

## How to unravel the youngest Universe: Introduction to the PTOLEMY project

**Chris Tully (Princeton University** 

4 DECEMBER 2024 GENOA, ITALY

Research supported

byohn Templeton ndation

## ERA OF AMAZING NEW TELESCOPES

#### James Webb Space Telescope

Source: Nasa



BBC

Event Horizon Telescope Supermassive Black Hole at Center of Milky Way Galaxy

https://www.space.com /james-webb-space-tel escope-ancient-black-h ole-quasar





At 330,000 years, the Universe is smooth to 10 parts per million

## **Cosmic Neutrino Background**

NOBEL PRIZE IN PHYSICS 2019



Dicke, Peebles<sup>\*</sup>, Roll, Wilk<mark>inson (1965)</mark>

Number density:  $n_{i} = 112/cm^{3}$ **Temperature:** T.~ 1.95K Time of decoupling: t ~ 1 second ~50% of the Total Energy Density of the Universe @ 1 sec neutron/proton ratio @start of nucleosynthesis

<sup>4</sup>He <sup>2</sup>H (<sup>3</sup>H e)

## Looking Back in Time with Photons



**Emission Time** -13.8x10<sup>9</sup> years  $\sim$  -4x10<sup>9</sup> years -200x10<sup>6</sup> years  $\sim$  -2x10<sup>6</sup> years All of this light arrives at the same time (t=0)

## Neutrino Masses from Oscillations



#### Theory developed by Bruno Pontecorvo



3 mass eigenstates X 3 flavors (electron, muon, tau)

## Neutrino Masses from Oscillations



#### Theory developed by Bruno Pontecorvo



3 mass eigenstates X 3 flavors (electron, muon, tau)

## Neutrino Mass Oscillation Observatory

KM3NeT/ORCA (Oscillation Research with Cosmics in the Abyss) determination of the neutrino mass hierarchy (E<sub>v</sub> ~ MeV - GeV) low energy neutrinos Depth -2500 m -offshore Toulon (France)



## Neutrino Mass Oscillation Observatory

KM3NeT/ORCA (Oscillation Research with Cosmics in the Abyss) determination of the neutrino mass hierarchy (E<sub>v</sub> ~ MeV - GeV) low energy neutrinos Depth -2500 m -offshore Toulon (France)



## Massive Neutrino Timeline



## **Einstein rings** Predicted by Einstein in 1936







#### IDEA OF ENRICO FERMI 90 year



Fermi, E. Versuch einer Theorie der β-Strahlen**a. Anderse Beschlerg!** (1934).

https://doi.org/10.1007/BF01351864



The neutrino masses are so tiny, their effects are smaller than atomic transitions in normal materials. (There is a reason that there are no units on this plot.)

## PTOLEMY: 2D MATERIAL - GRAPHENE



## PTOLEMY: 2D MATERIAL - GRAPHENE



#### MICRO-CALORIMETER



Based on the expertise of the INRiM an important results have been achieved on electron measurement with TES.

Key elements of the measurements: performing TES and new e-source based on nanostructures



**Design Goal (PTOLEMY):**  $\Delta E_{FWHM} = 0.05 \text{ eV} @ 10 \text{ eV}$ 

0.02

0.00

-0.04

-0.06

-0.08

-0.10

-0.14

 $\tau_{1} = 137 \text{ ns}$ 

(mV)

translates to  $\Delta E \propto E^{\alpha} \ (\alpha \le 1/3)$  $\Delta E_{FWHM} = 0.022 \text{ eV} @ 0.8 \text{ eV}$ 

 $\tau_{eff} = 147 \text{ ns}$ 

 $2\gamma$ 

 $\Delta E = 0.123 \text{ eV}$ 

1.0

# RF MEASUREMENTS NON-DESTRUCTIVE ELECTRON TAG

filter Can we detect the (semi-relativistic) electron on its target way to the micro-calorimeter? 20

# RECENT PROJECT 8 TRITIUM RF MEASUREMENT



RF measurement background levels extremely low.

No events observed above endpoint, Setting upper limit on background rate

< 3x10<sup>-10</sup> /eV/s (90% CL)

Background Rate < I event per eV in 100 years!

21

#### ACHIEVED !! KF MEASUREMENTS **NON-DESTRUCTIVE ELECTRON TAG** filter Can we detect the (semi-relativistic) electron on its target way to the micro-calorimeter? R ×10<sup>-6</sup> time [us] 2.4 200 2.2 180 2 160 1.8 140 1.6 120 1.4 1.2 100 80 0.8 60 0.6 40 0.4

0.2

20

535.5

536

536.5

537

537.5

538

538.5

freq.[MHz]

22

## TARGET FABRICATION

23





# Best in the World!

Hydrogenation of nano-porous graphene (left and center) showing over 90% coverage per carbon atom through the increase of sp<sup>3</sup> bonding (blue on right) <u>DOI:</u> <u>10.1021/acs.nanolett.2c00162</u>

#### T-chamber R side view

Sapienza



Quadrupole Mass Spectrometer: SRS RGA 100

JINST 17 (2022) 05, P05021

 A new electromagnetic filter idea based on RF detection and dynamic F setting



JINST 17 (2022) 05, P05021

 A new electromagnetic filter idea based on RF detection and dynamic E setting VB RF B Ē ANTENNA <sup>3</sup>H first measurement of the energy via cyclotron RF emission (~  $10\mu s$ By A. Esposito

JINST 17 (2022) 05, P05021

 A new electromagnetic filter idea based on RF detection and dynamic E setting



JINST 17 (2022) 05, P05021

- A new electromagnetic filter idea based on RF
  - detection and dynamic Epsetting









Zero B field saddle point key feature of the field map

## The PTOLEMY Collaboration



## **The PTOLEMY Collaboration**



#### 14 May 2024, Pollica, Italy

## CONCLUSION

- PTOLEMY's goal is to eventually detect the cosmic neutrino background
- The detector prototype will be ready at LNGS by the next year
- Prototype baseline option is: T embedded on graphene; New concept EM filter; electron energy resolution measured in several steps (MCP/SDD).
  Ultimately operating TES with sub-eV energy resolution.