

In situ Time Calibration of KM3NeT ARCA

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on behalf of the KM3NeT Collaboration

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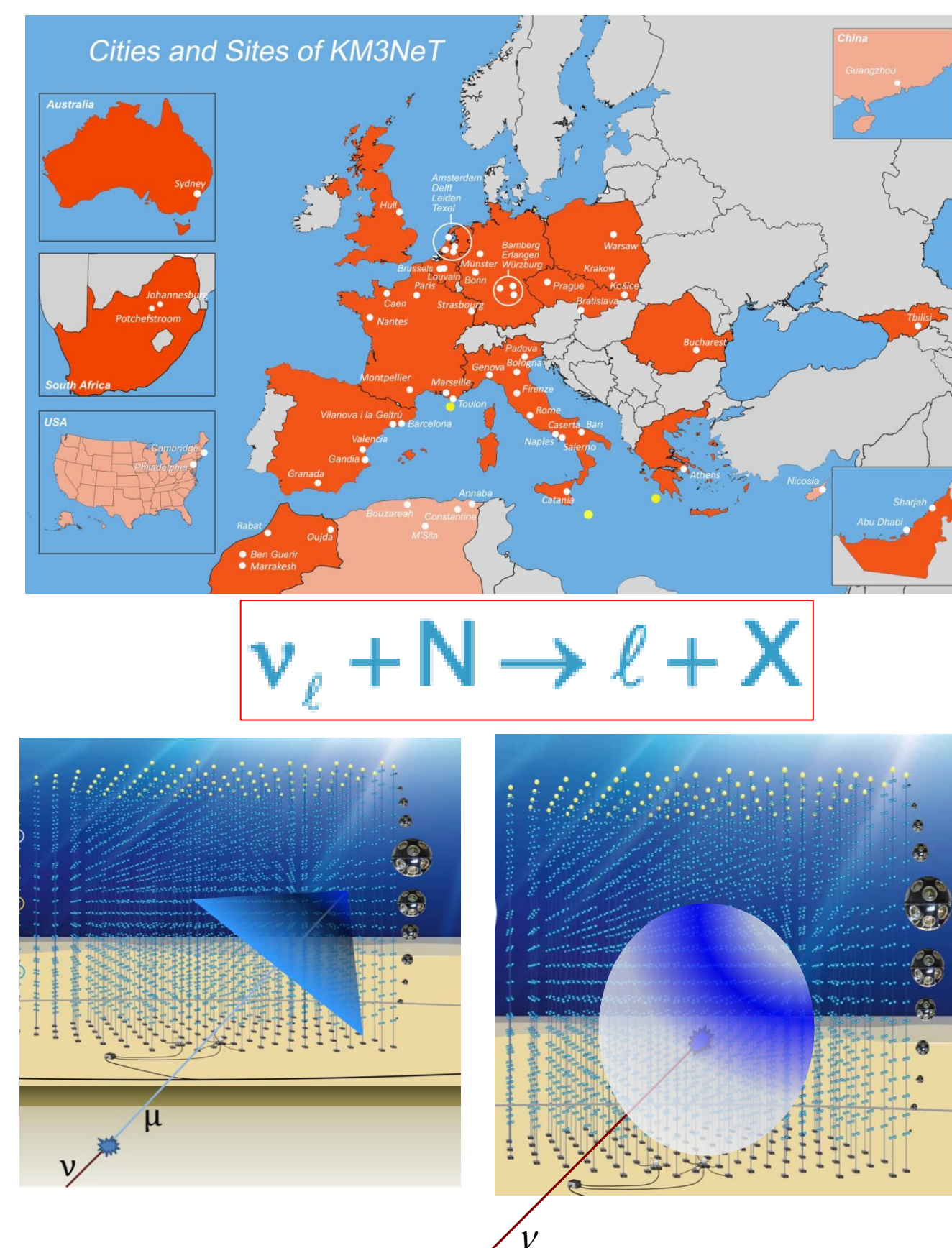


1 KM3NeT: Cubic Kilometre Neutrino Telescope

Large-scale underwater neutrino telescope designed to detect high-energy neutrinos (> 100 GeV) by observing the Cherenkov light emitted.

Two detectors:

- ARCA (Astroparticle Research with Cosmics in the Abyss), for high-energy astrophysical neutrinos.
- ORCA (Oscillation Research with Cosmics in the Abyss), for atmospheric neutrinos and determining the neutrino mass hierarchy.



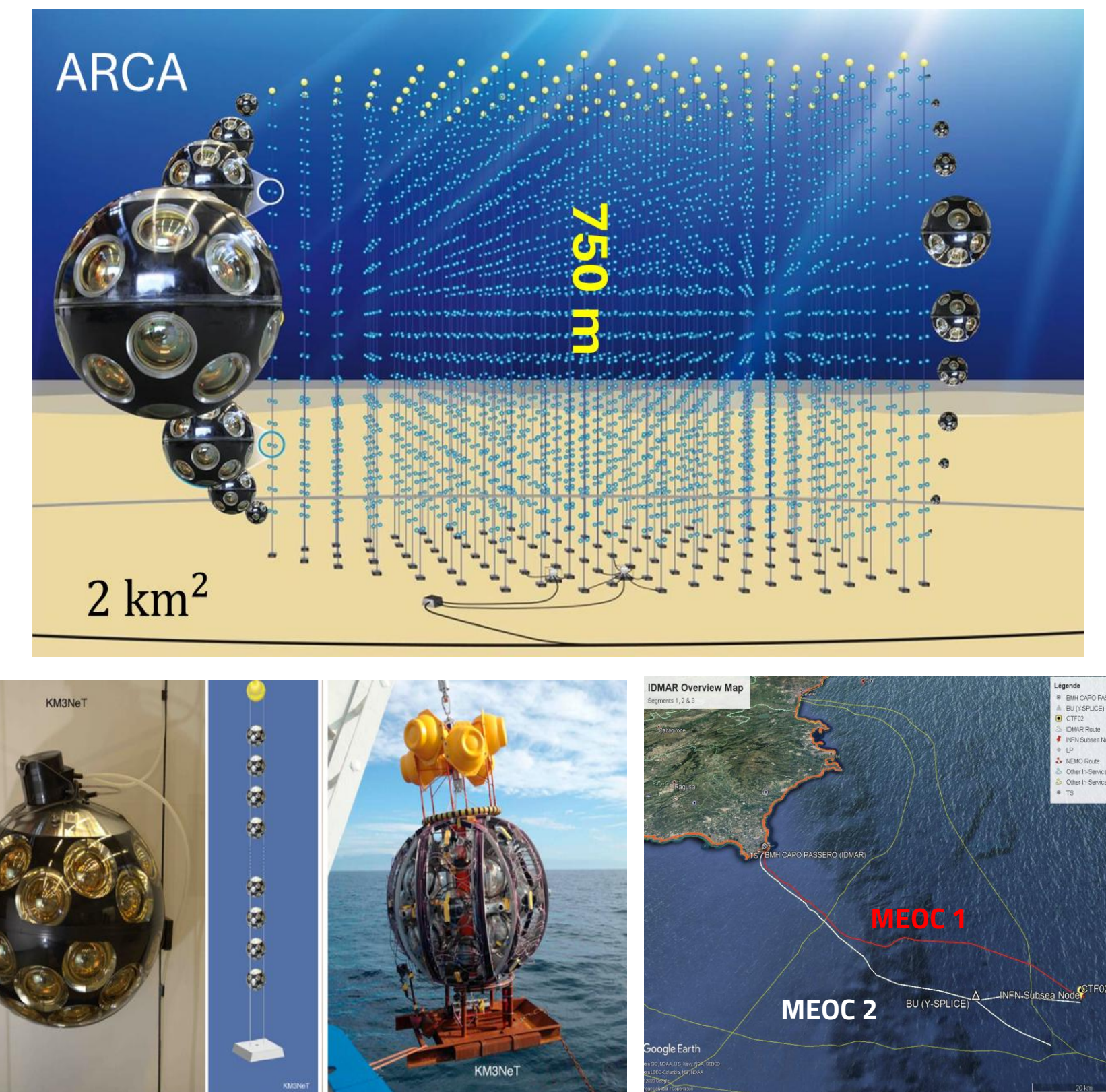
2 The KM3NeT ARCA detector

Located about 100 km off the coast of Sicily, at a depth of 3500 m.

Array of Digital Optical Modules (DOMs), anchored on the seabed, grouped into Detection Units (DUs).

Power distribution and data communication by Junction Boxes (JBs) connecting the system to the shore via two main electro-optical cables (MEOC 1 and 2).

Since October 2024 total of **33 strings** currently anchored on the seabed.



3 Calibration of ARCA detector

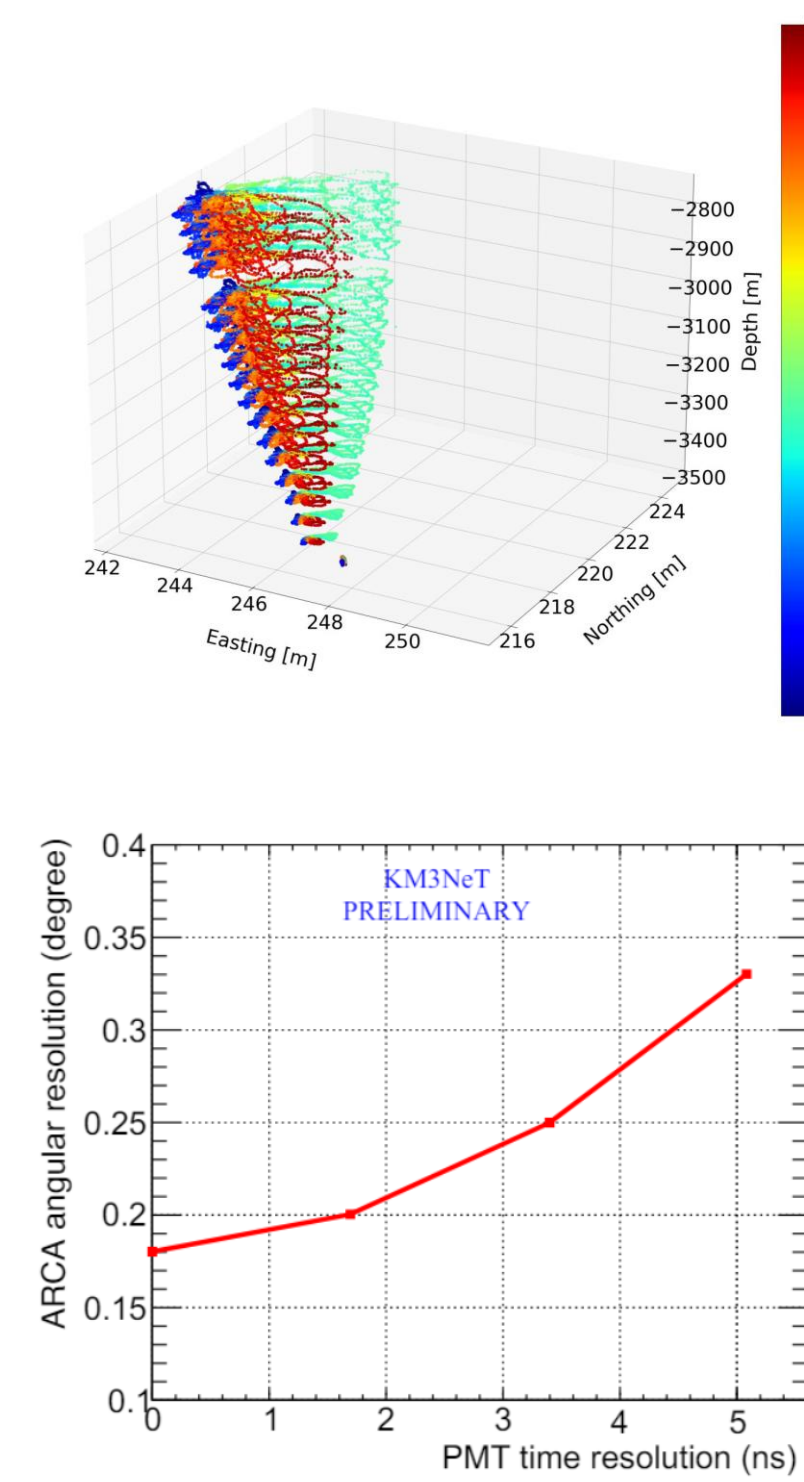
Angular resolution (< 0.1°) requirement for accurate neutrino event reconstruction and background suppression.

- Accurate spatial calibration through a relative **Acoustic Positioning System**.
- Accurate time calibration through a **nanosecond-level synchronization** across the detector.

Determination of the relative time offsets:

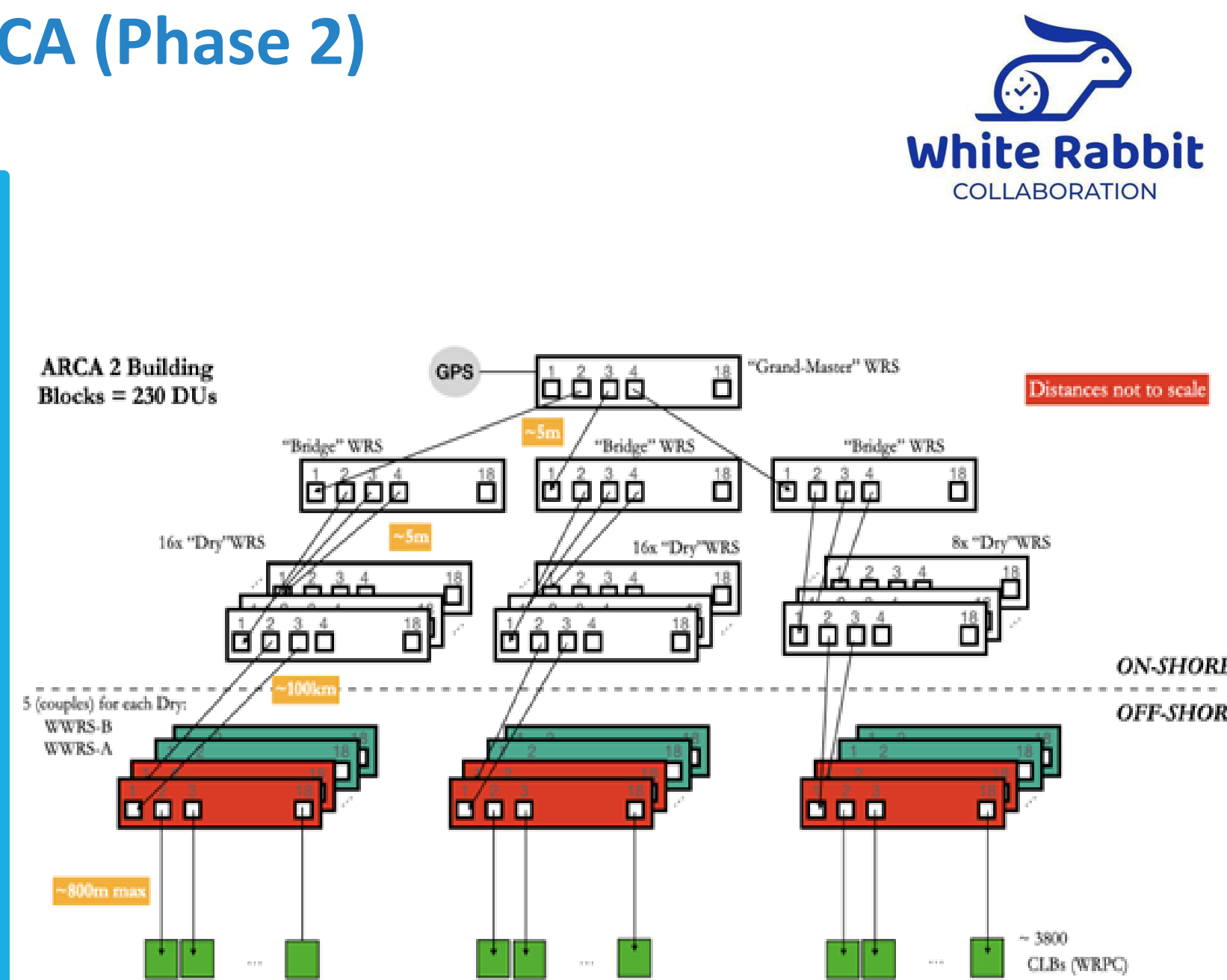
- between PMTs within the same DOM (**intra-DOM**),
- between DOMs in the same DU (**inter-DOM**),
- between different DUs (**inter-DU**).

In situ checks through two independent methods: atmospheric muons and laser beacons.



4 The novel White Rabbit time calibration system for KM3NeT-ARCA (Phase 2)

Sub-nanosecond synchronization through the WR protocol over **point-to-point** optical links, using round-trip-time measurements to a common reference. Shore master clock distributes timing through WR switch hierarchy to each DOM via CLB and WWS.

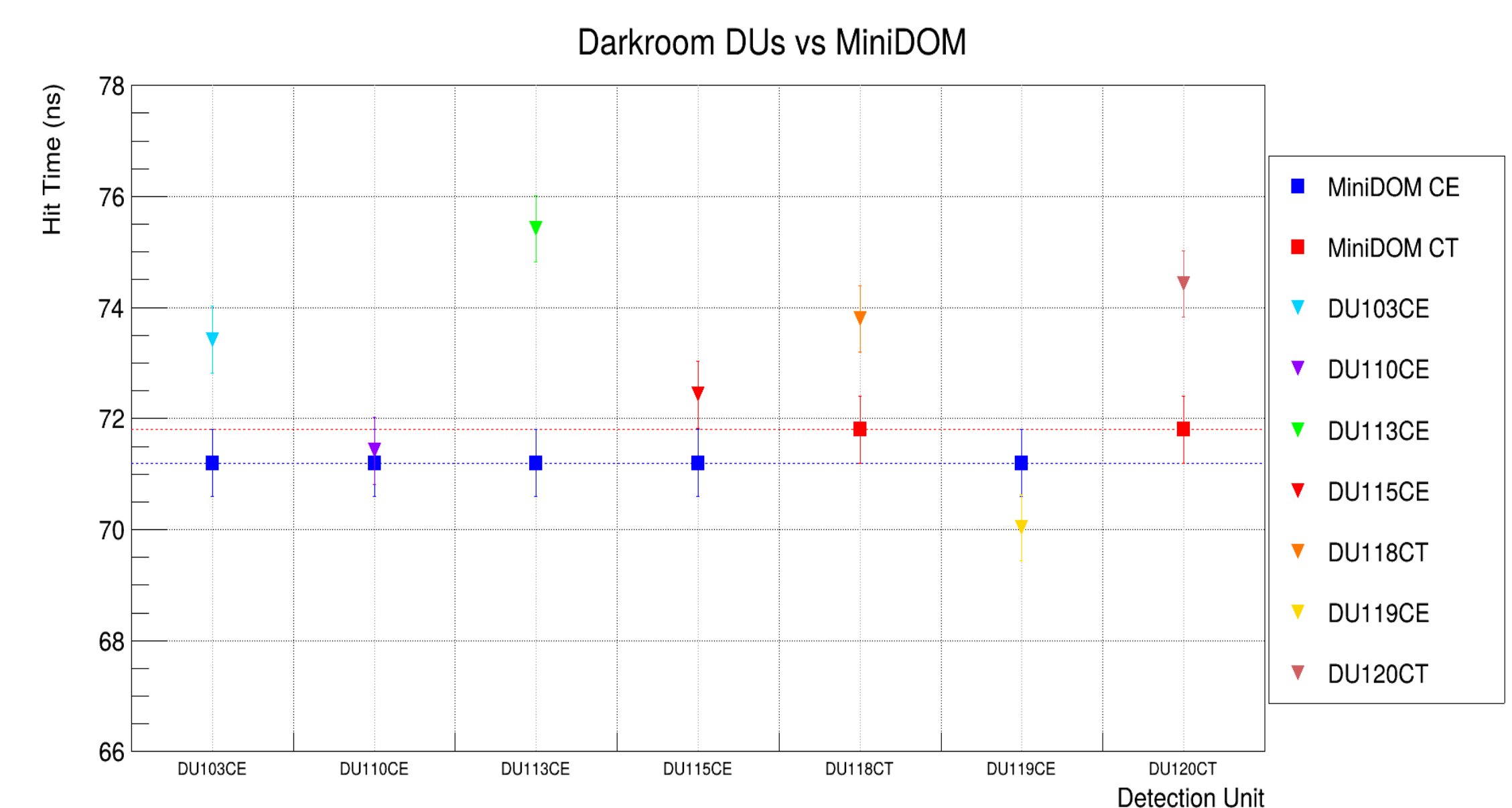


5 The Darkroom Calibration

Prerequisite for time calibration: High-Voltage (HV) tuning, setting the PMT gain to detect photons equal to **3x10⁶** and to ensure a uniform response among PMTs. DOM time calibration performed with a unique laser source illuminating DOMs on one DU. Laser runs then measure inter-DOM timing using two reference PMTs per DOM, correcting the average hit times for known delays.

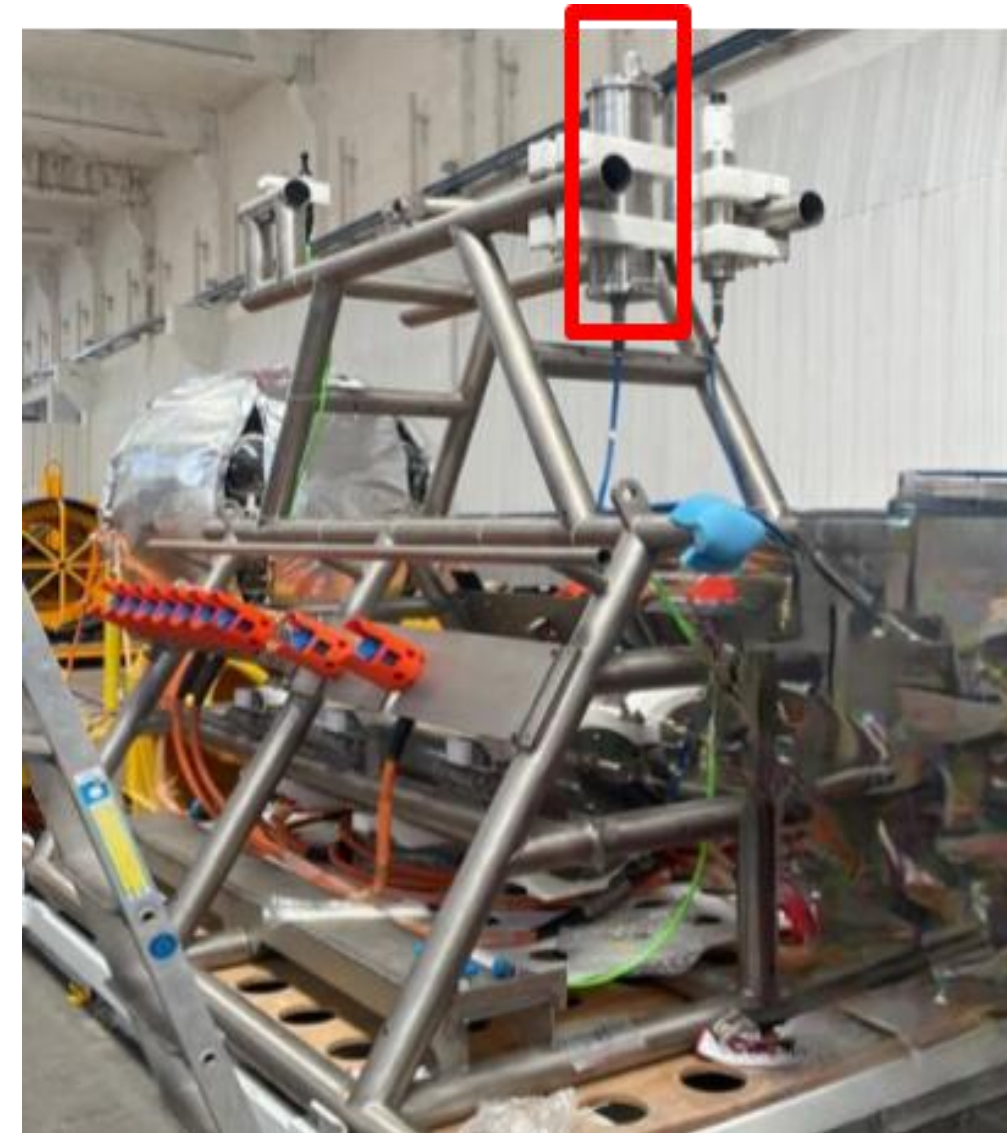
DU time calibration performed at three different darkrooms: Catania, Caserta and Genova with systematic differences due to the **different darkroom setups**.

Intercalibration of the three darkrooms performed with a portable DOM (the MiniDOM).



6 The Laser Beacon

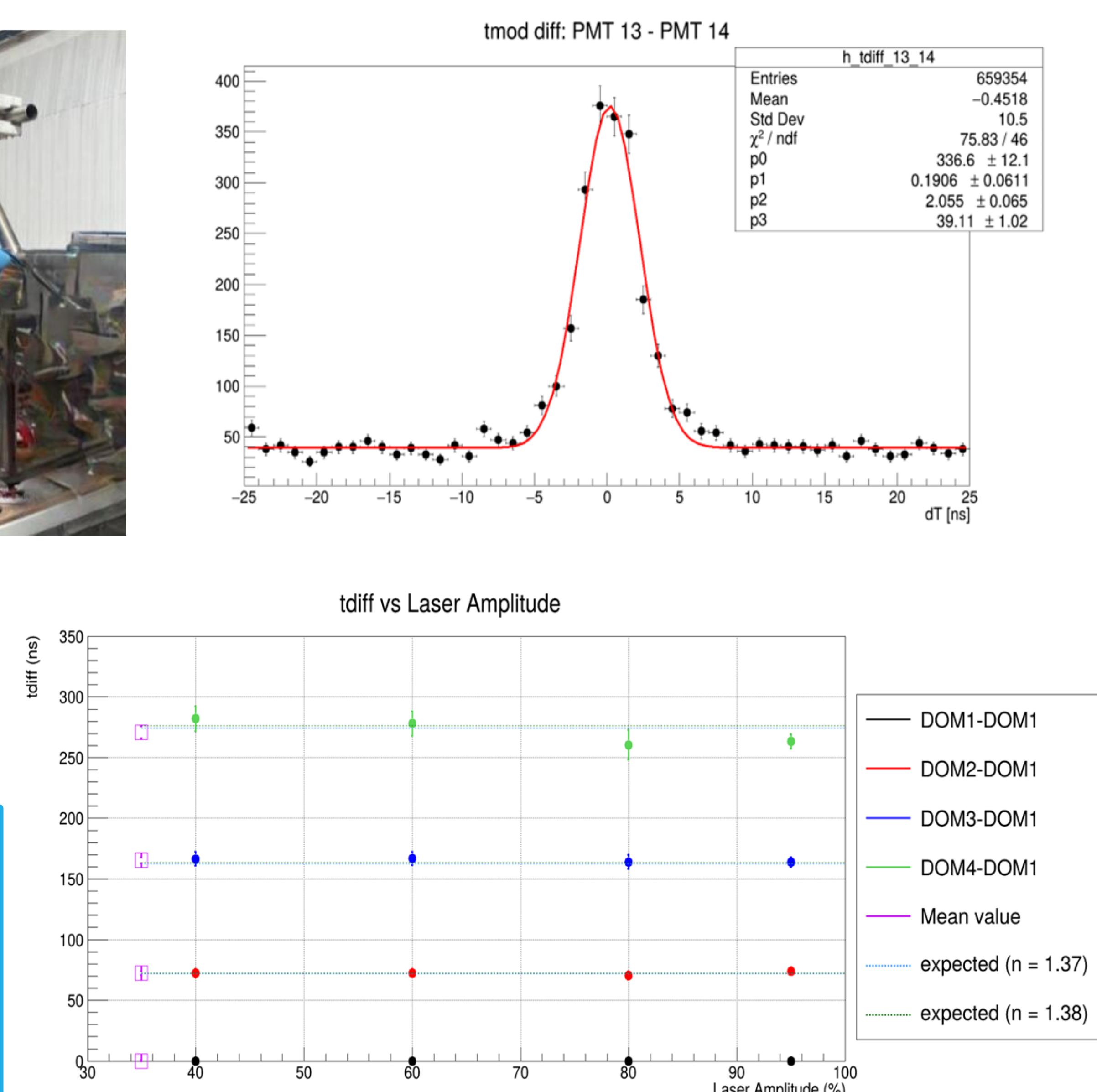
Auxiliary device, installed on the seafloor onboard ARCA-JBs, designed to perform **in situ time calibration** of the detector. Emits short, intense and isotropic green laser pulses (532 nm) with **sub nanosecond light width**.



Time differences **between PMT pairs** within the same DOM, finding **mean values close to 0**, in agreement with the one determined by intra-DOM time calibration and **sigma values** in a range around **2.2-2.3 ns**, consistent with the PMT Transit Time Spread (TTS).

Time differences calculated by the time of flight measurements, considering **the relative distances between each DOM and the LB** and varying laser amplitudes.

Using an index of refraction of 1.38 for **light group velocity**, the agreement between data and expectations is below 1.5 ns.



7 Conclusions

- White Rabbit Standard successfully implemented, improving time synchronization across the detector.
- A new device, called MiniDOM, developed and used to cross-calibrate the darkrooms laser systems.
- First in situ application of the Laser Beacon for: intra-DOM calibration and inter-DOM calibration. Results consistent with ⁴⁰K decays and muon track residuals.

References:

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7. S. Adrián-Martínez et al., Measurement of the group velocity of light in sea water at the ANTARES site, *Astroparticle Physics*, 2012.