Loffredo

SM Stefano Mastroianni

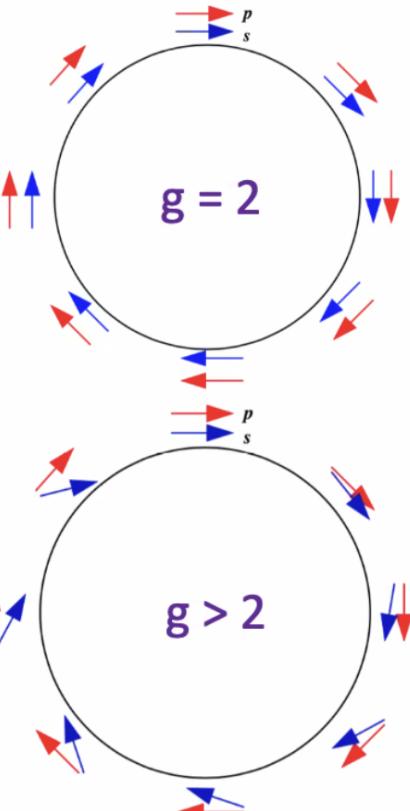
vi vincenzo izzo

Perché un anello di accumulazione con B uniforme?

- Il muone con momento \mathbf{p} (freccia rossa) e spin \mathbf{s} (freccia blu) subisce la rivoluzione del ciclotrone e la precessione di spin in un campo magnetico B (si assume che \mathbf{s} e \mathbf{p} siano perpendicolari a \mathbf{B})
- La differenza tra le velocità angolari della precessione di spin ω_s e rivoluzione ciclotrone ω_c è

$$\vec{\omega}_a = \vec{\omega}_s - \vec{\omega}_c = - \left(\frac{g_\mu - 2}{2} \right) \frac{q\vec{B}}{m} = -a_\mu \frac{q\vec{B}}{m}$$

- Se $g = 2$ esattamente, i vettori di spin e momento rimarrebbero bloccati insieme $\rightarrow \omega_a = 0$
- Invece lo spin precede più velocemente e, alle condizioni del Fermilab, fa una rotazione aggiuntiva di 12° per giro. Dopo circa 30 giri si riallinea all'impulso.



PADME

v1.pdf X Riunione Gruppo 1 - Napoli X Readout PCB drawings sig X PowerPoint Presentation - X Management - Reminders X DRD1 School 2024 Debrief X DRD1School-Followup - DR X + v https://agenda.infn.it/event/44906/contributions/252628/attachments/130296/193854/20250116-PADMEGr1NA.pdf ... Apple iCloud Yahoo Bing Google Wikipedia Facebook Twitter LinkedIn The Weather Channel Yelp TripAdvisor Allianz Portal Management - Paol... >> Other Bookmarks - + 140% <

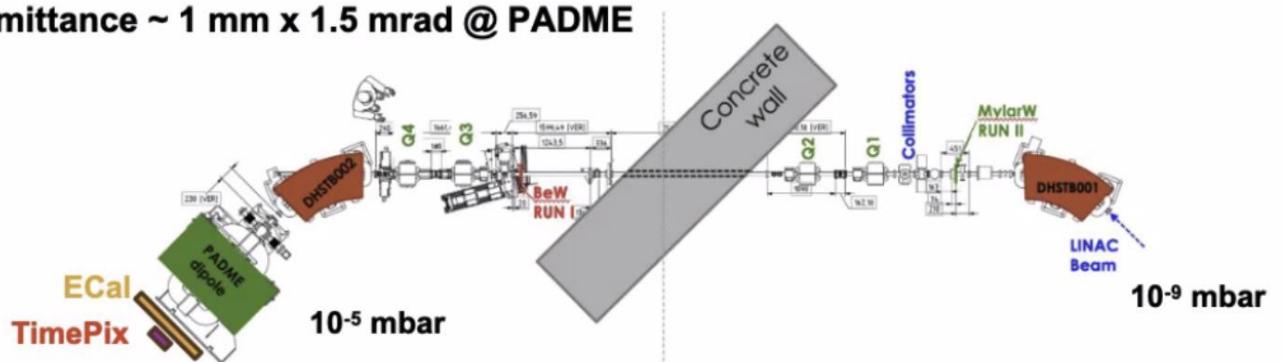
Positron Annihilation into Dark Matter Experiment (PADME)

Positrons from the DAFNE LINAC up to 550 MeV, O(0.5%) energy spread

Repetition rate up to 49 Hz, macro bunches of up to 300 ns duration

Intensity must be limited below $\sim 3 \times 10^4$ POT / spill against pile-up

Emittance ~ 1 mm x 1.5 mrad @ PADME



Past operations:

Run I e⁻ primary, target, e⁺ selection, 250 μm Be vacuum separation [2019]

Run II e⁺ primary beam, 125 μm Mylar™ vacuum separation, 28000 e⁺/bunch [2019-20]

Run III dipole magnet off, ~3000 e⁺/bunch, 47 scan points s^{1/2} ~ 17 MeV [2022]

- | | | | |
|----|---------------------|--|--|
| er | elvira rossi (me) | | |
| PI | Paolo lengo | | |
| ER | Elvira Rossi | | |
| BS | Bernardino Spisso | | |
| FD | Francesco Di Capua | | |
| g | giulio | | |
| IR | Ilaria Rosa | | |
| P | Paolo | | |
| SM | Stefano Mastroianni | | |
| vi | vincenzo izzo | | |

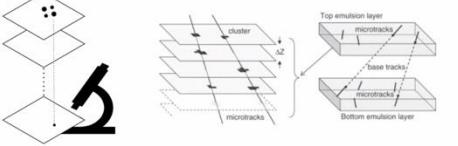
LIGHTNING TALKS

SPEAKER: FABIO ALICANTE

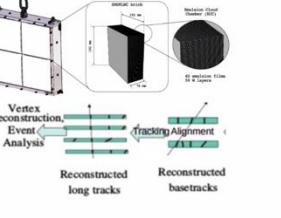
Emulsion Data reconstruction in SND@LHC

1. Emulsion films scanning (online cluster processing).
2. Correct microscope stage effects.
3. Link two layers of each emulsion film.
4. Align consecutive films.
5. Reconstruct tracks and vertices.

Local effects and corrections!



F. Alicante | Riunione Gruppo 1 - Napoli, 16 - 01 - 2025



ER	Elvira Rossi	✓
AD	Adelina D'Onofrio	✗
AL	Aurora Langella	✗
FC	Francesco Cirotto	✗
g	giulio	✗
IR	Ilaria Rosa	✗
mm	marco mirra	✗
Ol	Orso Iorio	✗
Pl	Paolo lengo	✗
SM	Stefano Mastroianni	✗
VT	Valeri Tioukov	✗
vi	vincenzo izzo	✗

ER	Elvira Rossi	✓
AD	Adelina D'Onofrio	✗
AL	Aurora Langella	✗
FC	Francesco Cirotto	✗
FD	Francesco Di Capua	✗
g	giulio	✗
Ol	Orso Iorio	✗
Pl	Paolo lengo	✗
SM	Stefano Mastroianni	✗
VT	Valeri Tioukov	✗
vi	vincenzo izzo	✗

Speaker: Aurora Langella

Astrophysical neutrino sources

Several astrophysical objects are suggested as sources of neutrinos, mainly those which are considered or confirmed to be cosmic rays acceleration sites (hadronic processes). Some of these are:



Core collapse supernovae (CC-SNe). First confirmed transient source of neutrinos (SN1987A)



Supernova remnants (SNRs) with shocks accelerating cosmic rays.



Binary systems involving compact objects producing jets and accreting matter (e.g. BNS mergers, Microquasars).



Pulsar wind nebulae (PWNe) powered by rapidly spinning neutron stars.

Most of these sources are particularly interesting for multi-messengers astrophysics.

16/01/2025

A. Langella

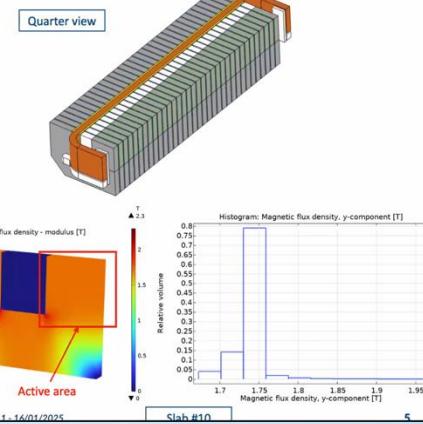
AL	Aurora Langella	✓
ER	Elvira Rossi	✗
AD	Adelina D'Onofrio	✗
FC	Francesco Cirotto	✗
FD	Francesco Di Capua	✗
g	giulio	✗
Ol	Orso Iorio	✗
Pl	Paolo lengo	✗
SM	Stefano Mastroianni	✗
VT	Valeri Tioukov	✗
vi	vincenzo izzo	✗

ER	Elvira Rossi	✓
AD	Adelina D'Onofrio	✗
AL	Aurora Langella	✗
FC	Francesco Cirotto	✗
FD	Francesco Di Capua	✗
g	giulio	✗
Ol	Orso Iorio	✗
Pl	Paolo lengo	✗
SM	Stefano Mastroianni	✗
VT	Valeri Tioukov	✗
vi	vincenzo izzo	✗

Speaker: Daniele Centanni

Electromagnetic design of the magnet

- ✓ Low power consumption
- ✓ Magnetised iron slabs
- ✓ Coil mass compliant with constraints on the crane maximum load
- ✓ Vertical 1.75 T uniform magnetic field
- ✓ High field homogeneity in the active region



ER	Elvira Rossi	✓
AD	Adelina D'Onofrio	✗
AL	Aurora Langella	✗
FC	Francesco Cirotto	✗
FD	Francesco Di Capua	✗
g	giulio	✗
Ol	Orso Iorio	✗
Pl	Paolo lengo	✗
SM	Stefano Mastroianni	✗
VT	Valeri Tioukov	✗
vi	vincenzo izzo	✗

Speaker: Alessandro Di Nola

Photodetectors

Hyper-Kamiokande will be instrumented with 20,000 20" photomultipliers

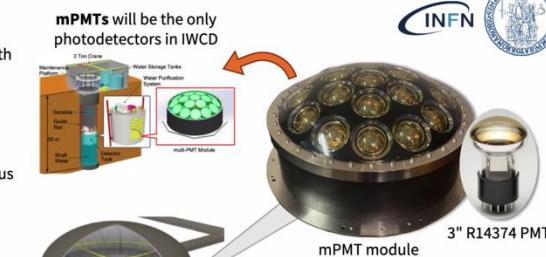
The PMT used are the newly developed Hamamatsu R12860

PMT improvements compared to the previous ones used in SK:

- higher pressure resistance
- double detection efficiency
- half time&charge resolutions



16/1/2025



Meeting INFN Gr. 1 Napoli

In addition to the 20" PMTs, Hyper-K will be equipped with 808 multi-PMT optical modules called mPMTs

Each mPMT is equipped with 19 Hamamatsu R14374 PMTs

ER	Elvira Rossi	✓
AD	Adelina D'Onofrio	✗
AL	Aurora Langella	✗
FC	Francesco Cirotto	✗
FD	Francesco Di Capua	✗
g	giulio	✗
Ol	Orso Iorio	✗
Pl	Paolo lengo	✗
SM	Stefano Mastroianni	✗
VT	Valeri Tioukov	✗
vi	vincenzo izzo	✗

2

Speaker: Leonardo Favilla

The Phase-2 Upgrade of the CMS Muon Detectors

❑ The muon system is challenged by high particle rates in the forward region of the CMS detector

- Inst. Lumi: $2 \times 10^{34} \text{ cm}^{-2} \text{s}^{-1}$ (LHC) $\rightarrow 5 \div 7.5 \times 10^{34} \text{ cm}^{-2} \text{s}^{-1}$ (HL-LHC)
- Pileup: 60 (LHC) $\rightarrow 140 \div 200$ (HL-LHC)
- Several upgrades taking place to handle such a harsh environment \rightarrow [link to CMS-TDR-016](#)

❑ Main purpose:

- $|\eta| < 2.4 \rightarrow$ Enhance identification and triggering capabilities
- $|\eta| > 2.4 \rightarrow$ Extend muon identification and triggering

❑ Three new GEM detectors:

- LS2: **GE1/1** ($1.5 < |\eta| < 2.2$)
- LS3: **GE2/1** ($1.6 < |\eta| < 2.4$) & **ME0** ($2.0 < |\eta| < 2.8$)

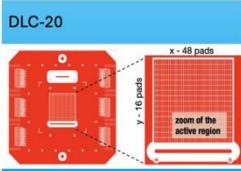
❑ CMS Napoli group involved in:

- GE1/1 operation
- GE2/1 and ME0 construction & commissioning
- High-Voltage Power System design & maintenance

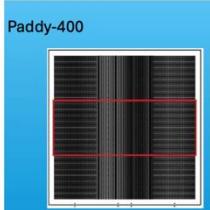


ER	Elvira Rossi
MN	Marco Napolitano
AD	Adelina D'Onofrio
FD	Francesco Di Capua
g	giulio
IR	Ilaria Rosa
PI	Paolo Iengo
SM	Stefano Mastroianni
vi	Vincenzo Izzo

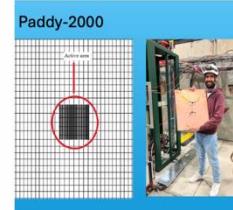
Test-Beam



- 768 pad di dimensioni $1 \times 3 \text{ mm}^2$.
- Gap drift a 7 mm.
- Gap ampl a $128 \mu\text{m}$.



- 4800 pad di dimensioni $1 \times 8 \text{ mm}^2$.
- Gap drift a 5 mm.
- Gap ampl a $128 \mu\text{m}$.



- Area Attiva: $50\text{cm} \times 40\text{cm}$
- Ha pad di dimensioni $1 \times 8 \text{ mm}^2$ nella zona centrale e $10 \times 10 \text{ mm}^2$ nella zona periferica.
- Gap drift di 6 mm.
- Gap ampl di $150 \mu\text{m}$.

Speaker: Lucrezia Borriello

Dual Readout with Crystals

- Include in the IDEA detector design an additional layer of homogeneous material for the Calorimeter that allows to improve the energy resolution

IDEA calorimeter baseline
Cherenkov and Scintillation
fiber based
 σ_E/E (EM) $\sim 13\%$ /VE
 σ_E/E (HAD) $\sim 31\%$ /VE

Calorimeter homogeneous
medium=crystal
 σ_E/E (EM) $\sim 3\%$ /VE
 σ_E/E (HAD) $\sim 30\%$ /VE

- discriminating the Cherenkov signal from the Scintillation signals inside different crystal samples:



Cherenkov signal: prompt
Scintillation signals: longer

	time	spectrum
Cherenkov	fast	$1/\lambda^2$
Scintillation	slow	peaked

ER	Elvira Rossi
AD	Adelina D'Onofrio
FD	Francesco Di Capua
g	giulio
IR	Ilaria Rosa
OI	Orso Iorio
PI	Paolo Iengo
R	Raffaele
SM	Stefano Mastroianni
vi	Vincenzo Izzo

Speaker: Antonio D'Avanzo

WHAT: SIMULATING CRYSTALS RESPONSE IN GEANT4

- Study of the expected response of crystals for dual-readout electromagnetic calorimeter at FCC-ee
 - Simulation of scintillation/cherenkov light detection with SiPMs
- Fully customizable simulation of test beam setup in July 2024
 - Provided with optical photons tracing, crystal rotation and several materials
 - Able to reproduce data taking configurations to compare with data
 - Built with Geant4 toolkit on INFN IBISCO cluster in Naples



ER	Elvira Rossi
AD	Adelina D'Onofrio
FD	Francesco Di Capua
gianpaolocarino	gianpaolocarino
giulio	giulio
Ilaria Rosa	Ilaria Rosa
Orso Iorio	Orso Iorio
Paolo Iengo	Paolo Iengo
Raffaele	Raffaele
Stefano Mastroianni	Stefano Mastroianni
Vincenzo Izzo	Vincenzo Izzo

Speaker: Gabriel Botogoske

Proto-DUNE HD @CERN

Proto-DUNE SP

- 0.42 kton fiducial LAr
- 6 APAs

Proto-DUNE HD

- 0.28 kton fiducial LAr
- 4 APAs
- Moved to X-ARAPUCA
- 2 different WLS and SiPMs (Hamamatsu and FBK)
- APA 1- Full streaming mode APA 2-4: Self Trigger mode

Successful

Elvira Rossi, Bernardino Spizzo, Francesco Di Capua, giulio, Ilaria Rosa, Paolo Iengo, Stefano Mastroianni

6

Speaker: Giovanni Gaudino

Event Selection and dataset

One B meson is fully reconstructed using a multivariate algorithm, Full Event Interpretation (FEI) with Hadronic Tagging.

- $\mathcal{O}_{FEI} > 10^{-2}$
- $-0.15 < \Delta E = E_B^* - \sqrt{s}/2 < 0.1 \text{ GeV}$
- $M_{bc}c^2 = \sqrt{s/4 - (p_B^*)^2} > 5.27 \text{ GeV}$

Backgrounds

- $e^+e^- \rightarrow q\bar{q}$
- $e^+e^- \rightarrow \tau^+\tau^-$
- $e^+e^- \rightarrow B^+B^-$
- $e^+e^- \rightarrow B^0\bar{B}^0$

Signal is searched through τ decays

- $\tau^+ \rightarrow e^+\nu_e\bar{\nu}_\tau$
- $\tau^+ \rightarrow \mu^+\nu_\mu\bar{\nu}_\tau$
- $\tau^+ \rightarrow \pi^+\bar{\nu}_\tau$
- $\tau^+ \rightarrow \rho^+\bar{\nu}_\tau$ with $\rho^+ \rightarrow \pi^+\pi^0$

Rest of Event (ROE)

- 0 Extra Tracks
- Extra Clusters clean-up

Giovanni Gaudino

2025, January 16

Speaker: Ilaria Rosa

Dark photon: a detailed view

New vector field $F'_{\mu\nu}$ symmetric under a new $U(1)$ symmetry feebly interacting with the SM fields

A minimal extension to the SM: kinetic mixing with the SM hypercharge

$-\frac{e}{2} F'^{\mu\nu} B_{\mu\nu}$

$M(A')$ and ϵ are free parameters

Decay width dominated by lepton antilepton final states for $M(A') < 700$ MeV

Elvira Rossi, Bernardino Spizzo, giulio

2

Speaker: Renato Fiorenza

$K \rightarrow \pi v\bar{v}$: Precision test of the SM

SM: Z-penguin & box diagrams

• $\mathcal{B}(K \rightarrow \pi v\bar{v})$ highly suppressed in SM

- GIM mechanism & maximum CKM suppression $s \rightarrow d$ transition: $\sim \frac{m_t}{m_W} |V_{ts}^* V_{td}|$
- Theoretically clean ⇒ high precision SM predictions, high sensitivity to BSM
 - Dominated by short distance contributions
 - Hadronic matrix element extracted from $\mathcal{B}(K \rightarrow \pi l \nu)$ decays via isospin rotation

Decay Mode BR	SM [Buras et al. EPJC 82 (2022) 7, 619]	SM [D'Ambrosio et al. JHEP 09 (2022) 149]	Experimental Status
$\mathcal{B}(K^+ \rightarrow \pi^+ v\bar{v})$	$(8.60 \pm 0.42) \times 10^{-11}$	$(7.86 \pm 0.61) \times 10^{-11}$	$(10.6^{+4.1}_{-3.5}) \times 10^{-11}$ (NA62)
$\mathcal{B}(K_L \rightarrow \pi^0 v\bar{v})$	$(2.94 \pm 0.15) \times 10^{-11}$	$(2.68 \pm 0.30) \times 10^{-11}$	$< 2.2 \times 10^{-9}$ (KOTO)

NA62 (2016-18 data): [JHEP 06 (2021) 093]
KOTO (2021 data): arXiv:2411.11237

INFN

Differences in SM calculations from choice of CKM parameters:
 [Eur.Phys.J.C 84 (2024) 4, 377]

2