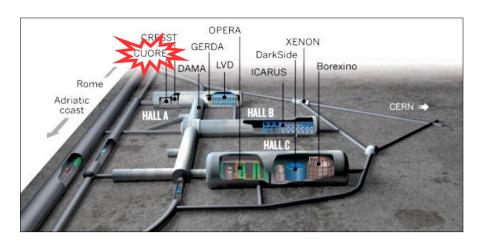


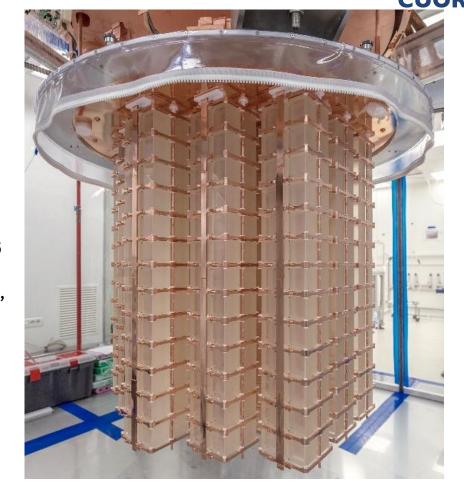
Cryogenic Underground Observatory for Rare

Events (CUORE)

- Array of 988 TeO₂ crystals (742 kg, 206 kg ¹³⁰Te)
 - Decay source = detector
 - Unenriched, large isotopic abundance (~1/3)
- Hall A of Laboratori Nazionali del Gran Sasso
 - Rock overburden reduces muon flux by factor of 10⁶ relative to sea level (BOREXINO, J. Cosmol. Astropart. Phys., 05:015, 2012)







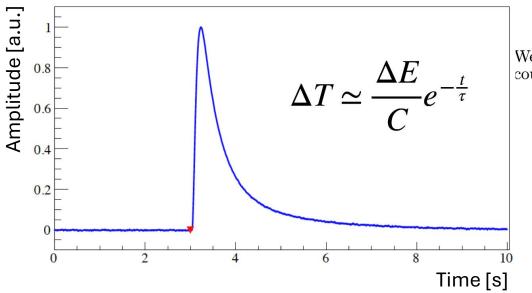


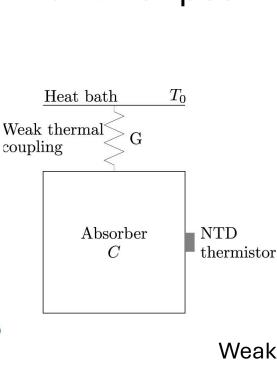
TeO₂ cryogenic calorimeters

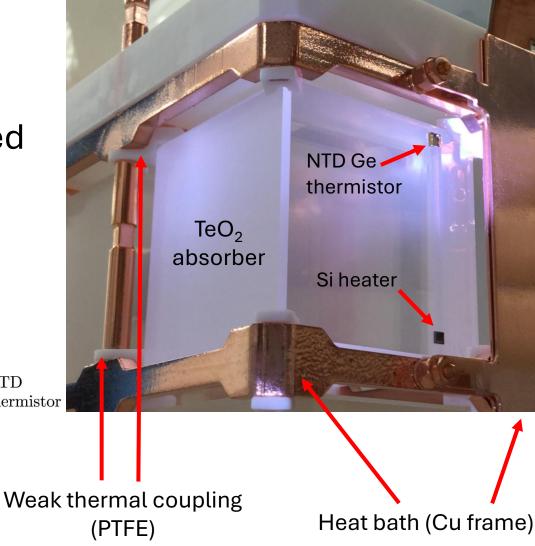


- Operated at O(10 mK)
- Energy depositions

 temperature change
- Read out using neutron transmutation doped (NTD) Ge thermistors





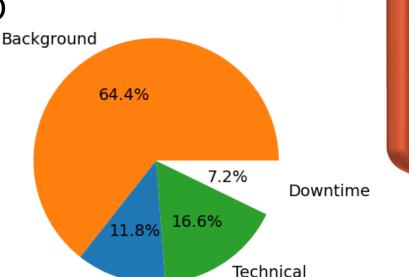




The CUORE cryostat

CUORE

- Large custom-built dilution refrigerator
 - Cool ~3 tonnes to 50mK
 - 5 pulse tubes, 6 thermal stages
- Over 5 years of stable data-taking, uptime >90%
- Data-taking planned through 2026 before upgrades for CUPID
 - Exposure goal: 3 t-yr TeO₂



Calibration

Plates:

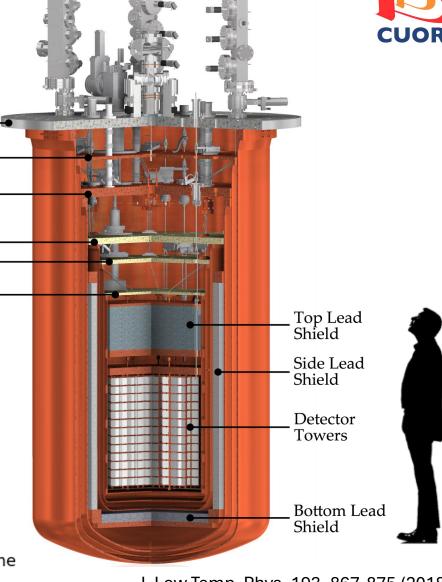
300 K

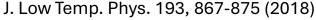
40 K-

4 K-

600 mK - 50 mK -

10 mK-



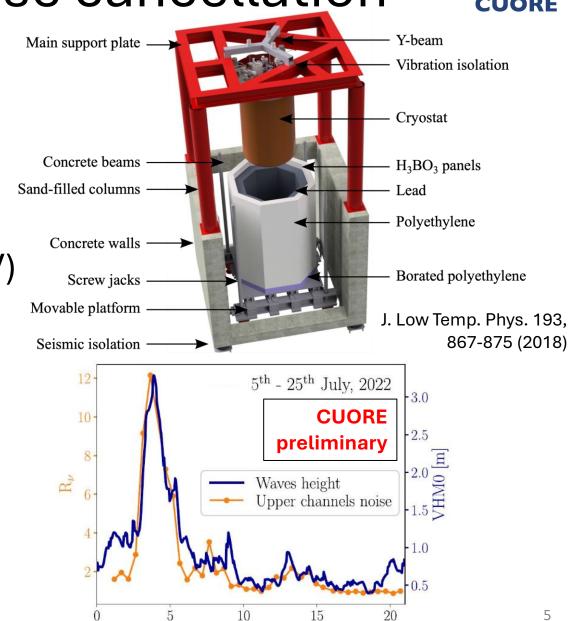




Vibrational isolation and noise cancellation

- Multiple stages of vibrational isolation
- Offline denoising
 - ~40% reduction in total RMS_{noise}
 - Eur. Phys. J. C 84, 243 (2024)
- Good resolution: 0.3% $\delta E/E$ at Q_{BB} (2.5MeV)
 - arXiv:2404.04453

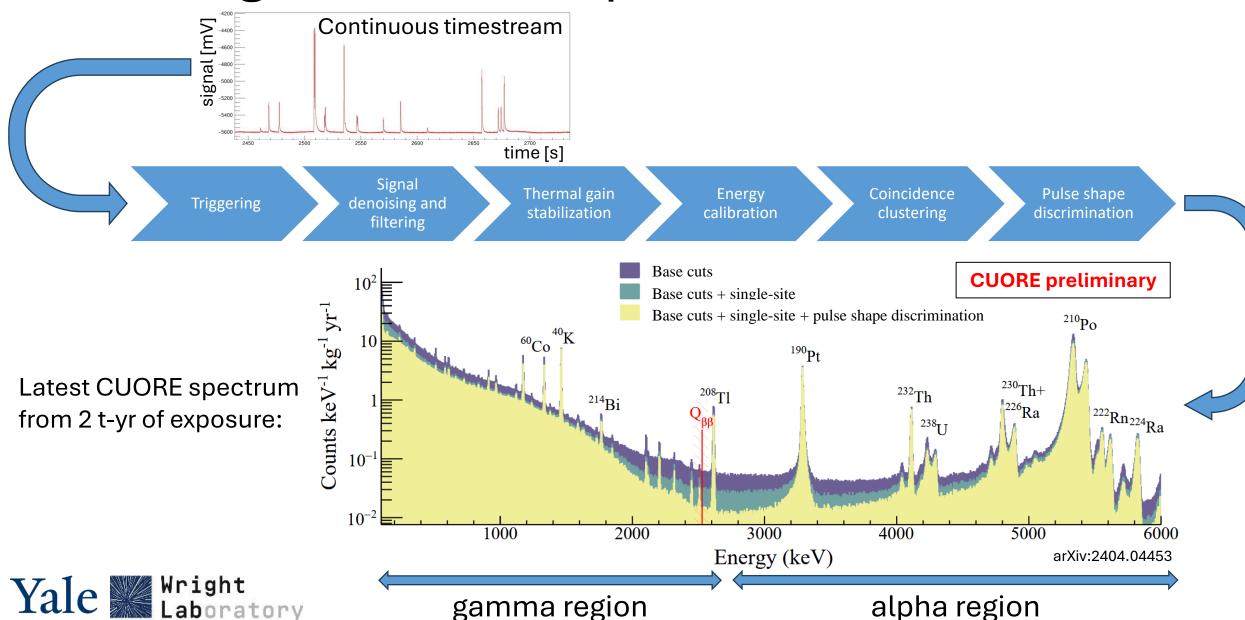
- Sensitive enough to see sea storm activity!
 - Eur. Phys. J. C 84, 728 (2024)





Producing the CUORE spectrum





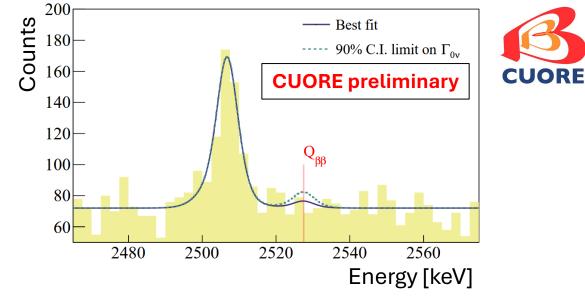
¹³⁰Te 0vββ search (latest!)

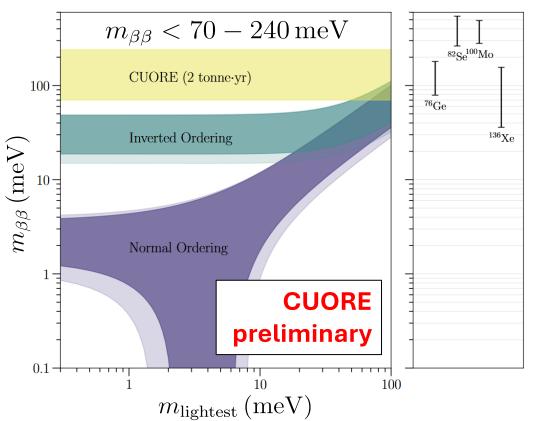
- Bayesian unbinned fit of ROI
 - Flat background
 - ⁶⁰Co peak
 - 0vββ peak at Q_{ββ}
- arXiv:2404.04453
- Resolution @ 2615 keV: ~7.5 keV
- Analysis efficiency: 93.4(18)%

No evidence of 0vββ

$$T_{1/2}^{0\nu} > 3.8 \times 10^{25} \text{yr} (90\% \text{ C.I.})$$



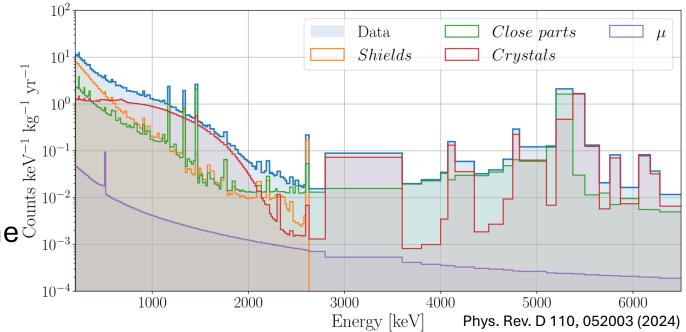




Backgrounds are well-understood



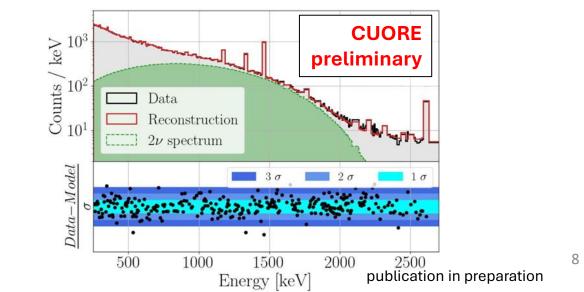
- Good reconstruction of spectral features in [200, 7000] keV
 - Verified by data and simulation
 - Work ongoing for below 200 keV
 - Important for CUPID which will use the same cryostat



 Most precise measurement of 2vββ in ¹³⁰Te to date (in preparation)

$$T_{1/2}(^{130}\text{Te}) = 9.323^{+0.052}_{-0.037} \text{ (stat)} \times 10^{20} \text{ y}$$





Beyond 0vββ: new analysis techniques



 Leverage good energy collection at keV-scale, detector segmentation, large exposure:

keV-scale "low energy" physics

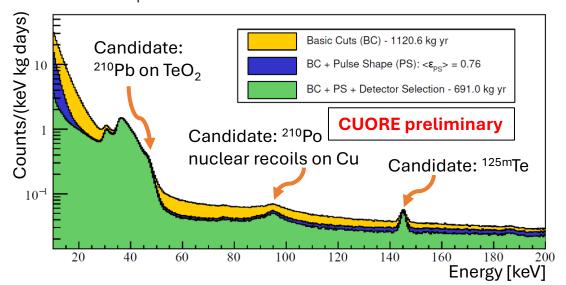
- 1. Analysis techniques (paper in preparation)
- 2. Solar axion searches (ongoing)
- 3. WIMP searches (ongoing)

Coincident multi-site physics

- 1. Decay to excited state of ¹³⁰Xe (Eur. Phys. J. C 81, 567 (2021))
- 2. Neutrinoless electron capture on ¹²⁰Te (Phys. Rev. C 105, 065504 (2022))
- 3. Fractionally-charged particles (Phys. Rev. Lett. 133, 241801 (2024))
- 4. Muon flux measurement (paper in preparation)

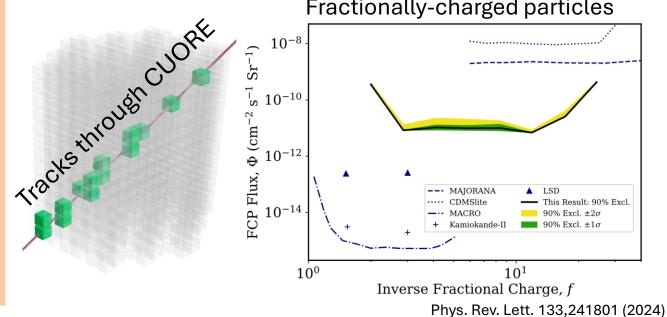


...and more!



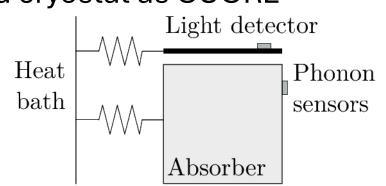
M1 spectrum of detector with OT threshold lower than 10 keV

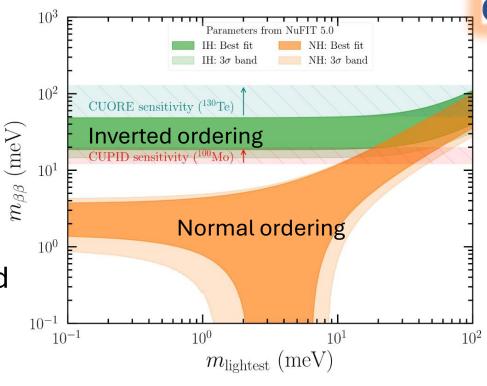


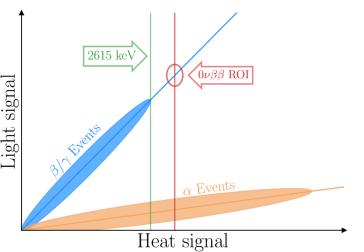


CUORE Upgrade with Particle IDentification

- Goal: fully explore inverted ordering parameter space
 - Reduce backgrounds in ROI by 100x
- 1596 scintillating crystals of $\text{Li}_2^{100}\text{MoO}_4$ 1710 Ge light detectors
 - 100 Mo higher Q-value \rightarrow less β/γ background
 - Less scintillation light for α 's
- Benefit from CUORE expertise
 - Same facilities and cryostat as CUORE









CUORE upgrades and CUPID R&D

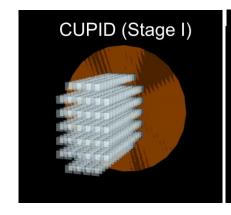
- CUORE data-taking through 2026
 - Plans for dedicated "low-energy" runs

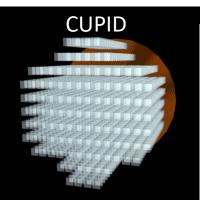
Cryogenic upgrades planned for 2026

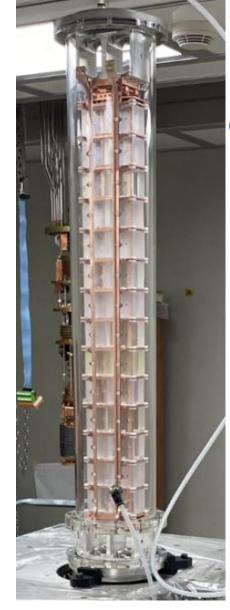
arXiv:2503.02894

arXiv:2503.04481

- CUPID full baseline tower R&D test underway
- Staged deployment for CUPID: physics data by 2030
 - Stage I: 532 crystals (1/3 of total array)
 - Stage II: full array







Test tower assembly

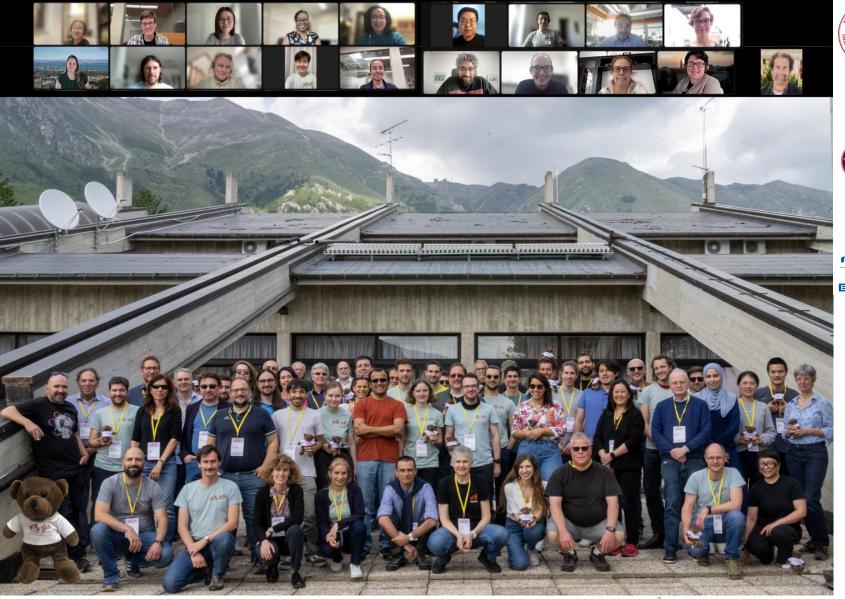


Summary and outlook

CUPID

- CUORE has taken data stably since 2019
- World-leading limit on 130 Te $0v\beta\beta$ half-life: $>3.8\times10^{25}$ yr
- Wide variety of physics possible to explore
- CUPID will improve on CUORE with background rejection using scintillation light detection
- CUPID benefits from CUORE experience:
 - Background model, operating procedures, etc.
- Upgrades expected in 2026 stay tuned!









































Institute of

























































Centro de Astropartículas y Física de Altas Energías

Universidad Zaragoza















Thank you!

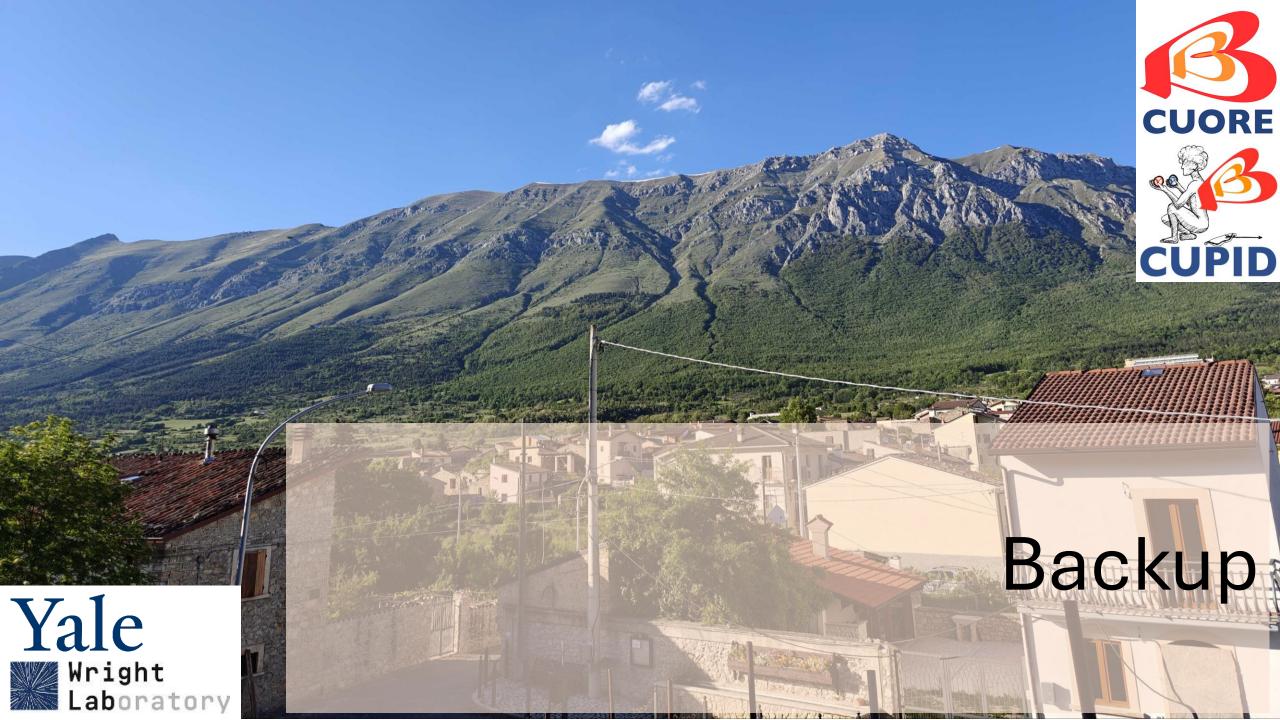
• The CUORE Collaboration thanks the directors and staff of the Laboratori Nazionali del Gran Sasso and the technical staff of our laboratories. This work was supported by the Istituto Nazionale di Fisica Nucleare (INFN); the National Science Foundation under Grant Nos. NSF-PHY-0605119, NSF-PHY-0500337, NSF-PHY-0855314, NSF-PHY-0902171, NSF-PHY-0969852, NSF-PHY-1307204, NSF-PHY-1314881, NSF-PHY-1401832, NSF-PHY-1913374, and NSF-PHY-2412377; Yale University, Johns Hopkins University, and University of Pittsburgh. This material is also based upon work supported by the US Department of Energy (DOE) Office of Science under Contract Nos. DE-AC02-05CH11231, and DE-AC52-07NA27344; by the DOE Office of Science, Office of Nuclear Physics under Contract Nos. DE-FG02-08ER41551, DE-FG03-00ER41138, DE-SC0012654, DE-SC0020423, DE-SC0019316, and DE-SC0011091. This research used resources of the National Energy Research Scientific Computing Center (NERSC). This work makes use of both the DIANA data analysis and APOLLO data

acquisition software packages, which were developed by the CUORICINO, CUORE, LUCIFER, and CUPID-0 Collaborations.

• The authors acknowledge Advanced Research Computing at Virginia Tech for providing computational resources and technical support that have contributed to the results reported within this paper. URL: https://arc.vt.edu/.The CUPID Collaboration thanks the directors and staff of the Laboratori Nazionali del Gran Sasso and the technical staff of our laboratories. This work was supported by the Istituto Nazionale di Fisica Nucleare, the Italian Ministry of University and Research (Italy), the European Research Council and European Commission, the US Department of Energy (DOE) Office of Science, the DOE Office of Science, Office of Nuclear Physics, the National Science Foundation (USA), the National Research Foundation (Ukraine), and the Russian Science Foundation (Russia). This research used resources of the National Energy Research Scientific Computing Center (NERSC). This work makes use of both the DIANA data analysis and APOLLO data acquisition software packages, which were developed by the CUORICINO, CUORE, LUCIFER and CUPID-0 Collaborations.



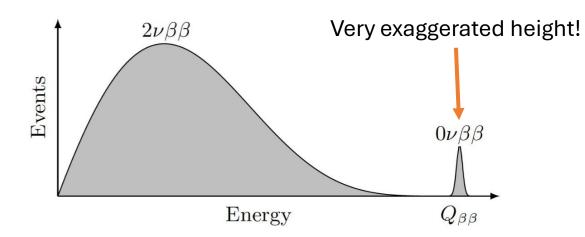


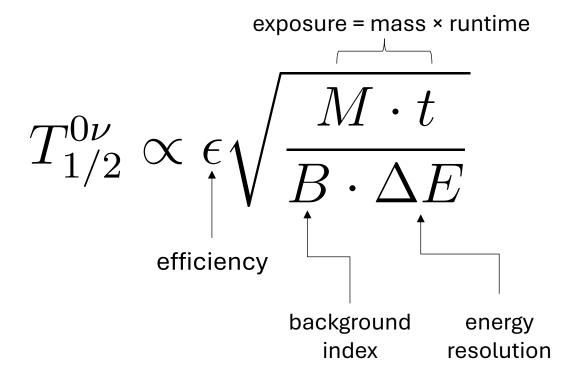


Double beta decay experiment design



- Peak in energy spectrum at tail of twoneutrino double-beta decay (2vββ)
- Half-life limits $> 10^{22} 10^{26} \text{ yr}$
- Sensitivity of experiment depends on:
 - Number of signal events you can see
 - Amount of active isotope
 - Runtime
 - Efficiency
 - Amount of background in ROI
 - Background index
 - Energy resolution (~width of ROI)





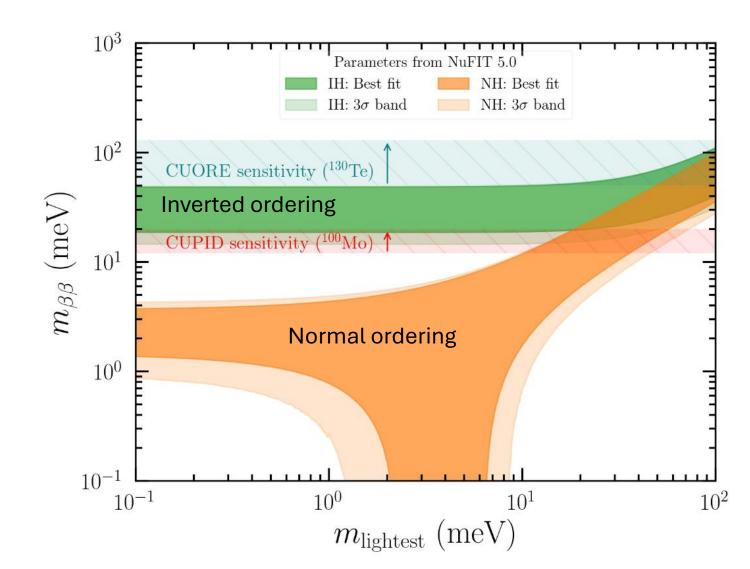


Effective neutrino mass



$$m_{\beta\beta} = \left| \sum_{i=1}^{3} |U_{ei}|^2 e^{i\phi_i} m_i \right|$$

$$T_{1/2}^{0\nu} = \left[G_{0\nu} |\mathcal{M}|^2 \frac{m_{\beta\beta}^2}{m_e^2} \right]^{-1}$$





¹³⁰Te 0vββ search (latest!)

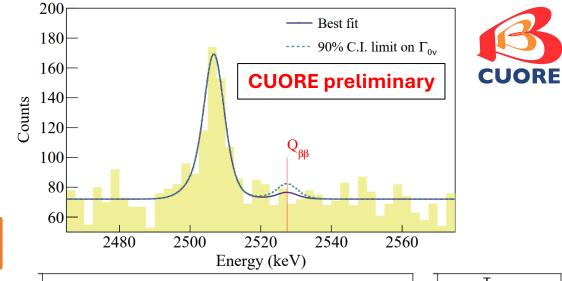
- Bayesian unbinned fit of ROI
 - Flat background
 - 60Co peak
 - $0v\beta\beta$ peak at $Q_{\beta\beta}$

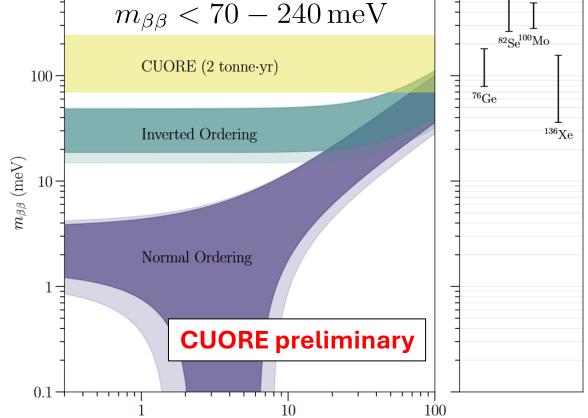
arXiv:2404.04453

Number of datasets	28
TeO ₂ exposure	2039.0 kg⋅yr
¹³⁰ Te exposure	567.0 kg·yr
FWHM at 2615 keV in calibration data	7.540(24) keV
FWHM at $Q_{\beta\beta}$ in physics data	7.320(24) keV
Total analysis efficiency (data)	93.4(18)%
Reconstruction efficiency	95.624(16)%
Anti-coincidence efficiency	99.80(5)%
PSD efficiency	97.9(18)%
Containment efficiency (MC)	88.35(9)% [34]

$$T_{1/2}^{0\nu} > 3.8 \times 10^{25} \text{yr} (90\% \text{ C.I.})$$







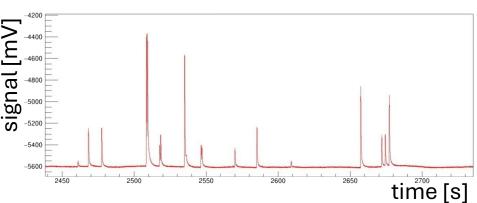
 $m_{\text{lightest}} \text{ (meV)}$

Analysis framework









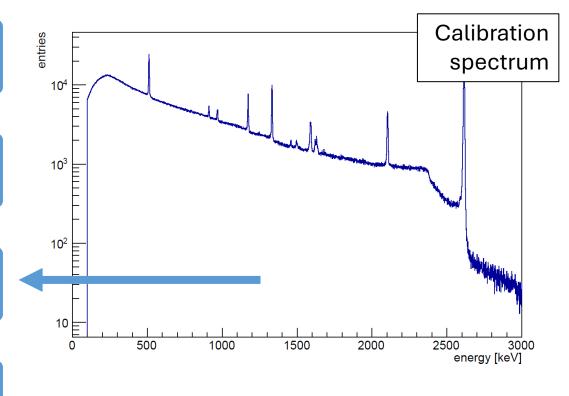
Signal denoising and filtering



Energy calibration

Coincidence clustering

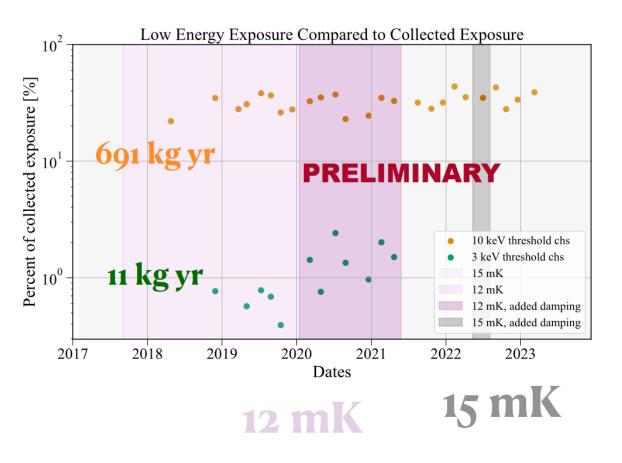
Pulse shape discrimination





keV-scale "low energy" physics runs

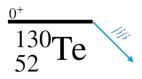




Sensitive to detector operating conditions

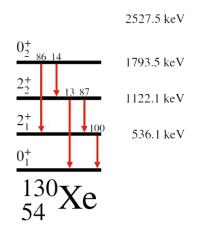


CUORE as a segmented detector



Decay to excited final state Eur. Phys. J. C 81, 567 (2021)

Updated analysis ongoing!



Fractionally-charged particles

