







Underground Site Characterization & Control of the GEIVIINI Platforms

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Experimental Framework: GEMINI



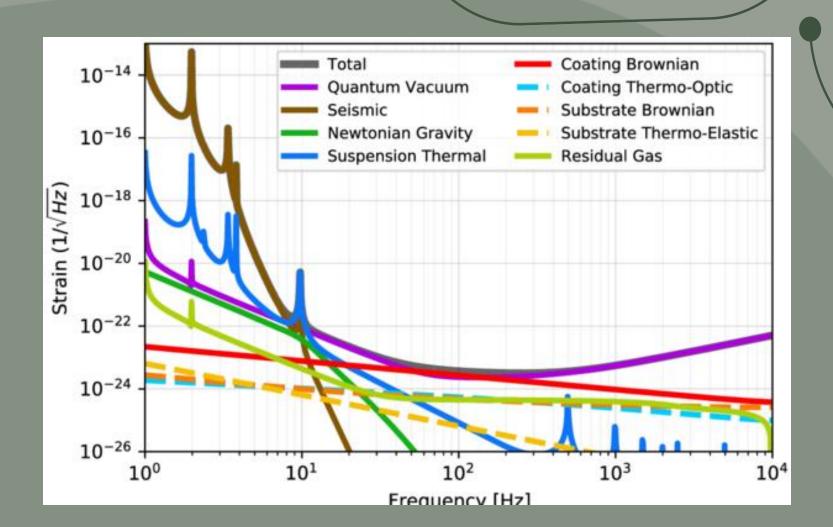




The Seismic VVall

Below 10 Hz,
gravitational wave
detectors are limited
by ground motion

Even the most advanced isolation systems cannot fully reject seismic noise



Ground displacement follows approximately:

$$S_x(f) = \alpha / f^2$$
, with $\alpha \approx 10^{-7} \text{ m} \cdot \text{Hz}^{3/2}$

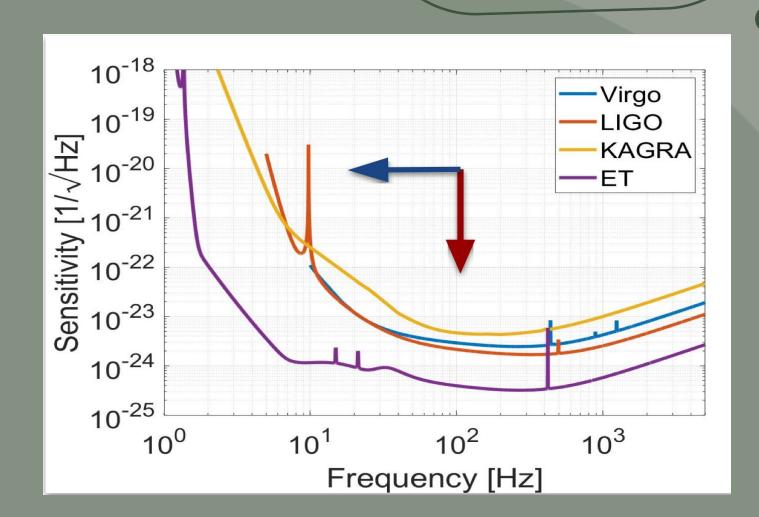
At 10 Hz: ground strain \approx 10^{-12} / $\sqrt{\text{Hz}}$, but required detector sensitivity is 10^{-21} / $\sqrt{\text{Hz}}$

Seismic vibrations must be suppressed by 10°!

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GEMINI I





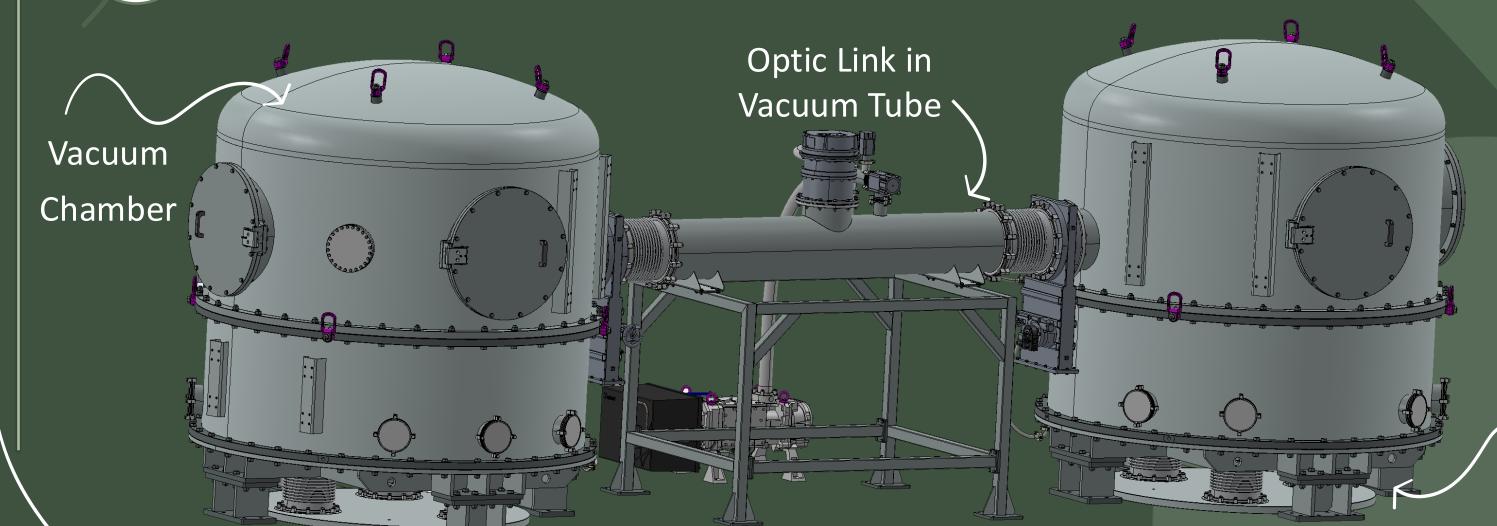
Testbed for validating new low frequency seismic isolation technologies.



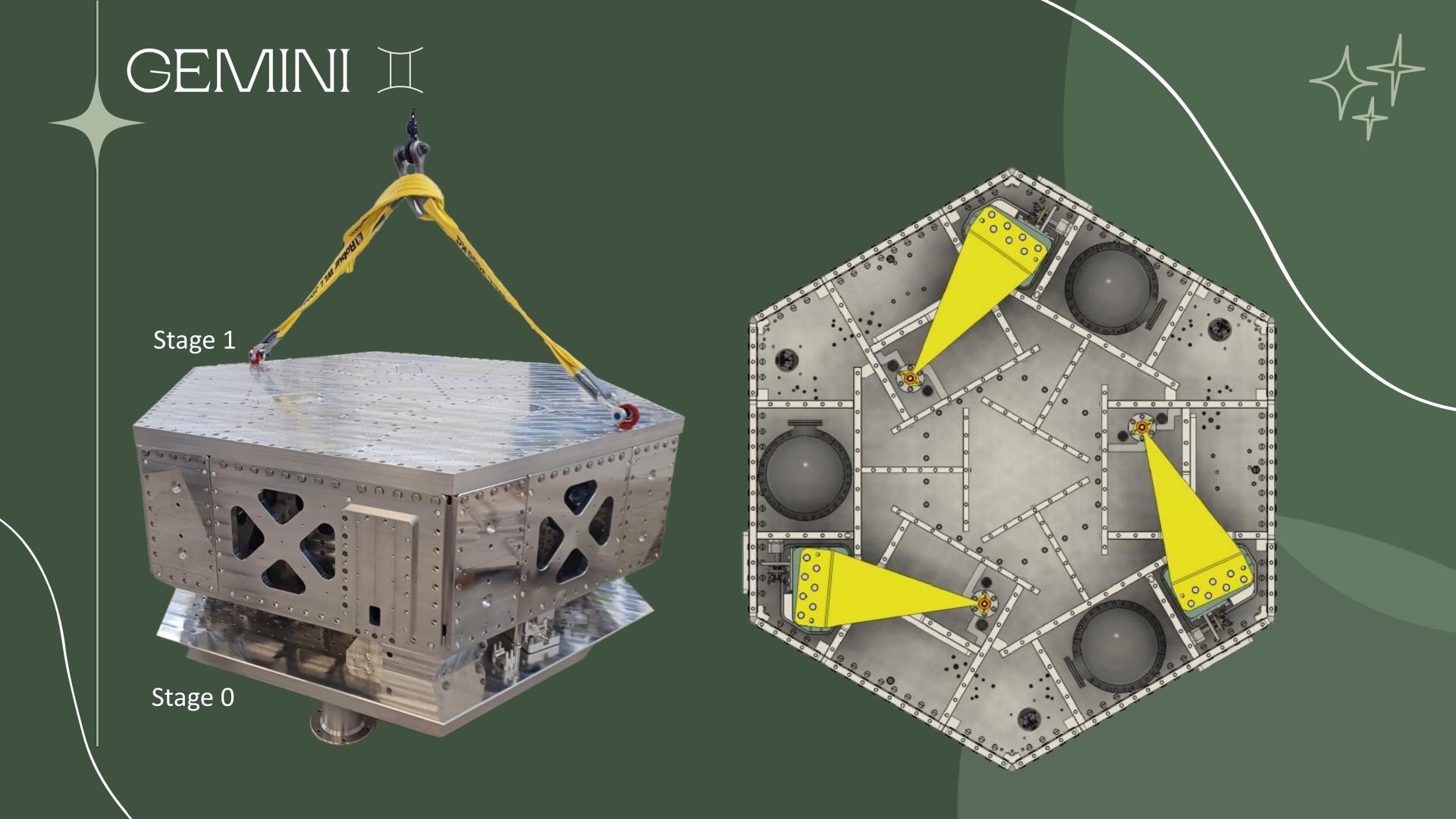
Combines passive isolation (spring blades, flexure rods) with active isolation (sensors, actuators).



Technology can then be scaled up and implemented in future GW facilities such as the Einstein Telescope and the Lunar GW Antenna.

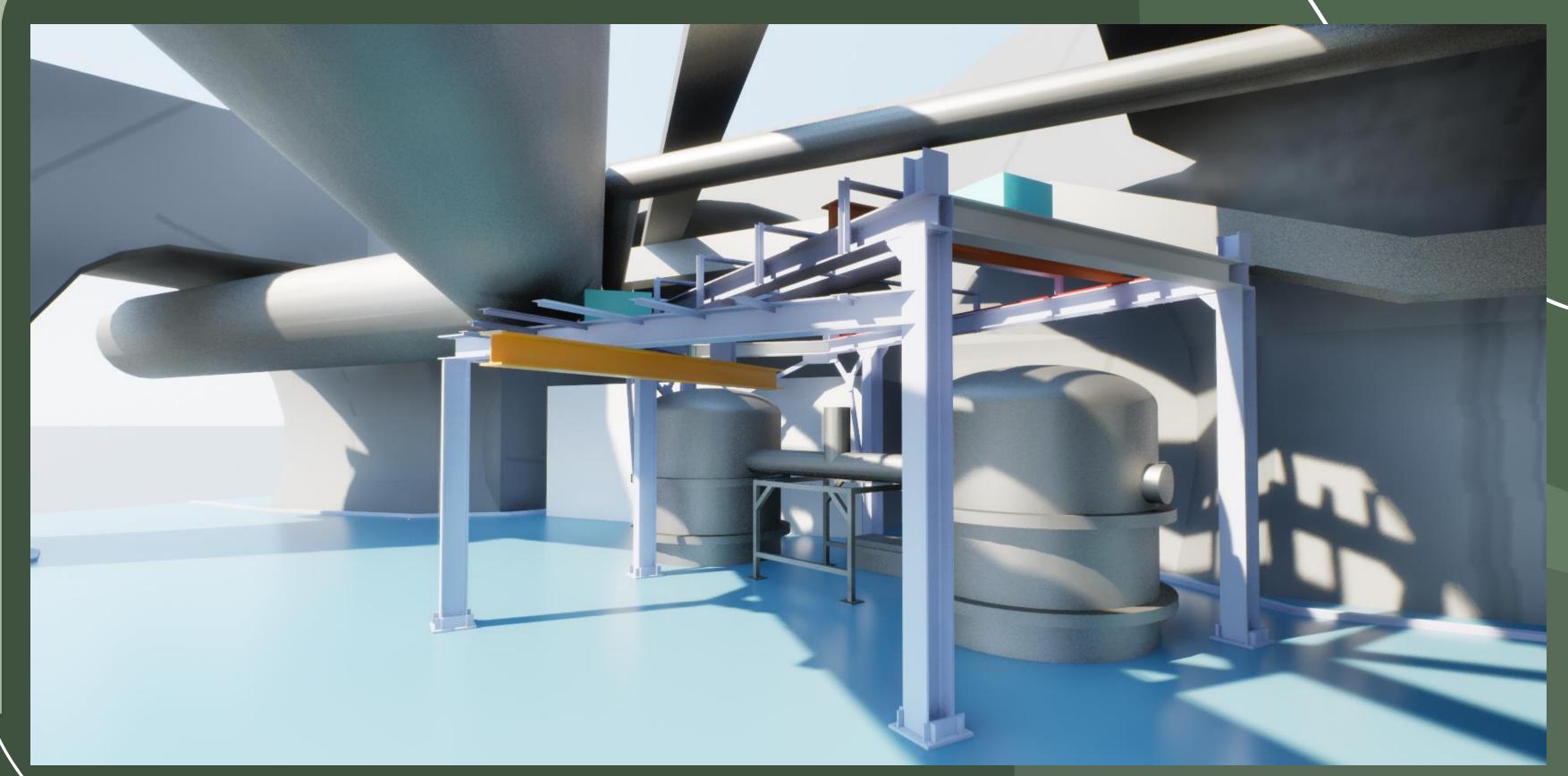


Seismic Isolation
Platform



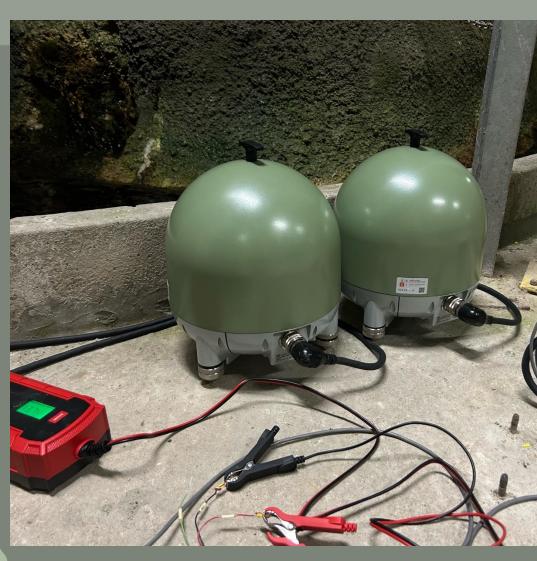
GEMINI I

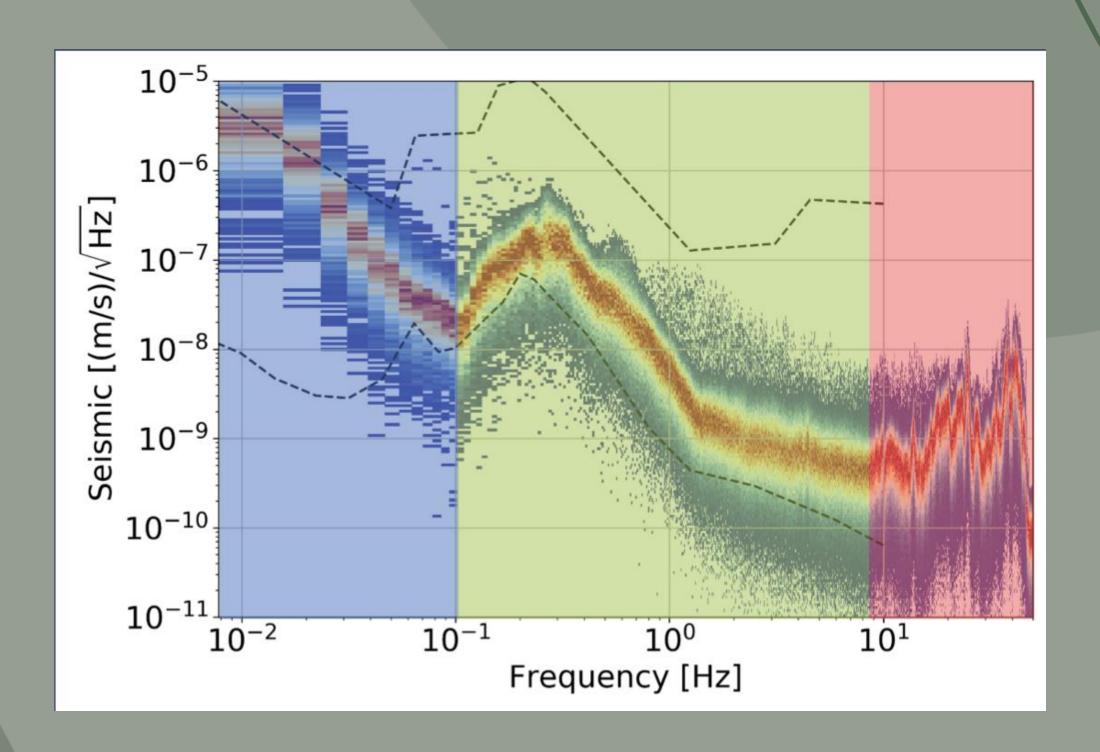




Motivation: previous measurements of two seismometers placed a few cm apart showed uncorrelated noise below 0.1 Hz.

Raised question: Is the absence of correlation due to the instruments themselves, or due to local temperature/air fluctuations?



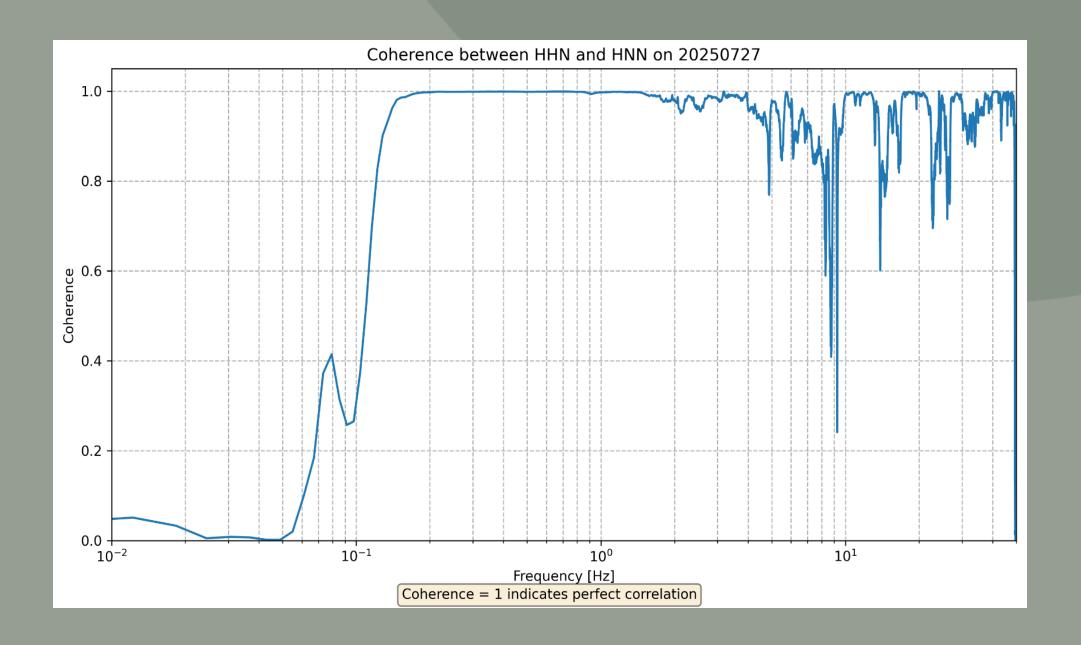




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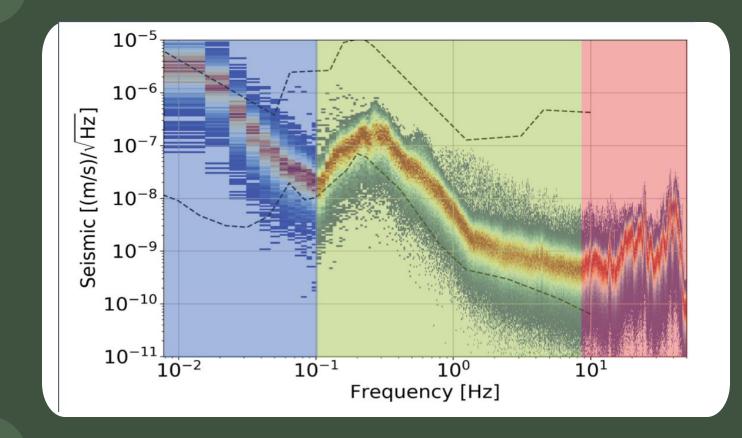




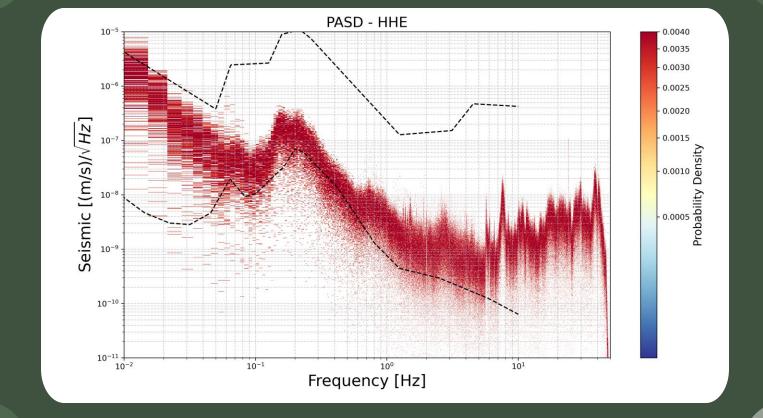


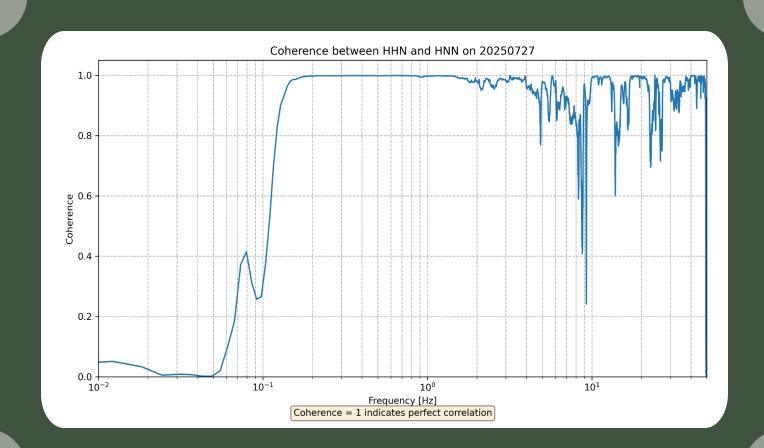
Data Analysis

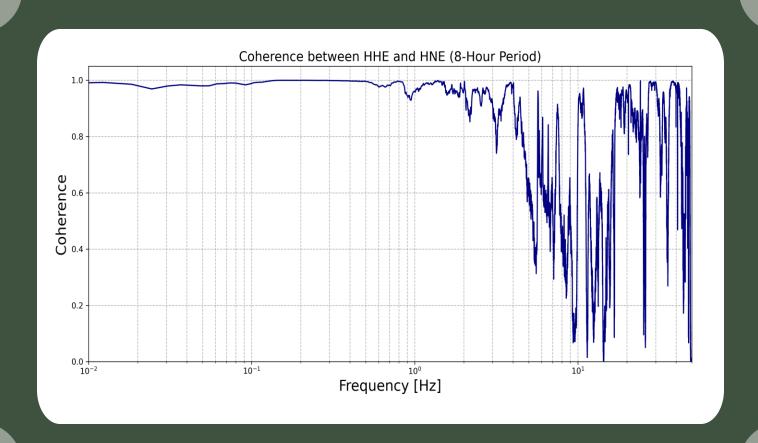
Original



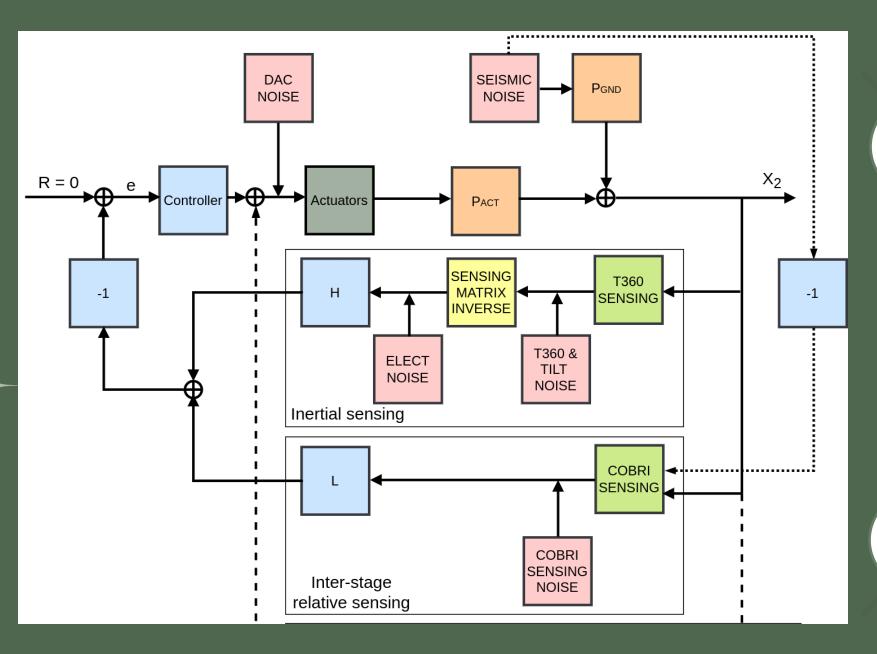






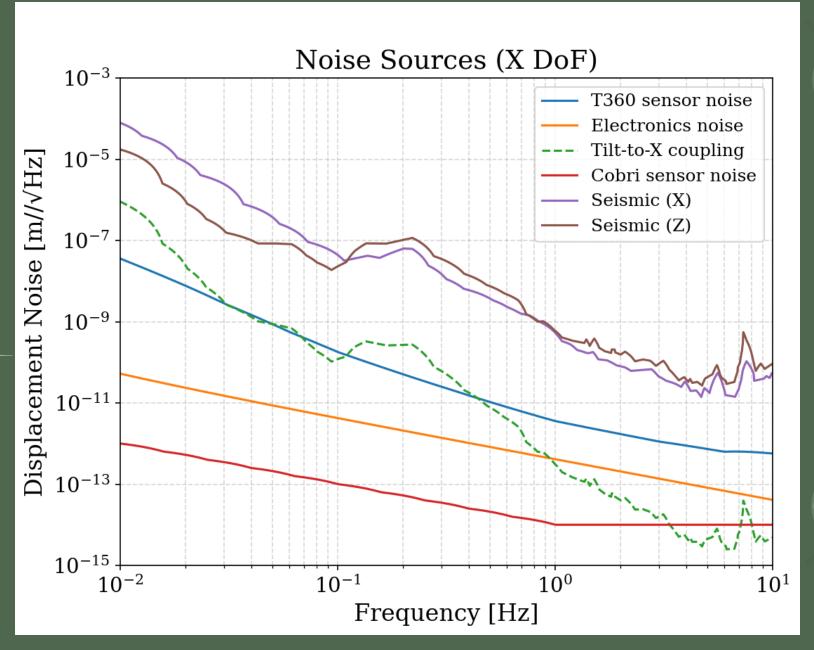


Time Domain Simulations



- Simulated GEMENI platform with displacement sensors
 - Feedback control for residual noise minimization and maximum stability
 - Controller for motion correction
 - Actuators for motion counteraction via applied force

Time Domain Simulations



Implementation of noise into code simulation for more realistic model

Seismic noise

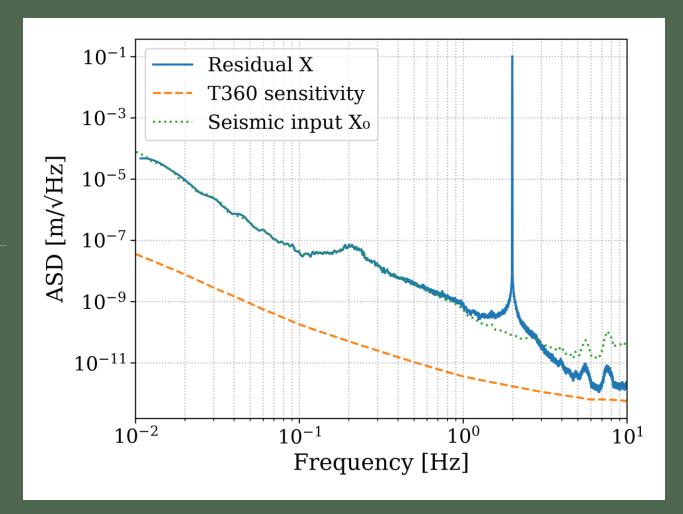
Electronic noise

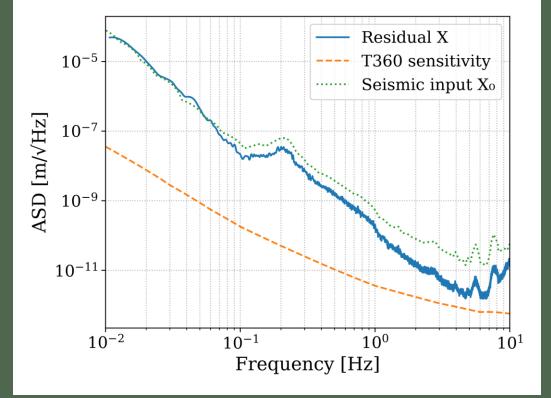
Tilt-to-length coupling

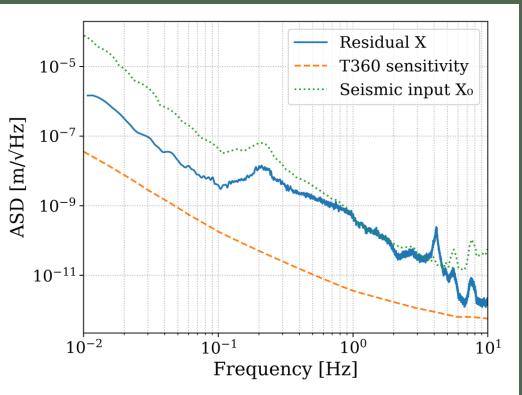
Sensing noises

Controller Implementation

Without Control







PID Controller

ZPK Controller



Trozzo, L., & Badaracco, F. (2022). Seismic and Newtonian Noise in the GW Detectors. Galaxies, 10, 20. https://doi.org/10.3390/galaxies10010020

https://doi.org/10.48550/arXiv.2509.04730 GEMINI: The First Underground Testbed for Seismic Isolation and Interplatform Control in Next-Generation Gravitational-Wave Detectors Harms, J. Terrestrial gravity fluctuations. Living Rev Relativ 22, 6 (2019). https://doi.org/10.1007/s41114-019-0022-2

Thank you for your attention!



Summary

Seismic Wall:

GEMINI:

Methods:

Results:

Gravitational wave detectors limited by ground motion below 10 Hz

Seismic vibrations must be

suppressed by many orders

of magnitude

below 10 Hz technologies

Combines passive isolation (platform) with active isolation (sensors)

Testbed for new **low**

frequency seismic isolation

Build enclosure for seismometer isolation to suppress environment effects

GEMINI simulation based on a feedback loop, so the platform continuously corrects itself to remain stable Seismometer isolation led to **higher coherence** of the data

Controller design implemented but **optimal tuning** requires more time