



Status of the LANTERN project

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Optical Calibration

Optical Calibration

- Based on **photon counting statistics**
- **Low energy calibration** (from ~10 eV)
- 400 nm photons (~3 eV)

Based on Poisson Distribution

$$\mu_\gamma = r \cdot N_\gamma$$

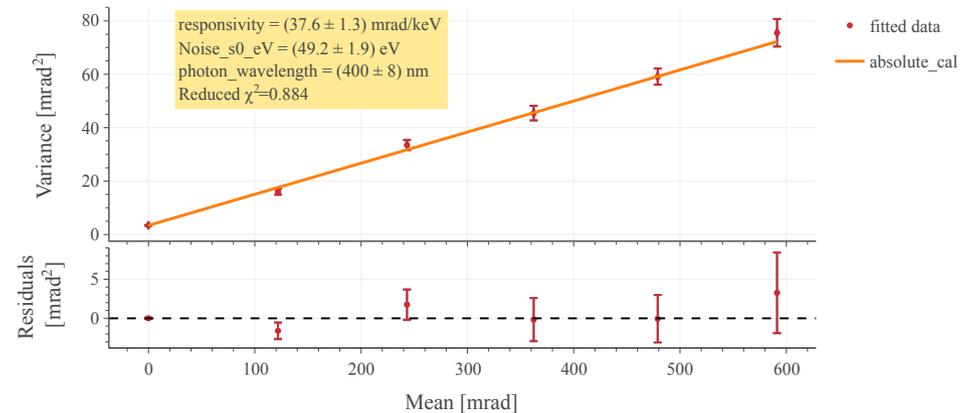
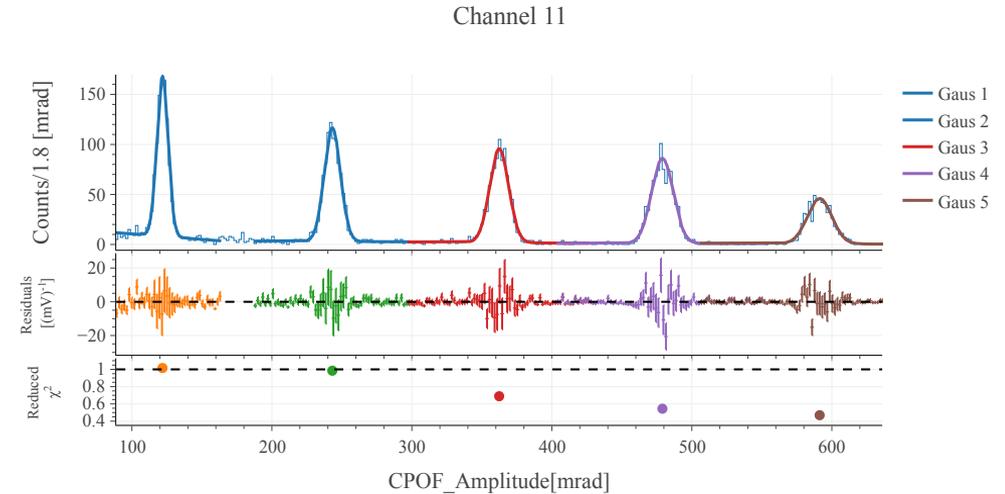
$$\sigma_\gamma = \sqrt{N_\gamma} \cdot r = \sqrt{\mu_\gamma \cdot r}$$

Fitting Function:

$$\sigma^2 = \sigma_0^2 + r \cdot \mu_\gamma$$

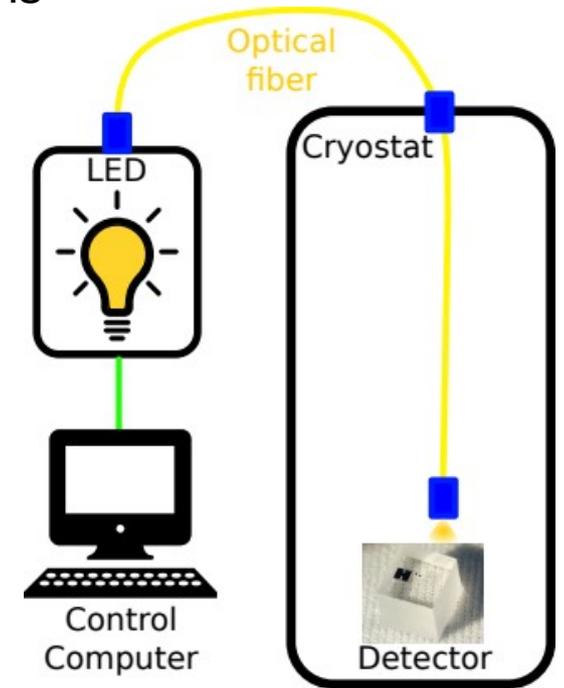
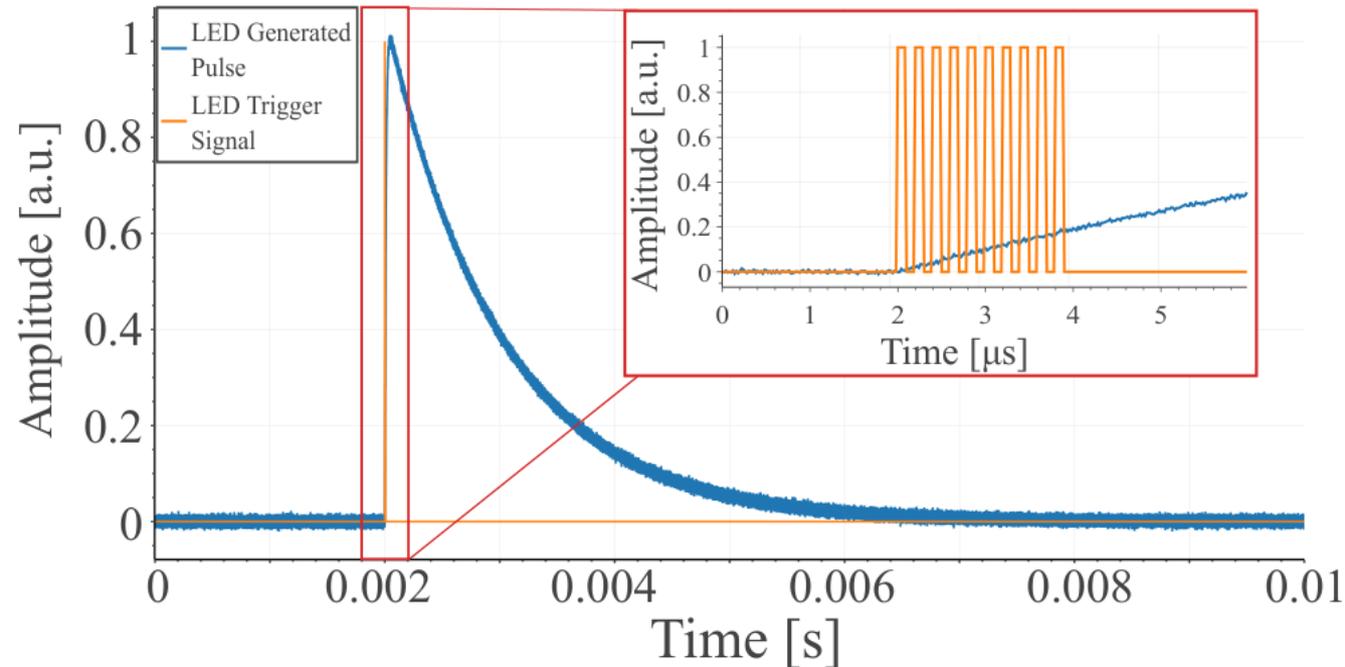
σ_0 = Detector Resolution

$$r : \frac{\text{mrad}}{\text{photon}} \rightarrow k = \frac{\text{photon energy}}{r} \left[\frac{\text{eV}}{\text{mrad}} \right]$$



Setup Concept

- **Fast light flashes (cycles)** for linear scaling of deposited energy
- Use monochromatic **LED (UV-Vis wavelength, photon energy of O(5 eV))**
- **Room temperature electronics**
- **No increase in particle background** → usable during physics runs



Non-Linearity Corrections

Linearize mean:

Number of cycles $\mu_\gamma = a \cdot n_c (1 + b \cdot a \cdot n_c)$

Non-linearity Parameter b

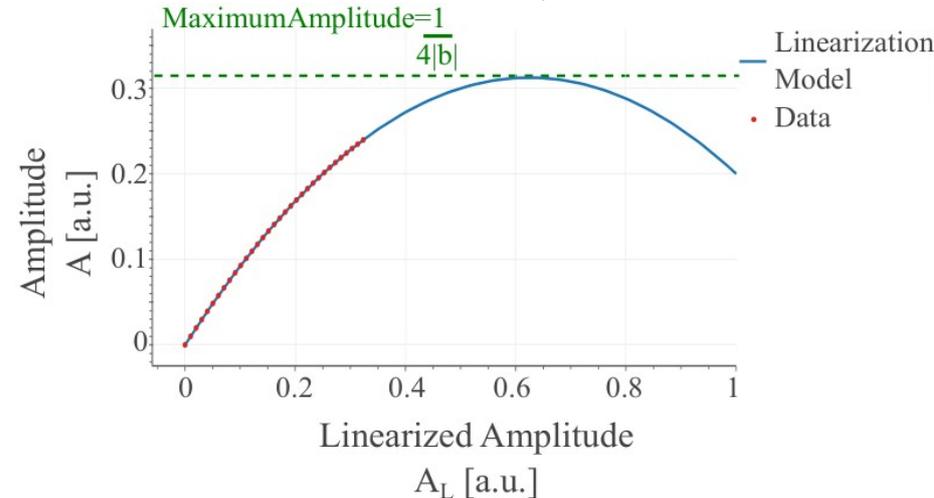
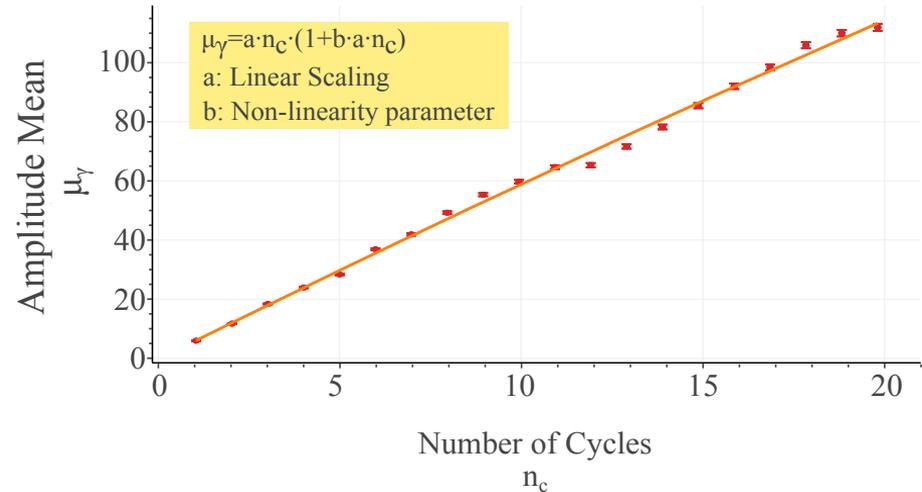
Average Linear Response $\mu_\gamma^L := \frac{\mu_\gamma}{1 + b \cdot a \cdot n_c} = a \cdot n_c$

How to convert event amplitude to energy

$$A_L(A) = \frac{-1 + \sqrt{1 + 4bA}}{2b}$$

$$E = k_L \cdot A_L \quad k_L = \frac{\text{photon energy}}{r_L} \left[\frac{\text{eV}}{\text{mrad}} \right]$$

Attention: Valid only in LED range

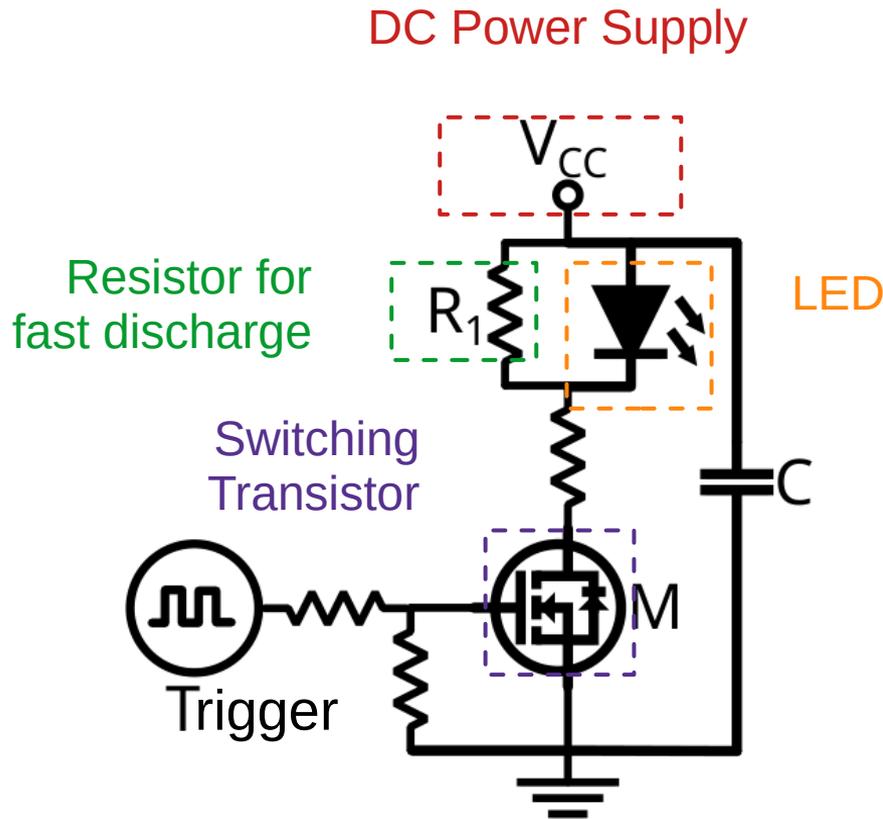


LANTERN: Concept and Tests

The LANTERN Project

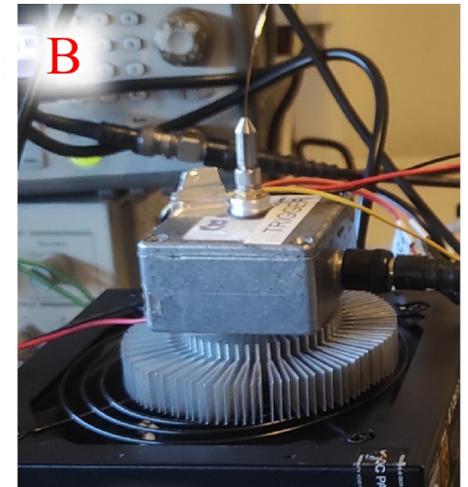
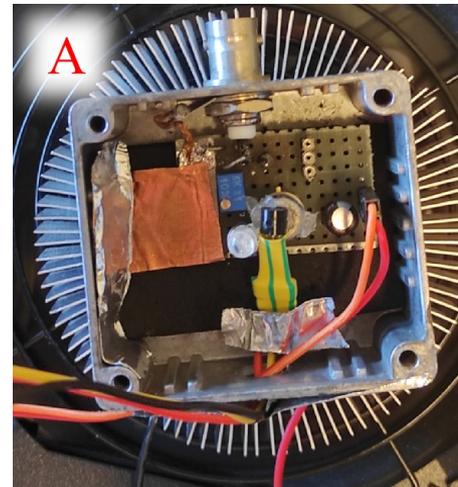
LANTERN: LED Array for Non-intrusive Tuning of the Energy Range with Nimbleness

* funded from Sapienza University PhD research-starting projects



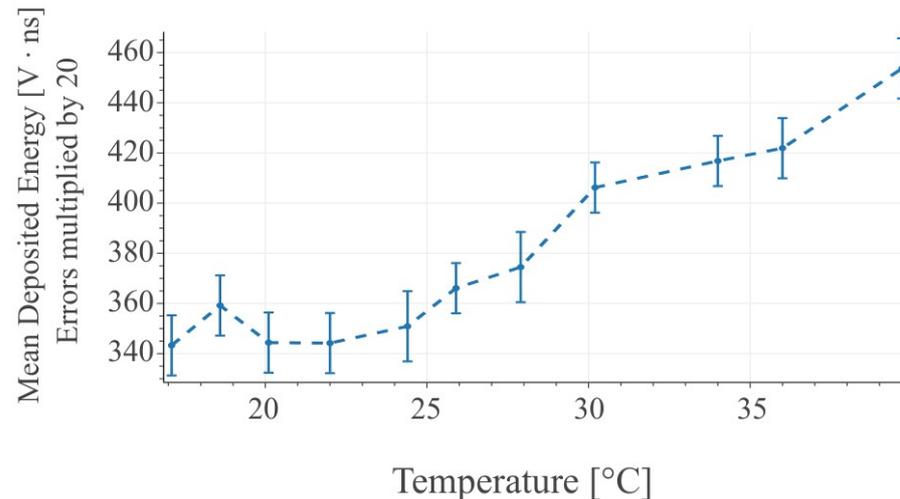
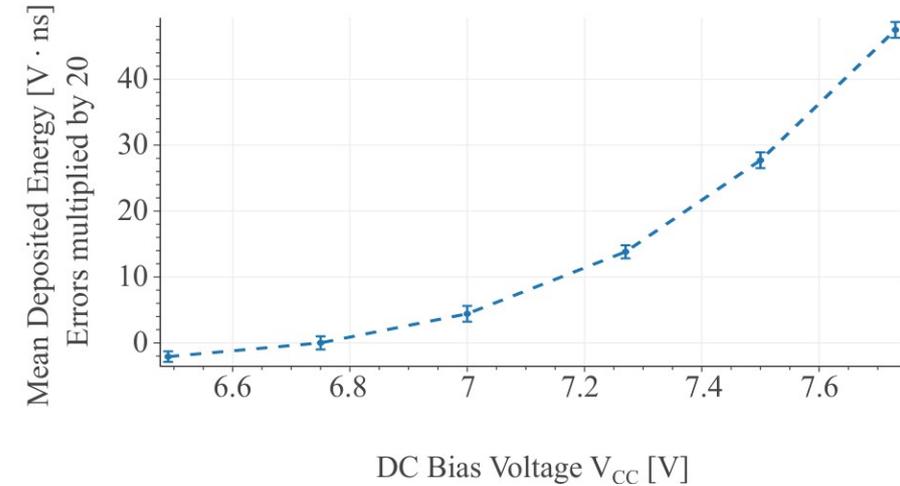
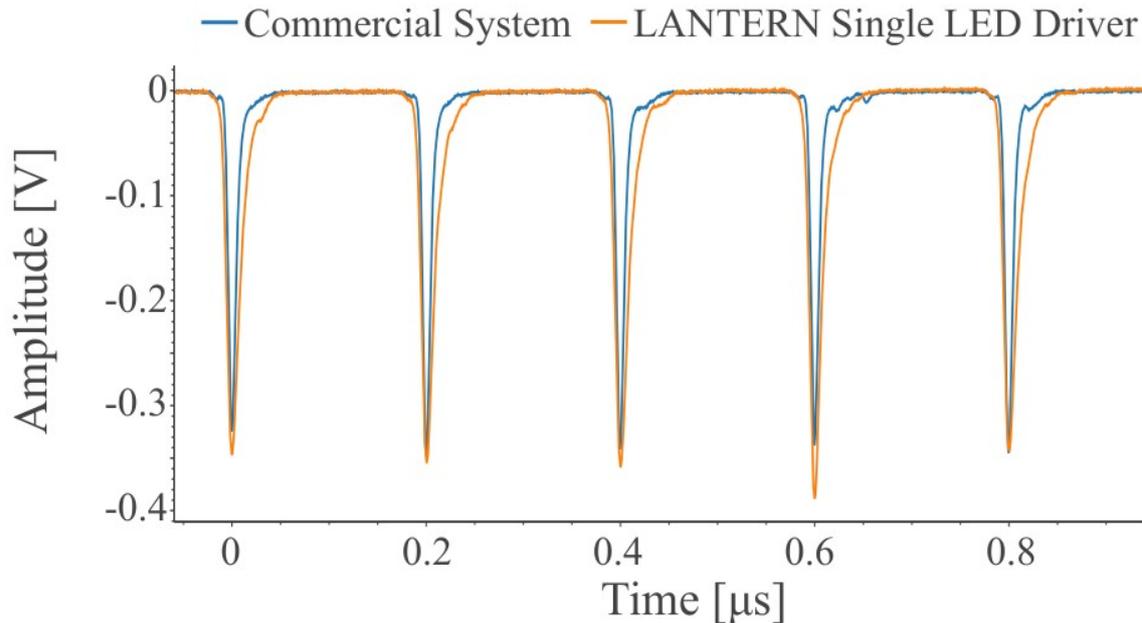
System Goals

- Scalability
- Wide dynamic range
- Reliability
- Stability



PMT Tests

- **Matched light emission** as commercial system
- Studied scaling with **DC bias voltage V_{CC}**
- Studied scaling with operation **temperature**

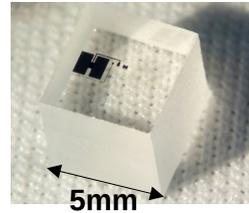


Validation of LANTERN on Cryodetectors

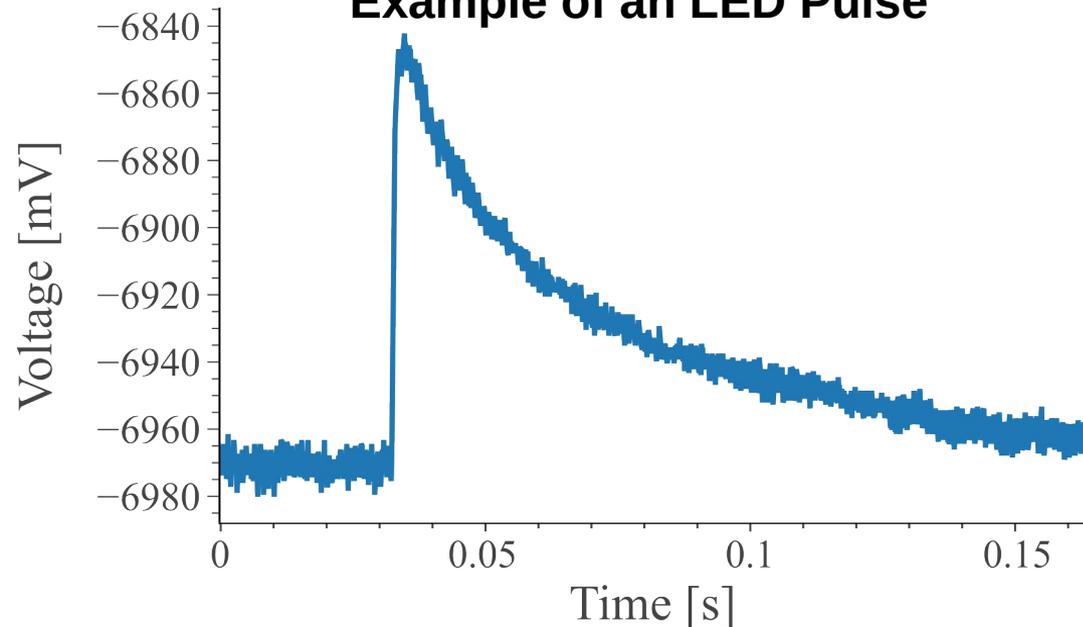
- **No pulse shape deformation** at high energies **observed**
- **Compatible calibration constant** between all tests



CaWO₄ and Al₂O₃ crystals
+
Transition Edge Sensors
(TES)



Example of an LED Pulse



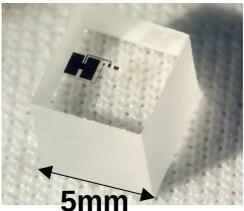
	Calibration Constant [eV/mV]	Mean Rise Time [ms]
Commercial	16.9 ± 0.6	0.94 ± 0.09
LANTERN	17.5 ± 0.7	0.92 ± 0.14

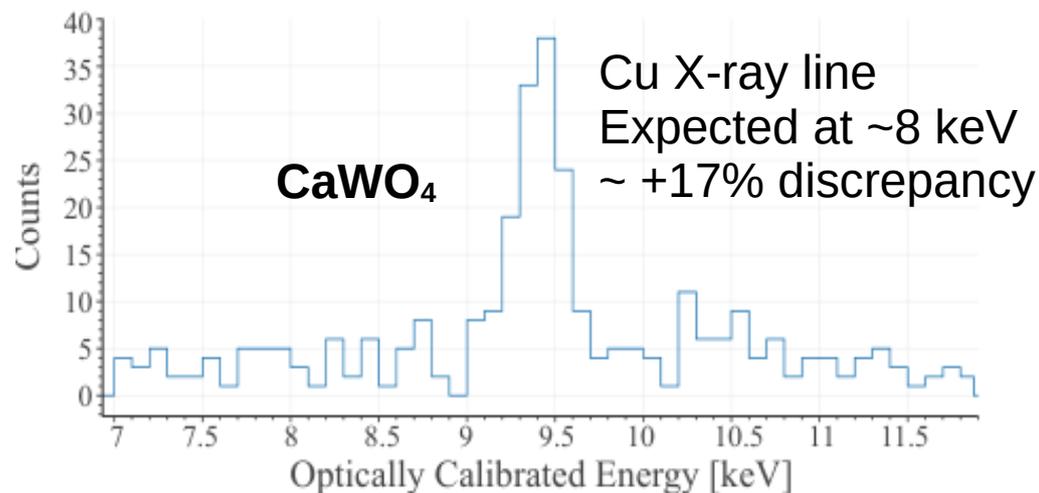
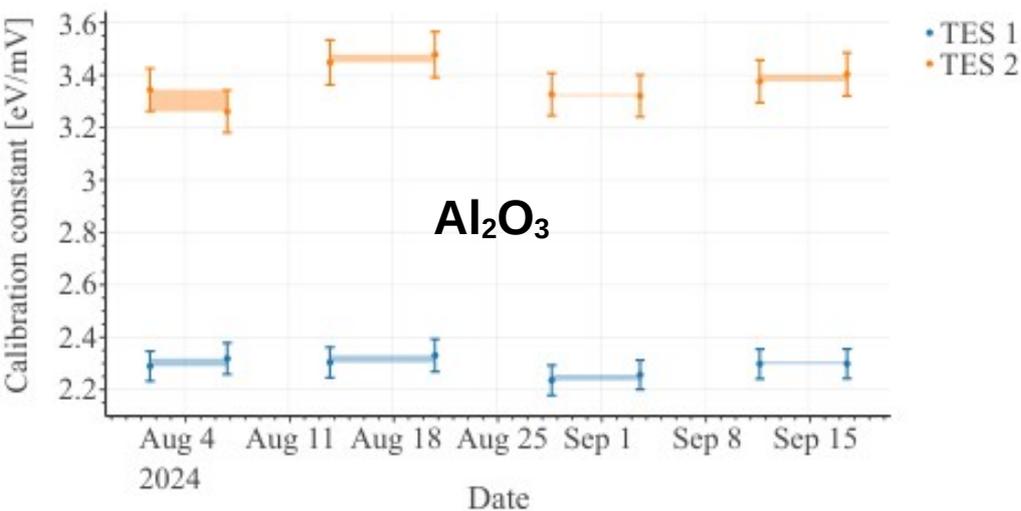
Cryodetector Tests on NUCLEUS detectors

- Performed with **LANTERN** single LED using **250 nm**
- **Stable calibration** on Al_2O_3 and CaWO_4
- Al_2O_3 dynamic range: 30 eV – 4 keV
- CaWO_4 dynamic range: 45 eV – 12 keV



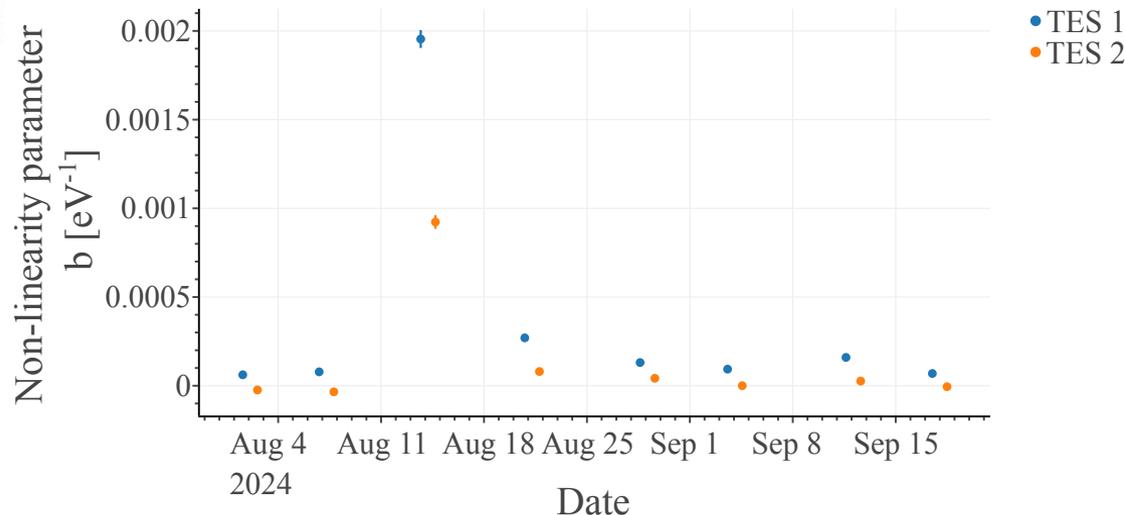
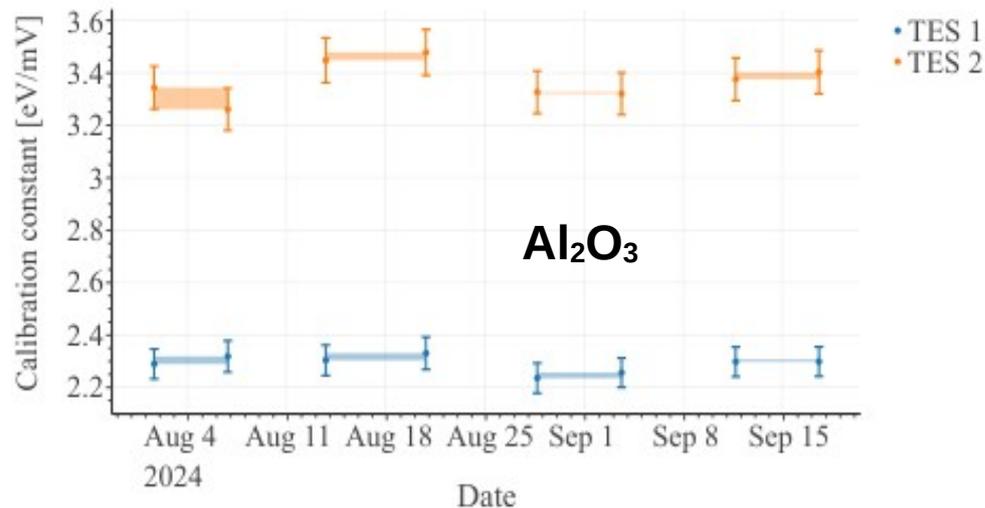
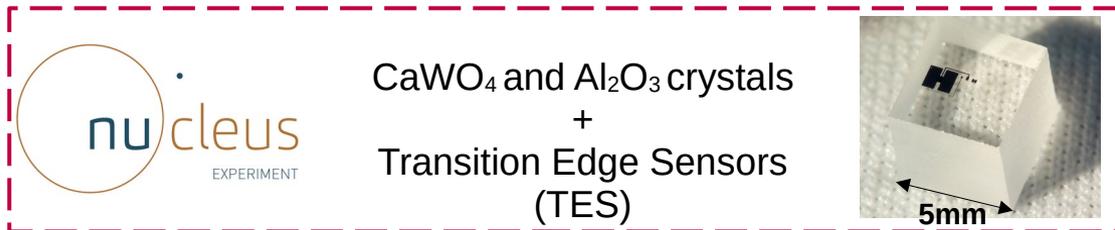
CaWO₄ and Al₂O₃ crystals
+
Transition Edge Sensors (TES)





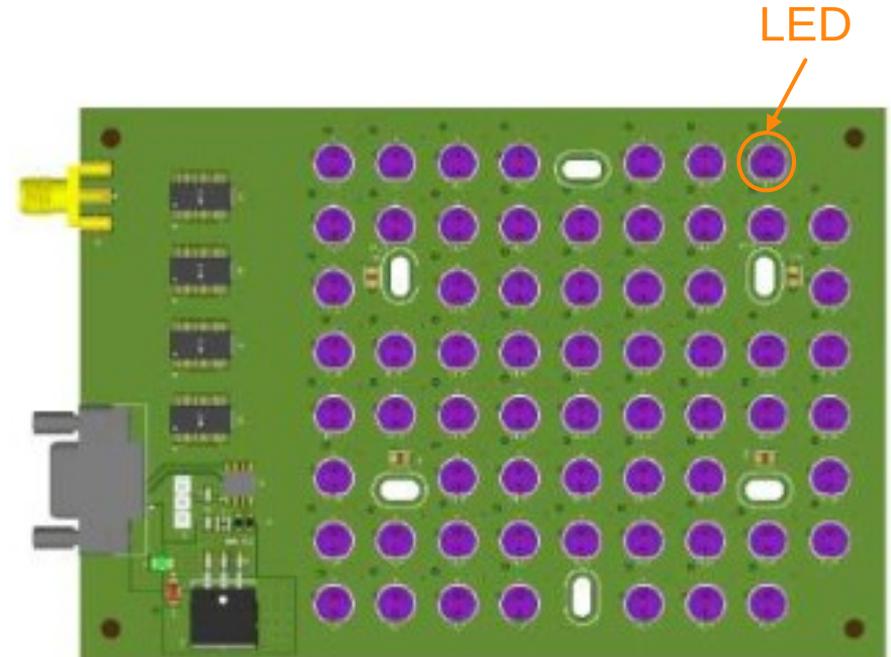
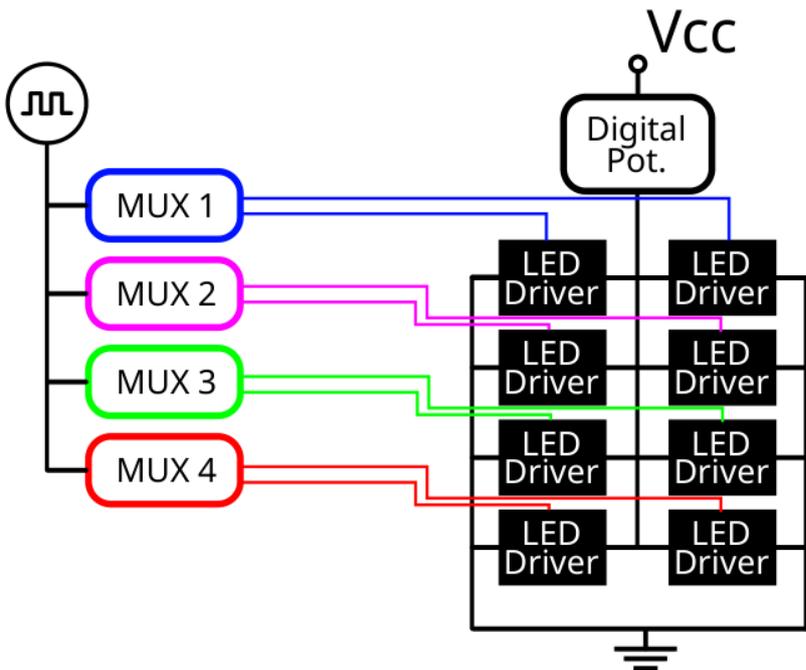
Linearization Issue

- **Damage on LANTERN made it non-linear** (does not affect calibration)
- Damage due to **over biasing**
- **Back to nominal values after fixing**
- **Calibration not affected!**

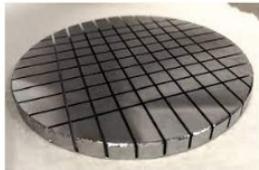
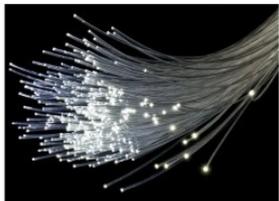


Multi-pixel LANTERN Electronics

- **Scaling concept:** Multiplexed LED drivers
- Targeted **64 LEDs**
- **On-chip voltage regulation**
- **Designed for in vacuum operation** → mounting inside cryostat at 300K stage.

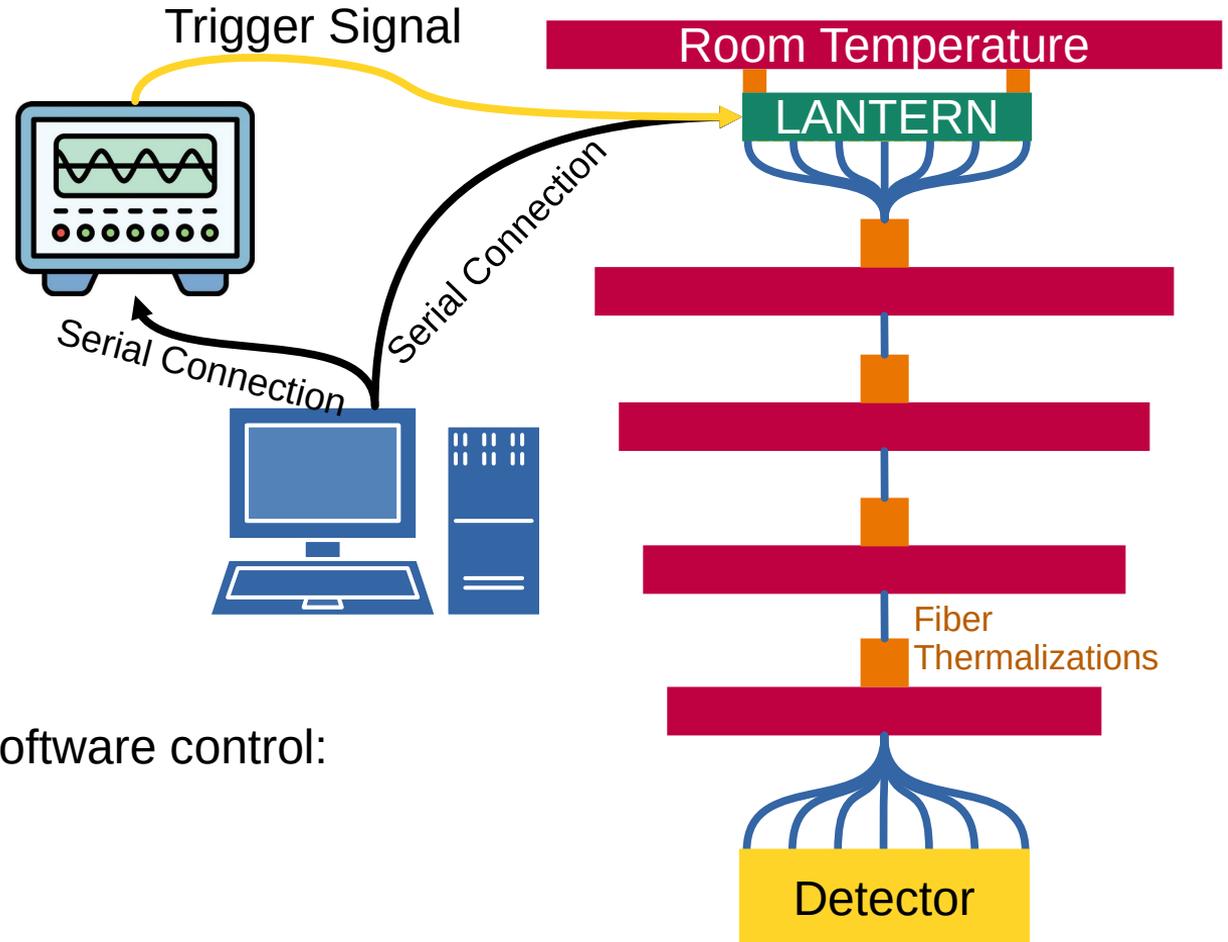


Mounting Concept



External electronics with a software control:

- LANTERN
- Signal Generator Control
- x64 Optical Channels

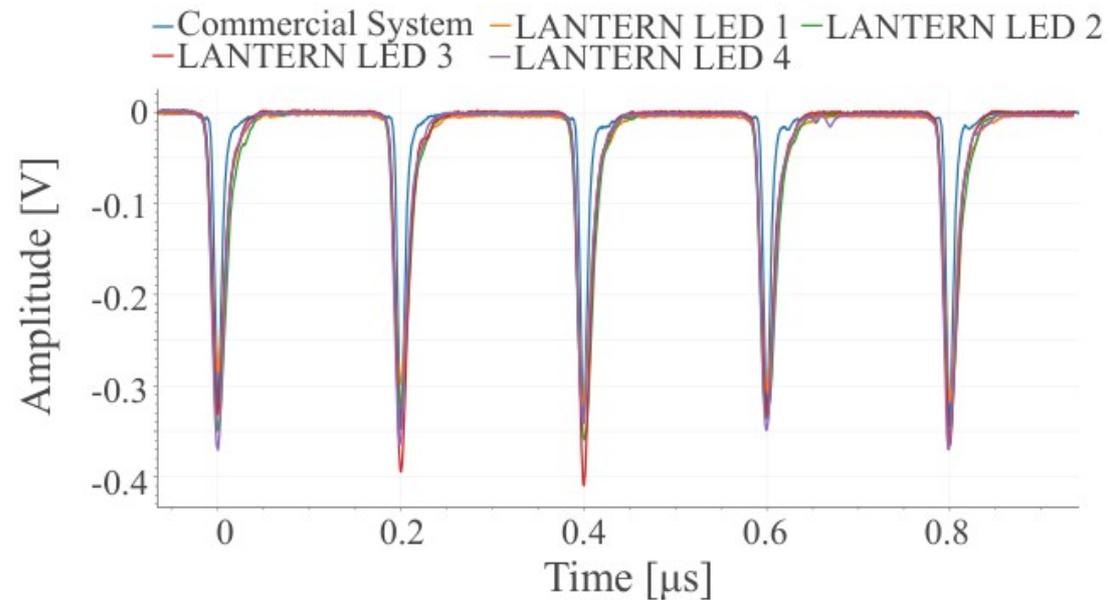


Where we left off: PMT Tests



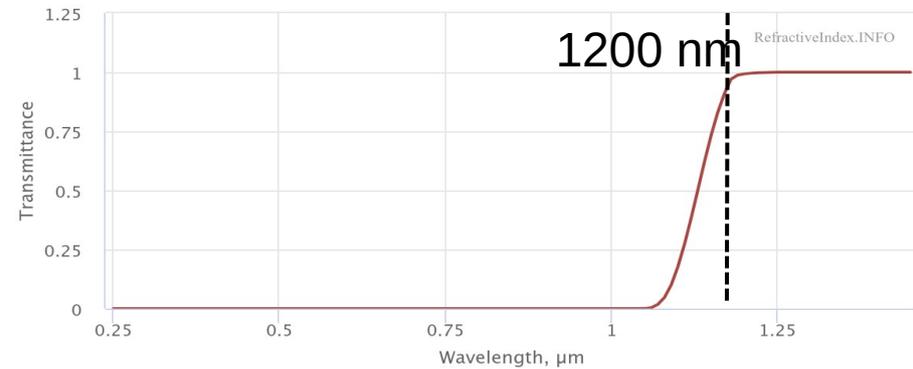
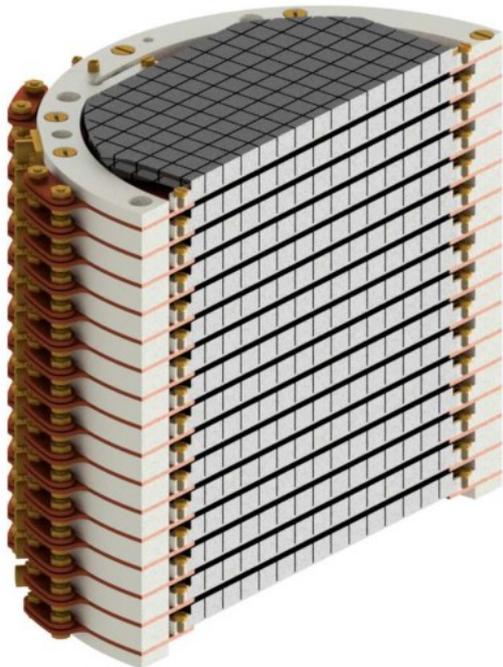
- Higher Luminosity
- **Dark Counting tests** (20 min duration): **No increase** observed
- **Matched light emission** as commercial system
- Studied scaling with **DC bias voltage V_{cc}**
- Studied scaling with operation **temperature**
- Tested in **vacuum operation**

Configuration	Counts
Dark	734
With Trigger - Enable False	731
Nearby LED shining	715
Far away LED shining	702

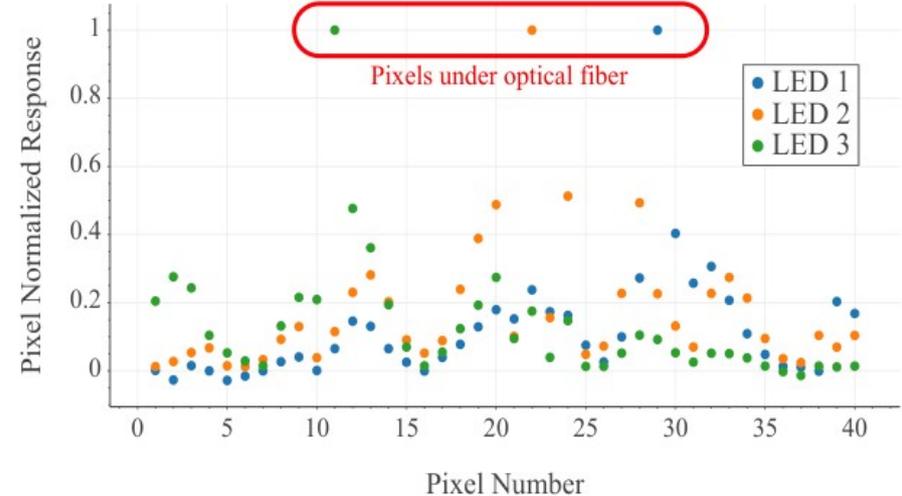


LANTERN for Pixel Identification

- 5mm of Si are semitransparent to IR
- Use them to penetrate the stack and perform **pixel ID**
- **Worked with 2 wafers** (needs further tests)



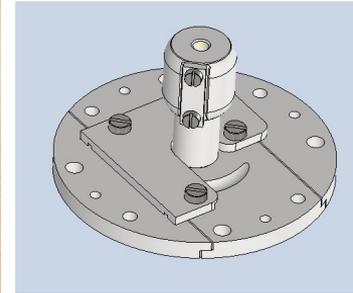
Pixel identification using IR light



Current Status



We have all the ingredients...
but we need to start cooking!!



Conclusion

The Road so Far ...

- **Promising tests** on both PMT and Cryodetectors
- LANTERN used **successfully in NUCLEUS** for physics run
- In NUCLEUS ~ **17% calibration discrepancy** X-rays (-10% on BULLKID see @Matteo Folcarelli Tomorrow)
- Produced **5 PCBs**, bought **fibers** and optical mountings, fiber **thermalization** elements are under production, **LEDs** are available



The Road Ahead ...

- **Mount LEDs on PCBs**
- Test the electrical and optical behavior of PCBs
- **Test thermalization of optical fibers !**
- Test in vacuum operation of PCBs (again)
- Develop control software (preliminary version already available)
- Deploy full system in Rome's cryostat





Thank you for
your
attention!!!

Variance Scaling with Energy

Optical Calibration

- Detector effects (position effects or others) can produce a quadratic scaling between variance and mean
- Fitting ω allows to take them into account in the calibration. The “r” parameter remains the one used for the calibration constant
- Saturation competes with this effect, so it is advised to perform the linearization first

Based on Poisson Distribution

$$\mu_\gamma = r \cdot N_\gamma$$

$$\sigma_\gamma = \sqrt{N_\gamma} \cdot r = \sqrt{\mu_\gamma \cdot r}$$

Fitting Function:

$$\sigma^2 = \sigma_0^2 + r \cdot \mu_\gamma + \omega \mu_\gamma^2$$

σ_0 = Detector Resolution

$$r : \frac{\text{mV}}{\text{photon}} \longrightarrow k = \frac{\text{photon energy}}{r} \left[\frac{\text{eV}}{\text{mV}} \right]$$

LED Spectral Width

$$\begin{aligned}\sigma_{LED}^2 &= \sigma_{N_{ph}}^2 \langle \epsilon_{ph} \rangle^2 + \sigma_{\langle \epsilon_{ph} \rangle}^2 \langle N_{ph} \rangle^2 = \langle N_{ph} \rangle \langle \epsilon_{ph} \rangle^2 + \frac{\sigma_{\epsilon_{ph}}^2}{\langle N_{ph} \rangle} \langle N_{ph} \rangle^2 = \\ &= \langle N_{ph} \rangle \langle \epsilon_{ph} \rangle^2 \left(1 + \frac{\sigma_{\epsilon_{ph}}^2}{\langle \epsilon_{ph} \rangle^2} \right) = E_{dep} \langle \epsilon_{ph} \rangle \left(1 + \frac{\sigma_{\epsilon_{ph}}^2}{\langle \epsilon_{ph} \rangle^2} \right) \approx E_{dep} \langle \epsilon_{ph} \rangle (1 + 10^{-4})\end{aligned}$$

