

# $^{30}\text{Si}(n,\gamma)$ & $^{64}\text{Ni}(n,\gamma)$ : Status of the Analysis

**Michele Spelta**



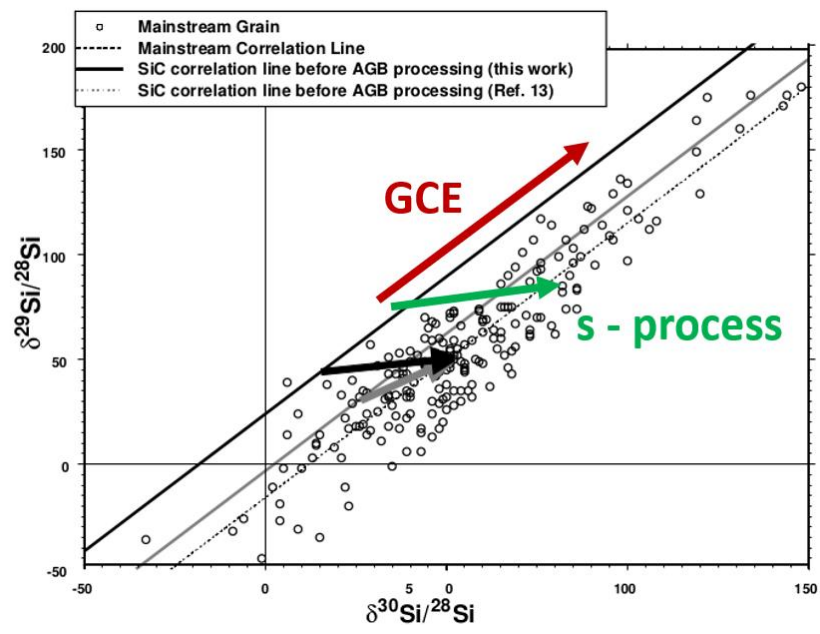
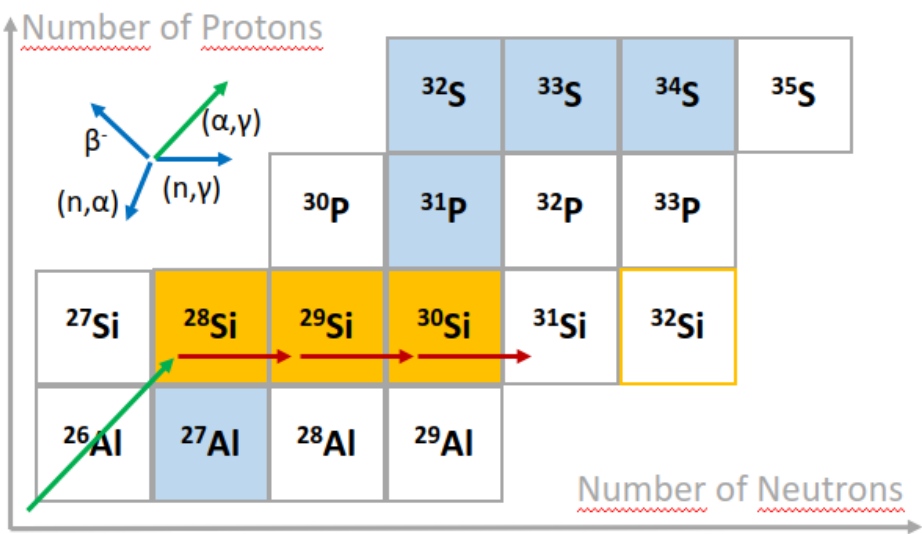
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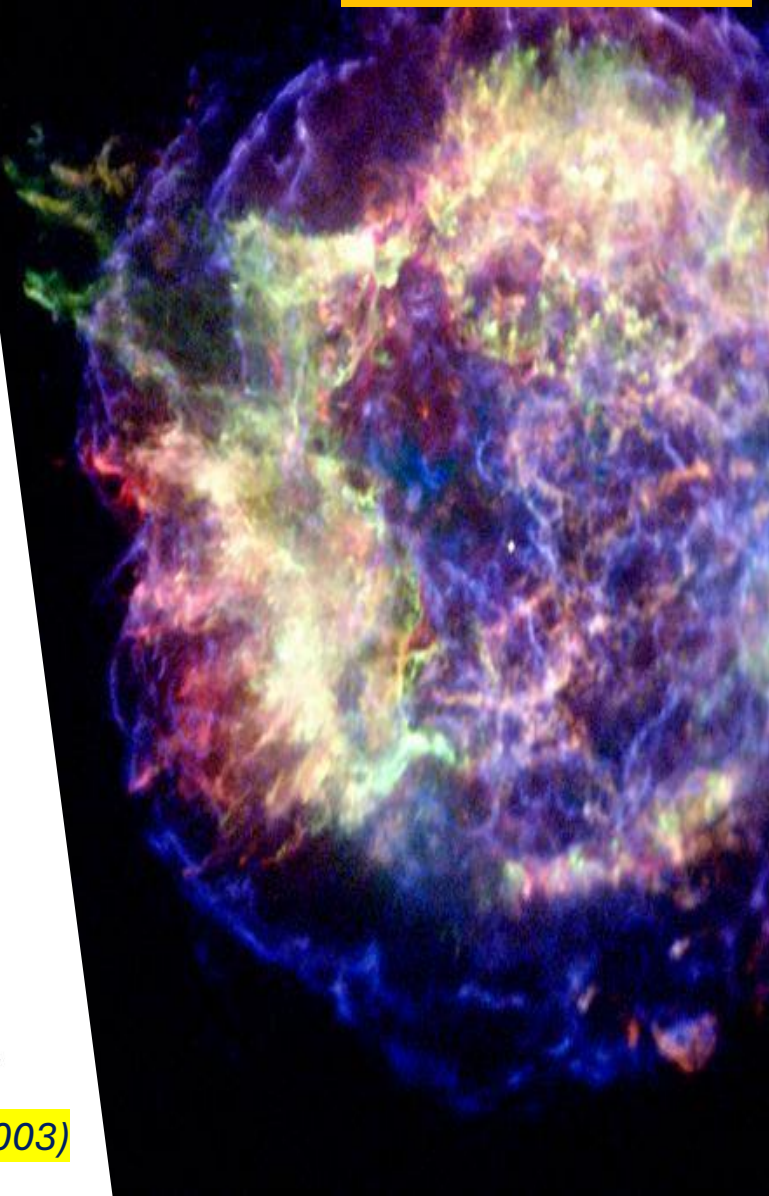
# $^{30}\text{Si}(n,\gamma)$ : Motivations

$^{30}\text{Si}(n,\gamma)$  cross section is important:

- to predict the abundance of Silicon isotopes produced in the convective carbon shell of **massive stars**
- to understand the Si isotopic abundances measured in **pre-solar SiC grains**, disentangling the contributions of **s-process** and **GCE**



K. Guber et al., Phys. Rev. C 67, 062802 (2003)



# $^{64}\text{Ni}(n,\gamma)$ : Motivations

RECAP

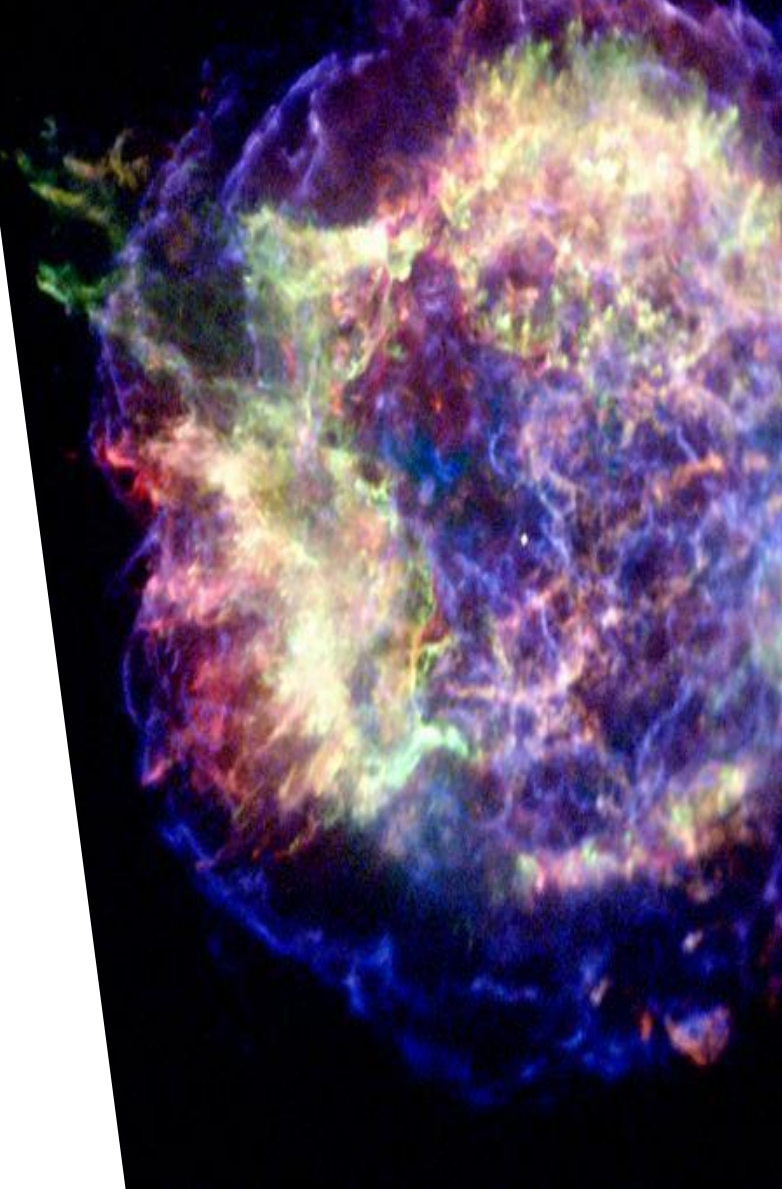
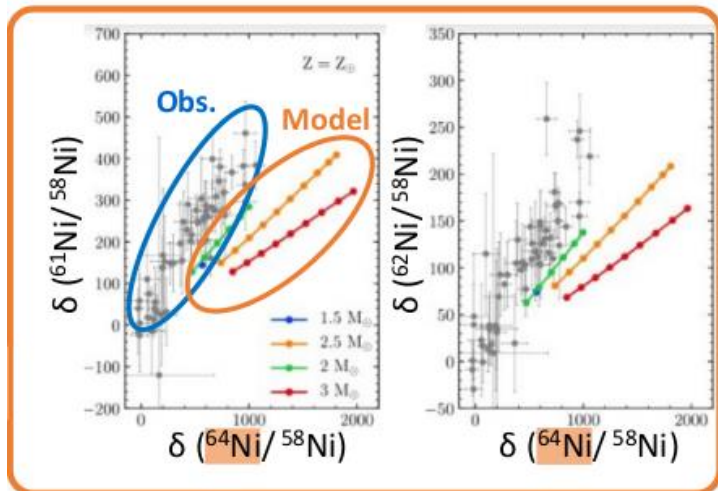
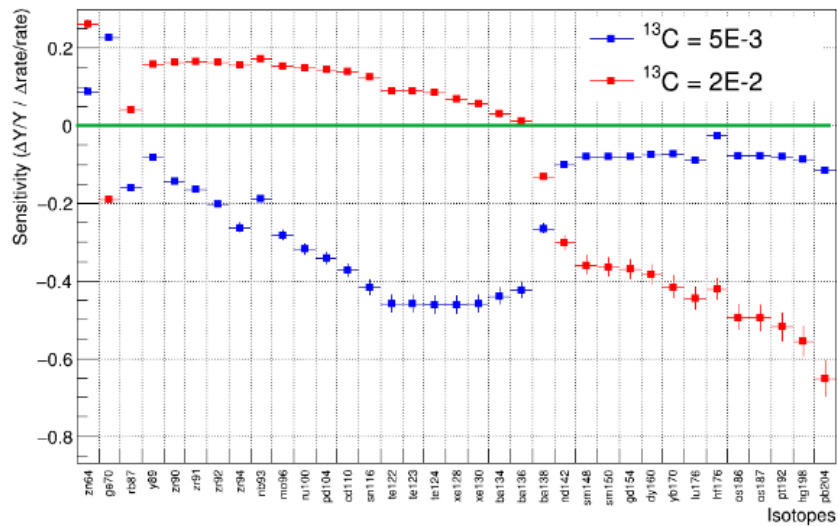
$^{64}\text{Ni}(n,\gamma)$  cross section is important:

- because, as a **seed of the s-process**, it affects the abundances of many isotopes synthesized in the process

*Cescutti et al., MNRAS 478, 4101 (2018)*

- to possibly explain the **discrepancy observed in SiC grains** between measured and predicted  $^{64}\text{Ni}$  isotopic abundances

*Vescovi et al., ApJ Lett 897, 25 (2020)*

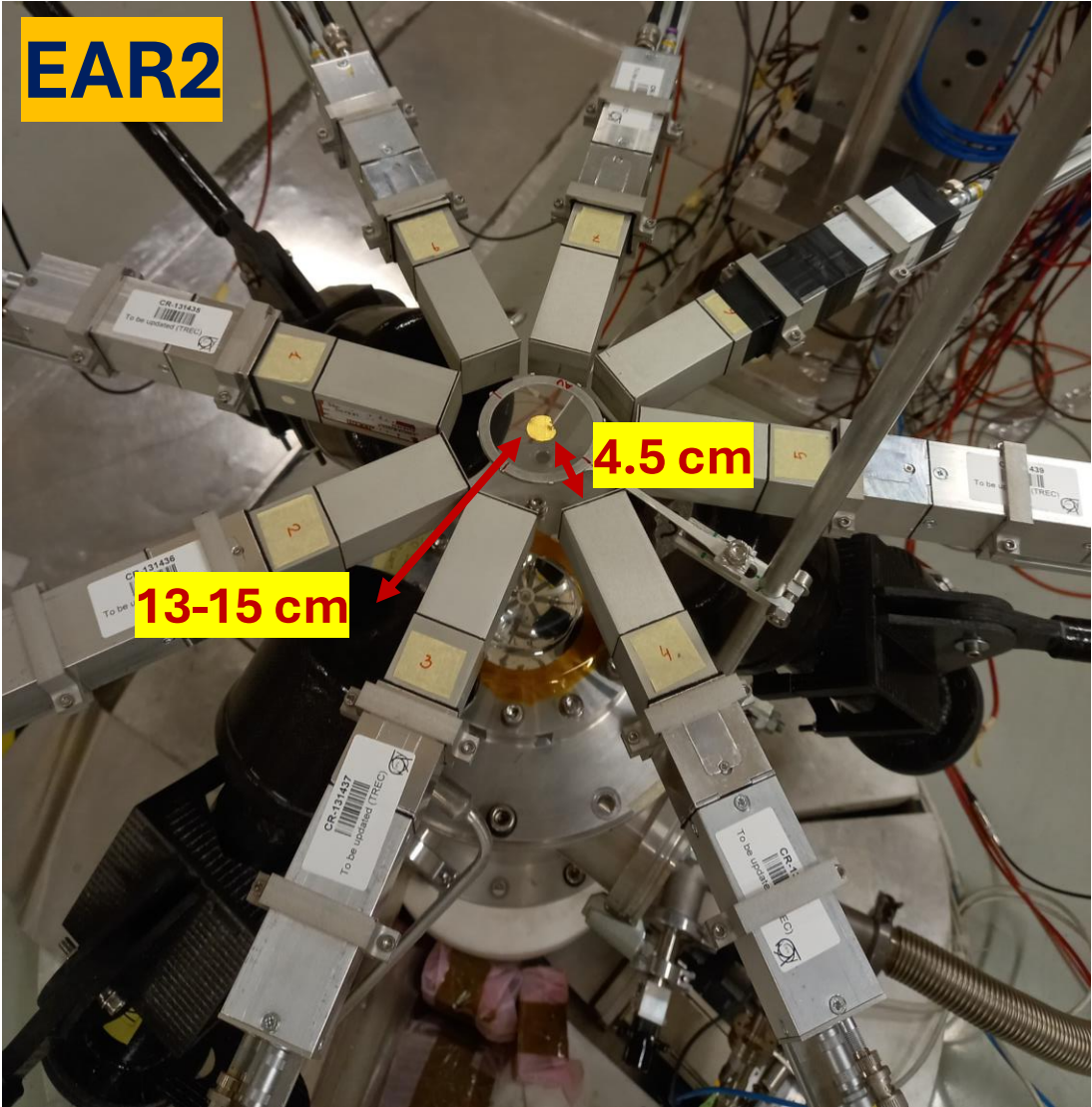
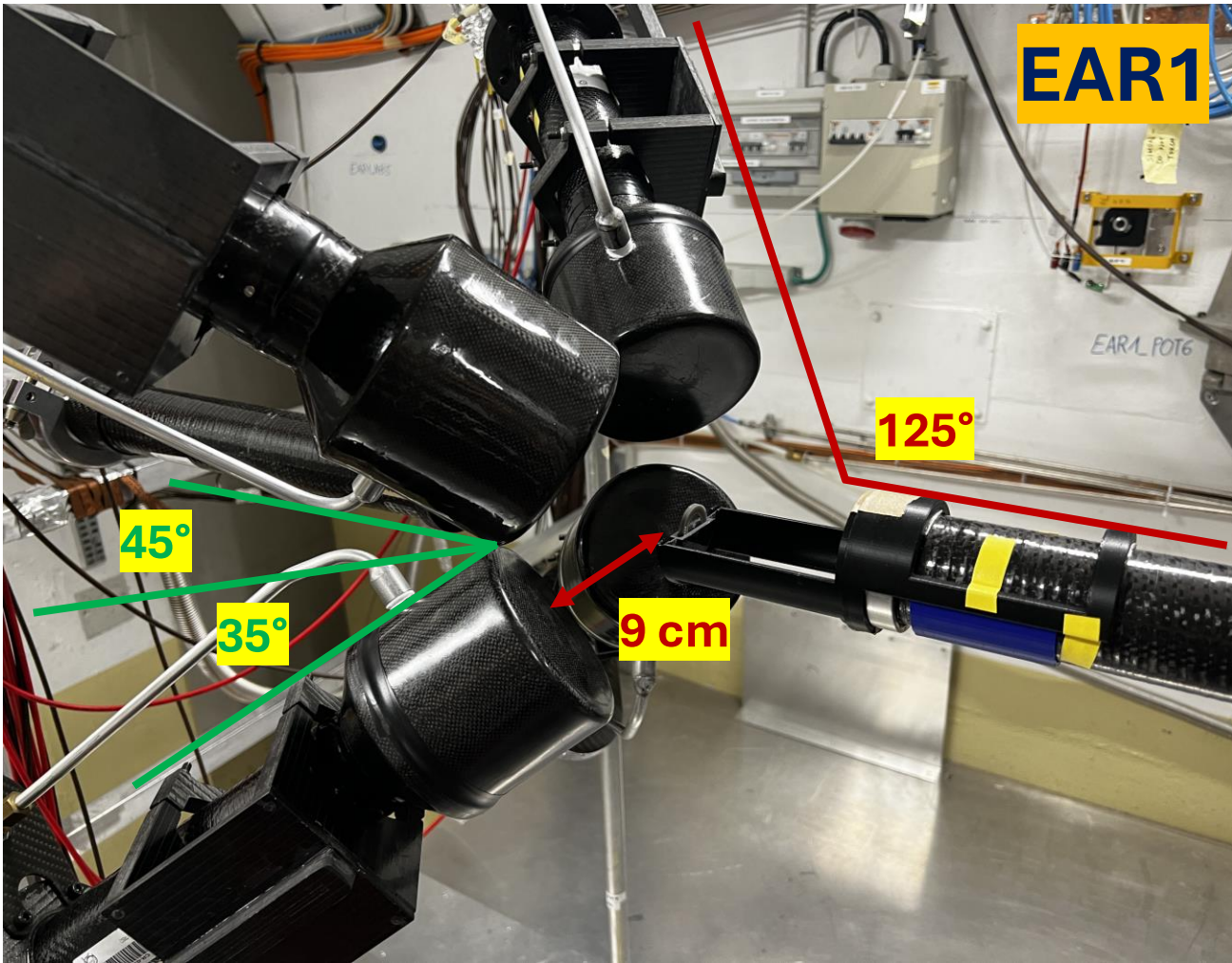




# Setup

# RECAP

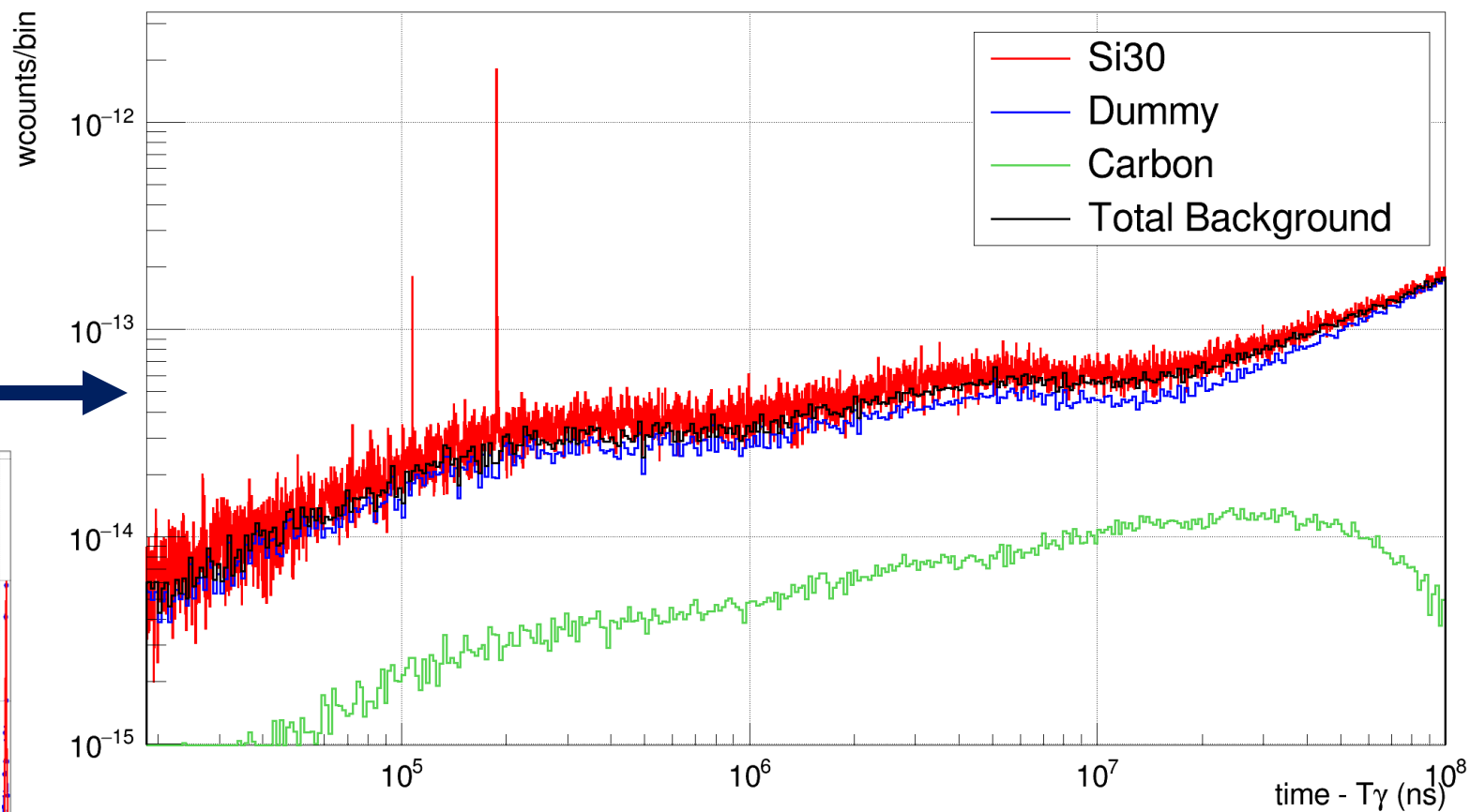
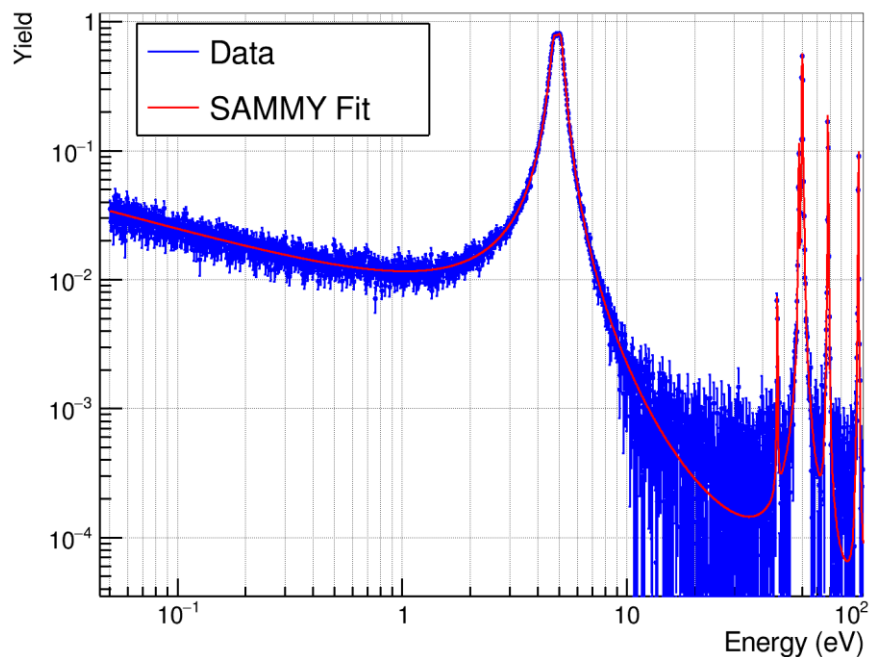
$^{30}\text{Si}(n, \gamma)$ : EAR1 + EAR2;  $^{64}\text{Ni}(n, \gamma)$ : only EAR2



# $^{30}\text{Si}(n,\gamma)$ EAR1: Status

RECAP

1. Calibration
2. Weighting Functions
3. Study of rebounds
4. Gold (reference)
5. **Background Subtraction**

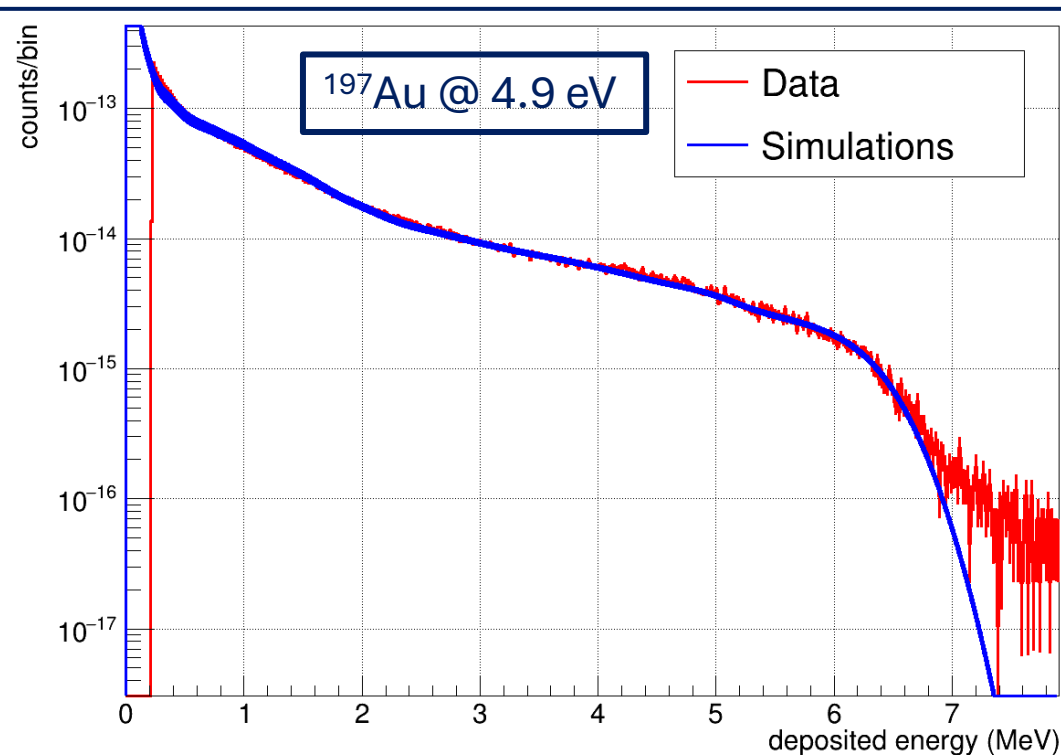
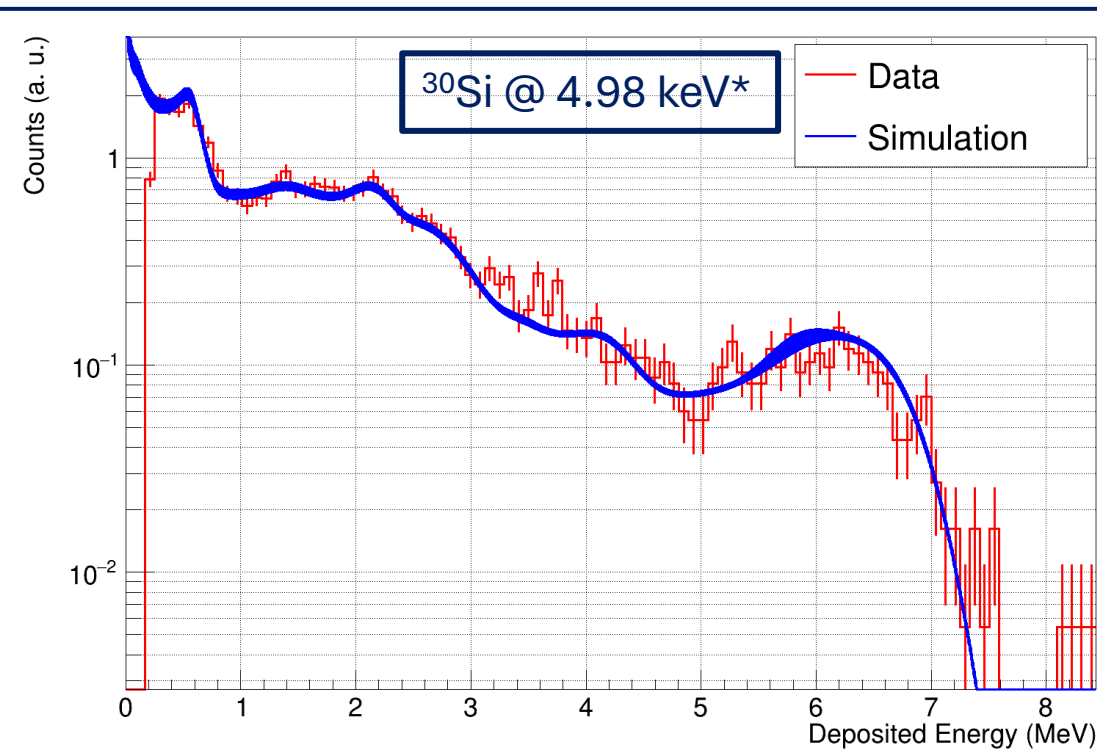


See previous collaboration meeting:  
<https://indico.cern.ch/event/1403789/>

# $^{30}\text{Si}(n,\gamma)$ EAR1: Yield Corrections

$^{30}\text{Si}$  and Gold de-excitation cascades have been simulated and tuned using **NUDEX** to correct for the **thresholds** and the **electron conversion** effects.

	C6D6 1	C6D6 2	C6D6 3	C6D6 4
Total w counts (Si30)	1.0052	1.0057	1.0040	1.0067
Total w counts (Au)	1.0503	1.0584	1.0446	1.0522
Correction $^{30}\text{Si}$ (Si/Au)	<b>0.957</b>	<b>0.950</b>	<b>0.961</b>	<b>0.957</b>

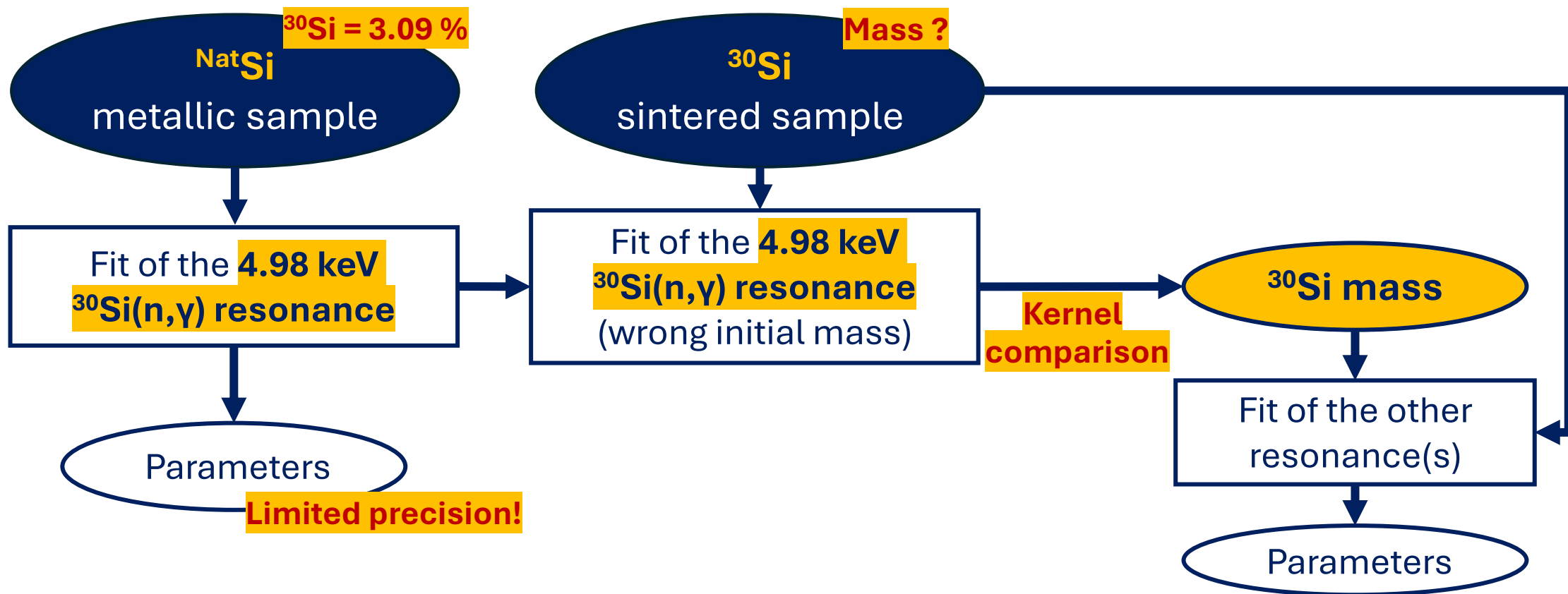


Ex: C6D6 4

\*assumed for all resonances

# $^{30}\text{Si}(n,\gamma)$ EAR1: Mass Problem

The mass of  $^{30}\text{Si}$  sample increased in the sintering procedure (oxidation), so a direct measurement of the actual mass of  $^{30}\text{Si}$  in-beam is not available, but it has been estimated using a **NatSi metallic sample**:



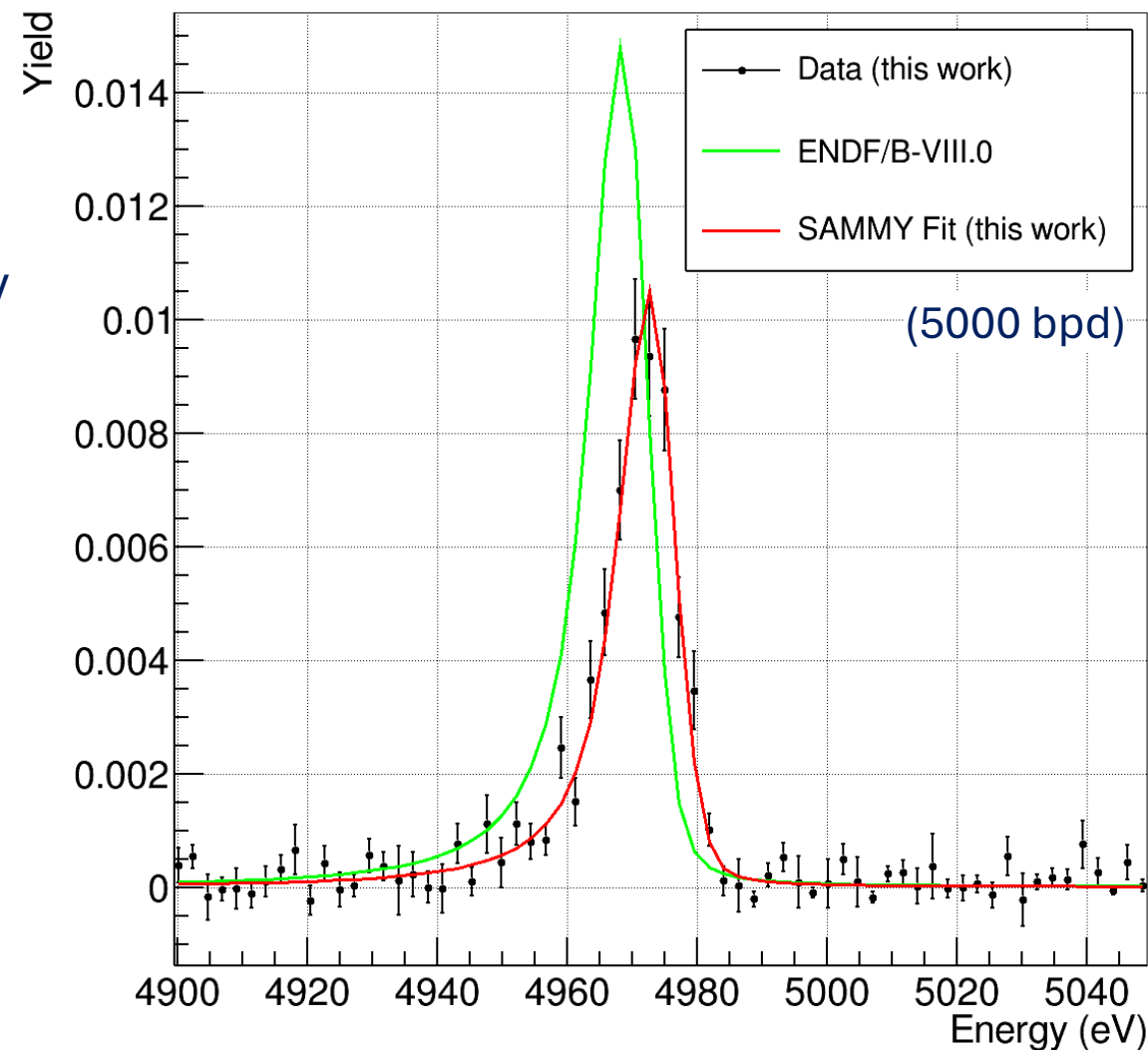


# $^{30}\text{Si}(n,\gamma)$ EAR1: Resonance Fitting

SAMMY fit from the  $^{\text{Nat}}\text{Si}$  sample

$J^\pi$	$1/2^-$
$E_R$ (eV)	$4981.3 \pm 0.3$
$\Gamma_n$ (meV)	$1000 \pm 500$ $\rightarrow 1300 \pm 40$ meV (Harvey 84)
$\Gamma_\gamma$ (meV)	$370 \pm 60$
<b><math>g \Gamma_n \Gamma_\gamma / \Gamma</math> (meV)</b>	<b><math>267 \pm 14</math> (5.2%)</b>

Prev. Measurements	$g \Gamma_n \Gamma_\gamma / \Gamma$ (meV)
<b>Guber et al.</b> <i>Phys. Rev. C</i> 67 (2003) 062802	<b>190</b> (- 29%)
<b>Beer et al.</b> <i>Nucl. Phys. A</i> 453 (2002) 062802	<b>410</b> (+ 54%)
<b>Boldeman et al.</b> <i>Nucl. Phys. A</i> 252 (1975) 62	<b><math>600 \pm 60</math></b> (+125%)





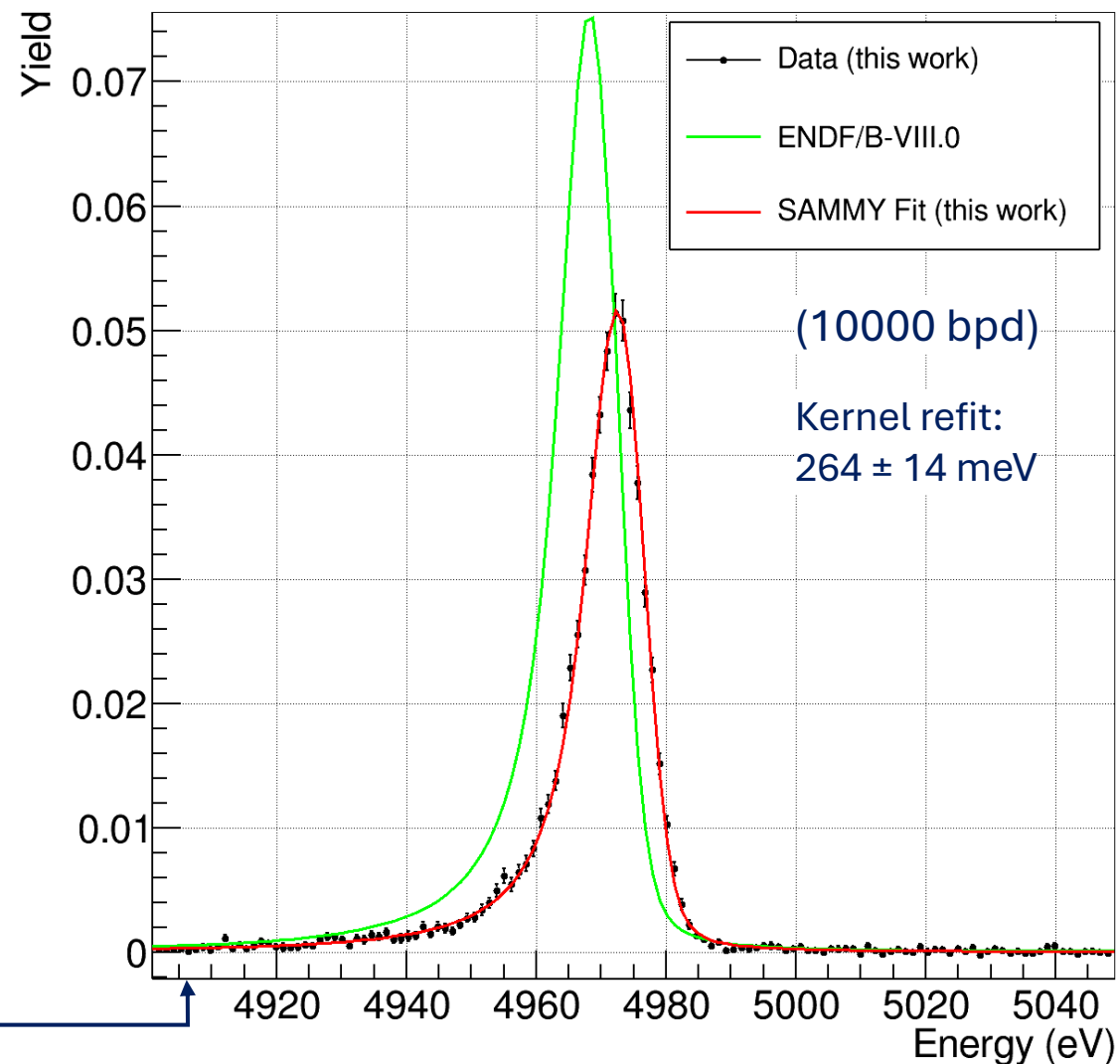
# $^{30}\text{Si}(n,\gamma)$ EAR1: Resonance Fitting

SAMMY fit from the  $^{30}\text{Si}$  sample to determine the actual  $^{30}\text{Si}$  mass in-beam

	$g \Gamma_n \Gamma_\gamma / \Gamma$ (meV)
NatSi sample	$267 \pm 14$ (5.2%)
$^{30}\text{Si}$ sample*	$247 \pm 3$ (1.2%)
<b>Ratio</b>	<b><math>0.92 \pm 0.05</math> (5.3%)</b>

	Mass (mg)
Pre-sintering*	$750.34 \pm 0.01$
Sample (total)	$992.5 \pm 0.1$ (+32%)
<b><math>^{30}\text{Si}</math> estimated</b>	<b><math>690 \pm 40</math> (5.8%) (-8%)</b>

In mass: 70%  $^{30}\text{Si}$ , 30% Oxygen  
 In atoms: 55%  $^{30}\text{Si}$ , 45% Oxygen

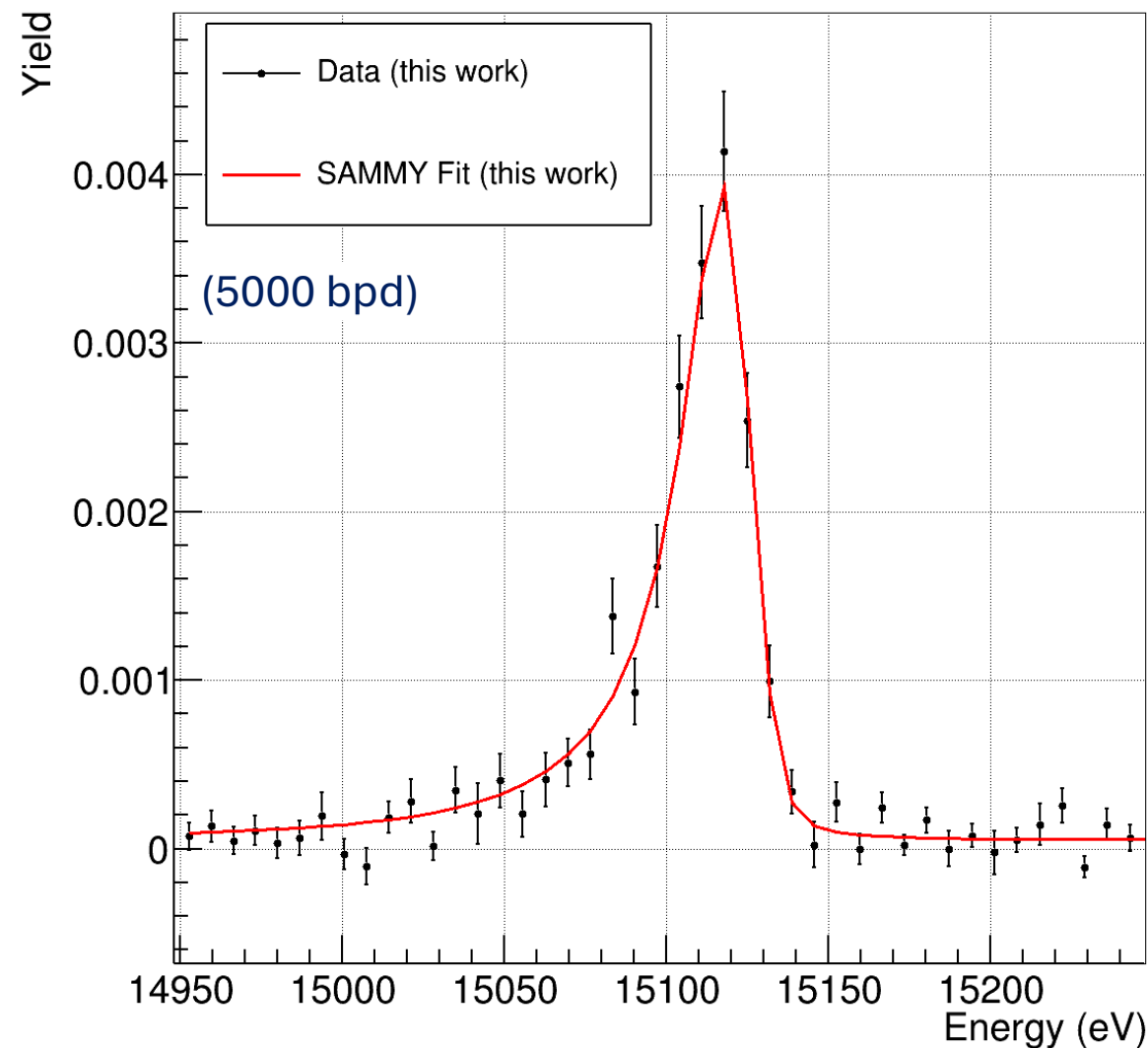


# $^{30}\text{Si}(n,\gamma)$ EAR1: Resonance Fitting

Resonance at 15.14 keV **not present in evaluated data libraries**, but already observed in previous measurements

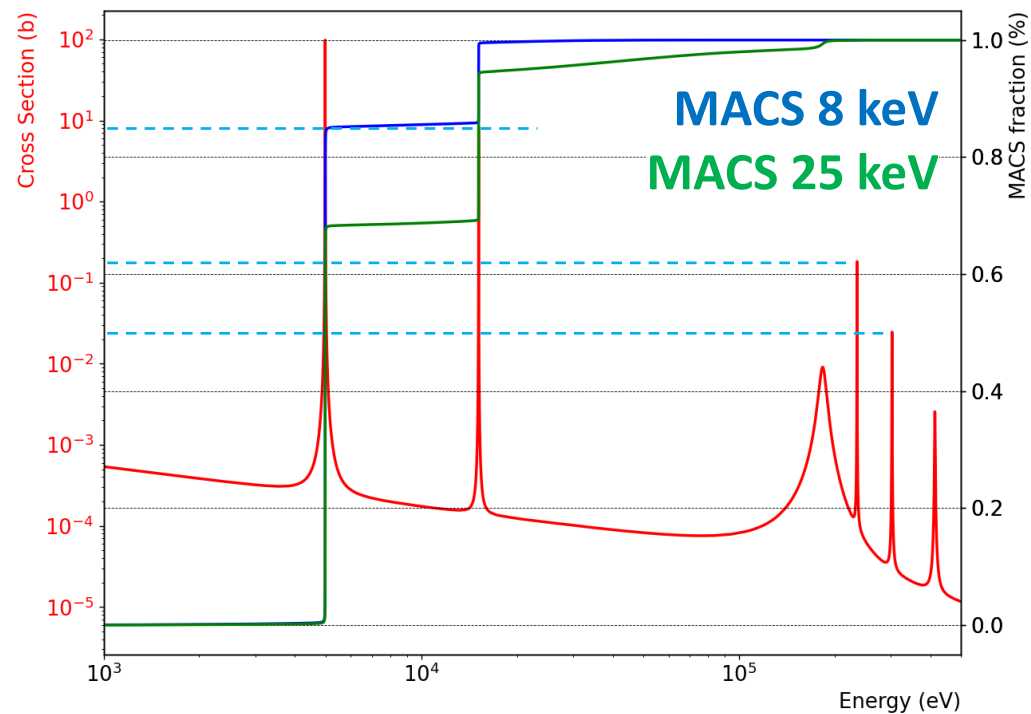
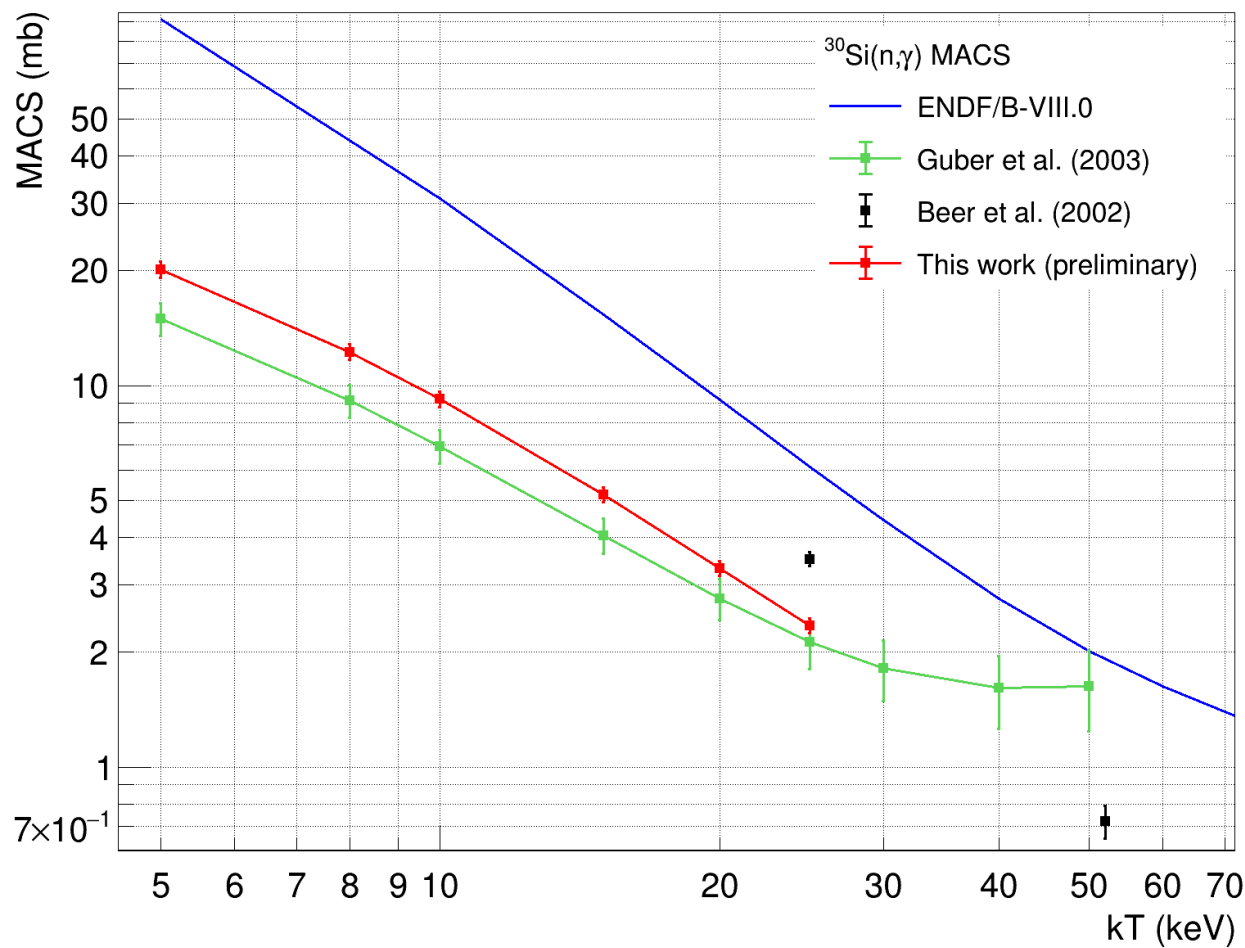
$J^\pi$	$(\frac{1}{2}^-)$
$E_R$ (eV)	$15144.1 \pm 0.6$
$\Gamma_n$ (meV)	$3000 \pm 1200$
$\Gamma_\gamma$ (meV)	$153 \pm 7$
$g \Gamma_n \Gamma_\gamma / \Gamma$ (meV)	<b><math>146 \pm 11</math> (7.5%)</b>

Prev. Measurements	$g \Gamma_n \Gamma_\gamma / \Gamma$ (meV)
<i>Guber et al.</i>	94 (-35%)
<i>Beer et al.</i>	120 (-18%)
<i>Boldeman et al.</i>	$170 \pm 20$ (+16%)



# $^{30}\text{Si}(n,\gamma)$ EAR1: Preliminary MACS

The first two resonances are the dominant contribution to the **MACS** up to 25 keV (**Direct capture\*** contribution still missing, estimated at 0.3 – 0.45 mb by Beer / Guber)

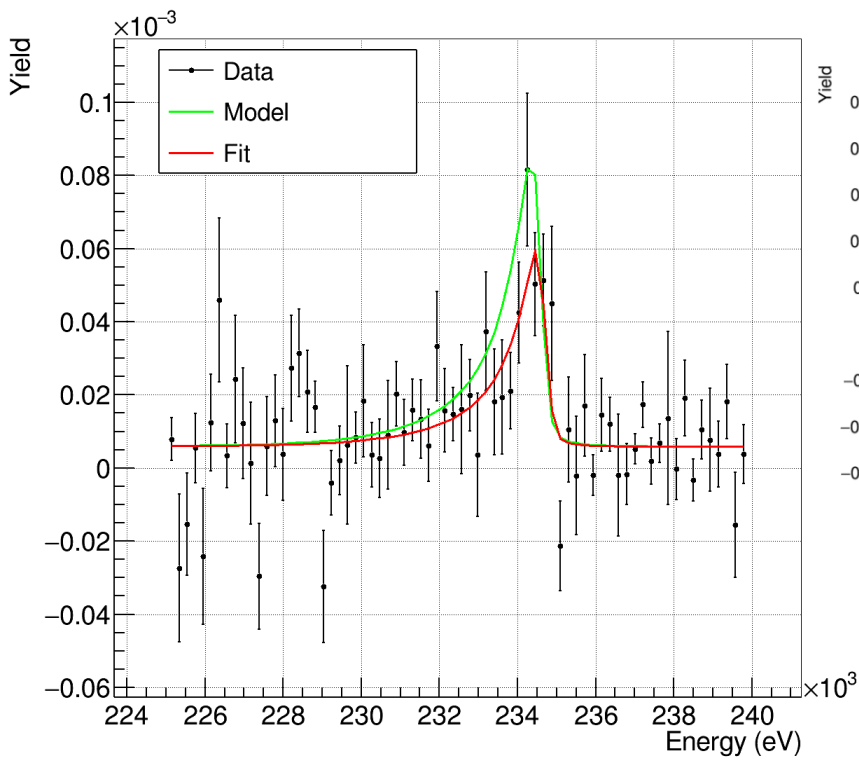


MACS @ 25 keV	This work	Guber et al.	Beer et al.
	2.35(10)**	2.1(3)	3.51(15)

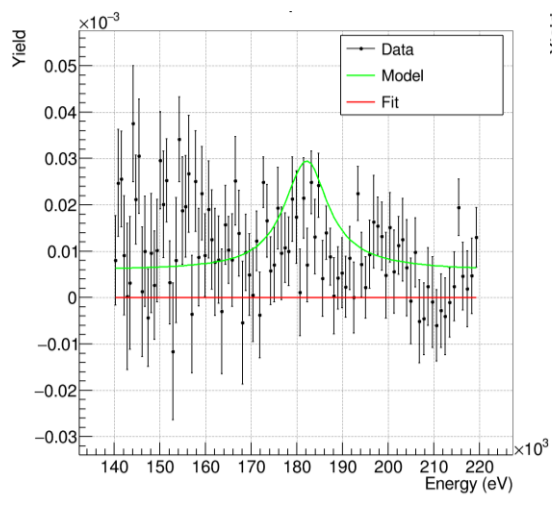


# $^{30}\text{Si}(n,\gamma)$ EAR1: High Energy Region

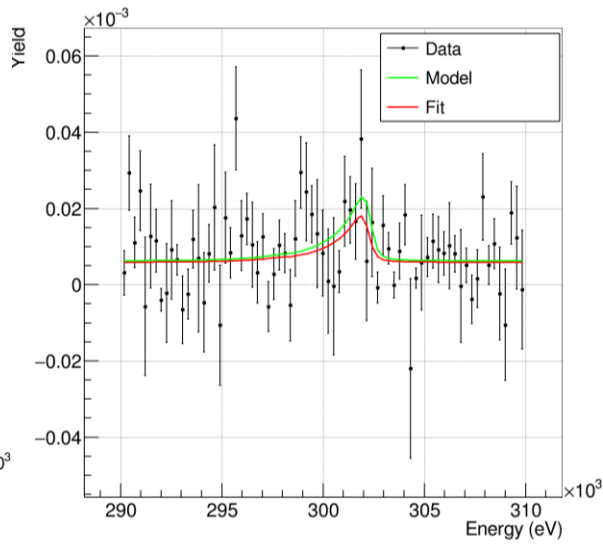
**Resonances after 100 keV** are smaller and suffer by large fluctuations and poor statistics. For the moment, **just some fitting attempts** have been performed...



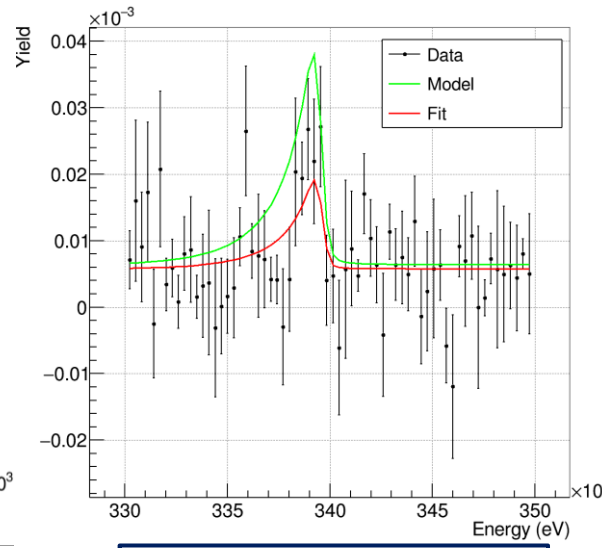
**235 keV**



**180 keV**  
Never measured in capture, upper limit?



**303 keV**

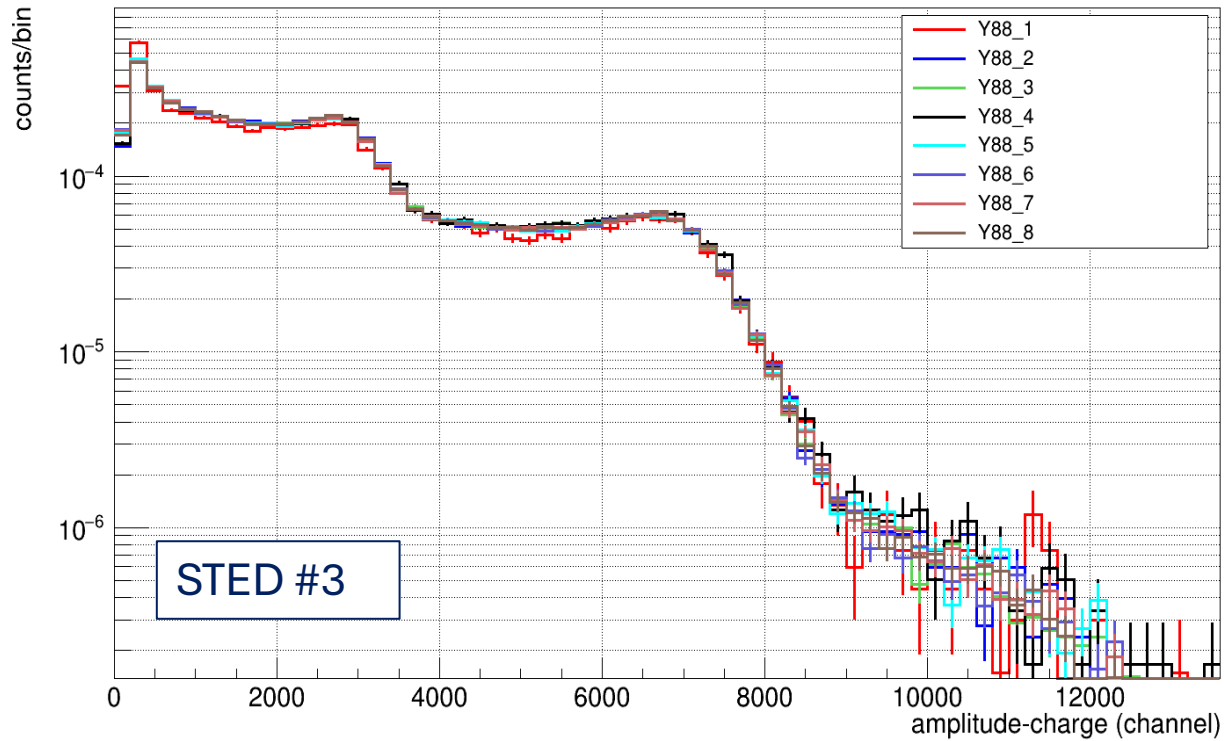


**340 keV**  
Not reported in libraries

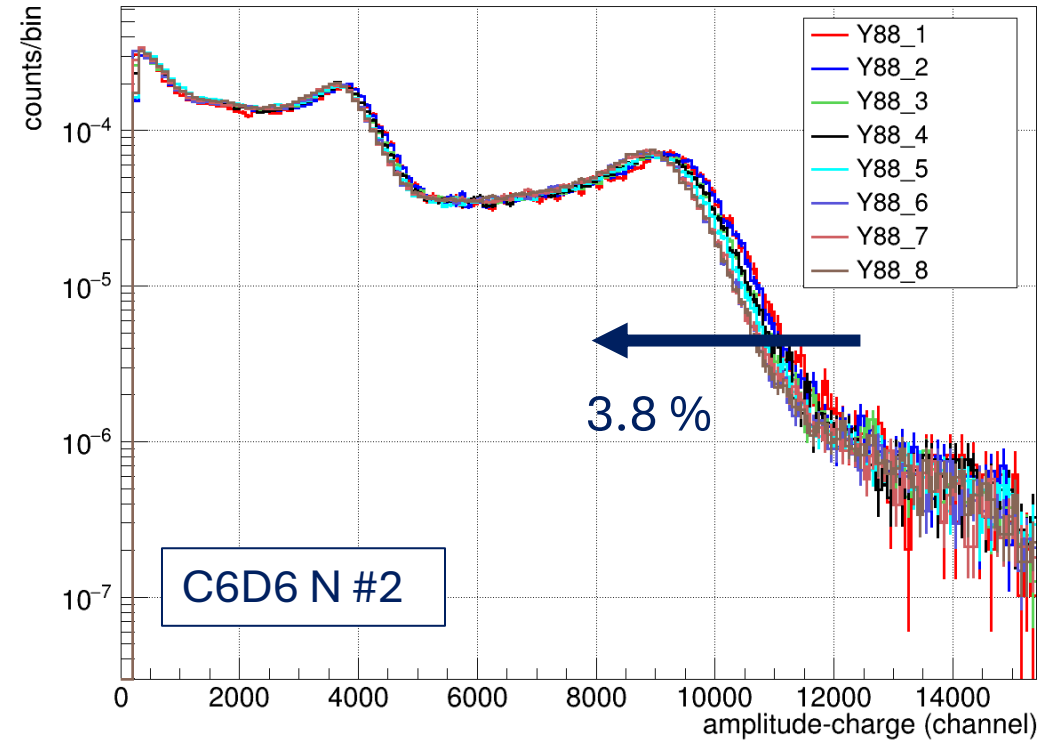
# EAR2: Calibrations

Gain shift has been comparing different Y88 and Au runs:

**STED: Stable**

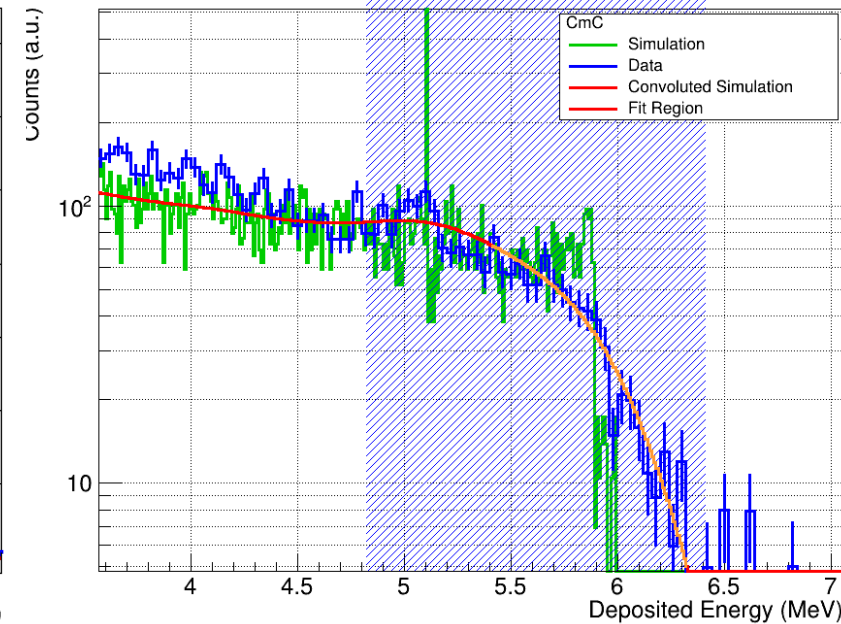
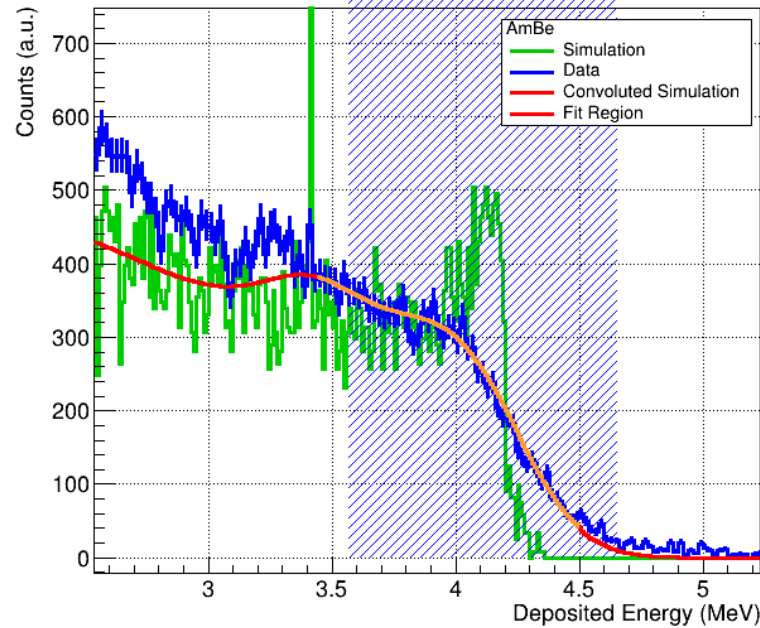
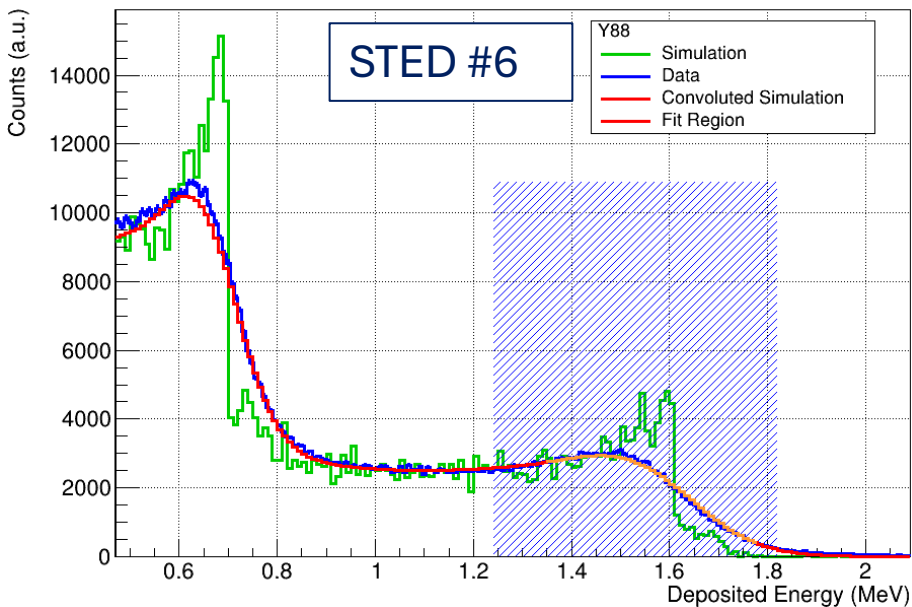
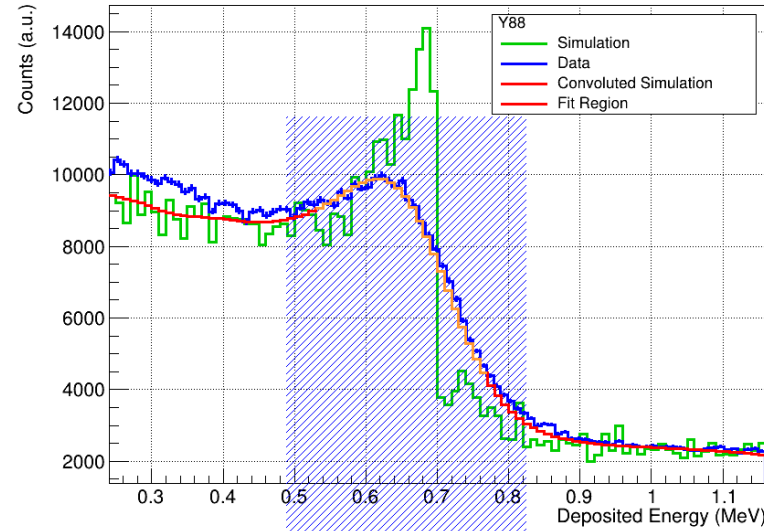
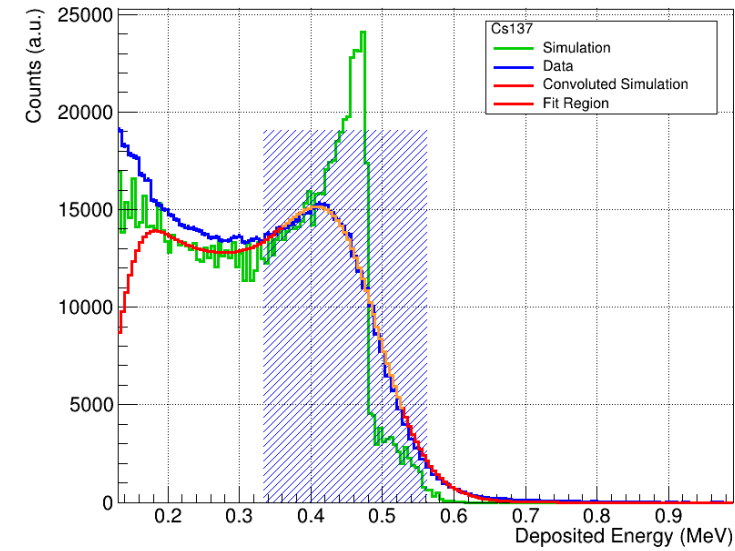


**C6D6: Gain Shift**



# EAR2: Calibrations STED

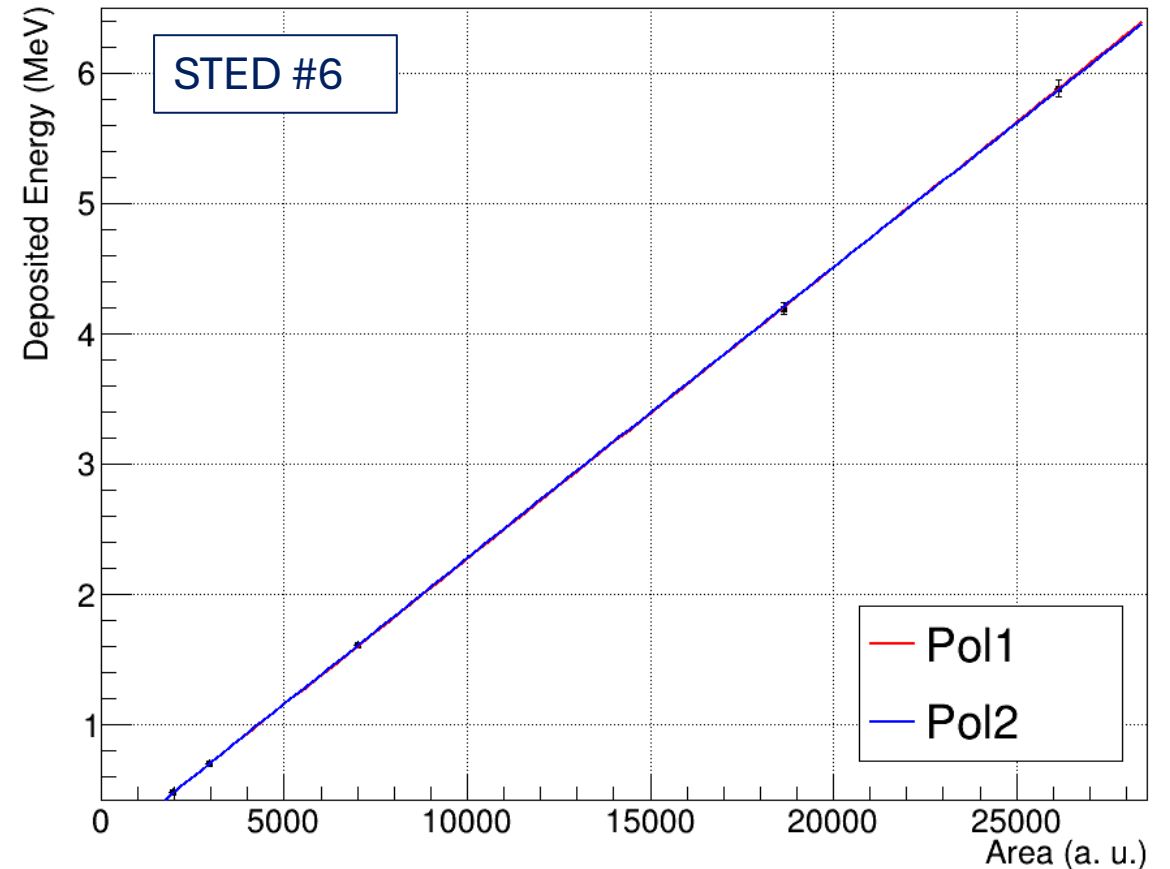
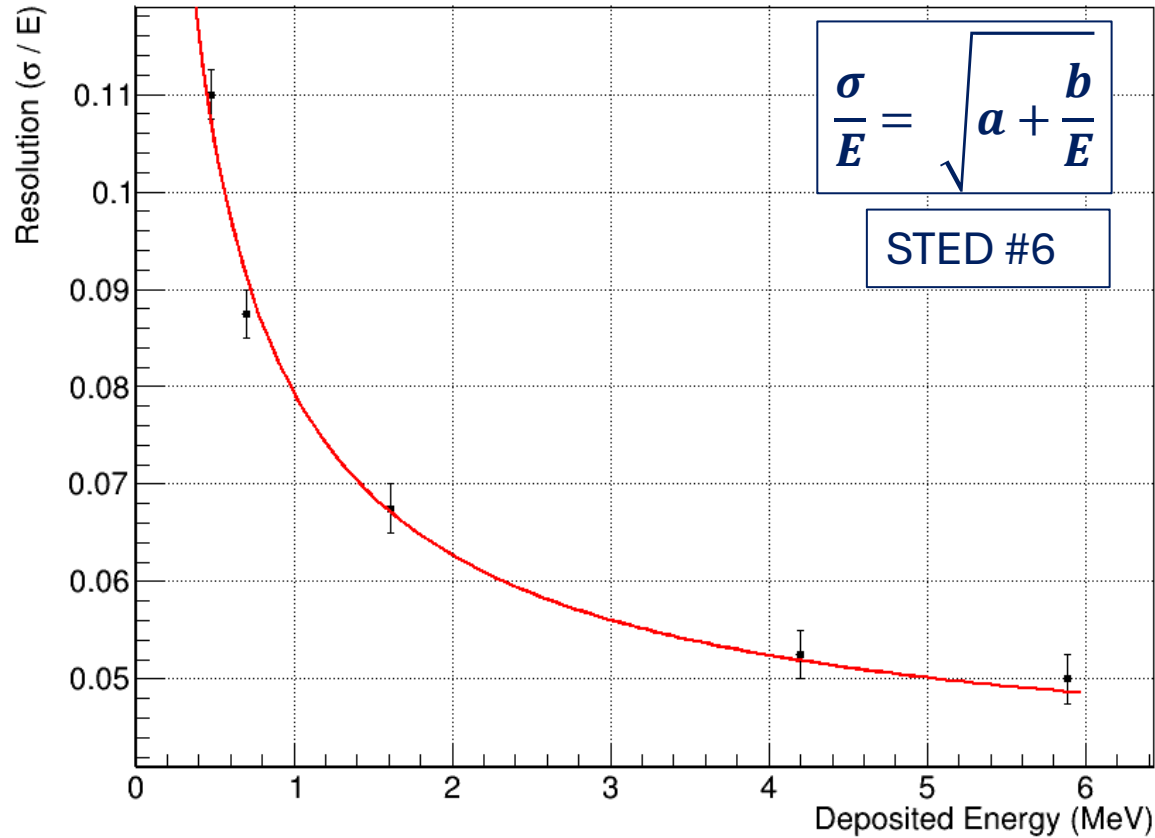
Calibration of the STED has been performed with a **direct comparison** between measured and simulated energy deposited spectra, tuning gain and resolution for each source:





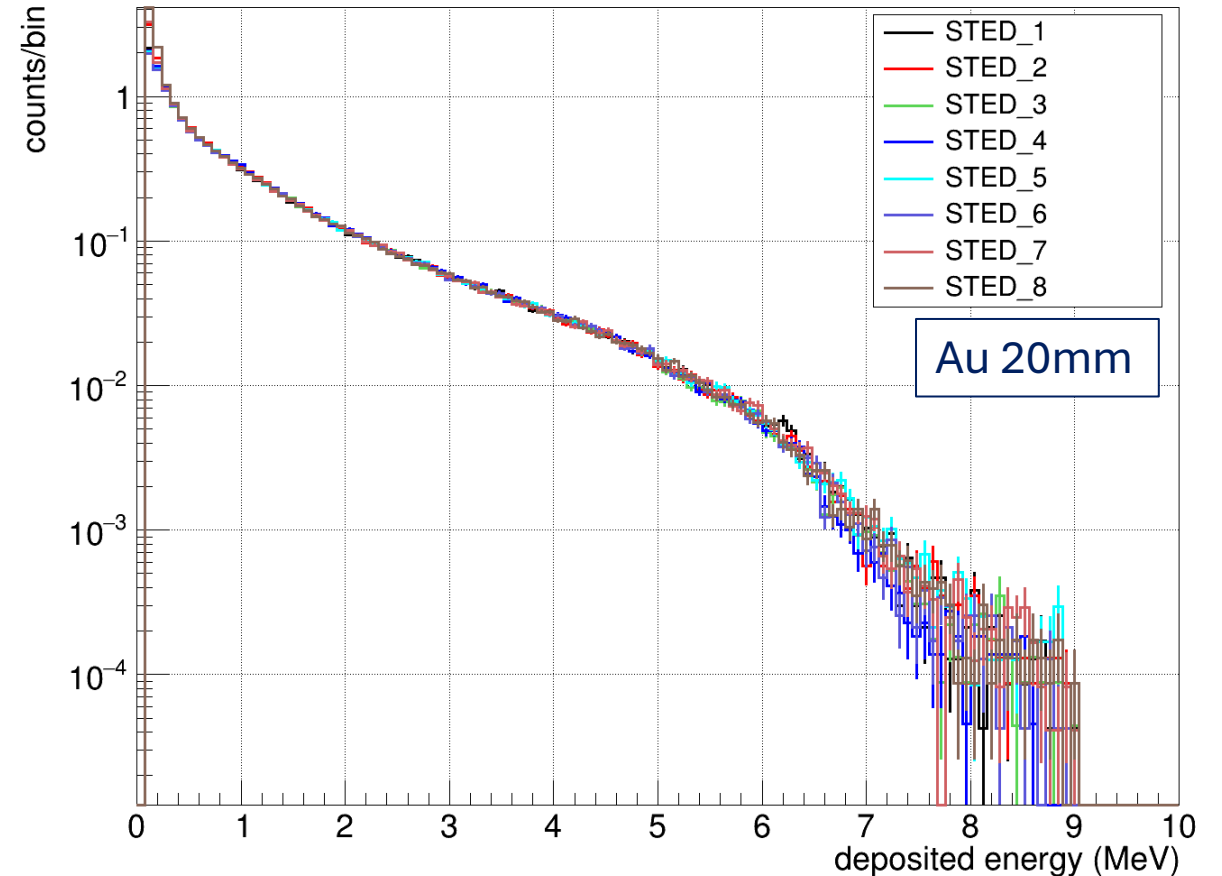
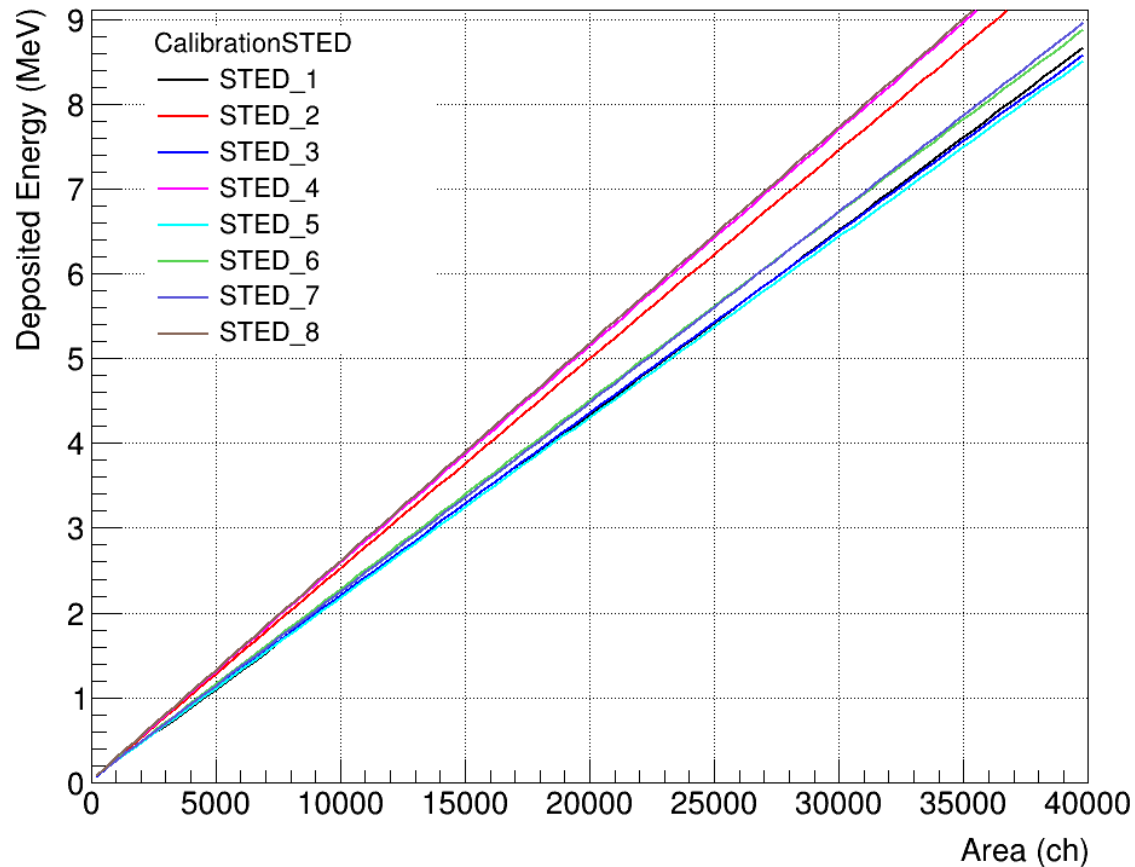
# EAR2: Calibrations STED

Calibration and resolution functions have been computed fitting the results from different sources. The calibration of the STEDs are pretty linear.



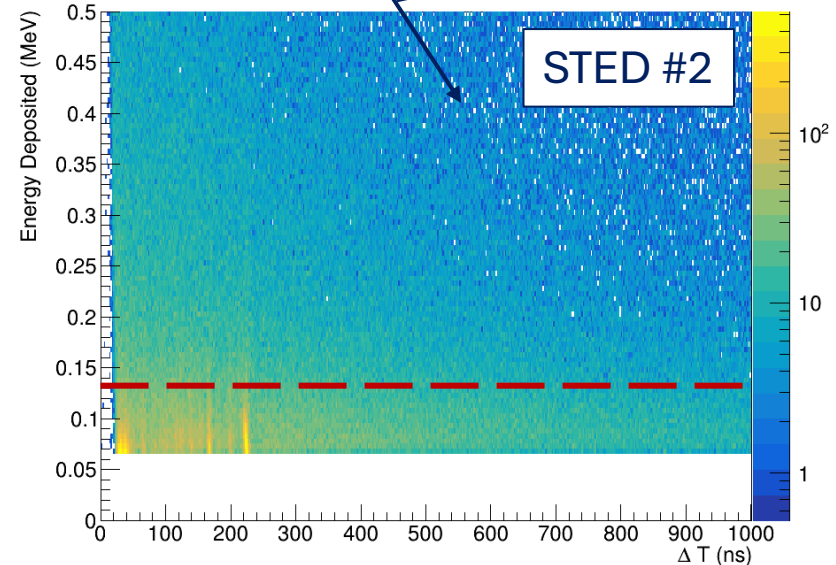
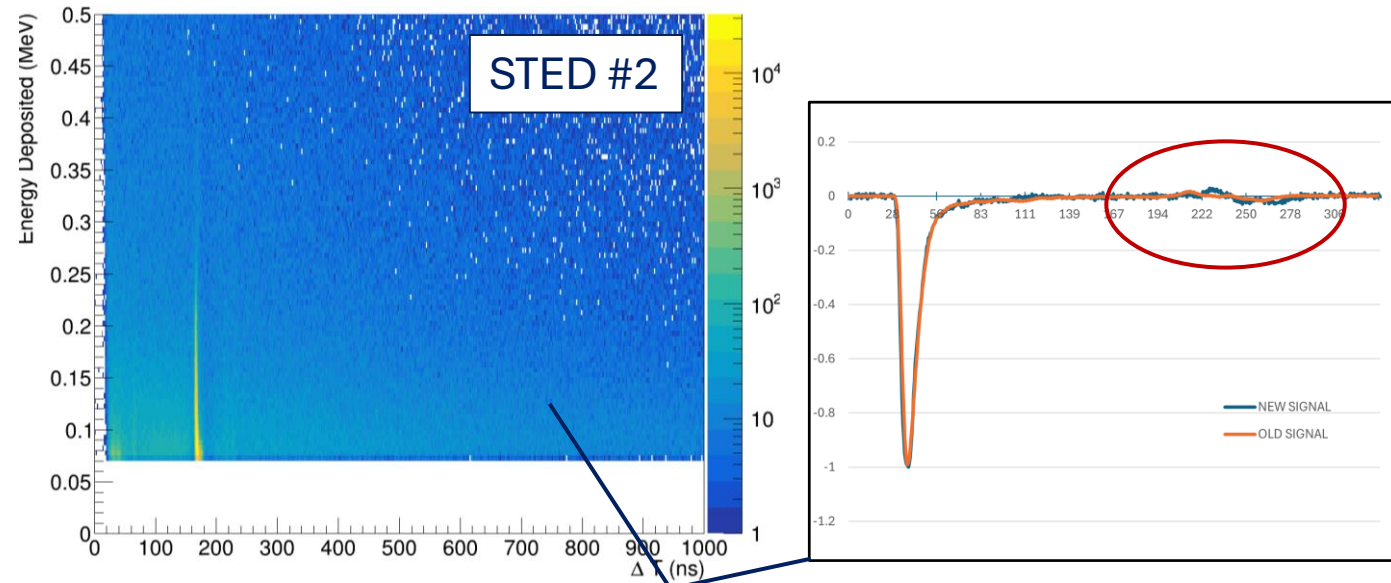
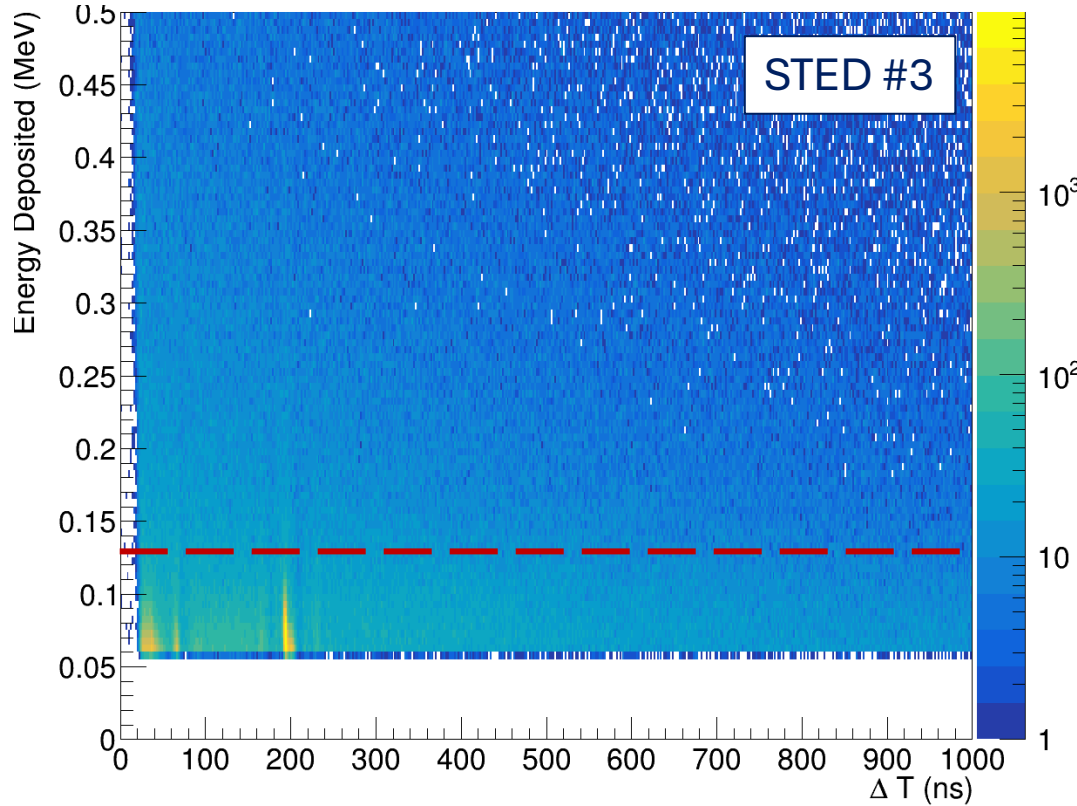
# EAR2: Calibrations STED

Calibration and resolution functions have been computed fitting the results from different sources. The calibration of the STEDs are pretty linear.



# EAR2: Thresholds STED

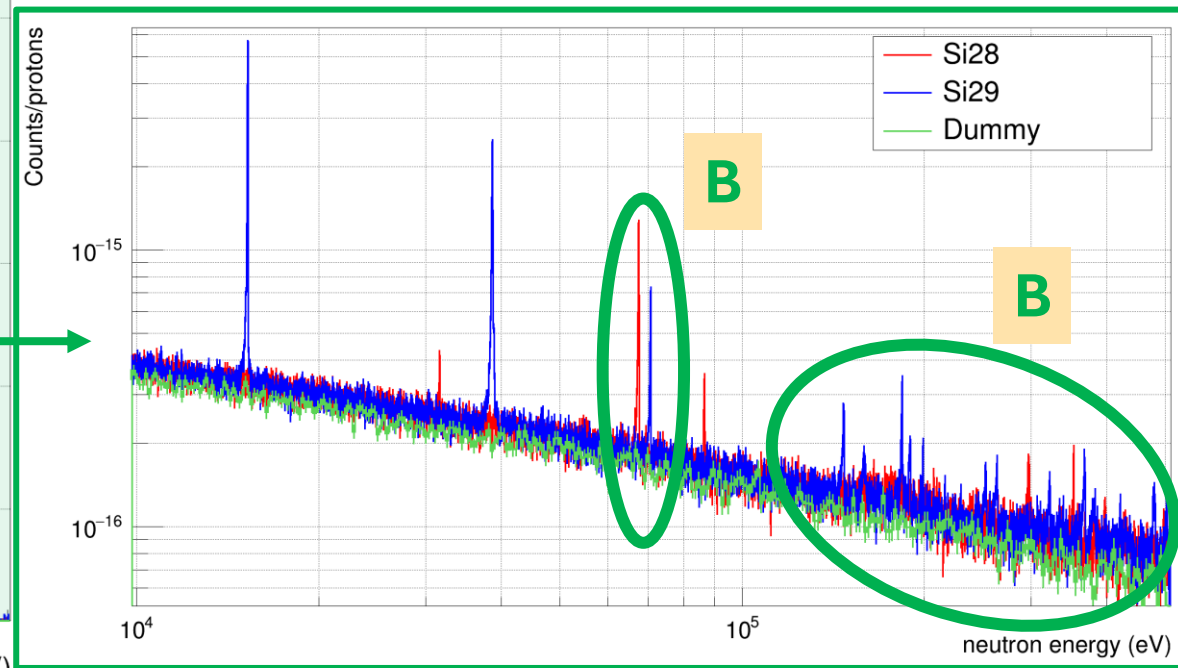
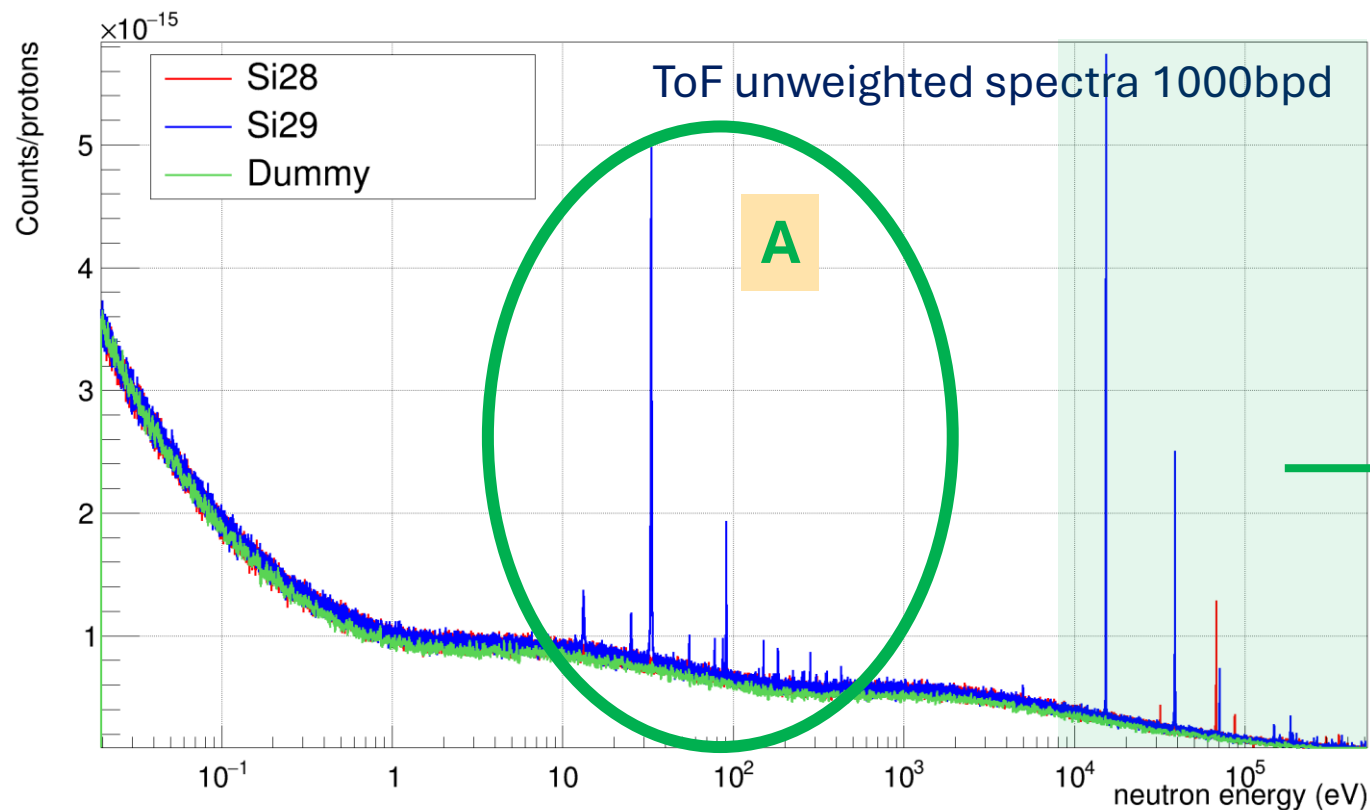
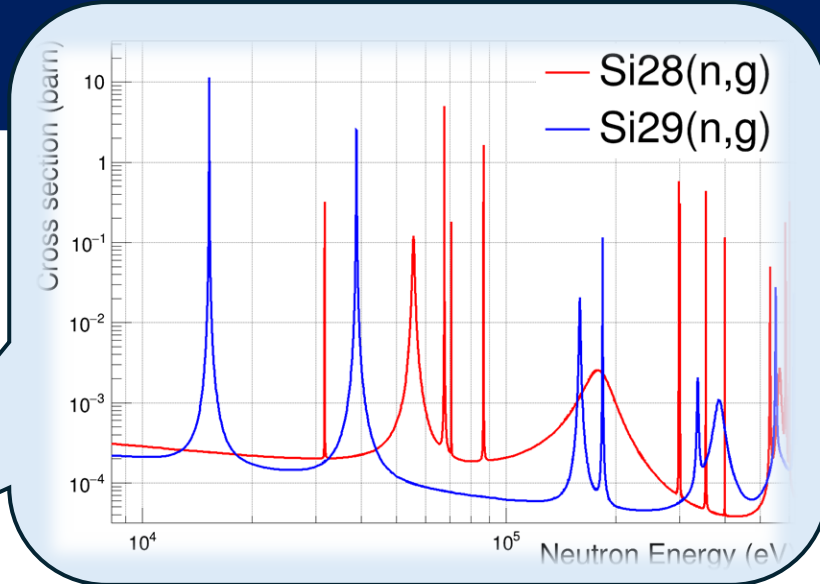
STED rebounds have been fixed adjusting the **PSA**. Energy deposited thresholds can be set at **130 keV**.



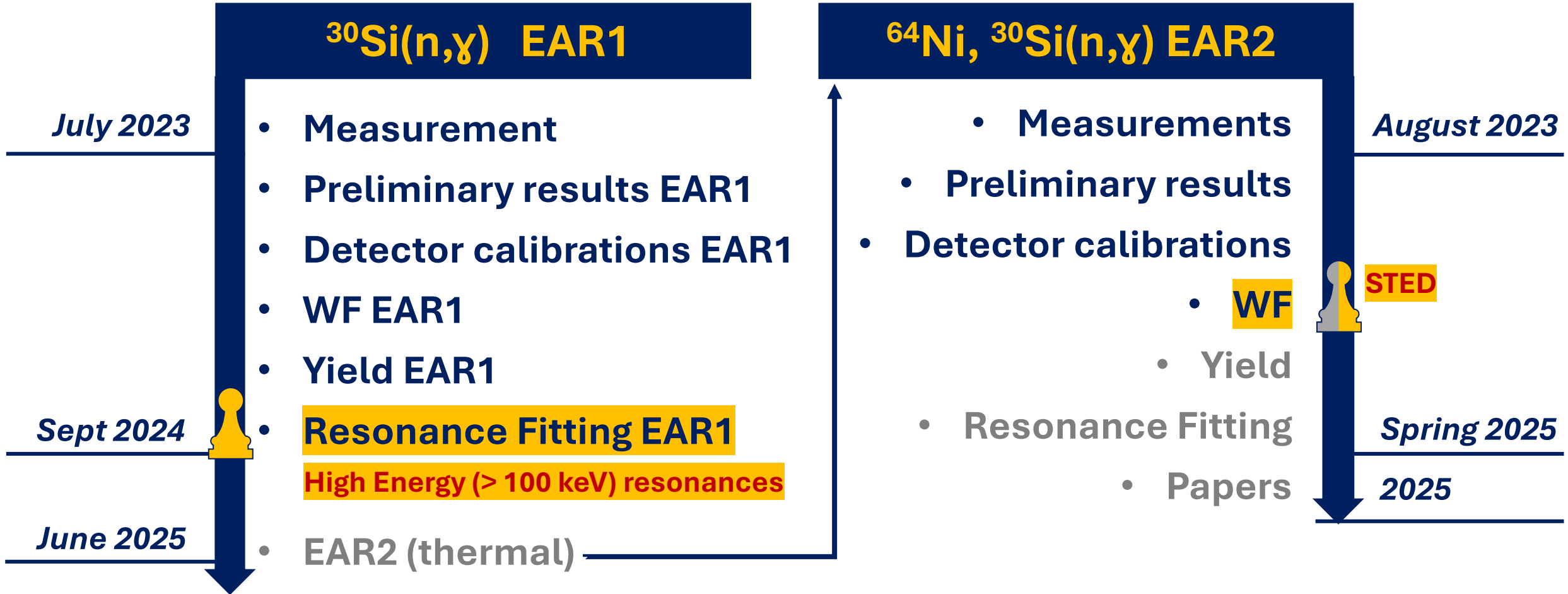


# 28, 29Si(n,γ) Preliminary results

- A. Contamination from **Pd** in Si29 sample ?
- B. Some “**new**” resonances for **Si29** ( $\approx 70$  keV,  $> 100$  keV) ?
- C. Some resonances not seen in **Si28**



# Summary



# $^{30}\text{Si}(n,\gamma)$ & $^{64}\text{Ni}(n,\gamma)$ : Status of the Analysis

**Michele Spelta**

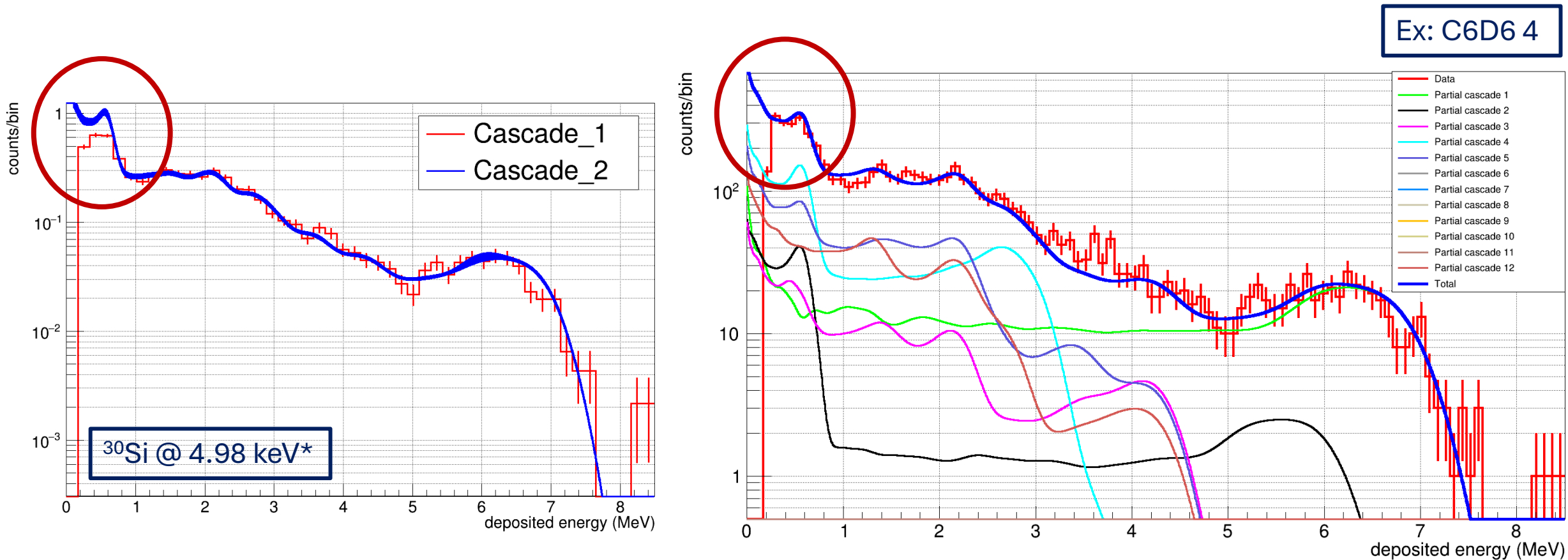


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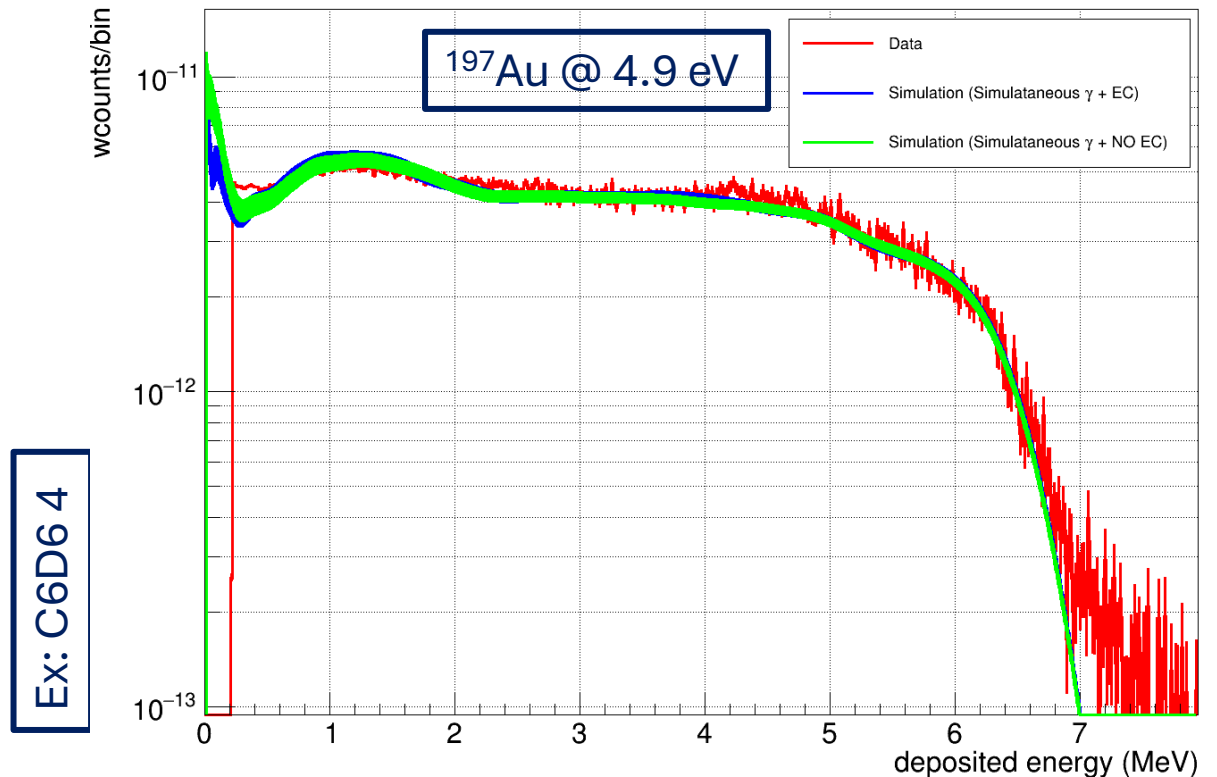
# $^{30}\text{Si}(n,\gamma)$ EAR1: Yield Corrections

$^{31}\text{Si}$  de-excitation cascade for the 4.98 keV resonance has been tuned using **NUDEX** to correctly reproduce the measured spectrum (tuning of the primary transition probability)



# $^{30}\text{Si}(n,\gamma)$ EAR1: Yield Corrections

Gold de-excitation cascade for the 4.9 eV resonance has been simulated using **NUDEX** to correct for the threshold and electron conversion.

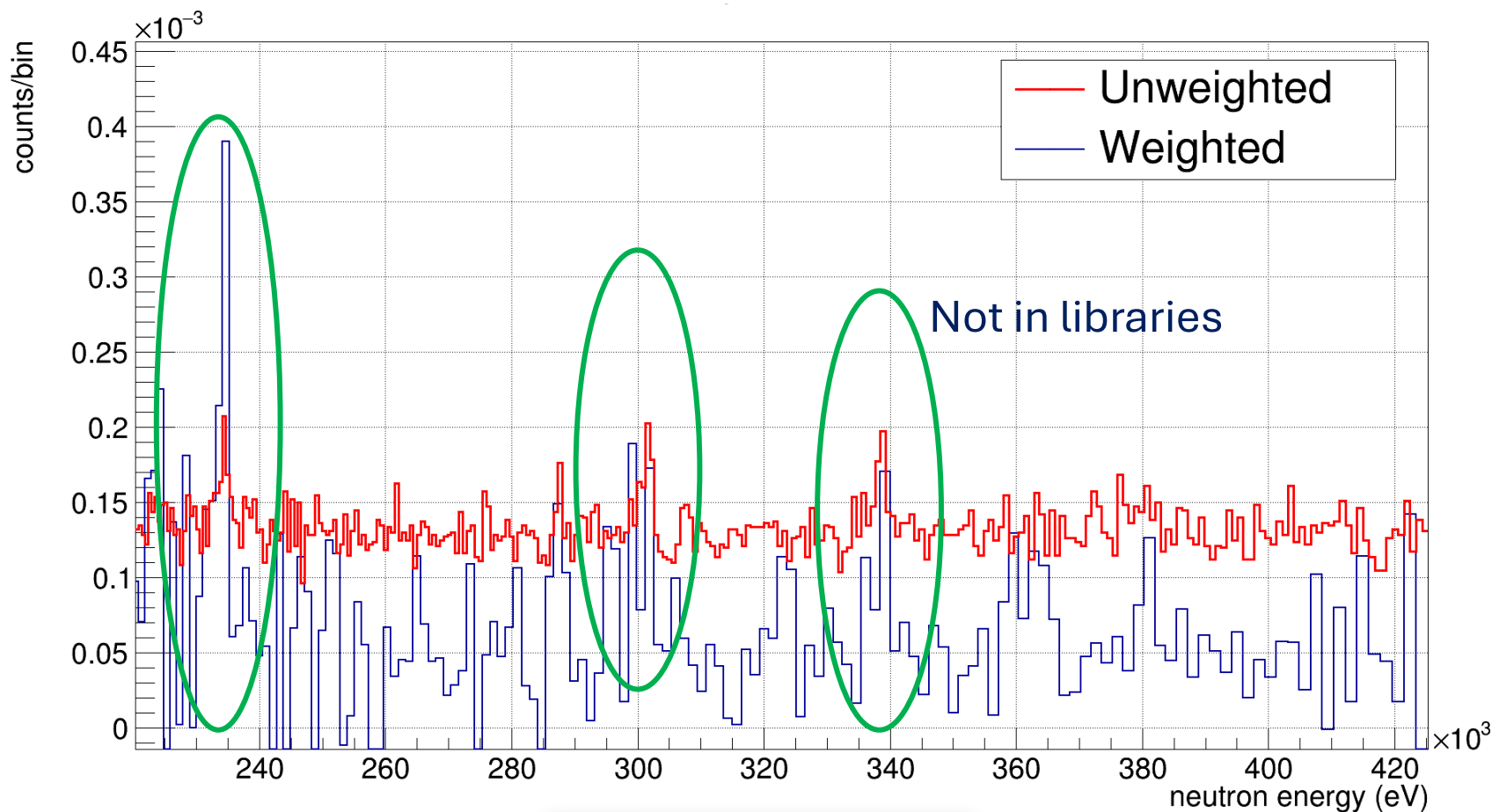


	C6D6 1	C6D6 2	C6D6 3	C6D6 4
Total w counts <b>no thrshold</b> (Au)	1.0396	1.0440	1.0364	1.0460
Total w counts + <b>EC</b> (Au)	1.0103	1.0138	1.0079	1.0060
Total w counts <b>no thr + EC</b> (Au)	1.0503	1.0584	1.0446	1.0522



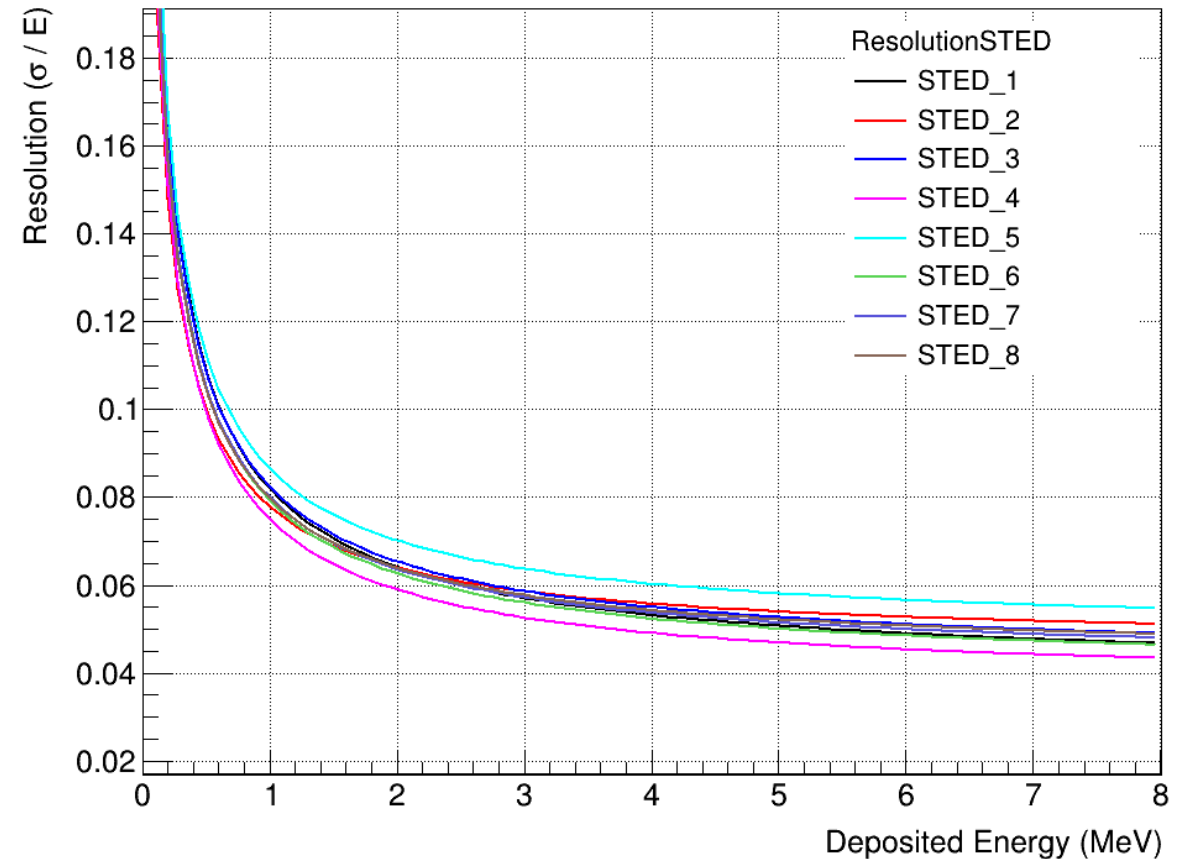
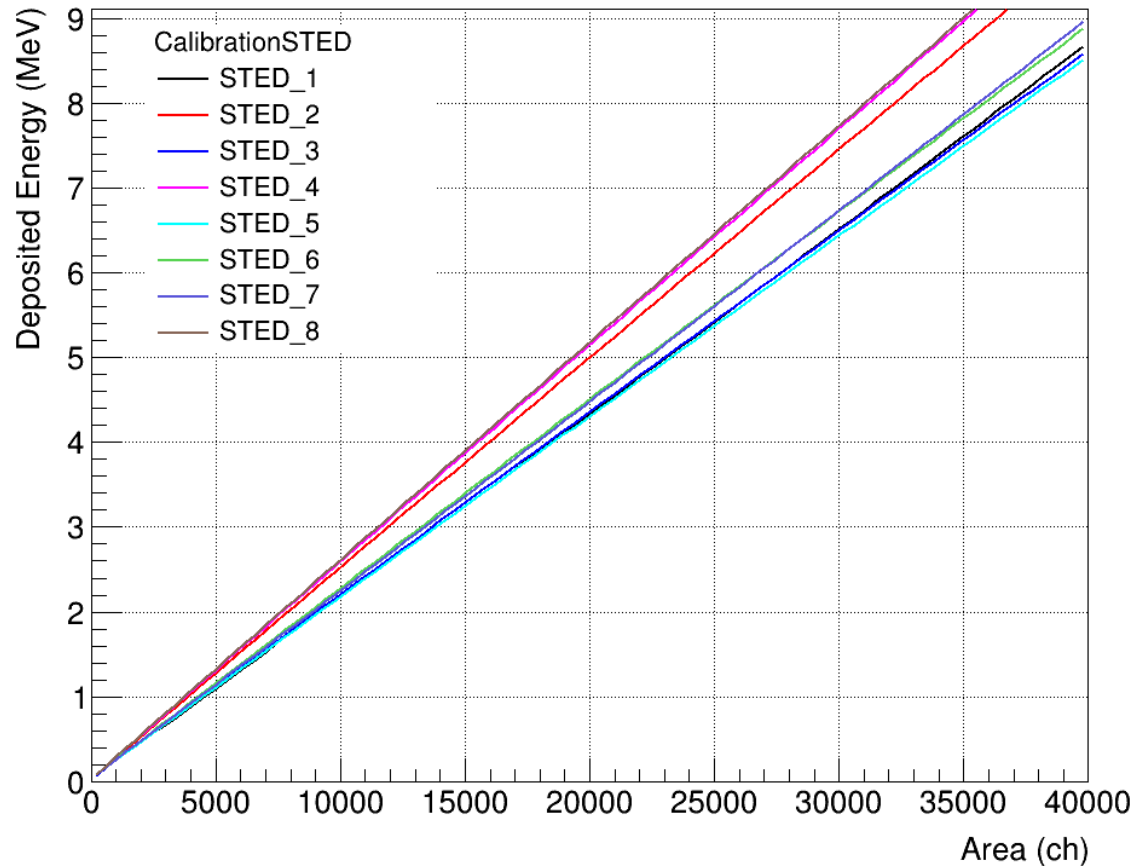
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Calibration and resolution functions have been computed fitting the results from different sources. The calibration of the STEDs are pretty linear.



# EAR2: Thresholds STED

STED rebounds have been fixed adjusting the **PSA**. Energy deposited thresholds can be set around **130 keV**.

