



Photons and pions off nuclei as probes for the nuclear matter equation of state

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Extreme matter in neutron stars



Image credits: MUSES Collaboration

Nuclear matter equation of state (EOS)

$$\frac{E}{A}(n,\alpha) = \frac{E}{A}(n,0) + \frac{S(n)}{n}\alpha^2 + \mathcal{O}[\alpha^4] \qquad n = n_p + n_n$$

symmetry energy

$$S(n) = J + \frac{L^n - n_0}{3n_0} + \dots$$

symmetry energy
at saturation density

$$Slope \text{ parameter,}$$

related to pressure of
pure neutron matter
at saturation density

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How to constrain the symmetry energy?

Neutron-skin thickness

$$R_{skin} = R_n - R_p$$



Data from M. Centelles et al, PRC 82, 054314 (2010), Hu et al, Nat Phys. 18, 1196–1200 (2022).

Ab initio nuclear theory



□ Building blocks: protons and neutrons.

□ Solve quantum many-body problem

 $H |\psi\rangle = E |\psi\rangle$ $H = T + V_{NN} + V_{3N}$

with controlled approximations.

□ 2 ingredients: nuclear interactions and many-body solver.

Chiral Effective Field Theory (EFT)



Coupled-cluster theory

lacksquare Starting point: Hartree-Fock reference state $\left|\Phi_{0}\right\rangle$

□ Add correlations via:

$$|\Psi_0
angle = e^T |\Phi_0
angle$$

with

$$T = \sum \mathbf{t}_i^a a_a^{\dagger} a_i + \sum \mathbf{t}_{ij}^{ab} a_a^{\dagger} a_b^{\dagger} a_j a_i + \sum \mathbf{t}_{ijk}^{abc} a_a^{\dagger} a_b^{\dagger} a_c^{\dagger} a_k a_j a_i + \dots$$

Coupled-cluster theory

 \square Starting point: Hartree-Fock reference state $|\Phi_0\rangle$

□ Add correlations via:

$$|\Psi_0
angle=e^T|\Phi_0
angle$$

with



G. Hagen, T. Papenbrock, M. Hjorth-Jensen, D. J. Dean, RPP 77, 096302 (2014).

Coupled-cluster theory

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Neutron skin measurements



FB, PhD Thesis, JGU Mainz (2024).

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Nuclear structure input for pion scattering

- Goal: predict differential cross sections for pion scattering off non-zero isospin nuclei, as ⁴⁸Ca.
- Structure input for reaction calculation: onebody densities at first order, and two-nucleon correlation functions at second order.
- Densities from coupled-cluster and correlation functions from Hartree-Fock using the ΔNNLO_{GO}(394) and NNLO_{sat} chiral forces.



V. Tsaran, F. Marino, S. Bacca, FB et al, arXiv:2505.18459 [nucl-th].

Pion scattering off ⁴⁸Ca



V. Tsaran, F. Marino, S. Bacca, FB et al, arXiv:2505.18459 [nucl-th].

Pion scattering off ⁴⁸Ca



V. Tsaran, F. Marino, S. Bacca, FB et al, arXiv:2505.18459 [nucl-th].

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Electric dipole polarizability



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Data from M. Centelles et al, PRC 82, 054314 (2010), X. Roca-Maza et al, PRC 88, 024316 (2013), Hu et al, Nat Phys. 18, 1196–1200 (2022).

Nuclear response functions



From bound to dipole-excited states

 $R(\omega)$



Continuum problem

S. Bacca, N. Barnea, G. Hagen, G. Orlandini, T. Papenbrock, PRL 111, 122502 (2013).

From bound to dipole-excited states



S. Bacca, N. Barnea, G. Hagen, G. Orlandini, T. Papenbrock, PRL 111, 122502 (2013).

The case of ^{40,48}Ca



R. Fearick, P. von Neumann-Cosel, S. Bacca, FB et al, Phys. Rev. Research 5, L022044 (2023).

Constraints on symmetry energy $S(n) = J + L \frac{n - n_0}{3n_0} + \dots$



FB, PhD Thesis, JGU Mainz (2024).

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FB, PhD Thesis, JGU Mainz (2024).









We need to extend our method beyond closed-shell nuclei!

Open-shell nuclei: two-particle-attached systems (2PA)



$\alpha_{\rm D}$ along the oxygen chain



$\alpha_{\rm D}$ along the calcium chain



$\alpha_{\rm D}$ along the calcium chain



FB et al., PRC 110, 044306 (2024).

Back to the EOS



Incompressibility of a finite nucleus

Incompressibility of a finite nucleus

 $K_A = \frac{M}{\hbar^2} R_m^2 E_{\rm monopole}^2$

from moments of the isoscalar monopole response

$$E_{\text{monopole}}^2 = \frac{m_1}{m_{-1}}$$



Y. Gupta et al, PLB 760, 482-485 (2016).

Incompressibility of nuclear matter

 $K_A = K_{\rm vol} + K_{\rm surf} A^{-1/3} + \dots$

incompressibility of nuclear matter

Incompressibility of nuclear matter



Conclusions

- Modern nuclear structure input can help in refining uncertainties in pion scattering differential cross sections.
- Electromagnetic observables cast light on the collective excitations of the nucleus as well as constraining the symmetry energy.
- U We extended ab initio reach of this observable to **nuclei in the vicinity of closed shells**.
- Estimates of the incompressibility of symmetric nuclear matter based on ab initio predictions of monopole moments in finite nuclei are consistent with nuclear matter calculations.

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