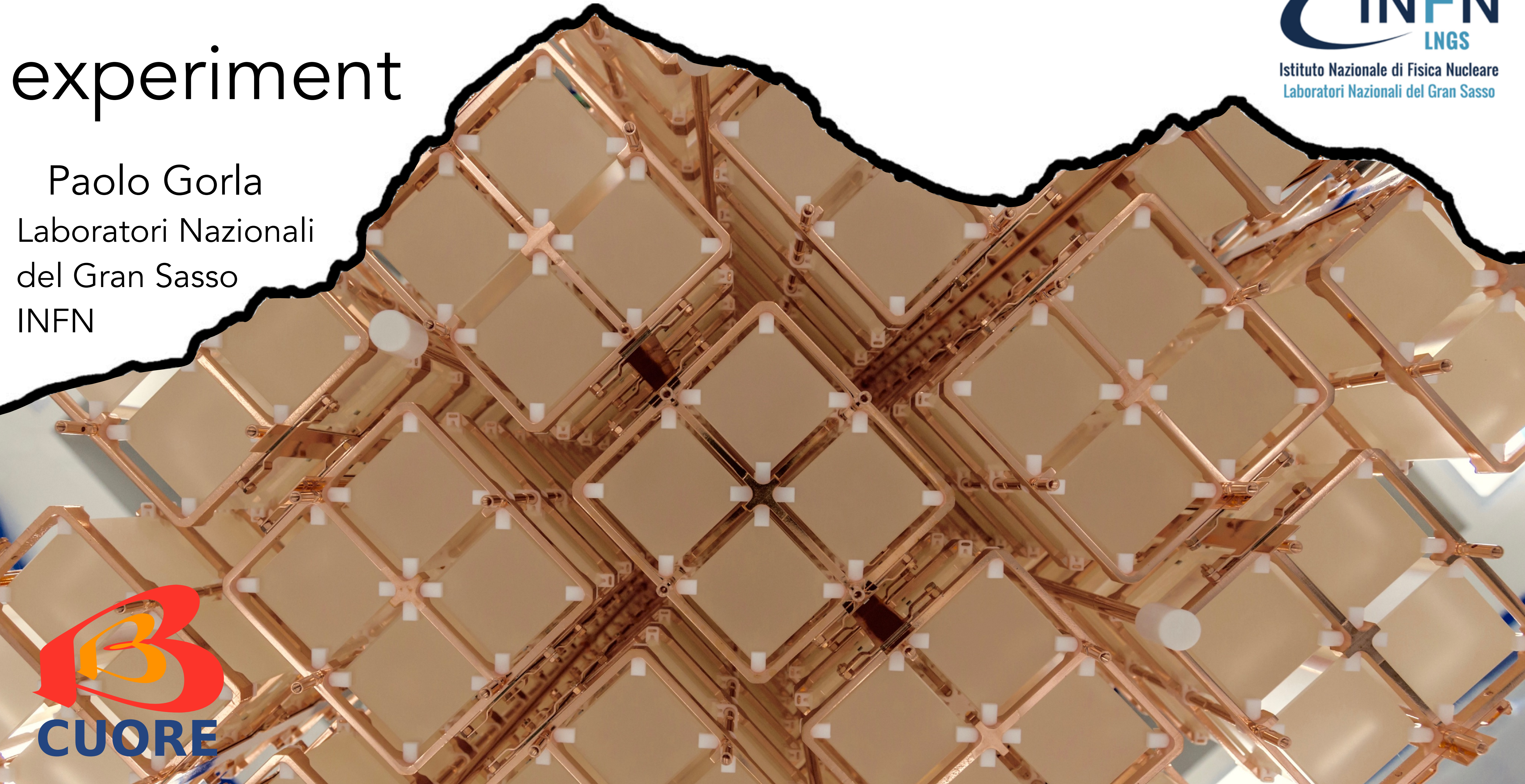


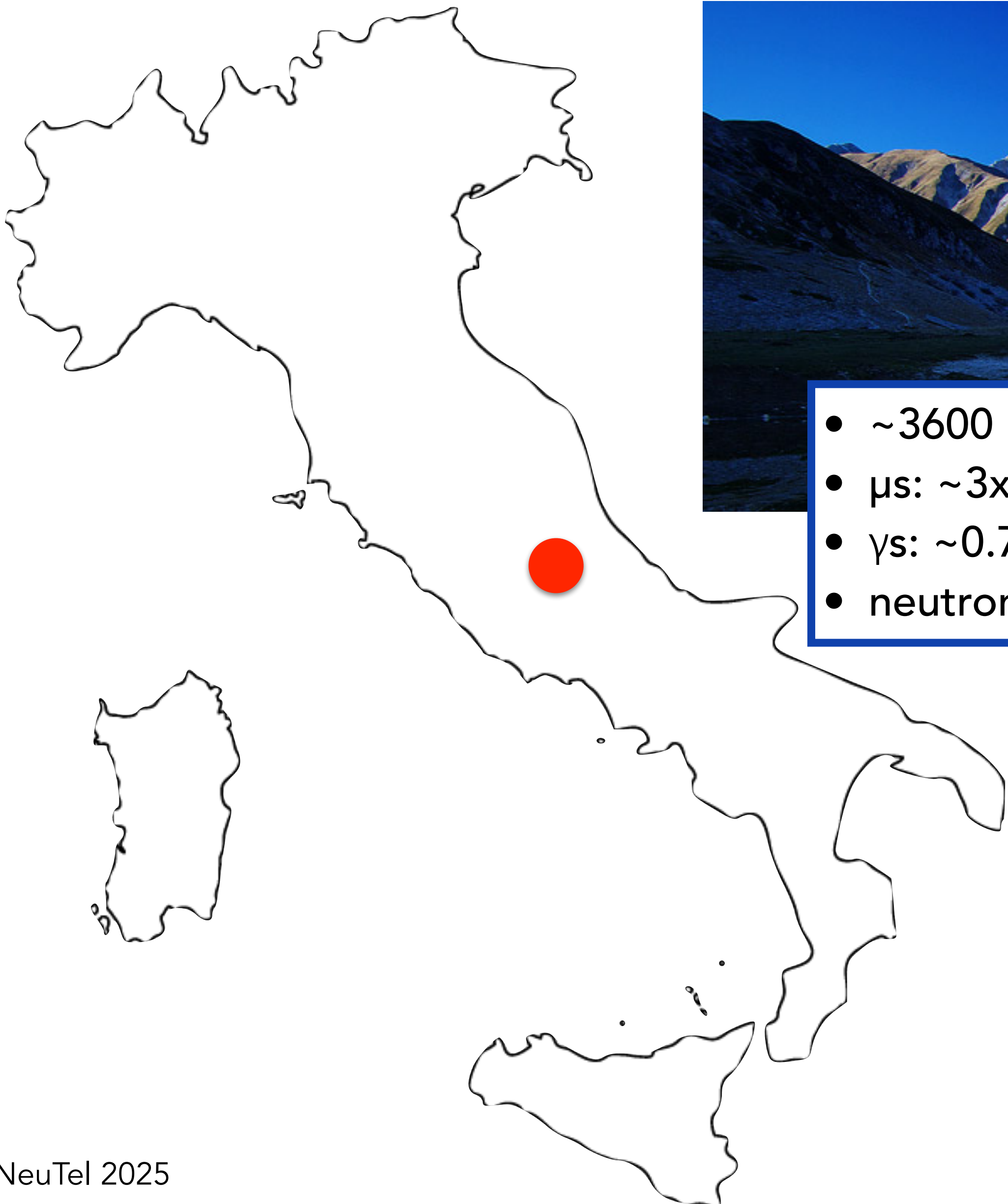
Latest results from the CUORE experiment

Paolo Gorla
Laboratori Nazionali
del Gran Sasso
INFN

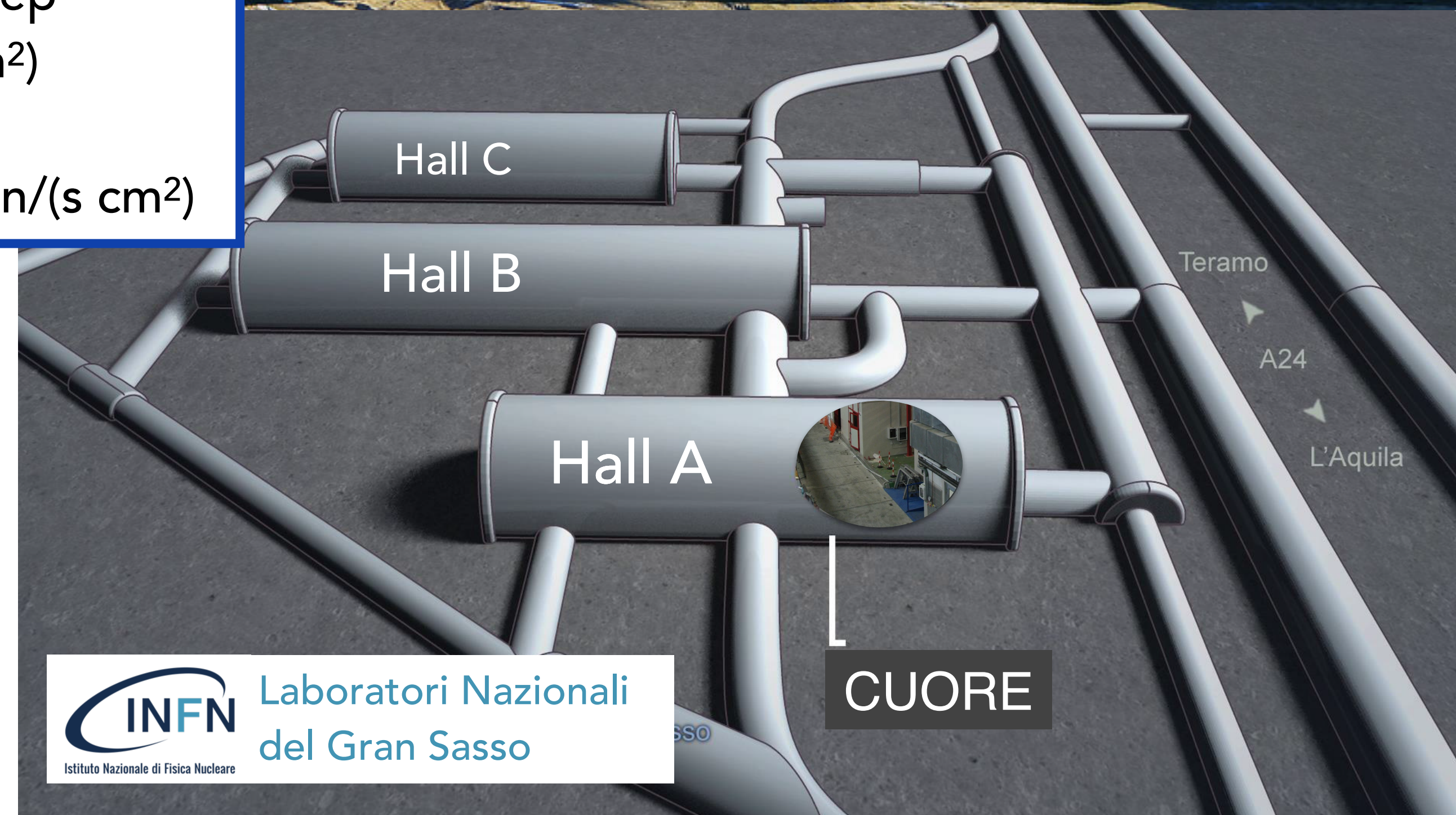


The CUORE collaboration





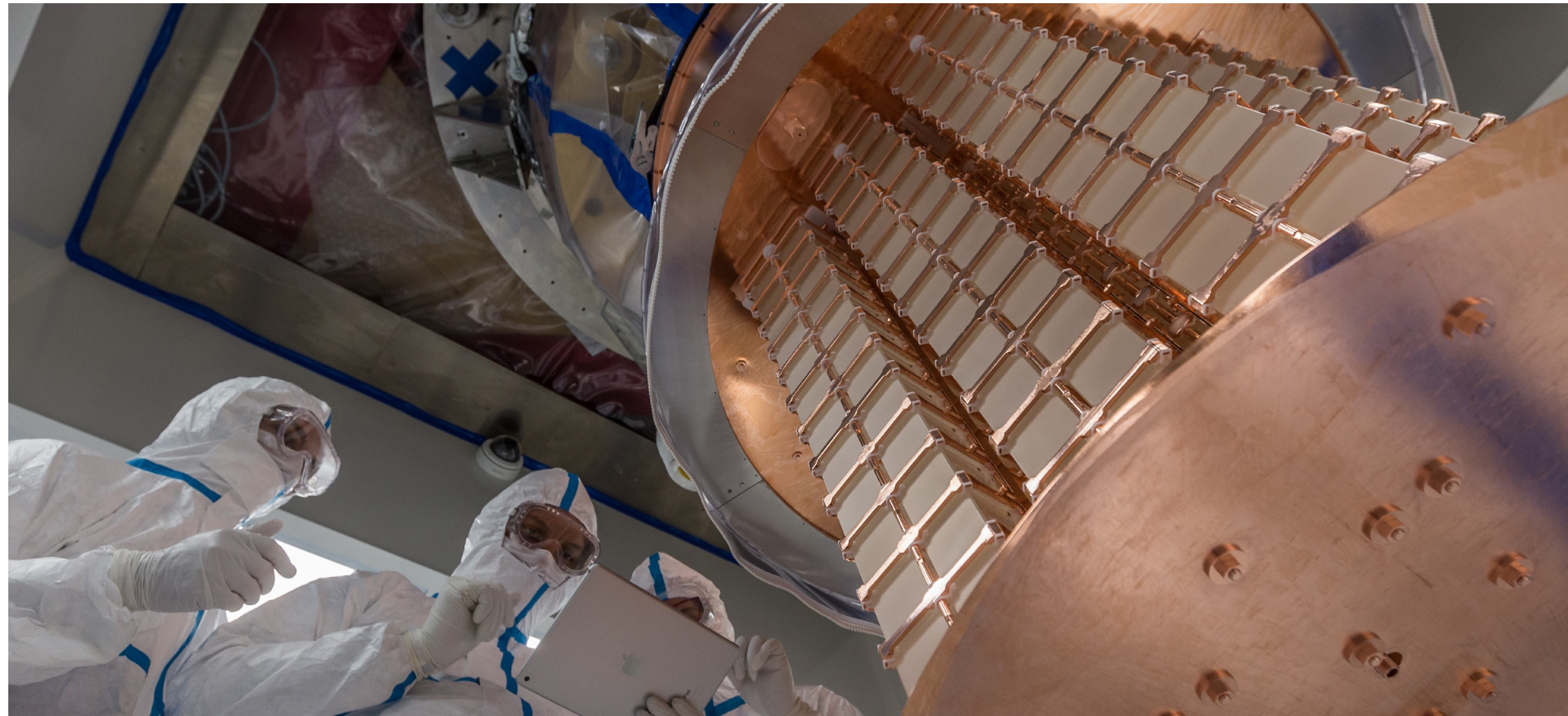
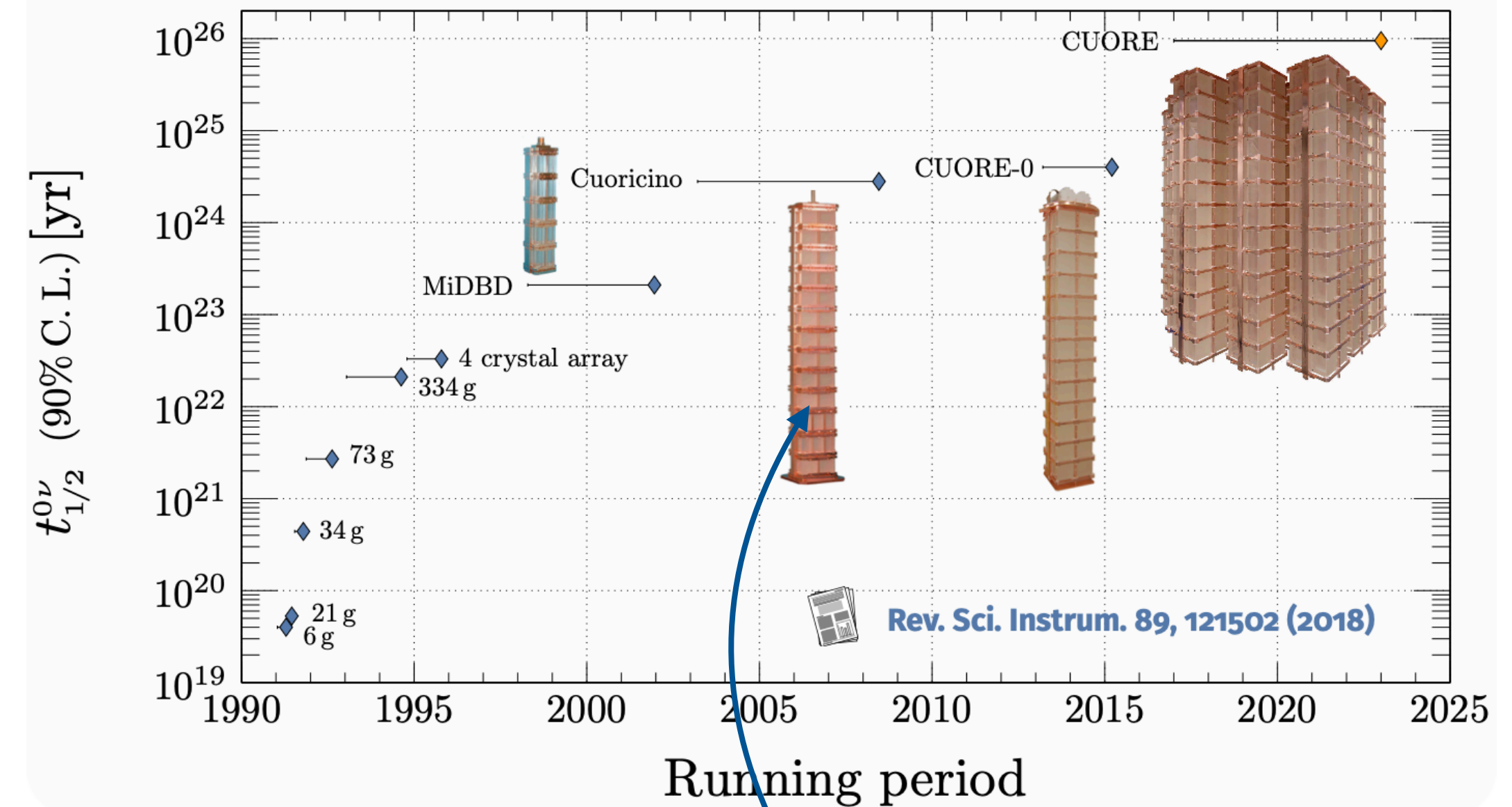
- ~ 3600 m.w.e. deep
- μ s: $\sim 3 \times 10^{-8}/(\text{s cm}^2)$
- γ s: $\sim 0.73/(\text{s cm}^2)$
- neutrons: $4 \times 10^{-6} \text{ n}/(\text{s cm}^2)$



The CUORE experiment in a nutshell

Cryogenic **U**nderground **O**bservatory for **R**are **E**vents

- Primary goal: search for $0\nu\beta\beta$ decay of ^{130}Te
 Eur. Phys. J. C77 (2017), 532
- **Largest cryogenic-calorimeter detector ever built**
19 towers of 13 floors of 4 crystals = **988 detectors**
- 1 t detector mass: **327 kg Cu + 742 kg TeO_2**
 $\rightarrow 206$ kg of ^{130}Te

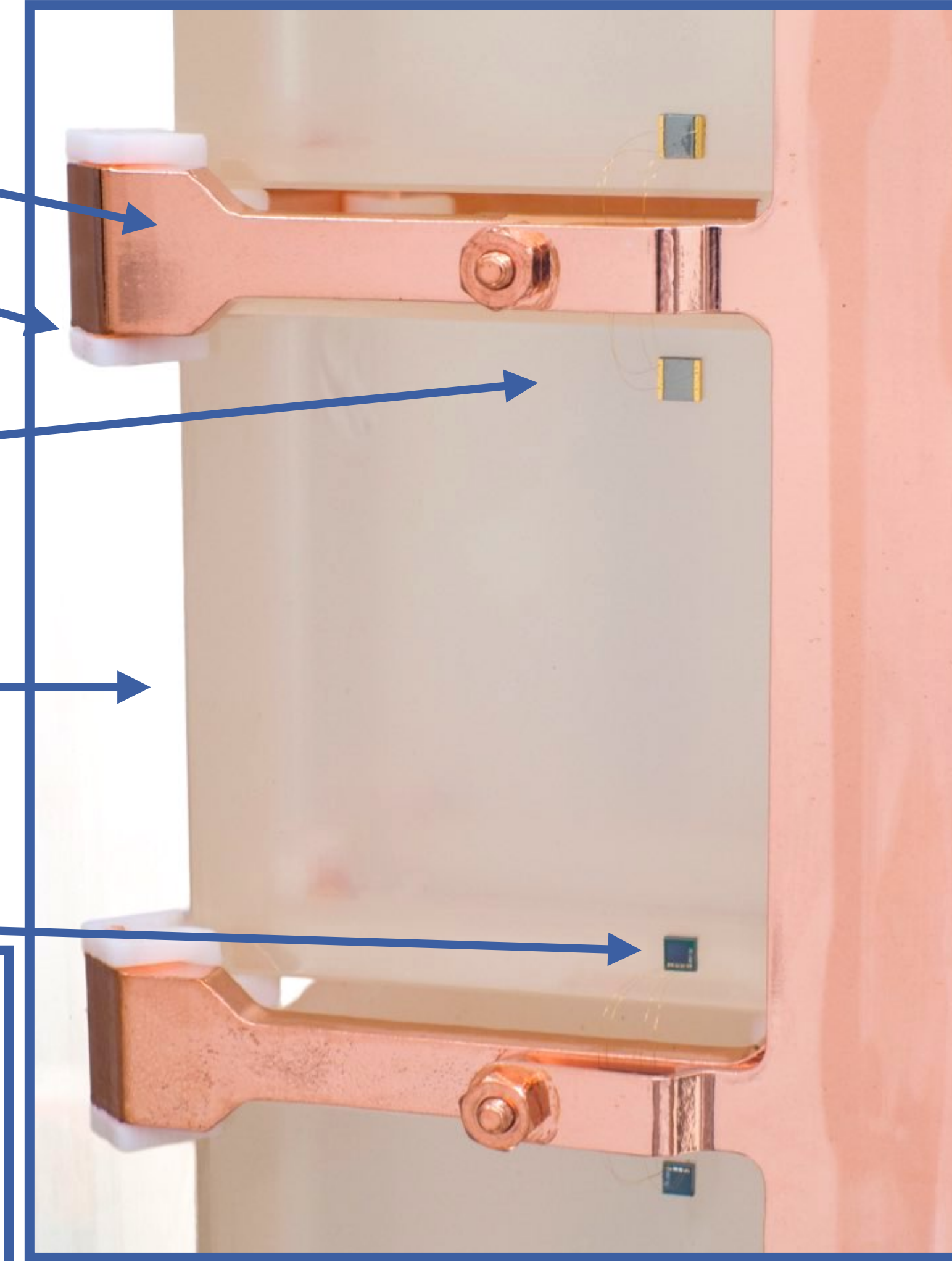
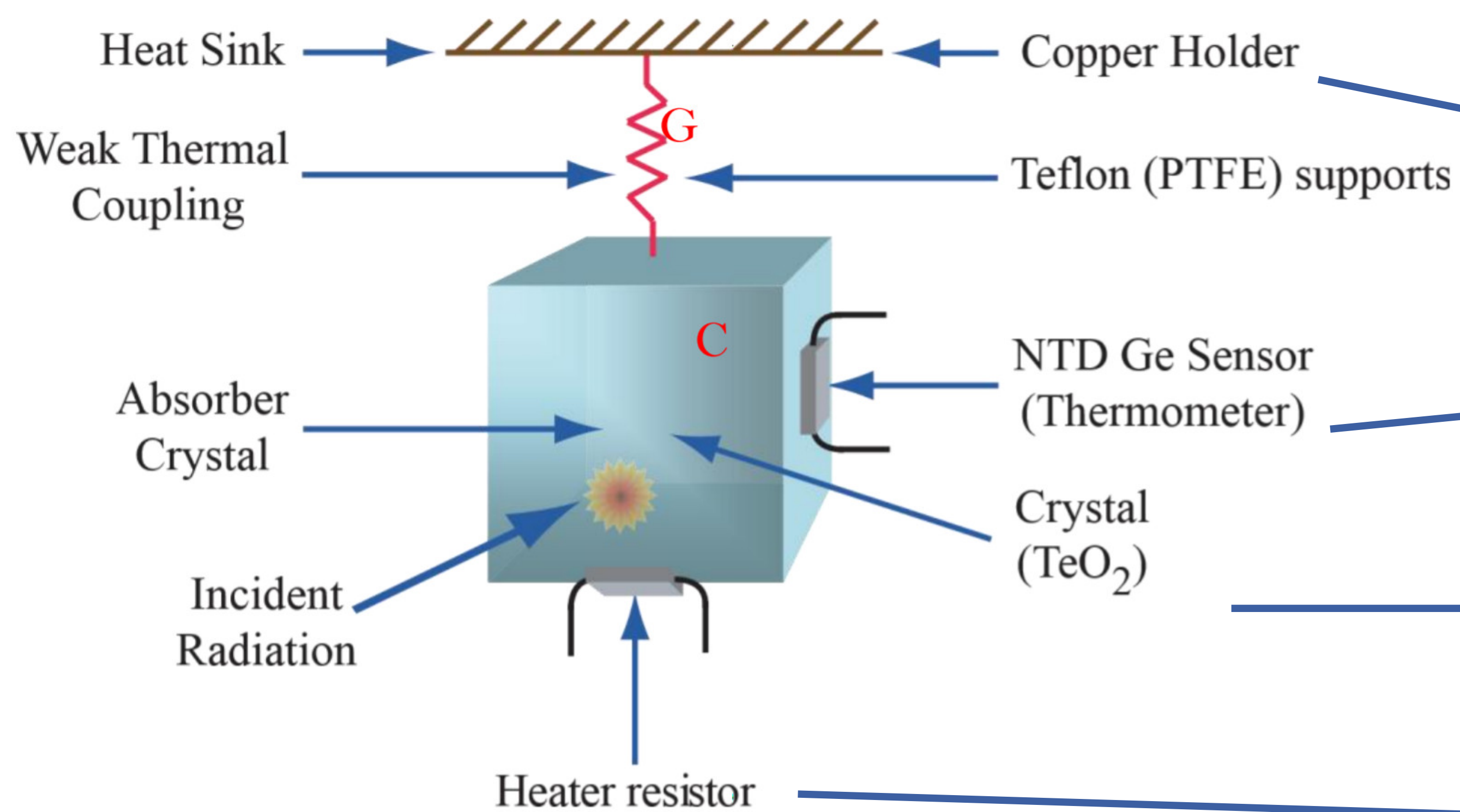


- Building on a 30-yr-long history of measurements
- Multi-year-long cryogenic stability at **T= 11–15 mK**
- FWHM resolution at $Q_{\beta\beta}$ of **7.3 keV**
- bkg in ROI **$1.4 \cdot 10^{-2}$ counts keV⁻¹ kg⁻¹ yr⁻¹**



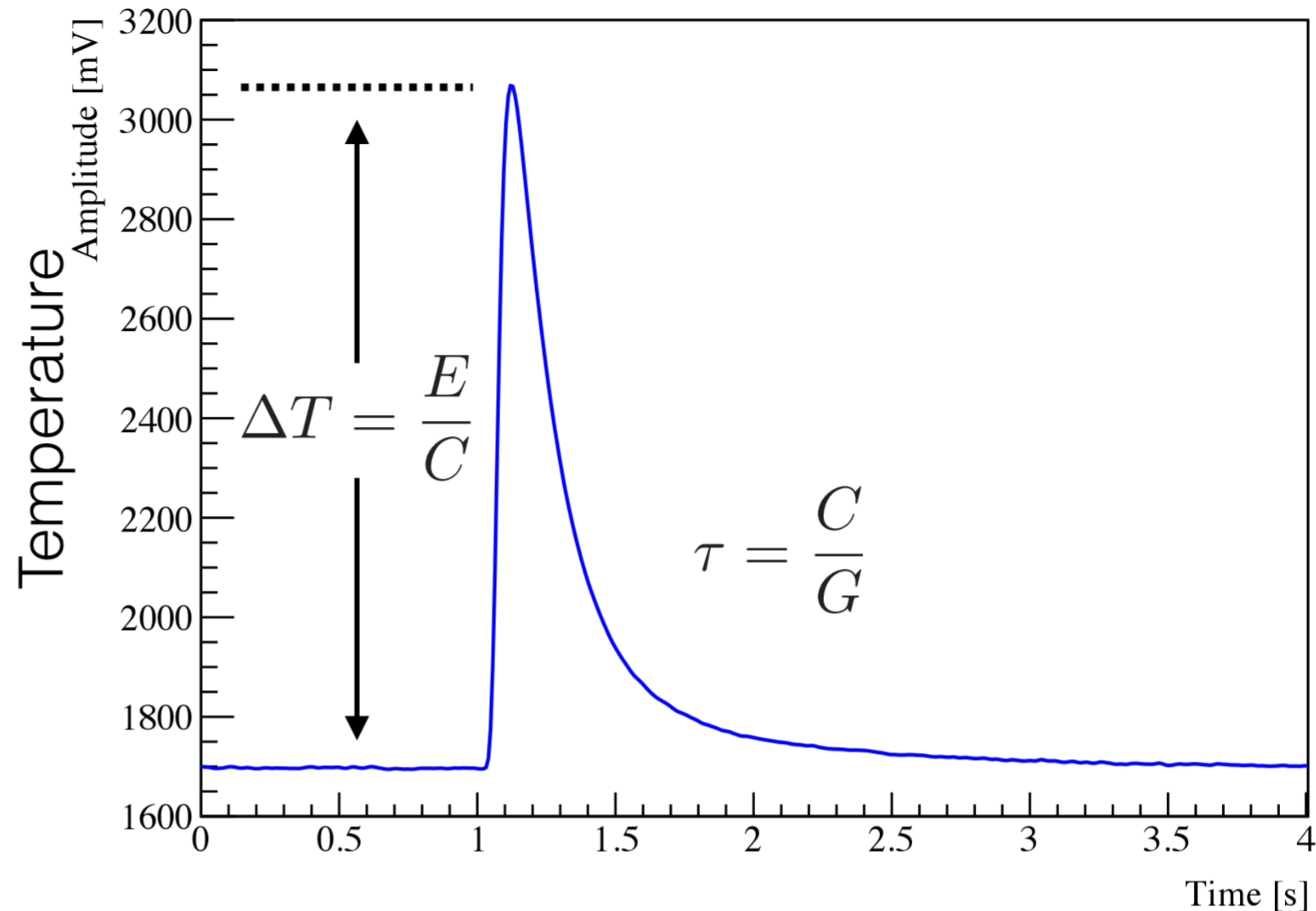
Adv. in High En. Phys. 2015, 879871

Working principle



- The **energy** released in a particle interaction is measured via **thermal excitations** (*phonons*)
- The **temperature increase** is converted into an **electric signal** by a cryogenic sensor (e.g. a thermistor)
- We use a Si heater to inject stable voltage pulses and do **thermal gain stabilization**

Working principle



$$\Delta T = \frac{\Delta E}{C} \sim \frac{100 \mu K}{MeV}$$

$$\tau = \frac{G}{C} \sim 1s$$

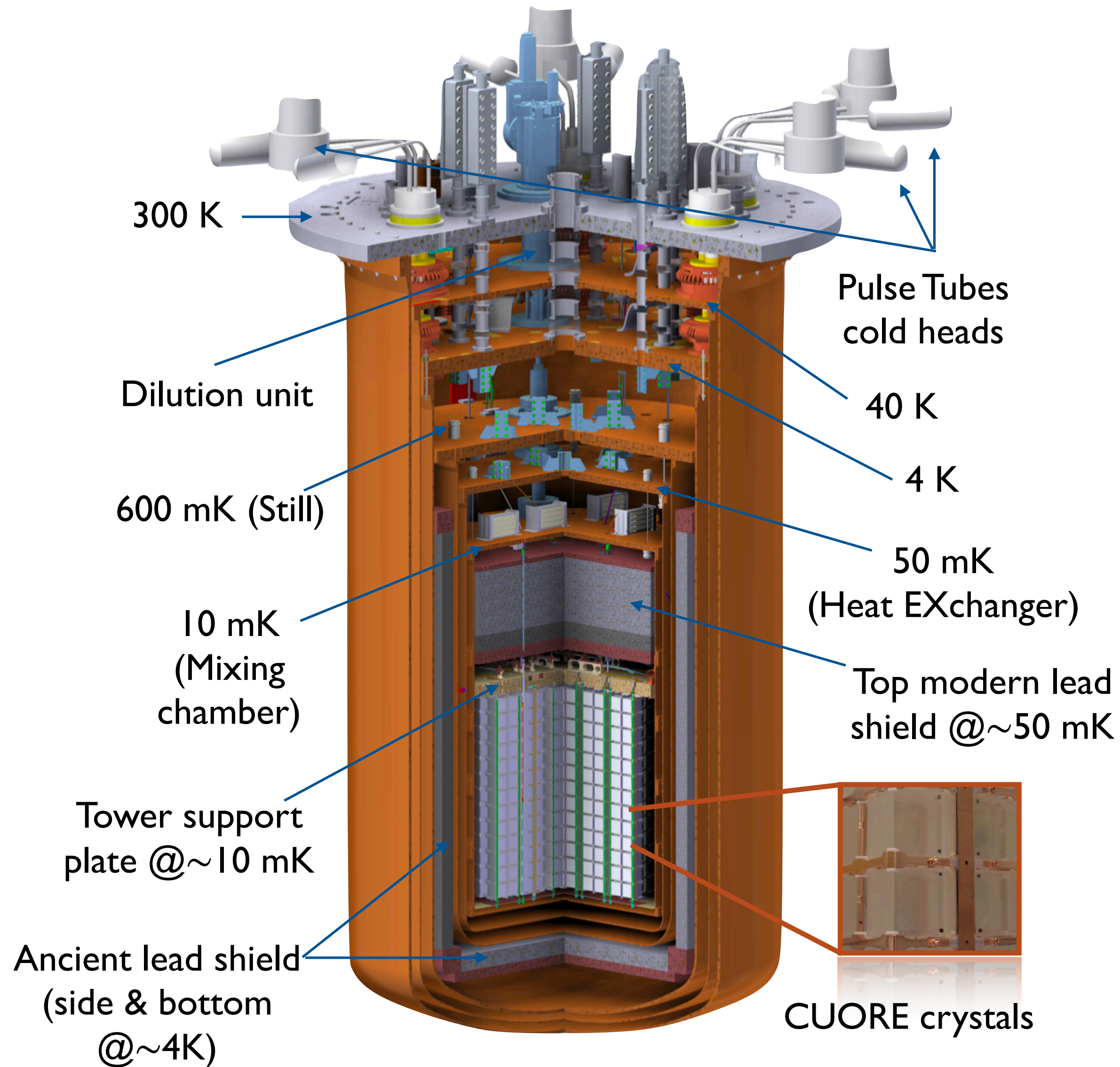
$$C(T) \propto T^3$$

$$R(T) = R_0 e^{\sqrt{T_0/T}}$$

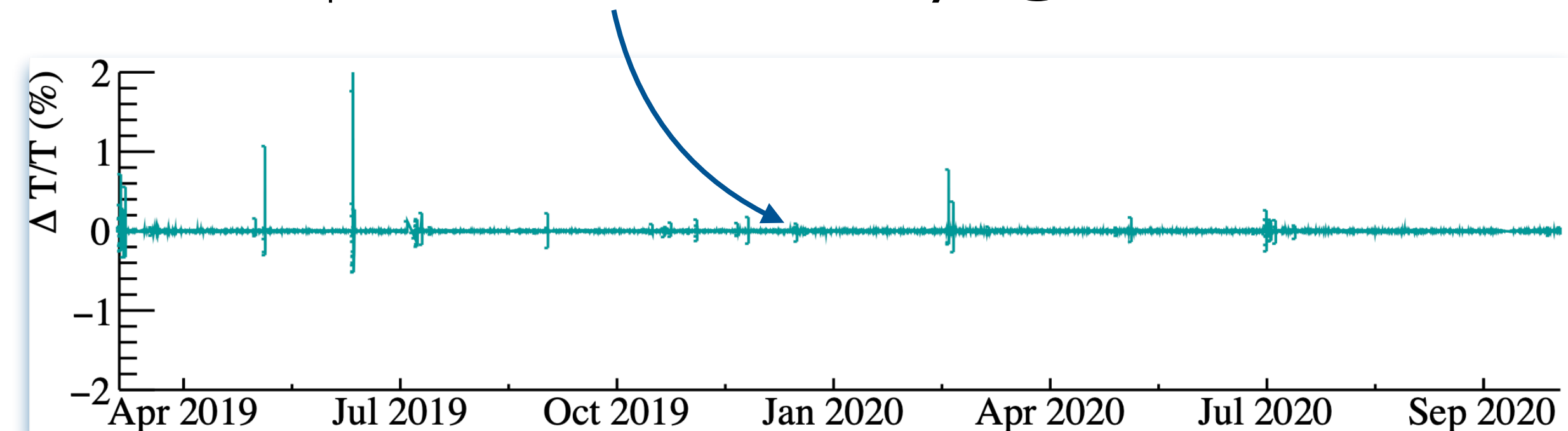
C: absorber capacity
 ΔT : temperature variation
 ΔE : energy deposition
G: thermal conductance
t: signal decay time

- Operating at a temperature of **10-15 mK**:
- **1 MeV** energy release causes **$\Delta T \sim 100 \mu K$** (**$\Delta R \sim 3 M\Omega / MeV$**)
- excellent energy resolution (few % FWHM @ $Q_{\beta\beta}$)
- no radiation discrimination
- slowness (suitable for rare event searches)

The challenges - part I: cryogenic infrastructure



- Custom-made dry (cryogen-free) dilution cryostat
- Stringent constraints on **radiopurity** - only few materials acceptable (Cu OFE/NOSV for plates and vessels) - **mechanical reliability** and **response** to **seismic events**
- 5 **pulse-tube** cryocoolers (one/two spares)
- 6 **nested vessels** at decreasing temperature to reach base temperature ($T_{\min} \simeq 7$ mK)
- Mass <4 K: 15 tons - mass @~50 mK: 3 tons
- Since 2019 the system is operating with **>90% uptime** in stable temperature conditions: **>5 yr @ <20 mK**



The challenges - part II: minimizing background

LNGS *natural* shielding



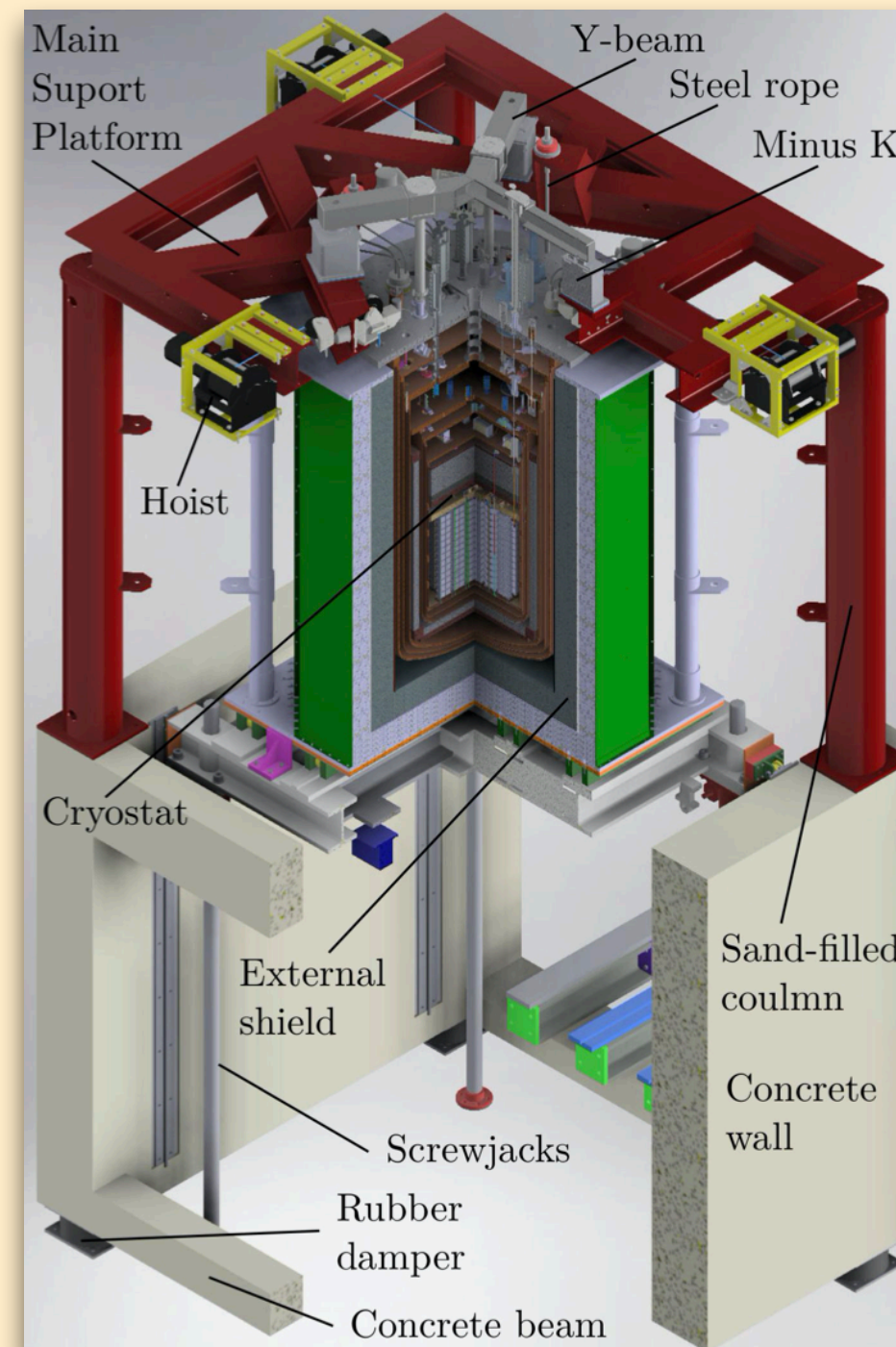
~3600 m water-equivalent rock overburden:
cosmic rays flux reduced of six orders of magnitude compared to the surface

$$\Phi_{\mu} = 3 \cdot 10^{-8} \text{cm}^{-2} \text{s}^{-1}$$

 Astropart Phys. 34 (2010) 18-24

$$\Phi_n = 4 \cdot 10^{-6} \text{cm}^{-2} \text{s}^{-1}$$

External shields



 Cryogenics 102 (2019) 9-21

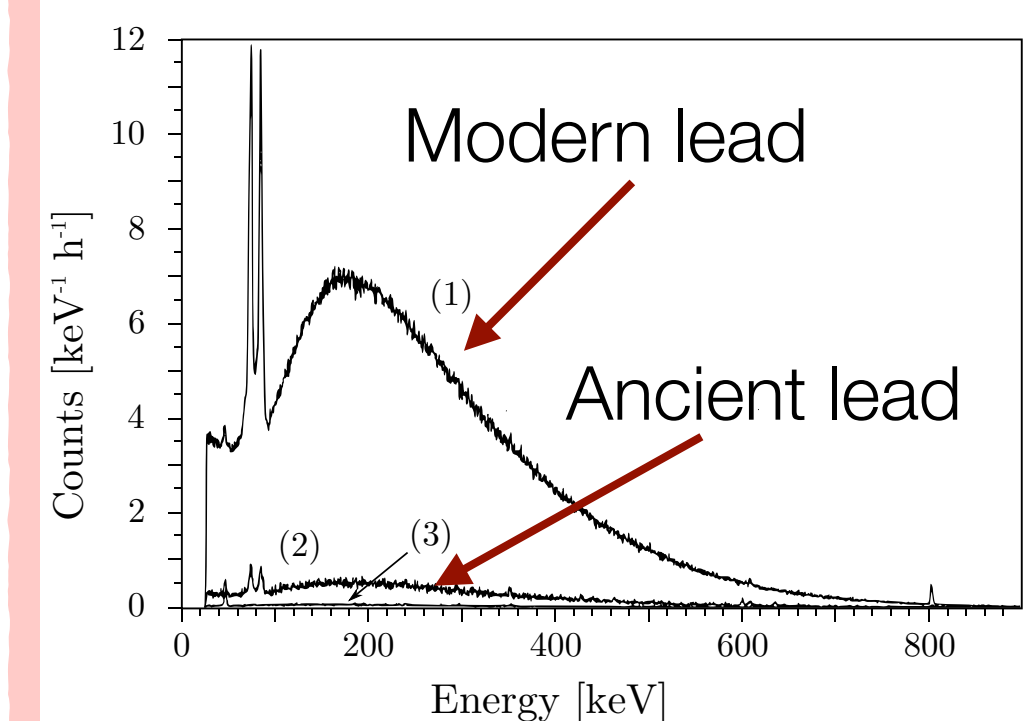
- ▶ From γ s: 25-cm thick Pb layer
- ▶ From neutrons: 20-cm layer in polyethylene + H_3BO_3 panels

Internal lead shields

Top: 30-cm modern lead

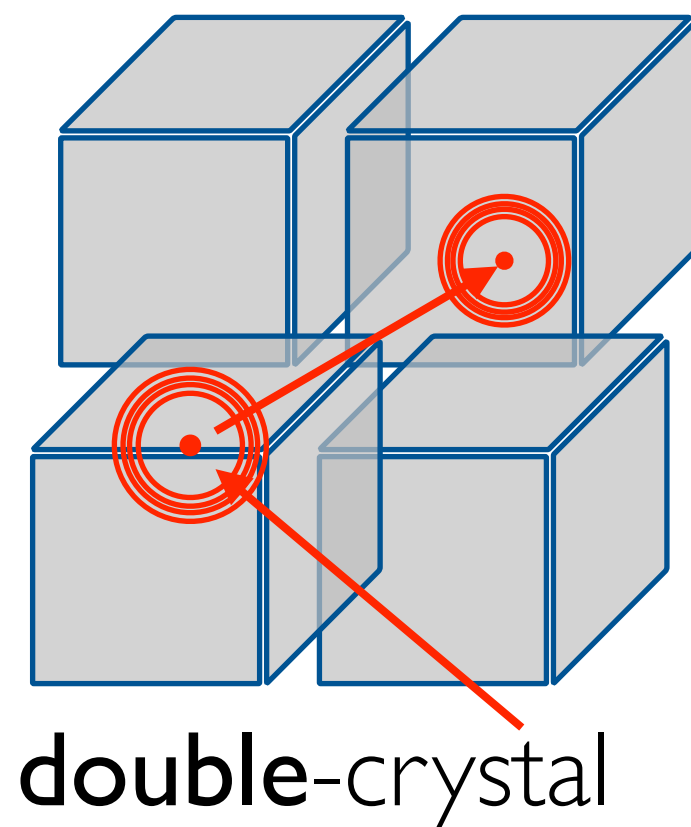
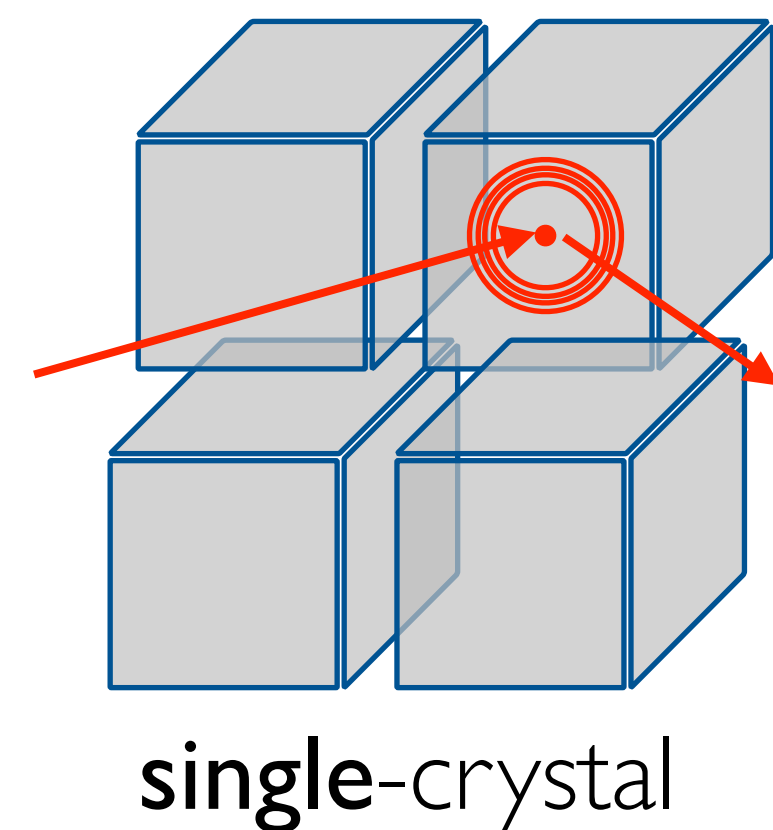


Side & bottom: 6-cm ancient roman lead from a shipwreck
 $^{210}\text{Po} < 4 \text{ mBq/kg}$



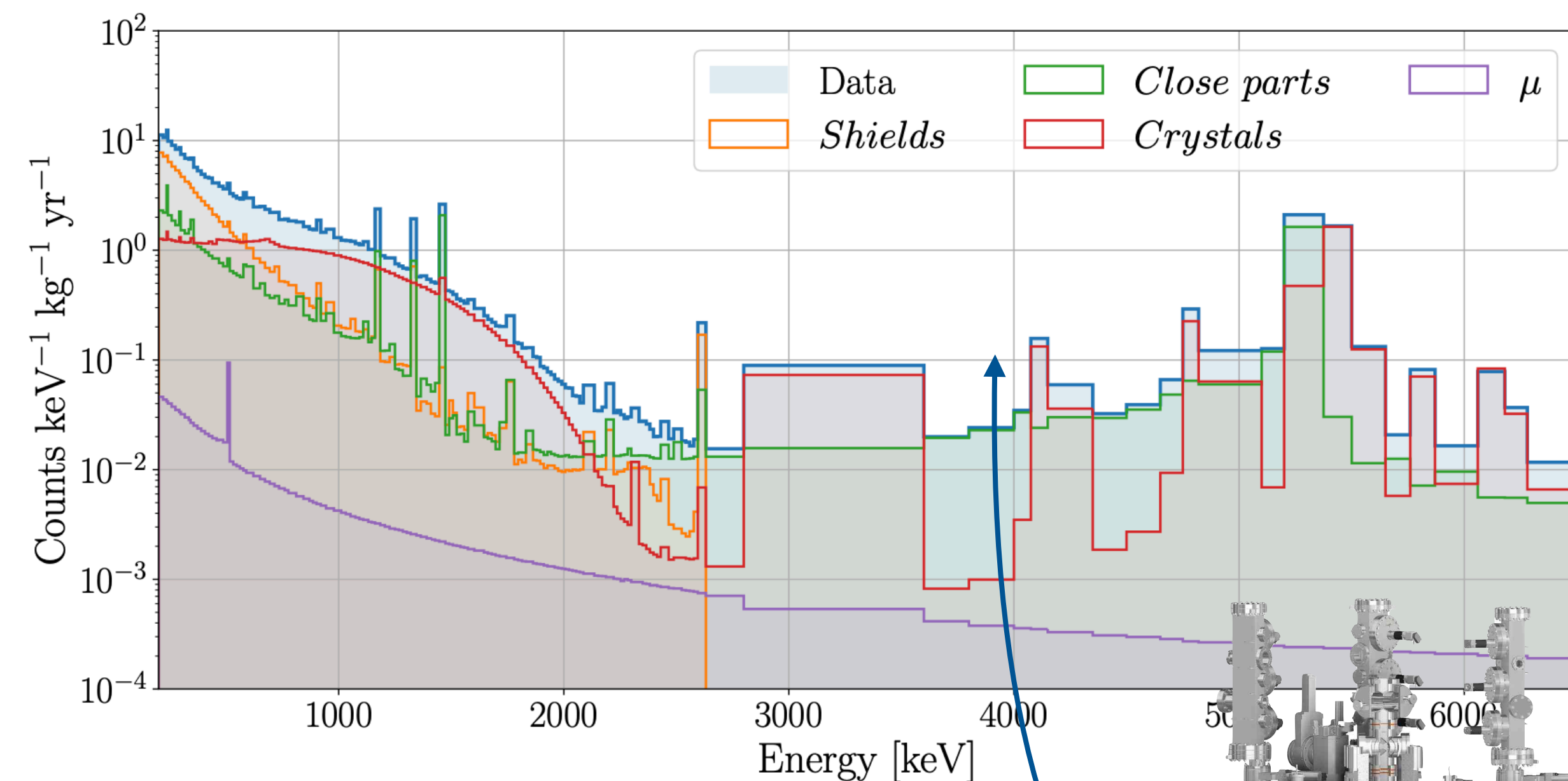
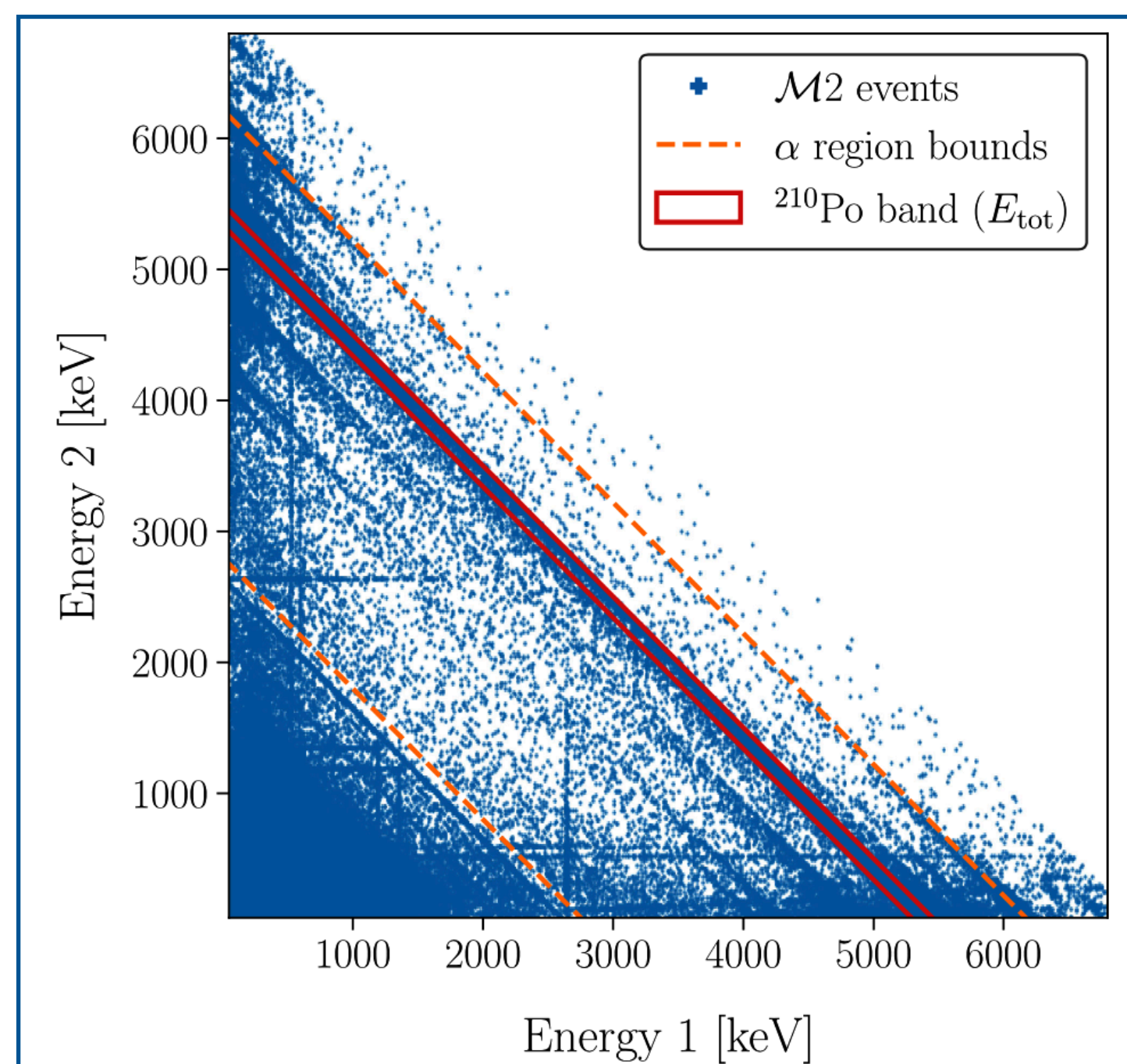
 Nucl. Instrum. Meth. B 142, 163 (1998)

The challenges - part II: modelling (leftover) background



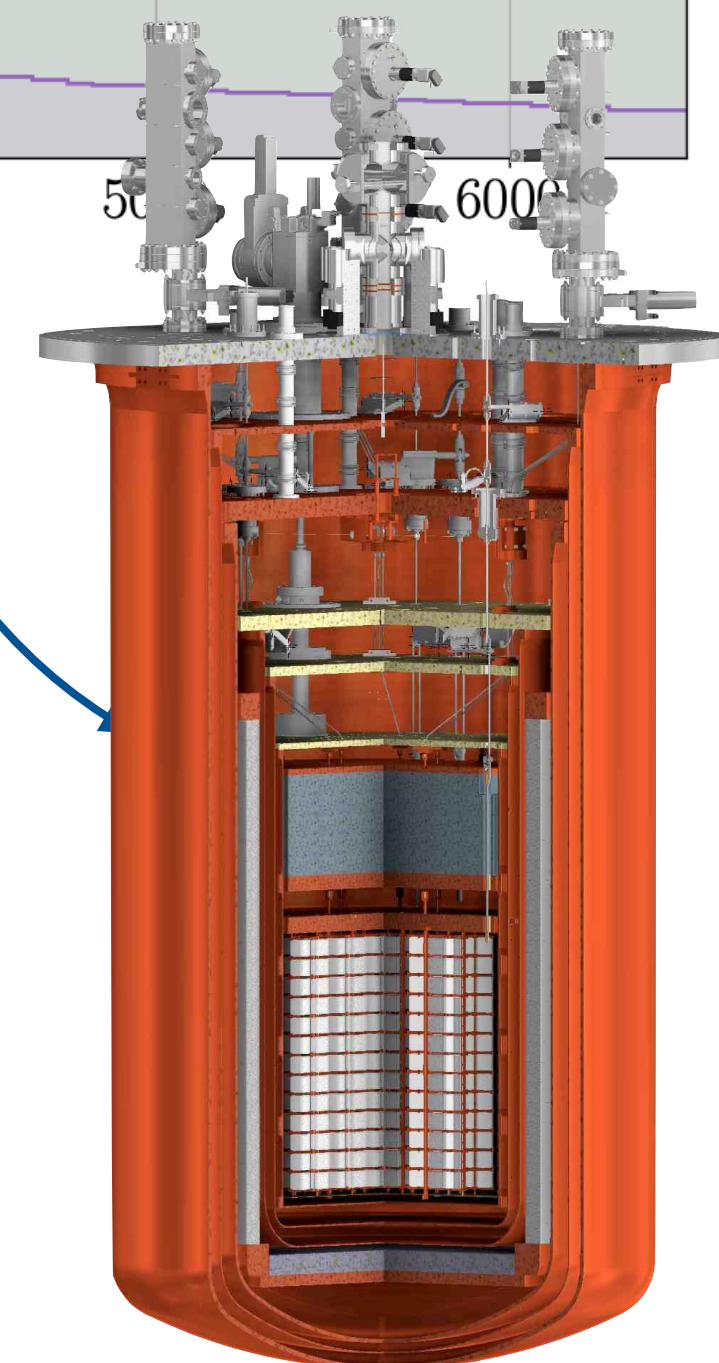
Priors extracted from radioassays and past experiments

Bayesian fit of *single*-crystal & *double*-crystal events



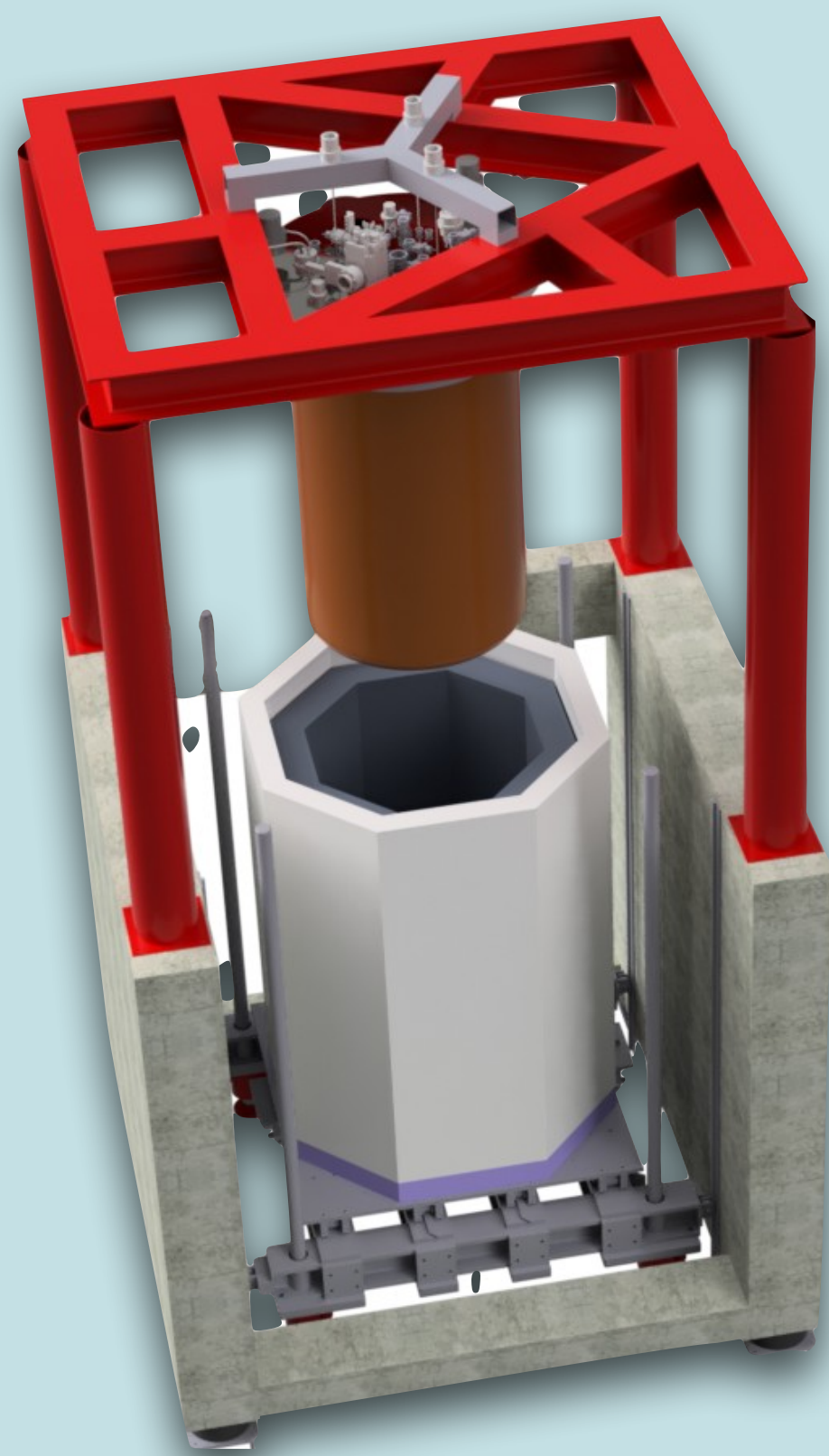
Accurate Geant-4 based background model with 80 sources thanks to high detector granularity, relating **sources** and specific **spectral contributions**

Precise measurement of ^{130}Te $2\nu\beta\beta$ decay

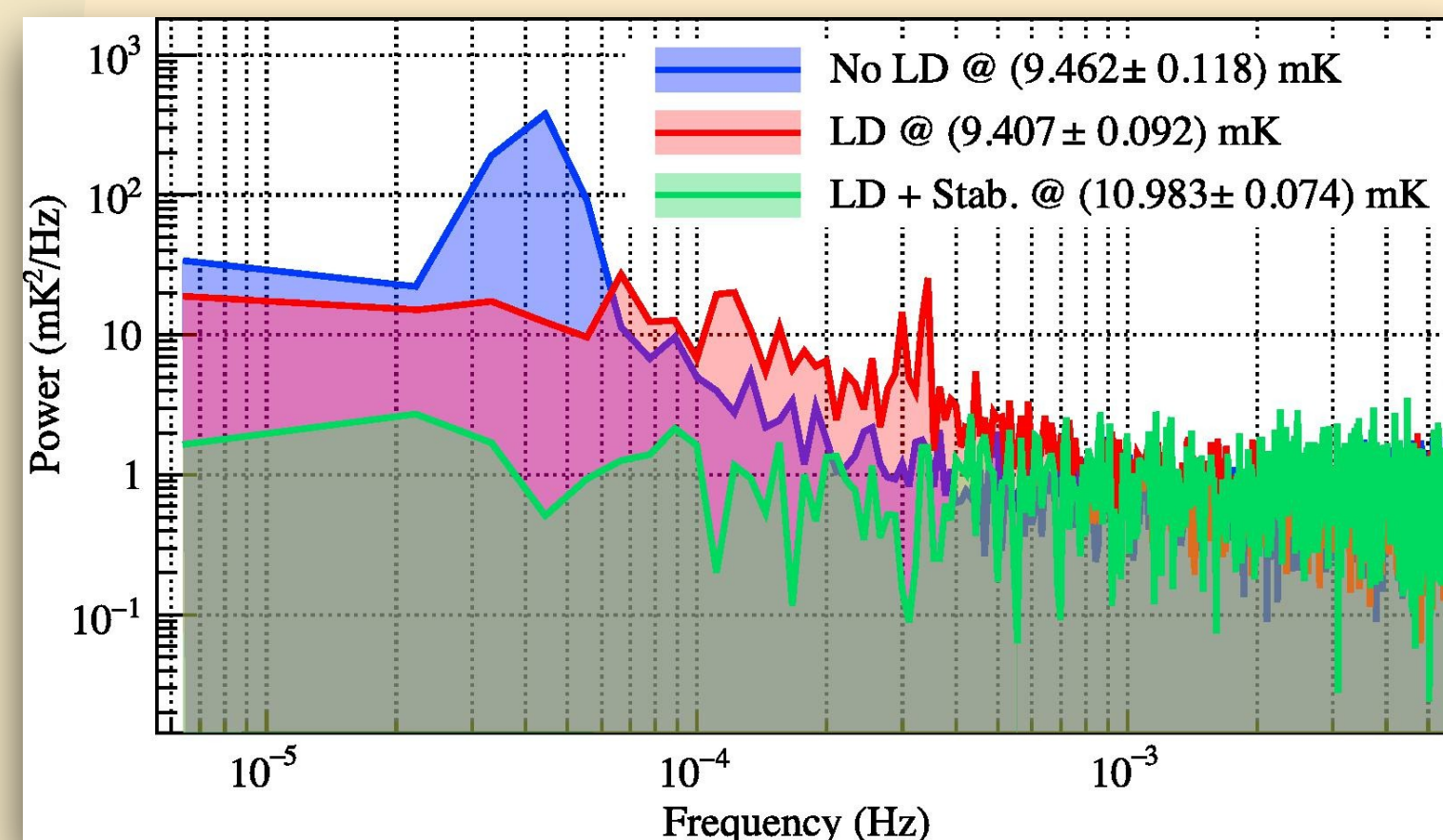


The challenges - part III: suppressing noise sources

External structure to decouple the detector from the cryostat



Linear drives and



active noise cancellation to minimize vibrations induced by the **pulse tubes**

 Cryogenics 93, 56-65 (2018)

Several **ancillary devices** installed in the CUORE hut to do data denoising and enhance the quality of data

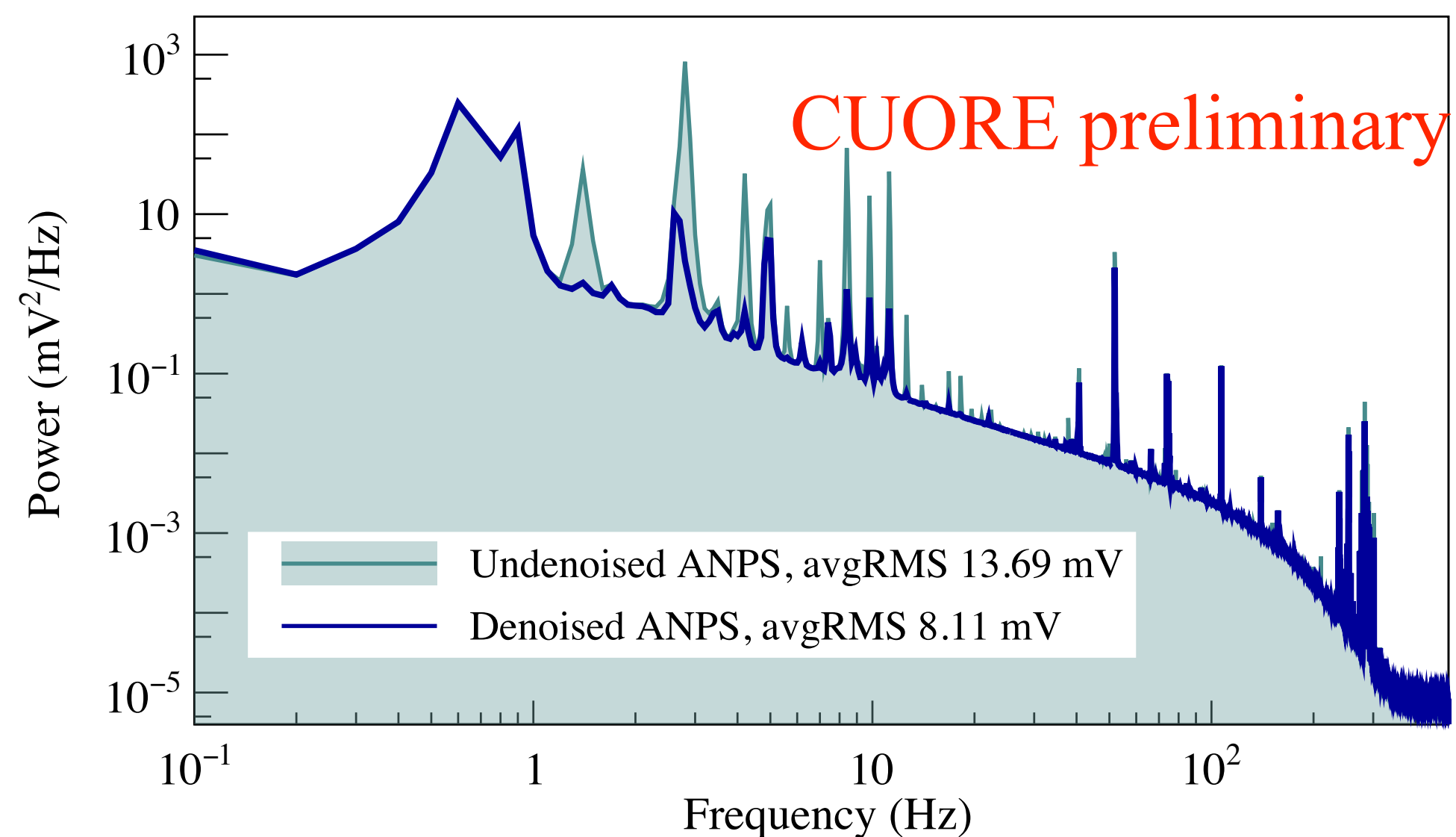


- seismometers, antennae, microphones and accelerometers

 Eur. Phys. J. C 84, 243 (2024)

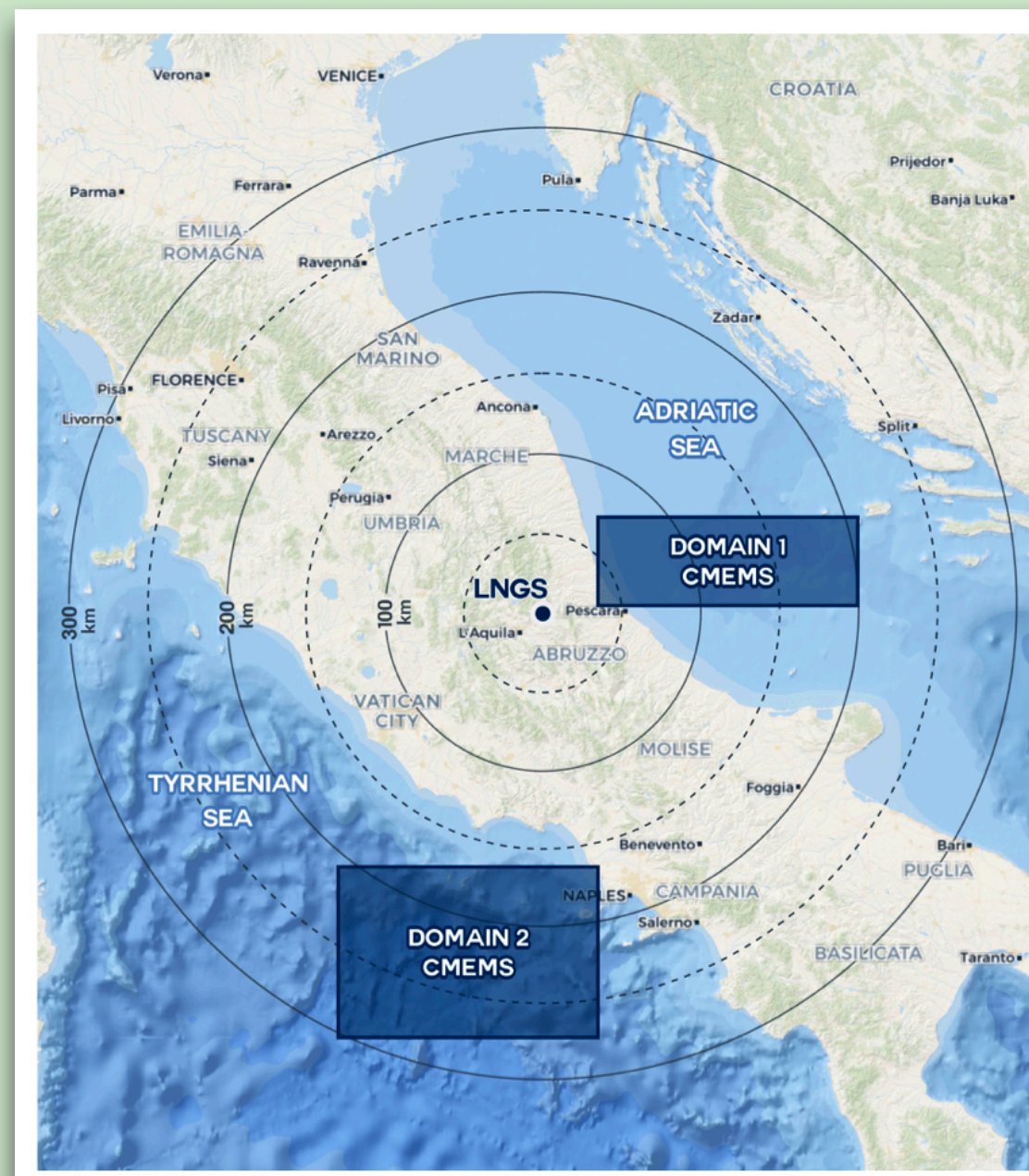
The challenges - part III: modelling (leftover) noise

- **Noise decorrelation algorithms** rely on the correlation between noise power spectra of detector channels and all diagnostic devices installed in the experimental hut:



40% total raw-RMS reduction w/ denoising

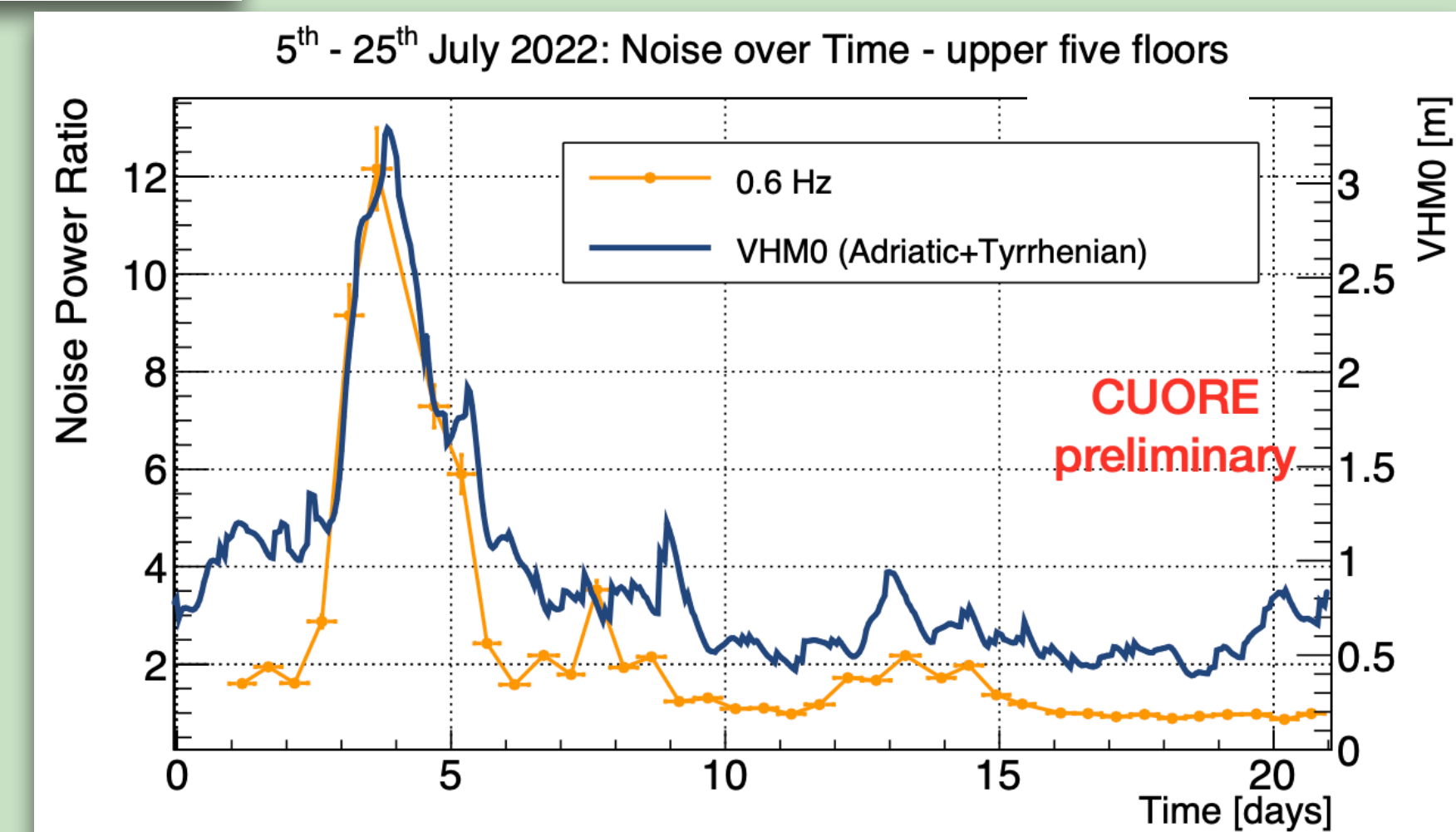
 [arxiv:2404.04453](https://arxiv.org/abs/2404.04453)



- Recent discovery: CUORE is sensitive to **microseismic activity** induced by the **sea waves** (0.2 - 0.3 Hz)
- Evident correlation **storms** ↔ low frequency noise

 [arxiv:2505.09652](https://arxiv.org/abs/2505.09652)

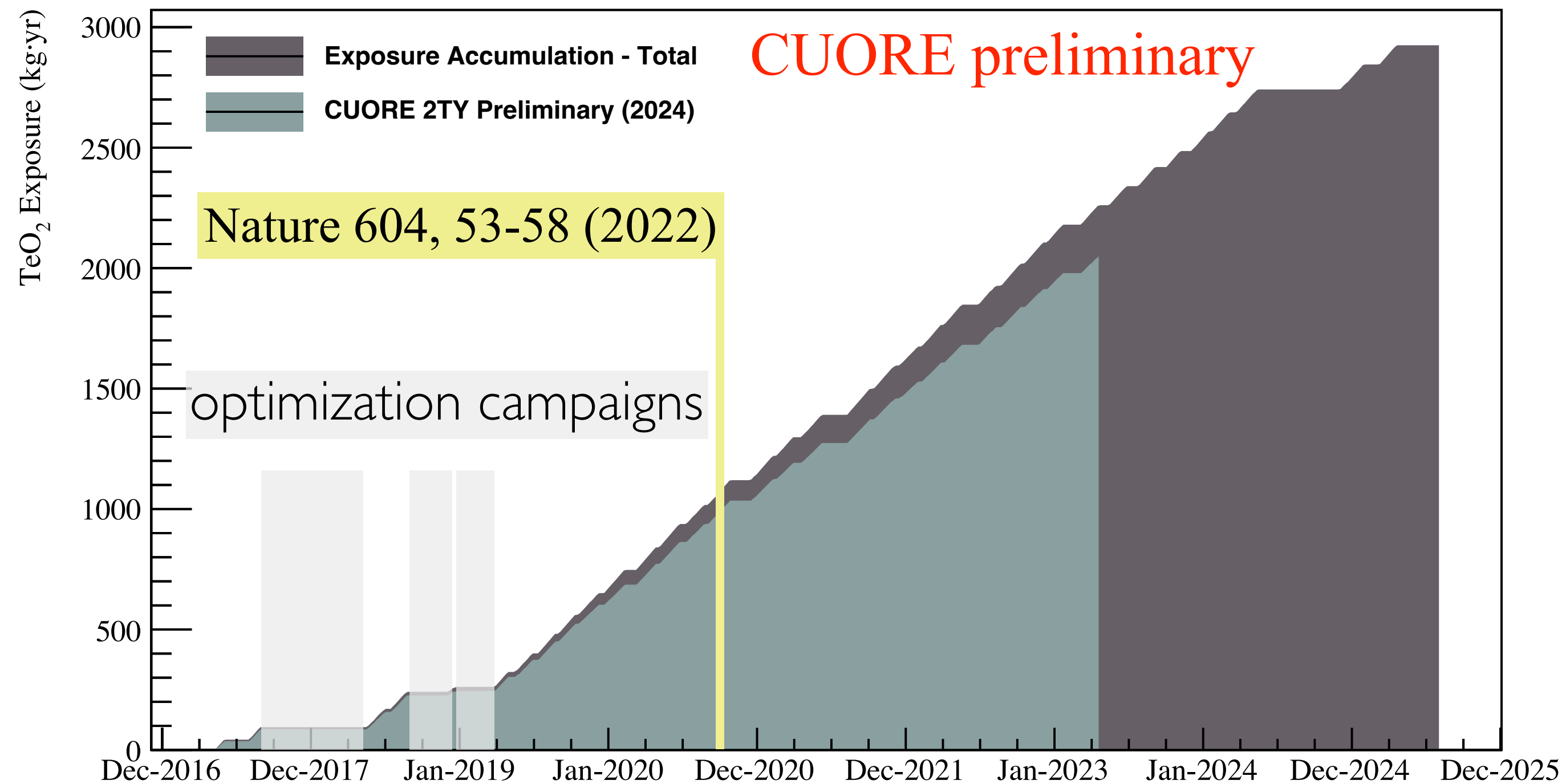
- Under investigation solutions to improve cryostat decoupling



The challenges - part IV: stable operation of 1000 detectors

- Data taking started in 2017

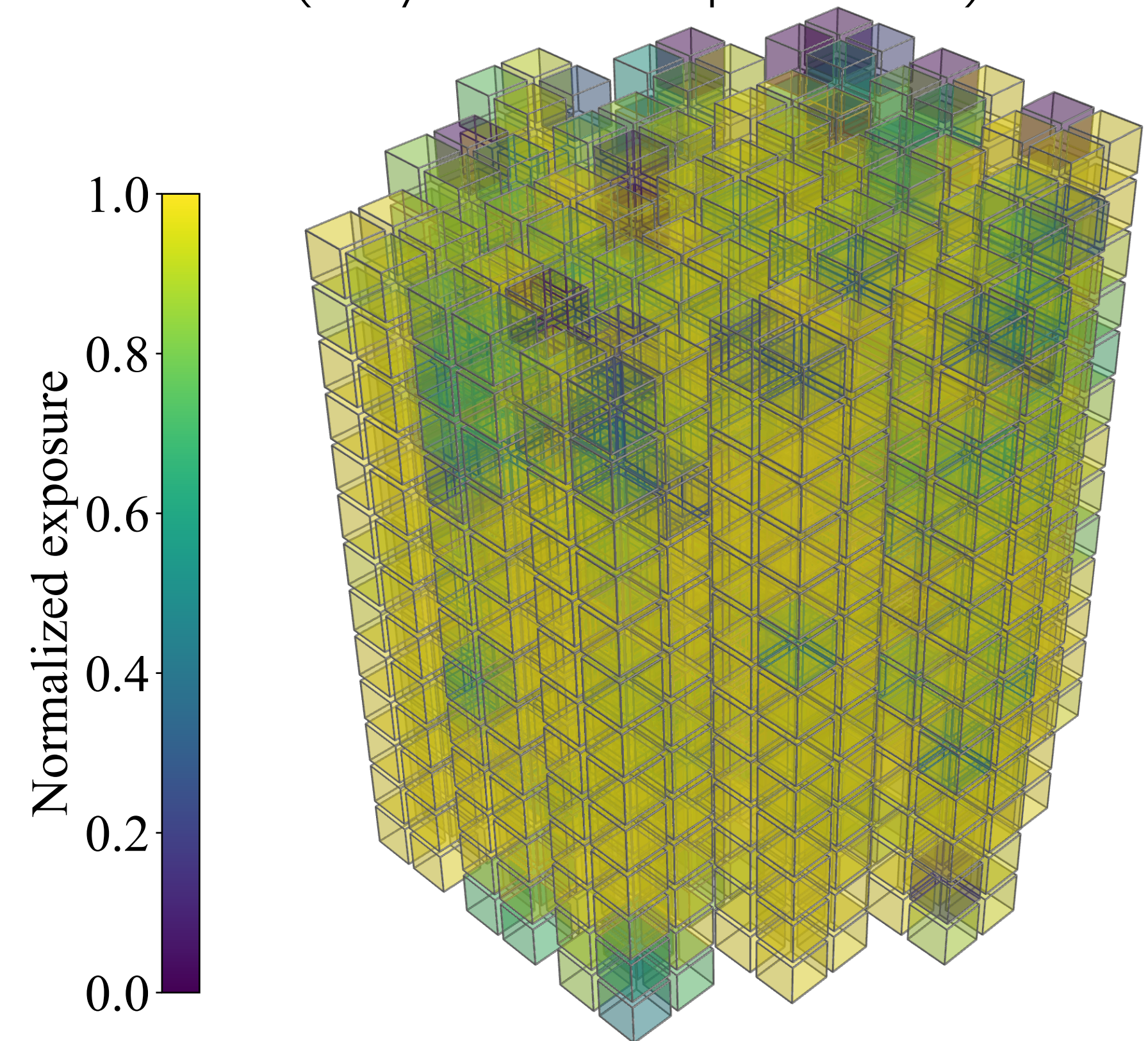
28 datasets (May 2017 - April 2023)



- Data split in *datasets*

calibration physics runs calibration

- Trigger rate 50 mHz (~ 6 mHz) in calibration (physics) runs
- Voltage across NTD Ge thermistors sampled at 1 kHz, a software trigger is applied offline



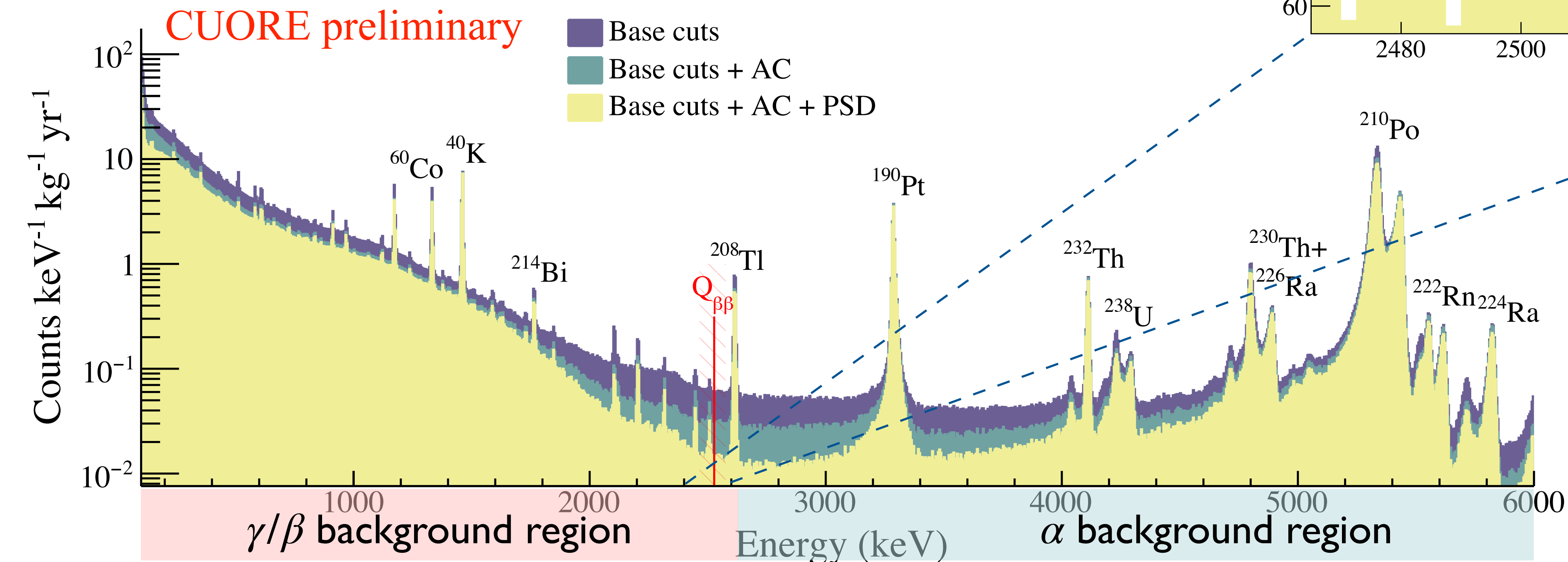
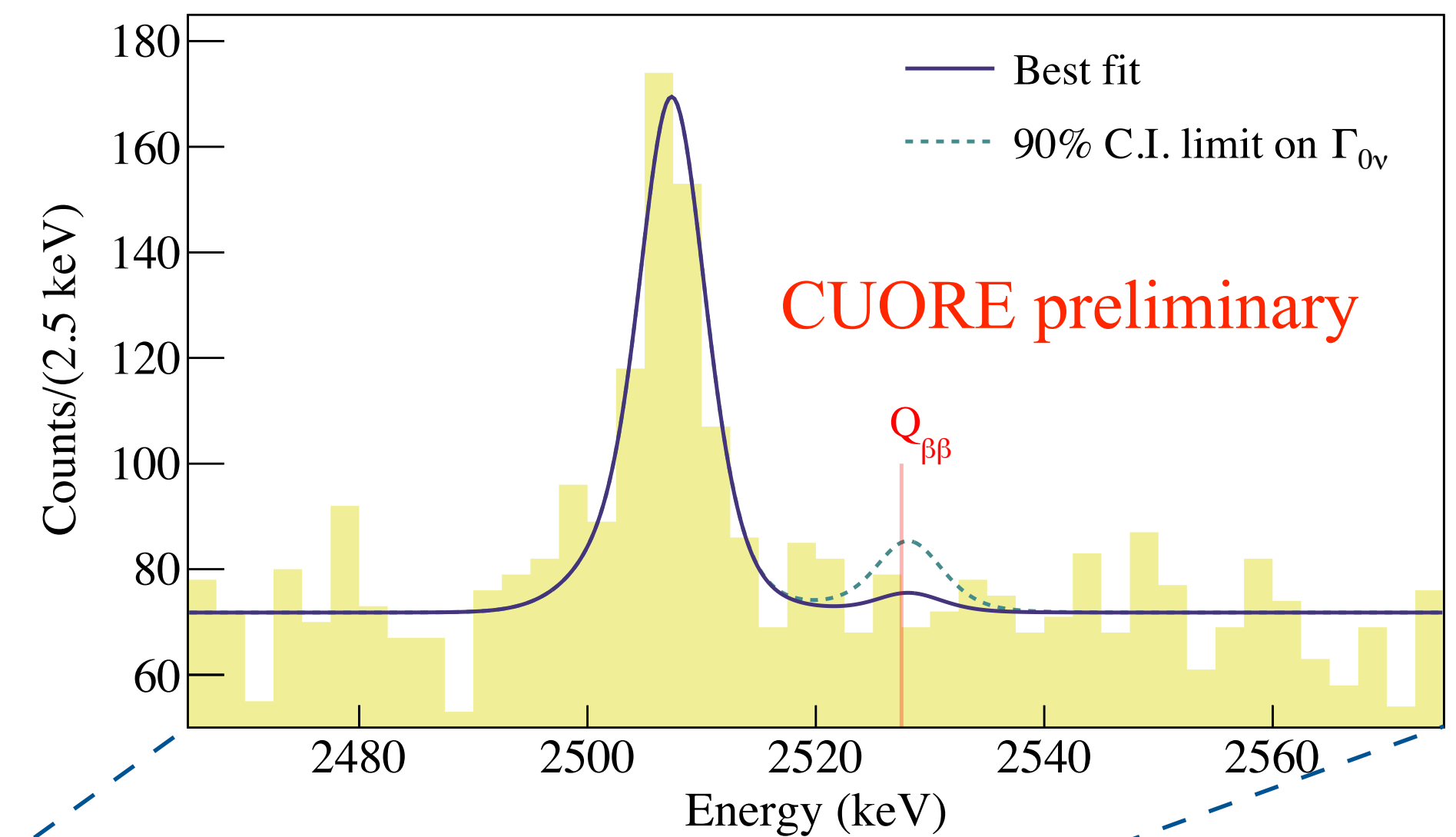
TeO₂ exposure: **2039.0 kg·yr**

¹³⁰Te exposure: **567.0 kg·yr**

uniformly distributed on the detector

The 2 tonne-yr data spectrum

- Several analysis cuts on top of each other:
 - Base cuts** (trigger, energy reconstruction, pile-up)
 - Anti-coincidence** (AC): only single-crystal events
 - Pulse shape discrimination** (PSD): only signal-like events
- Total analysis cut efficiency **93.4(18) %**

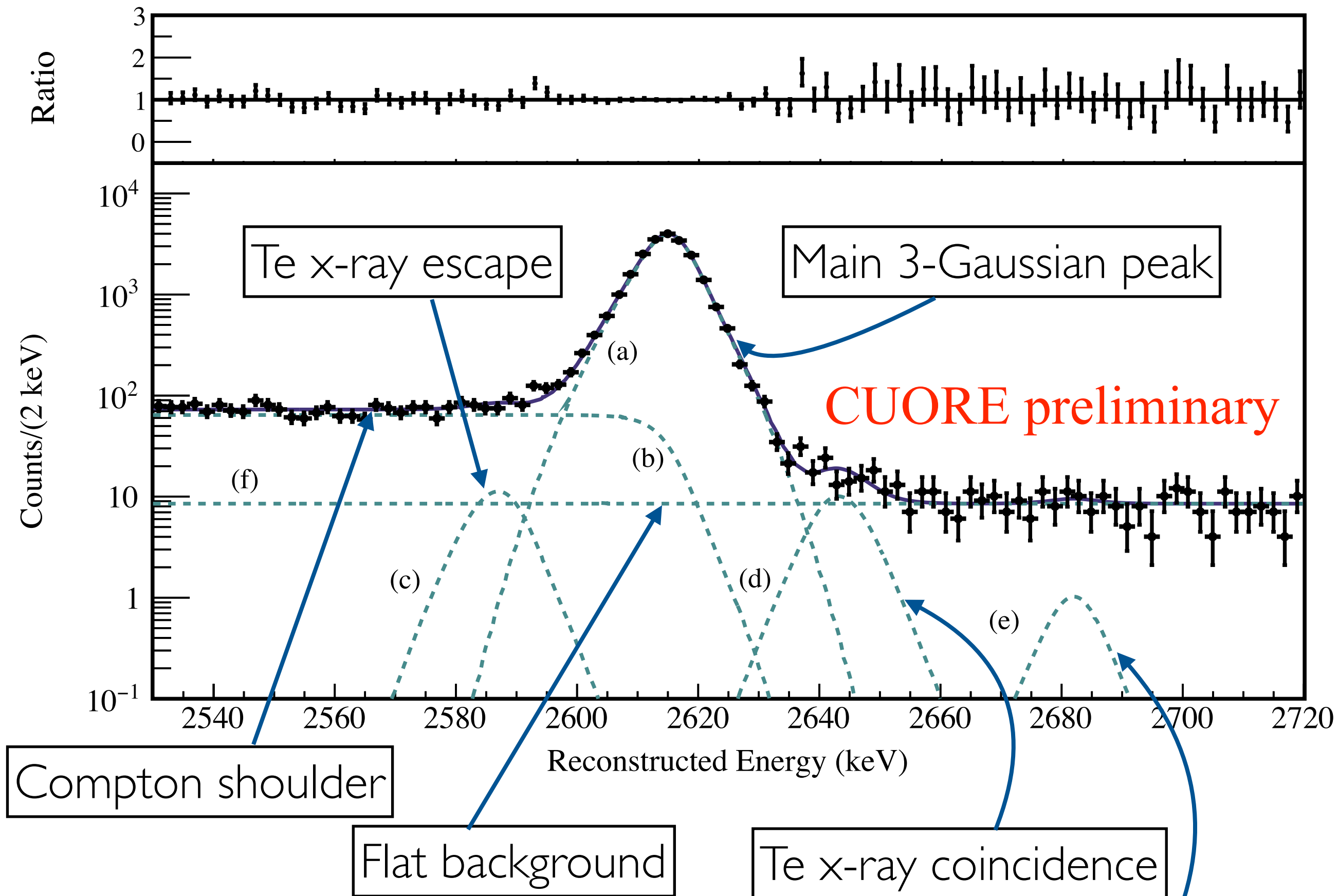


Average background index in the ROI

$$b = (1.42^{+0.03}_{-0.02}) \cdot 10^{-2} \text{ (counts/keV/kg/yr)}$$

Pending publication

Modeling detector response

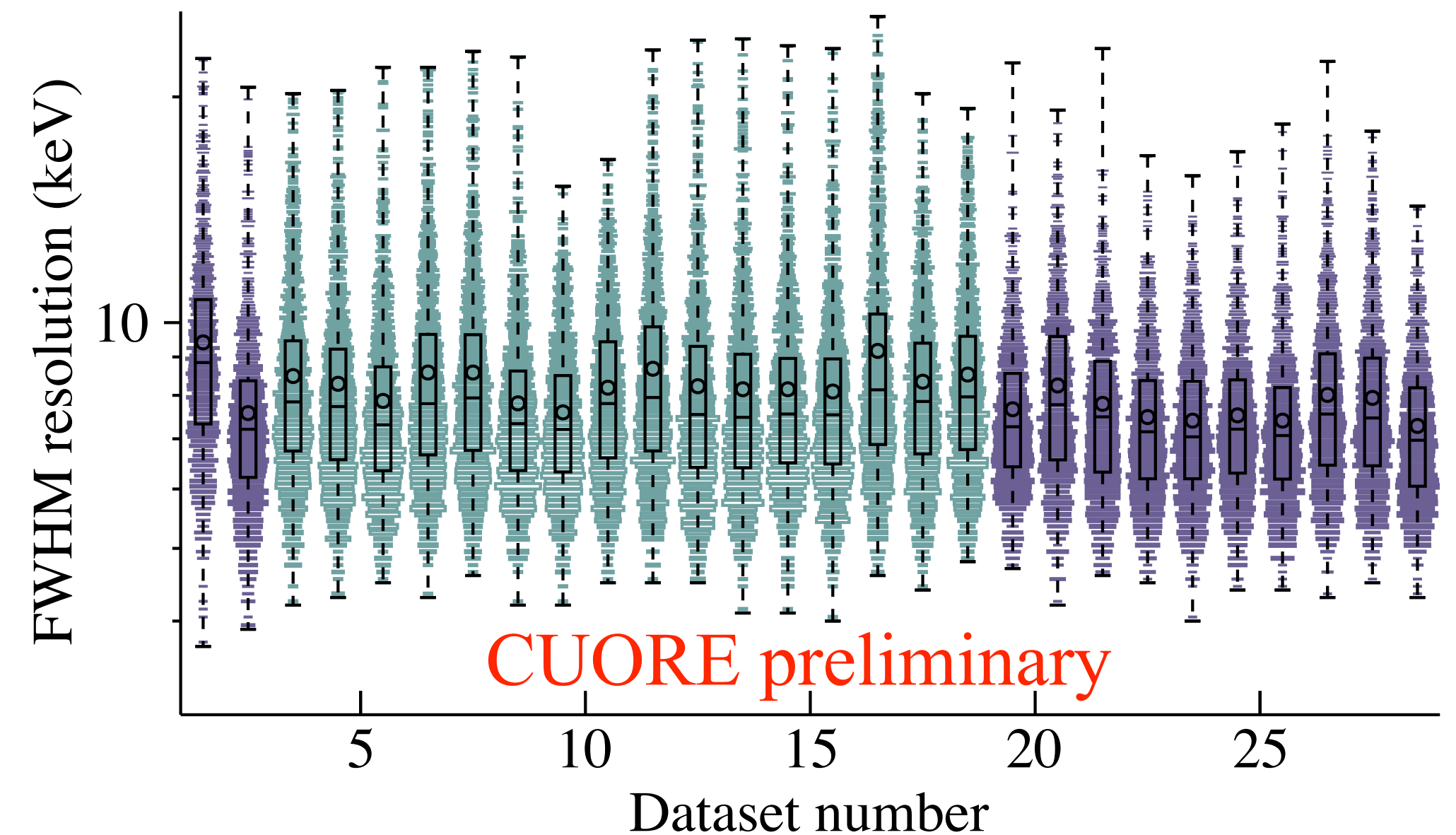
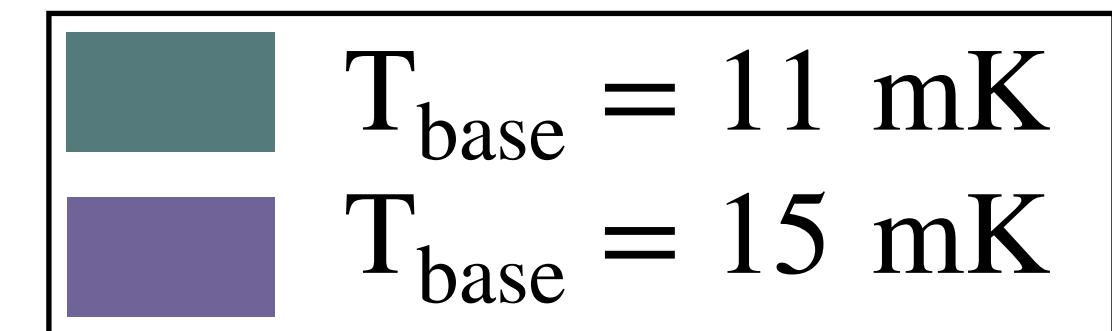


$$\text{FWHM} (^{208}\text{Tl}) = (7.540 \pm 0.024) \text{ keV}$$

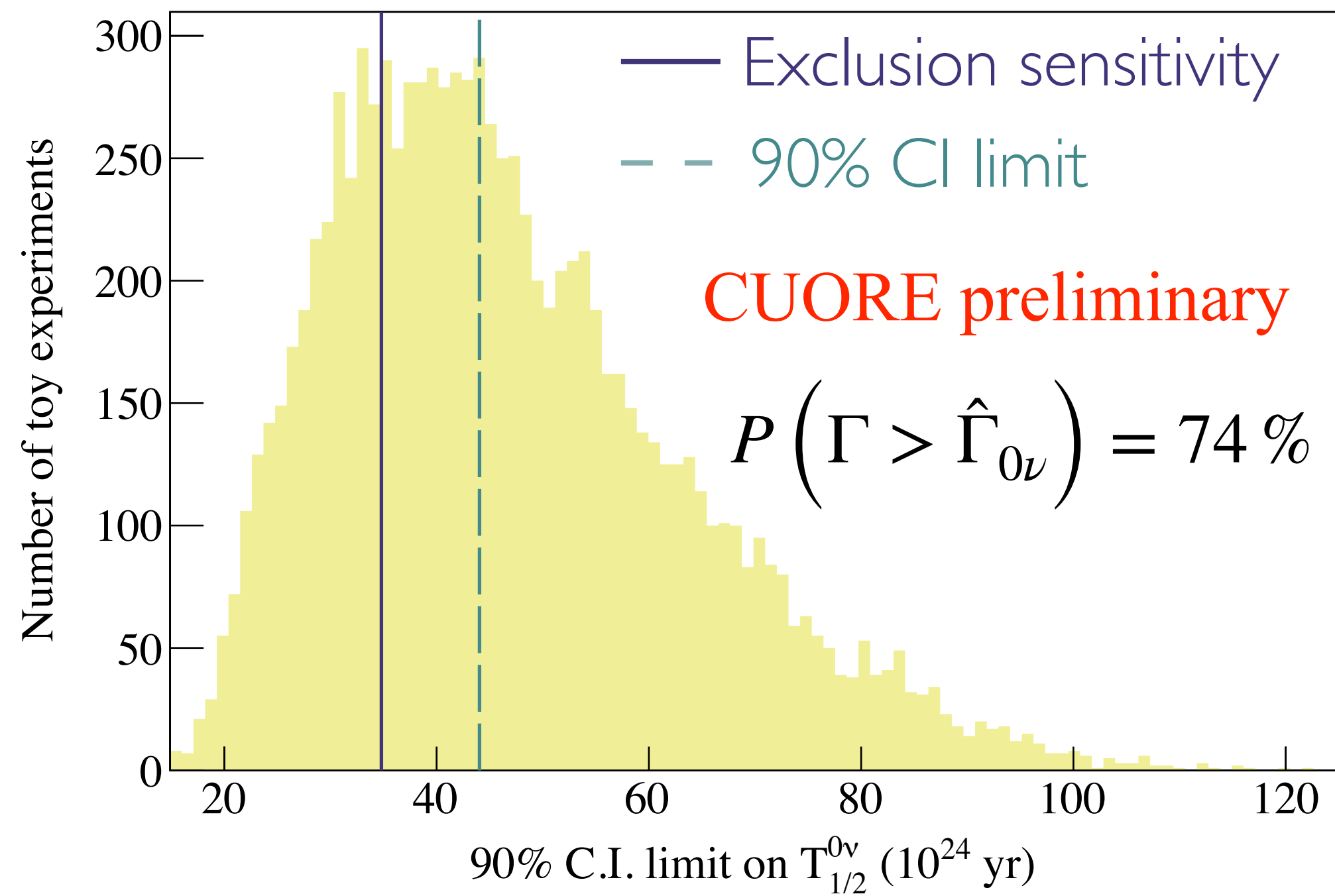
$$\text{FWHM} (Q_{\beta\beta}) = (7.310 \pm 0.024) \text{ keV}$$

$$\Delta E (Q_{\beta\beta}) = (0.40^{+0.21}_{-0.44}) \text{ keV}$$

- Detector response extracted on events from the ^{208}Tl line at 2615 keV in calibration data separately for each bolometer and dataset
- Fit of the most prominent γ lines in physics data to scale the energy resolution and calibration bias at $Q_{\beta\beta}$



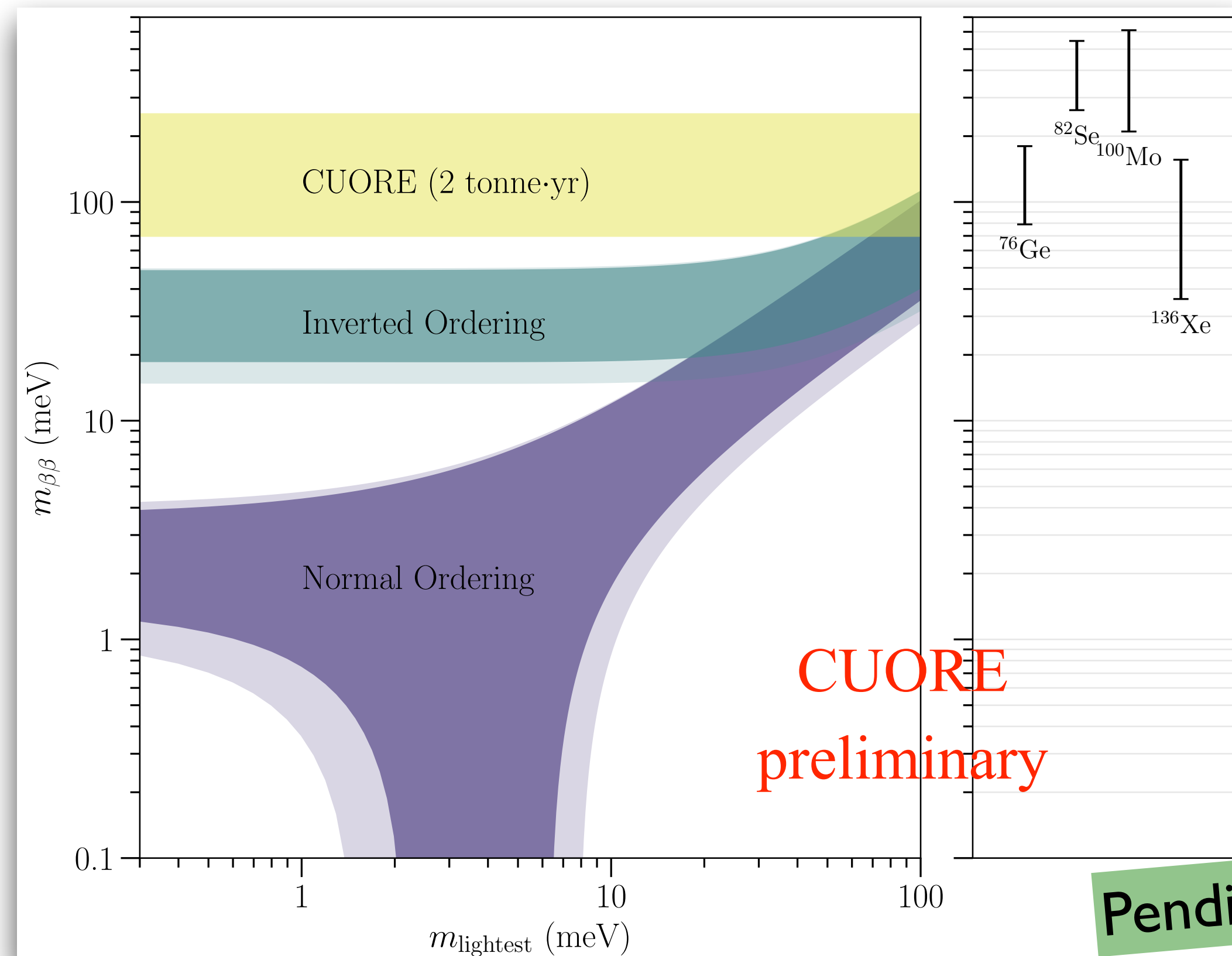
Fitting our data to extract $\Gamma_{0\nu\beta\beta}$



Frequentist result $T_{0\nu\beta\beta}^{1/2} > 3.4 \cdot 10^{25}$ yr (90 % C.L.)

Assuming the exchange of a light Majorana neutrino
the limit on the effective Majorana mass is

$$m_{\beta\beta} < 70 - 250 \text{ meV}$$

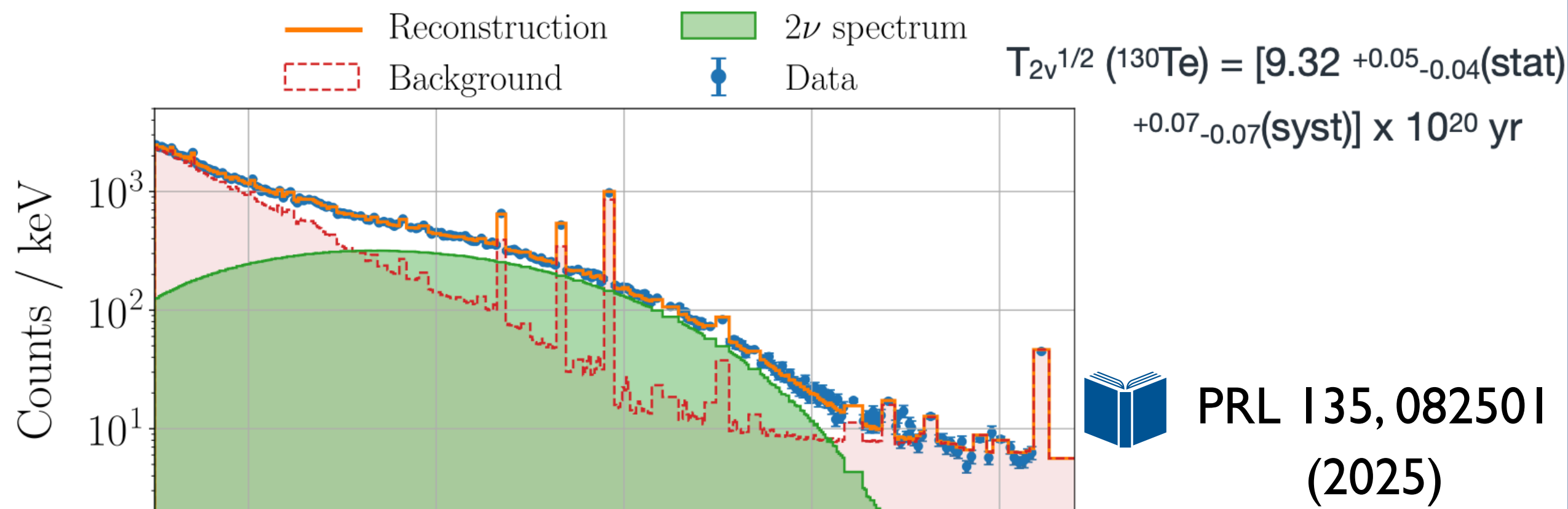


Pending publication

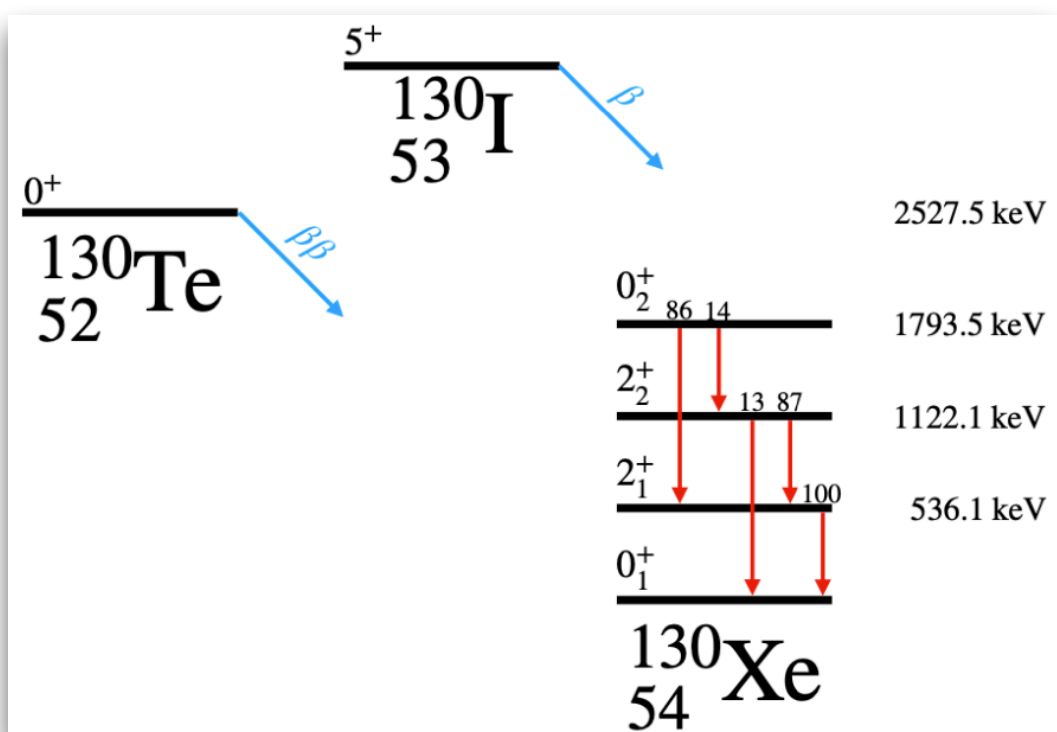
- Median exclusion sensitivity from toy MC experiments $T_{0\nu\beta\beta}^{1/2} = 4.4 \cdot 10^{25}$ yr (90 % C.I.)
- Unbinned Bayesian fit with $\Gamma_{0\nu\beta\beta} > 0$
- No evidence of $0\nu\beta\beta$ and new limit on ^{130}Te half-life $T_{0\nu\beta\beta}^{1/2} > 3.5 \cdot 10^{25}$ yr (90 % C.I.)

Not only ^{130}Te $0\nu\beta\beta$: other $\beta\beta$ searches

^{130}Te SM-allowed $2\nu\beta\beta$ decay



^{130}Te $\beta\beta$ decay to the 1st 0^+ excited state

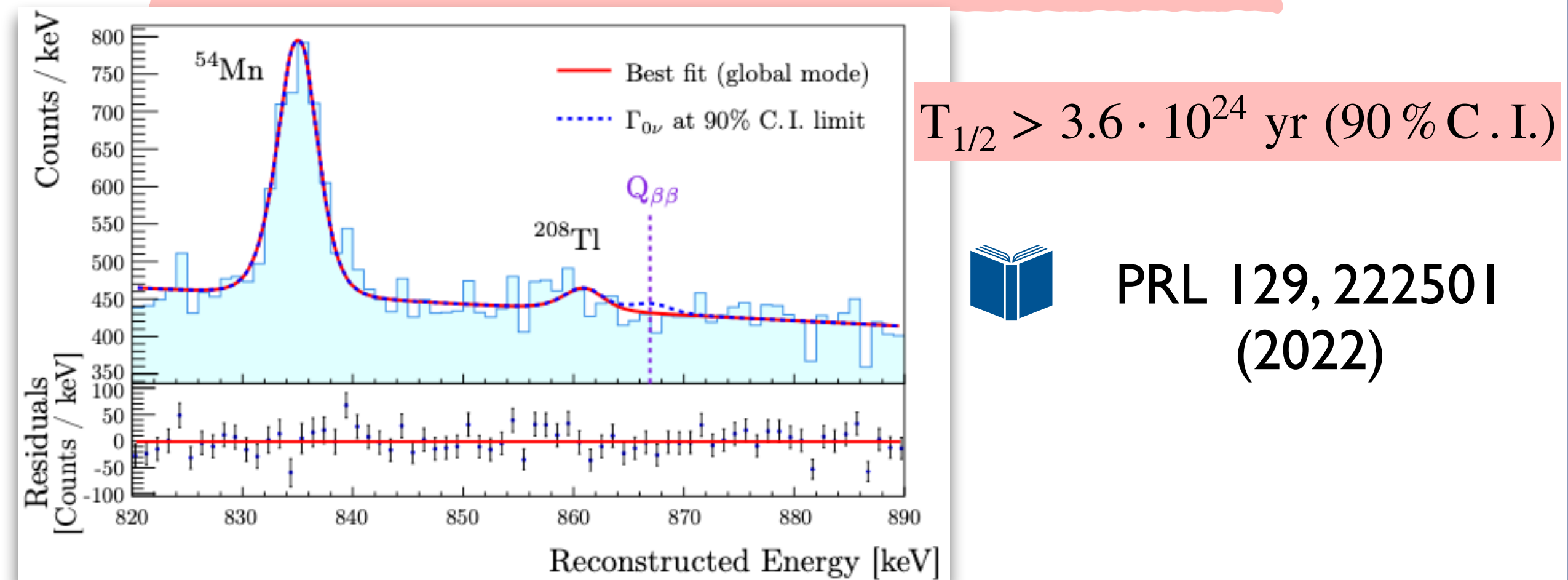


$$T_{0\nu}^{1/2} > 5.9 \cdot 10^{24} \text{ yr (90 \% C.I.)}$$

$$T_{2\nu}^{1/2} > 1.3 \cdot 10^{24} \text{ yr (90 \% C.I.)}$$

 Eur. Phys. J. C 81, 567 (2021)

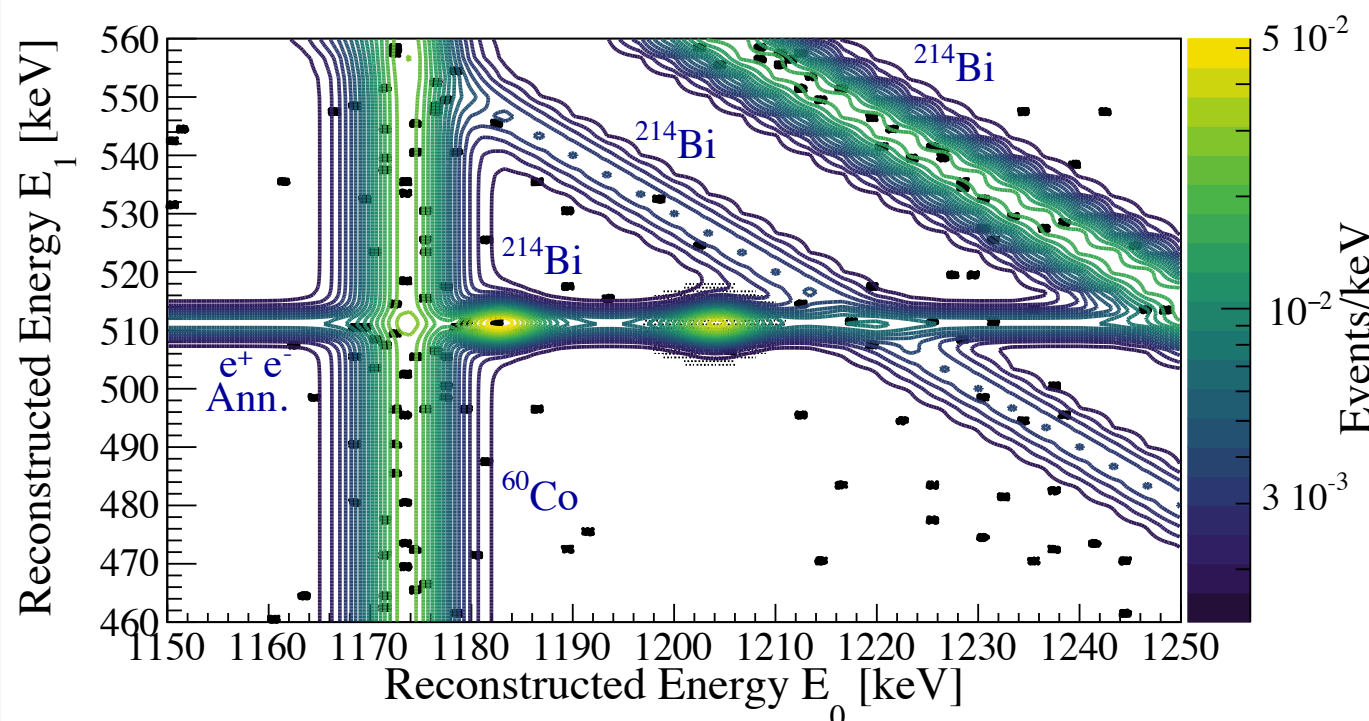
^{128}Te $0\nu\beta\beta$ decay to the ground state



^{120}Te $0\nu\beta^+\text{EC}$ decay to the ground state

$$T_{1/2} > 2.9 \cdot 10^{22} \text{ yr (90 \% C.I.)}$$

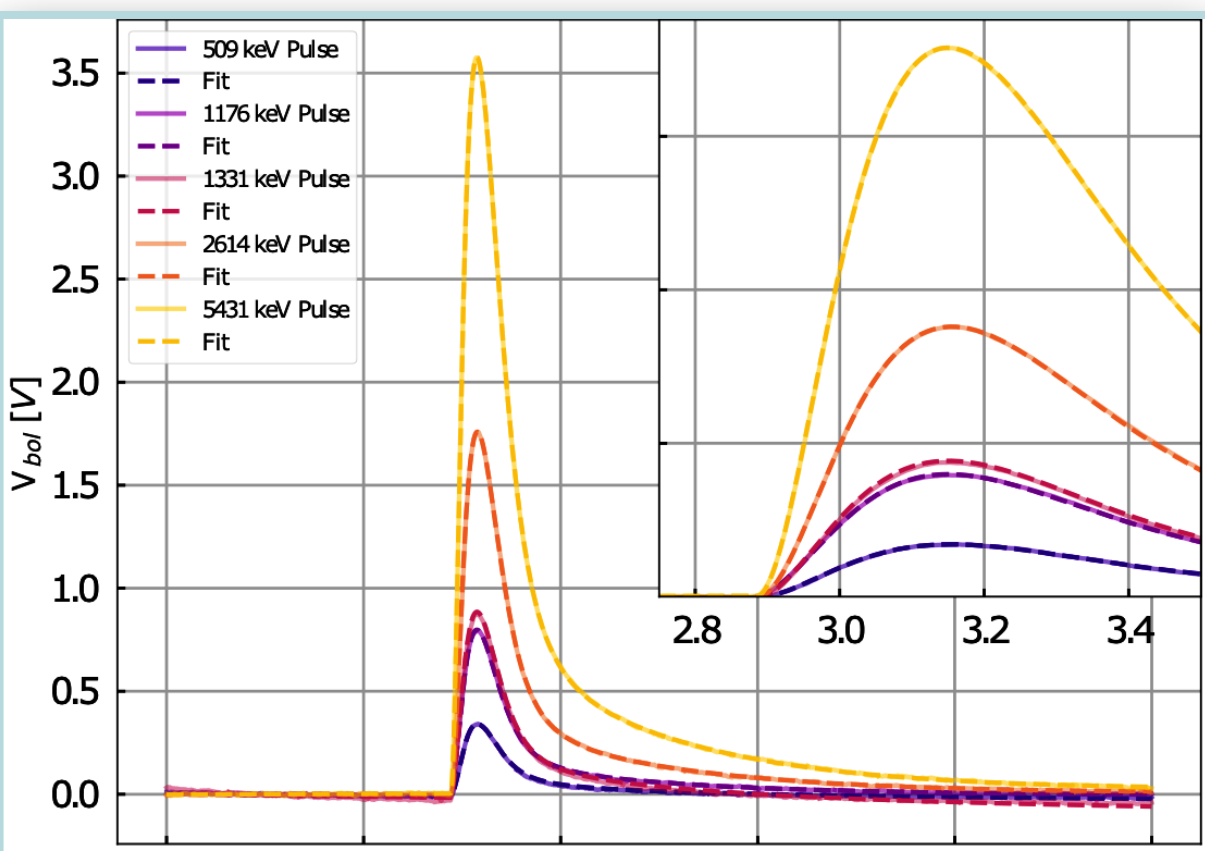
 Phys. Rev. C, 105 065504 (2022)



Not only $^{130}\text{Te } 0\nu\beta\beta$: other physics cases



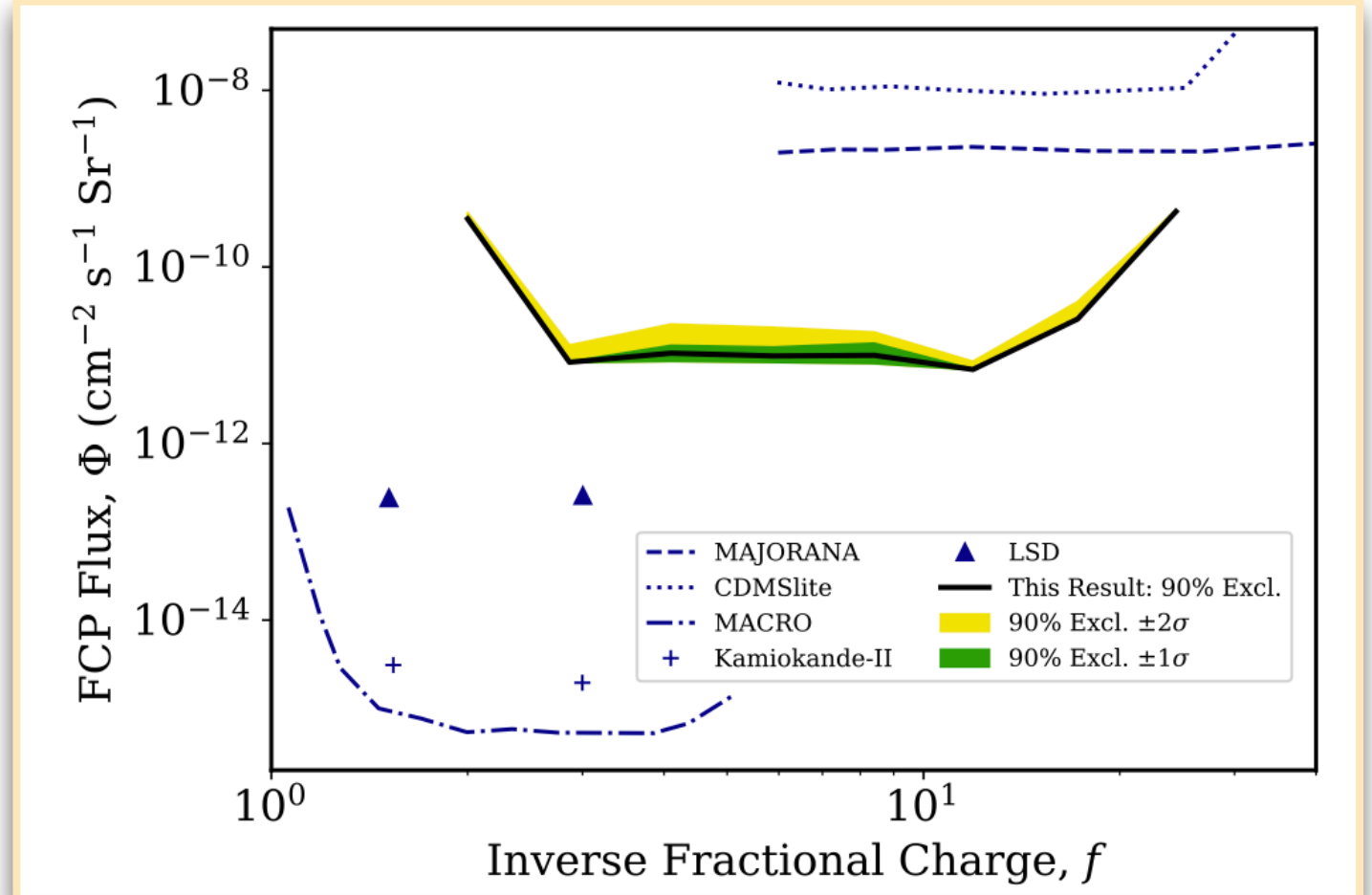
Thermal model of CUORE calorimeters



Dedicated study of environmental and antropic vibrational sources

JINST 17 (2022) 11, P11023

Search for fractionally charged particles

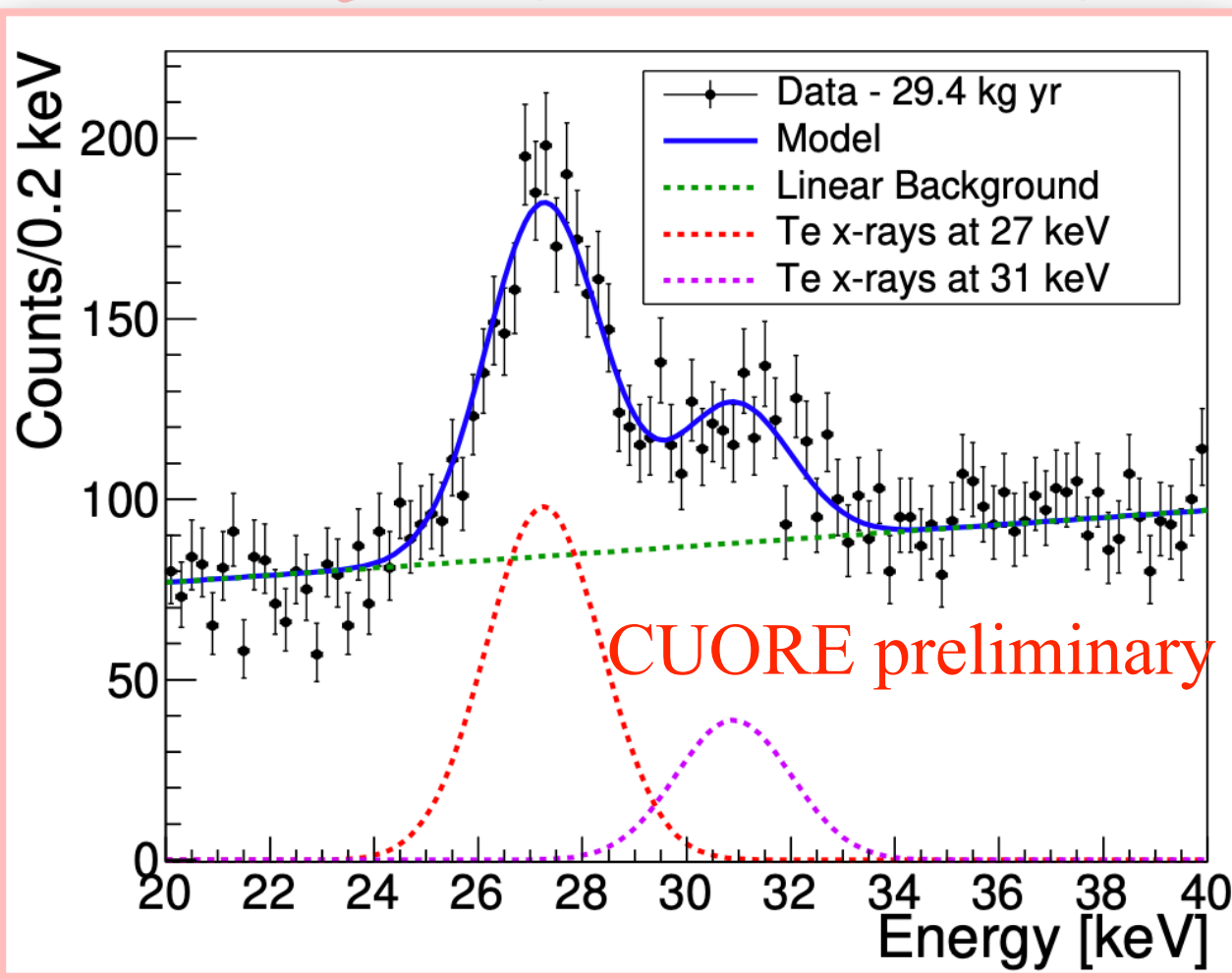


Multi-spectral search for rare events based on multi-crystal track-like topologies



Phys. Rev. Lett. 133, 241801

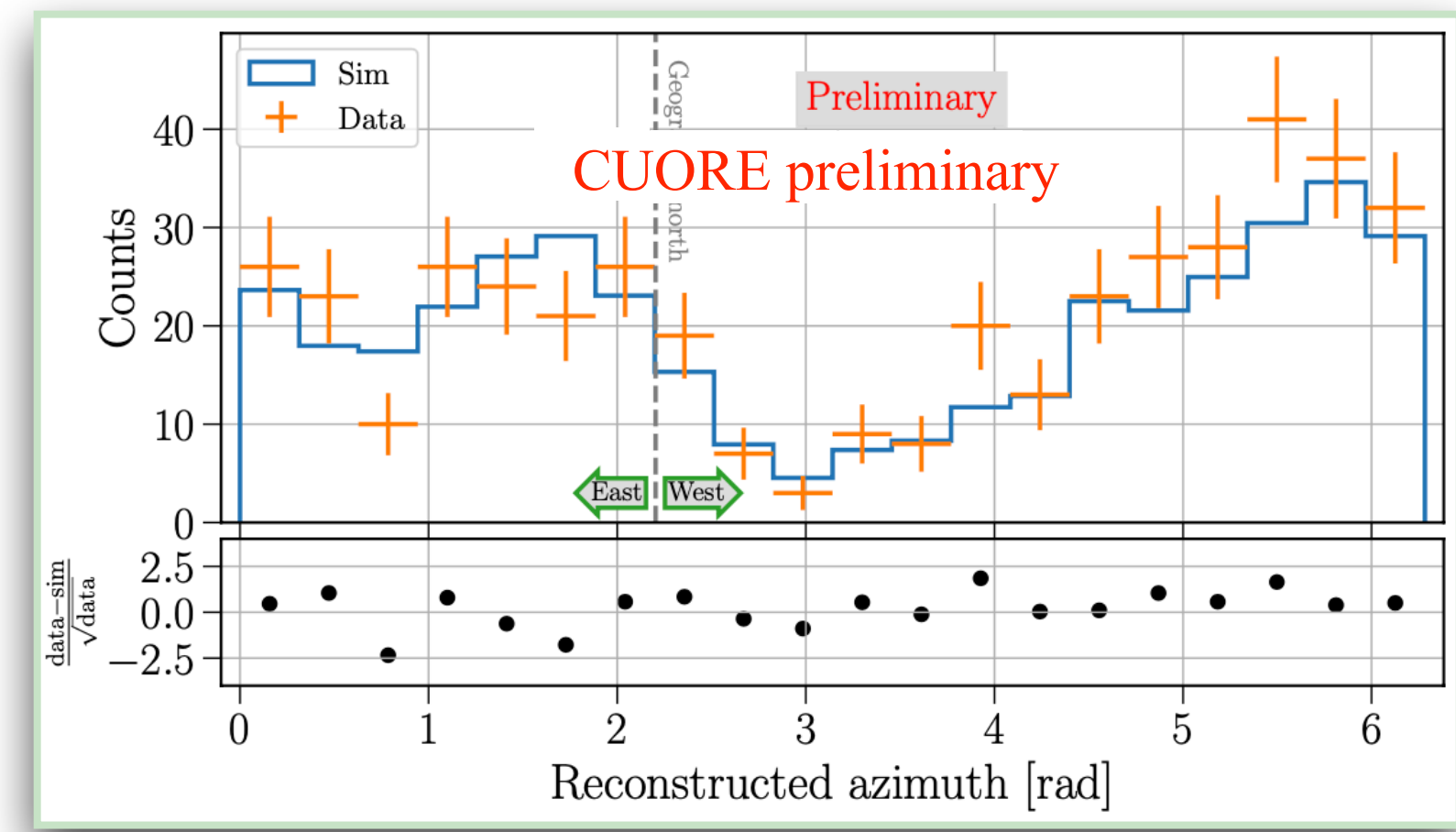
Low energy studies



Specific low-energy variables & event-level cuts to optimise sensitivity for solar axions, WIMP searches

arxiv:2505.23955

Muon event reconstruction (ongoing)



Track-like events ($N_{\text{crystals}} \geq 5$ & $E_{\text{dep}} \geq 9$ MeV) to study μ -induced background

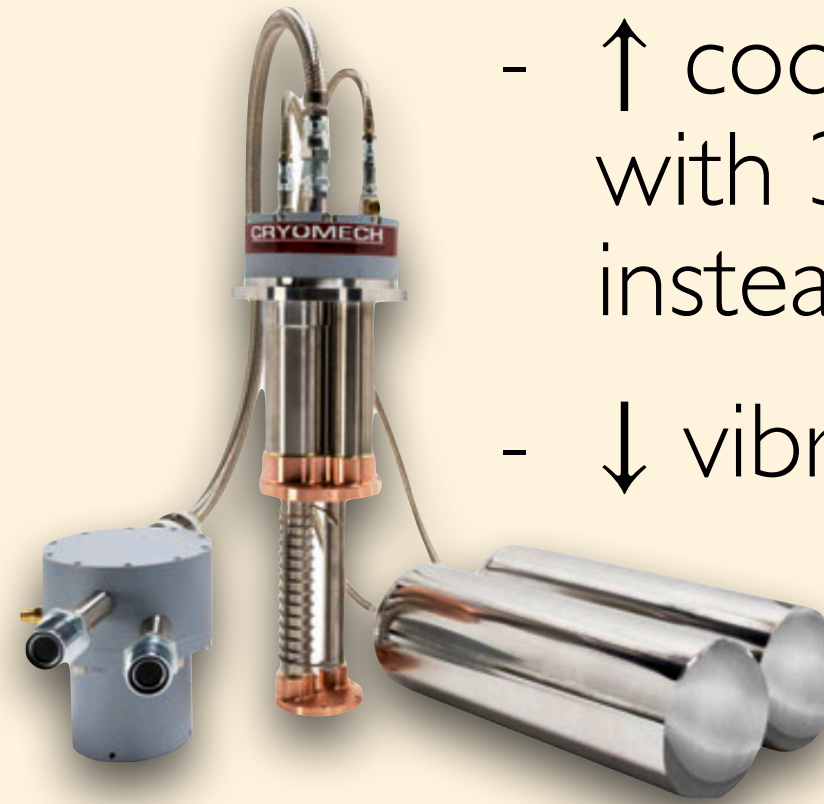
CUORE: what's next? Our path towards CUPID

CUORE (now)

- Continue data taking until meeting our goal:
3 tonne · yr TeO_2 exposure
(~ 1 tonne · yr of ^{130}Te)
- Estimate: end up data taking by mid 2026
- Large statistics to perform high sensitivity searches in several channels ($\beta\beta$ decay, dark matter, exotic phenomena, ...)

CUORE-phase II

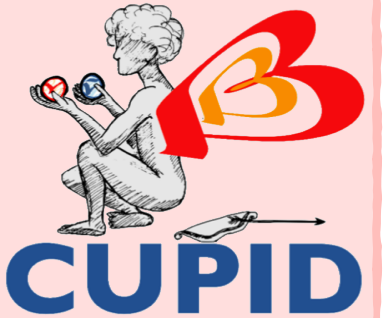
- Upgrade of the **cryogenic system** to improve **Pulse Tubes** performance and coupling to the cryostat



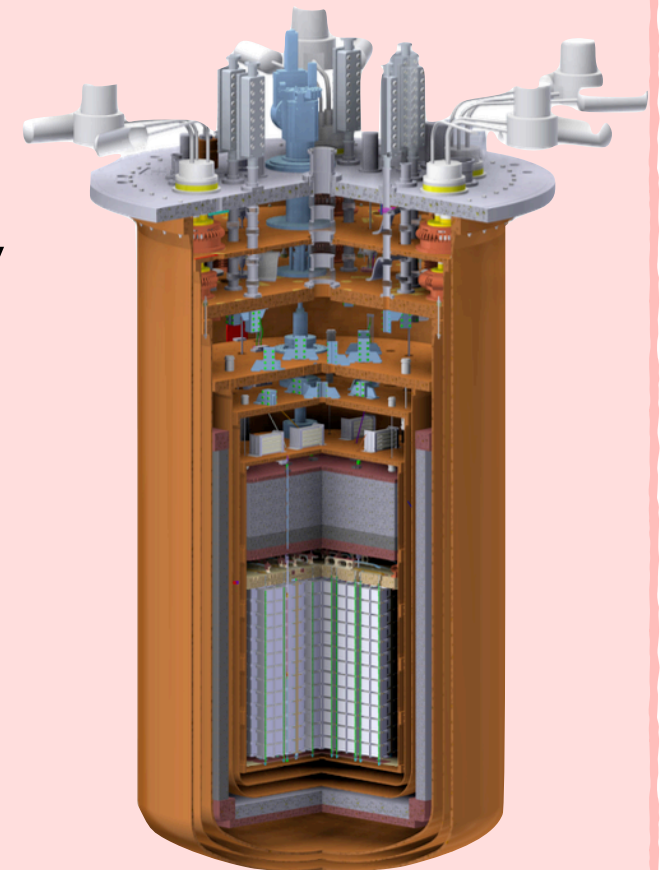
- $\uparrow$ cooling power with 3+1 setting instead of 4+1
- $\downarrow$ vibrations
- Lower thresholds $\rightarrow$ high sensitivity low energy studies (axions, WIMPS, ...)

CUPID

CUORE Upgrade with Particle IDentification



- Scintillating cryogenic calorimeters to overcome CUORE-sensitivity-limiting α background
- $^{130}\text{Te} \rightarrow ^{100}\text{Mo}$
2528 $\rightarrow$ 3034 keV
- 10^{-4} cts/keV/kg/yr target background
- Same cryogenic infrastructure



{See S.Fu talk in this session}

Conclusions and perspectives

- CUORE proved the scalability of the cryogenic calorimeters technique to tonne-scale detectors thereby paving the way to rare decay searches with cryogenic calorimeters
 - ✓ We exceeded 2 tonne · yr TeO_2 analyzed exposure and data collection is progressing towards our goal of a final 3 tonne · yr TeO_2 exposure (corresponding to ~ 1 tonne · yr ^{130}Te)
 - ✓ We found no evidence of $0\nu\beta\beta$ decay with 2039.0 kg · yr TeO_2 exposure and set a new limit on the half life for such decay of $T_{0\nu\beta\beta}^{1/2} > 3.5 \cdot 10^{25}$ yr (90 % C.I.)
 - ✓ Many interesting analyses ongoing also beyond $\beta\beta$ decay: background-related studies (e.g. muon reconstruction), multispectral (search for decays in multiple-crystal events) and low energy studies
 - ✓ Important feedback for CUPID, both for the cryogenics and background budget
 - ✓ After interventions on the cryogenics and before the CUPID detector installation, a CUORE phase II dedicated to low energy studies (dark matter searches, e.g. WIMPs, axions, ...) is planned