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Aula Conversi

ISTITUTO NAZIONALE DI FISICA NUCLEARE SEZIONE DI ROMA

CORSO DI FORMAZIONE INFN







A PRELUDE - Experimental Surprises?: MAGIC (2005) et al. & OPERA (2011) Playing with Light: Quantum Field Theories in non-trivial Vacua & Light (superluminal) Refraction Causality & Superluminality – not necessarily incompatible

ACT I QUANTUM GRAVITY (QG) as a non-trivial Vacuum (``Medium'') CONCEPTS & MODELS

#### PHENOMENOLOGY OF QG FOAM VACUA (``FOAM-OLOGY''):

Some aspects Non-trivial Optical properties (refractive index, birefringence(?))

ACT II A STRINGY MODEL OF SPACE-TIME D-(efect) FOAM

#### **ENTR'ACTE**

ACT III QUANTUM GRAVITY, DECOHERENCE & ENTANGLED PARTICLE STATES: QG Decoherence & CPT Violation

#### ACT IV PHENOMENOLOGY OF CPT VIOLATION

- (i) EPR Modifications the ω-effect-Order of Magnitute Estimates/Model dependence
- (ii) Kaon vs B-systems and the  $\omega$ -effect

**EPILOGUE:** Reconciling MAGIC et al. with OPERA & DAΦNE-2: **D-FOAM**, Further QG Tests? 10-13-1

# A PRELUDE Experimental Surprises Playing with Light Quantum Field Theory in non-trivial Vacua

**Refractive Index** 

## **MAGIC results (2005)**



First Interesting result... *in conflict* with Conventional Astrophysical acceleration AGN Models (e.g. Crab Nebula) TeV Photons from Active Galactic Nucleus (AGN) Mkn 501 at red-shift z = 0.03

More energetic photons (1.2 - 10 TeV) delayed by O(1 min) compared to E < 0.6 TeV



#### **Other Observed Photon Delays (H.E.S.S, FERMI)**



#### **OPERA RESULTS – Superluminal neutrinos ?**



2011 : Another Surprise, more mysterious

NEUTRINOS IN OPERA have been argued to propagate with superluminal velocities

 $v/c - 1 = (2.48 \pm 0.28 \pm 0.30) 10^{-5}$ 

(but independent of the energy, at least in the range of the experiment)

OPERA v2

 $(v-c)/c = (2.37 \pm 0.32 \text{ (stat.)} ^{+0.34}_{-0.24} \text{ (sys.)}) \times 10^{-5}$ 

overall significance more 6.2  $\sigma$ 

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Can MAGIC *et al.* Photon events be reconciled with OPERA within a fundamental new physics framework?

## **GAMES WITH LIGHT**



Speed of Light in Vacuo = c (Universal constant)

But in MATERIAL SYSTEMS light propagates with different speed,  $v_{light} \neq c$ 

**NON TRIVIAL REFRACTIVE INDEX** 

Phase velocity  $v_{\text{light}} = \frac{c}{\eta}$ 

Group Velocity  $v_{\text{group}} = \frac{\partial \omega}{\partial k} \neq c$ 

## **GAMES WITH LIGHT**



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NON TRIVIAL REFRACTIVE INDEX

$$n_1 \sin \alpha = n_2 \sin \beta$$

## **QUANTUM MECHANICS & REFRACTIVE INDEX**

electrons (mass m) of the medium as forced quantum oscillators, with force exerted by electromagnetc field photons interact with this background

Feynman

Electron area density  $n_e = \rho_e \Delta z$  ( $\rho_e = volume density of electrons$ )



Distance  $\Delta z$  traversed by photons

$$m\left(\frac{d^2}{dt^2}x + \omega_0^2 x\right) = eE_0 e^{i\omega t}$$

Excited-atoms produced electric field

$$E_a = -\frac{en_e}{\epsilon_0 c} i \frac{eE_0}{m(\omega^2 - \omega_0^2)} e^{i\omega(t-z)},$$

Speed of photons in medium c/nsuppressed by refractive index n causing *delay*  $\Delta t = (n-1)\Delta z / c$ 

$$(n-1)\Delta z = \frac{n_e e^2}{2\epsilon_0 m(\omega_0^2 - \omega^2)}$$

$$n = 1 + \frac{\rho_{e_e} e^2}{2\epsilon_0 m (\omega_0^2 - \omega^2)_1}$$

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Depending on Light Polarization

# BIREFRINGENCE







Courtesy Univ. of Surrey





Conventional Refraction

Courtesy Univ. of Surrey



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Distance L

**CASIMIR VACUUM**: Force due to Quantum fluctuations of Photons in space between neutral capacitor plates.

Quantum effects lead to ``vacuum polarization' ' (emission of virtual electron-positron pairs as a photon propagates in this non-trivial vacuum Between the plates) and to a modified group Velocity of photons, larger than c:

$$v_{\rm gr}^{\rm Casimir} = 1 + \frac{11\pi^2}{8100} \alpha^2 \frac{1}{L^4 m_{\rm e}^4} > 1$$

Scharnhost (1990), Barton (1990) Lattore, Pascual, Tarrach (1995)

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 $\alpha$  = e<sup>2</sup> / 4 $\pi$  = Fine structure constant

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Boundary Conditions Break Lorentz Symmetry in vacuo<sup>26</sup>

#### **FINITE TEMPERATURE**

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**PLASMA VACUUM:**Plasma is a state of Matter at very high temperature where Matter is ionized. **Temperature T** 

Quantum effects related to vacuum polarization in this non-trivial vacuum lead to a modified Group Velocity for photons, larger than c

(low T: analogy with Casimir Vacuum:  $L^{-1}$   $\longleftrightarrow$  2T)

High T modes **Subluminal** 

$$v_{\rm gr}(kT \gg m_{\rm e}^2) = 1 - \frac{\alpha^2}{6} \left(\frac{T}{k}\right)^2 \ln^2\left(\frac{kT}{m_{\rm e}^2}\right) < 1$$

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The Vacuum (i.e. lowest energy state) of QUANTUM SYSTEMS may also be characterized by strange properties of light, e.g. Lorentz invariant Breaking ....



**Subluminal** 

### **FINITE TEMPERATURE-NEUTRINOS**

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Quantum effects related to vacuum polarization in this non-trivial vacuum lead to a modified Group Velocity for *neutrinos, smalle*r than c

(low T: analogy with Casimir Vacuum:  $L^{-1}$   $\longleftrightarrow$  2T)

$$v(qT \ll M_z^2) = 1 - g_W^2 \frac{7\pi^2}{45} \left(\frac{T}{M_z}\right)^4 + \mathcal{O}\left(\frac{q^2 T^6}{M_z^8}\right)$$

High T modes Subluminal

$$v(qT\gg M_z^2)=1-g_W^2\frac{1}{24}\left(\frac{T}{q}\right)^2+\mathcal{O}\left(\frac{M_z^4}{q^4}\right)\ ,\quad T\leq q$$

# **CURVED SPACE-TIME BACKGROUNDS**

Drummond & Hathrell (1980)

#### VACUUM POLARIZATION EFFECTS IN QED IN CURVED BACKGROUNDS

**FEEECTIVE HIGHER DERIVATIVE ACTION AFTER ELECTRON FIELDS ARE INTEGRATED OUT IN A PATH INTEGRAL** 

$$S_{\text{eff}} = \int d^4x \sqrt{-g} \left( -\frac{1}{4} F_{\mu\nu} F^{\mu\nu} + W_{1-\text{loop}} \right) \qquad \mathbf{v} \qquad \mathbf{e} \qquad \mathbf{v}$$
$$W_{1-\text{loop}} = \int d^4x \sqrt{-g} \left( aR F_{\mu\nu} F^{\mu\nu} + bR_{\mu\nu} F^{\mu\lambda} F^{\nu}_{\lambda} + cR_{\mu\nu\rho\lambda} F^{\mu\nu} F^{\rho\lambda} + dD_{\mu} F^{\mu\nu} D_{\lambda} F^{\lambda} \nu \right)$$



**Consequence** : modified (by higher derivative terms) photon propagator & dispersion relations, Coefficients proportional to the product of Newton's constant with the fine structure constant  $G_N \alpha$  & inversely proportional to electron mass squared  $m_2^2$ 

**Example:** Photon group velocity in Expanding Friedman-Robertson-Walker (FRW) **Backgrounds** 

$$v = 1 + \frac{11}{45} \alpha G_N \frac{\rho + p}{m_e^2} \qquad > 1$$

*Superluminal* for ordinary matter, radiation fluids **p** + **p** > **0** Light like for Cosmological constant (Lorentz invariant) vacua **p** + **p** = **0** 30

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Shore, Hollowood

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Latorre, Pascual, Tarrach

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energy density of ground state

Latorre, Pascual, Tarrach



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$$n = 1 + a \, \frac{\tilde{\rho} \, g^2}{m_e^4}$$

Energy density of non-trivial vacua may be lower w.r.t. normal

()

e.g. FRW gravitational energy

$$\rho_G = -\tilde{\rho} < 0$$

# **CAUSALITY & SUPERLUMINALITY**

No problems with Causality if Superluminality is Observer dependent

*i.e.* Causality paradox arise if signals travel with the **same** speed V > c in **two different** frames



Liberati, Sonego, Viser (2002)
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Previous examples are fine in this respect, the effects on photon dispersion relation can be represented by `` effective " metrics ,

$$p^{\mu}p^{\nu}g_{\mu\nu} = 0$$

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e.g. in Casimir photons in a cavity ( $n^{\mu}$  unit space-like vector orthogonal to plates)

$$g_{\mu\nu} = \eta_{\mu\nu} - \frac{\xi}{1+\xi} n_{\mu} n_{\nu} \qquad \qquad \xi = \frac{11\pi^2 \alpha^2}{4050 L^4 m_e^2}$$

Parallel Plates

Liberati, Sonego, Viser (2002)

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$$x^{\mu} \to x^{\mu} + \xi^{\mu}(x, u)$$
$$g_{\mu\nu} \to g_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$$

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Quantization of Gravitational Interaction is still a mystery

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> Quantization of Gravitational Interaction is still a mystery

String theory is one approach to quantum gravity.
The theory predicts extra dimensions
But also two ``gravitational'' scales:
(i) one on our world, The Planck Mass scale, M<sub>P</sub> = 1.2 x 10<sup>19</sup> GeV
&

(ii) the bulk (extra dimensional) one, which is the *string mass scale* characterizing string theory itself, and may be as low *as a few TeV* 

Towards Background independence via AdS/CFT

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There are other approaches to QG (Loop Quantum Gravity) with background independence built in but not so advanced though to incorporate the Standard Model of particle physics in their framework

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Quantization of Gravitational Interaction is still a mystery

There are other approaches to QG (Loop Quantum Gravity) with background independence built in but not so advanced though to incorporate the Standard Model of particle physics in their framework

Also : *Doubly* or *Deformed Special Relativities* might be effective theories of QG — their quantization still in progress

cf. Amelino – Camelia talk 45

 QUANTUM FLUCTUATIONS OF SPACE TIME METRIC AT PLANCK SCALES MAY RESULT IN MICROSCOPIC FOAMY SPACE-TIME STRUCTURE

$$\int Dg_{\mu\nu}(x)D(\dots)e^{-\frac{1}{\kappa}\int d^4xR(g)+\dots}$$

$$\kappa \propto G_N = M_P^{-2}$$
  
 $\ell_P = M_P^{-1} = 10^{-35} \text{ m}$ 

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## **Space-time Foam**

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## **Space-time Foam**

Space-Time at Planck scales may have a ``foamy'' structure (J. A. Wheeler ), with possible coordinate non-commutativity or Lorentz Violation at microscopic scales



Quantum Gravity then may behave as a medium, with non-trivial ``optical'' properties:

Vacuum Refractive Index induced by QG ! Energy dependent speed of light, effects increase with energy of photon, due to increase in distortion of space time. Contrast with Matter-induced ordinary refractive indices.

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ACT I ``FOAM'' OLOGY PHENOMENOLOGY OF QG FOAM VACUA

Some Aspects



Modified dispersion due to QG induced space-time (metric) distortions (c=1 units):

$$p^{\mu}p^{\nu}G_{\mu\nu}(\vec{p},E) = 0$$
,  $p^{\mu} = (E,\vec{p})$ 

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Space-time Metric describing space-time Distortions induced by Interactions of Photons with space-time defects NB: momentum dependent metric (Finsler) GEOMETRY OF PHASE-SPACE IMPORTASNT IN QG?



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$$p^{\mu}p^{\nu}G_{\mu\nu}(\vec{p},E) = 0 , \quad p^{\mu} = (E,\vec{p})$$
$$E = p\left(1 + \sum_{n=1}^{\infty} a_n \left(\frac{|\vec{p}|}{M_{QG}}\right)^n\right)$$

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$$V_{\text{phase}} = \frac{E}{|\vec{p}|} = \frac{1}{\eta} , \quad V_{\text{group}} = \frac{\partial E}{\partial |\vec{p}|}$$

 $\eta(|\vec{p}|) = \text{refractive index in vacuo}$ 

subluminal :  $\eta > \mathbf{1}$  , superluminal  $\eta < \mathbf{1}$ 

n



Quantum Gravity then may behave as a medium, with non-trivial ``optical'' properties:

If Vacuum Refractive Index induced by QG non trivial...

....then, it would be Manifested through delays (subluminal) or advances (superluminal) in arrival times of the the more energetic photons in a wave packet or in general a multi frequency group of ``simultaneously" emitted photons.



Subluminal QG-induced Refractive Index: Higher energy photons arrive later

Courtesy: N. Doltsinis@kcl.ac.uk

### **Early Theoretical Predictions**

Time as a Renormalization Group Irreversibly flowing scale in (Liouville, non-critical) String Theory Ellis, NEM, Nanopoulos (1992)

Time – Space different behaviour, different symmetries Lorentz Violation natural consequence of this microscopic approach

``Environment'' of Quantum Gravity (QG) (stringy) d.o.f. inaccessible to low-energy observer scattering experiment; Induced Decoherence and hence Microscopic Time Arrow

Off-shell stringy matter excitations due to interaction with the QG ``environment''

Induced Modified Dispersion Relations (c(E) refractive index) for photon probes due to propagation in QG ``medium''

Amelino Camelia, Ellis, NEM, Mavromatos (1996/05)

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### Gamma Ray Bursts (GRB)

GRB light curves: Fine structure in sharp photon arrival peaks at different energy channels



 Red
 25-50 keV

 Yellow
 50-100 keV

 Green
 100-300 keV

 Blue
 > 300 keV

Arrival Time



Use them to Test energy dependent speed of photons ....

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Amelino Camelia, Ellis, NEM, Nanopoulos, Sarkar (1997/12) Ellis, Farakos, Mitsou, NEM, Nanopoulos (1999/07) Ellis, NEM, Nanopoulos, Sakharov, Sarkisyan (2002- 2011)



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#### Gamma Ray Bursts (GRB)

#### Arrival time delays or advances

$$\Delta t = \frac{1+n}{2H_0} \int_0^z dz \frac{E^n}{M_{\rm QG}^n} \frac{(1+z)^n}{\sqrt{\Omega_M (1+z)^3 + \Omega_\Lambda}} \ , n \ge 1$$

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Ellis, NEM, Nanopoulos, Sakharov, Sarkisyan (2002-2007) Jacob Piran (2007)



## **MAGIC results (2005)**



First Interesting result... *in conflict* with Conventional Astrophysical acceleration AGN Models (e.g. Crab Nebula) TeV Photons from Active Galactic Nucleus (AGN) Mkn 501 at red-shift z = 0.03

More energetic photons (1.2 - 10 TeV) delayed by O(1 min) compared to E < 0.6 TeV

Interestingly can fit QGsubluminal refractive index with linear  $M_{QG}$  suppression with  $M_{QG(1)} = 0.2 \times 10^{18} \text{ GeV}$ 

or, if astrophysics at source taken into account  $M_{QG(1)} > 0.2 \times 10^{18} \text{ GeV}$ 

### **Other Observed Photon Delays (H.E.S.S, FERMI)**



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**YES** *e.g.* IN A SPACE-TIME FOAM MODEL INSPIRED FROM STRING THEORY





# **``Foamy Structures'' in String Theory**

Foamy space-time structures may also be provided by higher-dimensional space-time
 ``real Defects'' in the modern version of string theory involving branes (D-Foam)

Ellis, NEM, Nanopoulos, Sarben Sarkar, Szabo, Westmuckett...

# **BRANE/STRING -THEORY**



#### **BRANE-WORLDS with D-PARTICLE (POINT-LIKE BRANE) DEFECTS**





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**Colliding Brane world model of Space-Time with point-like space-time defects** 



### Two kinds of Interactions



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(I) Just **RECOIL** of massive defect distortion of surrounding space time

Problem Equivalent to Strings propagating in Local ``electric field'' backgrounds Time-Space non-commutativity

$$[X^i, t] \propto F^{0i}(k, x) \equiv u^i(k, x)$$

But electric field is on phase space



D-particle's recoil velocity depends on momentum transfer  $\Delta k_i$ 



 $M_s/g_s$  arbitrary D-particle mass (bulk string scale  $M_s$ not determined, can be as low as a few TeV,  $g_s < 1$ ).

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$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$

Explicit local breaking of SO(3,1) down to SO(2,1) rotation and boosts in transverse directions Problem Equivalent to Strings propagating in Local ``electric field'' backgrounds Time-Space non-commutativity

$$[X^i, t] \propto F^{0i}(k, x) \equiv u^i(k, x)$$

Induced metric depends on momenta as well as coordinates (Finsler type) : e.g. u || X<sub>1</sub>

$$h_{00} = -h_{11} = |u_1|^2$$
$$h_{01} = g_s \frac{\Delta k_i}{M_s} \equiv u_1$$

**Local Lorentz Violation due to** direction of Defect recoil velocities<sub>5</sub> Locally induced metric distortions seen by ``open strings''

$$ds^{2} = g_{\mu\nu}dx^{\mu} dx^{\nu} = -(1 - |\vec{u}|^{2})dt^{2} + (1 - |\vec{u}|^{2})dx_{1}^{2} + dx_{2}^{2} + dx_{3}^{2} + 2\vec{u} \cdot d\vec{x} dt$$

Modified photon dispersion relations due to locally induced metric

$$p^{\mu}p^{\nu}g_{\mu\nu} = 0$$
$$p^{\mu} = (E, \vec{p}) \implies 0 < E = -\vec{p} \cdot \vec{u} + p\left(1 + \frac{1}{2}|\vec{u}|^2 + O(|\vec{u}|^3)\right)$$

(Sub)Superluminal Group photon velocities, depending on relative directions u, p

$$v_g = \frac{\partial E}{\partial p} = 1 - |\vec{u}| \cos\vartheta + \frac{1}{2} |\vec{u}|^2 + \mathcal{O}(|\vec{u}|^3)$$







CHARGE CONSERVATION MUST BE RESPECTED DURING CAPTURE (*STRING SPLITTING*, INTERMEDIATE STRING CREATION & STRETCHING):

ONLY ELECTRICALLY NEUTRAL EXCITATIONS INTERACT VIA CAPTURE DOMINANTLY WITH FOAM DEFECT RECOIL OCCURS

### *Time Delays* due to Intermediate String Creation & Oscillations

J ELLIS, NEM, NANOPOULOS





*Time Delays* due to Intermediate String Creation & Oscillations *Amplitude Pole Analysis* 

ELLIS, NEM, NANOPOULOS, ...+ LI, XIE

Veneziano Amplitude is proportional to  $\pi$ 

$$\frac{\pi}{\sin(\pi s \alpha')} \to e^{i\pi s \alpha' - \epsilon}$$
Poles at s = n/\alpha' (\alpha' = Regge slope)

#### **Shift pole prescription** $s \rightarrow s + i\epsilon$

Delay:

$$\Delta t \sim \alpha' E$$

90







### DELAYS COMPATIBLE WITH STRINGY UNCERTAINTIES

Time Delays due to Intermediate String Creation, growth up to length L & N Oscillations →

$$E = \frac{L}{\alpha'} + \frac{N}{L}$$

Minimise right-hand-size w.r.t. L. End of intermediate string on D3-brane Moves with *speed of light in vacuo c=1* Hence TIME DELAY (causality) during Capture.  $\Delta t \sim \alpha' E$ 



Compatible with the Stringy Uncertainties

 $\Delta t \Delta x \ge \alpha'$ 

$$\Delta t \Delta p \geq 1 + \alpha' (\Delta p)^2$$



#### J ELLIS, NEM, DV NANOPOULOS









### **QUANTUM MECHANICS & REFRACTIVE INDEX**

electrons (mass m) of the medium as forced quantum oscillators, with force exerted by electromagnetc field photons interact with this background

Feynman

Electron area density  $n_e = \rho_e \Delta z$  ( $\rho_e = volume density of electrons$ )



Distance  $\Delta z$  traversed by photons

$$m\left(\frac{d^2}{dt^2}x + \omega_0^2 x\right) = eE_0 e^{i\omega t}$$

Excited-atoms produced electric field

$$E_a = -\frac{en_e}{\epsilon_0 c} i \frac{eE_0}{m(\omega^2 - \omega_0^2)} e^{i\omega(t-z)},$$

Speed of photons in medium c/nsuppressed by refractive index n causing *delay*  $\Delta t = (n-1)\Delta z / c$ 

$$(n-1)\Delta z = \frac{n_e e^2}{2\epsilon_0 m(\omega_0^2 - \omega^2)}$$

$$n = 1 + \frac{\rho_{e_e} e^2}{2\epsilon_0 m (\omega_0^2 - \omega^2)}$$



*Time Delays* due to Intermediate String Creation & Oscillations increase due to recoil





Delays are Independent of photon polarization NO BIREFRINGENCE

## Stringy Uncertainties & D-Foam

- D-foam: populations of defects encountered by probe
- D-foam captures neutral probes & re-emits them
- Time Delay (Causal) in **each** Capture:

$$\Delta t \sim lpha' p^0$$
 p<sup>0</sup> = E

- Independent of photon polarization (no Birefringence)
- Total Delay from emission of photons till observation over a distance D (assume n<sup>\*</sup> defects per string length):

$$\Delta t_{\text{total}} = \alpha' p^0 n^* \frac{D}{\sqrt{\alpha'}} = \frac{p^0}{M_s} n^* D$$

# **COSMOLOGICAL D-FOAM**

Universe Expansion may affect density of defects – n\*(z) Red-shift Dependent

$$\Delta t_{\rm obs} = \int_0^z dz \frac{n(z) E_{\rm obs}}{M_s H_0} \frac{(1+z)}{\sqrt{\Omega_M (1+z)^3 + \Omega_\Lambda}}$$

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$$M_{\rm QG}^{\rm Eff} = \frac{M_{s}}{n^{\star}(z)}$$
Account for MAGIC
(& HESS) events for low z and ALSO for GRB 090510 (short burst) at high z =1 Higher z GRBs delays partly due to D from partly due to D

If brane moves in inhomogeneous bulk

D-loam, partly due to **Sourcce Delayed Emission** 

### **Observed Photon Delays (H.E.S.S, FERMI)**



#### Space time Foam situations – Average over both populations of defects & quantum fluctuations

Isotropic & (in)homogeneous foam

for a brane observer:

$$\langle u_i \rangle \equiv \frac{g_s}{M_s} \langle \Delta k_i \rangle = 0$$

Lorentz Invariance on Average



$$rac{g_s^2}{M_s^2}\langle \Delta k_i \Delta k_j 
angle = \sigma^2 \delta_{ij}$$
 Violated in flcts

 $v_g = 1 + \frac{1}{2} \langle |\vec{u}|^2 \rangle + \mathcal{O}(|\vec{u}|^3) = 1 + \frac{1}{2} \sigma^2 > 1$ 

Superluminal propagation IF NOT CAPTURE

$$\left(\text{cf. locally } v_g = \frac{\partial E}{\partial p} = 1 - |\vec{u}| \cos\vartheta + \frac{1}{2} |\vec{u}|^2 + O(|\vec{u}|^3)\right)_{\text{\tiny 103}}$$

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c.f. Stochastic Foam, through coherent graviton states

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}$$
$$\langle h_{\mu\nu} \rangle = 0$$

 $\langle h_{\mu\nu}h_{\rho\sigma}\rangle \neq 0$ 

leading to light cone fluctuations energy-dependent photon pulse broadening

Ford (95) ....



Subluminal QG-induced Refractive Index: Higher energy photons arrive later Stochastic Light-Cone fluctuations: Energy dependent width of photon pulses (e.g. D-particle (stringy) foam, width proportional to photon energy)

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STOCHASTIC FLUCTUATIONS ALSO INDUCE *QUANTUM DECOHERENCE* FOR LOW-ENERGY MATTER










# **FOAM & QUANTUM DECOHERENCE**

- SPACE-TIME FOAM d.o.f. (metric fluctuations, defect recoil etc.) cannot be detected by a low-energy observer performing only scattering experiments.
- This implies an ``*environment''* for low-energy matter (*open system*) propagating in a *Quantum Gravity* (QG) foamy background.
- INDUCED (LOW-ENERGY) DECOHERENCE
- Coupling of matter system to environment is expected to be generically ``weak'' due to the weakness of quantum graviational interactions.
- *Perturbative* treatment of related *QG effects*

#### MATTER EVOLUTION MASTER EQUATIONS

Time evolution of matter density matrix  $\rho = |\psi\rangle \langle \psi|$ Highly dependent on details of the microscopic model

 $\partial_t \rho = i[\rho, H] + \delta H(\rho)$ 

Basic phenomenological assumption: LINDBLAD form (i) Linear evolution  $\partial_t \rho = i[\rho, H] + \not A H \rho$ (ii) total probability conservation  $Tr(\rho) = 1$ (iii) complete positivity of  $\rho$  eigenvalues  $\geq 0$ (iv) entropy increase  $S = -Tr(\rho \ln(\rho))$ ,  $dS/dt \geq 0$ (v) energy conservation (on the average)

$$\rho = \operatorname{Tr}_{\mathcal{M}} |\Psi \rangle \langle \Psi|$$

ENVIRONMENT OPERATORS  $\mathbf{B}_{\mathbf{m}}$ ,  $\mathbf{B}_{\mathbf{m}}^+$ 

$$\dot{\rho} \equiv \partial_t \rho = i[\rho, H] - \sum_m \{B_m^{\dagger} B_m, \rho\}_+ + 2\sum_m B_m \rho B_m^{\dagger}$$

In terms of state vectors  $|\psi\rangle$  :

$$|d\Psi\rangle = -\frac{i}{\hbar}H|\Psi\rangle + \sum_{m}(\langle B_{m}^{\dagger}\rangle_{\Psi} B_{m} - \frac{1}{2}B_{m}^{\dagger}B_{m} -$$

 $\frac{1}{2} < B_m^{\dagger} >_{\Psi} < B_m >_{\Psi}) |\Psi > dt + \sum_m (B_m - < B_m >_{\Psi}) |\Psi > d\xi_m$ 

$$\rho = \operatorname{Tr}_{\mathcal{M}} |\Psi \rangle \langle \Psi|$$

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(i) Linear evolution

(ii) total probability conservation : Tr(ρ) = 1
but evolution of pure to mixed states: Tr(ρ<sup>2</sup>) ≠ 1
(iii) complete positivity of ρ(t) : Eigenvalues of ρ(t) ≥ 0

(iv) monotonic (Von Neumann) entropy increase:  $S=-Tr(\rho Ln \rho)$ ,  $dS/dt \ge 0$ (v) energy conservation (on the average) (?)

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(iv) monotonic (Von Neumann) entropy increase:  $S = -Tr(\rho \ln \rho)$ ,  $dS/dt \ge 0$ (v) energy conservation (on the average) (?) If (iv) & (v)  $B_j^{\dagger} = B_j$ ,  $[B_j, H] = 0$ self-adjointness

LINDBLAD EVOLUTION  

$$\dot{\rho} = i[\rho, H] + \mathcal{D}[\rho]$$
  
 $\mathcal{D}[\rho] = \sum_{j} [B_j, [B_j, \rho]]$ 

Adler & Horwitz

So Double commutator form for Lindblad term in master equation if: (iv) monotonic (Von Neumann) entropy increase:  $S=-Tr(\rho \ln \rho)$ ,  $dS/dt \ge 0$ and

 $B_j^{\dagger} = B_j, \quad [B_j, H] = 0$ 

(v) energy conservation (on the average)

Gisin & Percival

Assumption: Hamiltonian : block diagonal form in channels  
independent of ``measurement"  
$$H = \begin{pmatrix} H_1 \dots \dots \\ \dots H_2 \dots \\ \dots \dots H_k \end{pmatrix}$$
  
Channel k projection operator:  $P_k$  [P<sub>k</sub>, H] = 0

Localisation (``collapse") is quantified by rate of decrease of Quantum Dispersion Entropy

$$\frac{d}{dt}(\mathcal{MK}) = -\sum_{j} \frac{1 - \langle P_k \rangle_{\psi}}{\langle P_k \rangle_{\psi}} \sum_{m} |\langle P_j B_m P_j \rangle_{\psi}|^2 \le 0$$

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$$\overset{\text{Mean over}}{\swarrow}$$

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Some times: Localization may stop before it is copmplete: Pointer state from decoherence

Zurek

Energy driven Lindblad decoherence

$$D = H$$

Adler & Horwitz

$$d\rho = -i[H,\rho]dt - \frac{1}{8}\sigma^{2}[D,[D,\rho]]dt + \frac{1}{2}\sigma[\rho,[\rho,D]]dW_{t}$$

$$dW_t^2 = dt \ , \ \ dt dW_t = 0$$

(white noise conditions)

Decoherence damping :

$$e^{-\mathcal{D}t}$$

Order of magnitude estimates:

$$\mathcal{D} = \frac{\left(\Delta m^2\right)^2}{E^2 M_{\rm QG}}$$

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Suppressed compared to naïve dimensional estimates of D= E<sup>2</sup>/M<sub>QG</sub>

### D-particle (stringy) Foam:



 $\overline{\partial_t \rho_{Matter}} = i \left[ \rho_{Matter}, H \right] - \Omega \left[ \overline{u}_\ell, \left[ \overline{u}^\ell, \rho_{Matter} \right] \right]$   $\overline{u}_x \to g_s \frac{r}{M_s} \widehat{p}$ 

NEM, Sarkar

Generalization Including damping  $\gamma$ :

$$i\frac{\partial}{\partial t}\rho = \frac{1}{2m}\left[\widehat{p}^2,\rho\right] - i\Lambda\left[\widehat{x},\left[\widehat{x},\rho\right]\right] + \frac{\gamma}{2}\left[\widehat{x},\{\widehat{p},\rho\}\right] - i\Omega r^2\left[\widehat{p},\left[\widehat{p},\rho\right]\right]$$

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Local Lorentz Violation due to direction of Defect recoil velocities

D-particle (stringy) Gaussian Foam:



 $\overline{\partial_t \rho_{Matter}} = i \left[ \rho_{Matter}, H \right] - \Omega \left[ \overline{u}_\ell, \left[ \overline{u}^\ell, \rho_{Matter} \right] \right]$   $\overline{u}_x \to g_s \frac{r}{M_s} \widehat{p} \qquad \left\langle r \right\rangle = 0, \quad \left\langle r^2 \right\rangle = \sigma^2$ 

NEM, Sarkar

Generalization Including damping  $\gamma$ :

$$\begin{aligned} &i\frac{\partial}{\partial t}\rho = \frac{1}{2m}\left[\widehat{p}^{2},\rho\right] - i\Lambda\left[\widehat{x},\left[\widehat{x},\rho\right]\right] + \frac{\gamma}{2}\left[\widehat{x},\left\{\widehat{p},\rho\right\}\right] - i\Omega r^{2}\left[\widehat{p},\left[\widehat{p},\rho\right]\right] \\ &\text{Decoherence damping} \\ &\text{in two-level systems:} \quad e^{-\mathcal{D}t} \qquad \mathcal{D} = \sigma^{2}g_{s}\frac{\left(\Delta m^{2}\right)^{2}}{E^{2}M_{s}} \end{aligned}$$

NEM, Alexandre, Farakos...

# **D-particle Foam:** Cauchy-Lorentz D-particle recoil velocities distribution



$$f(x) = \frac{1}{\pi} \frac{\gamma}{x^2 + \gamma^2}$$

$$\langle \langle \dots \rangle \rangle_{\rm CL} = \int_{-\infty}^{\infty} dx f(x) \dots$$

Decoherence damping in two-level systems:

$$e^{-\mathcal{D}t}$$

$$\mathcal{D} = \gamma \frac{\left(\Delta m^2\right)}{E}$$





Jost, Pauli, Bell, Schwinger

• **CPT Theorem:** Invariance of the Lagrangian of a

(relativistic) field theory under the action (generated by $\theta$ ) of: **C**(harge conjugation),**P**(arity=reflexion) ,**T**(ime reversal) in any order

$$\theta \mathcal{L}(x)\theta^{\dagger} = \mathcal{L}(-x)$$

 $\boldsymbol{\theta}$  is anti-unitary in view of T-reversal

- It is proven for Relativistic field theories, *in FLAT space-times*, upon the assumptions of:
  - (i) Lorentz invariance
  - (ii) Locality of interactions
  - (iii) Unitarity

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### is it fundamental?

#### GREENBERG

Assumed: Covariant T-ordered product, well-defined S-matrix

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Local CPTViolating (CPTV) but LORENTZ INVARTIANT models explicitly constructed but with non covariant T-ordered product and not well defined S-matrix Chaichian et al. PLB669, 177 (2011)

Jost, Pauli, Bell, Schwinger

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May not be valid in highly-curved space-times

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It is proven for Relativistic field theories, *in FLAT space-times*, upon the assumptions of:

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nitarity

(ii) Locality of interactions

May not be valid in highly-curved space-times of QG, e.g. space-time foam



Jost, Pauli, Bell, Schwinger

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## θis anti-unitary in view of T-reversal

 It is proven for Relativistic field theories, in FLAT space-times, upon the assumptions of:

(i) Lorentz invariance
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 Ill-defined CPT operator Wald 1979

Wald (1979)

**Theorem:** If there is quantum *decoherence*, then there is a strong form of CPT Violation, in the sense that the *quantum generator* of CPT symmetry is *ill-defined* 

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**Proof:** If  $\Theta$  acting on density matrices is well-defined then  $\Theta$  is a *unitary operator* ( $\theta$  acting on sate vectors is anti-unitary)

$\Theta \sim \theta \theta^{\dagger}$	$\Theta ho\sim heta ho heta^{\dagger}$
$\Theta \sim \theta \theta'$	$\Theta ho\sim \theta ho  heta$

$$\theta^{\dagger} = -\theta^{-1}$$

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$$\begin{split} \Theta &\sim \theta \theta^{\dagger} & \Theta \rho \sim \theta \rho \theta^{\dagger} \\ \rho_{\text{out}} &= \$ \rho_{\text{in}} & \theta^{\dagger} = -\theta^{-1} \end{split}$$

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Hawking super-scattering operator (linear) If decoherence, \$ \$ \$ \$ \$ \$ (\$ = e<sup>iHt</sup>) has NO INVERSE

Wald (1979)

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# DETAILED PROOF

$$\begin{split} \Theta \rho_{in} &= \overline{\rho}_{out} & \text{was} \\ \rho_{in} &= \Theta^{-1} \overline{\rho}_{out} & \text{uni} \\ \rho_{out} &= \$ \rho_{in} & \Theta \rho \\ \overline{\rho}_{out} &= \$ \overline{\rho}_{in} & \$ \Theta^{-1} \rho \\ \$ \Theta^{-1} \rho & \$ \Theta^{-1} \rho \\ \$ \Theta^{-1} \rho & \flat \Theta^{-1} \rho \\ \bullet \Theta^{-1} \rho & \flat \Theta^{-1} \rho \\ \bullet \Theta^{-1} \rho & \flat \Theta^{-1} \rho \\ \bullet \Theta^{-1} \rho & \bullet \Theta^{-1} \rho \\ \bullet$$

WELL-DEFINED unitary CPT generator  $\Theta^{-1} = \Theta^+$ 

Wald (1979)

**Theorem:** If there is quantum *decoherence*, then there is a strong form of CPT Violation, in the sense that the *quantum generator* of *CPT* symmetry is *ill-defined* 

**Proof:** If  $\Theta$  acting on density matrices is well-defined then  $\Theta$  is a *unitary operator* ( $\theta$  acting on sate vectors is anti-unitary)

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### CPT symmetry without CPT invariance Wald (79)



#### But....nature may be tricky: WEAK FORM OF CPT **INVARIANCE** might exist, such that the fundamental "arrov of time" does not show up in any experimental measurements (scattering experiments).

Probabilities for transition from  $\psi =$ initial pure state to  $\phi = final state$ 

$$P(\psi \to \phi) = P(\theta^{-1}\phi \to \theta\psi)$$

where  $\theta: \mathcal{H}_{in} \to \mathcal{H}_{out}$ ,  $\mathcal{H} = Hilbert$  state space,  $\Theta \rho = \theta \rho \theta^{\dagger}, \quad \theta^{\dagger} = -\theta^{-1} \text{ (anti - unitary)}.$ 

In terms of superscattering matrix \$:

$$\$^{\dagger} = \Theta^{-1} \$ \Theta^{-1}$$

Here,  $\Theta$  is well defined on pure states, but \$ has no inverse, hence  $\stackrel{\dagger}{=} \stackrel{\neq}{=} \stackrel{-1}{=} ($ full CPT invariance:  $\stackrel{\$}{=} SS^{\dagger}, \stackrel{\$^{\dagger}}{=} \stackrel{\$^{-1}}{=} ).$ 

## Weak CPT invariance in Black-Holes? Wald (79)

Supporting evidence for Weak CPT from Black-hole thermodynamics: Although white holes do not exist (strong CPT violation), nevertheless the CPT reverse of the most probable way of forming a black hole is the most probable way a black hole will evaporate: the states resulting from black hole evaporation are precisely the CPT reverse of the initial states which collapse to form a black hole.

### EXPERIMENT CAN TELL...

#### IF CPT IS ILL-DEFINED (INTRINSIC VIOLATION) QUITE DIFFERENT PHENOMENOLOGY FROM ORDINARY LORENTZ- AND CPT-VIOLATION CASES WHERE $\Theta$ IS WELL DEFINED BUT $[\Theta, H] \neq 0$

Kostelecky et al. Standard Model extension

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In principle... depends on the order of magnitude of the effect Highly microscopic model dependent...





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  - Lorentz & CPT Violation in the Hamiltonian
    - Neutral Mesons & Factories, Atomic Physics, Anti-matter factories, Neutrinos, ...
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$$K^{0}\rangle = |d\,\overline{s}\rangle$$
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Physical (observable) states are the ``weak" (or CP) eigenstates (having definite life times under weak interaction decays):

$$\begin{aligned} |K_L\rangle &\propto (1+\epsilon-\delta) \left| K^0 \right\rangle - (1-\epsilon+\delta) \left| \bar{K}^0 \right\rangle \quad \text{CP}_{=+1} \\ |K_S\rangle &\propto (1+\epsilon+\delta) \left| K^0 \right\rangle + (1-\epsilon-\delta) \left| \bar{K}^0 \right\rangle \quad \text{CP}_{=-1} \end{aligned}$$

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e.g. due to Lorentz Violation

$$\begin{split} |K^{0}\rangle &= |d\,\overline{s}\rangle & \text{masses:} & m_{K^{0}} = 497.614 \pm 0.024 \text{ MeV} \\ \text{Bounds on CPTV:} \\ |\overline{K}^{0}\rangle &= |\overline{d}\,s\rangle & \frac{m_{K^{0}} - m_{\overline{K}^{0}}}{m_{K^{0}}} < 8 \times 10^{-19} \end{split}$$

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mean life:  $K_s$ : (0.8958 ± 0.0005) × 10<sup>-10</sup> s

 $K_L$ : (5.116 ± 0.020) × 10<sup>-8</sup> s

mass difference:  $m_{K_L} - m_{K_S} = (3.483 \pm 0.006) \times 10^{-15} \, \text{GeV}$ 

#### QG DECOHERENCE IN NEUTRAL KAONS: SINGLE STATES

Quantum Gravity (QG) may induce decoherence and oscillations  $K^0 \rightarrow \overline{K}^0 \Rightarrow$  could use Lindblad-type approach (one example) (Ellis, Hagelin, Nanopoulos, Srednicki, Lopez, NM):

 $\partial_t \rho = i[\rho, H] + \delta H \rho$ 

where

$$H_{\alpha\beta} = \begin{pmatrix} -\Gamma & -\frac{1}{2}\delta\Gamma & -\mathrm{Im}\Gamma_{12} & -\mathrm{Re}\Gamma_{12} \\ -\frac{1}{2}\delta\Gamma & -\Gamma & -2\mathrm{Re}M_{12} & -2\mathrm{Im}M_{12} \\ -\mathrm{Im}\Gamma_{12} & 2\mathrm{Re}M_{12} & -\Gamma & -\deltaM \\ -\mathrm{Re}\Gamma_{12} & -2\mathrm{Im}M_{12} & \deltaM & -\Gamma \end{pmatrix}$$

and

positivity of  $\rho$  requires:  $\alpha, \gamma > 0$ ,  $\alpha \gamma > \beta^2$ .

 $\alpha, \beta, \gamma$  violate CPT (Wald : decoherence) & CP:  $CP = \sigma_3 \cos \theta + \sigma_2 \sin \theta$ ,  $[\delta H_{\alpha\beta}, CP] \neq 0$ 

• Complete Positivity Different parametrization of Decoherence matrix (Benatti-Floreanini) (in  $\alpha, \beta, \gamma$  framework:  $\alpha = \gamma, \beta = 0$ )

**Current Experimental Bounds:** 

 $\alpha \approx 10^{-17} \text{ GeV}$ 

 $\gamma \approx 10^{-21} \text{ GeV}$ 

 $\beta \approx 10^{-19} \text{ GeV}$ 

Can test complete positivity experimentally (in principle!)

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Current Experimental Bounds:

 $a \approx 10^{-17} \text{ GeV}$ 

 $\gamma \approx 10^{-21} \text{ GeV}$ 

NB: decoherence Damping of oscillations  $e^{-\mathcal{D}t}$ 

 $\mathcal{D} \sim \alpha + \gamma$ 

e.g. Gaussian D- foam

$$\mathcal{D} \sim g_s \sigma^2 \frac{(\Delta m^2)^2}{E^2 M_s}$$

 $\beta \approx 10^{-19} \text{ GeV}$ 

 $\Delta m \sim 3.5 \times 10^{-15} {
m GeV}$ 

 Complete Positivity Decoherence matrix
 (Benatti-Floreanini)
 (in a 6 a framework: a = 0)

(in  $\alpha, \beta, \gamma$  framework:  $\alpha = \gamma, \beta = 0$ )



Current Experimental Bounds:

 $a \approx 10^{-17} \text{ GeV}$ 

 $\gamma \approx 10^{-21} \text{ GeV}$ 

NB: decoherence Damping of oscillations  $-\mathcal{D}t$ 

e.g. Cauchy – Lorentz

 $\mathcal{D} \sim \alpha + \gamma$ 

$$\mathcal{D} \sim \gamma \frac{(\Delta m^2)}{E}$$

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#### EPR correlated states and particle physics

What are EPR correlations?

Einstein-Podolsky-Rosen (EPR) effect proposed originally as a PARADOX testing foundations of Quantum Theory.

Correlations between spatially separated events, instant transport of information? contradicts relativity?

#### NO, NO PARADOX

EPR has been confirmed EXPERIMENTALLY:

- (i) pair of particles can be created in a definite quantum state,
- (ii) move apart,
- (iii) decay when they are widely separated (spatially).



EPR CORRELATIONS between different decay modes should be taken into account, when interpreting any experiment. (Lipkin (1968))

- CPT Violation Consequences for Neutral mesons
- Einstein Podolsky Rosen (EPR) correlators



Neutral Kaon, anti-Kaon mesons treated as indistinguishable particles, Bose-statistics applies

Bernabeu, NEM, Papavassiliou (04)



Bernabeu, NEM, Papavassiliou (04)







$$|K_{S}\rangle \propto (1 + \epsilon + \delta) |K^{0}\rangle + (1 - \epsilon - \delta) |\bar{K}^{0}\rangle$$

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#### $\phi$ Decays and the $\omega$ Effect

Consider the  $\phi$  decay amplitude: final state X at  $t_1$  and Y at time  $t_2$  (t = 0 at the moment of  $\phi$  decay)



Amplitudes:

$$A(X,Y) = \langle X|K_S \rangle \langle Y|K_S \rangle \mathcal{N} \ (A_1 + A_2)$$

with

$$A_{1} = e^{-i(\lambda_{L}+\lambda_{S})t/2} [\eta_{X}e^{-i\Delta\lambda\Delta t/2} - \eta_{Y}e^{i\Delta\lambda\Delta t/2}]$$
  

$$A_{2} = \omega[e^{-i\lambda_{S}t} - \eta_{X}\eta_{Y}e^{-i\lambda_{L}t}]$$

the CPT-allowed and CPT-violating parameters respectively, and  $\eta_X = \langle X | K_L \rangle / \langle X | K_S \rangle$  and  $\eta_Y = \langle Y | K_L \rangle / \langle Y | K_S \rangle$ .

The "intensity"  $I(\Delta t)$ : ( $\Delta t = t_1 - t_2$ ) is an observable

$$I(\Delta t) \equiv \frac{1}{2} \int_{|\Delta t|}^{\infty} dt \, |A(X,Y)|^2$$

 $\omega$ -Effect & Intensities

$$I(\Delta t) \equiv \frac{1}{2} \int_{|\Delta t|}^{\infty} dt \, |A(\pi^{+}\pi^{-}, \pi^{+}\pi^{-})|^{2} = |\langle \pi^{+}\pi^{-}|K_{S}\rangle|^{4} |\mathcal{N}|^{2} |\eta_{+-}|^{2} \Big[ I_{1} + I_{2} + I_{12} \Big]$$

$$\begin{split} I_{1}(\Delta t) &= \frac{e^{-\Gamma_{S}\Delta t} + e^{-\Gamma_{L}\Delta t} - 2e^{-(\Gamma_{S}+\Gamma_{L})\Delta t/2}\cos(\Delta M\Delta t)}{\Gamma_{L} + \Gamma_{S}} \\ I_{2}(\Delta t) &= \frac{|\omega|^{2}}{|\eta_{+-}|^{2}} \frac{e^{-\Gamma_{S}\Delta t}}{2\Gamma_{S}} \\ I_{12}(\Delta t) &= -\frac{4}{4(\Delta M)^{2} + (3\Gamma_{S}+\Gamma_{L})^{2}} \frac{|\omega|}{|\eta_{+-}|} \times \\ \left[ 2\Delta M \left( e^{-\Gamma_{S}\Delta t}\sin(\phi_{+-}-\Omega) - e^{-(\Gamma_{S}+\Gamma_{L})\Delta t/2}\sin(\phi_{+-}-\Omega + \Delta M\Delta t) \right) \right. \\ \left. - (3\Gamma_{S}+\Gamma_{L}) \left( e^{-\Gamma_{S}\Delta t}\cos(\phi_{+-}-\Omega) - e^{-(\Gamma_{S}+\Gamma_{L})\Delta t/2}\cos(\phi_{+-}-\Omega + \Delta M\Delta t) \right) \right] \right] \end{split}$$

 $\Delta M = M_S - M_L \text{ and } \eta_{+-} = |\eta_{+-}| e^{i\phi_{+-}}.$ NB: sensitivities up to  $|\omega| \sim 10^{-6}$  in  $\phi$  factories, due to enhancement by  $|\eta_{+-}| \sim 10^{-3}$  factor.  $\omega$ -Effect & Intensities

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Characteristic cases of the intensity  $I(\Delta t)$ , with  $|\omega| = 0$  (solid line) vs  $I(\Delta t)$ (dashed line) with (from top left to right): (i)  $|\omega| = |\eta_{+-}|$ ,  $\Omega = \phi_{+-} - 0.16\pi$ , (ii)  $|\omega| = |\eta_{+-}|$ ,  $\Omega = \phi_{+-} + 0.95\pi$ , (iii)  $|\omega| = 0.5|\eta_{+-}|$ ,  $\Omega = \phi_{+-} + 0.16\pi$ , (iv)  $|\omega| = 1.5|\eta_{+-}|$ ,  $\Omega = \phi_{+-}$ .  $\Delta t$  is measured in units of  $\tau_S$  (the mean life-time of  $K_S$ ) and  $I(\Delta t)$  in units of  $|C|^2 |\eta_{+-}|^2 |\langle \pi^+ \pi^- |K_S \rangle|^4 \tau_S$ .

Bernabeu, NEM, Papavassiliou (04)



**W** - effect : can be distinguished from conventional C-even background effects

 $e^+ e^- \to 2\gamma \to K^0 \overline{K}^0$ 

Different interference effects

### **CPTV & EPR-correlations modification**

(Bernabeu, NM and Papavassiliou, hep-ph/0310180 (PRL 92) )

If CPT is broken via Quantum Gravity (QG) decoherence effects on  $\$ \neq SS^{\dagger}$ , then: CPT operator  $\Theta$  is ILL defined  $\Rightarrow$  Antiparticle Hilbert Space INDEPENDENT OF particle Hilbert space.

Neutral mesons  $K^0$  and  $\overline{K}^0$  SHOULD NO LONGER be treated as IDENTICAL PARTICLES.  $\Rightarrow$  initial Entangled State in  $\phi$  (B) factories  $|i\rangle$  (in terms of mass eigenstates):

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NB!  $K_S K_S$  or  $K_L - K_L$  combinations, due to CPTV  $\omega$ , important in decay channels. There is contamination of C(odd) state with C(even). Complex  $\omega$  controls the amount of contamination by the "wrong" (C(even)) symmetry state.

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NB2: Also disentangle  $\omega$  from non-unitary evolution ( $\alpha = \gamma$  ...) effects (different structures) (Bernabéu, NM, Papavassiliou, Weldron NP B744.180-206,2006)

### Disentangling $\omega$ -effects from Background

**CPTV**  $K_L K_L$ ,  $\omega K_S K_S$  terms originate from  $\Phi$ -particle , hence same dependence on centre-of-mass energy s. Interference proportional to real part of amplitude, exhibits peak at the resonance....


# **Disentangling ω-effects from Background**



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Alvarez, Bernabeu NEM, Nebot, Papavassiliou

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- B-systems do not have such a "good'' channel but have the *advantage* of statistics  $\rightarrow$  Interesting limits of  $\omega$ -effects there
- Flavour tagging: Knowledge that one of the two-mesons in a meson factory decays at a given time through *flavour-specific* "channel'? Unambiguously determine the *flavour* of the other meson at the *same time*.

**Not True if intrinsic CPTV – ω-effect present** : Theoretical limitation ("demise'') of flavour tagging

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EquaL-Sign di-lepton charge asymmetry Δt dependence

ALVAREZ, BERNABEU, NEBOT

- But Interesting tests of the  $\omega$ -effect can be performed by looking at the equal-sign di-lepton decay channels

a first decay  $B \to X\ell^{\pm}$  and a second decay,  $\Delta t$  later,  $B \to X'\ell^{\pm}$ 

### EquaL-Sign di-lepton charge asymmetry Δt dependence

$$I(X\ell^{\pm}, X'\ell^{\pm}, \Delta t) = \int_0^\infty \left| \langle X\ell^{\pm}, X'\ell^{\pm} | U(t_1) \otimes U(t_1 + \Delta t) | \psi(0) \rangle \right|^2 dt_1$$

$$\begin{split} I(X\ell^{\pm}, X'\ell^{\pm}, \Delta t) &= \frac{1}{8} e^{-\Gamma\Delta t} |A_X|^2 |A_{X'}|^2 \left| \frac{(1+s_{\epsilon} \epsilon)^2 - \delta^2/4}{1-\epsilon^2 + \delta^2/4} \right|^2 \\ & \left\{ \begin{bmatrix} \frac{1}{\Gamma} + a_{\omega} \frac{8\Gamma}{4\Gamma^2 + \Delta m^2} Re(\omega) + \frac{1}{\Gamma} |\omega|^2 \end{bmatrix} \cosh\left(\frac{\Delta\Gamma\Delta t}{2}\right) + \\ - \frac{1}{\Gamma} + b_{\omega} \frac{8\Gamma}{4\Gamma^2 + \Delta m^2} Re(\omega) - \frac{\Gamma}{\Gamma^2 + \Delta m^2} |\omega|^2 \end{bmatrix} \cos(\Delta m\Delta t) + \\ & \left[ d_{\omega} \frac{4\Delta m}{4\Gamma^2 + \Delta m^2} Re(\omega) + \frac{\Delta m}{\Gamma^2 + \Delta m^2} |\omega|^2 \right] \sin(\Delta m\Delta t) \right\}, \end{split}$$



$$\mathbf{A}_{sl}(\Delta \mathbf{t}_{peak}) = \mathbf{0.77} \cos(\mathbf{\Omega}) \qquad \Delta t_{peak} = \frac{1}{\Gamma} \sqrt{\frac{2}{1 + x_d^2}} \left| \boldsymbol{\omega} \right| + \mathcal{O}(\boldsymbol{\omega}^2) \approx \frac{1}{\Gamma} 1.12 \left| \boldsymbol{\omega} \right|$$

### EquaL-Sign di-lepton charge asymmetry



Approximate Periodicity of  $A_{sl}$  in  $\Delta t \Delta m$ : terms  $\cosh(\Delta \Gamma \Delta t)$  almost constant for small  $\Delta \Gamma$ 

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EquaL-Sign di-lepton charge asymmetry Δt dependence

1

Current Experimental measurements of A<sub>sl</sub> have been performed for

### $0.8 \Gamma^{-1} \leq \Delta t \leq 10 \Gamma^{-1}$

(BaBar, Belle Collaborations)

# **Current Experimental Limits**





EquaL-Sign di-lepton charge asymmetry Δt dependence

Current Experimental measurements of A<sub>si</sub> have been performed for

### 0.8 $\Gamma^{-1}$ ≤ Δt ≤ 10 $\Gamma^{-1}$

(BaBar, Belle Collaborations)

At large times region, where peak repeats itself, the amount of events is suppressed by:

e<sup>-8.2</sup> ~ 10<sup>-4</sup>

Future super B experiments could exhibit sensitivity in such regions (?)

1

(Bernabéu, Sarben Sarkar, NM, hep-th/0606137)

Theoretical models using interactions of particle-probes with specific space-time defects (e.g. D-particles, inspired by string/brane theory); Use stationary perturbation theory to describe gravitationally dressed 2-meson state - medium effects like MSW  $\Rightarrow$  initial state:

$$|\psi\rangle = |k,\uparrow\rangle^{(1)} |-k,\downarrow\rangle^{(2)} - |k,\downarrow\rangle^{(1)} |-k,\uparrow\rangle^{(2)} + \xi |k,\uparrow\rangle^{(1)} |-k,\uparrow\rangle^{(2)} + \xi' |k,\downarrow\rangle^{(1)} |-k,\downarrow\rangle^{(2)}$$

NB:  $\xi = -\xi'$ : strangeness conserving  $\omega$ -effect  $(|K_L\rangle = |\uparrow\rangle$ ,  $|K_S\rangle = |\downarrow\rangle$ .).

In recoil D-particle stochastic model: (momentum transfer:  $\Delta p_i \sim \zeta p_i$ ,  $\langle \Delta p_i \rangle = 0$ ,  $\langle \Delta p_i \Delta p_j \rangle \neq 0$ )

$$|\omega|^2 \sim \frac{\zeta^2 k^4}{M_P^2 (m_1 - m_2)^2}$$

NB: For neutral kaons, with momenta of the order of the rest energies  $|\omega| \sim 10^{-4} |\zeta|$ . For  $1 > \zeta \ge 10^{-2}$  not far below the sensitivity of current facilities, such as DA $\Phi$ NE Constrain  $\zeta$  significantly in upgraded facilities.

Perspectives for KLOE-2 at DA $\Phi$ NE-2 (A. Di Domenico home page) : Re( $\omega$ ), Im( $\omega$ )  $\longrightarrow 2 \times 10^{-5}$ .

NB:  $\omega$ -Effect also generated by propagation through the medium, but with time-dependent (sinusoidal)  $\omega(t)$ -terms, can be (in principle) disentangled from initial-state ones...



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• If foam, ``flavour'' changes Bernabeu, NEM, Sarkar



# $\omega$ -Effect estimates in D-particle Foam

 $M_{\rm QG}$  = 2. 10<sup>19</sup> GeV  $\Delta m \sim 3.5 \times 10^{-15} {
m GeV}$ 

### ω-Effect estimates in D-particle Foam



In above estimates: QCD effects & sub-structure in neutral mesons ignored, and D-foam acts as if they were structureless particles, then for  $M_{\rm QG} \simeq 10^{19}~{\rm GeV}$ 

the estimate for  $\omega$ :  $|\omega| \sim 10^{-4} |\zeta|$ , for  $1 > |\zeta| > 10^{-2}$  (natural)

Not far from sensitivity of upgraded meson factories (e.g. DAFNE2)

Re( $\omega$ ), Im( $\omega$ ) = O(10<sup>-5</sup>)

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Not far from sensitivity of upgraded meson factories (e.g. DAFNE2)

In D-foam: quarks (charged) do not directly interact with D-particles only gluons inside the Kaons do, suppressed effects? QCD details matter? *need to estimate* 

(Bernabéu, Sarben Sarkar, NM, hep-th/0606137)

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# $\omega\text{-effect}$ as discriminant of foam models

Bernabeu, NEM, Sarben Sarkar

•  $\omega$ -effect not generic, depends on details of foam

• (I) D-foam: 
$$\widehat{H_I} = -(r_1\sigma_1 + r_2\sigma_2)\,\widehat{k}$$

### Features: direction of k violates local Lorentz symmetry, flavour non conservation non-trivial ω-effect

(II) Quantum Gravity Foam as "thermal isotropic Bath"  

$$\mathcal{H} = \nu a^{\dagger}a + \frac{1}{2}\Omega\sigma_{3}^{(1)} + \frac{1}{2}\Omega\sigma_{3}^{(2)} + \gamma \sum_{i=1}^{2} \left(a\sigma_{+}^{(i)} + a^{\dagger}\sigma_{-}^{(i)}\right)$$
Bath frequency  
"atom" (matter) frequency





#### **OPERA & MAGIC RESULTS CAN BE RECONCILED : e.g. IN D-FOAM**

**TO RECAP**: (I) RECOIL OF MASSIVE DEFECTS DISTORS SPACE TIME , MODIFIES DISPERSION RELATIONS FOR MATTER PROBES

#### (II) CAPTURE OF MATTER BY DEFECT LEAD TO EXTRA TIME DELAYS DUE TO STRINGY NATURE/MINIMUM LENGTH OF STRINGS, WHICH ARE ENHNACED BY RECOIL

$$\Delta t \sim \alpha' E$$

FOR CHARGE CONSERVATION REASONS CAPTURE (HENCE DELAYS) UNDERGO ONLY ELECTRICALLY NEUTRAL PARTICLES (PHOTONS...)

NEUTRINOS IN STRING THEORIES MAY HAVE ADDITIONAL ``CHARGES''/FLUXES DUE TO EXTRA U(1)' GROUP. IN SUCH A CASE NEUTRINOS WILL ALSO AVOID CAPTURE AND WILL THUS BE SUBJECTED ONLY TO (superluminal) GRAVITATIONAL D-foam RECOIL EFFECTS

$$v_g = 1 + \frac{1}{2} \langle |\vec{u}|^2 \rangle + O(|\vec{u}|^3) = 1 + \frac{1}{2} \sigma^2 > 1$$



# **MAGIC results (2005)**



First Interesting result... *in conflict* with Conventional Astrophysical acceleration AGN Models (e.g. Crab Nebula) TeV Photons from Active Galactic Nucleus (AGN) Mkn 501 at red-shift z = 0.03

More energetic photons (1.2 - 10 TeV) delayed by O(1 min) compared to E < 0.6 TeV

Interestingly can fit QG subluminal refractive index with linear  $M_{QG}$  suppression with  $M_{QG(1)} = 0.2 \times 10^{18}$  GeV

or, if astrophysics at source taken into account  $M_{QG(1)} > 0.2 \times 10^{18} \text{ GeV}$ 

### **Observed Photon Delays (H.E.S.S, FERMI)**



### **TWO PROCESSES IN D-FOAM/MATTER STRING INTERACTIONS**

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#### **Space time Foam** situations – Average over both populations of defects & quantum fluctuations

### Isotropic & (in)homogeneous foam

for a brane observer:

$$\langle u_i \rangle \equiv \frac{g_s}{M_s} \langle \Delta k_i \rangle = 0$$

Lorentz Invariance on Average



$$rac{g_s^2}{M_s^2} \langle \Delta k_i \Delta k_j 
angle = \sigma^2 \delta_{ij}$$
 Violated in flcts

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 Violated in flcts

$$g = 1 + \frac{1}{2} \langle |\vec{u}|^2 \rangle + O(|\vec{u}|^3) = 1 + \frac{1}{2} \sigma^2 > 1$$
  
Superluminal propagation  
IF NOT CAPTURE

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Superluminal propagation IF NOT CAPTURE

$$\left(\text{cf. locally } v_g = \frac{\partial E}{\partial p} = 1 - |\vec{u}| \cos\vartheta + \frac{1}{2} |\vec{u}|^2 + O(|\vec{u}|^3)\right)_{\text{\tiny 213}}$$

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## **POSSIBLE TO RECONCILE**



**OPERA** ``ADVANCES''

**DUE TO RECOIL-INDUCED SPACE-TIME DISTORTION** FOR NEUTRINOS, **NO CAPTURE BY FOAM POSSIBLY DUE** TO U'(1) GAUGE SYMMETRIES

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**DUE TO CAPTURE BY D-FOAM** 



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**DUE TO CAPTURE BY D-FOAM** 

**NO PROBLEMS WITH CAUSALITY** 


# **CAUSALITY & SUPERLUMINALITY**

#### No problems with Causality if Superluminality is Observer dependent

*i.e.* Causality paradox arise if signals travel with the **same** speed V > c in **two different** frames



Liberati, Sonego, Viser (2002)

Previous examples are fine in this respect, the effects on photon dispersion relation can be represented by `` effective '' metrics ,  $p^{\mu}p^{\nu}g_{\mu\nu}=0$ 

e.g. in Casimir photons in a cavity (  $n^{\mu}$  unit space-like vector orthogonal to plates )

$$x^{\mu} \to x^{\mu} + \xi^{\mu}(x, u)$$
$$g_{\mu\nu} \to g_{\mu\nu} + \partial_{(\mu}\xi_{\nu)}$$

As such, speed of signal depends on observer velocities  $u^{\mu}$  relative to system



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DUE TO CAPTURE BY D-FOAM

> NO NEUTRINO CHERENKOV Radiation (No Cohen-Glashow pair-production effect)



### NO NEUTRINO CHERENKOV Radiation in D-foam (No Cohen-Glashow pair-production effect)

 $\nu_{\tau} \to \nu_{\tau} + e^+ + e^-$ 

Kinematically allowed if neutrino speed is faster than c=1 in vacuo and there is a preferred frame

$$p^{\mu}p^{\nu}g_{\mu\nu} = 0$$

$$x^{\mu} \to x^{\mu} + \xi^{\mu}(x, u)$$
$$g_{\mu\nu} \to g_{\mu\nu} + \partial_{(\mu} \xi_{\nu)}$$

HOWEVER CHERENKOV RADIATION (DECAY) IS OBSERVER INDEPENDENT ONE CAN GO (WITHIN STRING THEORY) TO THE REST FRAME OF NEUTRINO, WHERE SUCH A CHERENKOV PROCESS IS NOT ALLOWED KINEMATICALLY.



**OPERA** ``ADVANCES''

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WITH ``MAGIC" PHOTON ``DELAYS"

$$\Delta t \sim \alpha' E$$

DUE TO CAPTURE BY D-FOAM

> OTHER EXPLANATIONS (including weird effective geometries & boundary conditions in the conditions of OPERA experiment ) POSSIBLE OF COURSE.....



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... & WITH EPR MODIFICATIONS IN ENTANGLED PARTICLE STATE (ω-effect)

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$$\sigma^2 \sim O(10^{-5})$$

**DAΦNE-2**?

... & WITH EPR MODIFICATIONS IN ENTANGLED PARTICLE STATE (ω-effect)

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### Conclusions

- Quantum Gravity may affect the matter quantum mechanical behaviour by ``opening up'' the matter subsystem in certain models of space-time foam
- Low-energy experimentalists do not have access to QG d.o.f., hence effective *decoherence*, affects Quantum Mechanical time *evolution* of matter
- This may induce an ill-defined CPT operator (perturbatively, antiparticle exists) for the low-energy matter subsystem with
   ``smoking-gun'' evidence in entangled particle states experiment
   Modified EPR correlations ...model dependent

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### Outlook

- Plethora of tests if Lorentz invariance is violated but not so many if conserved on average but violated in quantum fluctuations
- Decoherence may be compatible with Lorentz Invariance on average – in such a case ω-effect may be a smoking gun, but for next generation facilities to have *sensitivity* one needs densities of defects at present as *one per string volume*
- Such a situation *affects early Universe cosmology*
- Could also lead to observable vacuum refraction
   Astrophysical & Terrestrial tests of the latter from cosmic
   photons and/or neutrinos...(Energy dependent) Arrival-time
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- **CALCULATE EFFECTS IN DETAILED MODELS... KEEP SEARCHING....**













### D-particle Recoil & the "Flavour" Problem

Not all particle species interact the same way with D-particles e.g. electric charge symmetries should be preserved, hence electrically-charged excitations cannot split and attach to neutral D-particles....

Neutrinos (or neutral mesons) are good candidates...

But there may be flavour oscillations during the capture/recoil process, i.e. wavefunction of recoiling string might differ by a phase from incident one....

In statistical populations of D-particles, one might have isotropic situations, with  $\ll u_i \gg = 0$ , but stochastically fluctuating  $\ll u_i u^i \gg \neq 0$ . For slow recoiling heavy D-particles the resulting Hamiltonian, expressing interactions of neutrinos (or "flavoured" particles, including oscillating neutral mesons), reads:

D-particle Recoil & the "Flavou	NB: direction	
Not all particle species interact the same way with D- e.g. electric charge symmetries should be preserved, l electrically-charged excitations cannot split and attack neutral D-particles <i>Neutrinos (or neutral mesons) are good candidates</i> But there may be flavour oscillations during the captur function of recoiling string might differ by a phase from In statistical populations of D-particles, one might have $u_i >> = 0$ , but stochastically fluctuating $<< u_i u^i >> \neq 0$ For slow recoiling heavy D-particles the resulting Ham interactions of neutrinos (or "flavoured" particles, inc	of recoil dependence LIV+ Stochastically flct. Environment Decoherence, CPTV ill defined	- <<
$\widehat{H} = g^{01} \left( g^{00} \right)^{-1} \widehat{k} - \left( g^{00} \right)^{-1} \sqrt{\left( g^{01} \right)^2 k^2 - g^{00} \left( g^{11} k^2 + m^2 \right)}$		
$g^{\sigma_{I}} = g^{r_{0}} = r_{0}1 + r_{1}\sigma_{1} + r_{2}\sigma_{2}$ $H_{I} = -(r_{1}-r_{1})$ $\langle r_{\mu}\rangle = 0, \ \langle r_{\mu}r_{\nu}\rangle = \Delta_{\mu}\delta_{\mu\nu}$	$\Delta_{\mu} \sim O\left(\frac{E}{M}\right)$	$\left(\frac{2}{\frac{2}{P}}\right)$



**Apply non-degenerate perturbation theory to construct "gravitationally** dressed' states from  $|k,\uparrow\rangle^{(i)}, |k,\downarrow\rangle^{(i)}, i = 1,2$  $\left|k^{(i)},\downarrow\right\rangle_{QG}^{(i)} = \left|k^{(i)},\downarrow\right\rangle^{(i)} + \left|k^{(i)},\uparrow\right\rangle^{(i)}\alpha^{(i)} \qquad \alpha^{(i)} = \frac{(i)\left\langle\uparrow,k^{(i)}\right|\widehat{H_{I}}\left|k^{(i)},\downarrow\right\rangle^{(i)}}{E_{0}-E_{1}}$  $\widehat{H}_{I} = -(r_{1}\sigma_{1} + r_{2}\sigma_{2})\widehat{k} \qquad \langle r_{\mu}\rangle = 0, \ \langle r_{\mu}r_{\nu}\rangle = \Delta_{\mu}\delta_{\mu\nu}$ 

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angle^{(i)}$  dressed state  $\left|\downarrow\right\rangle\leftrightarrow\left|\uparrow\right\rangle$  and  $\alpha\rightarrow\beta$ **Similarly for**  $\beta^{(i)} = \frac{{}^{(i)} \left\langle \downarrow, k^{(i)} \middle| \widehat{H_I} \middle| k^{(i)}, \uparrow \right\rangle^{(i)}}{E_1 - E_2}$ 

Apply non-degenerate perturbation theory to construct "gravitationally dressed'' states from  $|k,\uparrow\rangle^{(i)}, |k,\downarrow\rangle^{(i)}, i=1,2$  $\left|k^{(i)},\downarrow\right\rangle_{OG}^{(i)} = \left|k^{(i)},\downarrow\right\rangle^{(i)} + \left|k^{(i)},\uparrow\right\rangle^{(i)}\alpha^{(i)} \qquad \alpha^{(i)} = \frac{\left(i\right)\left\langle\uparrow,k^{(i)}\right|\widehat{H_{I}}\left|k^{(i)},\downarrow\right\rangle^{(i)}}{E_{2}-E_{1}}$  $\widehat{H}_{I} = -(r_{1}\sigma_{1} + r_{2}\sigma_{2})\widehat{k} \qquad \langle r_{\mu}\rangle = 0, \ \langle r_{\mu}r_{\nu}\rangle = \Delta_{\mu}\delta_{\mu\nu}$  $|k,\uparrow\rangle_{QG}^{(1)}|-k,\downarrow\rangle_{QG}^{(2)}-|k,\downarrow\rangle_{QG}^{(1)}|-k,\uparrow\rangle_{QG}^{(2)} = |k,\uparrow\rangle_{QG}^{(1)}|-k,\downarrow\rangle_{QG}^{(1)}|-k,\uparrow\rangle_{QG}^{(2)} = |k,\downarrow\rangle_{QG}^{(1)}|-k,\uparrow\rangle_{QG}^{(2)} = |k,\downarrow\rangle_{QG}^{(1)}|-k,\uparrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)} = |k,\downarrow\rangle_{QG}^{(1)}|-k,\uparrow\rangle_{QG}^{(2)} = |k,\downarrow\rangle_{QG}^{(1)}|-k,\uparrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG}^{(2)}|-k,\downarrow\rangle_{QG$  $+ |k,\downarrow\rangle^{(1)} |-k,\downarrow\rangle^{(2)} (\beta^{(1)} - \beta^{(2)}) + |k,\uparrow\rangle^{(1)} |-k,\uparrow\rangle^{(2)} (\alpha^{(2)} - \alpha^{(1)})$  $+\beta^{(1)}\alpha^{(2)}|k,\downarrow\rangle^{(1)}|-k,\uparrow\rangle^{(2)}-\alpha^{(1)}\beta^{(2)}|k,\uparrow\rangle^{(1)}|-k,\downarrow\rangle^{(2)}$ 

Apply non-degenerate perturbation theory to construct "gravitationally dressed'' states from  $|k,\uparrow\rangle^{(i)}, |k,\downarrow\rangle^{(i)}, i=1,2$  $\left|k^{(i)},\downarrow\right\rangle_{OG}^{(i)} = \left|k^{(i)},\downarrow\right\rangle^{(i)} + \left|k^{(i)},\uparrow\right\rangle^{(i)}\alpha^{(i)} \qquad \alpha^{(i)} = \frac{\left(i\right)\left\langle\uparrow,k^{(i)}\right|\widehat{H_{I}}\left|k^{(i)},\downarrow\right\rangle^{(i)}}{E_{2}-E_{1}}$  $\widehat{H}_{I} = -\left(r_{1}\sigma_{1} + r_{2}\sigma_{2}\right)\widehat{k} \qquad \langle r_{\mu}\rangle = 0, \ \langle r_{\mu}r_{\nu}\rangle = \Delta_{\mu}\delta_{\mu\nu}$  $\begin{array}{c} |k,\uparrow\rangle_{QG}^{(1)} |-k,\downarrow\rangle_{QG}^{(2)} - |k,\downarrow\rangle_{QG}^{(1)} |-k,\uparrow\rangle_{QG}^{(2)} = \\ |k,\uparrow\rangle_{QG}^{(1)} |-k,\downarrow\rangle_{QG}^{(2)} - |k,\downarrow\rangle_{QG}^{(1)} |-k,\uparrow\rangle_{QG}^{(2)} \end{array}$  $|-k,\downarrow\rangle^{(2)} \left(\beta^{(1)} - \beta^{(2)}\right) + \langle k,\uparrow\rangle^{(1)} |-k,\uparrow\rangle^{(2)} \left(\alpha^{(2)} - \alpha^{(1)}\right)$  $-\alpha^{(1)}\beta^{(2)}$