



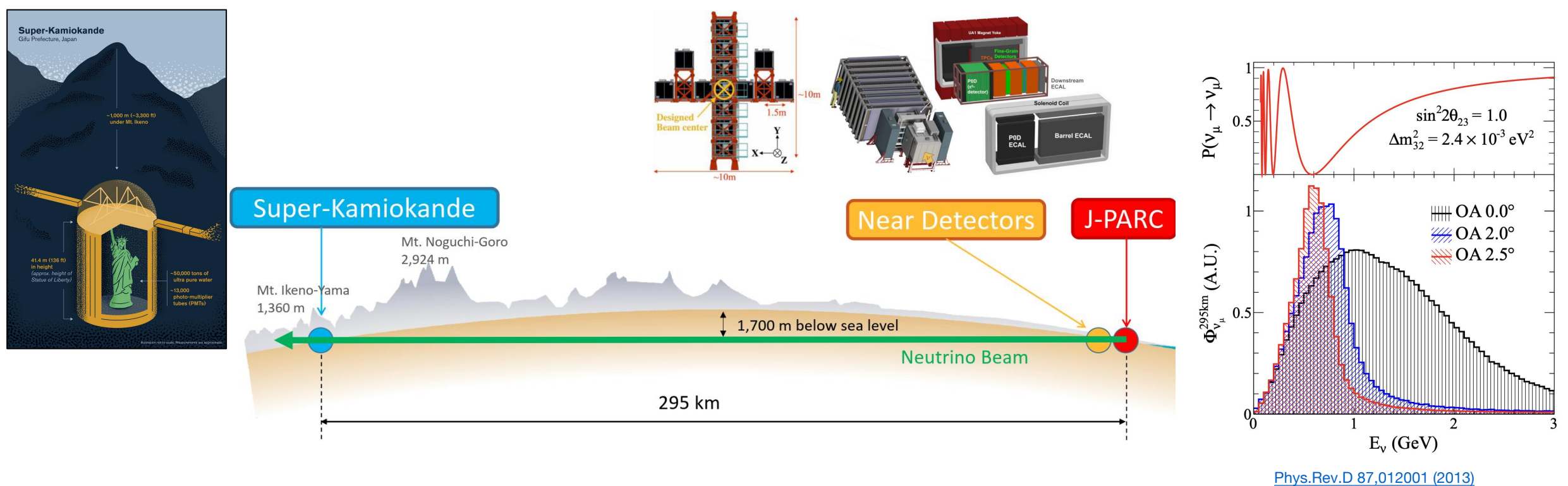
# NEUTRON PRODUCTION IN NEUTRINO-NUCLEUS INTERACTIONS WITH T2K

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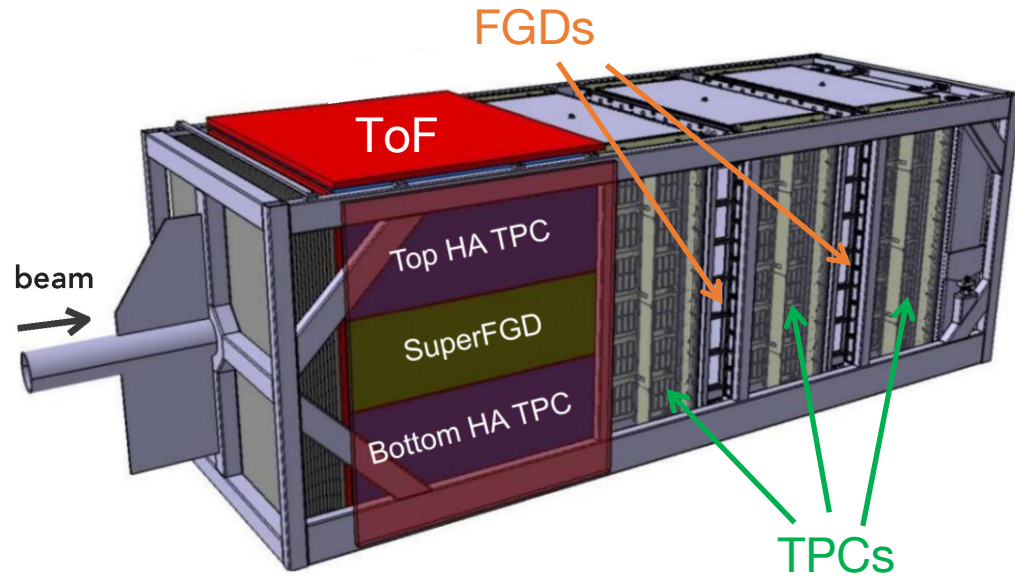
**XXI International Workshop on Neutrino Telescopes**

# The T2K Experiment



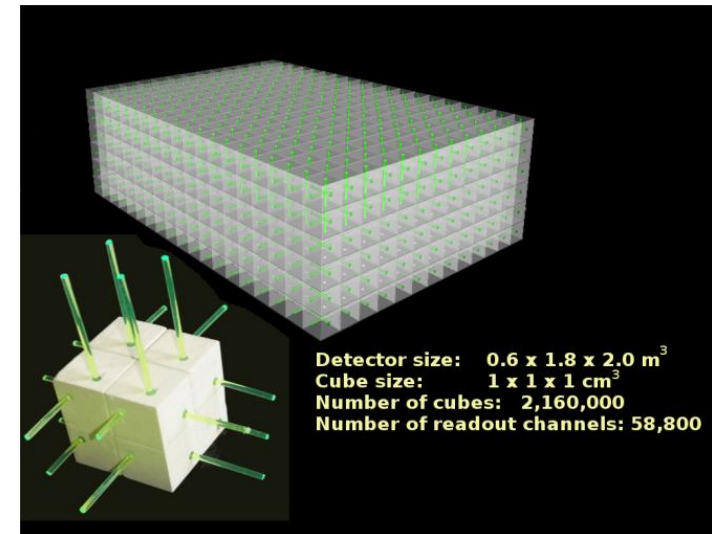
- Long-baseline experiment (295 km)
  - Near detectors (ND280, INGRID, WAGASCI-BabyMIND) at 280m from J-PARC
  - $E_\nu$  peaks at 0.6 GeV at ND280 (2.5° off-axis)
- Related Talks: [S. King](#), [L. Machado](#)

# Near Detector ND280 Upgrade



- HA-TPC does the identification and reconstruction of particles traveling at high angles.
- TOF, along with SFGD, can differentiate between incoming and outgoing particles

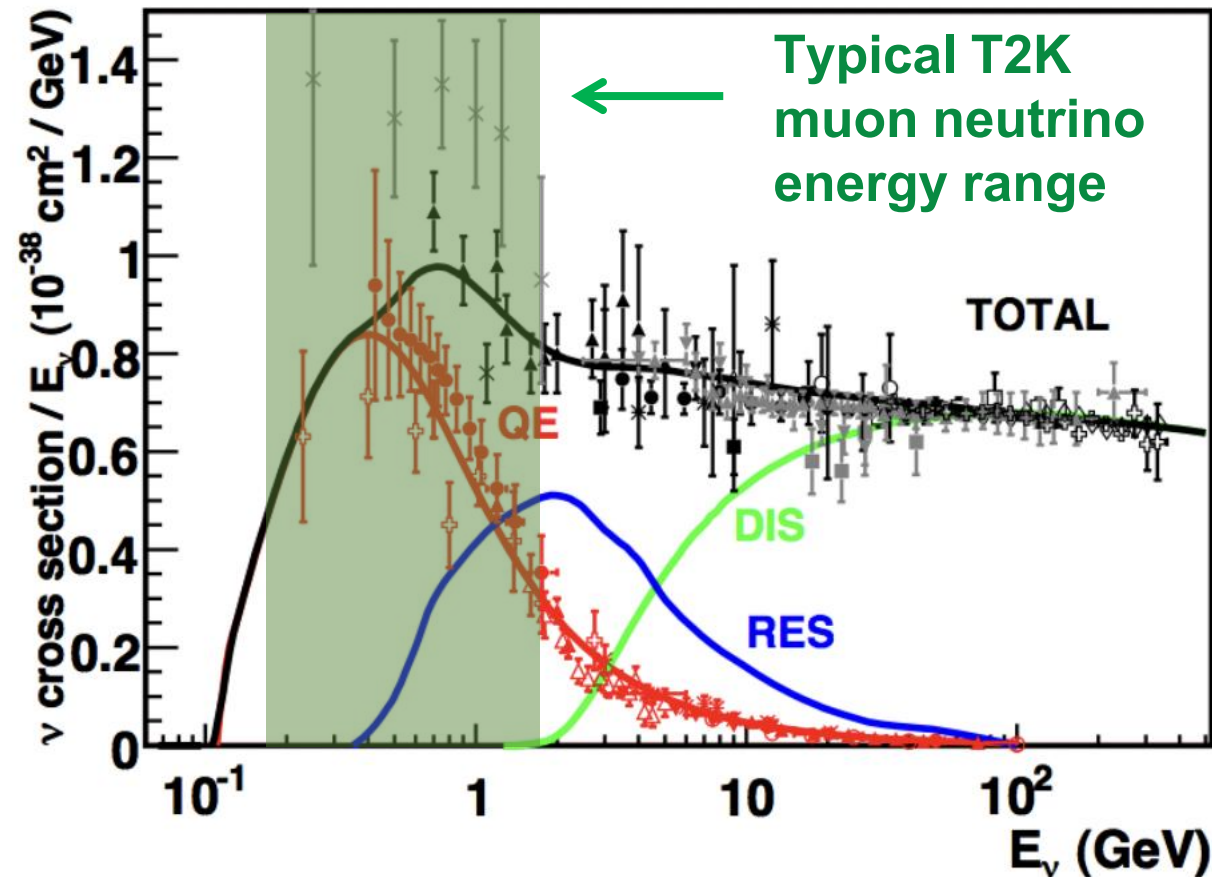
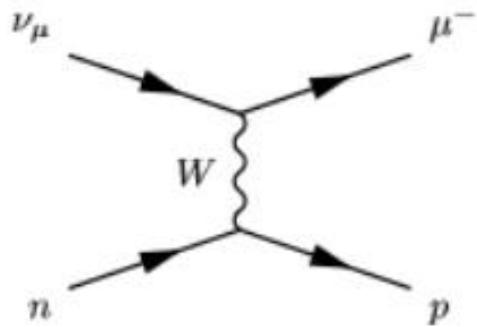
- Super-FGD is made of ~ 2 million scintillator cubes of size 1 cm which provide sub-ns timing resolution.



Related Talks: [G. Reina](#), [M. Feltre](#), [W. Okinaga-san](#)

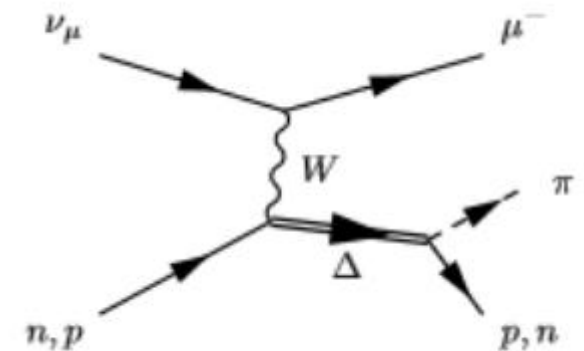
# Neutrino-Nucleus Interactions

Quasi-Elastic Scattering (CCQE)



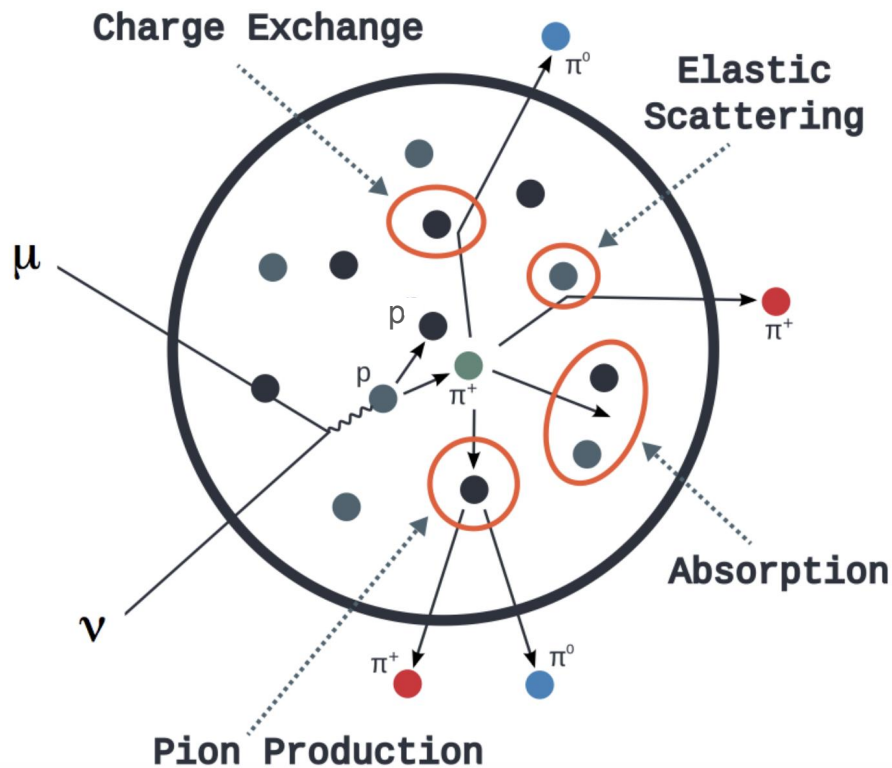
<https://doi.org/10.2172/1042577>

Delta Resonance



Related Talk: [E. Sandford](#)

# Final State Interactions (FSI)



Picture Credits: T. Golan

- Particles produced after neutrino-nucleus interaction can undergo FSI.
- This makes identification of specific interactions in detector difficult.
- Identifying topology is a better choice as it reduces model dependence.

Some topologies:

CC0 $\pi$ , CC1 $\pi$ , CC-multi  $\pi$ , **CC0 $\pi$ Nn**

# Interactions in CC0 $\pi$ Nn Topology

**CCQE ( $\nu_\mu n \rightarrow \mu^- p$ ) : With FSI**

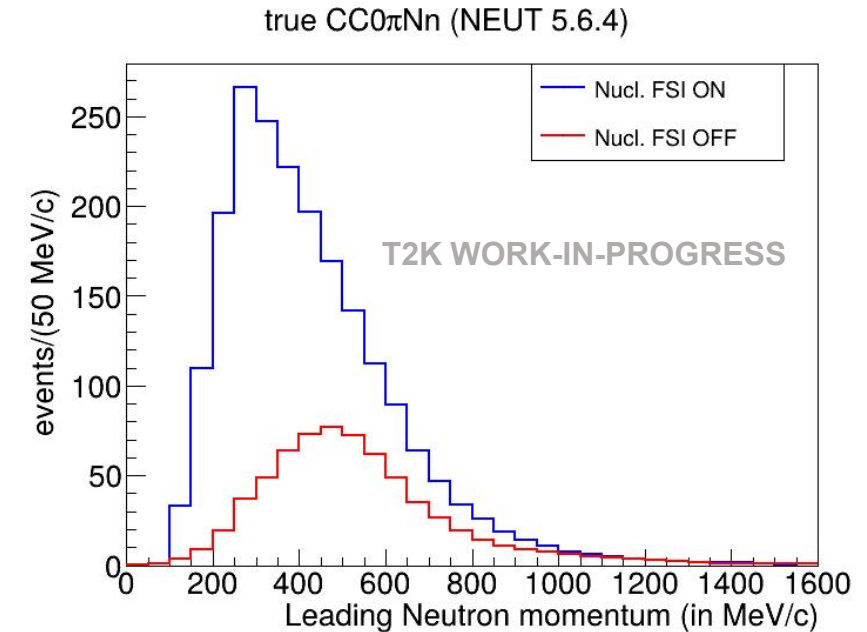
2p2h ( $\nu_\mu nn \rightarrow \mu^- np$ ) : With or without FSI

( $\nu_\mu np \rightarrow \mu^- pp$ ) : With FSI

Delta resonance : With or without FSI

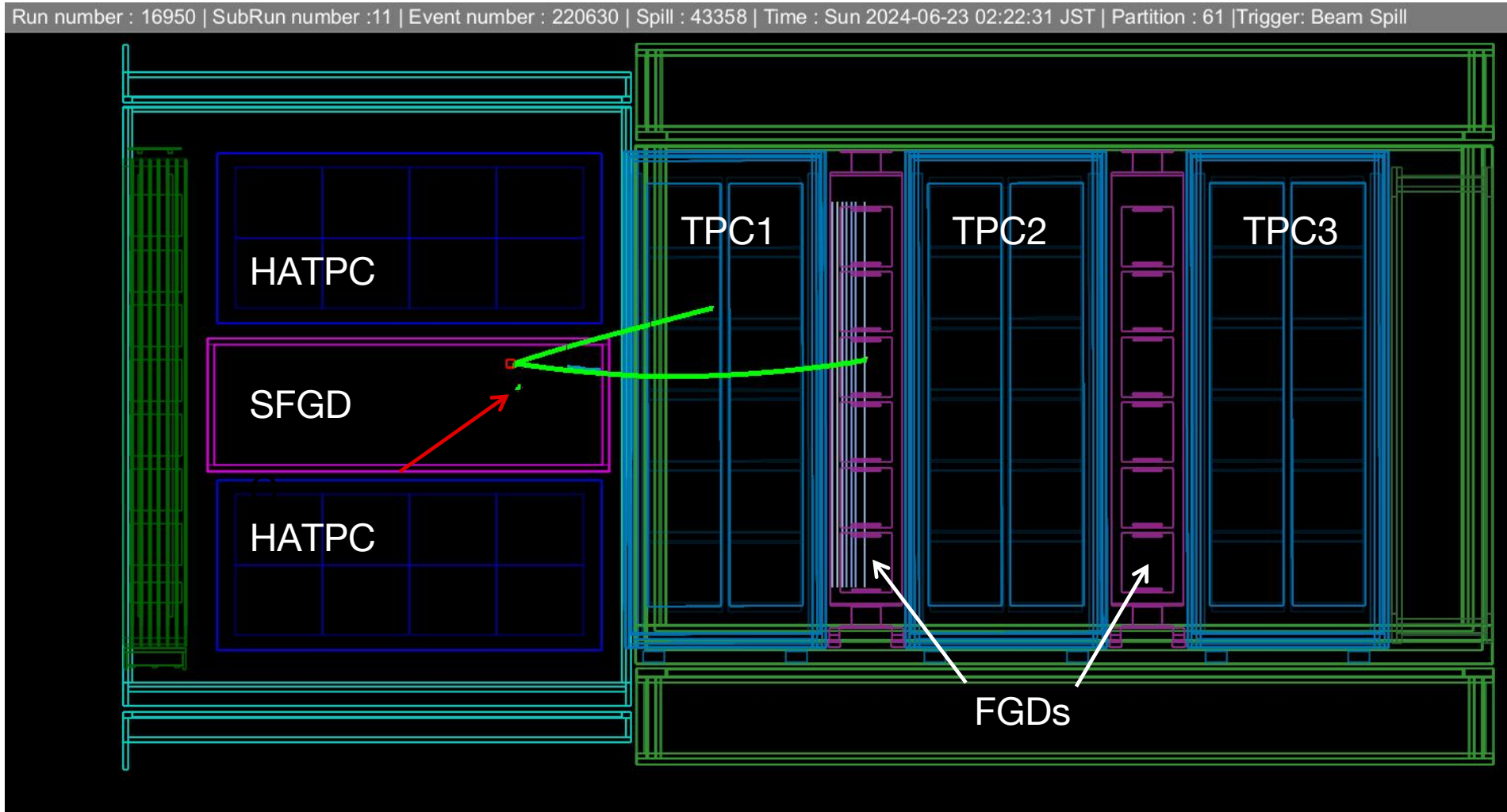
Type of interaction	% in CC0 $\pi$ Nn (before FSI*)	% in CC0 $\pi$ Nn (after FSI*)
CCQE	0% (0)	54.80% (1119)
2p2h	36.08% (245)	22.82% (466)
Delta res.	59.65% (405)	21.11% (431)
Others	4.27% (29)	1.27% (26)

\*FSI turned ON or OFF here is Nucleon FSI



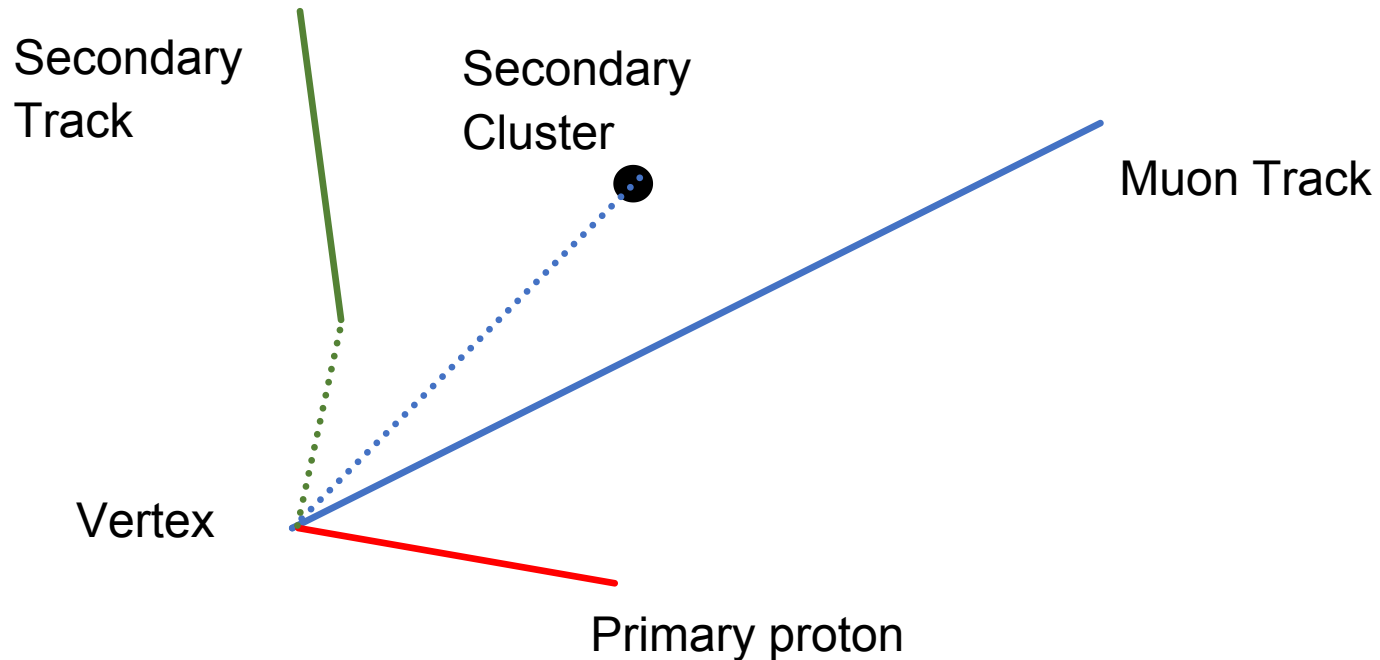
- Large increase in events due to CCQE and 2p2h undergoing FSI
- CC0 $\pi$ Nn topology can be used to test FSI models

# Neutron candidate in an actual event



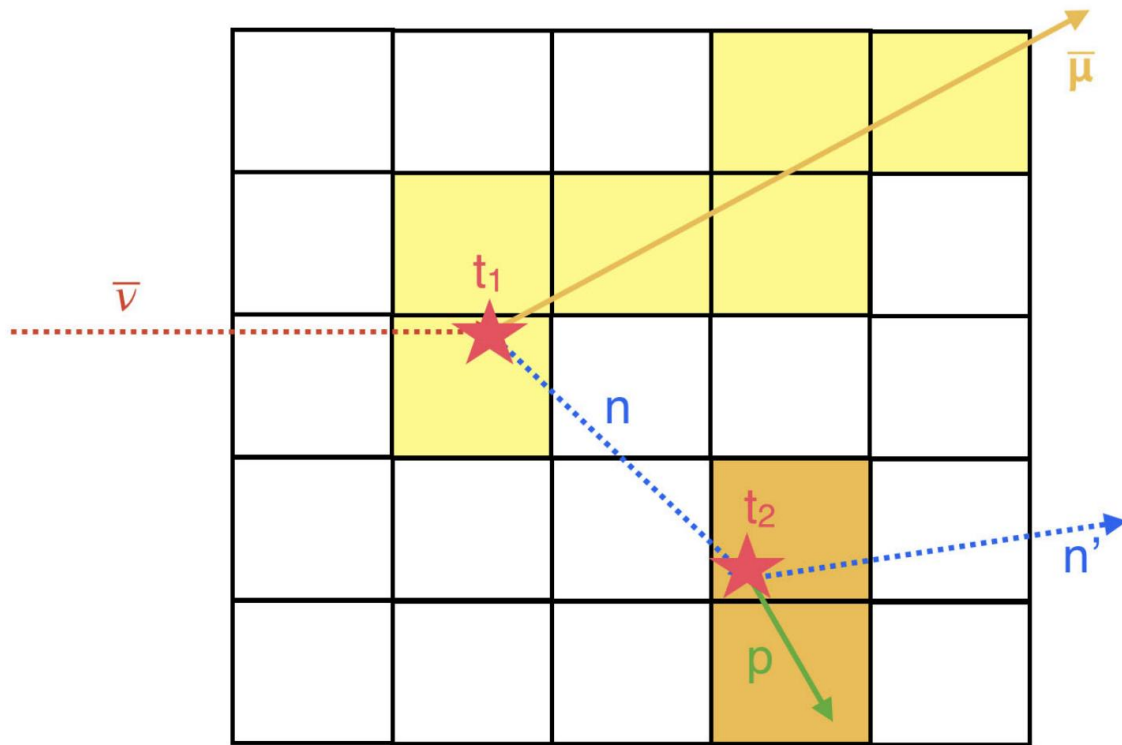
One neutron candidate in this event

# Selection of CC0 $\pi$ events



- One muon track
- No pion track
- Any number of proton tracks

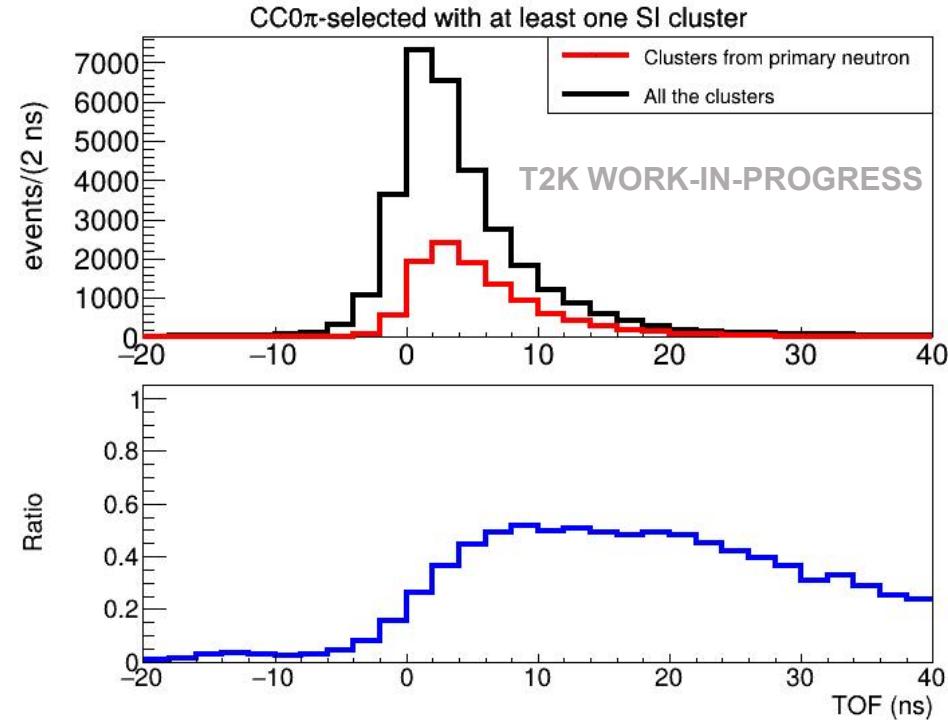
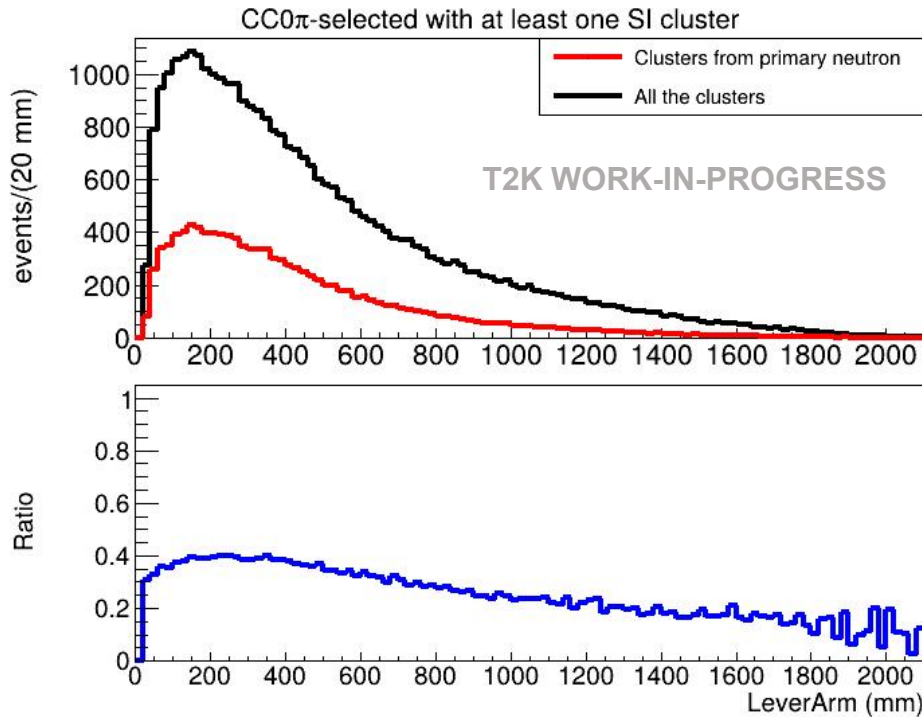
# Reconstructing a neutron



[Phys. Rev. D 101, 092003 \(2020\)](#)

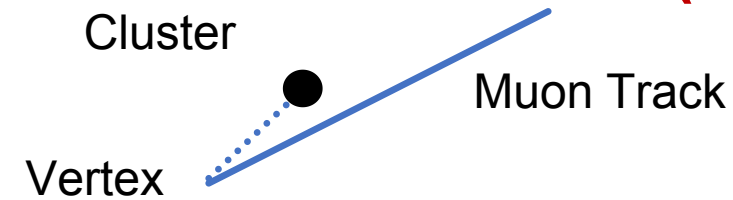
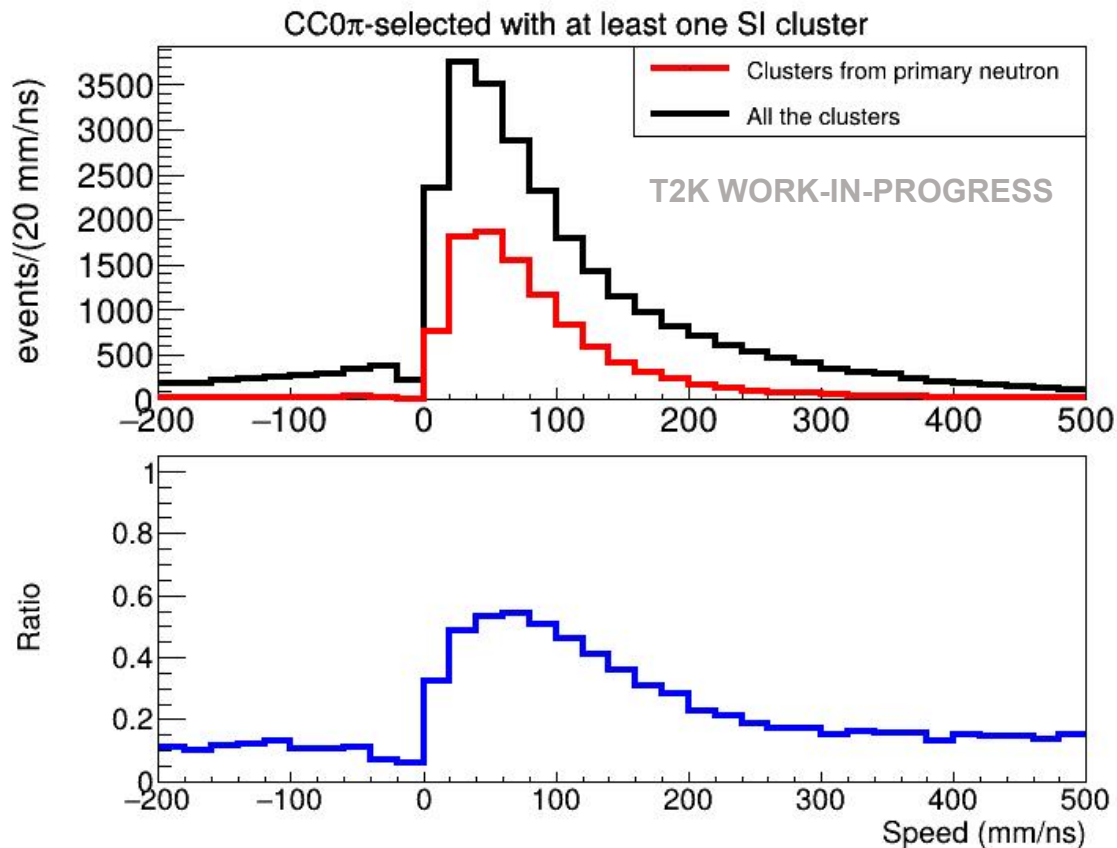
- Neutrons deposit energy away from the vertex.
- Lever arm and Time-of-flight are tools to identify a neutron and reconstruct its energy.

# CC0 $\pi$ pre-selected with secondary clusters



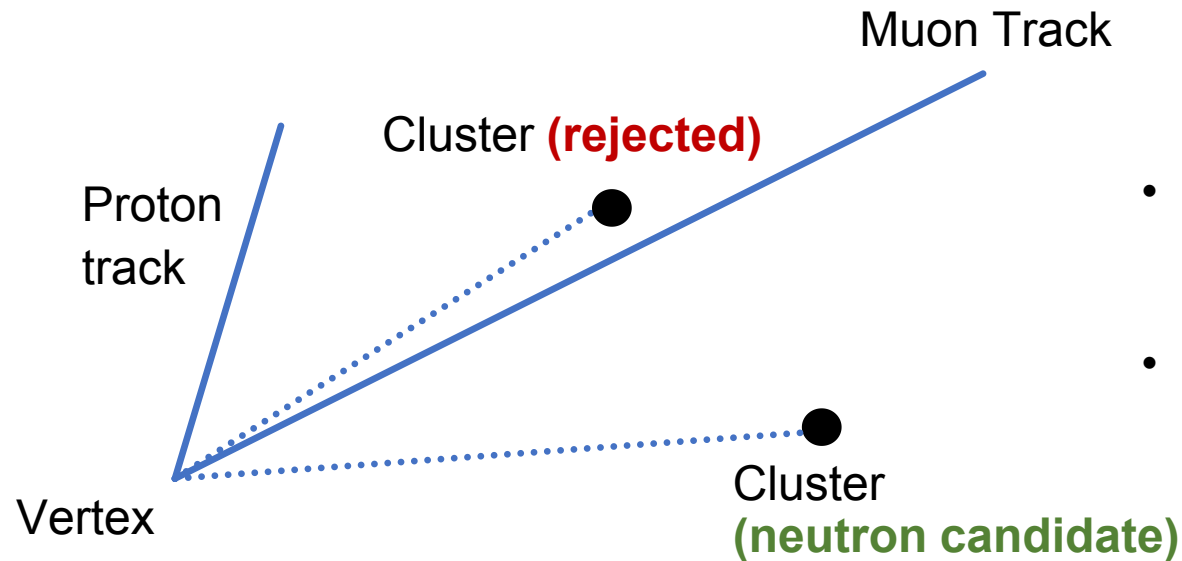
- Neutron (and background) cluster distribution peaks around 15 cm from the neutrino vertex.
- Time-of-flight can be negative due to timing resolution of SFGD. Applying a cut on the distribution, such as  $0 < \text{TOF (ns)} < 20$ , can provide a pure sample of neutrons.

# CC0 $\pi$ pre-selected with secondary clusters (speed)



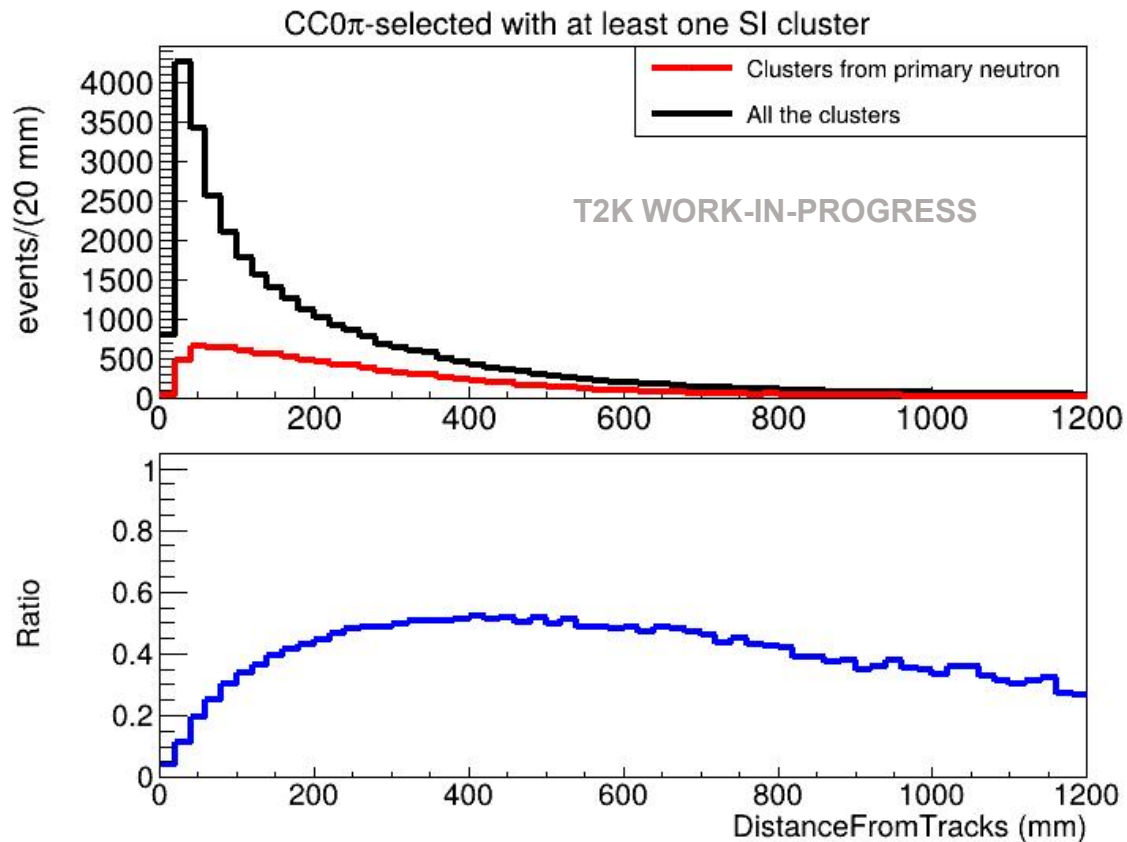
- Speed = Lever Arm / Time-of-Flight
- Reconstructed speed would not be more than 300 mm/ns (speed of light) if the detectors were perfect.
- Neutrons are heavy particles with low momentum, so speed peaks at around 20% of the speed of light ( $\sim 200$  MeV/c momentum) and significantly drops around 67% of the speed of light ( $\sim 850$  MeV/c momentum) for T2K flux.

# Minimum distance from all tracks in SFGD



- Clusters close to any track could be produced by that track and not primary neutrons.
- They should be rejected as neutron candidates.

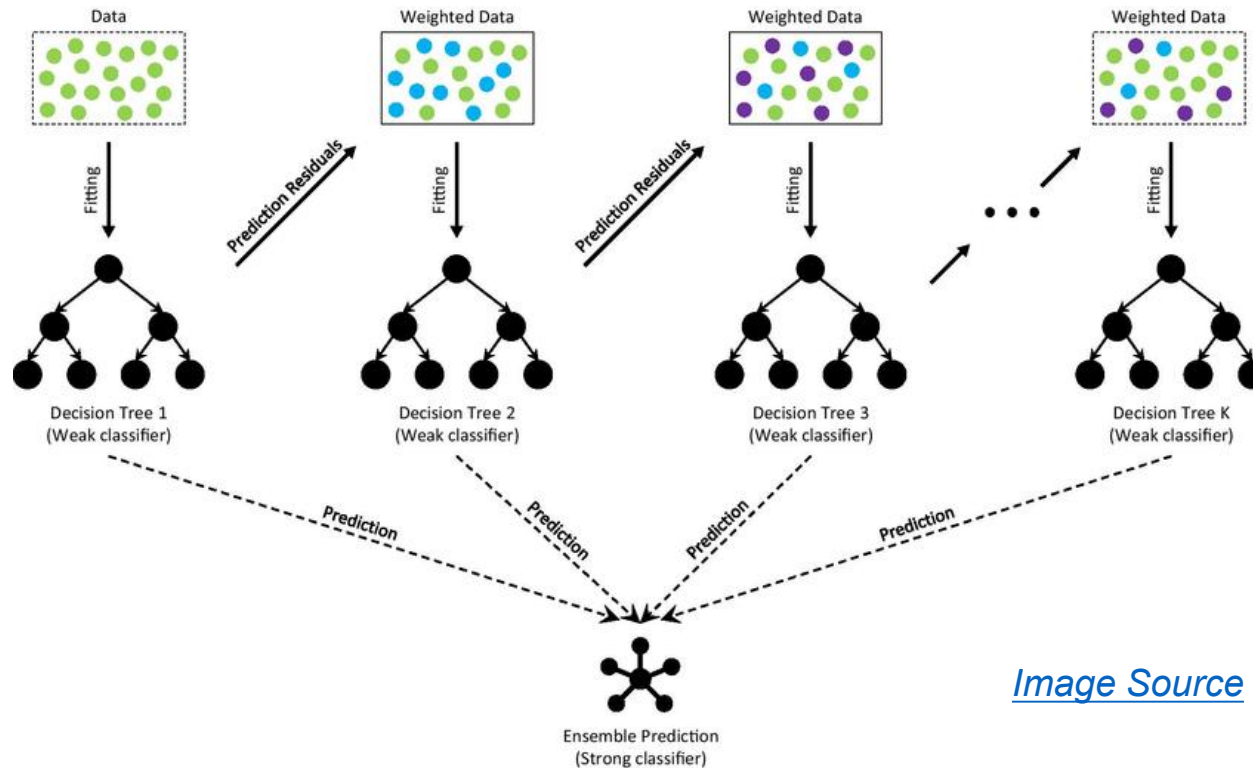
# CC0 $\pi$ pre-selected with secondary clusters (minimum distance from all tracks)



- Background distribution peaks for clusters close to other tracks ( $\sim 5$  cm) and leads to impure neutron sample.
- These are primarily due to muon hits not used in track reconstruction and delta electrons produced by primary muons.

# Boosted Decision Trees (BDTs)

- Training: Trees are trained sequentially, and each new tree focuses on correcting the errors made by the previous one.



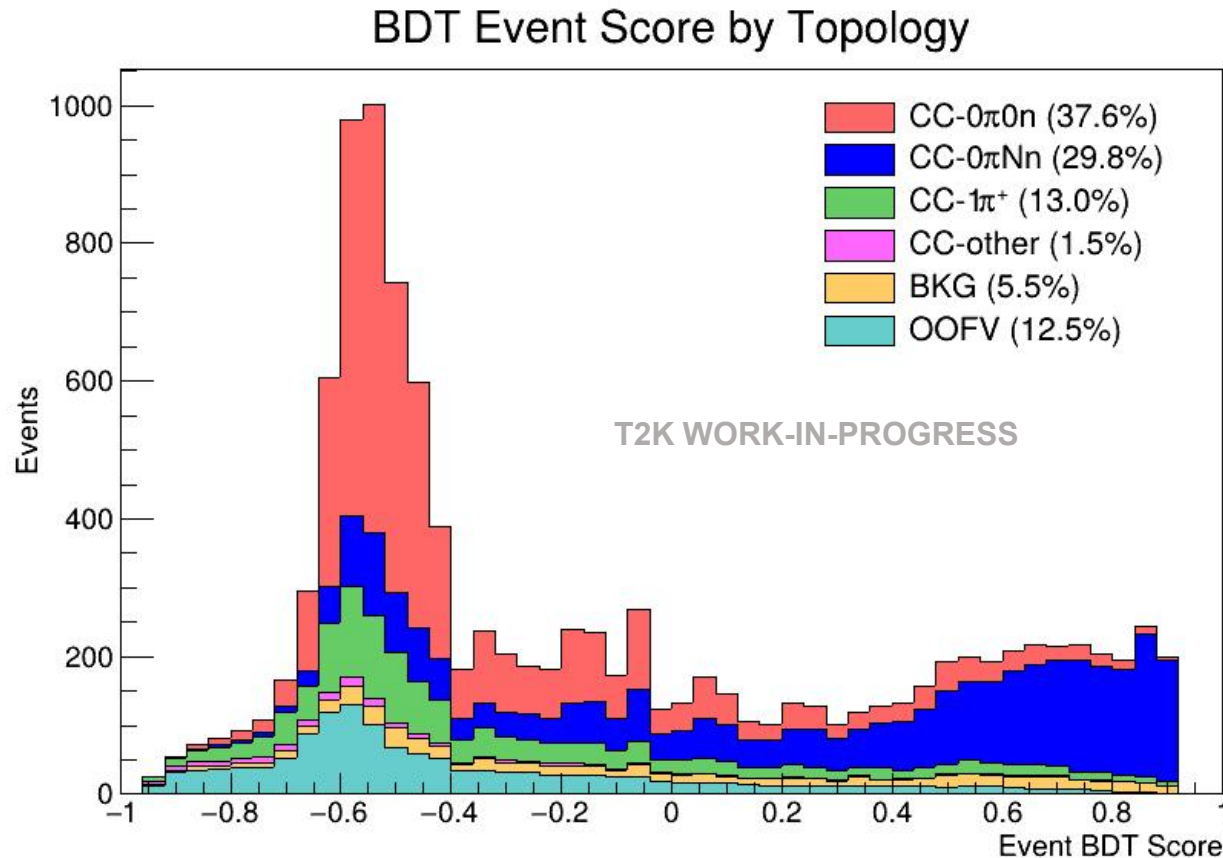
Gradient BDT

[Image Source](#)

# BDT to Select $CC0\pi Nn$ events

- BDT1: Cluster training (Signal: clusters from neutrons)
- BDT2: Track training (Signal: tracks from neutrons)
- BDT3: Event level to find  $CC0\pi Nn$  events (takes inputs from BDT1 and BDT2)

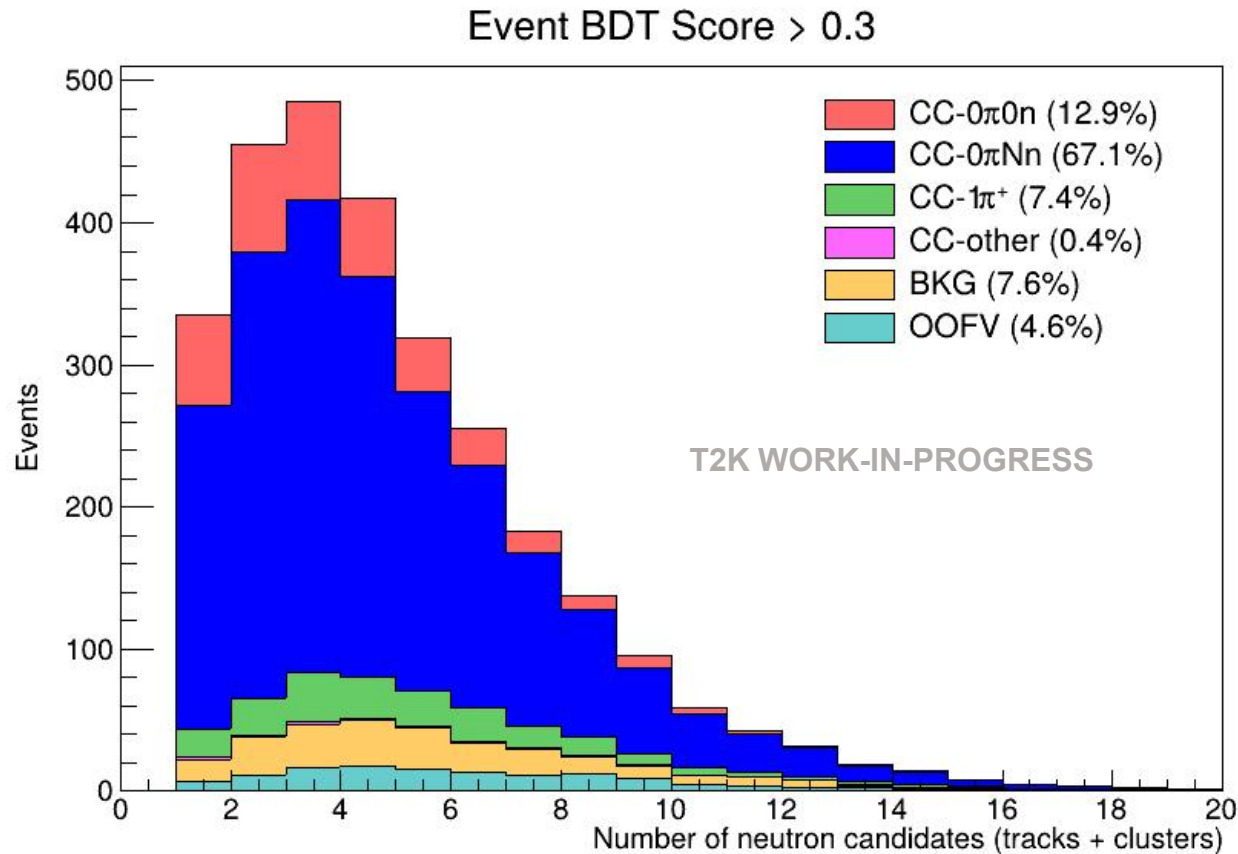
# Event BDT Score



Initial purity of CC0 $\pi$ Nn:  $\sim 30\%$

With high BDT scores,  $\sim 90\%$  purity is achievable.

# Number of neutron candidates



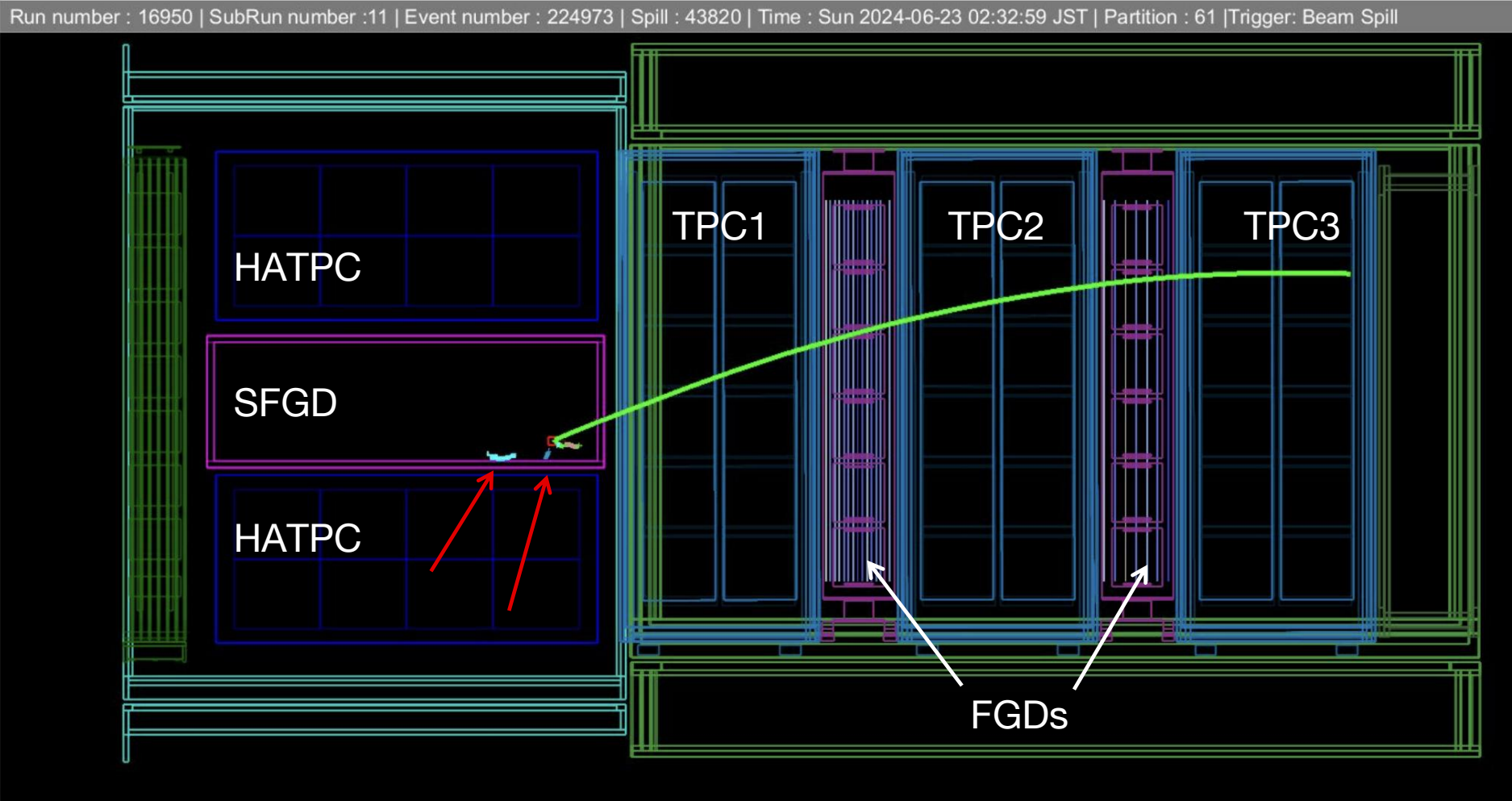
Distribution of neutron candidates (tracks + clusters) peaks around 2 - 4 neutron candidates.

# Summary

- ND280 upgrade has made **reconstruction of particles with shorter tracks** or smaller energies even better.
- CC0 $\pi$ Nn topology can be instrumental in **testing FSI models**.
- Neutrons leave an energy deposit in the form of a track or a cluster **away from the vertex**.
- To identify neutron events, BDTs are trained and they give **high signal purity**.
- Currently, work is in progress to **develop systematics** and **study model dependence** in CC0 $\pi$ Nn selection.

Thanks for your attention!

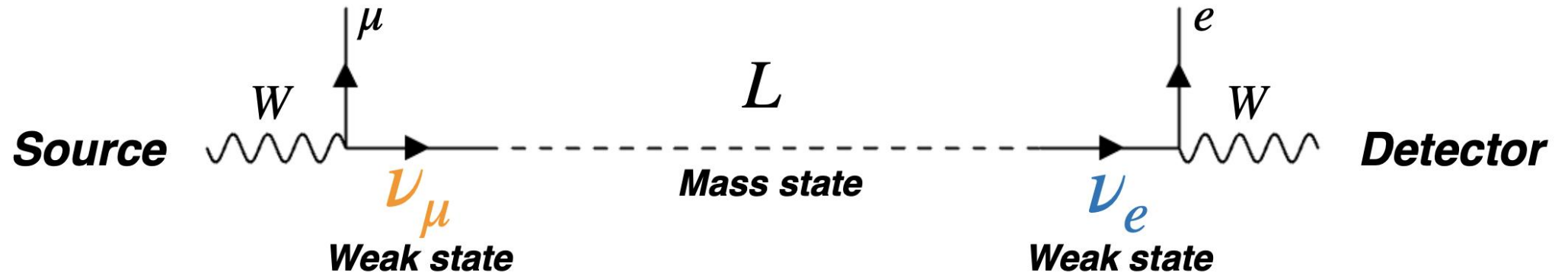
# Neutron candidate in an actual event



Two neutron candidates in this event

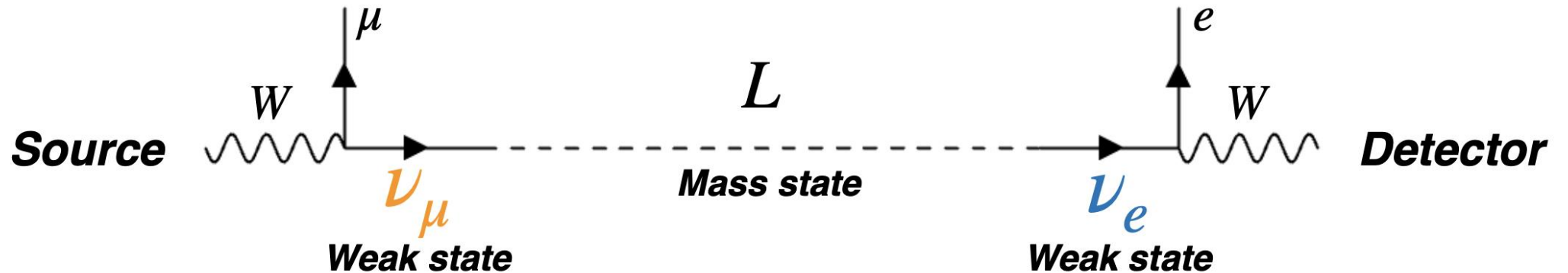
**Backup**

# Neutrino Oscillations



- Neutrinos produced in a specific flavor state ( $\nu_\mu$ ) with energy  $E$
- After traveling a distance  $L$ ,  
 $\nu_\mu$  decrease in number  $\Rightarrow \nu_\mu$  disappearance  
 $\nu_e$  increase in number  $\Rightarrow \nu_e$  appearance
- Sensitivity to maximum oscillation depends on  $L/E$

# Neutrino Oscillations



- Neutrinos produced in a specific flavor state ( $\nu_\mu$ ) with energy  $E$
- Oscillation probability depends on:

- Neutrino energy
- Travelling distance (“baseline”)
- PMNS mixing parameters
- Difference between mass states

Need to know

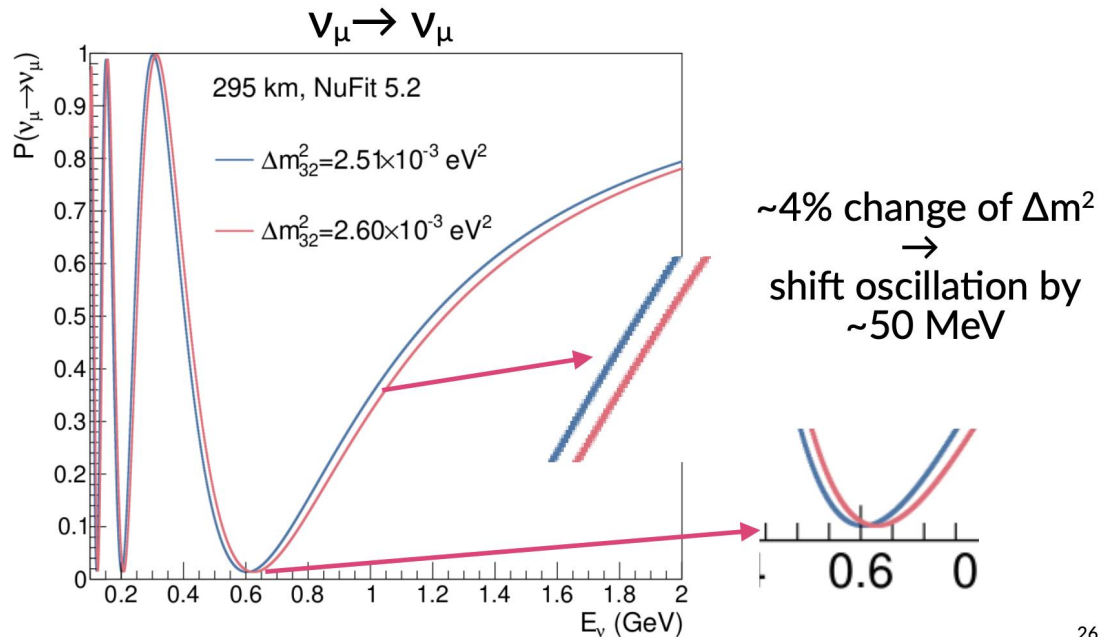
Known

What is measured

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \text{PMNS} \\ \text{matrix} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

Parameters in PMNS matrix  
( $\theta_{23}, \theta_{13}, \theta_{12}, \delta_{CP}$ )

# Neutrino Energy as Uncertainty



[C. Wret, NuSTEC2024](#)

- For precise measurement of oscillation parameters, neutrino energy must be well-reconstructed.
- One of the leading sources of uncertainty in neutrino energy reconstruction are neutrons (calorimetric) and interaction models (kinematic).
- They have a high chance of escaping the detector without depositing any energy, hence, creating a problem in reconstructing neutrino energy.

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# PONTECORVO- MAKI-NAKAGAWA-SAKATA (PMNS) MATRIX

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} \text{PMNS} \\ \text{matrix} \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta_{CP}} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta_{CP}} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} P$$

$$= \begin{pmatrix} c_{12}c_{13} & s_{12}c_{13} & s_{13}e^{-i\delta_{CP}} \\ -s_{12}c_{23} - c_{12}s_{13}s_{23}e^{i\delta_{CP}} & c_{12}c_{23} - s_{12}s_{13}s_{23}e^{i\delta_{CP}} & c_{13}s_{23} \\ s_{12}s_{23} - c_{12}s_{13}c_{23}e^{i\delta_{CP}} & -c_{12}s_{23} - s_{12}s_{13}c_{23}e^{i\delta_{CP}} & c_{13}c_{23} \end{pmatrix} P.$$

- PMNS matrix describes mixing between flavour states and mass states

- Six parameters in PMNS matrix  
 $(\theta_{23}, \theta_{13}, \delta_{CP}, \theta_{12}, \alpha_1, \alpha_2)$



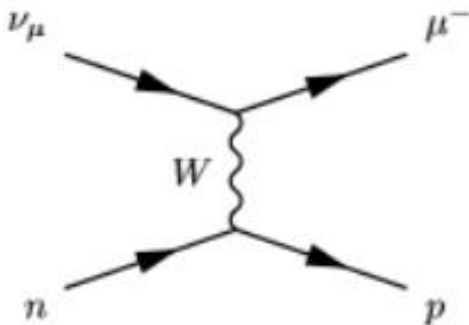
T2K is sensitive to these three parameters

$$P_{\text{Majorana}} = \begin{pmatrix} e^{i\alpha_1} & 0 & 0 \\ 0 & e^{i\alpha_2} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

- CP phase (or “Dirac phase”)  $\delta_{CP}$ : a measure of asymmetry between neutrino and antineutrino oscillations
- Majorana phase ( $\alpha_1, \alpha_2$ ): Plays a role in neutrinoless double beta decay but not in neutrino oscillations

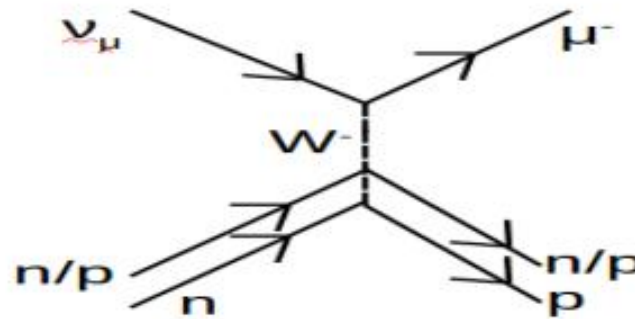
# NEUTRINO-NUCLEUS INTERACTIONS

Quasi-Elastic  
Scattering (CCQE)



$$\nu_\mu n \rightarrow \mu^- p$$

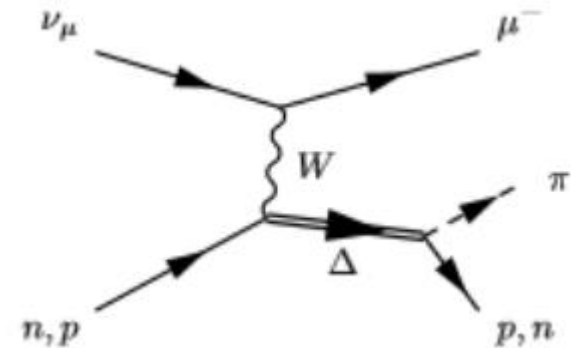
Two-particle two-hole  
process (2p2h)



$$\nu_\mu nn \rightarrow \mu^- np$$

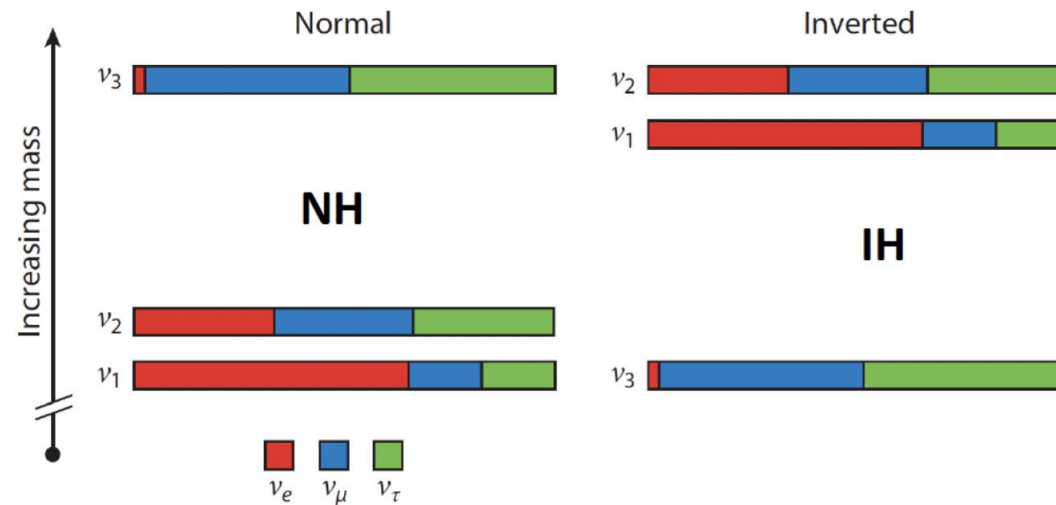
$$\nu_\mu np \rightarrow \mu^- pp$$

Delta Resonance



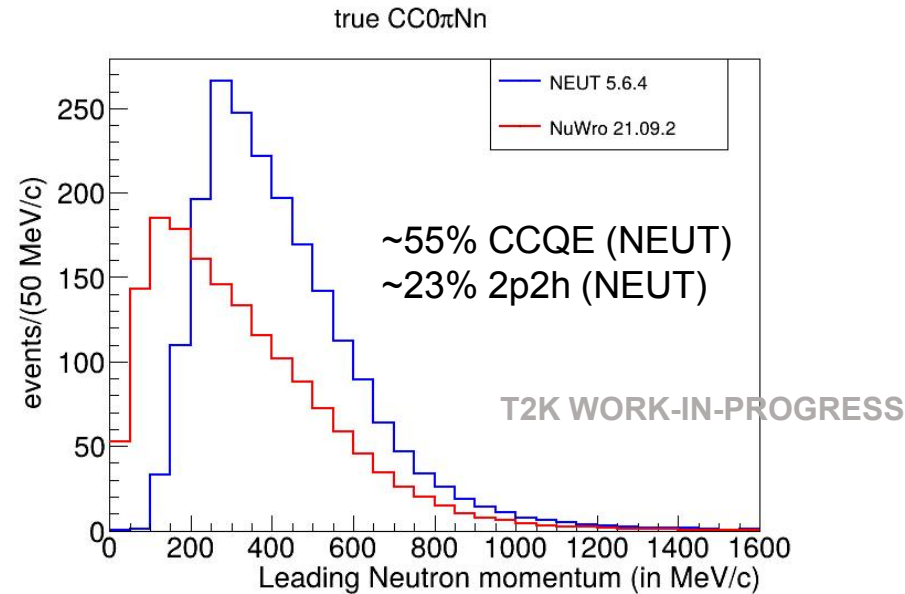
$$\begin{aligned} \nu_\mu n &\rightarrow \mu^- \Delta^+ \rightarrow \mu^- \pi^+ n \\ &\rightarrow \mu^- \pi^0 p \\ \nu_\mu p &\rightarrow \mu^- \Delta^{++} \rightarrow \mu^- \pi^+ p \end{aligned}$$

# Neutrino Mass Ordering



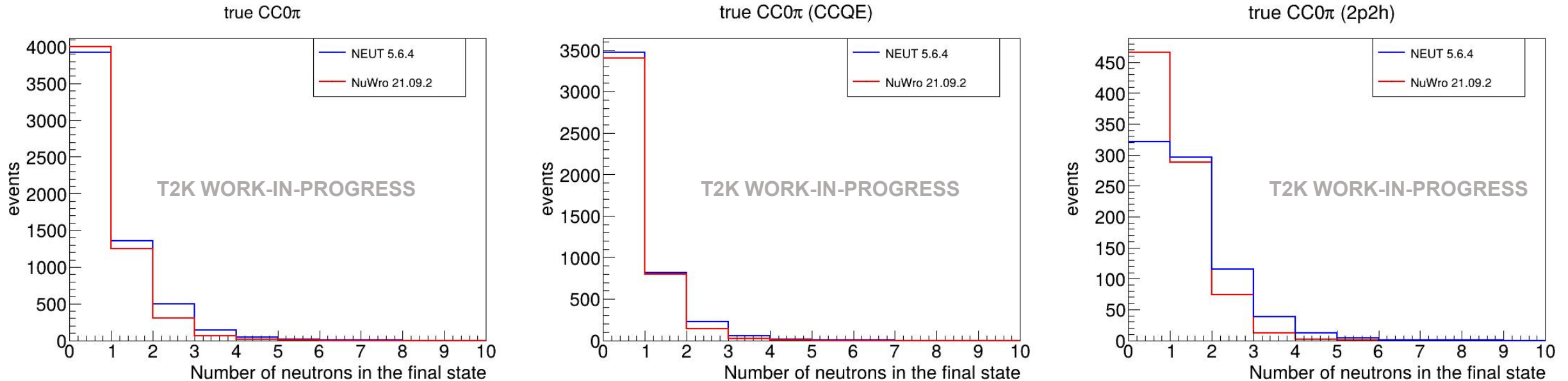
<https://arxiv.org/pdf/1610.05533.pdf>

# Comparison between NEUT and NuWro



- All the CCQE events enter this topology through FSI only.
- This topology can differentiate between FSI models in NEUT and NuWro using T2K data.

# Neutron Multiplicity in NEUT and NuWro



- NEUT 5.6.4 and NuWro 21.09.2 use the same 2p2h model but the fraction pp/np is different between the two which has an impact on neutron multiplicity.

# Event Rate Calculation

Number of  $\nu_e$  events at Super-Kamiokande

$$N(E_{reco}) = \int dE_{true} P(\nu_\mu \rightarrow \nu_e)(E_{true}) \sigma(E_{true}) \Phi(E_{true}) S(E_{true}, E_{reco}) \epsilon(E_{true})$$

$N(E_\nu)$  = Event rate

$P(\nu_\mu \rightarrow \nu_e)(E_\nu)$  = Oscillation probability

$\sigma(E_\nu)$  = Interaction cross-section

$\Phi(E_\nu)$  = Neutrino flux

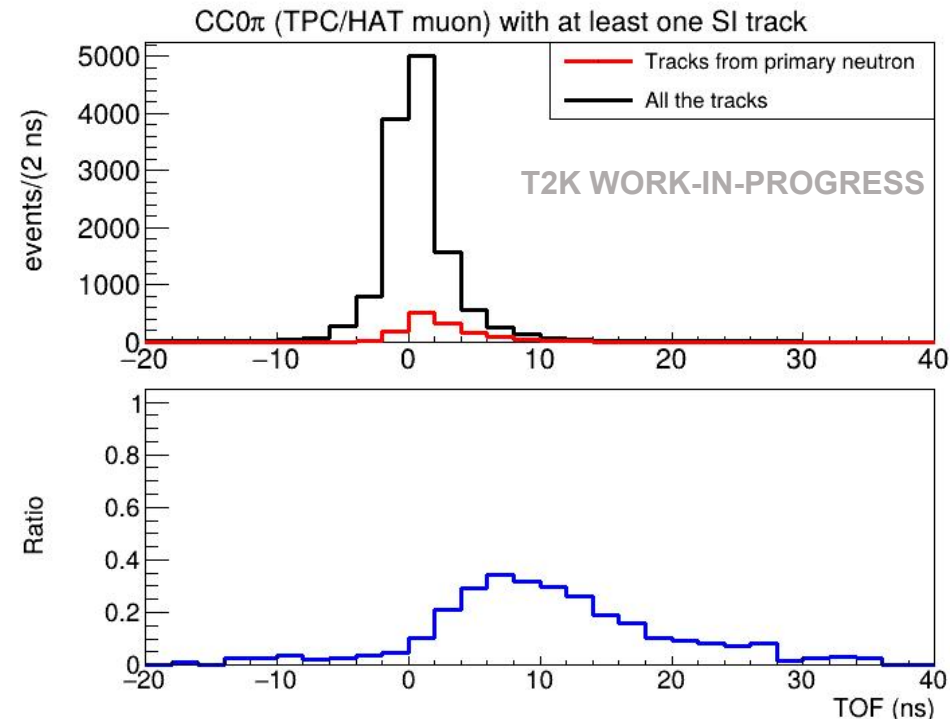
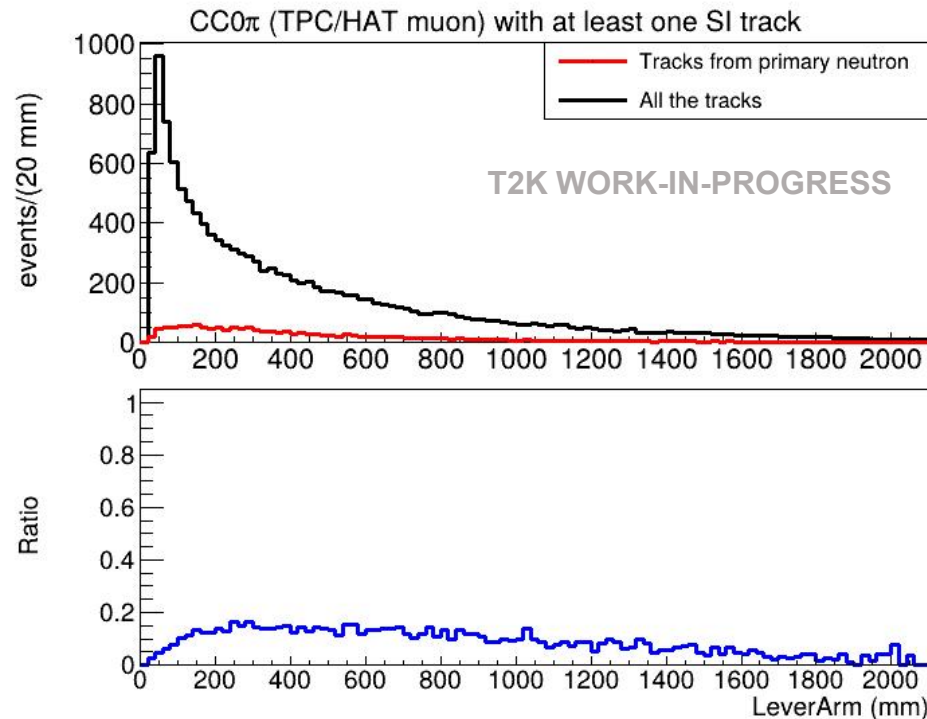
$S(E_{true}, E_{reco})$  = Smearing matrix

$\epsilon(E_\nu)$  = Detector efficiency

Reconstructed energy can be different from true energy

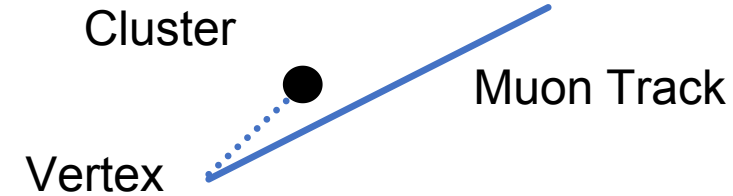
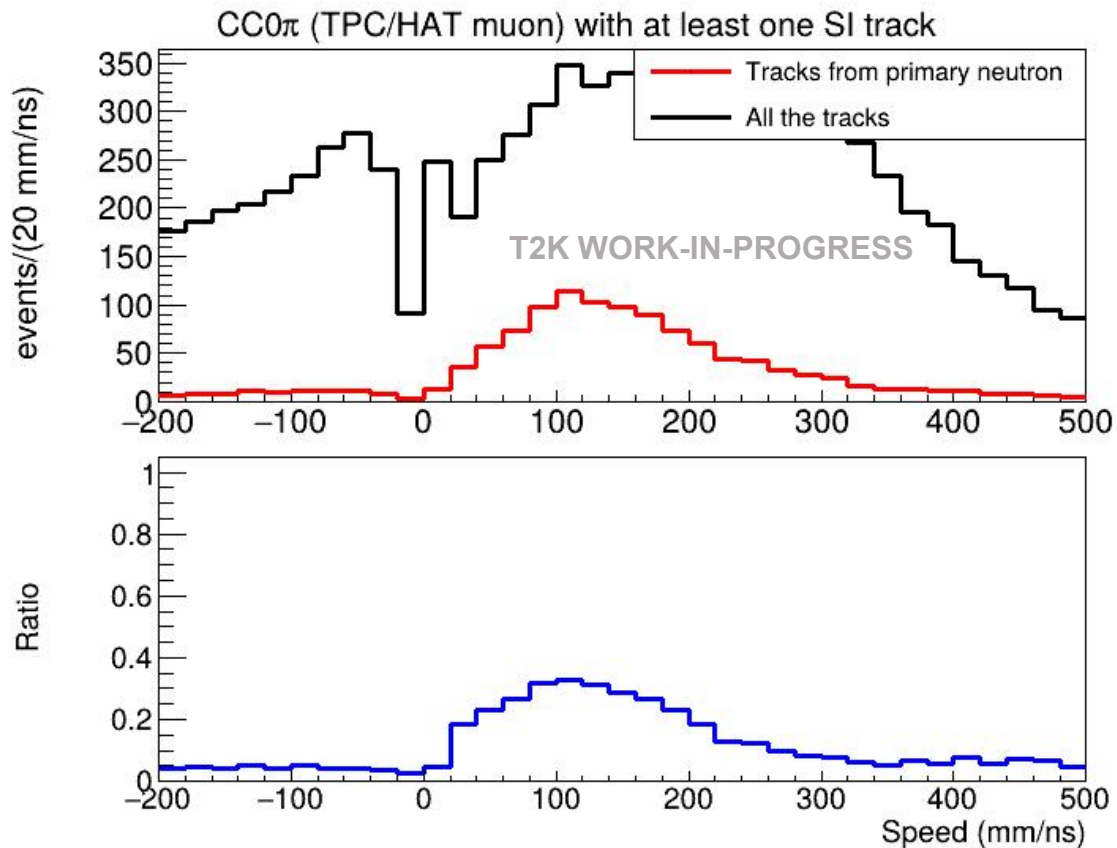
Neutrino-nucleus interaction models

# CC0 $\pi$ pre-selected with secondary tracks



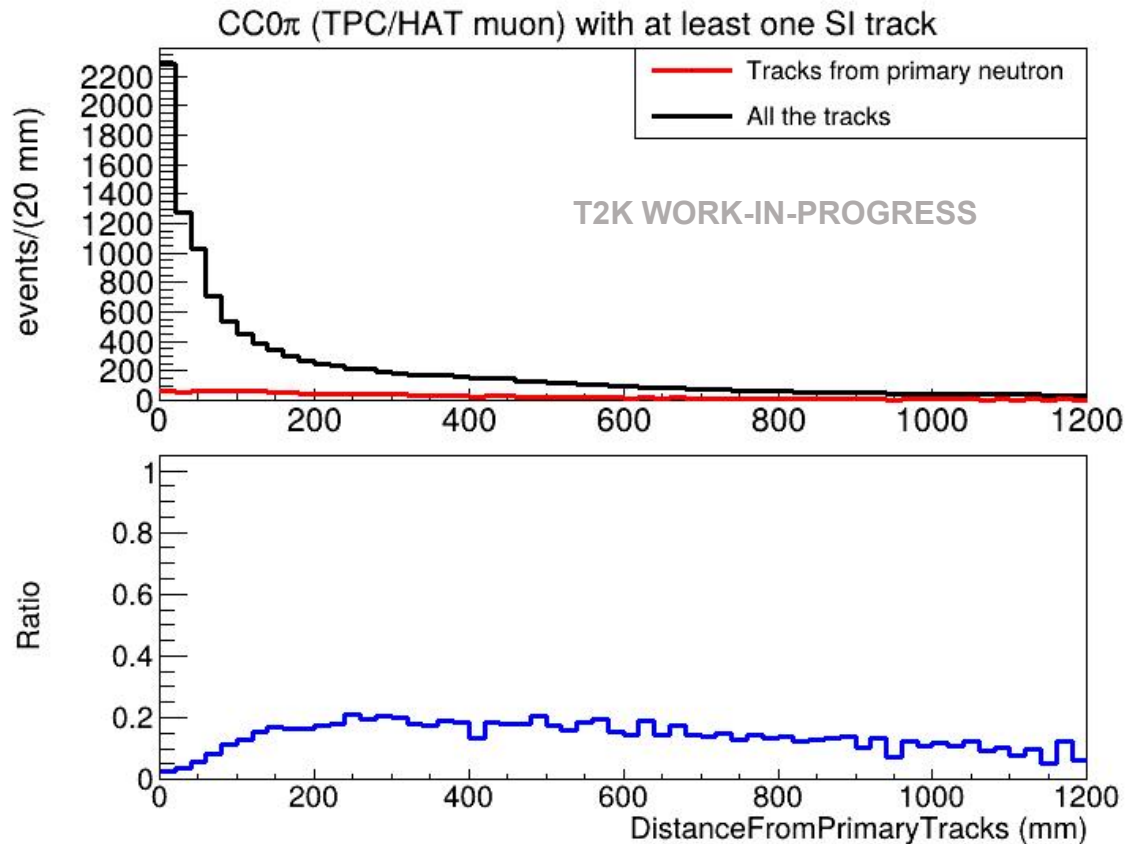
- Neutron (and background) cluster distribution peaks around 50-60 cm from the neutrino vertex.
- Time-of-flight can be negative due to timing resolution of SFGD. Applying a cut on the distribution, such as  $0 < \text{TOF (ns)} < 10$ , can provide a pure sample of neutrons.

# CC0 $\pi$ pre-selected with secondary tracks (speed)



- Speed = Lever Arm / Time-of-Flight
- Reconstructed speed would not be more than 300 mm/ns (speed of light) if the detectors were perfect.
- Speed drops around the speed of light and peaks around 40% of the speed of light.
- This is higher than cluster distribution since tracks are generally produced by high momentum neutrons.

# CC0 $\pi$ pre-selected with secondary tracks (minimum distance from all tracks)



- Background distribution peaks for clusters close to other tracks ( $\sim 5$  cm) and leads to impure neutron sample.
- These are primarily due to muon hits not used in track reconstruction and delta electrons produced by primary muons.

# Decision Tree

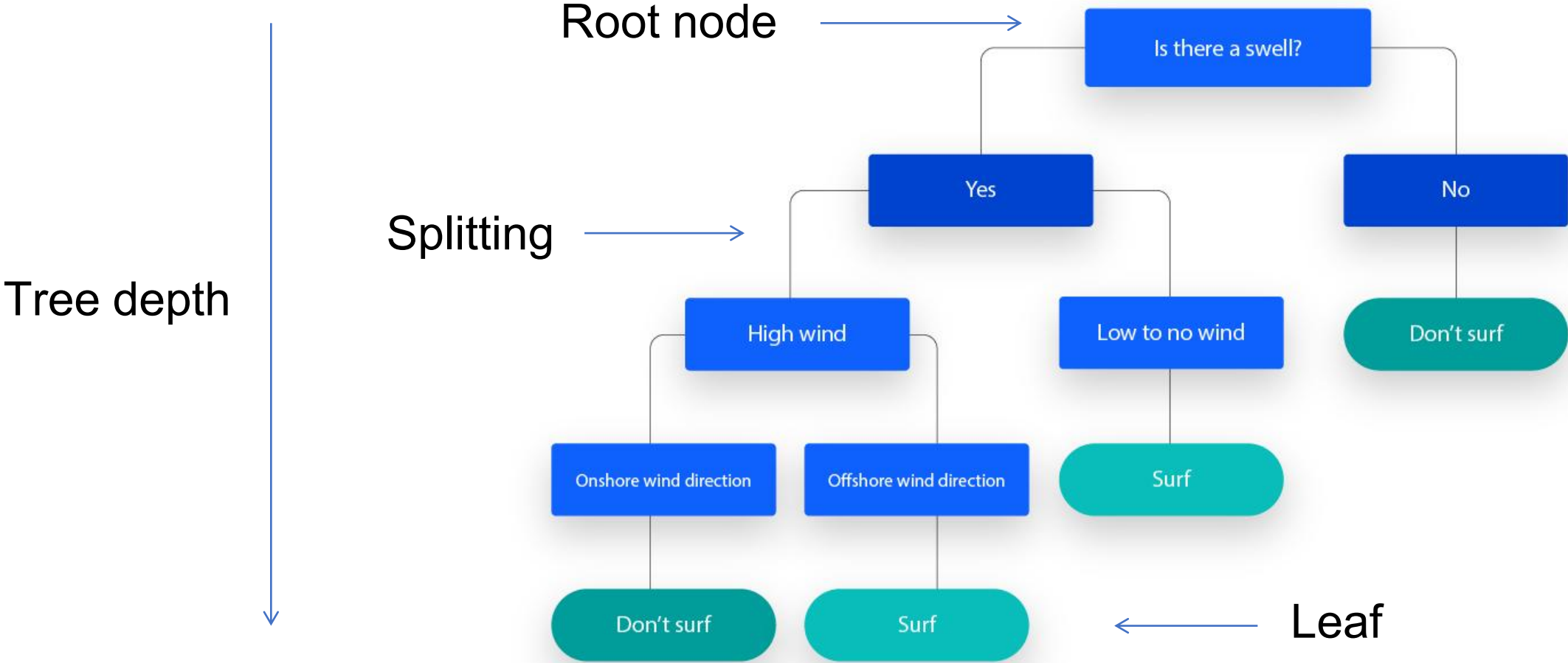
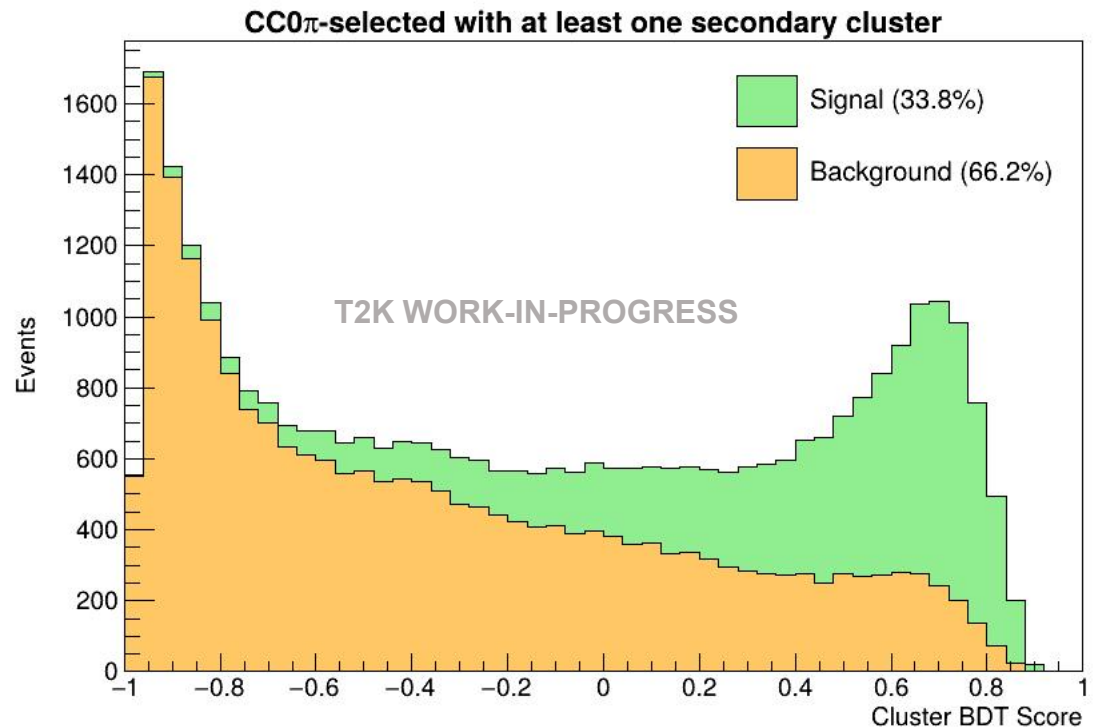
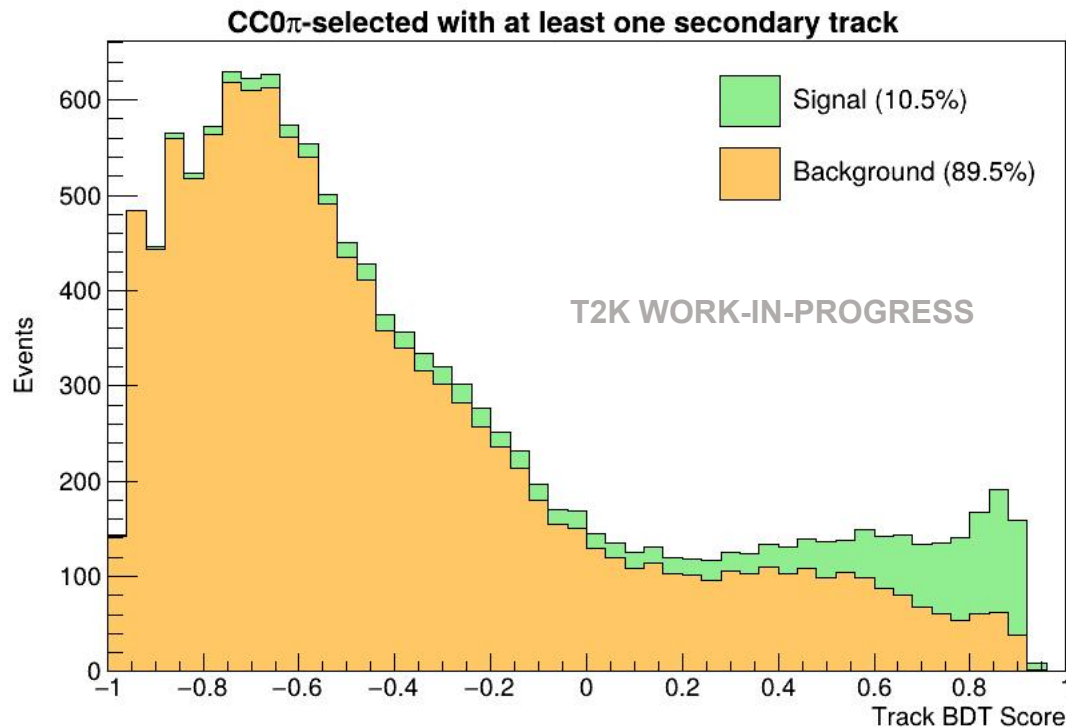


Image Source

# Problems with Single Decision Trees

- Single decisions tree are fast and easy but have problems:
  1. Can lead to overfitting, especially if deep, which means difficult to generalize to new events.
  2. Might not describe the data very well if shallow.
  3. Instability: Small changes in data can change the whole tree.

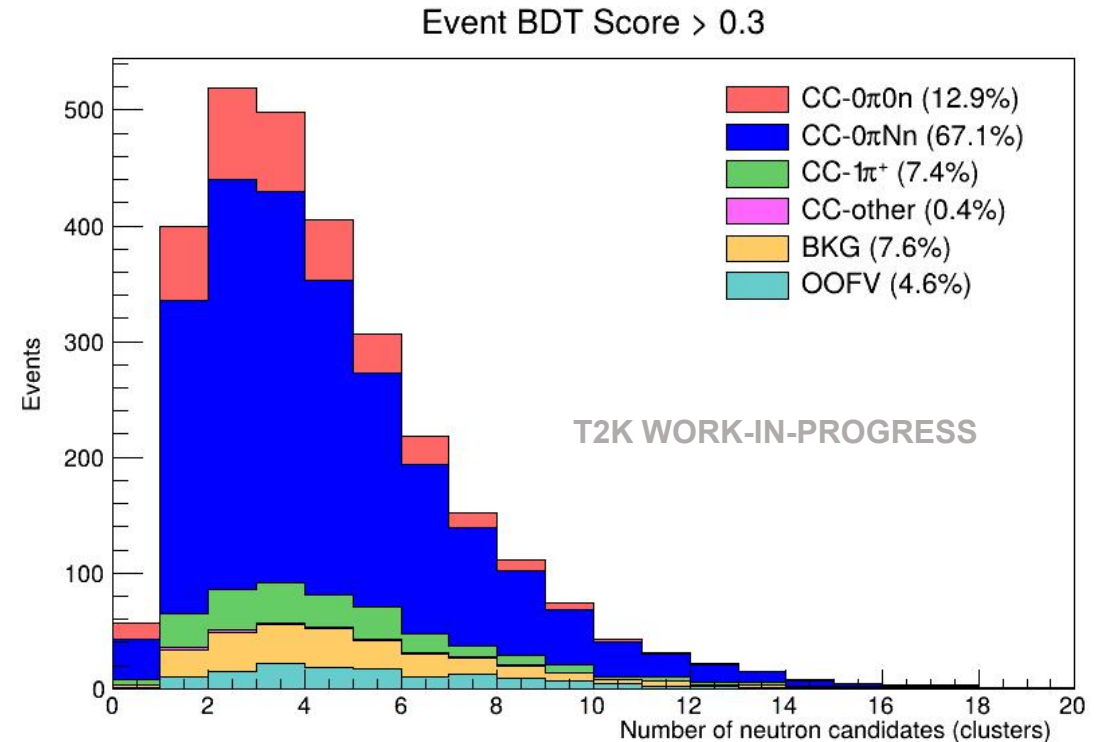
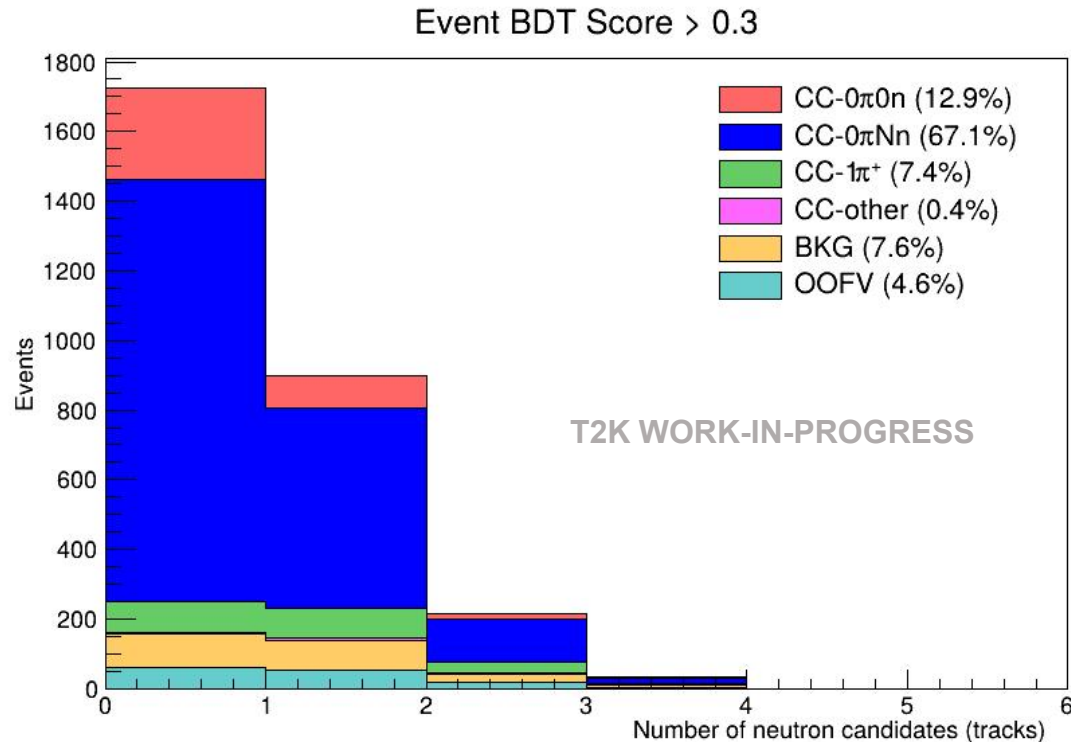
# Track and Cluster BDT Score



Track BDT score is for all the secondary tracks (at least 3 cm from the reconstructed primary vertex) in all CC0 $\pi$ -selected events. Same for clusters.

Signal means that the secondary object is connected to a primary neutron through one or more scatterings and that neutron has the same vertex as that of the muon.

# Number of neutron (track and cluster) candidates



There is at least one neutron candidate in these events. So, no cluster means that there is at least one track and vice versa.