

# Gamma-ray identification with Imaging Atmospheric Cherenkov Telescopes

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## Imaging Atmospheric Cherenkov Telescopes

TeV gamma-ray ideal to probe the most energetic Universe

Supernova

remnants

Pulsar wind

nebulae

Neutrino

counterparts

Active Galactic Nuclei

Gamma-

**Ray Bursts** 

GW counterparts

Dark matter

 $\gamma$ -ray enters the atmosphere

Electromagnetic cascade

10 nanosecond snapshot

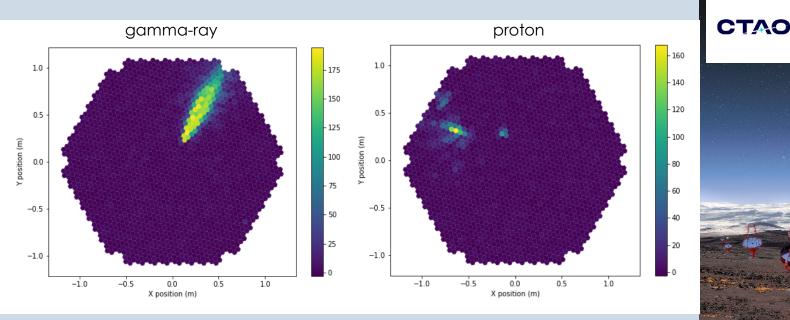
0.1 km<sup>2</sup> "light pool", a few photons per m<sup>2</sup>.

Primary

### **Particle identification**

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- Gamma rays and cosmic rays produce particle showers in atmosphere which emit Cherenkov light
- Shower images detected by fast high-resolution cameras
- ML algorithms used for the particle identification and the measurement of direction and energy of the primary particle





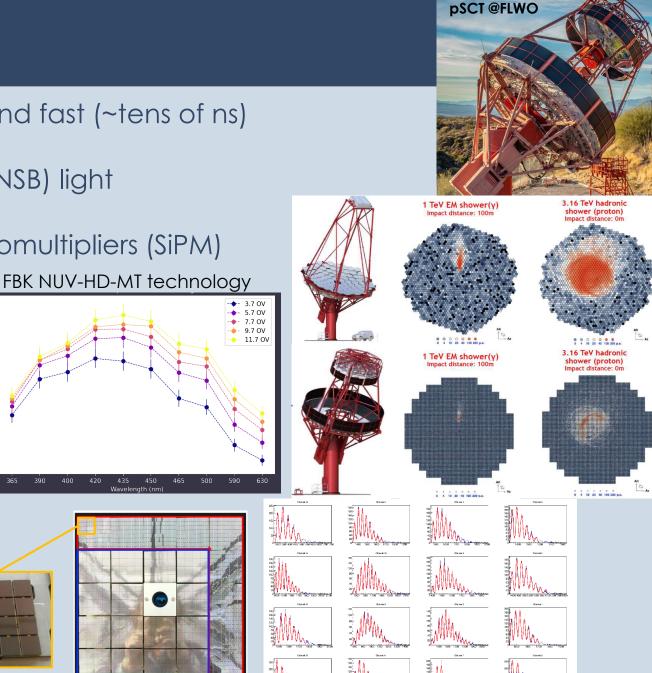
https://www.cta-observatory.org

### **IACT** camera

- Need to detect faint (down to few p.e.) and fast (~tens of ns) Cherekov light
- Need to deal with night sky background (NSB) light
- Photon detectors: Photomultiplier Tubes (PMT) → Silicon Photomultipliers (SiPM)
- Pros:
  - o Single p.e. resolution
  - NSB tolerant  $\rightarrow$  Operable under full moon
  - o High PDE (> 50% peak)
  - Small pixels  $\rightarrow$  easy to make arrays
  - Low bias voltage (<100V)</li>
- Cons:
  - High sensitivity to NSB in > 550 nm range
  - o Correlated noise
  - high dark count rate → usually below the NSB rate

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Ambrosi+2022 https://doi.org/10.1016/j.nima.2022.167359

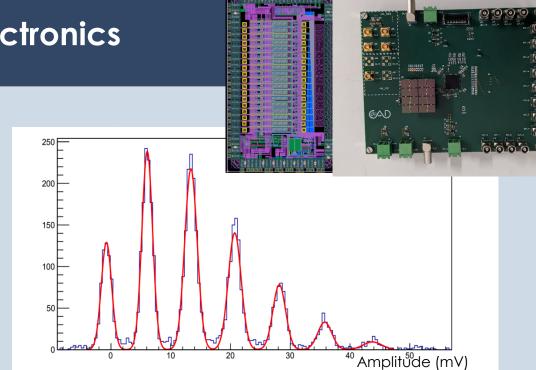


### Fast and single p.e. resolution frontend electronics

#### SMART: a SiPM Multichannel Asic for high Resolution Cherenkov Telescopes

#### Features:

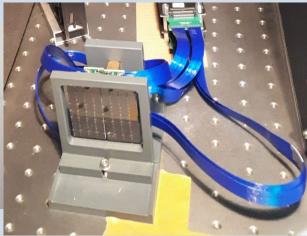
- 3.3 V & 1.2 V power supply
- SiPM fine bias adjust range: 1.7 V
- Channel current consumption 3.6 mA
- Fast-path output dynamic range: up to 1.7 V on 50 Ohm load (AC coupled)
- Slow-path input dynamic range: up to 2 mA of mean SiPM current (1 pC/pe at 2 GHz)
- Programmable gain (R), bandwidth (C) and tail suppression (PZ)
  - Gain is adjustable for each channel to be combined with SiPM bias fine adjustment to improve gain equalization



#### Designed by F. Licciulli & G. De Robertis at the Electronics CAD INFN Bari

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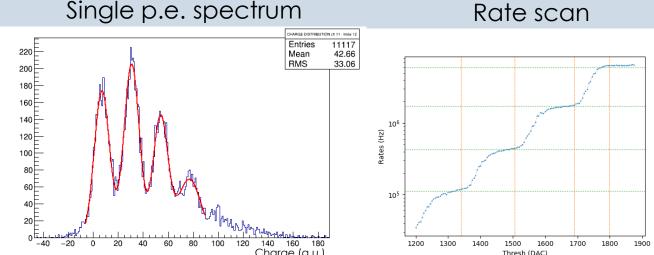


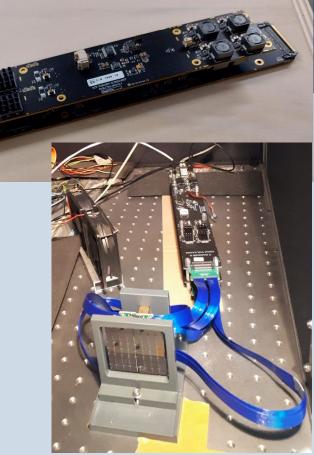


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#### **Readout electronics**

- Readout electronics to digitize fast signals and generate trigger signals at pixel level -> TARGET ASICs
  - o CTC ASIC: 16-channel 1GSa/s digitizer
    - Analog buffer with 16k cells per channel  $\rightarrow$  16 us storage depth
  - o CT5TEA ASIC: 16-channel trigger ASIC
    - Channels are summed in groups of 4 to obtain 4 trigger pixels
      per ASIC
      Single p.e. spectrum
      Perte segn





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