

Neutrino and Beyond Standard Model Physics Searches at the CERN ProtoDUNE Detectors

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On behalf of the *DUNE Collaboration*
NuTel 2025
30/09/2025



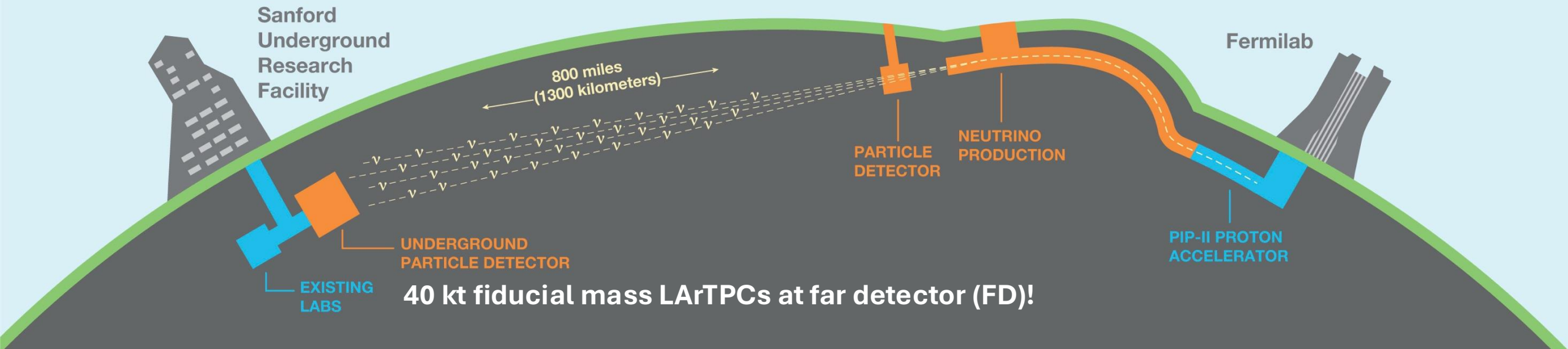
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Deep Underground Neutrino Experiment (DUNE)

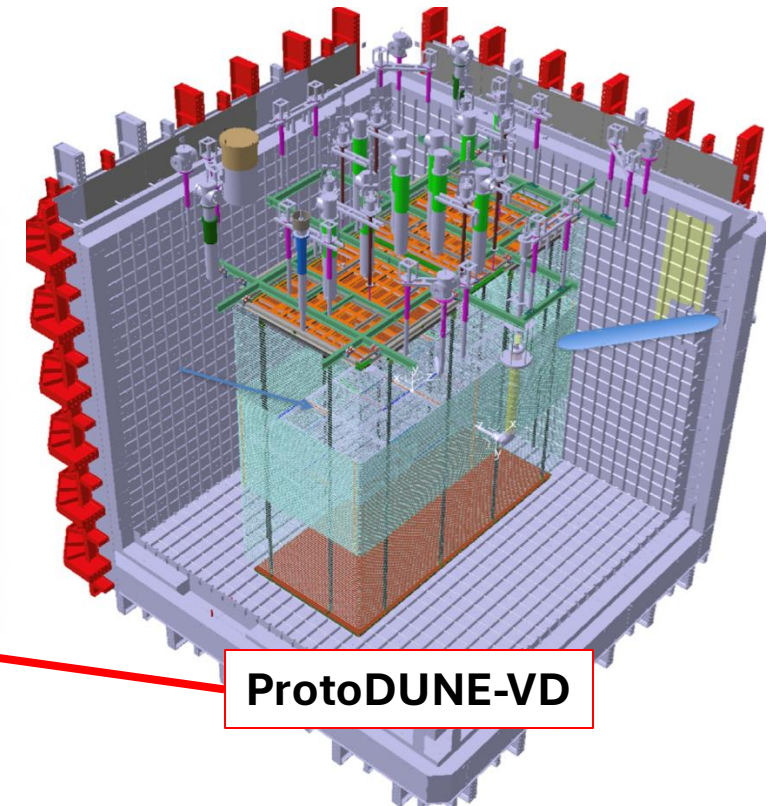
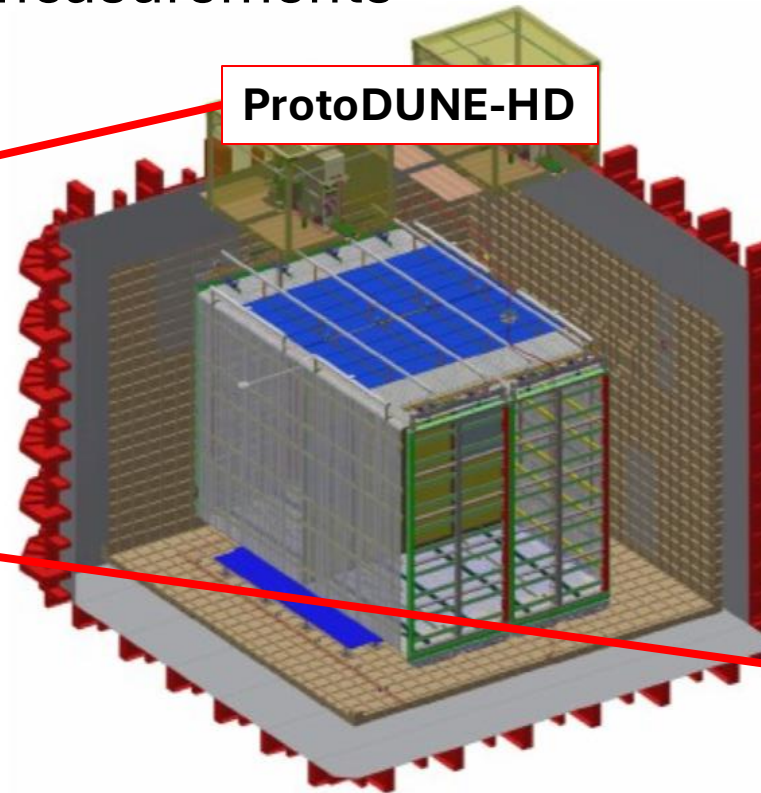
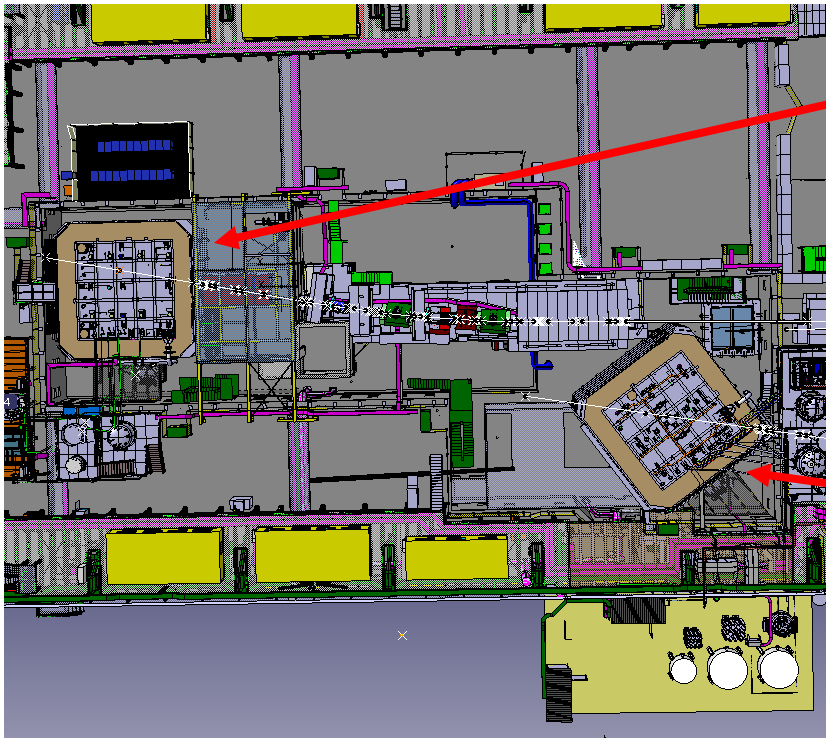
- See DUNE plenary talk by David Rivera on Thursday for more details

Observe $\nu_\mu \rightarrow \nu_\mu$, $\nu_\mu \rightarrow \nu_e$, $\bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$ and $\bar{\nu}_\mu \rightarrow \bar{\nu}_e$
Measure δ_{CP} , Δm_{32}^2 , θ_{23} , θ_{13} , mass ordering



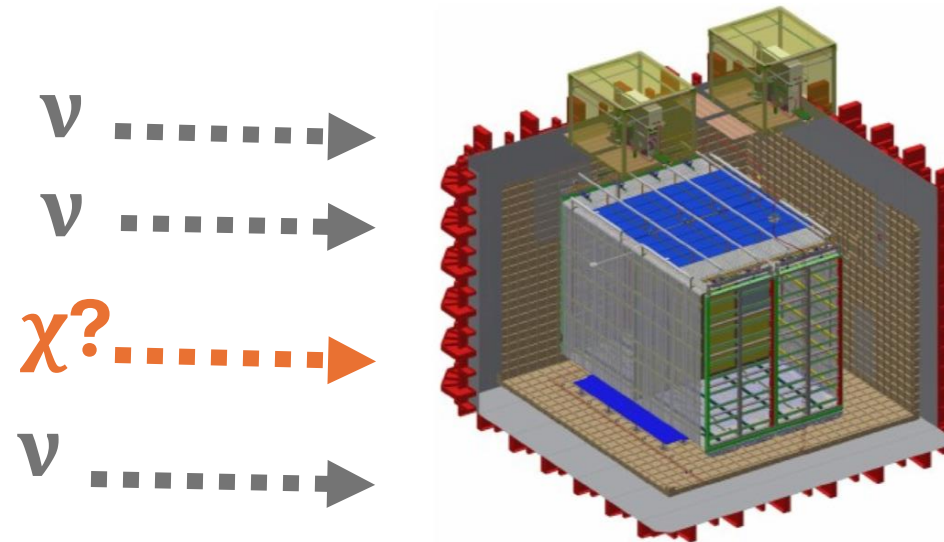
ProtoDUNE at CERN Neutrino Platform

- Two large **prototypes** of the DUNE FD at **CERN Neutrino Platform**
- Built to **test LArTPC detector technology** for the DUNE FD and measure hadron-Ar cross section measurements



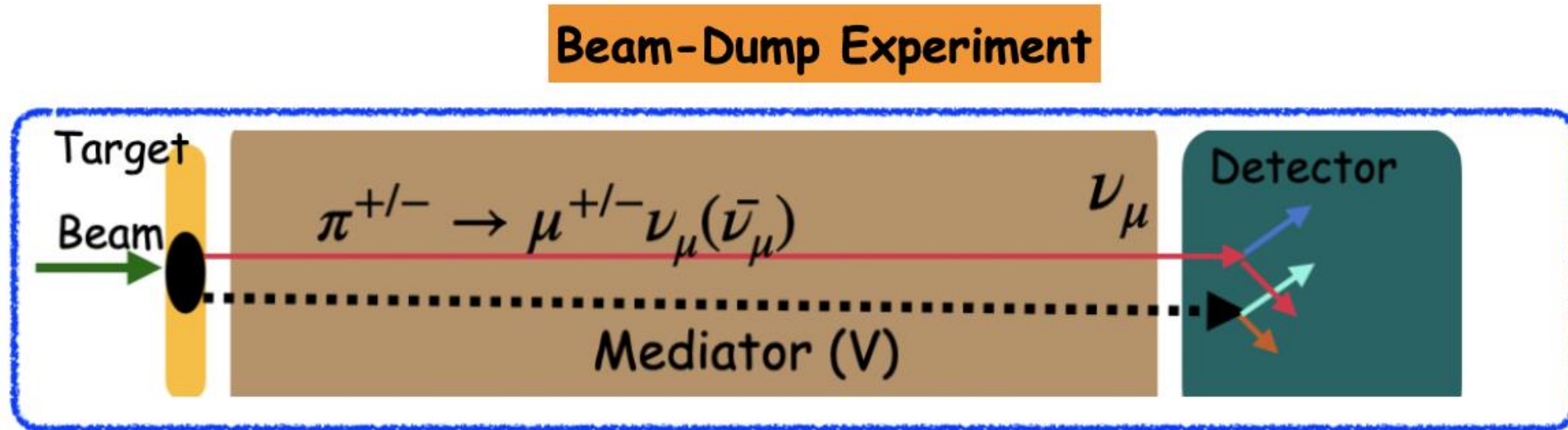
Neutrinos and BSM Physics in ProtoDUNE

- New idea: search for **neutrinos** and **beyond-standard model (BSM)** particles at **ProtoDUNE** from the **CERN Super Proton Synchrotron (SPS)**
- Details of proposal and estimated physics sensitivity of a ProtoDUNE BSM search published: [J. High Energ. Phys. 2024, 134 \(2024\)](#)
- ProtoDUNE becomes a beam-dump experiment – common way to search for weakly interacting long-lived particles or neutrinos



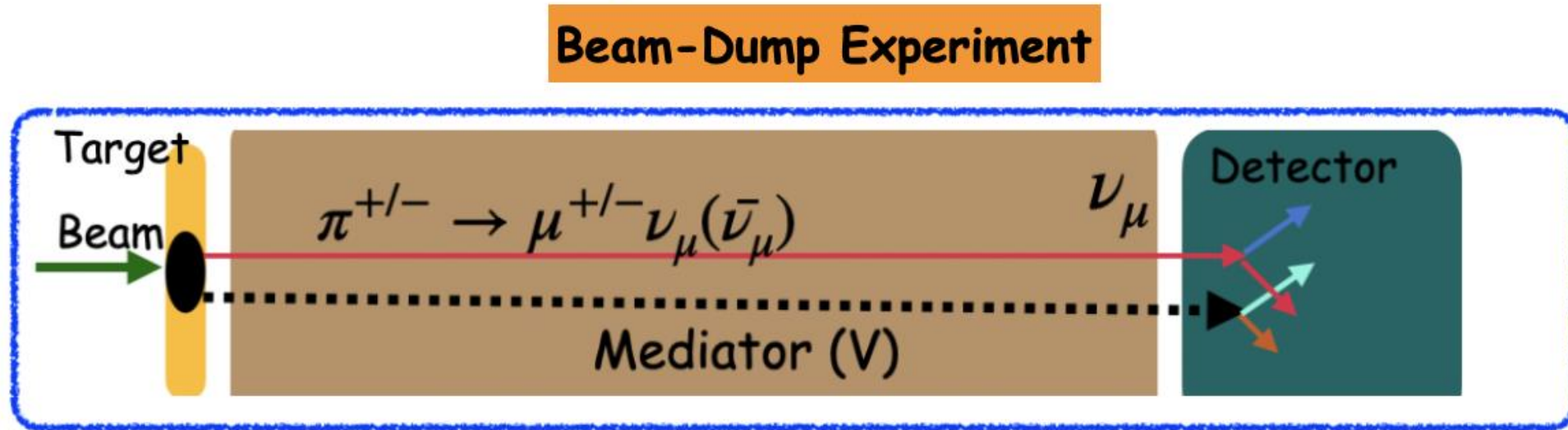
Beam-Dump Experiment

- A beam (usually protons) hitting a **fixed target** is a common way to produce **neutrinos** – or possibly something more exotic (e.g. Heavy Neutral Lepton)
- Pions, kaons etc. are produced in the target – **decay downstream** to neutrinos/BSM
- Place a detector in path of the neutrinos/BSM to try and observe them



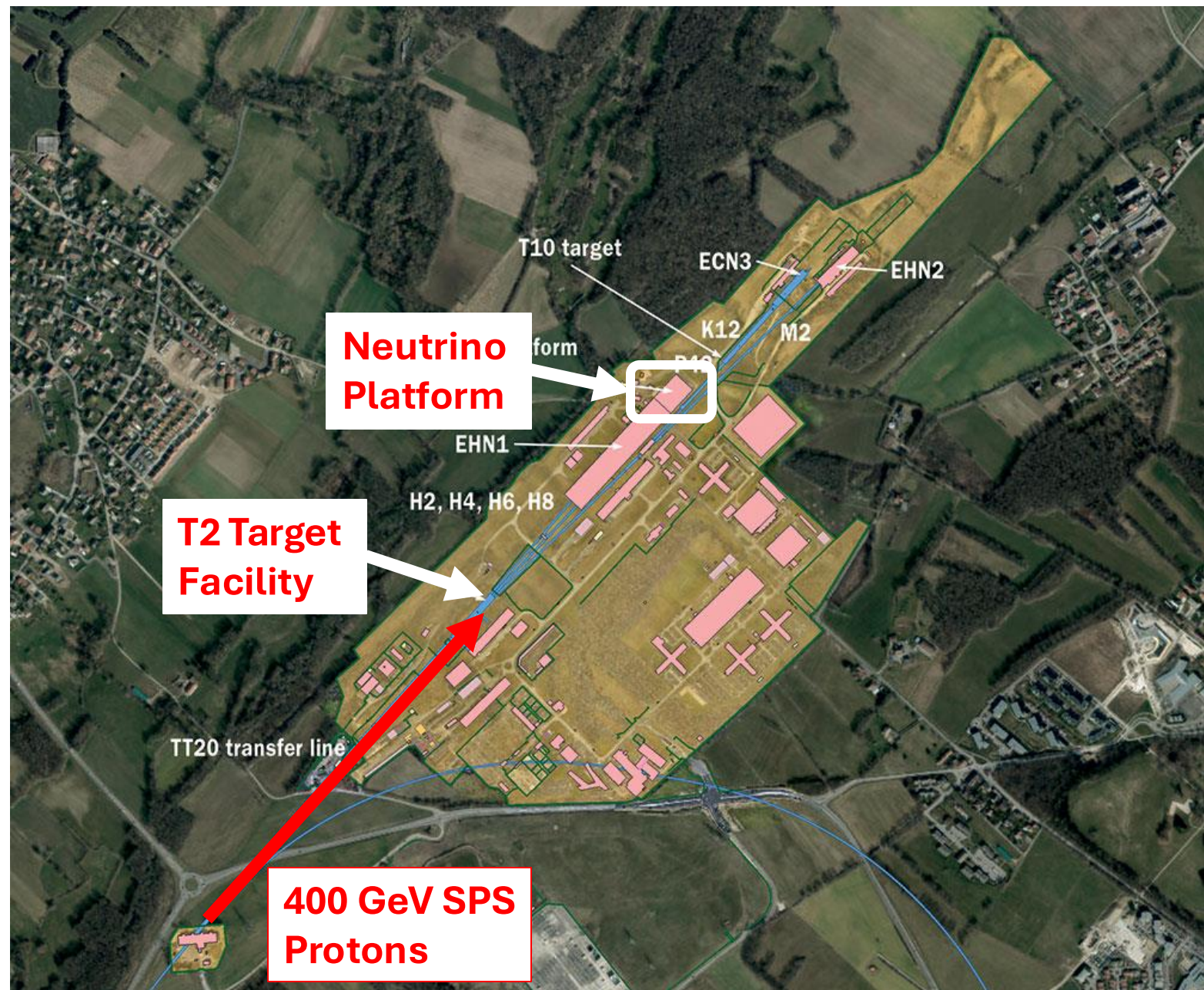
Beam-Dump Experiment

How can the ProtoDUNE detectors be a beam-dump experiment?

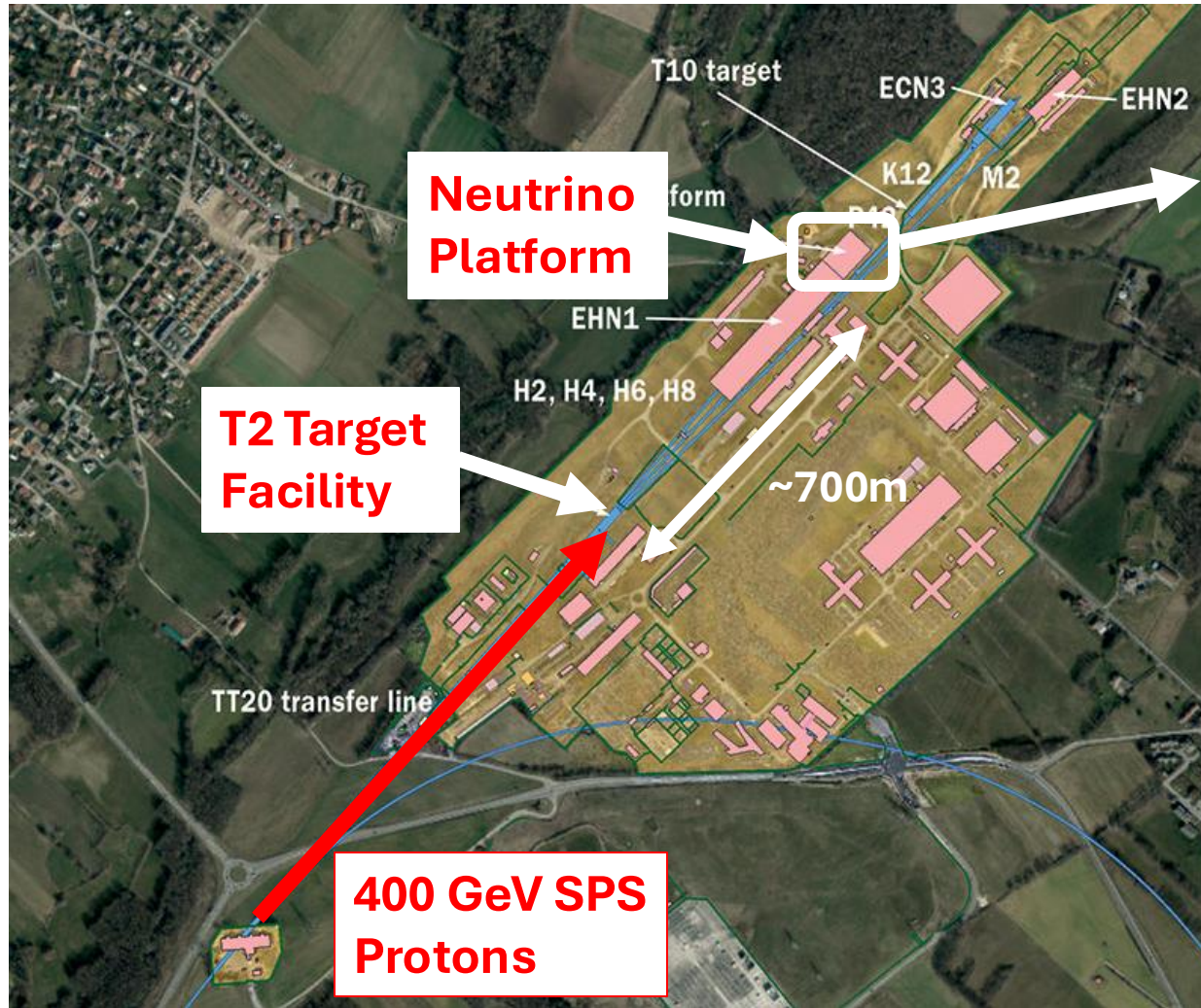


CERN SPS and Neutrino Platform

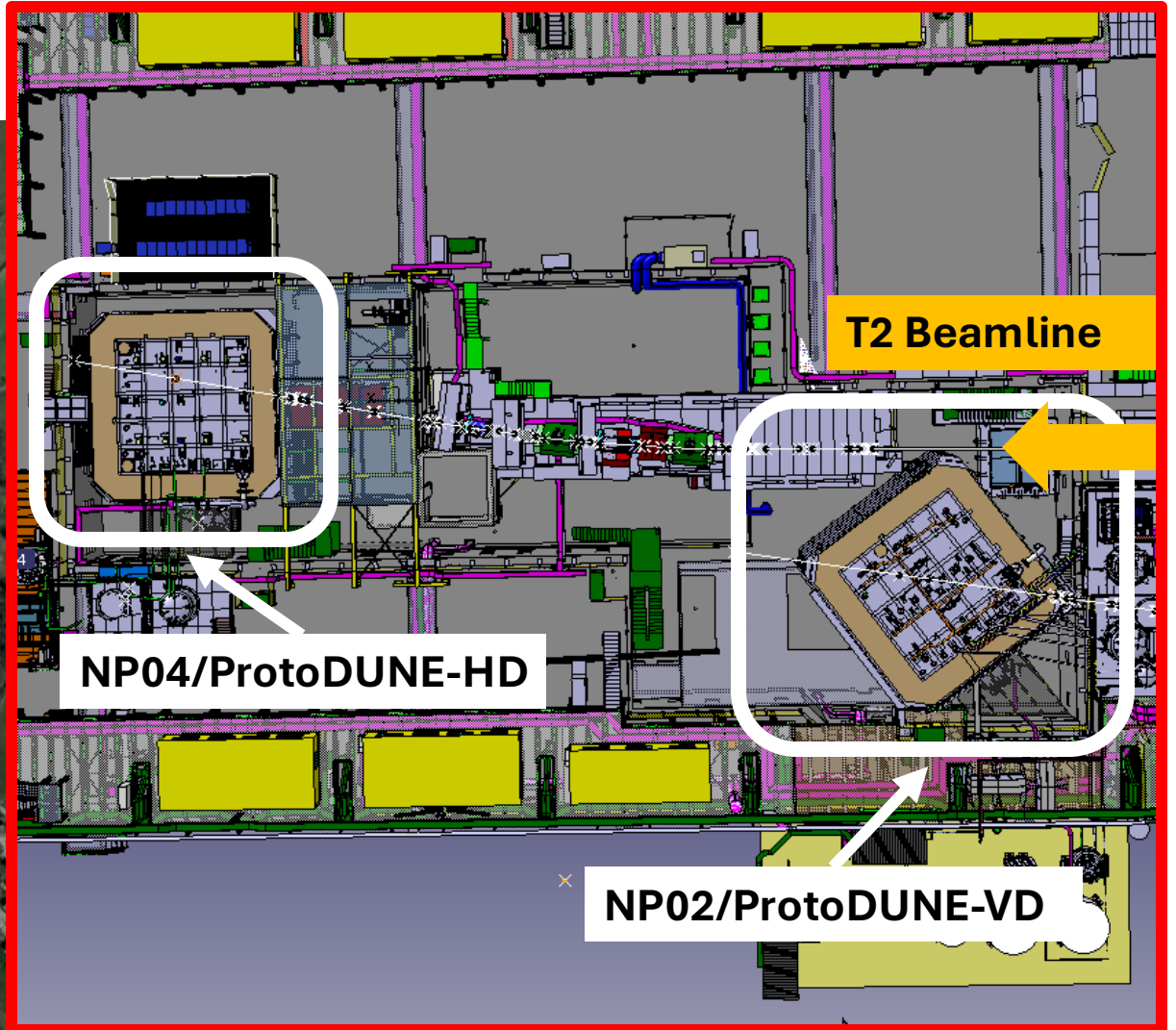
- CERN SPS beam accelerates protons to **400 GeV**
- SPS protons hit the **“T2” target**
- Protons hitting T2 are **aligned in the direction of the Neutrino Platform**



The T2 Beamline

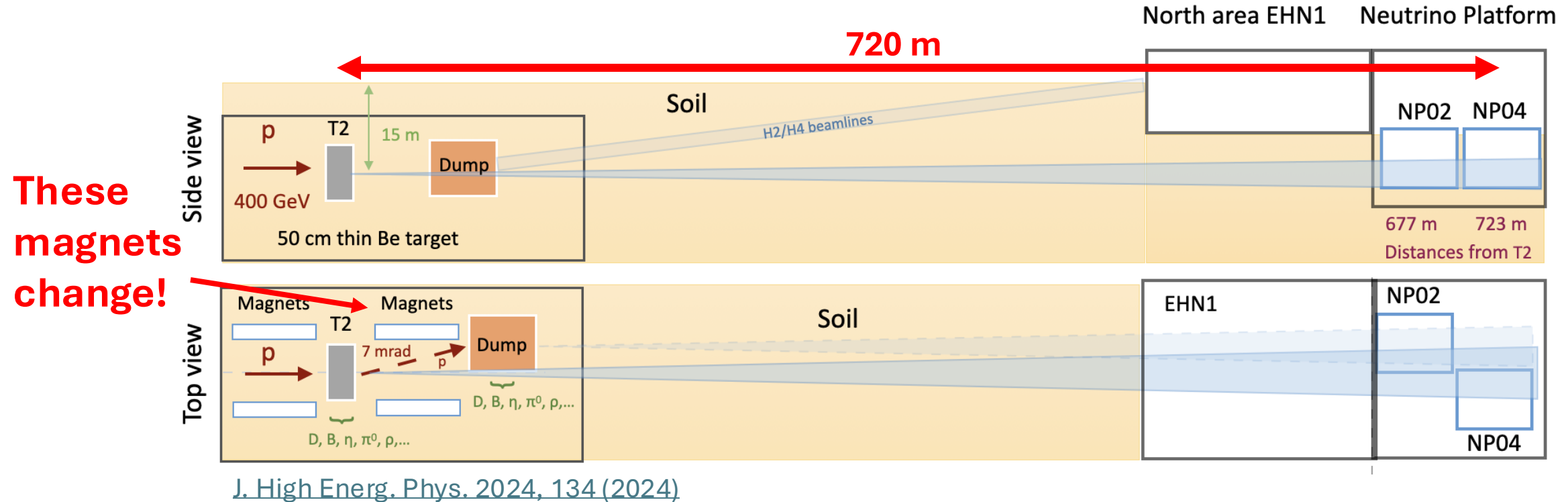


Neutrino Platform Top View



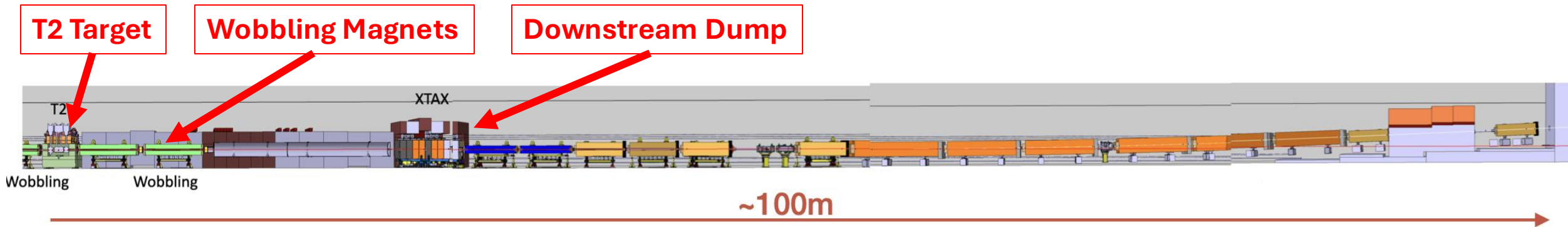
The T2 beamline

- We have a **670-720m beamline (VD or HD ProtoDUNE)** – neutrinos/BSM produced in T2 or downstream dump (TAX)



T2 Target Area G4 Simulation

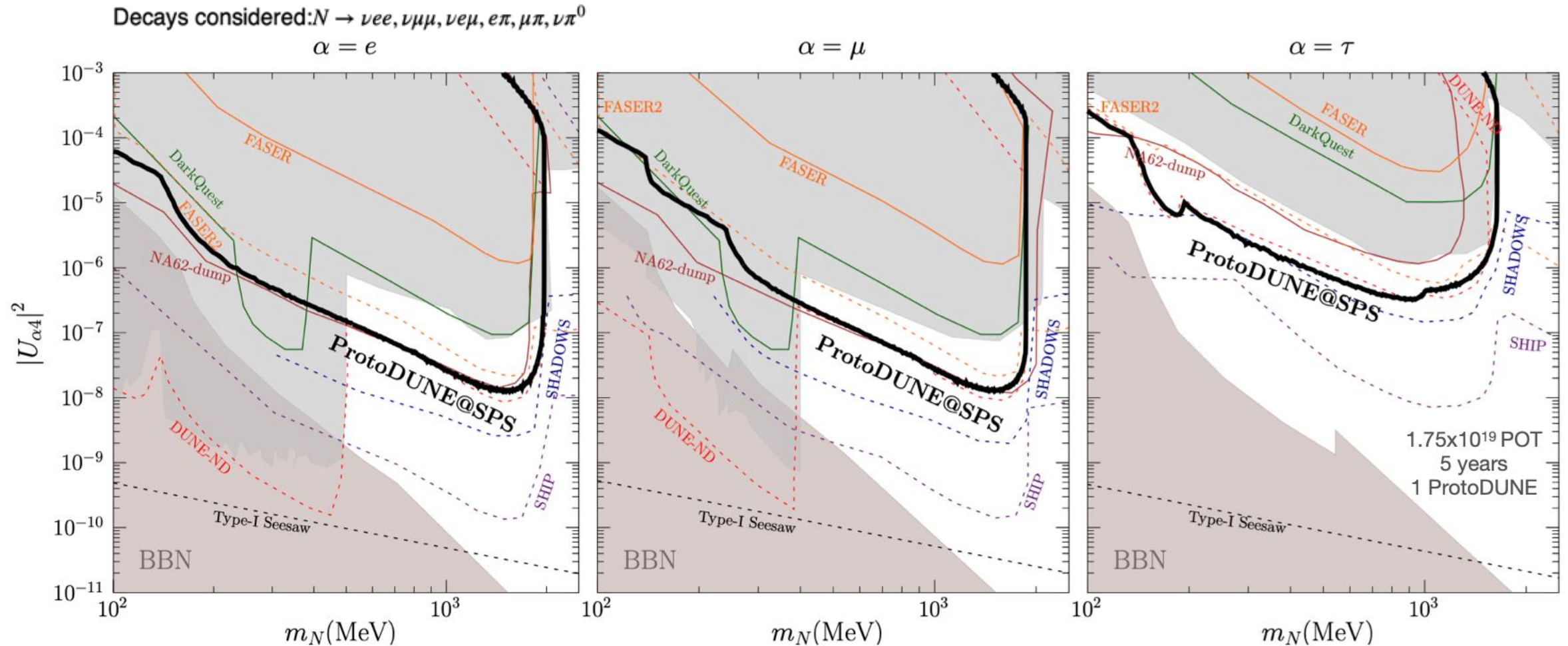
- A full **Geant4 model** of the **target facility** around T2 has been created
- Simulate 400 GeV SPS protons hitting this facility and the subsequent hadrons produced
- Produce fluxes of BSM particles, such as Heavy Neutral Leptons (HNLs)
- Dump, magnets and ~500m of soil absorb remaining hadrons



Potential Sensitivity of ProtoDUNE BSM Program

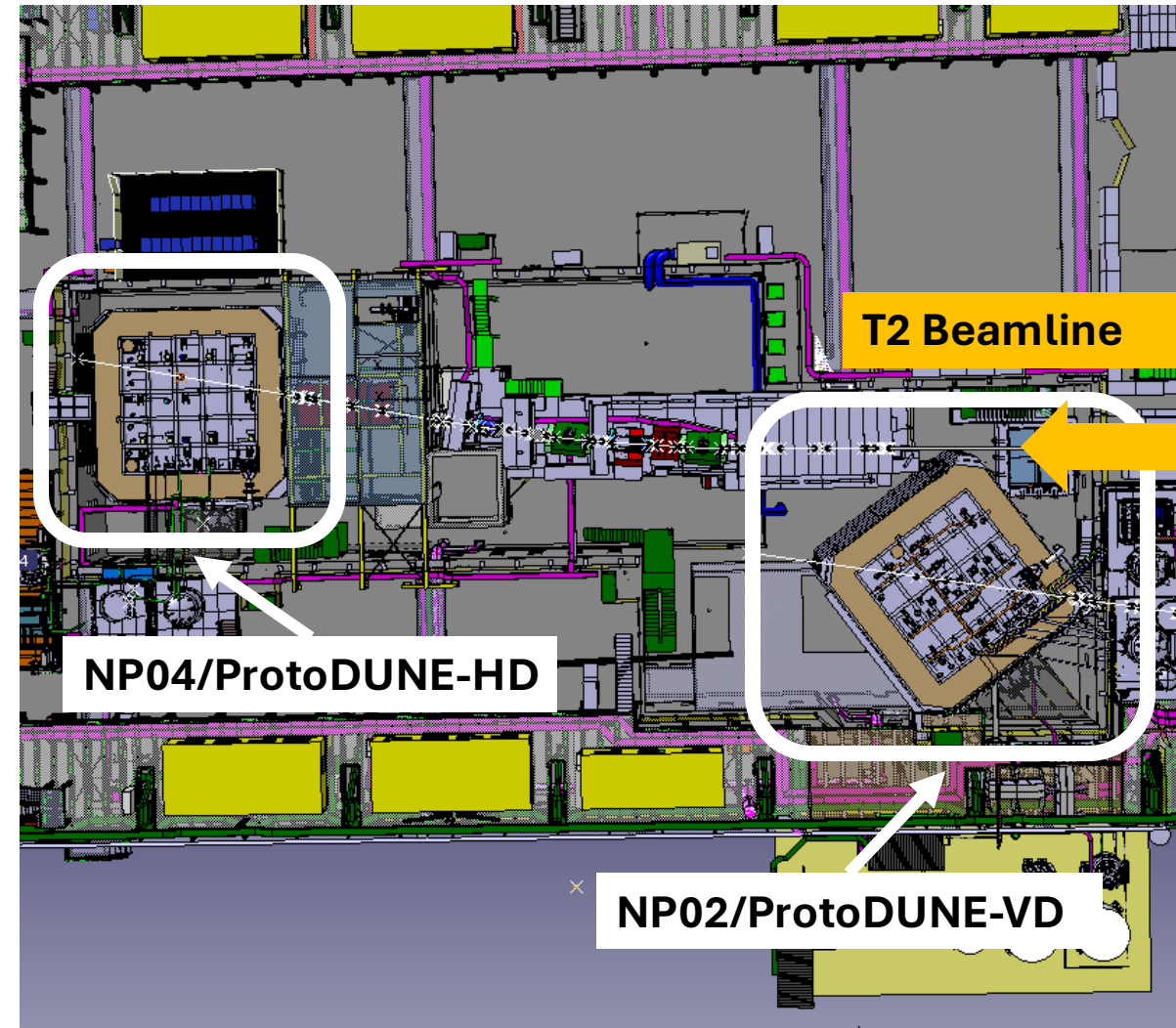
Assuming negligible background in NP02

- Sensitivity for HNL benchmark model: [J. High Energ. Phys. 2024, 134 \(2024\)](#)



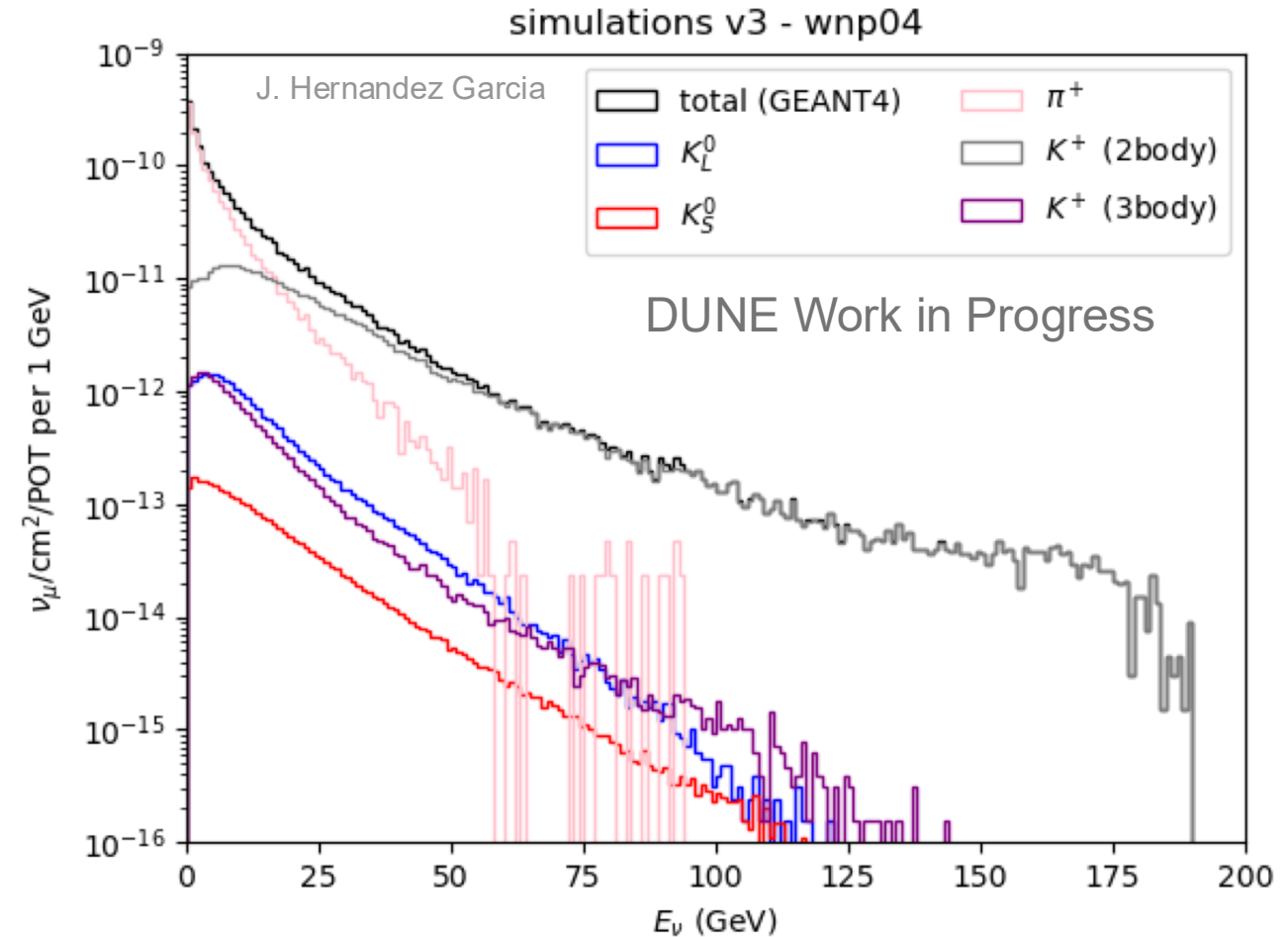
Proof of Principle – Observing Neutrinos at ProtoDUNE

- The aim is to do a **BSM search at ProtoDUNE**
- First search for something that **exists** – **neutrinos**!
- Proof of concept: observation of **neutrinos** from the **T2 target** – becomes background in BSM search
- A sample of neutrinos in a FD-like LArTPC will be useful – **test energy reconstruction methods** or possibly a cross-section measurement



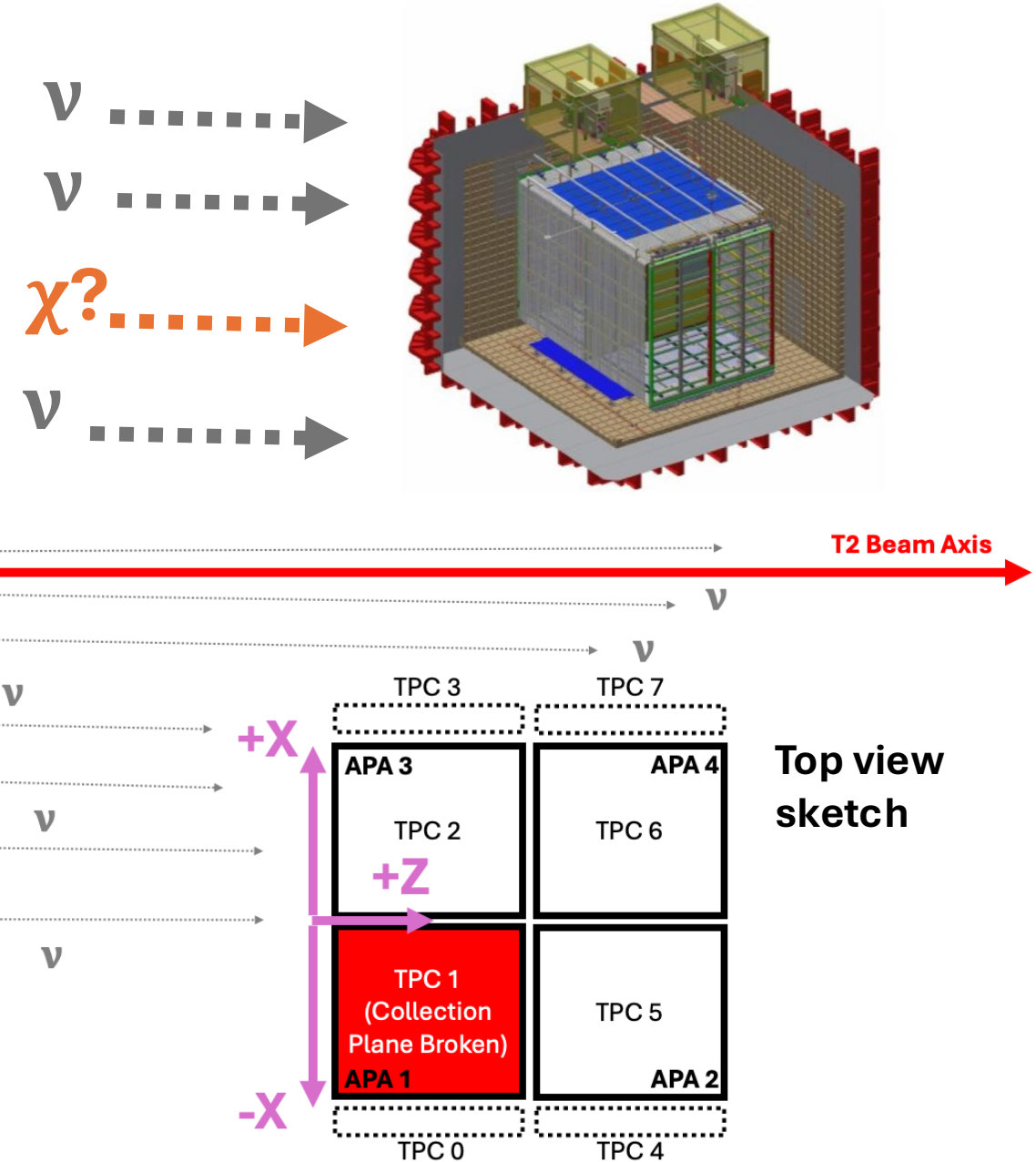
What is the Neutrino Flux?

- Geant4 simulation of SPS protons hitting T2 target area and downstream facilities
- Flux peaks at **0-10 GeV** and long tail to **more than 100 GeV**
- Very **high-energy** neutrinos!
- Simulate separate **flux** for **each T2 magnet configuration**



ProtoDUNE-HD (NP04)

- Horizontal drift configuration
- 4 drift volumes – 2 wire planes either side of cathode
- We are **~720m** downstream from T2 and only **~7m** off the central axis of the beam
- Neutrinos (or BSM particles) are moving **totally parallel** to the **central cathode**
- **Many of them very high energy**

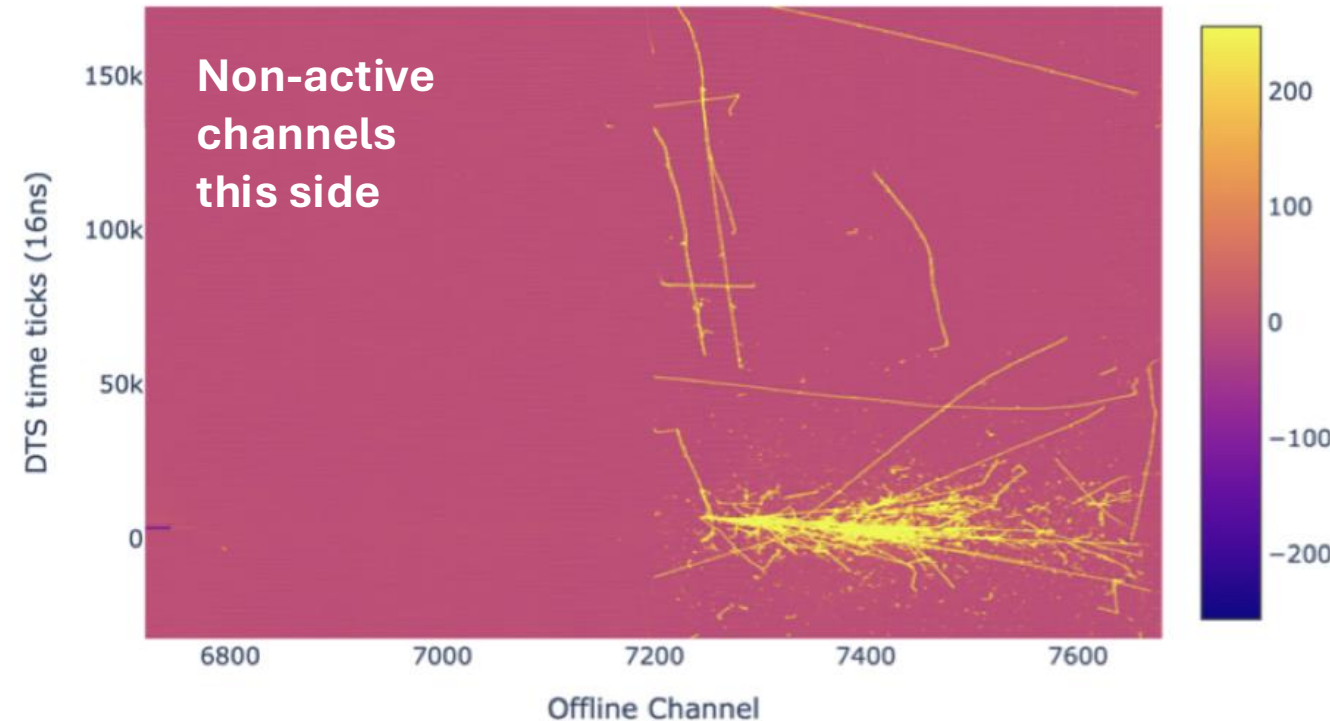


Data Overview at ProtoDUNE-HD

- Collected **~1 week of data**
- Trigger algorithm that selects **high energy events parallel to the cathode**
- Need to self-trigger detector – **rare events** and long **4.8s beam spill**
- Data taken non-intrusively – no effect on other CERN operations

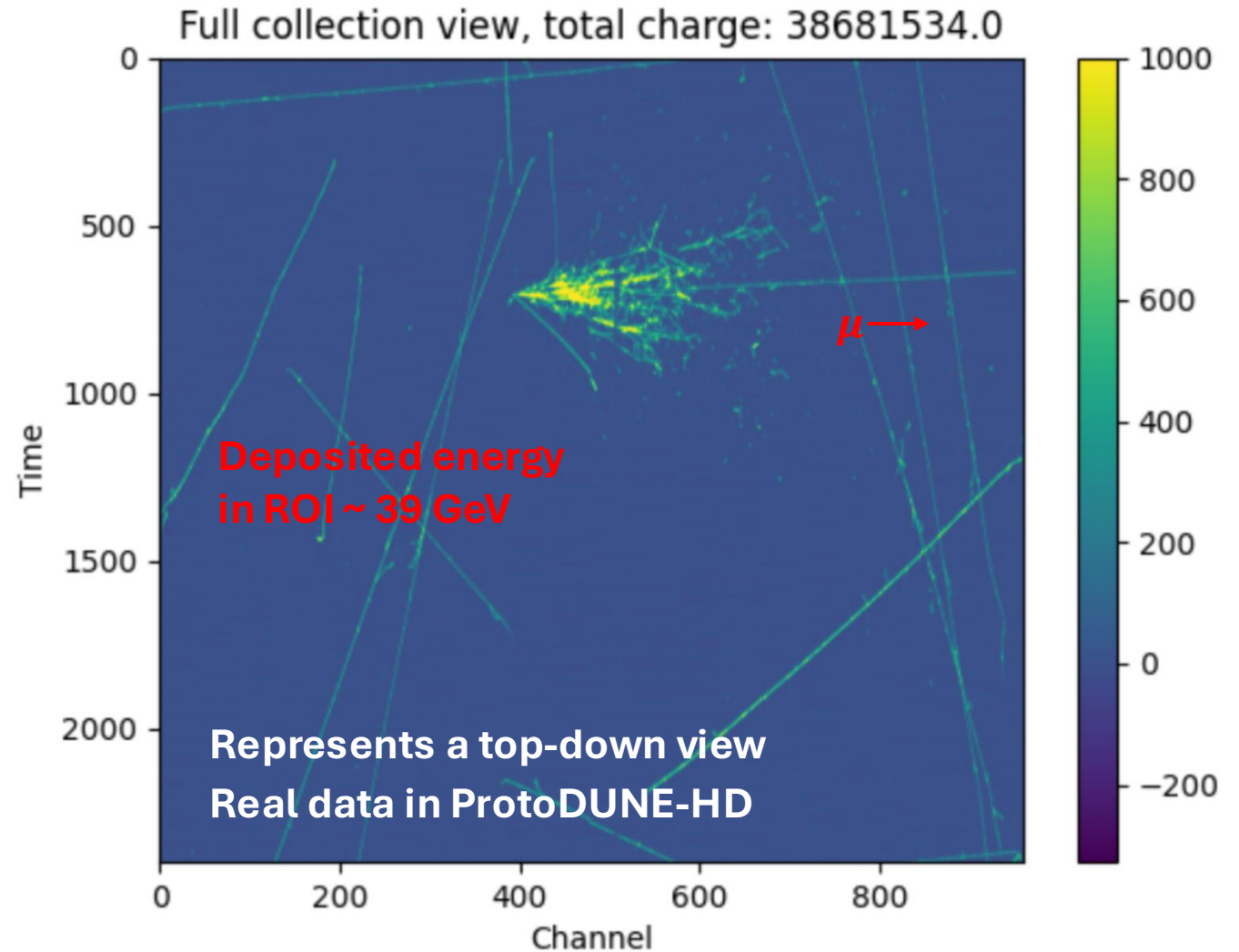
Run 29416, Trigger 401, APA2 Plane 2

Trigger Type (Supernova), 2024-10-03 12:30:14+02:00 (CERN)



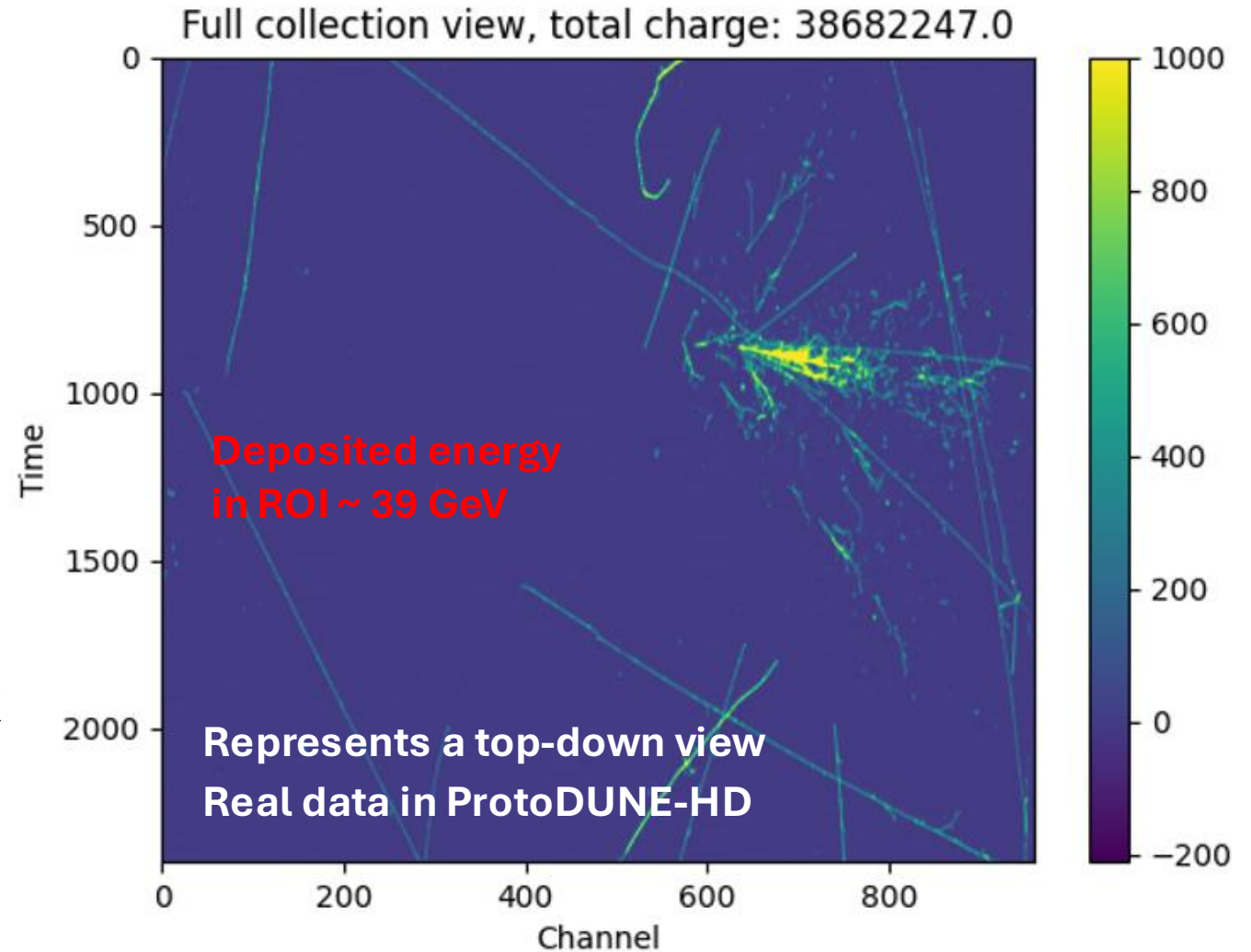
First Neutrino Candidates in Data

- Here is an identified **neutrino candidate**
- Identified by initial eye-scanning process by CERN Student Dario Pullia
- **Real data taken when the SPS beam spill was on**
- Forward-going shower with vertex within fiducial volume



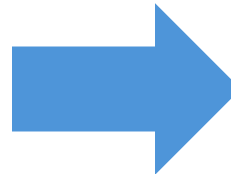
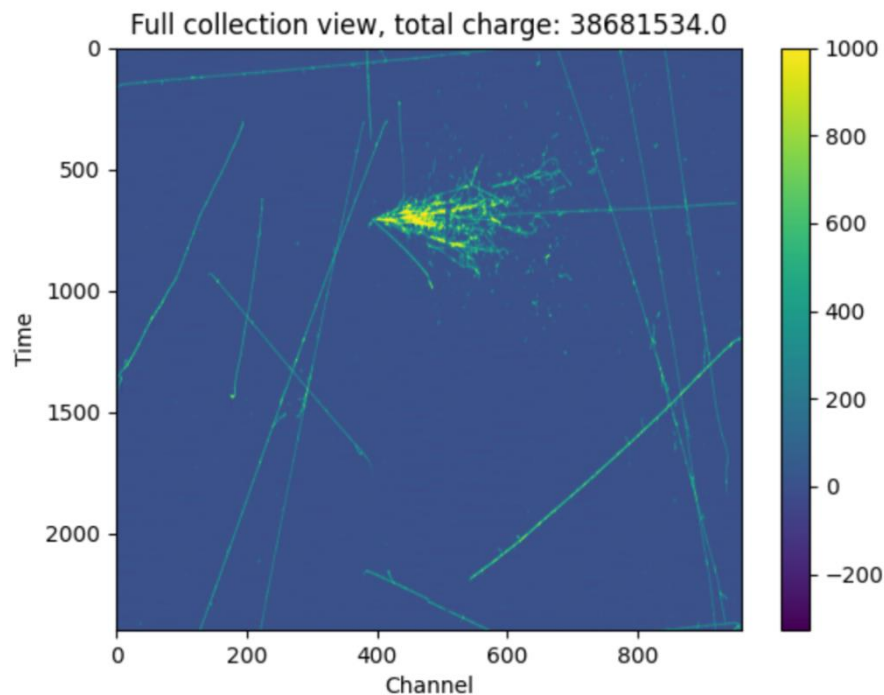
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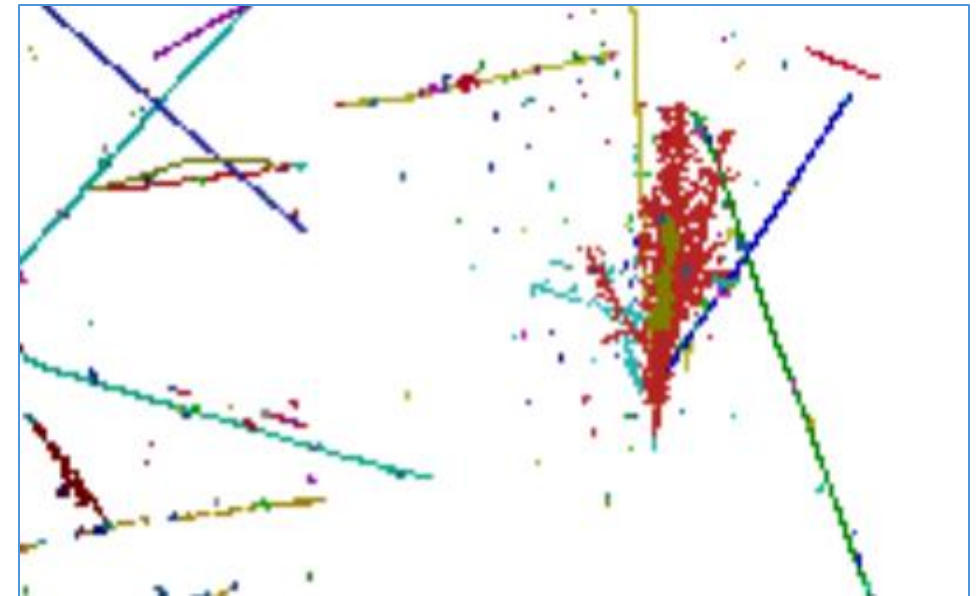


Data Processing to Reconstruction

- End-to-end **data processing implemented**
- Data has been **reconstructed with Pandora** – analysis with cut-based neutrino selection near completion



Real Pandora Reconstructed Data

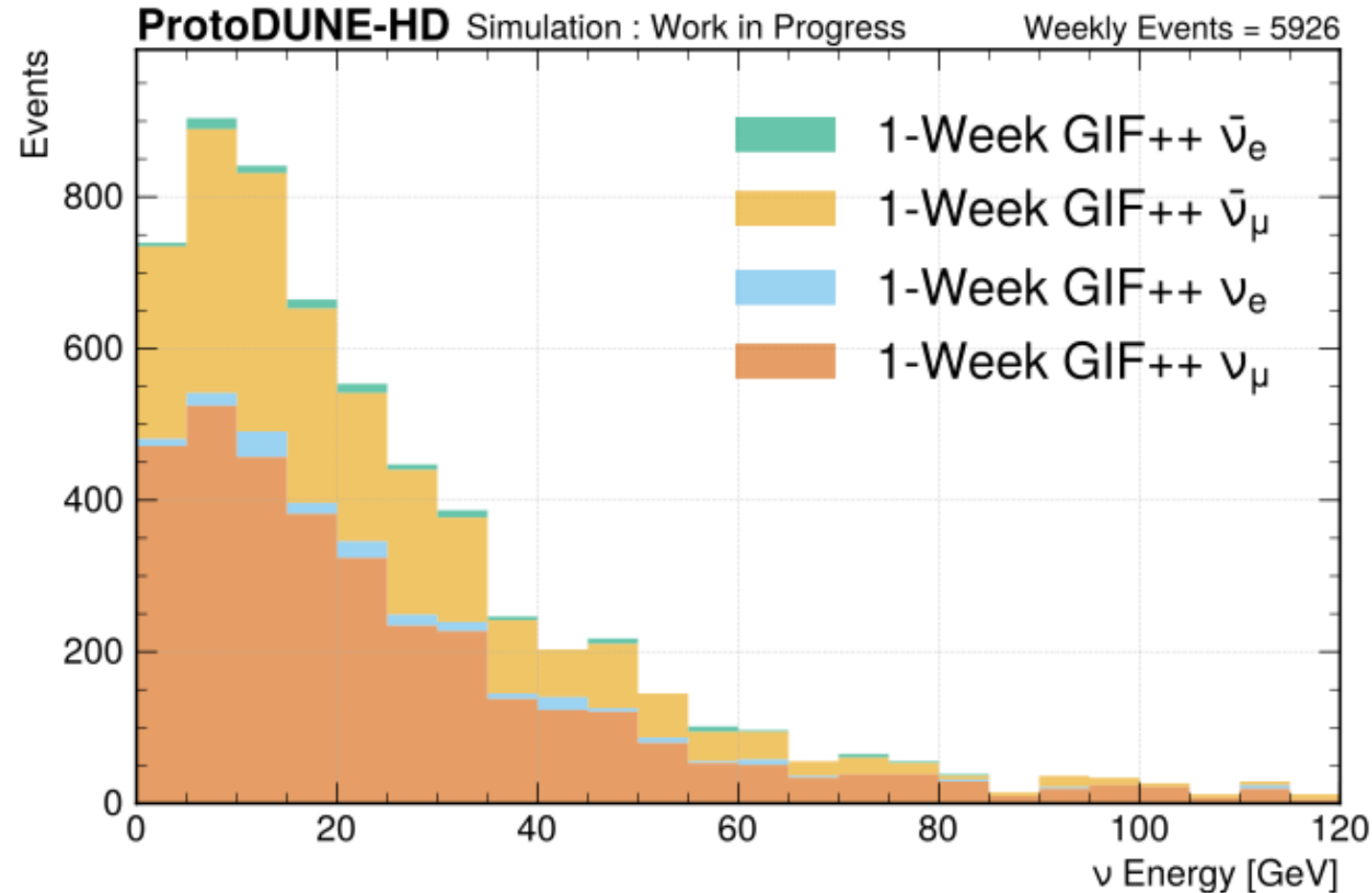


Neutrino MC Event Simulation

- Neutrino flux passed to **GENIE**
- neutrino interactions in a G4 model of ProtoDUNE-HD
- **Thousands of interactions in active volume per week**

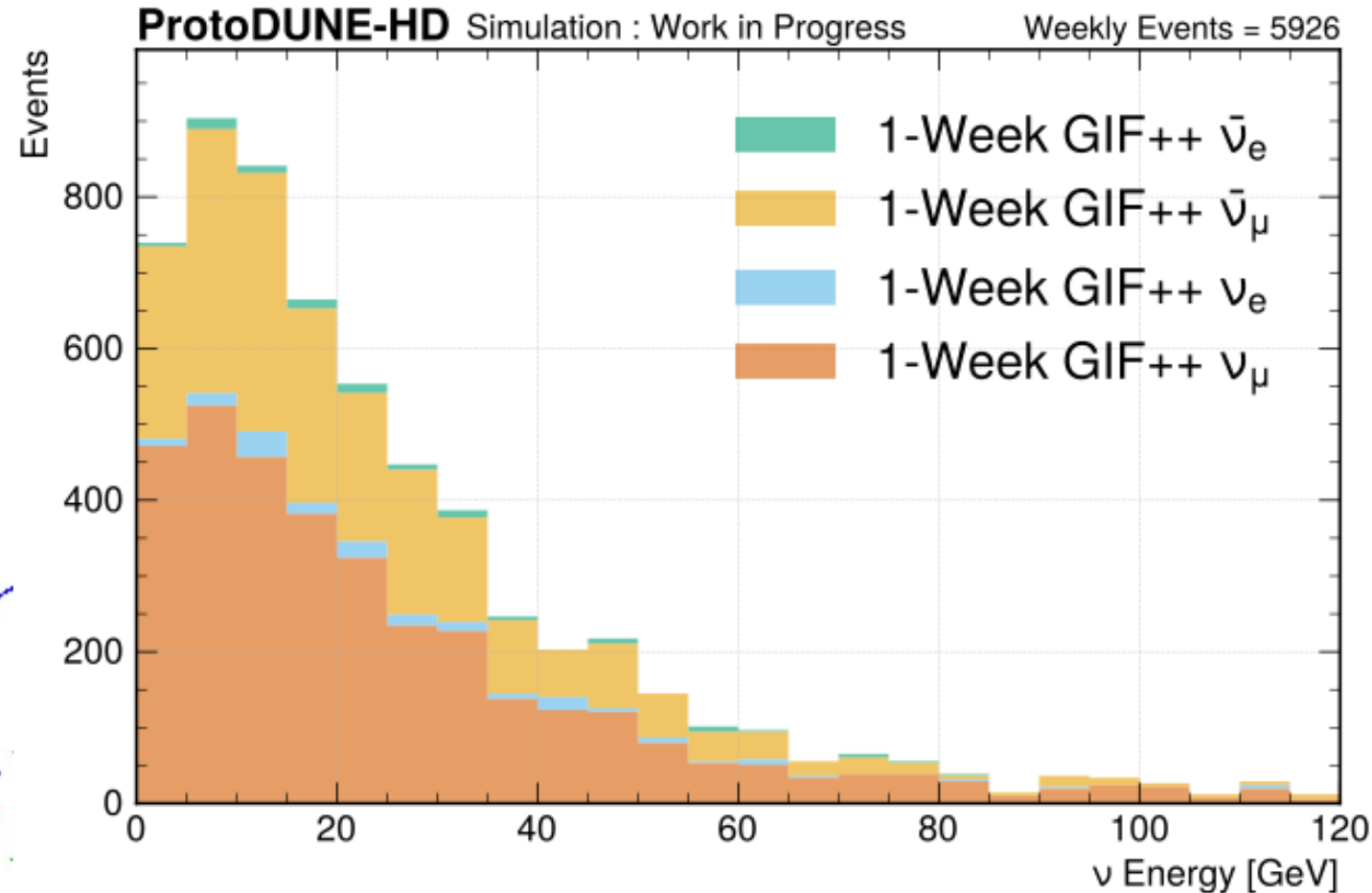
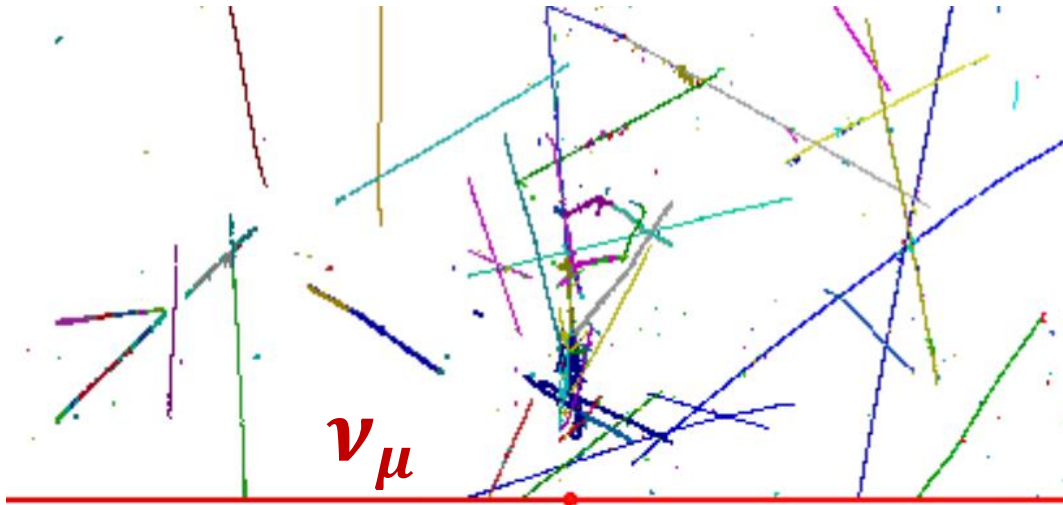
Magnet Conf.	Weekly interactions		
	w000	w133	wNP04
$\nu_{\mu}^{-40}\text{Ar} - \text{CC}$	2983	464	2597
$\nu_{\mu}^{-40}\text{Ar} - \text{NC}$	955	150	823
$\bar{\nu}_{\mu}^{-40}\text{Ar} - \text{CC}$	1248	305	1574
$\bar{\nu}_{\mu}^{-40}\text{Ar} - \text{NC}$	495	121	615
Total	5681	1040	5609

J. Hernandez Garcia



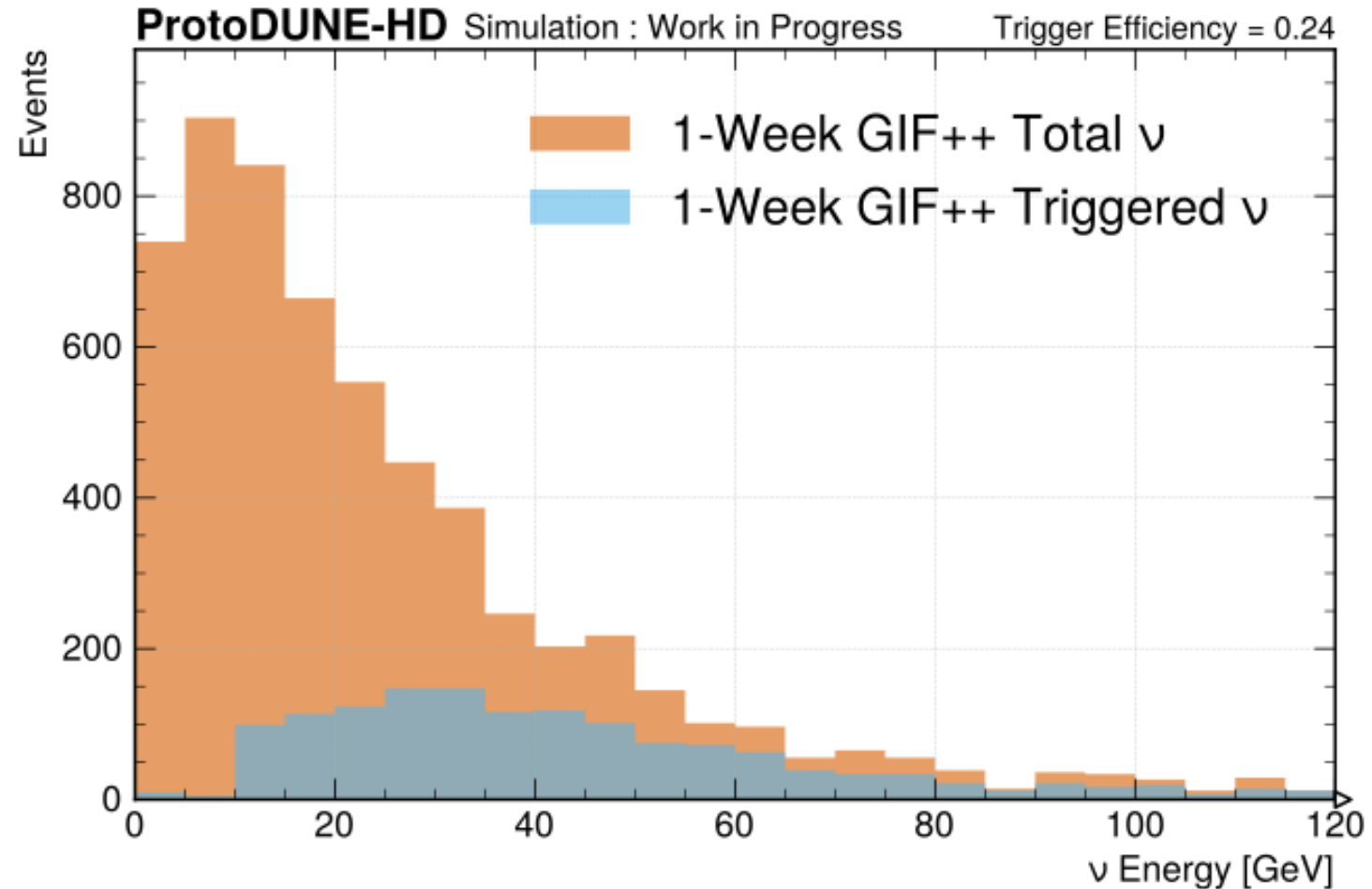
Neutrino MC Event Simulation

- Neutrino flux passed to GENIE - neutrino interactions in a G4 model of ProtoDUNE-HD
- Simulated neutrinos with cosmic overlay **reconstructed with Pandora**



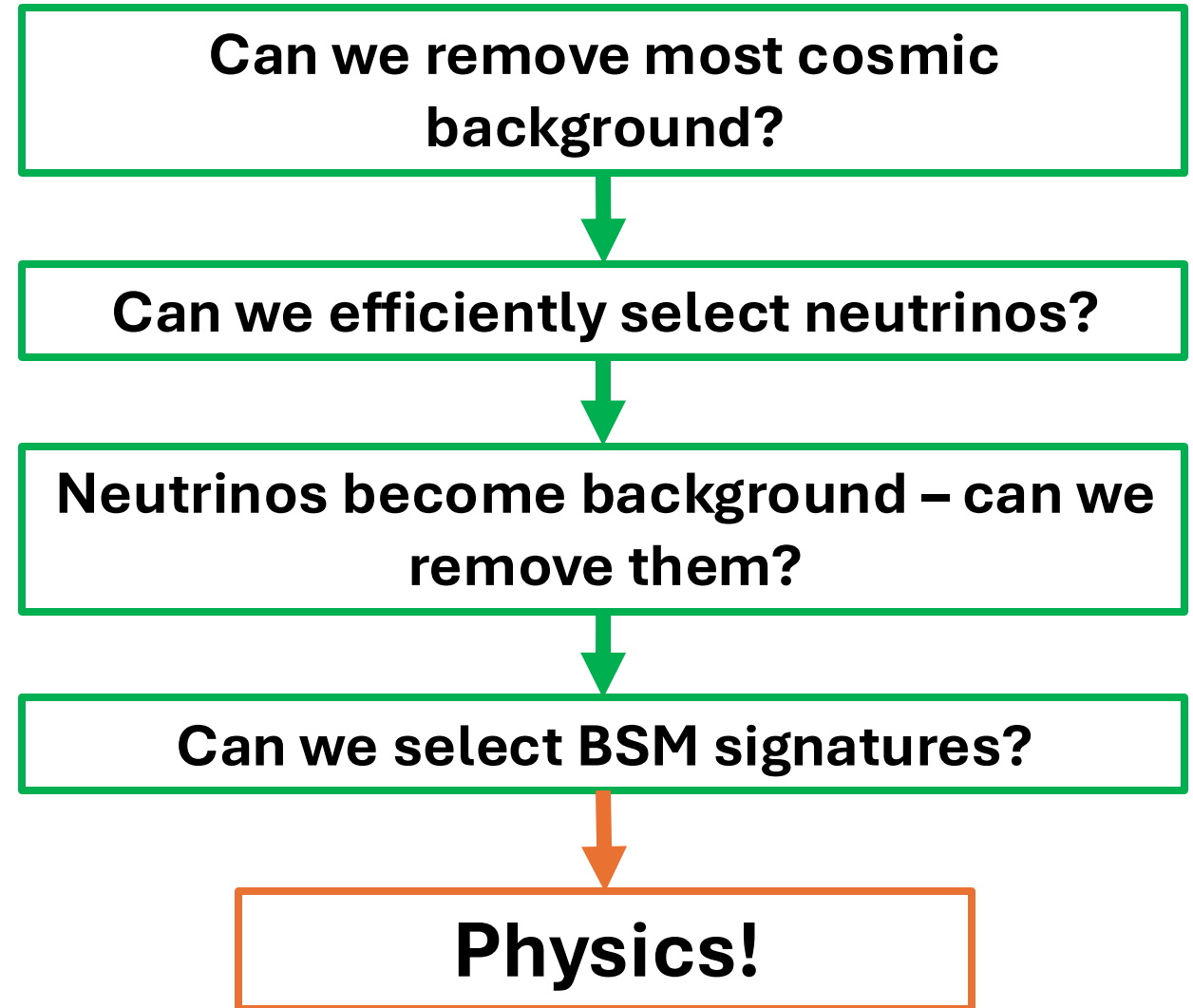
Estimate Rate of Triggering Neutrinos with Simulation

- Simulation has realistic detector and trigger simulation
- Estimate **proportion of neutrinos we trigger for**
- Same trigger algorithm simulated as used for data
- Achieve **~24% trigger efficiency** for ν and $E > 10$ GeV
- ~1400 triggering neutrinos per week with this magnet configuration



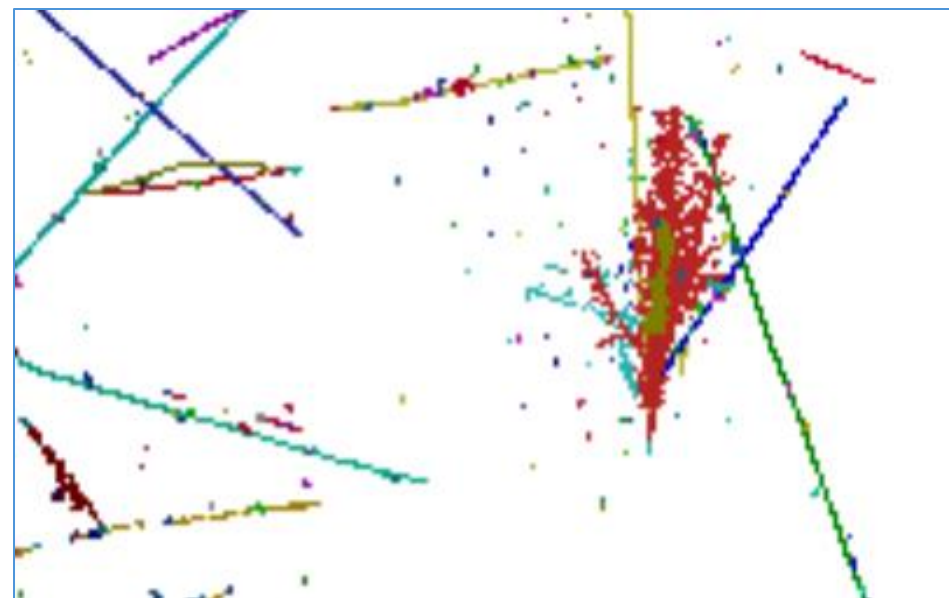
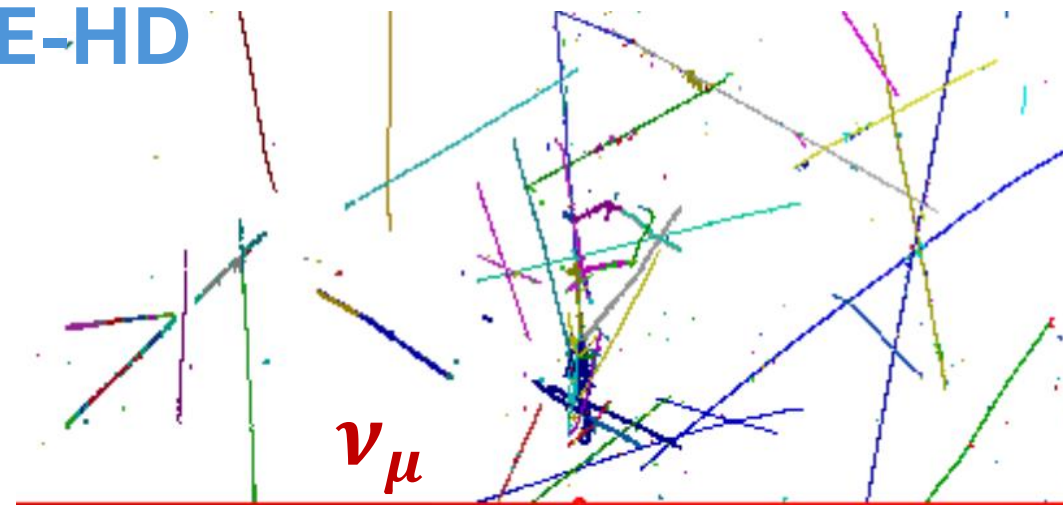
Analysis Strategy for ProtoDUNE-HD

- Trigger due to a cosmic muon **1-2 times per second**
- Trigger due to **neutrino once every few minutes** (depending on magnet configuration)
- First remove dominant cosmic background
- Measure our cosmic background rate from our **Spill OFF data**



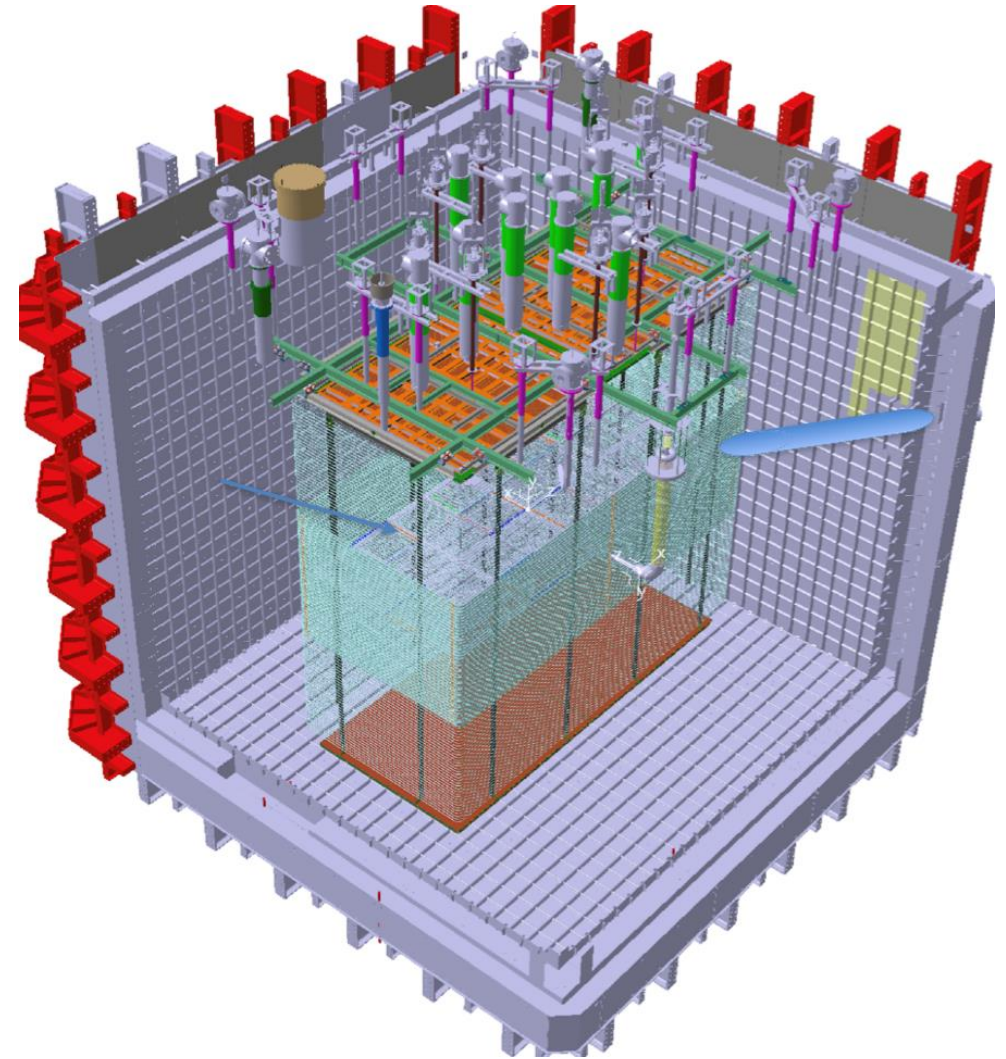
Analysis Strategy for ProtoDUNE-HD

- Fortunately, neutrinos look quite **different from most cosmics**
- Event selection aims to show an **excess of beam-aligned high-energy showers** arising from a vertex
- Selection also under development to distinguish HNL decays from neutrinos



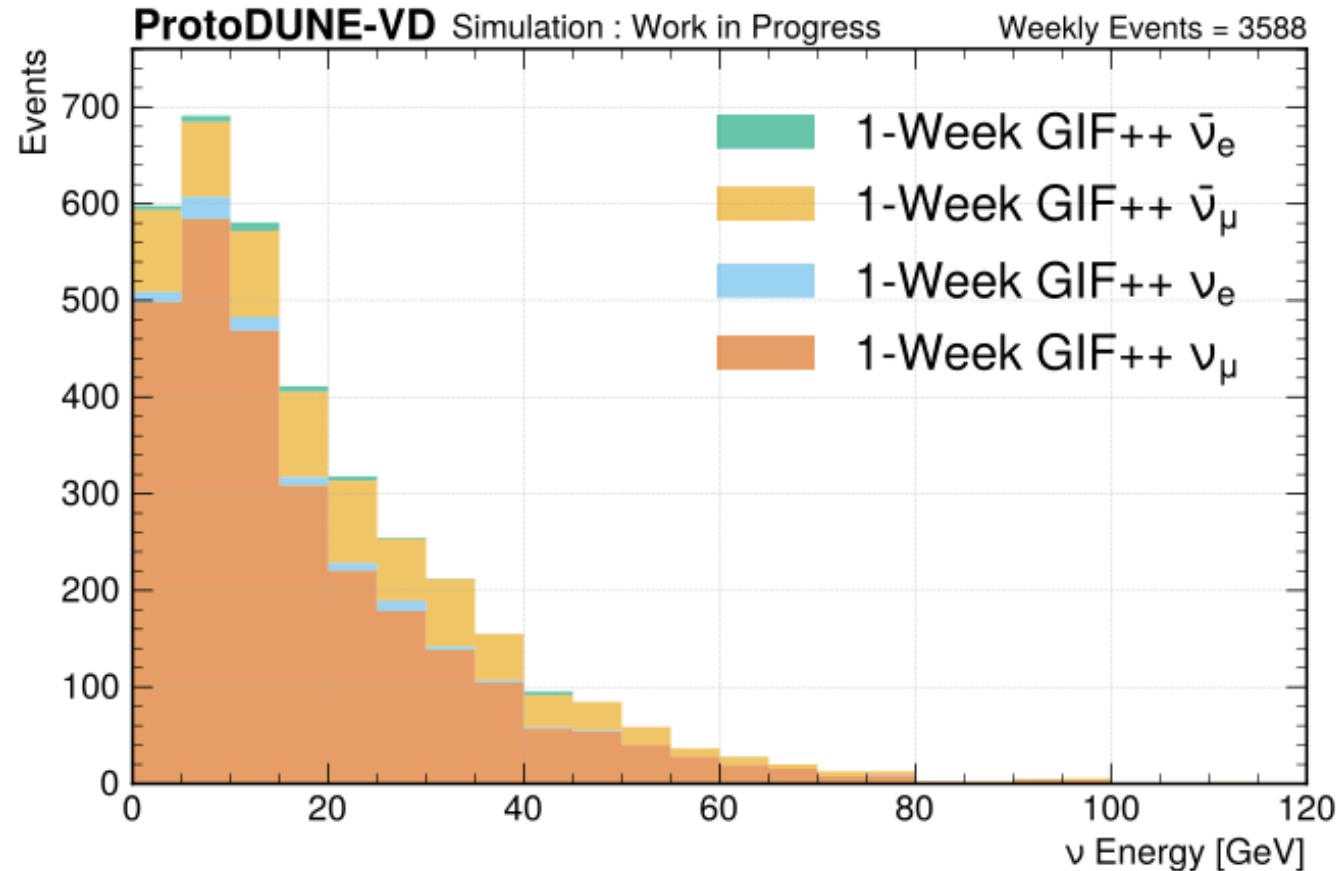
Looking to the Future – Vertical Drift

- Argon has been moved to **ProtoDUNE-VD** and is **taking data now!**
- Investigate searching for **neutrinos/BSM physics in NP02**
- Take data whilst the SPS is on
- Simulate neutrinos to estimate the rate
- Event selection to search for beam-aligned showers from a vertex



Looking to the Future – Vertical Drift

- Neutrino flux simulation repeated for ProtoDUNE-VD
- Similar position off-axis, but fiducial volume ~70% the size of ProtoDUNE-HD
- Still estimating **thousands of interaction in active detector volume per week of SPS beam**
- Expect similar trigger efficiency
- Developing an upgraded ML-based trigger algorithm



Conclusions

- Opportunity to search for neutrinos and BSM physics at ProtoDUNE detectors
- These searches can be run non-intrusively
- We have **~1 week of data** with at ProtoDUNE-HD – seen first neutrino candidates in ProtoDUNE!
- End-to-end **data processing for neutrino search complete**
- End-to-end **neutrino simulation complete**
- Aim to neutrino search analysis first as a **proof-of-principle of a future BSM program**
- Large sample of neutrinos in a DUNE FD-like detector very useful for testing energy reconstruction methods for the future FD

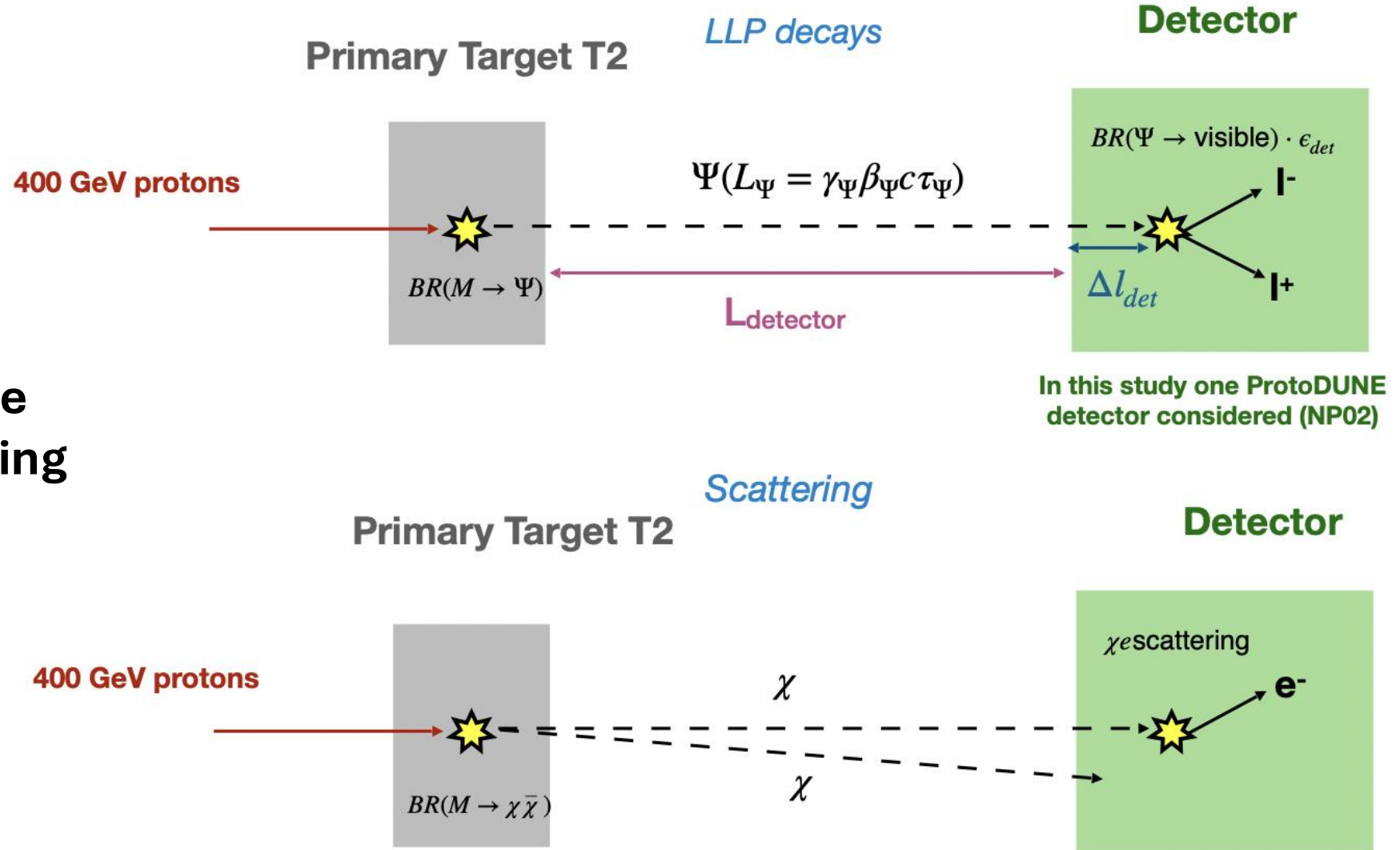
Thank you for listening!

The ProtoDUNE-BSM Group:

H. Amar, C. Andreopoulos, A. Chatterjee, S. Bianco, E. van Campenhout, P. Coloma, A. De Roeck, G. Gurung, C. Hasnip, J. Hernández García, W. Ketchum, F. Lanni, J. López-Pavón, J. Martin-Albo, L. Molina-Bueno, O. Palamara, J. K. Plows, F. Pietropaolo, D. Pullia, F. Resnati, P. Sajitha, H. Sieber, C. Touramanis, S. Urrea

Backup

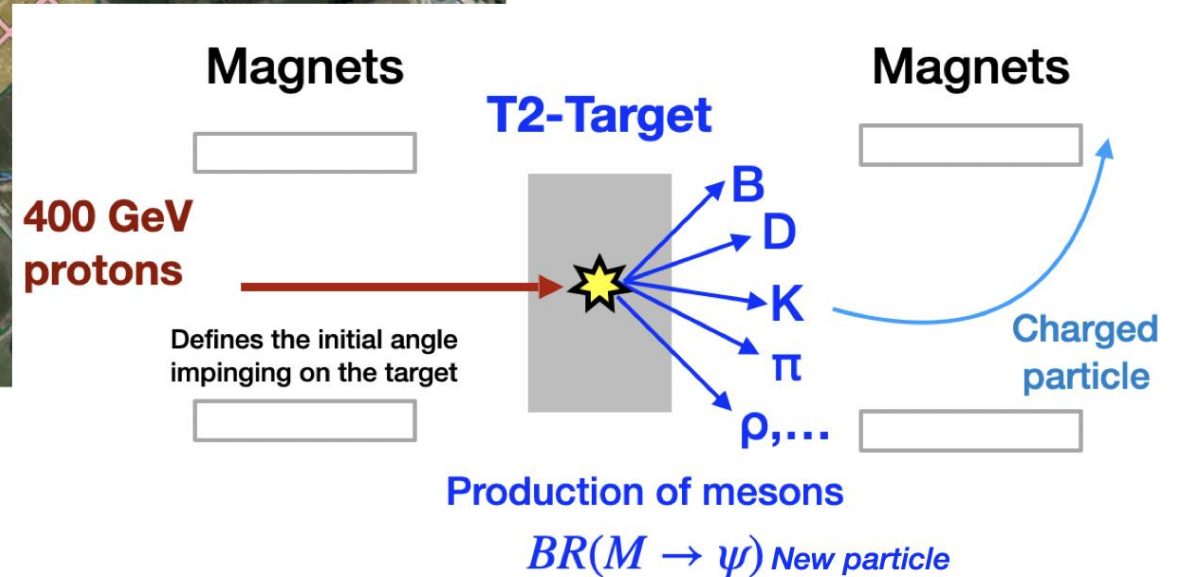
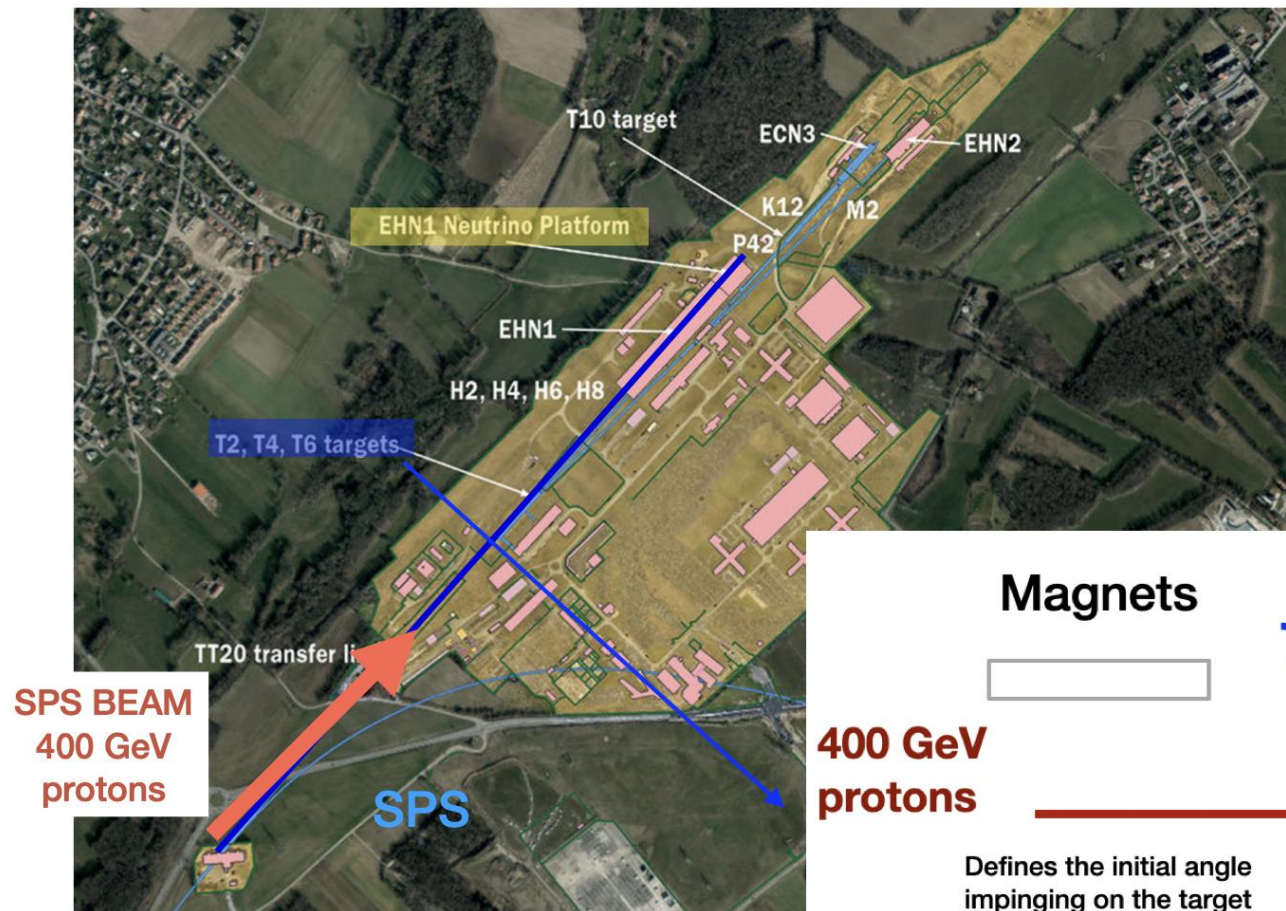
Could investigate
decay or scattering
channels



P. Coloma, J. López-Pavón, L. Molina-Bueno and S. Urrea, JHEP 01 (2024), 134 doi:10.1007/JHEP01(2024)134

Backup

From slides by L. Molina-Bueno

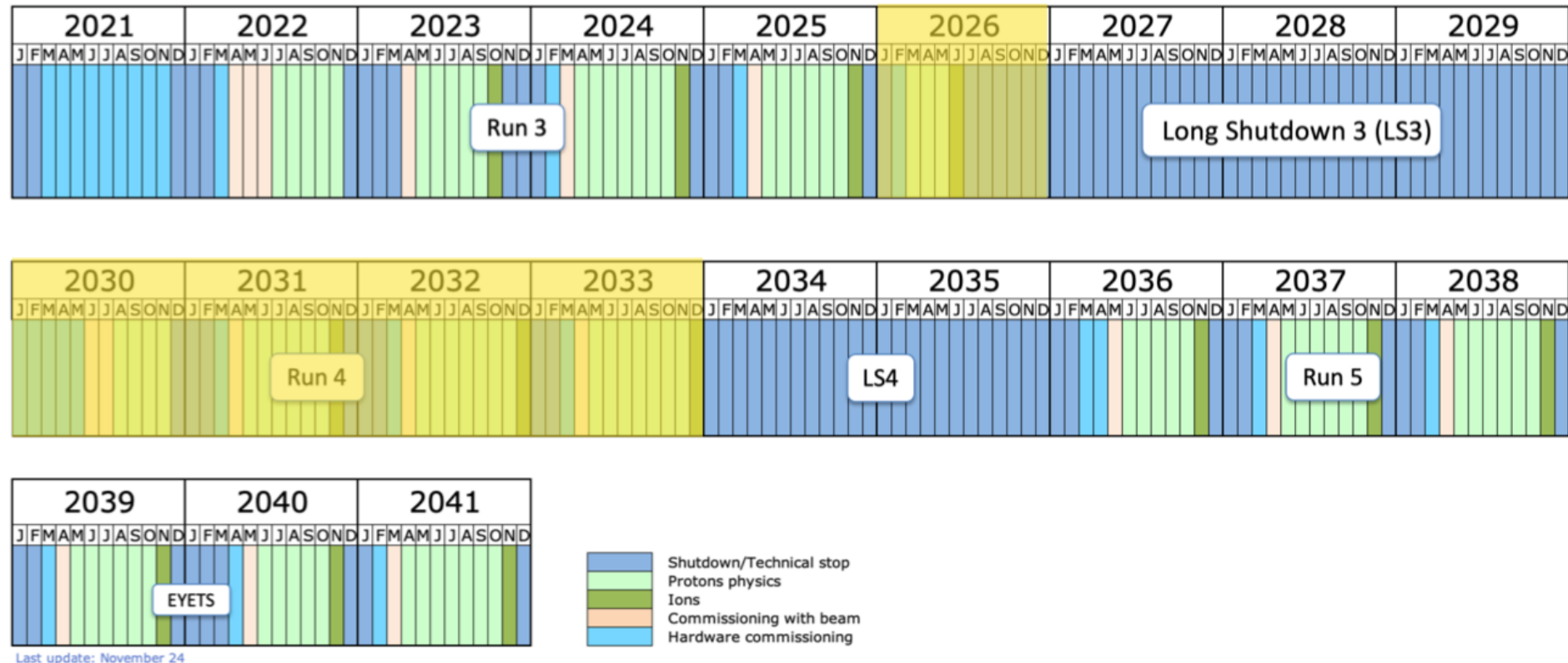


The current on these magnets depend on H4/H2 user
They are called “wobbling”

Backup: Possible Schedule for BSM Program

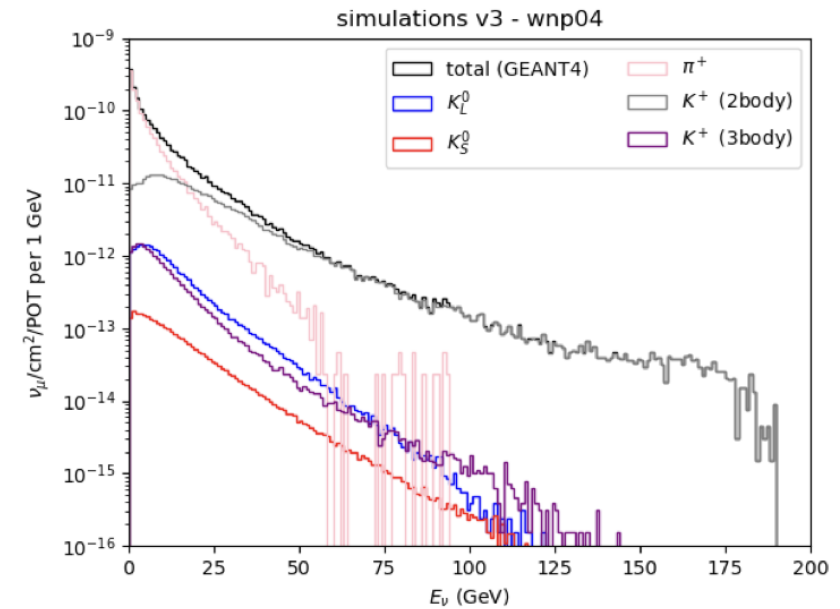
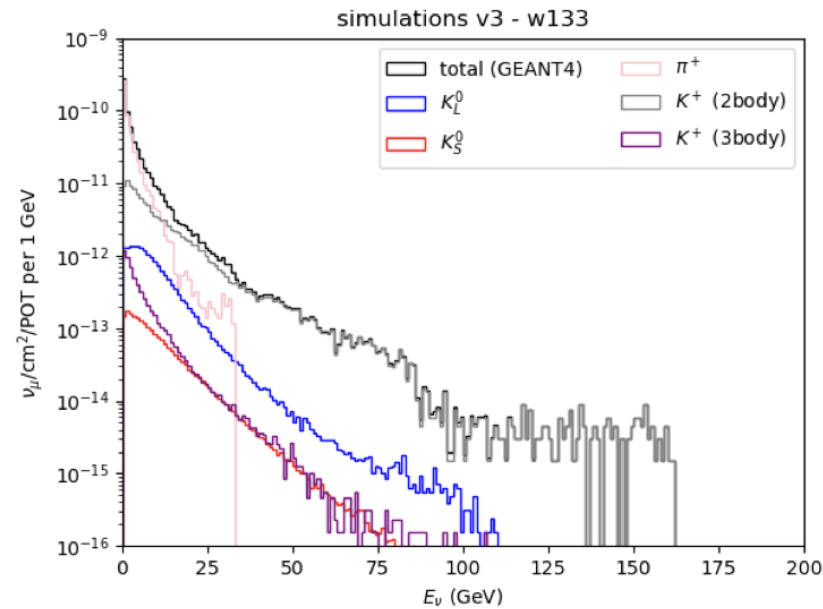
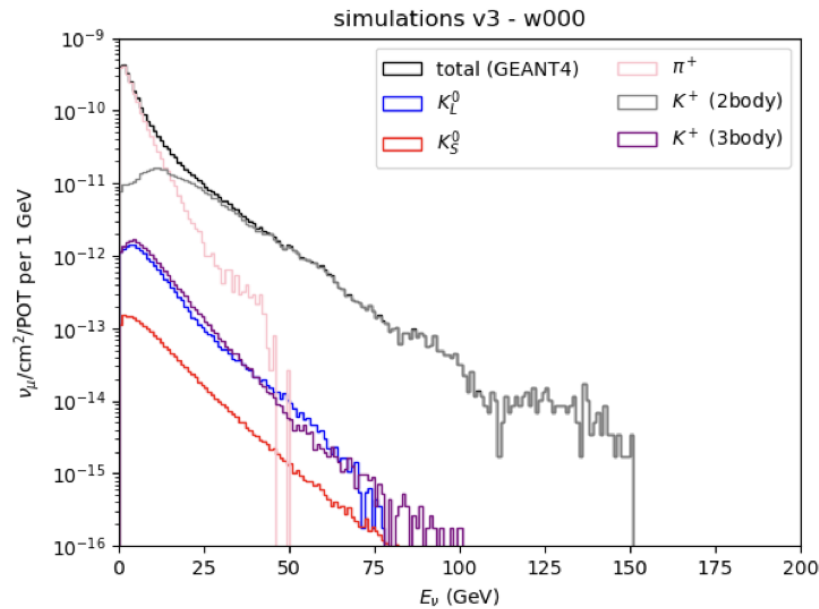
- Run parasitically at SPS (not interfering with other experiments at H4/H2)
- Proton run at SPS can be up to 6 months depending on the year
- We consider 5 years to collect 1.75×10^{19} POT

From: <http://lhc-commissioning.web.cern.ch/schedule/LHC-long-term.htm>



Backup: Possible Schedule for BSM Program

- Separate neutrino flux simulation for each magnet wobbling configuration



Data Processing Procedure

- End-to-end **data processing implemented** and running through to fully reconstructed events
- Reconstruction uses **FD neutrino pandora + PD-HD cosmic tagging**
- Produce a flag for whether SPS spill was ON/OFF
- Filter for **ground shake removal**
- 1 week of data processed to reconstruction and cut-based analysis on-going

