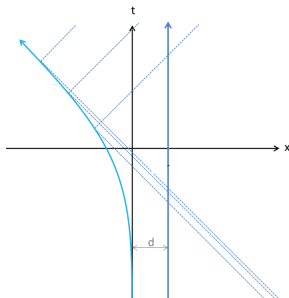


# Detectors and moving boundaries

**Jorma Louko**

School of Mathematical Sciences, University of Nottingham

SIGRAV International School  
Vietri sul mare, 18 February 2025



# Plan

## Part A

- ▶ Historical remarks
- ▶ Localised quantum system interacting with a non-localised quantum field: detector!
- ▶ Time dependent systems: switching interaction on and off
- ▶ Six examples 1977–2025

## Part B

At most one of:

- ▶ Vacua in locally de Sitter cosmologies, and how to distinguish them
- ▶ Temperature of acceleration, in spacetime and in the laboratory

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Wilson et al 2011

Lähteenmäki et al 2013

(electrical simulation of motion)

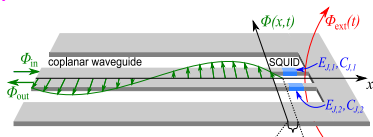


Image: Johansson et al 2009



# History

Cosmological 'particle creation'

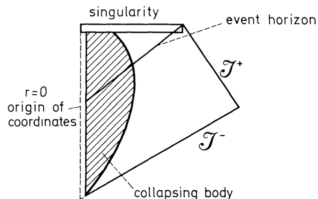
Parker 1968...

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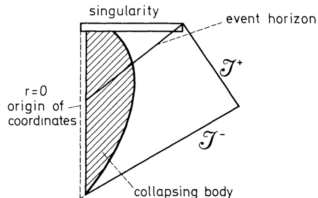
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Receding mirror in Minkowski

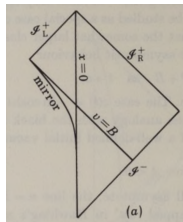
Tailored late time trajectory  
⇒ Hawking-type radiation  
from the mirror

Parker 1968...

Hawking 1974, 1975



Davies and Fulling 1977



# Arena: Fock space

## Input

- ▶ “vacuum plus excitations” (old school)
- ▶ relies on notion of “positive frequency”
  - ⇒ different positive frequency choices related by Bogoliubov transformations
- ▶ typically employs spread-out Fourier modes
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## Quantum field

$D$  spacetime dimension

$\phi$  real scalar field

$|\Psi\rangle$  field initial state

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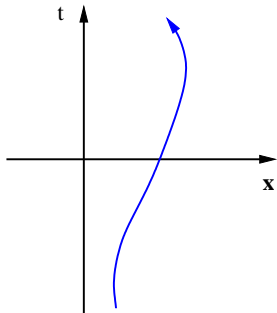
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$$H_{\text{int}}(\tau) = c\chi(\tau)\mu(\tau)\phi(x(\tau))$$

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$W$ : field initial state  $|\Psi\rangle$  and detector motion

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$E$ : detector energy gap

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**Theorem** (Hörmander 1971)

If  $|\Psi\rangle$  is Hadamard and  $x(\tau)$  is smooth, then  $W(\tau', \tau'')$  is a well-defined distribution on  $\mathbb{R} \times \mathbb{R}$

**Corollary**

If  $\chi$  is  $C_0^\infty$  and  $x(\tau)$  is smooth,  $F_{\chi}(E)$  is **well defined!**



## Switching?

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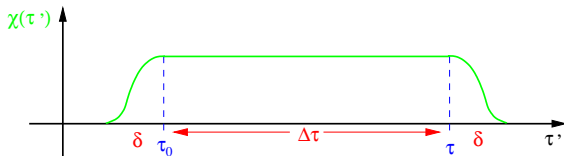
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**Stationary** motion:  $W(\tau', \tau'') = W(\tau' - \tau'', 0)$

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- static observer outside Schwarzschild black hole, ...

Long time limit  $\rightarrow$  **transition rate**:



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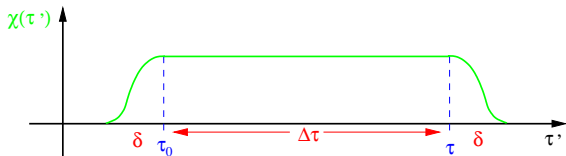
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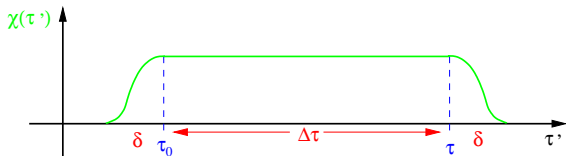
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Long time limit  $\rightarrow$  **transition rate**: yet within linear perturbation theory!

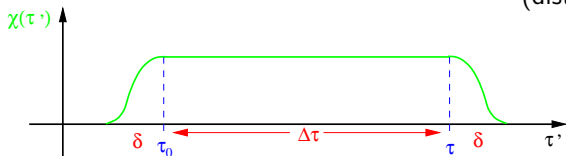


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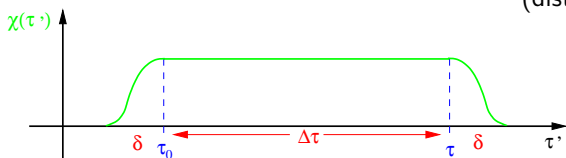
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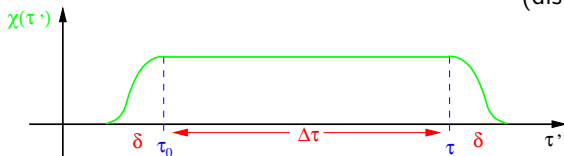
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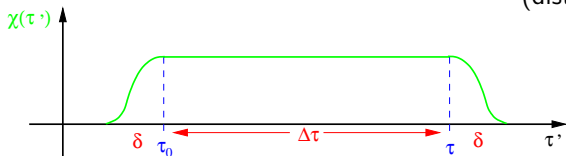
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Instantaneous transition rate divergent

(for generic states/trajectories)



## Remarks

- ▶  $W(\tau', \tau'')$  sometimes infrared divergent (eg  $D = 2$  massless)  
Cure:  $H_{\text{int}}(\tau) \propto \frac{d}{d\tau} \phi(\mathbf{x}(\tau))$  (but more singular at sharp switching)
- ▶ Fermions:  $H_{\text{int}}(\tau) \propto \overline{\psi(\mathbf{x}(\tau))} \psi(\mathbf{x}(\tau))$  **quadratic** Takagi 1986  
 $\Rightarrow$  more singular JL and Toussaint 2016
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Needed in entanglement harvesting by multiple detectors
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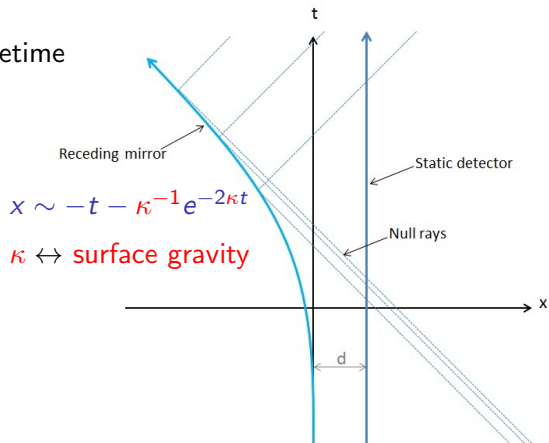
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- ▶ Moving cavities Bruschi et al 2012, Lorek et al 2015, ...
- ▶ Complementary track: open quantum systems approach  
Long time, weak coupling, Markovian approximation  
 $\Rightarrow$  Sharp switch-off well defined (after renormalisation)  
Benatti and Floreanini 2004, ...
- ▶ Pointlike two-state detector  $\rightarrow$  pointlike SHO  
 $\Rightarrow$  Gaussian nonperturbative methods Lin and Hu 2007, ...

# Example: 1+1 receding mirror

Juárez-Aubry and JL 2015

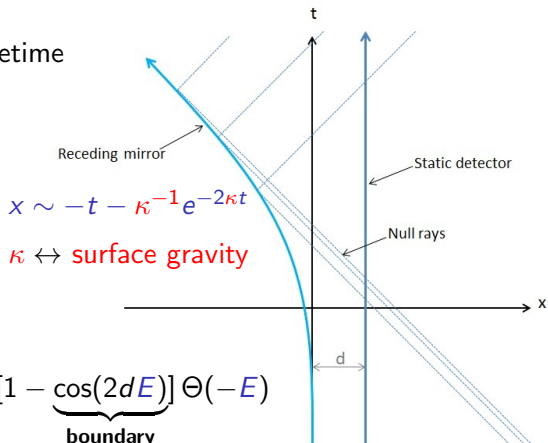
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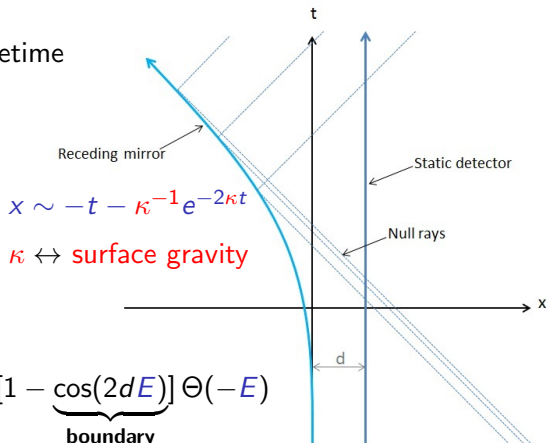
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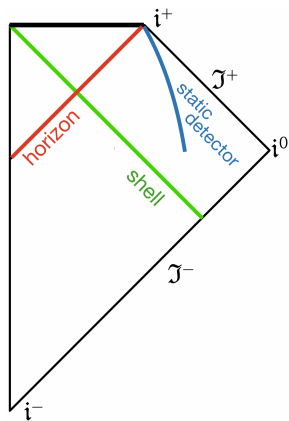
Future:

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Onset of  
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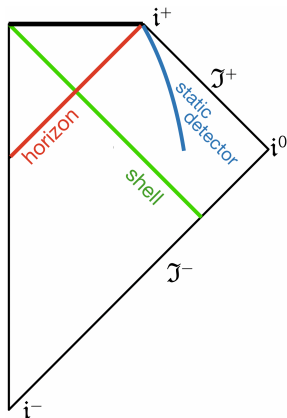
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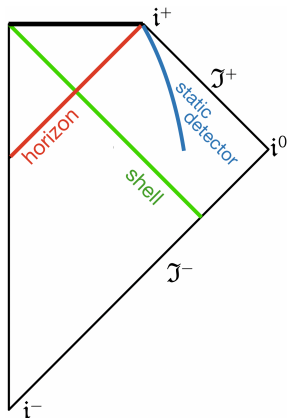
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Late time state approaches the Unruh state on Kruskal (stationary outgoing flux)



Massless field

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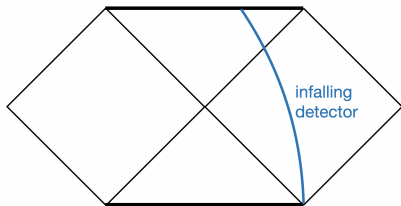
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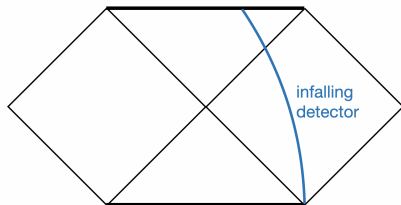
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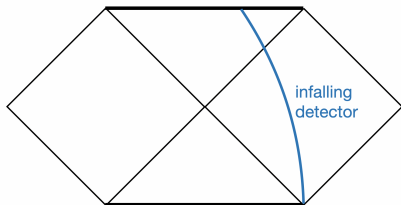


**Vacua** (stationary wrt Killing time)

- ▶ **Boulware**: Minkowski vacuum at infinity
- ▶ **Hartle-Hawking-Israel**: thermal **equilibrium**  $T_\infty = 1/(8\pi M)$
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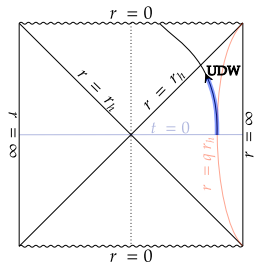
## Outcomes (massless field)

- ▶ At infinity **Boulware** empty, **HHI** thermal, **Unruh** half-thermal
- ▶ **HHI** and **Unruh**:
  - ▶ **(half-)thermality** gradually lost during the fall
  - ▶ Horizon-crossing non-drastric
  - ▶  $\dot{F}_\tau \sim r^{-3/2}$  near the singularity

# Example: 2+1 BTZ infall

Conformal field, **HHI** state

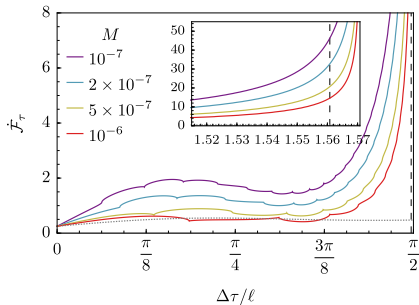
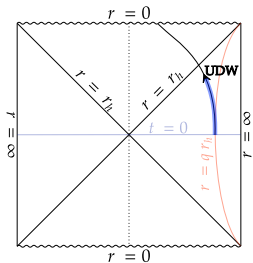
**Spinless** Hodgkinson and JL 2012, Preciado-Rivas et al 2024



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Conformal field, **HHI** state

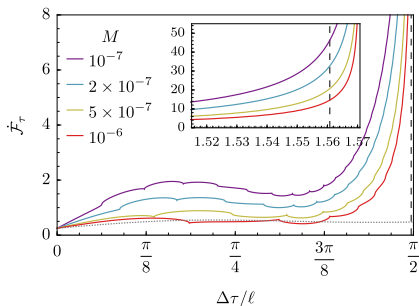
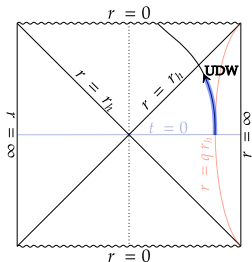
**Spinless** Hodgkinson and JL 2012, Preciado-Rivas et al 2024



## Example: 2+1 BTZ infall

Conformal field, **HHI** state

**Spinless** Hodgkinson and JL 2012, Preciado-Rivas et al 2024

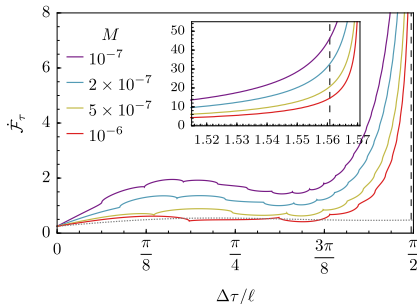
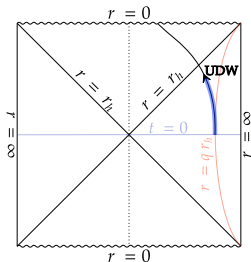


- ▶  $\dot{F}_\tau$  has discontinuous derivative ("glitch") when switch-off moment is on the lightcone of the switch-on moment
- ▶ Happens already outside the horizon!

## Example: 2+1 BTZ infall

Conformal field, **HHI** state

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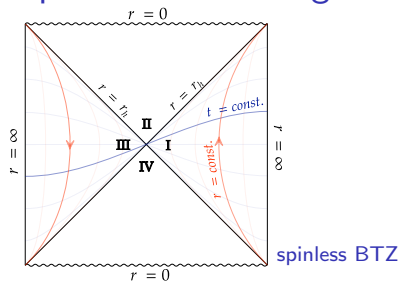


- ▶  $\dot{F}_\tau$  has discontinuous derivative (“glitch”) when switch-off moment is on the lightcone of the switch-on moment
- ▶ Happens already outside the horizon!

**Spinning** Wang et al 2024

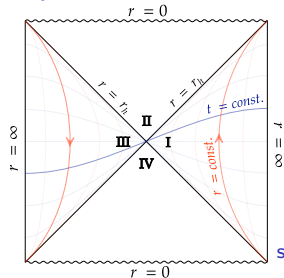
- ▶ **HHI** state has singularities beyond the inner horizon Steif 1994  
⇒ severe glitches there

# Example: infall in a single-exterior BTZ hole Spadafora et al 2024





# Example: infall in a single-exterior BTZ hole Spadafora et al 2024

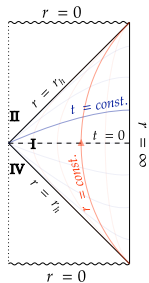


spinless BTZ

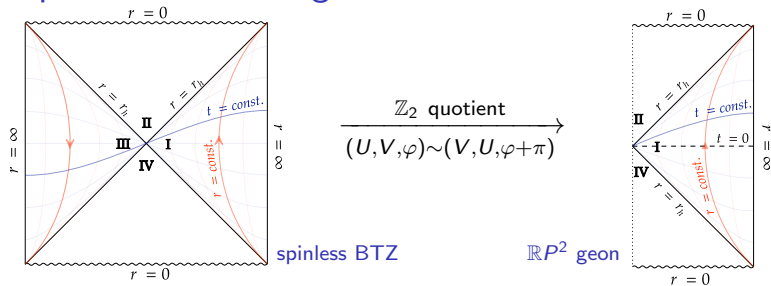
$$\xrightarrow{\mathbb{Z}_2 \text{ quotient}}$$

$$(U, V, \varphi) \sim (V, U, \varphi + \pi)$$

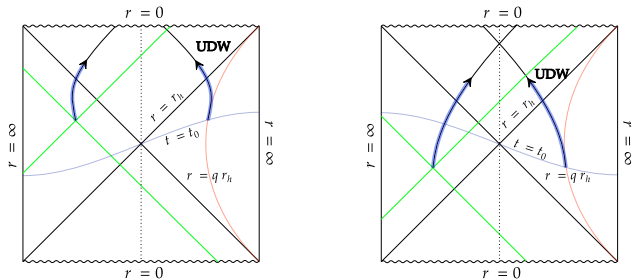
$\mathbb{RP}^2$  geon



# Example: infall in a single-exterior BTZ hole Spadafora et al 2024



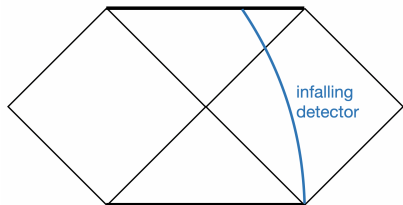
Infalling detector configurations on the geon:



► New glitches after horizon-crossing

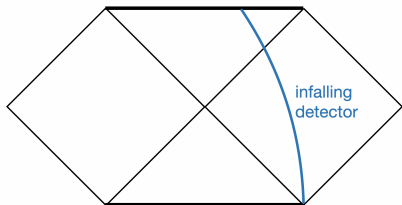
# Example: 3+1 Schwarzschild infall

Ng et al 2022; Shallue and Carroll 2025



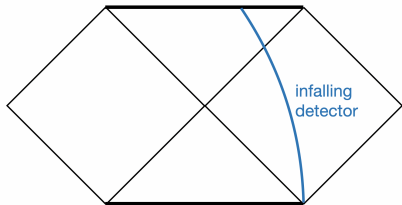
# Example: 3+1 Schwarzschild infall Ng et al 2022; Shallue and Carroll 2025

$\dot{F}_\tau$  numerically difficult!



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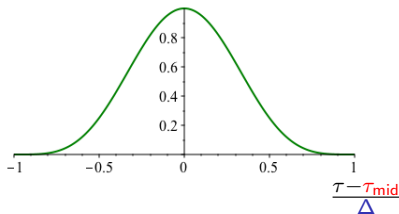


Consider  $F_\chi$  with

$$\chi(\tau) = \begin{cases} \cos^4\left(\frac{\pi(\tau - \tau_{\text{mid}})}{2\Delta}\right) & \text{for } |\tau - \tau_{\text{mid}}| < \Delta \\ 0 & \text{otherwise} \end{cases}$$

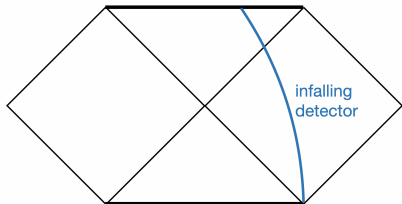
$\Delta$ : half the interaction duration

$\tau_{\text{mid}}$ : interaction interval midpoint



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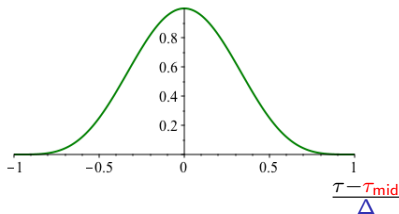


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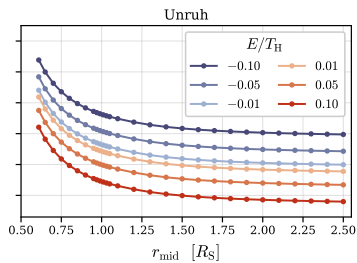
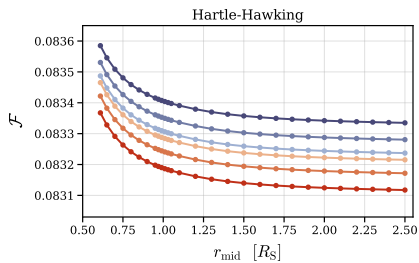
$\tau_{\text{mid}}$ : interaction interval midpoint



- ▶ Finite duration,  $C^3$
- ▶ Close to Gaussian

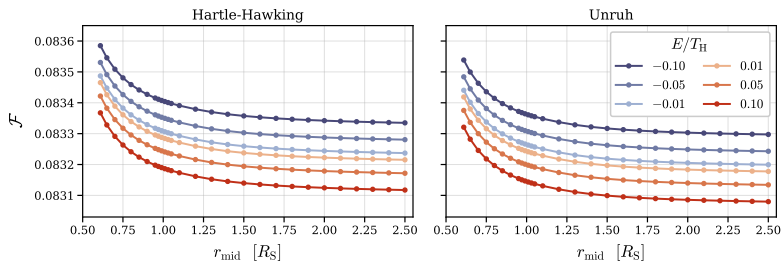
# 3+1 Schwarzschild infall (cont'd) Shallue and Carroll [arXiv:2501.06609]

$F_{\chi}$



# 3+1 Schwarzschild infall (cont'd) Shallue and Carroll [arXiv:2501.06609]

$F_\chi$



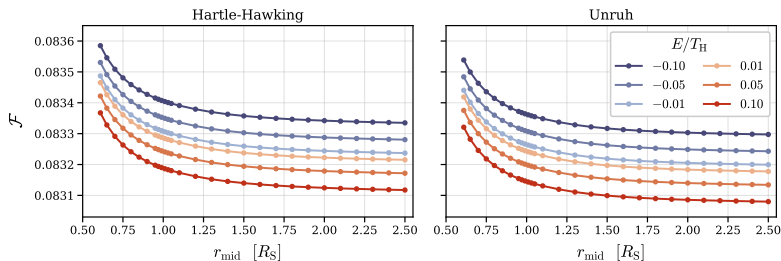
$\Delta \lesssim M$  needed to resolve timescales near and beyond the horizon

$\Rightarrow F_\chi$  dominated by switching effects



# 3+1 Schwarzschild infall (cont'd) Shallue and Carroll [arXiv:2501.06609]

$F_\chi$



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**Notion of 'effective temperature' during the infall continues to be under debate!**