



Dark sector search at BESIII

Xiaoxuan Ding

Peking University

On behalf of the BESIII collaboration

Apr 11, Genova, Italy

BESIII



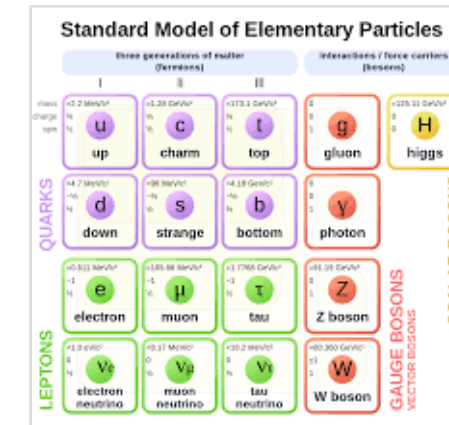
北京大学
PEKING UNIVERSITY

LDMA'25



Dark sector

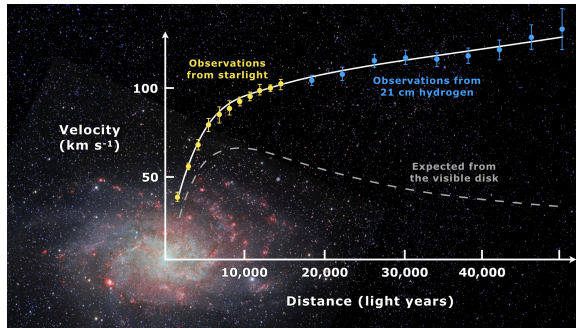
There is nothing new to be discovered in physics now.



Dark sector

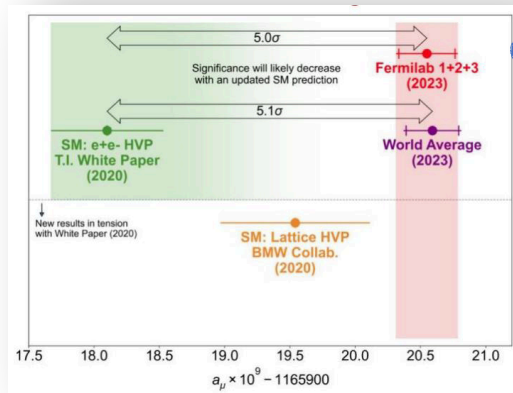
There is nothing new to be discovered in physics now.

Galactic rotation curve



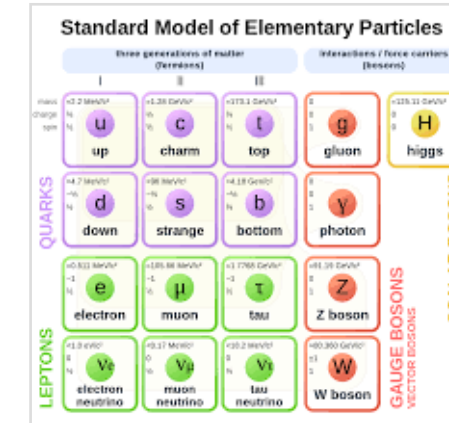
arXiv:astro-ph/0403324

$g_\mu - 2$ anomaly



PRL 131, 161802(2023)

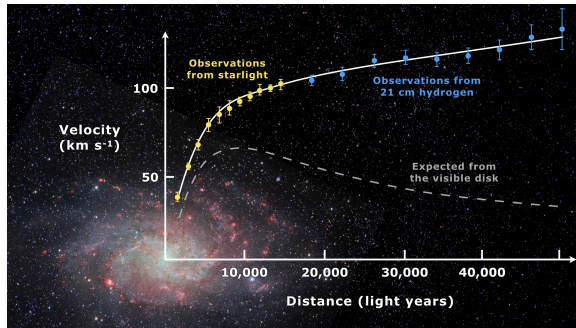
Xiaoxuan Ding



Dark sector

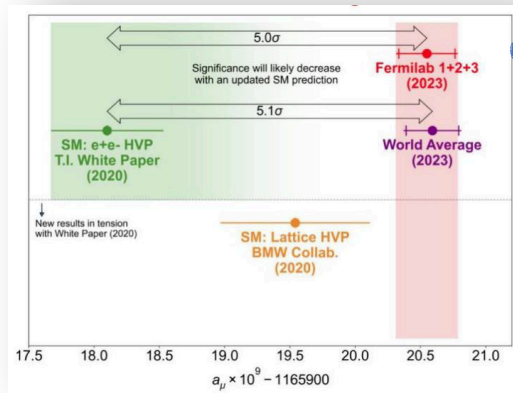
There is nothing new to be discovered in physics now.

Galactic rotation curve



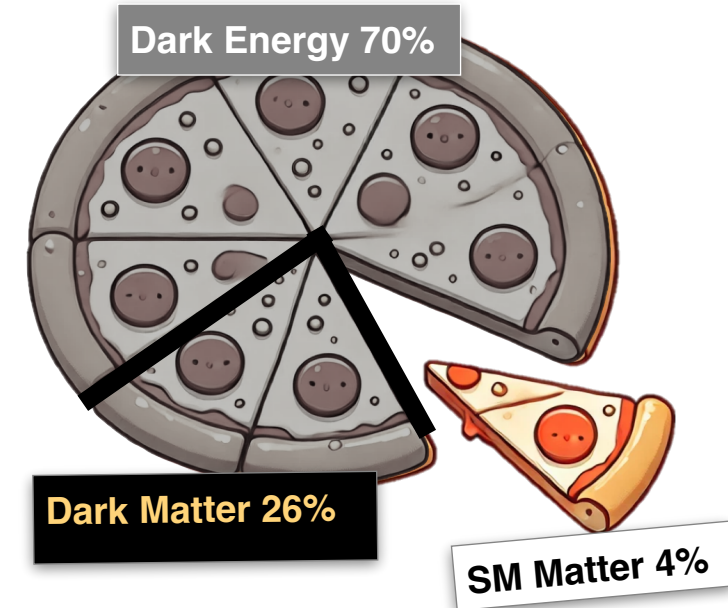
arXiv:astro-ph/0403324

$g_\mu - 2$ anomaly



PRL 131, 161802(2023)

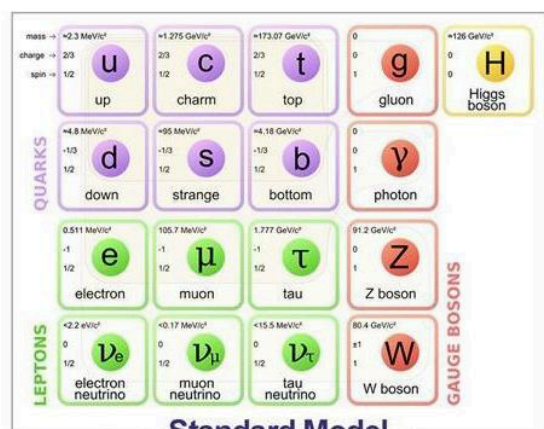
Xiaoxuan Ding



Physics BSM must exist!

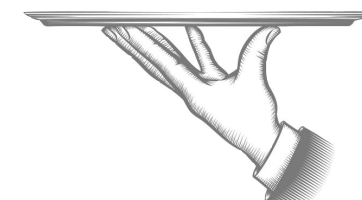
Dark sector

- A **collection of particles** that are not charged directly under the SM strong, weak, or electromagnetic forces.
 - May interact with SM particle through **portal** interactions, not through gravitational effect only.



The main of
this talk

- **Invisible decay of K_S^0**
arXiv: 2501.06426 accepted by JHEP
- **$\Sigma^+ \rightarrow p + \text{invisible}$**
Phys.Lett.B 852 (2024) 138614
- **Muonphilic particle in $J/\psi \rightarrow \mu^+ \mu^- X$**
Phys.Rev.D 109 (2024) 3, L031102
- **(Massless) Dark photon in $D^0 \rightarrow \omega \gamma', \gamma \gamma'$**
Phys.Rev.D 111 (2025) 1, L011103



➡ Dark sector particles

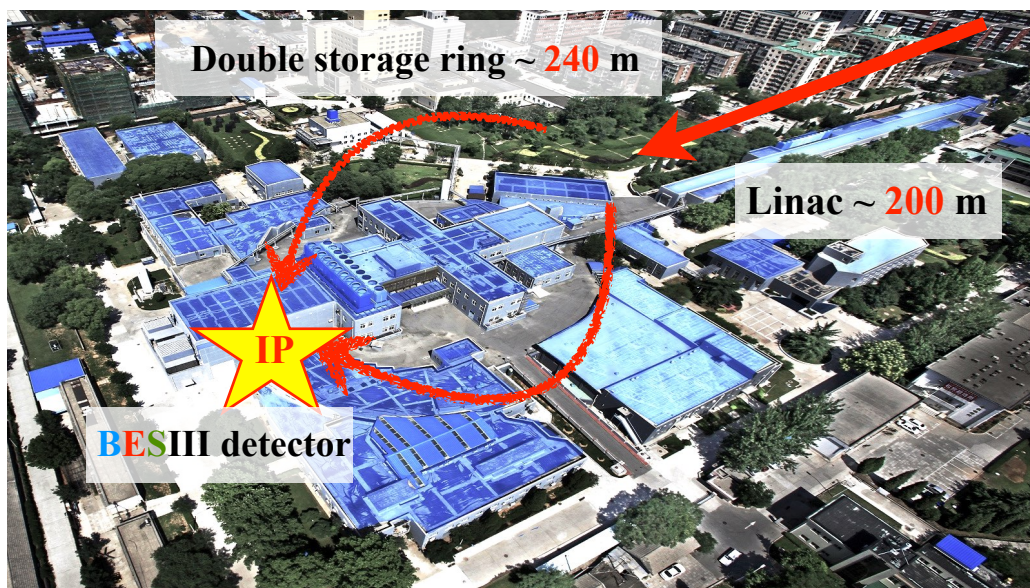
- **New invisible particles**
- **New particle decay to SM particles**
- **.....**

BEPCII

○ A symmetric e^-e^+ collider running at **tau-charm region** | $E_{CM} = 1.84 \sim 4.95 \text{ GeV}$ |

- BEPCII: Electron-positron colliders: accelerate the e^+, e^-

✓ **Beijing Electron
Positron Collider II**



- Double rings
- $E_{CM} = 1.84 \sim 4.95 \text{ GeV}$
- Peak luminosity @ $E_{cm}=3.77 \text{ GeV}$
 $\mathcal{L}_{\text{peak}} = 1.1 \times 10^{33} \text{ cm}^{-2} \cdot \text{s}^{-1}$
- 2009~today: BESIII physics runs;

BESIII

BEijing Spectrometer **III**

*Can't do B physics



END

12 KM

BESIII detector

Merits of BSM search at BESIII:

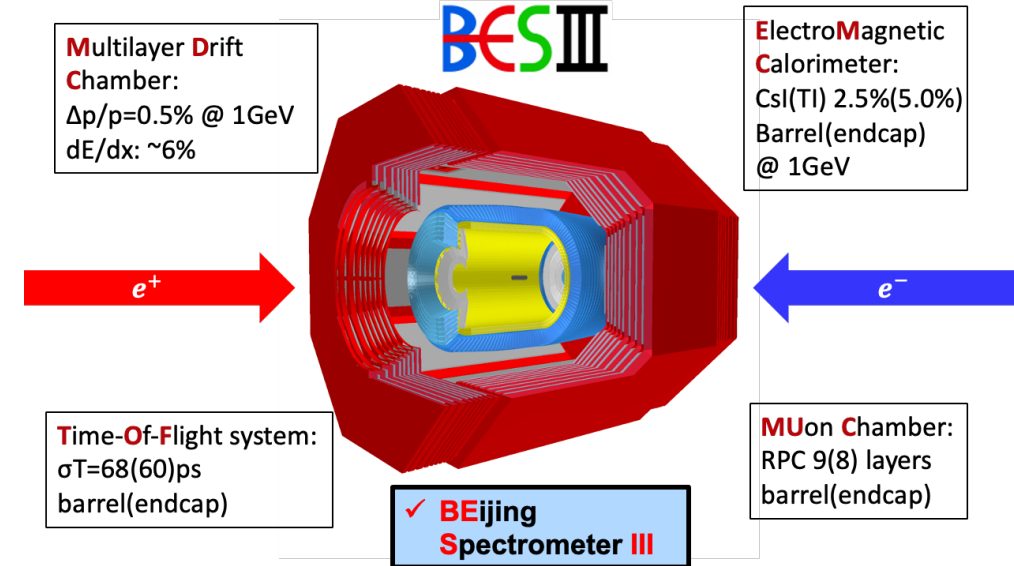
- High statistic of data
- Clean background environment
- Angular coverage is almost 4π
- Data collected at the **threshold**

*Also mentioned by
Enrico Graziani*

Sensitive to invisible signal!

BESIII has collected large data samples in τ -c energy region! which can benefit the search for the LDM in the **MeV-GeV** range

$$E_{CM} = 1.84 \sim 4.95 \text{ GeV}$$





Invisible decay of K_S^0

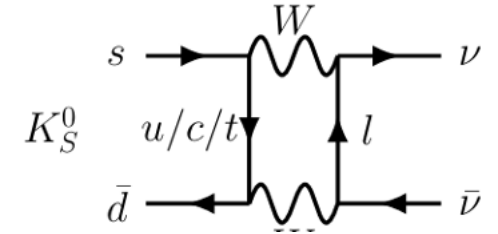
Light Dark Matter at
Accelerators (LDMA) 2025

WCCFGLRGLR (EDWW) 5052



Phys.Rev.D 91 (2015) 1, 015004

FCNC && helicity suppression, $\text{BF} < 10^{-16}$

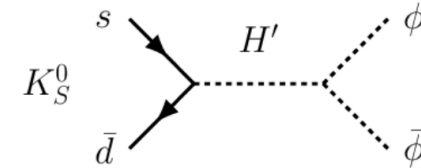
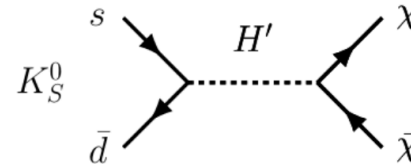


SM decay

$$K_S^0 \rightarrow \nu \bar{\nu}$$

Decay to DM

$$K_S^0 \rightarrow \chi \bar{\chi}$$



2HDM model
 $\text{BF} \sim \mathcal{O}(10^{-6})$

Natural Sci.Rev. 1 (2024) 5

**Ordinary-mirror
particle oscillations**

$$K_S^0 \leftrightarrow K_S^{0'}$$

Mirror matter model
 $\text{BF} \sim \mathcal{O}(10^{-6})$

arXiv: 2006.10746

Input for CPT test

Bell-Steinberger relation **connects** CPTV to the amplitudes of all decay channels of neutral kaons. BUT currently assumes no invisible modes

Phys.Rev.D 91 (2015) 1, 015004



Phys.Rev.D 91 (2015) 1, 015004

FCNC & helicity suppression, $\text{BF} < 10^{-16}$

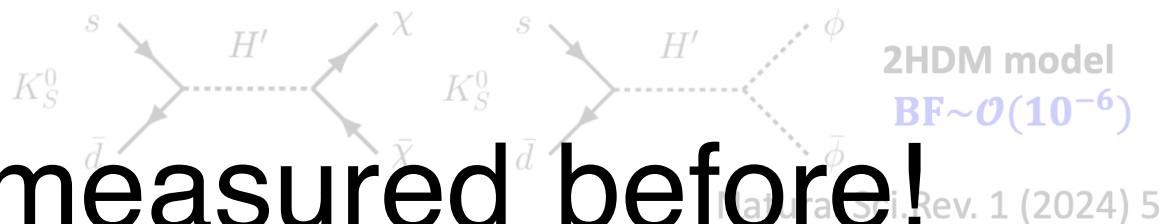
Decay to SM

$$K_S^0 \rightarrow \nu \bar{\nu}$$

Decay to DM

$$K_S^0 \rightarrow \chi \bar{\chi}$$

Never been measured before!



Phys.Rev. 1 (2024) 5

$K_S^0 \rightarrow K_S^{0'}$ Mirror matter model
 $\text{BF} \sim \mathcal{O}(10^{-6})$ arXiv: 2006.10746

Ordinary-mirror
particle oscillations

Input for CPT test

Bell-Steinberger relation connects CPTV to the amplitudes of all decay channels of neutral kaons. BUT currently assumes no invisible modes

Phys.Rev.D 91 (2015) 1, 015004

K_S^0



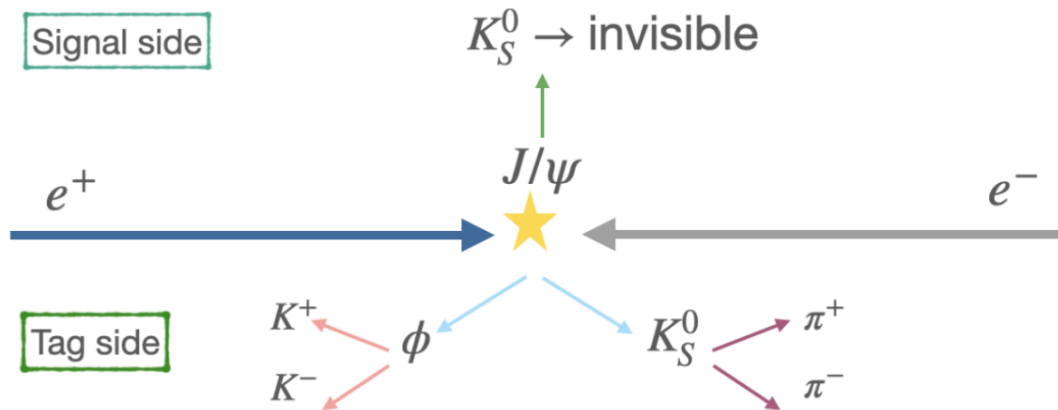
Invisible



Analysis strategy

K_S^0 source

- $J/\psi \rightarrow \phi K_S^0(\text{tag}) K_S^0, K_S^0(\text{tag}) \rightarrow \pi^+ \pi^-$ from **10 Billion** J/ψ events



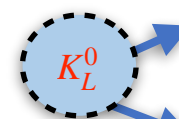
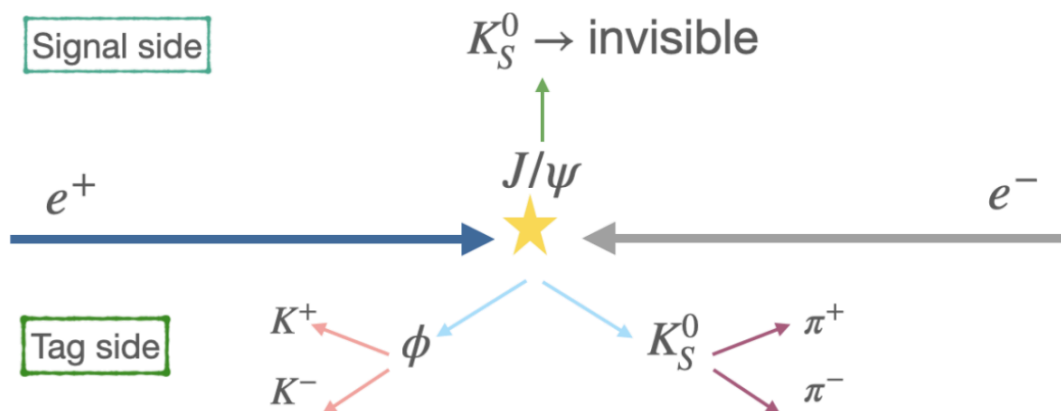
Why this channel?



Analysis strategy

K_S^0 source

- $J/\psi \rightarrow \phi K_S^0(\text{tag}) K_S^0, K_S^0(\text{tag}) \rightarrow \pi^+ \pi^-$ from 10 Billion J/ψ events



Long life, mostly fly through the detector

Rarely interact with the detector

Can mimic the fake signal!



Tag a clean K_S^0 sample:

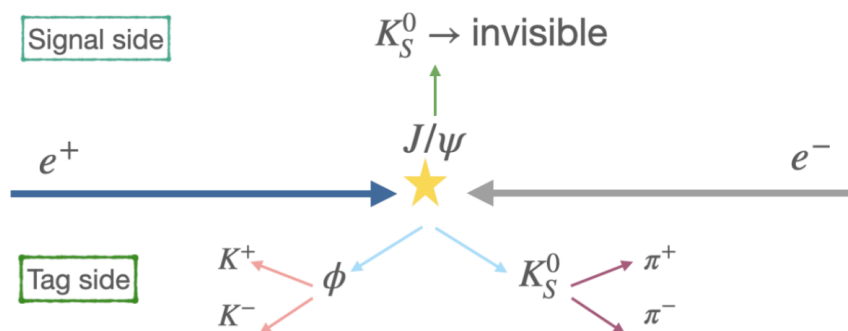
- One of the dominant K_L^0 backgrounds, $J/\psi \rightarrow \phi K_S^0 K_L^0$ is **forbidden** by C-parity conservation
- But still have $J/\psi \rightarrow K^+ K^- K_S^0 K_L^0$



Analysis strategy

Tag a clean K_S^0 sample:

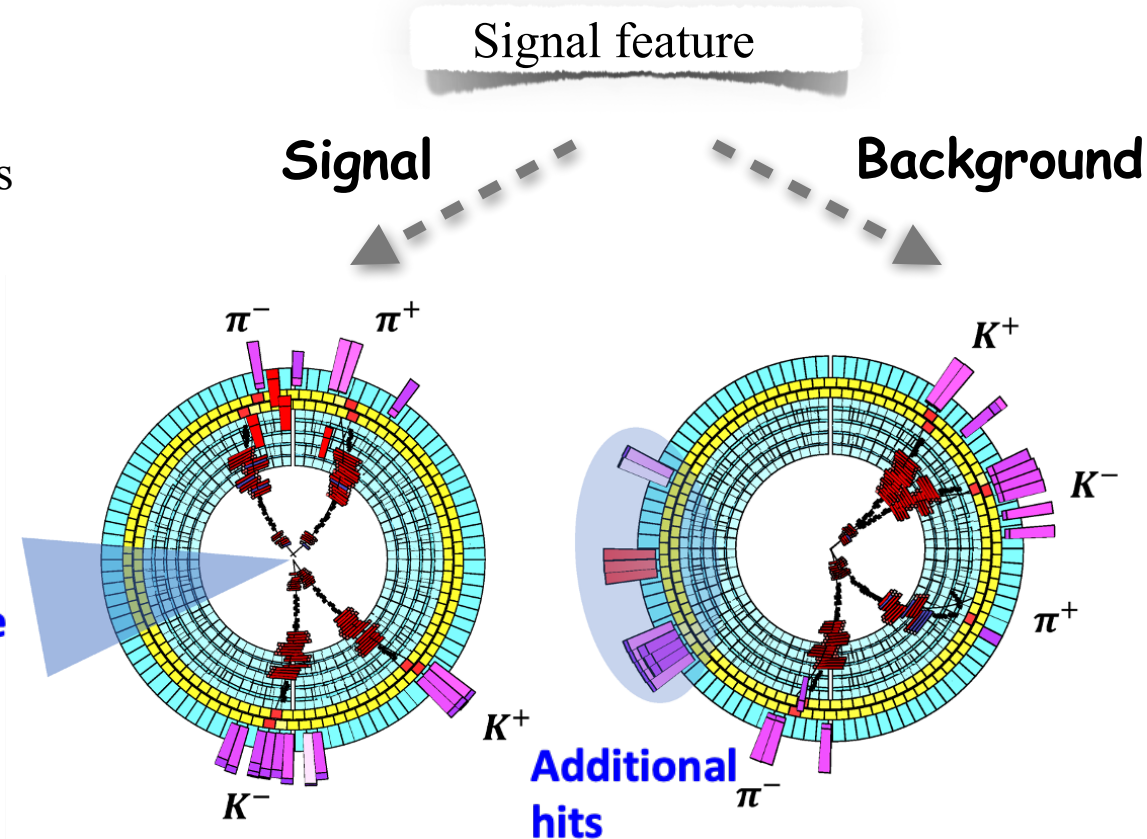
- $J/\psi \rightarrow \phi K_S^0(\text{tag}) K_S^0$, $K_S^0(\text{tag}) \rightarrow \pi^+ \pi^-$ from 10 Billion J/ψ events



- $J/\psi \rightarrow \phi K_S^0 K_L^0$ is forbidden by C-parity conservation, **lower background**

- But still have $J/\psi \rightarrow K^+ K^- K_S^0 K_L^0$

Invisible in EMC

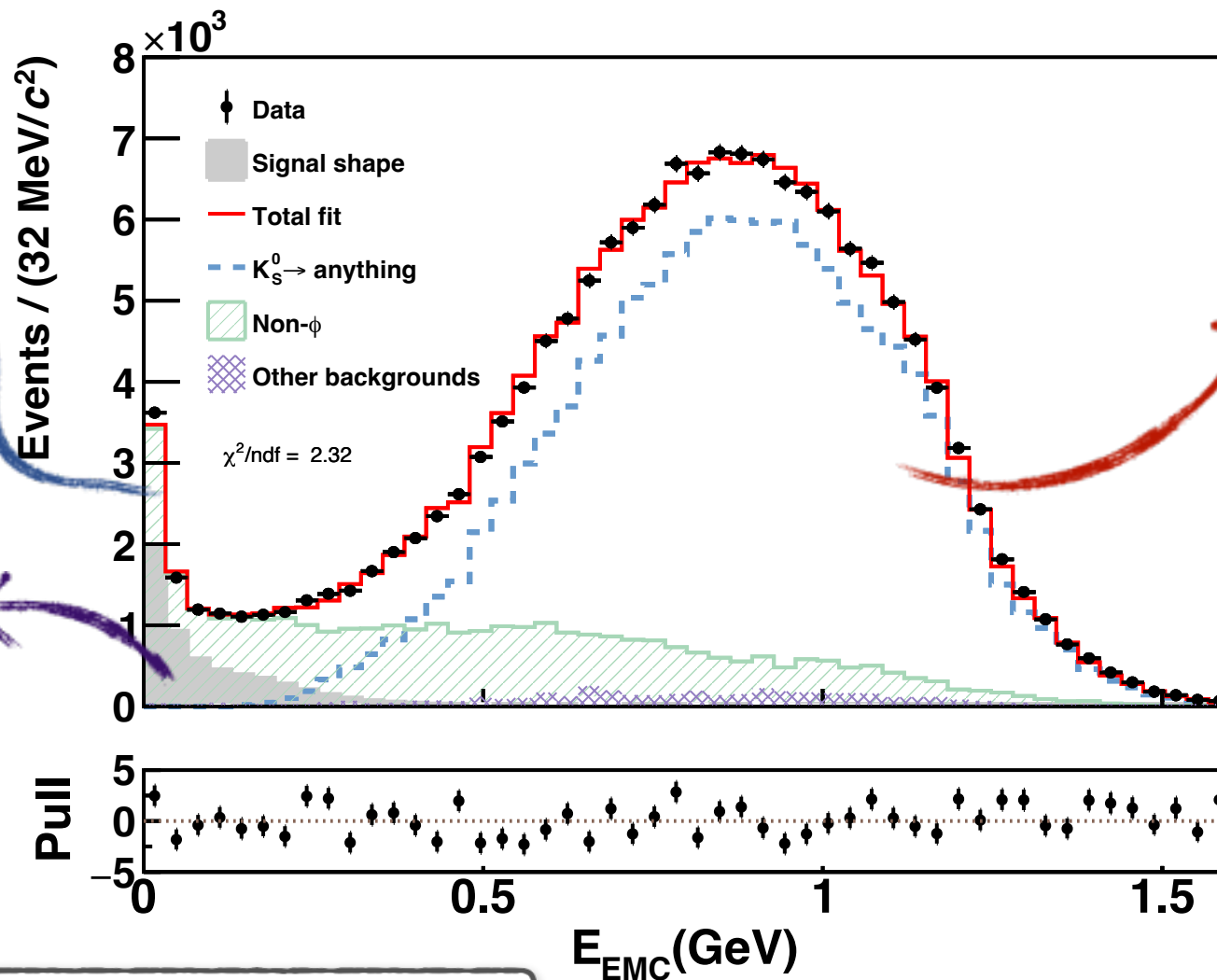


Using the deposited energy in **EMC**(energy calorimeter) to identify the invisible signal

Signal extraction

Peaking background from non- ϕ processes,
 $J/\psi \rightarrow K^+ K^- K_S^0 K_L^0$,
 Shape from ϕ sideband

Invisible Signal peaks around zero



Other background modeled with MC simulation, such as $K_S^0 \rightarrow \pi^0 \pi^0$

$N_{\text{invisible}} = 56 \pm 201$,
 $\mathcal{B}(K_S^0 \rightarrow \text{invisible}) < 8.4 \times 10^{-4}$
 (90% C.L.)

Accepted by JHEP
 during the LDMA2025

First direct measurement of
 $K_S^0 \rightarrow \text{invisible}$!

the UL still lies above the NP prediction





$\Sigma^+ \rightarrow p + \text{invisible}$

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Accelerators (LDMA) 2025

WCCF6187012 (EDMW) 5052



ROPP. 86, 016201 (2023)

Decay to SM

$$s \rightarrow d \nu \bar{\nu}$$

FCNC && GIM suppression, $\text{BF} < 10^{-11}$

Contributions from BSM can enhance the BF ($\sim 10^{-4}$)

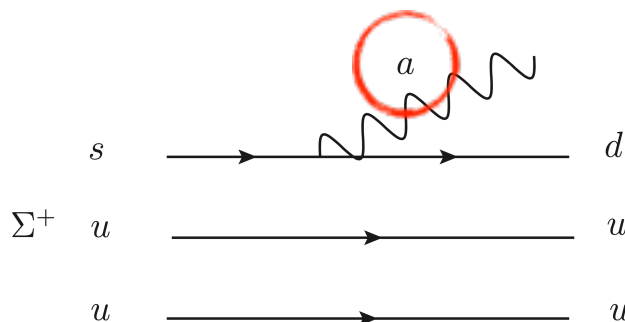
JHEP 04, 104 (2019).

Decay to BSM particle

- **Massless** dark photon γ' && QCD axion, a

$$m_a \propto 1/f_a \rightarrow m_a \ll 1\text{eV}$$

Phys. Rev. D 102, 015023 (2020).



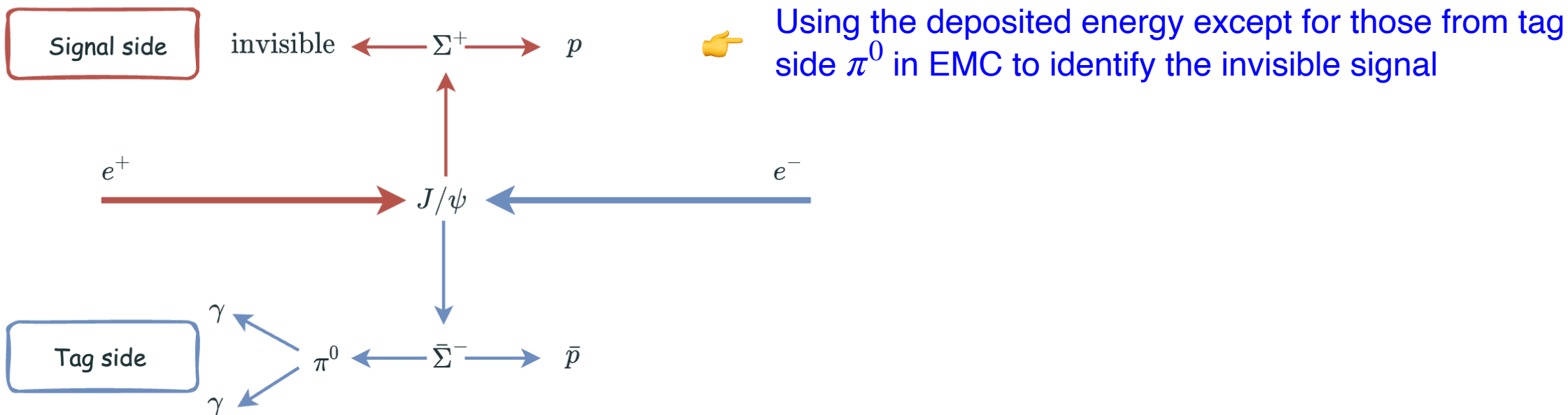
$$\Gamma(\Sigma^+ \rightarrow pa) = \frac{M_{\Sigma^+}^3}{16\pi} \left(1 - \frac{M_p^2}{M_{\Sigma^+}^2}\right)^3 \left(\frac{(-1)^2}{|F_{sd}^V|^2} + \frac{0.34^2}{|F_{sd}^A|^2} \right)$$

Our measurement can set limits on axion-fermion effective decay constants

Analysis strategy

- 10B J/ψ events, $J/\psi \rightarrow \Sigma\bar{\Sigma}$ ($\sim 10^7 \Sigma^+ \bar{\Sigma}^-$ pairs), double tag method

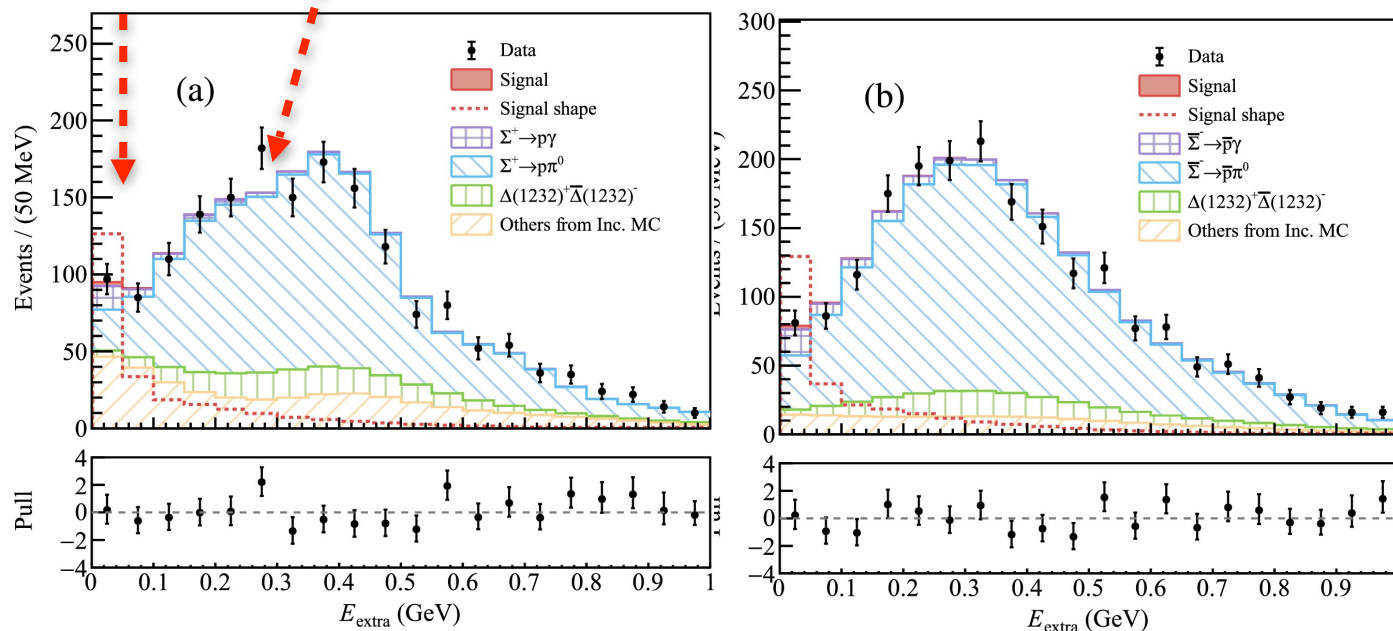
Kinematic fit to constraint the invisible axion mass to **zero**



Signal extraction

Background modeled with MC simulation, such as $\Sigma^+ \rightarrow p\pi^0$

Invisible Signal peaks around zero



$\text{BF}(\Sigma^+ \rightarrow p + \text{invisible}) < 3.2 \times 10^{-5} @ 90\% \text{ CL.}$

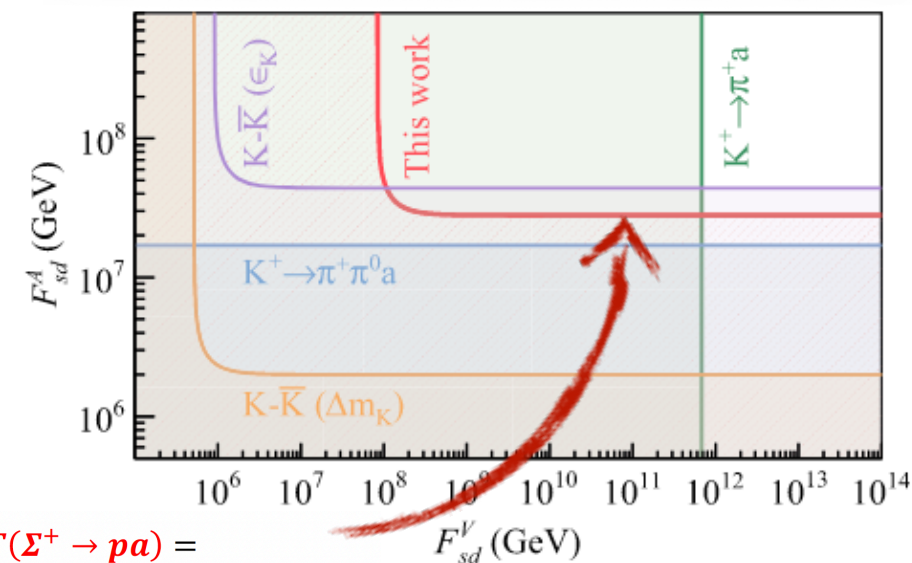
- Our result lies below the maximum BF allowed for $\Sigma^+ \rightarrow p + \text{invisible}$ in certain scenarios, under the hypothesis of massless γ'

$$\rightarrow 3.8 \times 10^{-5}$$

Phys. Rev. D 101, 035044 (2020)

- Competitive limit on the axial-vectorial part of axion-fermion effective decay constant F_{sd}^A :

$$\rightarrow F_{sd}^A > 2.8 \times 10^7 \text{ GeV}$$



$$\Gamma(\Sigma^+ \rightarrow pa) = \frac{M_{\Sigma^+}^3}{16\pi} \left(1 - \frac{M_p^2}{M_{\Sigma^+}^2}\right)^3 \left(\frac{(-1)^2}{|F_{sd}^V|^2} + \frac{0.34^2}{|F_{sd}^A|^2} \right)$$





Muon-philic particle

Light Dark Matter at
Accelerators (LDMA) 2025

Accelerators (LDMA) 2025

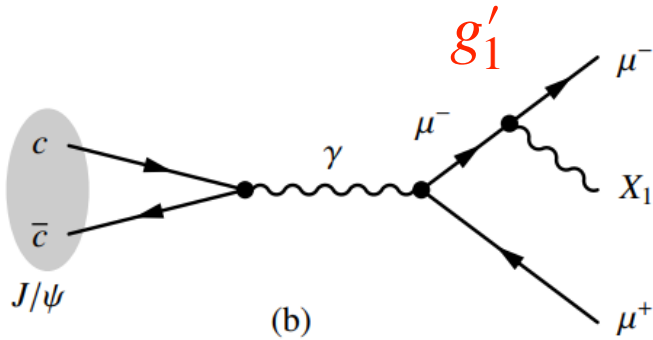
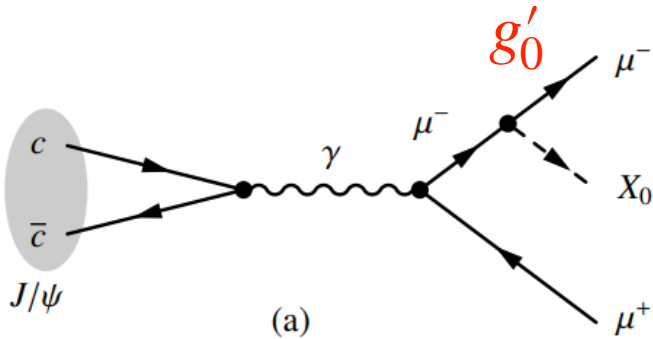
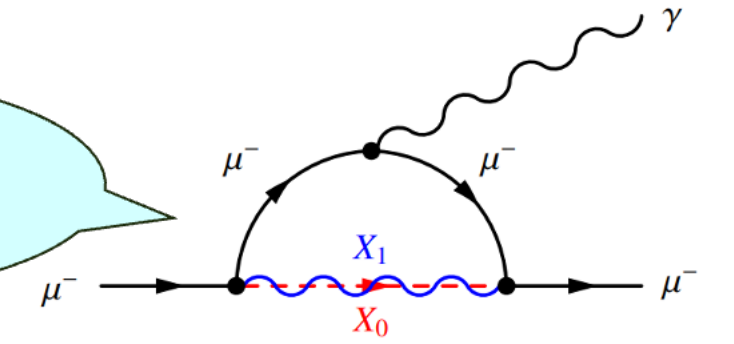


- $U(1)_{L_\mu-L_\tau}$ model : a new massive **vector boson X_1 or scalar boson X_0 only**
couple to the second or third generations of leptons ($\mu, \nu_\mu, \tau, \nu_\tau$) with the coupling strength g'_1, g'_0

➤ **The main muon source at BESIII:**
 $J/\psi \rightarrow \mu^+ \mu^-$

➤ **Search for $J/\psi \rightarrow \mu^+ \mu^- X_{0,1}$** → **Invisible $X_{0,1}$ at this time**

explain $(g - 2)_\mu$ anomaly

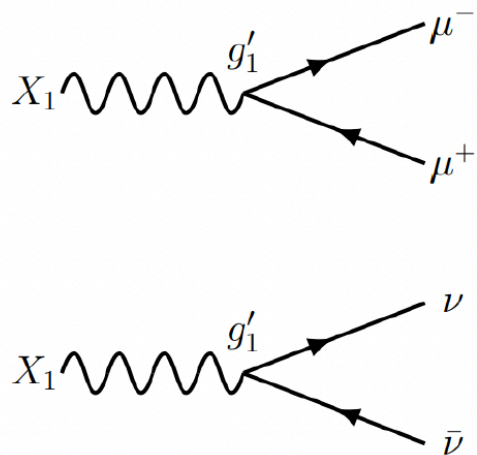


JHEP10(2020)207



Three cases of muon-philic particles

“vanilla” $L_\mu - L_\tau$ model

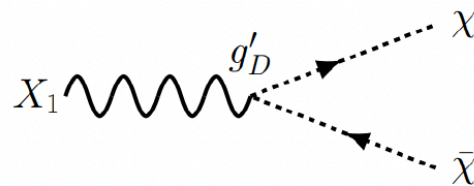


Large mass of dark matter kind:

$$m_\chi > m_{\chi_1}/2$$

- $\mathcal{B}(X_1 \rightarrow \nu\bar{\nu}) = 33 \sim 100\%$ with different m_{χ_1}

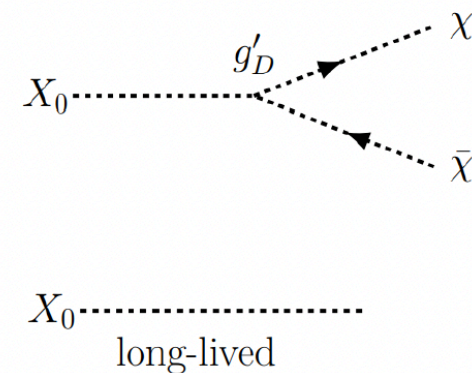
“invisible” $L_\mu - L_\tau$ model



Light dark matter kind $m_\chi < m_{\chi_1}/2$

- $g_D \gg g_{1'}$
- $\mathcal{B}(X_1 \rightarrow \chi\chi) \sim 100\%$

“scalar” $U(1)$ model



- Assume the X_0 is long lived or only decay to invisible final states

Search for X_0, X_1 via $J/\psi \rightarrow \mu^+\mu^-X_{0,1}$ with $X_{0,1}$ is invisible



Recoil mass spectrum of $J/\psi \rightarrow \mu^+\mu^-X_{0,1}$

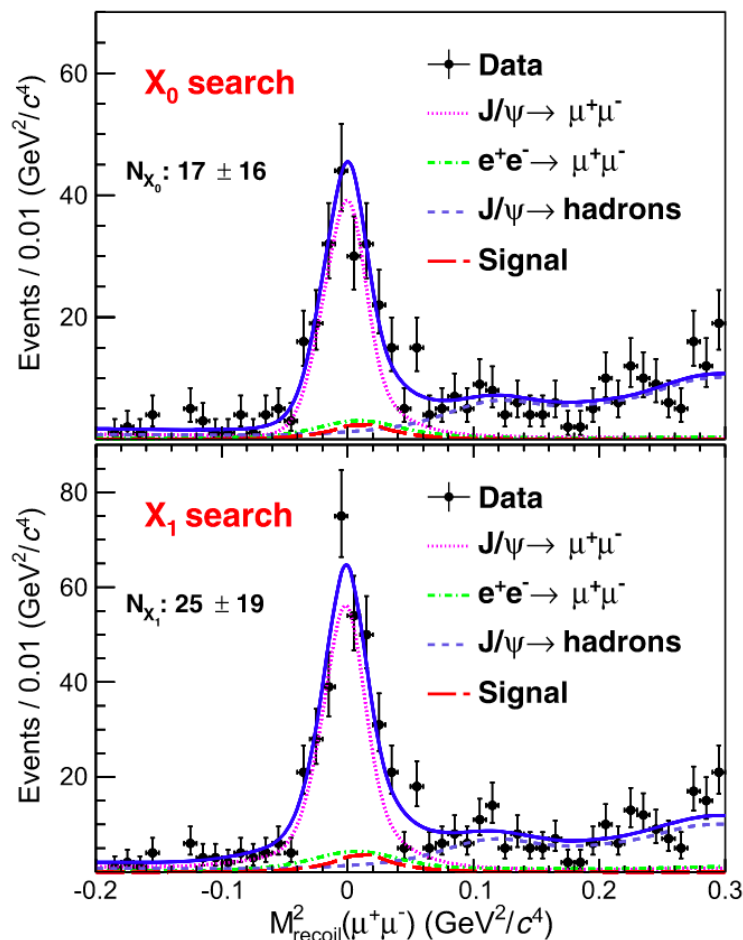
- Data samples: **9B** J/ψ events

58 different $X_{0,1}$ mass values, ranging from 1 to 1000 MeV/c^2 , in step of 10~20 MeV/c^2

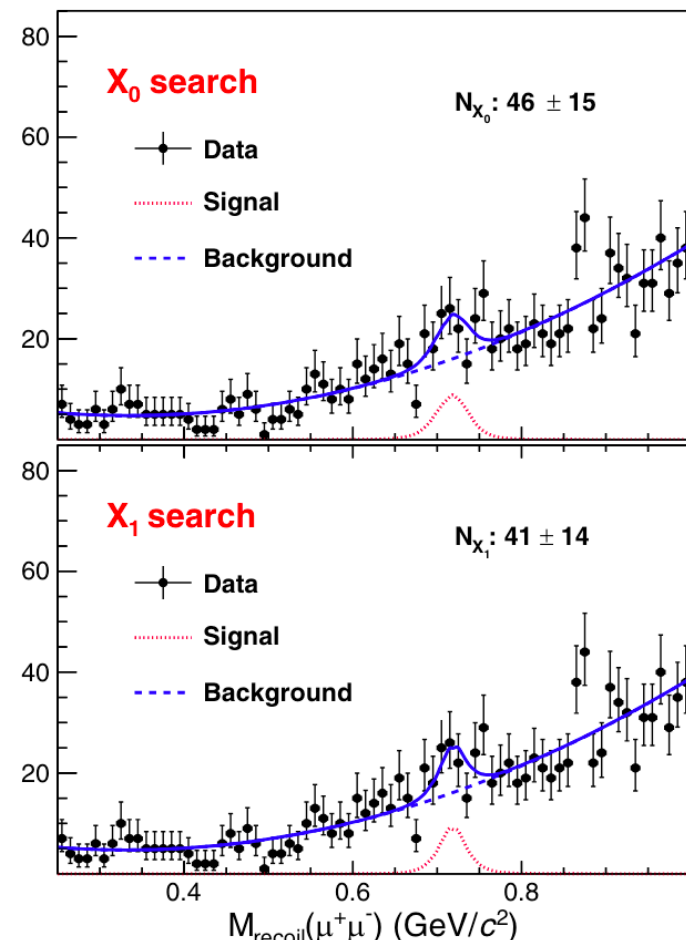
No evidence for $J/\psi \rightarrow \mu^+\mu^- + X_{0,1}$ signals

- The maximum local significance is 2.5σ at $M(X_{0,1}) = 720 \text{ MeV}/c^2$

■ **Low mass region**, with $M(X_{0,1}) = 120 \text{ MeV}/c^2$

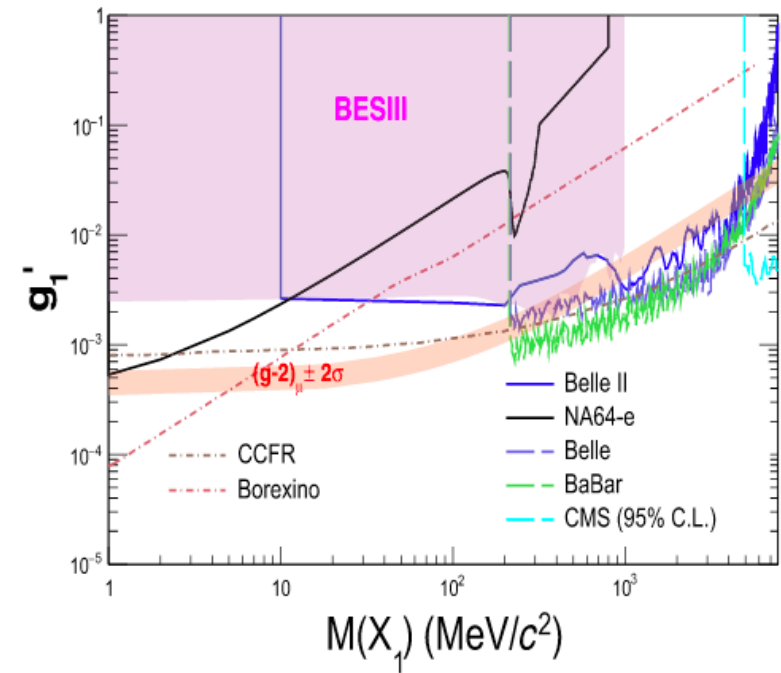


■ **High mass region**, with $M(X_{0,1}) = 720 \text{ MeV}/c^2$



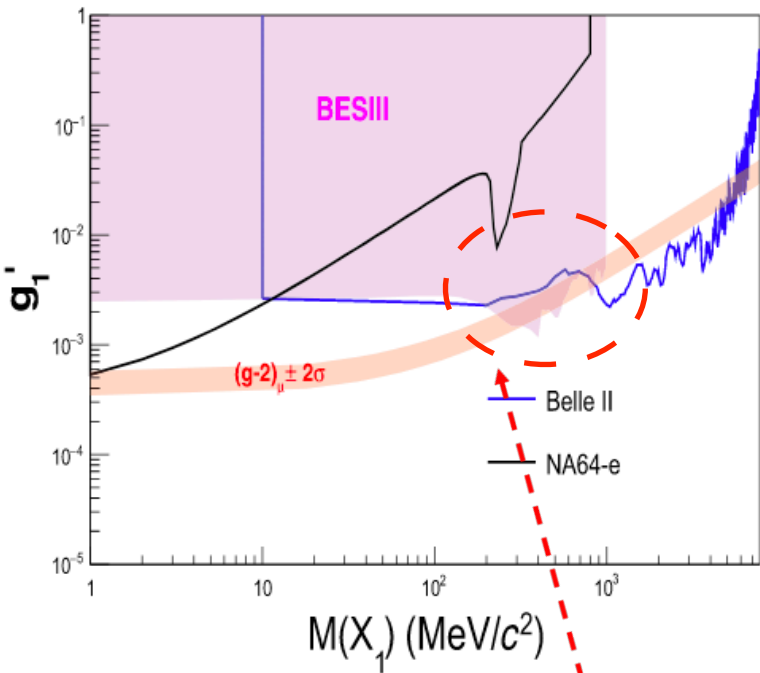
Coupling constraint

“vanilla” $L_\mu - L_\tau$ model



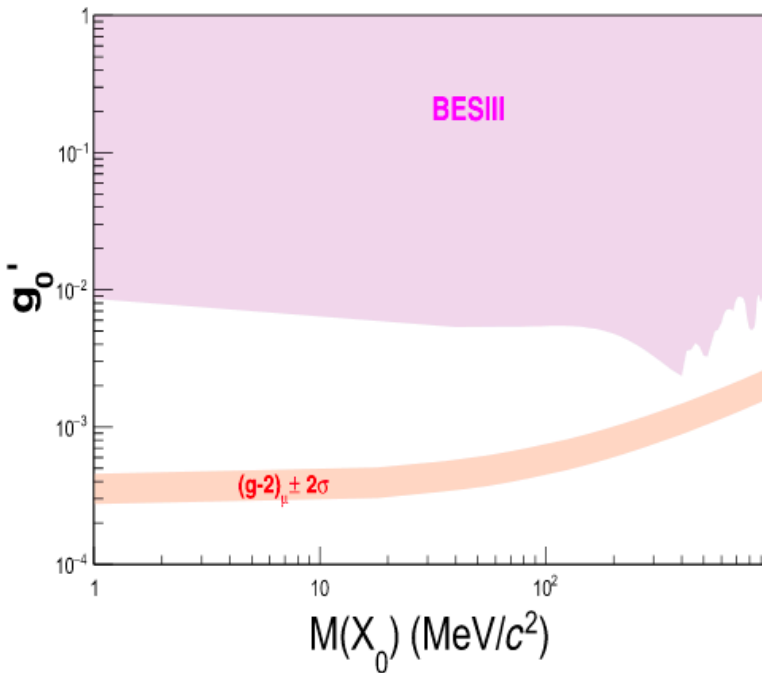
BarBar, CMS, Belle : $X_1 \rightarrow \mu^+ \mu^-$
 BelleII and **BESIII**: $X_1 \rightarrow \nu \bar{\nu}$
 (Taking $\mathcal{B}(X_1 \rightarrow \nu \bar{\nu})$ into account)

“invisible” $L_\mu - L_\tau$ model



Better sensitivity in the range 200–860 MeV/c^2 is obtained

“scalar” $U(1)$ model



First constraint for the “scalar” invisible X_0 case .
 Belle II can also give the constraint



Dark photon in $D^0 \rightarrow \omega\gamma', \gamma\gamma'$

Light Dark Matter at
Accelerators (LDMA) 2025

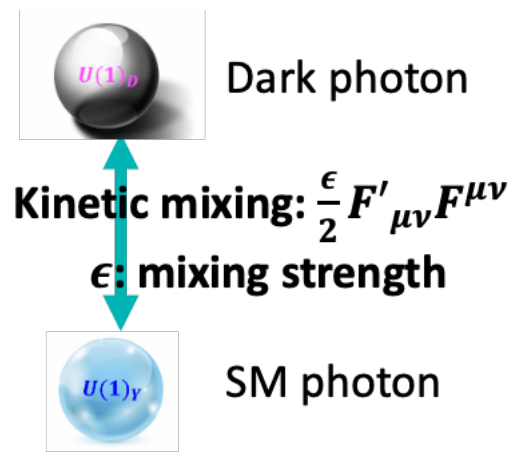
WCCF6187012 (EDMW) 5052



Massless Dark photon

See talks by Fernando Arias Aragon

- A minimal extension to SM, $U(1)_D$, causing the associated spin-one boson, the dark photon



- Massive γ' , if the symmetry is spontaneously **broken**
 - **Massless γ'** , if the symmetry is **unbroken** This work
- 👉 Only couple to SM particles through operators of higher dimension

Dimension-**six**
operator

PRL 94, 151802 (2005)

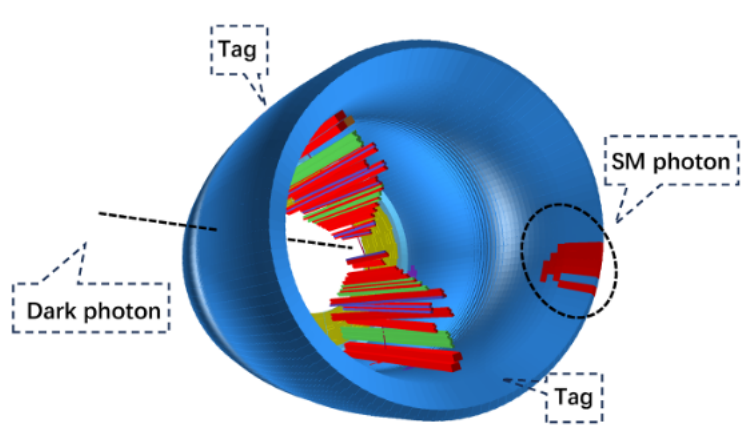
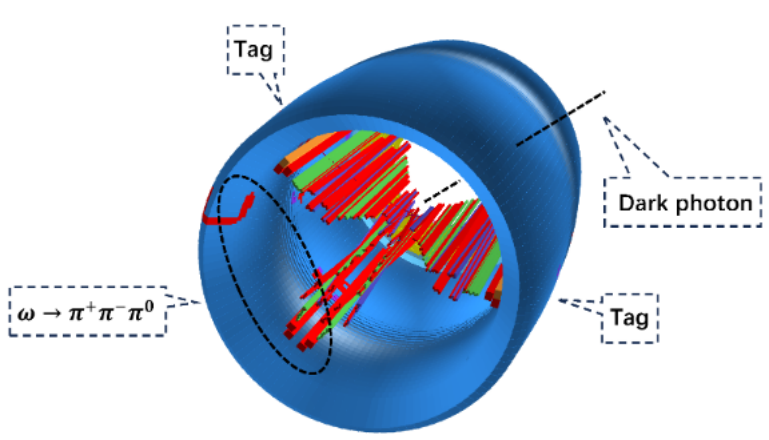
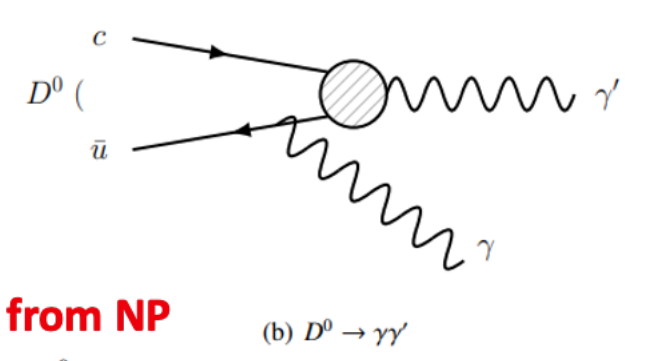
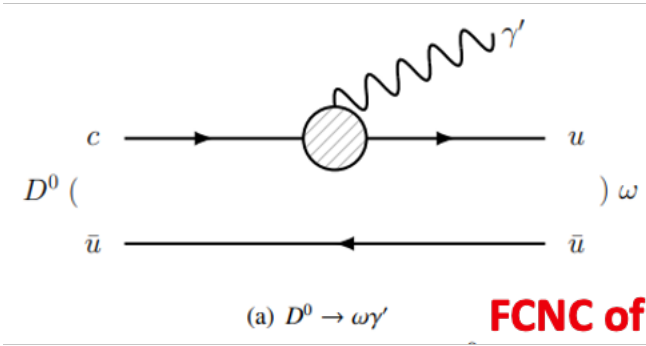
$$\mathcal{L}_{NP} = \frac{1}{\Lambda_{NP}^2} (\underbrace{C_{jk}^U \bar{q}_j \sigma^{\mu\nu} u_k \tilde{H}}_{\text{Up type quarks coupling}} + \underbrace{C_{jk}^D \bar{q}_j \sigma^{\mu\nu} d_k H}_{\text{Down type quarks coupling}} + \underbrace{C_{jk}^L \bar{l}_j \sigma^{\mu\nu} e_k H}_{\text{Charged leptons coupling}} + h.c.) \bar{F}_{\mu\nu}$$

Massless dark photon

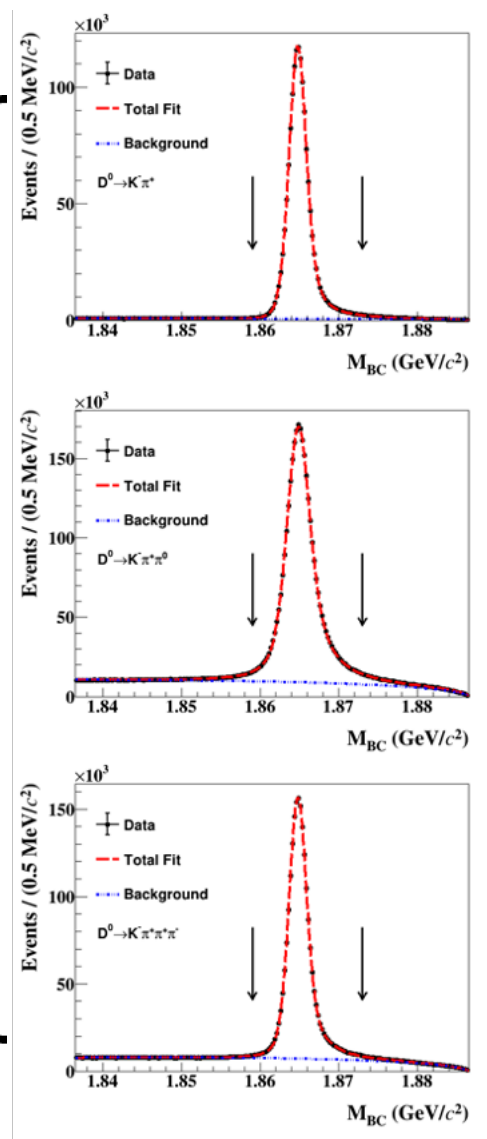
- Experimental** search:
- $H \rightarrow \gamma\gamma'$
 - $\mu \rightarrow e\gamma'$
 - $\Lambda_c^+ \rightarrow p\gamma'$
- } •No signal observed

Search for $D^0 \rightarrow \omega \gamma'$ and $D^0 \rightarrow \gamma \gamma'$

- Double tag method: $\sim 6 \times 10^6 D^0(\bar{D}^0)$ are tagged with 7.9fb^{-1} data @3.77 GeV
- The massless dark photon is invisible

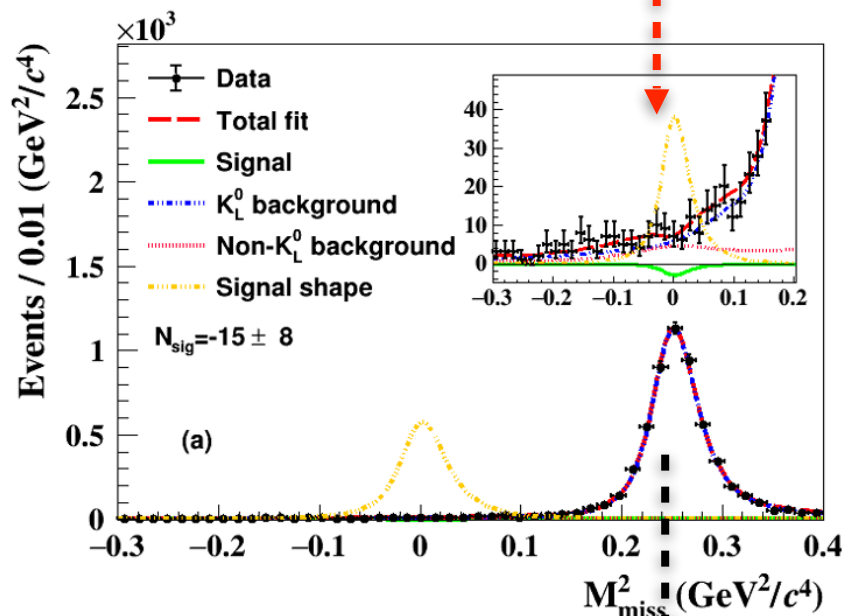


Tag



Signal extraction

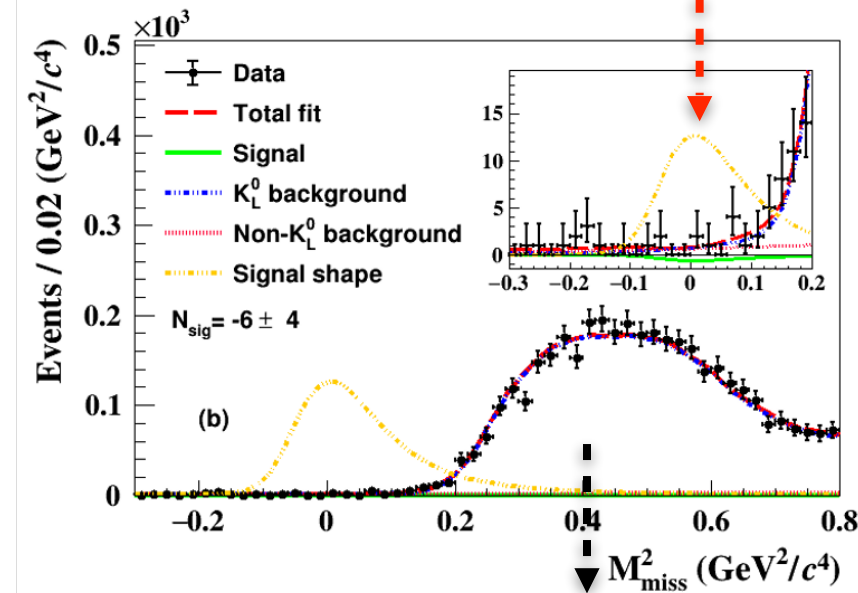
Signal peaks around
zero



$D^0 \rightarrow \omega K_L^0$ background

Signal extraction of $D^0 \rightarrow \omega \gamma'$

Signal peaks around
zero



$D^0 \rightarrow \pi^0 K_L^0$ background

Signal extraction of $D^0 \rightarrow \gamma \gamma'$



Coupling constraint

This work {

$$\mathcal{B}(D \rightarrow V\gamma') = \frac{\tau_D f_{DV}^2 (m_D^2 - m_V^2)^3}{2\pi m_D^3} (|\mathbb{C}|^2 + |\mathbb{C}_5|^2)$$

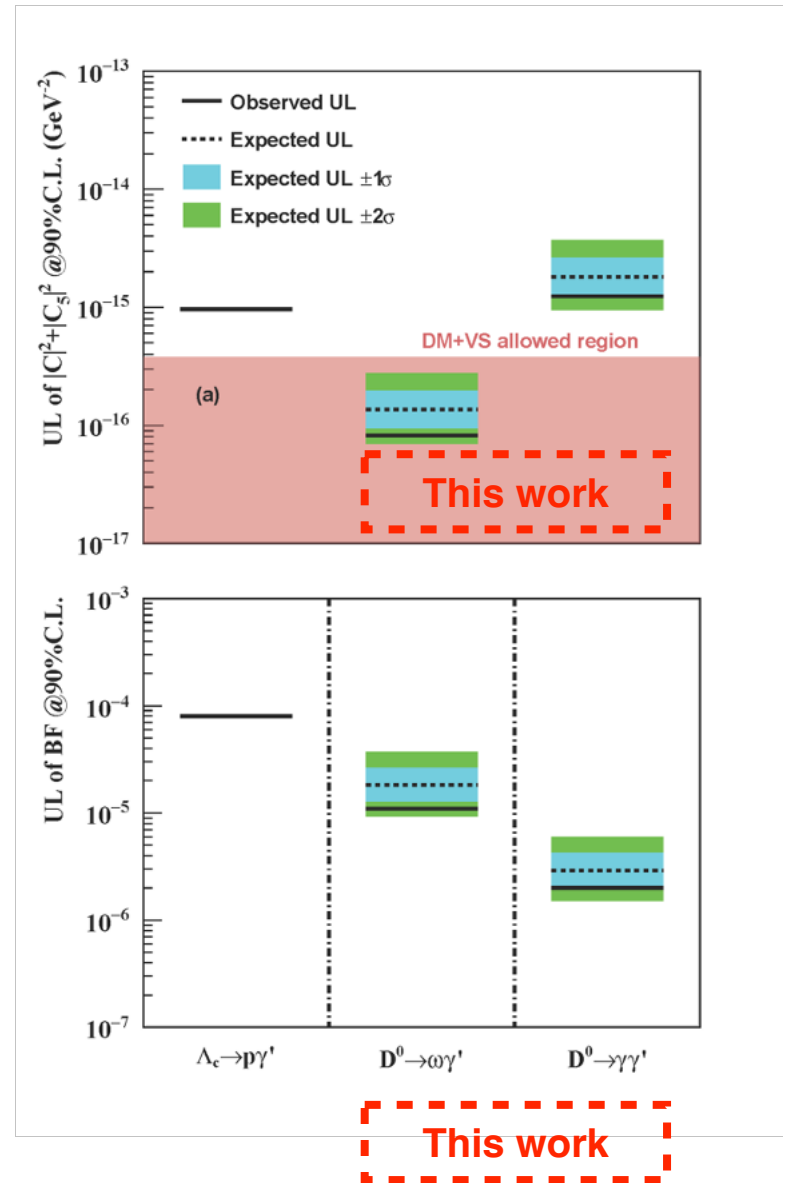
$$\mathcal{B}(D \rightarrow \gamma\gamma') = \frac{\alpha_e}{2} \tau_D f_{D\gamma}^2 m_D^3 (|\mathbb{C}|^2 + |\mathbb{C}_5|^2)$$

Previous BESIII work {

$$\mathcal{B}(\Lambda_c \rightarrow p\gamma') = \frac{\tau_{\Lambda_c} f_{\Lambda_c p}^2 (m_{\Lambda_c}^2 - m_p^2)^3}{2\pi m_{\Lambda_c}^3} (|\mathbb{C}|^2 + |\mathbb{C}_5|^2)$$

- $\mathbb{C} = \Lambda_{NP}^{-2} (\mathbf{C}_{12}^U + C_{21}^{U*}) v / \sqrt{8}$
- $\mathbb{C}_5 = \Lambda_{NP}^{-2} (C_{12}^U - C_{21}^{U*}) v / \sqrt{8}$

The constraint from $D^0 \rightarrow \omega\gamma'$ goes into the dark matter (DM) and vacuum stability (VS) allowed region for the first time, **improved by more than 1 order**



More recent results on LDM search...

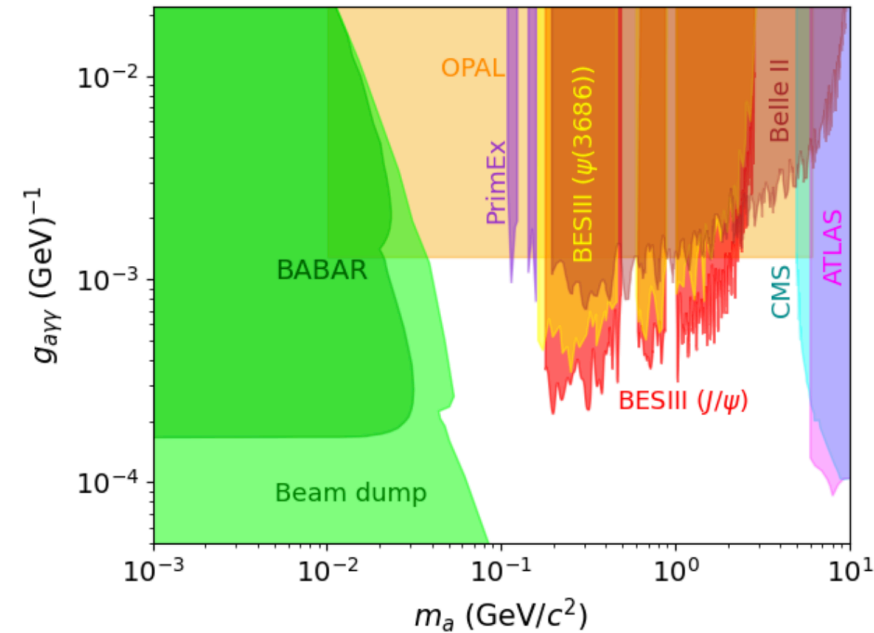
○ ALP in $J/\psi \rightarrow \gamma a, a \rightarrow \gamma\gamma$

✓ Using 2.7B $\psi(3686)$ data PLB 838 137698 (2023)

✓ Using 10B J/ψ data PRD.110.L 031101 (2024)

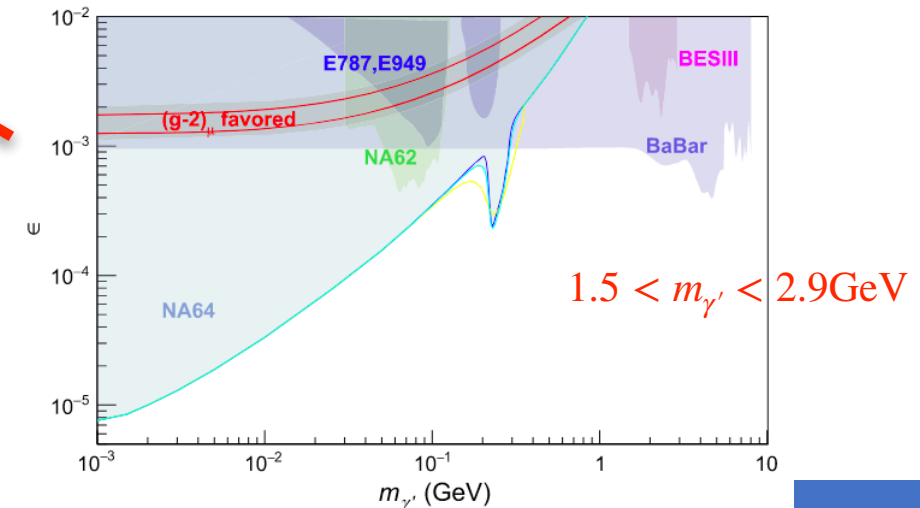
Most stringent constraints to date
for $0.18 < m_a < 2.85 \text{ GeV}/c^2$

ALPs-photon Coupling



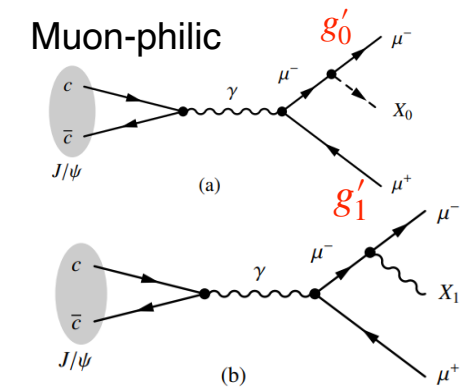
○ Search for massive dark photon with $e^+e^- \rightarrow \gamma\gamma'$

PLB 839 (2023) 137785

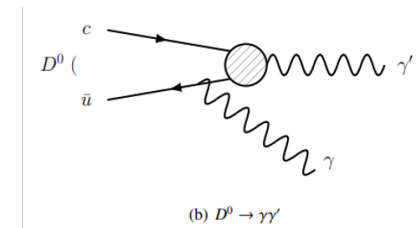
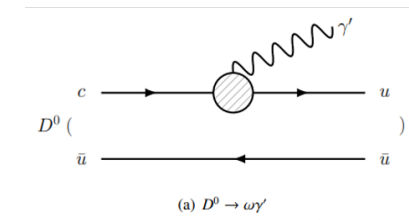


- ➡ New results of K_S^0 (Σ^+) invisible decay, muonphilic particle, dark photon, ALP at BESIII
- ➡ No evidence and more stringent constraint on NP
- ➡ BESIII has collected $10^{10} J/\psi$, $2.7 \times 10^9 \psi$, $20 \text{ fb}^{-1} \psi(3770)$ and more
- ➡ More & better results are coming soon

Thanks you!



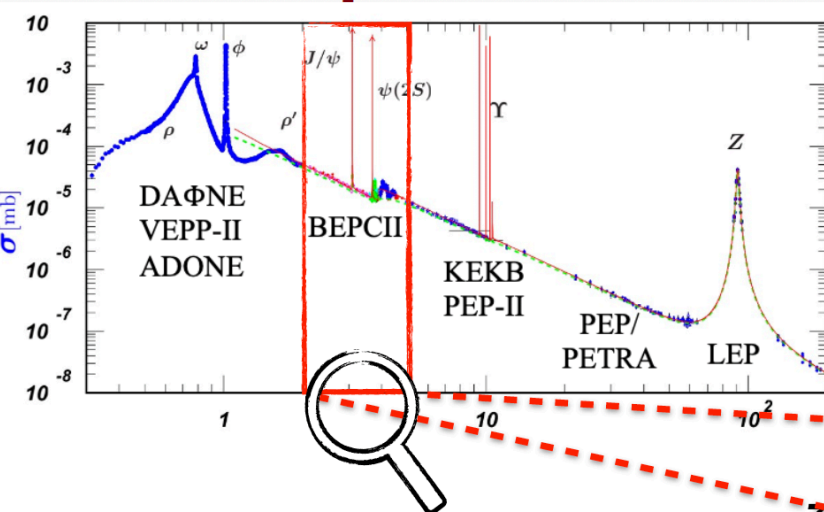
massless- γ'



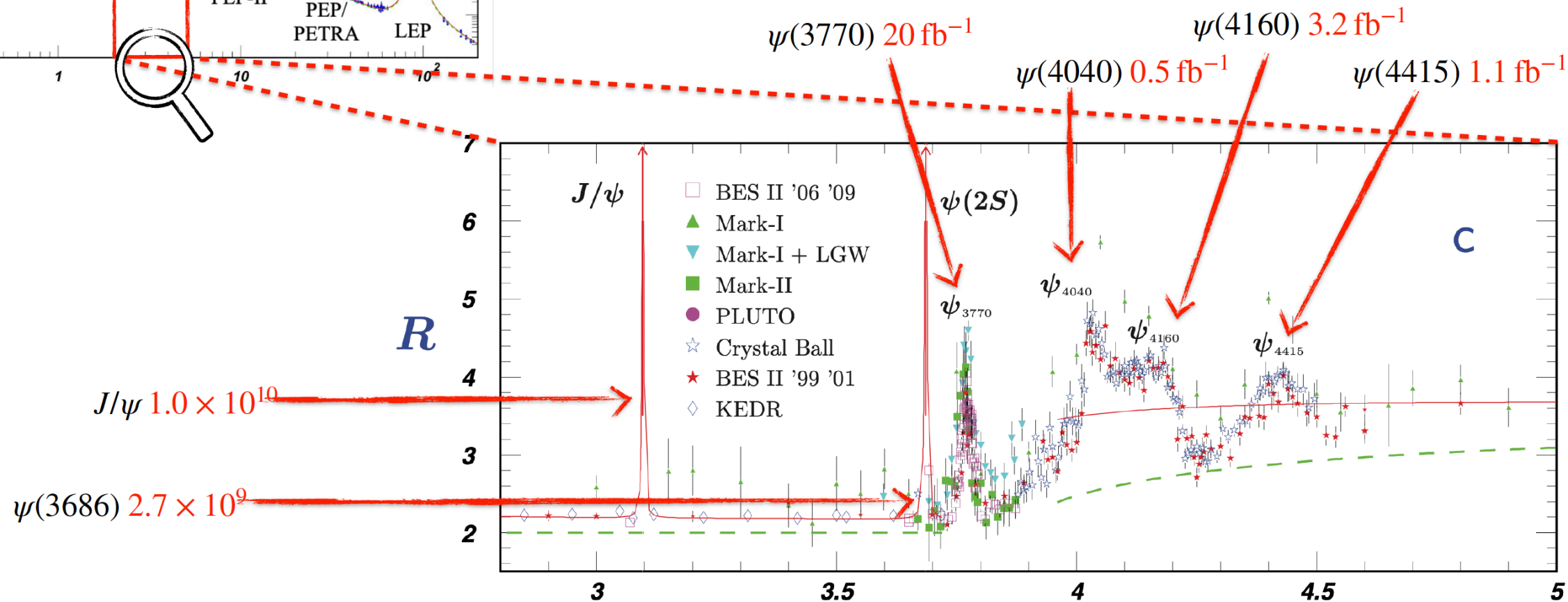
Thanks for your attention!

Back UP

Data samples at BESIII



- BESIII has collected the largest data samples of J/ψ and $\psi(3686)$ on the threshold in the world!
- $> 20 \text{ fb}^{-1}$ data above 4.0 GeV in total



Massless dark photon

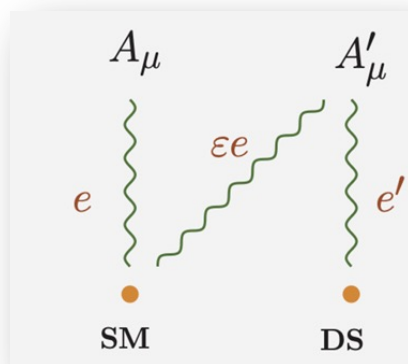
Massive dark photon

Symmetry broken spontaneously

$$A^\mu \rightarrow A^\mu + \epsilon A'^\mu$$

$$\mathcal{L} = eJ_\mu A^\mu + \underline{\epsilon e J_\mu A'^\mu} + e' J'_\mu A'^\mu$$

- Dark photon couples to the SM matter



Dark photon

Kinetic mixing: $\frac{\epsilon}{2} F'_{\mu\nu} F^{\mu\nu}$
 ϵ : mixing strength



SM photon

Massless dark photon

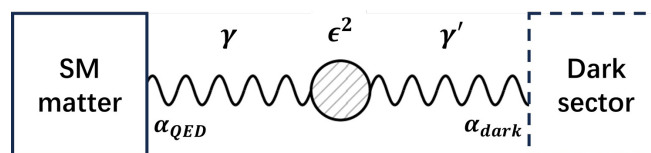
Symmetry remains unbroken

PRL 94, 151802 (2005)

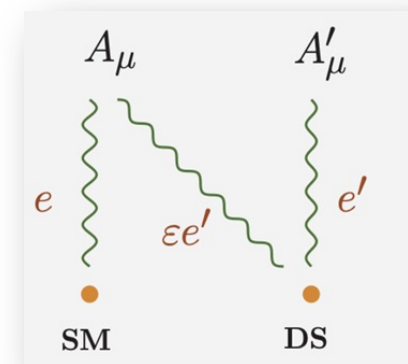
$$A'^\mu \rightarrow A'^\mu + \epsilon A^\mu$$

$$\mathcal{L} = eJ_\mu A^\mu + \underline{\epsilon' e' J'_\mu A'^\mu} + e' J'_\mu A'^\mu$$

- SM photon couples to the dark sector particles



A portal to connect the SM matter and the dark sector



Massless dark photon has no interaction with the SM matter in the dimension-4 operator

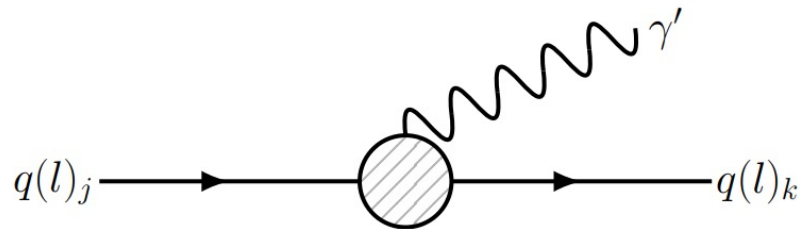
The interaction of γ'

$$\mathcal{L} = eJ_\mu A^\mu + e'\epsilon J'_\mu A^\mu + e'J'_\mu A'^\mu \text{ (no interaction between } \gamma' \text{ and SM matter)}$$

PRL 94, 151802 (2005)

$$\mathcal{L}_{NP} = \frac{1}{\Lambda_{NP}^2} (\underbrace{C_{jk}^U}_{\text{Up type quarks coupling}} \bar{q}_j \sigma^{\mu\nu} u_k \tilde{H} + \underbrace{C_{jk}^D}_{\text{Down type quarks coupling}} \bar{q}_j \sigma^{\mu\nu} d_k H + \underbrace{C_{jk}^L}_{\text{Charged leptons coupling}} \bar{l}_j \sigma^{\mu\nu} e_k H + h.c.) \bar{F}_{\mu\nu}$$

Dimension-six operator Up type quarks coupling Down type quarks coupling Charged leptons coupling Massless dark photon



Experimental search:

- $H \rightarrow \gamma\gamma'$
 - $\mu \rightarrow e\gamma'$
 - $\Lambda_c^+ \rightarrow p\gamma'$
- } • **No signal observed**
• **The sensitivity still lie outside the theoretically predicted allowed region**

Decay width of $J/\psi \rightarrow \mu^+ \mu^- X$

Probing new physics scenarios of muon $g - 2$ via J/ψ decay at BESII

J. High Energy Phys. 10 (2020) 207

$$|\mathcal{M}_{\mu\mu X_0}|^2 = \left(\frac{2}{3}e^2 g_0 \frac{f_J}{m_J}\right)^2 \frac{-8}{3 m_J^2(m_J - 2 E_-)^2(-2 E_- - 2 E_X + m_J)^2} \left(-4 m_\mu^2(4 E_-^2(m_X^2 - 2 E_X m_J) \right. \\ + E_-(-8 E_X^2 m_J + 4 E_X(m_X^2 + 2 m_J^2) - 4 m_X^2 m_J) - E_X^2(m_X^2 - 6 m_J^2) - 2 E_X m_J(m_X^2 + m_J^2) + m_X^2 m_J^2) \\ + 4 E_-^2(2 E_X^2 m_J^2 + m_X^2 m_J(m_J - 2 E_X) + m_X^4) \\ + 4 E_-(2 E_X^3 m_J^2 - 2 E_X^2 m_J(m_X^2 + m_J^2) + E_X(m_X^4 + 3 m_X^2 m_J^2) - m_X^2 m_J(m_X^2 + m_J^2)) \\ \left. - 16 E_X^2 m_\mu^4 + m_J(-4 E_X^3 m_J^2 + 2 E_X^2(3 m_X^2 m_J + m_J^3) - 2 E_X(m_X^4 + 2 m_X^2 m_J^2) + m_X^2 m_J(m_X^2 + m_J^2)) \right),$$

where E_- , the energy of μ^- and E_X , the energy of X_0 are measured in the rest frame of J/ψ .

$$\Gamma_{\mu\mu X_{0,1}} = \int_{E_X^{\min}}^{E_X^{\max}} \int_{E_-^{\min}}^{E_-^{\max}} \frac{|\mathcal{M}_{\mu\mu X_{0,1}}|^2}{64\pi^3 m_J} dE_- dE_X,$$

$$|\mathcal{M}_{\mu\mu X_1}|^2 = \left(\frac{2}{3}e^2 g_1 \frac{f_J}{m_J}\right)^2 \frac{-16}{3 m_J^2(m_J - 2 E_-)^2(-2 E_- - 2 E_X + m_J)^2} \left(16 E_-^4 m_J^2 + 32 E_-^3 m_J^2(E_X - m_J) \right. \\ + 2 m_\mu^2(4 E_-^2(m_J(m_J - 2 E_X) + m_X^2) - 4 E_-(2 E_X^2 m_J - E_X(m_X^2 + 3 m_J^2) + m_J(m_X^2 + m_J^2)) \\ + 2 E_X^2(m_X^2 + 3 m_J^2) - 2 E_X m_J(m_X^2 + 2 m_J^2) + m_J^2(m_X^2 + m_J^2)) \\ + 4 E_-^2(m_J^2(6 E_X^2 - 14 E_X m_J + 7 m_J^2) + m_X^2 m_J(3 m_J - 2 E_X) + m_X^4) \\ + 4 E_-(2 E_X^3 m_J^2 - 2 E_X^2 m_J(m_X^2 + 4 m_J^2) + E_X(m_X^4 + 5 m_X^2 m_J^2 + 9 m_J^4) - m_J(m_X^4 + 3 m_X^2 m_J^2 + 3 m_J^4)) \\ \left. + 8 E_X^2 m_\mu^4 + m_J(-4 E_X^3 m_J^2 + 2 E_X^2(3 m_X^2 m_J + 5 m_J^3) - 2 E_X(m_X^2 + 2 m_J^2)^2 + m_J(m_X^4 + 3 m_X^2 m_J^2 + 2 m_J^4)) \right),$$

where E_- , the energy of μ^- and E_X , the energy of X_0 are measured in the rest frame of J/ψ .

$K_S^0 \rightarrow \text{invisible}$: Strategy

Signal sample

► $J/\psi \rightarrow \phi K_S^0(\text{tag}) K_S^0$, $K_S^0(\text{tag}) \rightarrow \pi^+ \pi^-$, $K_S^0 \rightarrow \text{invisible}$

$$N_{\text{signal}} = 2 \times N_{J/\psi \rightarrow \phi K_S^0 K_S^0} \times \mathcal{B}(\phi \rightarrow K^+ K^-) \times \mathcal{B}(K_S^0 \rightarrow \pi^+ \pi^-) \\ \times \mathcal{B}(K_S^0 \rightarrow \text{invisible}) \times \varepsilon_{\text{signal}}, \quad (3.2)$$

Normalization sample

► $J/\psi \rightarrow \phi K_S^0(\text{tag}) K_S^0$, $K_S^0(\text{tag}) \rightarrow \pi^+ \pi^-$, $K_S^0 \rightarrow \text{non} - \pi^+ \pi^-$

$$N_{\text{norm.}} = 2 \times N_{J/\psi \rightarrow \phi K_S^0 K_S^0} \times \mathcal{B}(\phi \rightarrow K^+ K^-) \times \mathcal{B}(K_S^0 \rightarrow \pi^+ \pi^-) \\ \times (1 - \mathcal{B}(K_S^0 \rightarrow \pi^+ \pi^-)) \times \varepsilon_{\text{norm.}}, \quad (3.1)$$

$$\mathcal{B}(K_S^0 \rightarrow \text{invisible}) = \frac{N_{\text{signal}}}{N_{\text{norm.}} (\varepsilon_{\text{signal}}/\varepsilon_{\text{norm.}})} (1 - \mathcal{B}(K_S^0 \rightarrow \pi^+ \pi^-)).$$

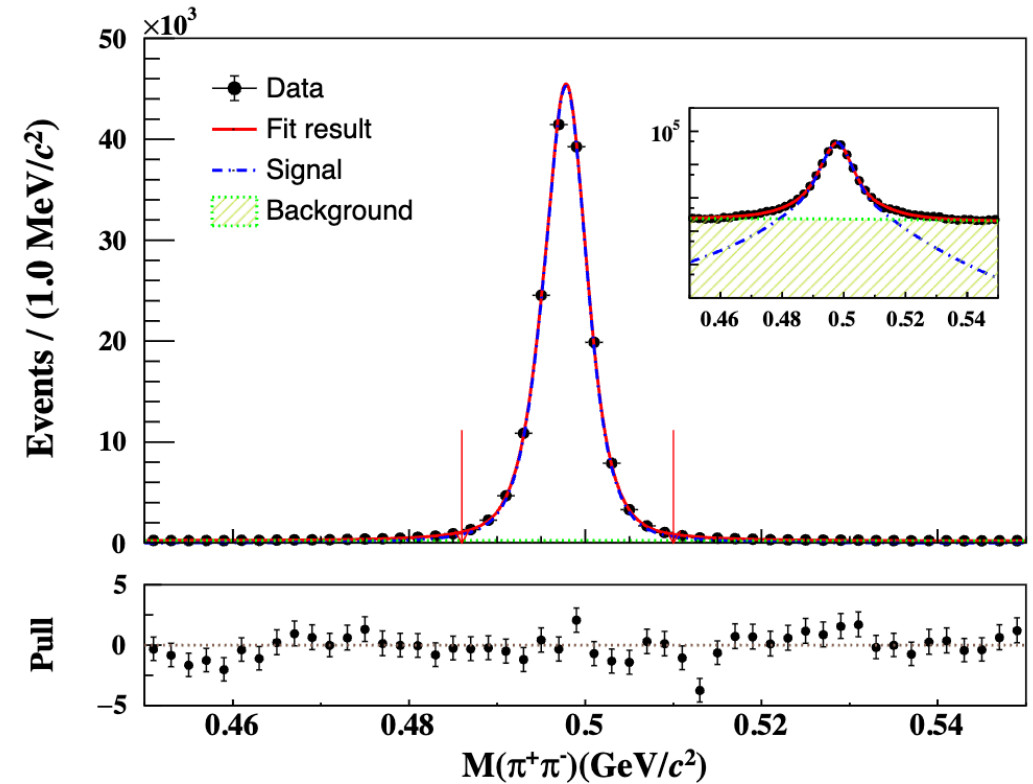
$K_S^0 \rightarrow \text{invisible}$: Strategy

Normalization sample

$$\triangleright J/\psi \rightarrow \phi K_S^0(\text{tag}) K_S^0, \quad K_S^0(\text{tag}) \rightarrow \pi^+\pi^-, \quad K_S^0 \rightarrow \text{non} - \pi^+\pi^-$$

$$N_{\text{norm.}} = 2 \times N_{J/\psi \rightarrow \phi K_S^0 K_S^0} \times \mathcal{B}(\phi \rightarrow K^+ K^-) \times \mathcal{B}(K_S^0 \rightarrow \pi^+ \pi^-) \\ \times (1 - \mathcal{B}(K_S^0 \rightarrow \pi^+ \pi^-)) \times \varepsilon_{\text{norm.}},$$

Fit to $M(\pi^+\pi^-)$ to extract the yields for normalization sample

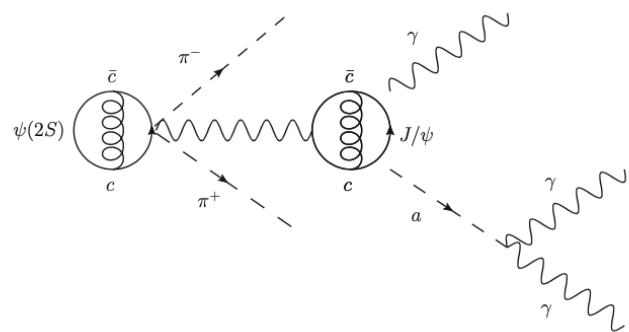


Search for an axion-like particle in radiative J/ψ decays

PLB 838 137698 (2023)

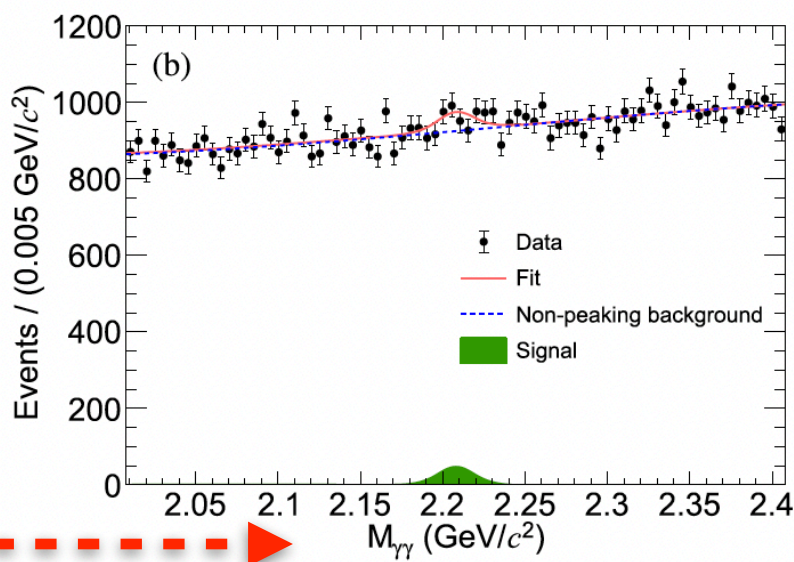
- Data samples: 2.7B $\psi(3686)$ events
- Strategy: Search for $J/\psi \rightarrow \gamma a, a \rightarrow \gamma\gamma$ with J/ψ sample obtained from $\psi(3686) \rightarrow \pi^+\pi^- J/\psi$

- $\psi(3686)$ decay: avoid the pollution from non-resonant production $e^+e^- \rightarrow \gamma\gamma\gamma$



$$\frac{\mathcal{B}(J/\psi \rightarrow \gamma a)}{\mathcal{B}(J/\psi \rightarrow e^+ e^-)} = \frac{m_{J/\psi}^2}{32\pi\alpha} g_{a\gamma\gamma}^2 \left(1 - \frac{m_a^2}{m_{J/\psi}^2}\right)^3$$

- Decay width of $a \rightarrow \gamma\gamma$: $\Gamma_a = \frac{g_{a\gamma\gamma}^2 m_a^3}{64\pi}$
- Taking $g_{a\gamma\gamma} \sim 10^{-4} \text{ GeV}^{-1}$, $m_a \sim \text{GeV}$, the lifetime of ALP is short in the detector
 - ➔ Assume $\text{BF}(a \rightarrow \gamma\gamma) \sim 100\%$
- Three $\gamma\gamma$ combinations per event, **extract signal from $M_{\gamma\gamma}$ distribution**
- 674 mass hypotheses



The maximum local significance is 2.6σ at $m_a = 2208 \text{ MeV}/c^2$

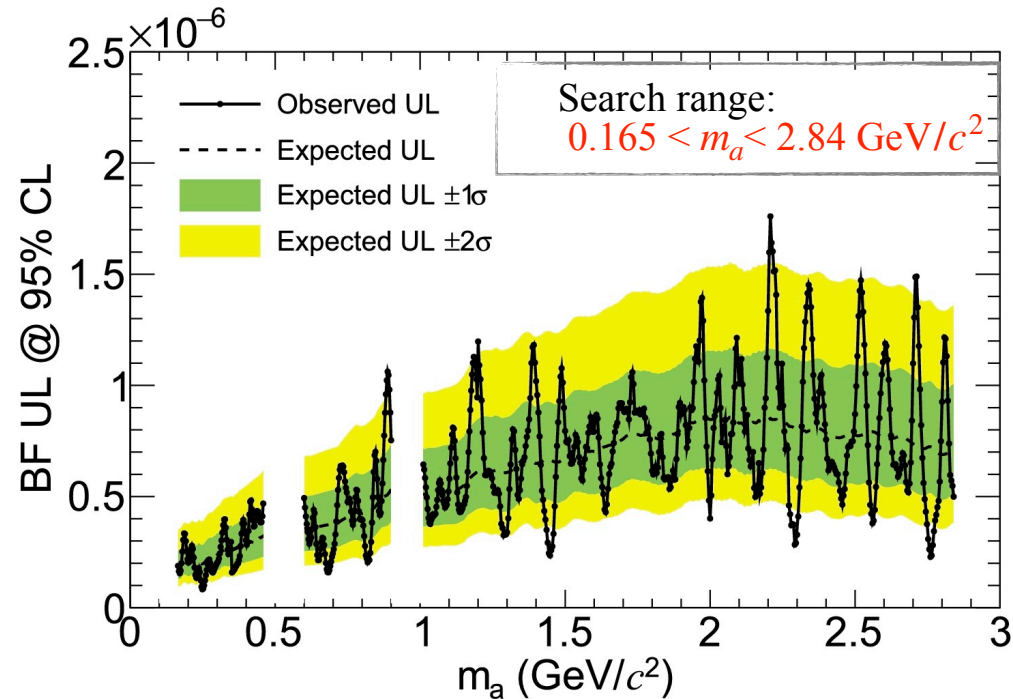
Search for an axion-like particle in radiative J/ψ decays

- No significant ALP signal is observed

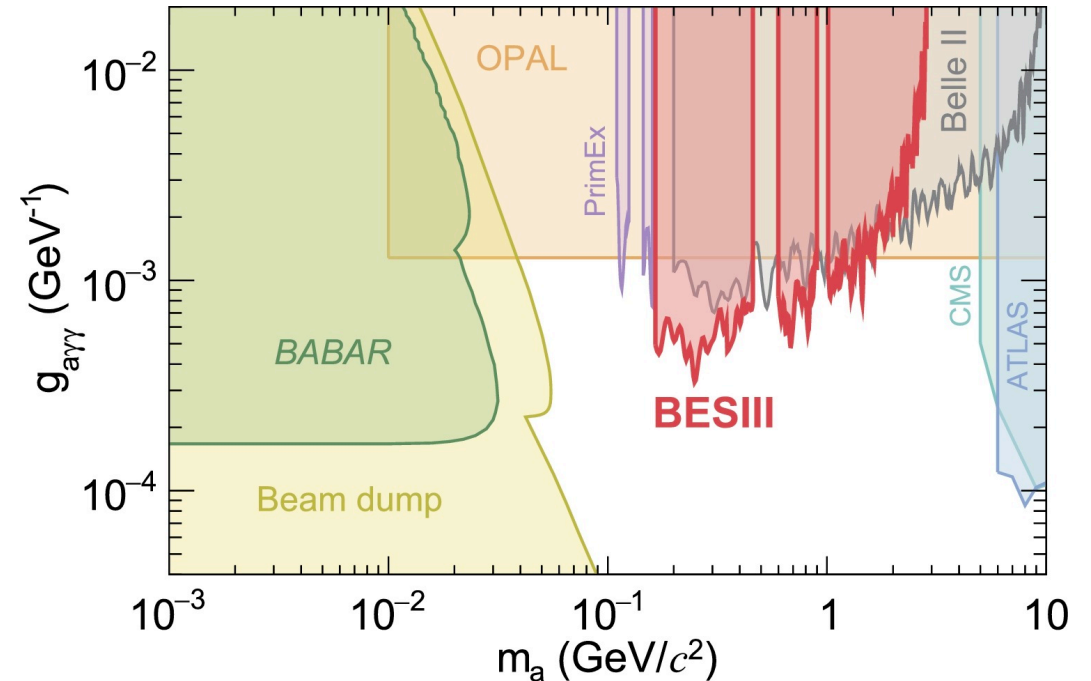
$$\frac{\mathcal{B}(J/\psi \rightarrow \gamma a)}{\mathcal{B}(J/\psi \rightarrow e^+ e^-)} = \frac{m_{J/\psi}^2}{32\pi\alpha} g_{a\gamma\gamma}^2 \left(1 - \frac{m_a^2}{m_{J/\psi}^2}\right)^3$$



BF @ 95% CL



ALPs-photon Coupling



Search for an axion-like particle in radiative J/ψ decays

arXiv:2404.04640

○ Data samples: 10B J/ψ events

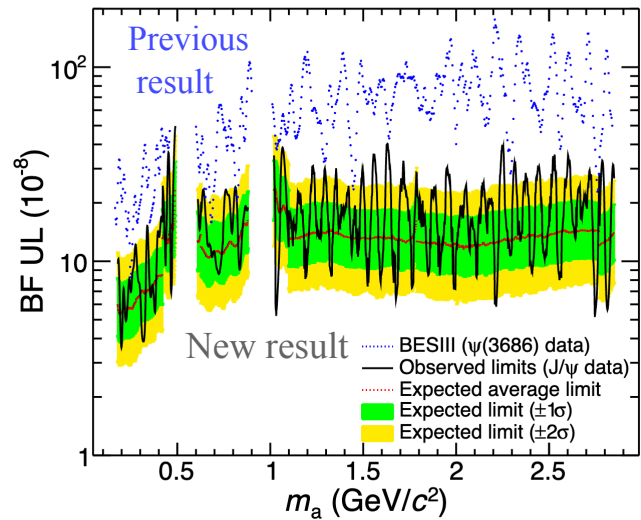
NEWEST RESULT!

○ Strategy: Search for $J/\psi \rightarrow \gamma a, a \rightarrow \gamma\gamma$ with J/ψ data on threshold

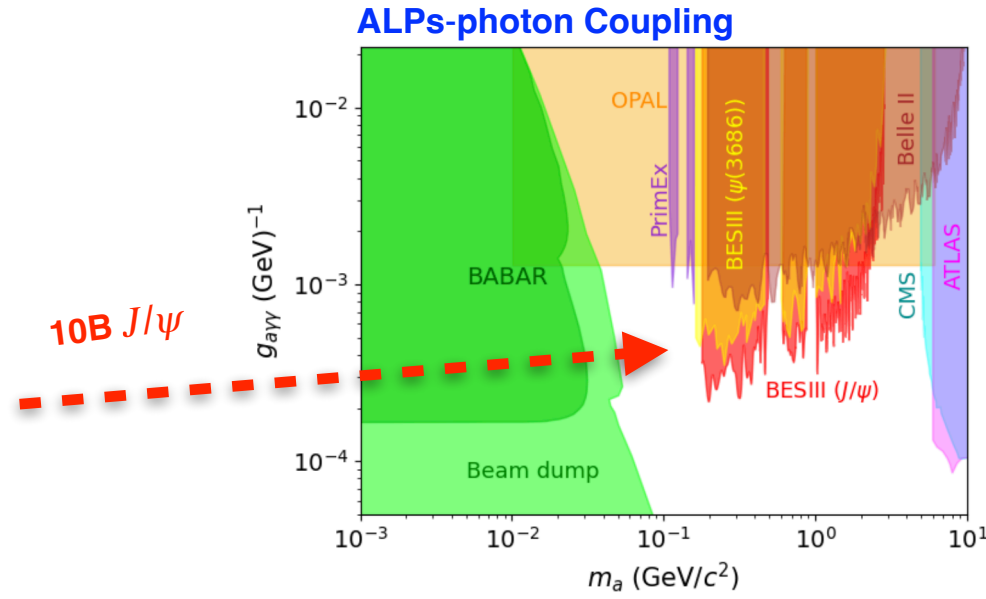
- Non-resonant process $e^+e^- \rightarrow \gamma a$ is indistinguishable from $J/\psi \rightarrow \gamma a$
- The ratio, $\sigma_{non-res}/\sigma_{res}$ is calculated to be 4.4%
- The contribution from $e^+e^- \rightarrow \gamma a$ is subtracted from the signal yields (interference between $\sigma_{non-res}, \sigma_{res}$ is negligible).

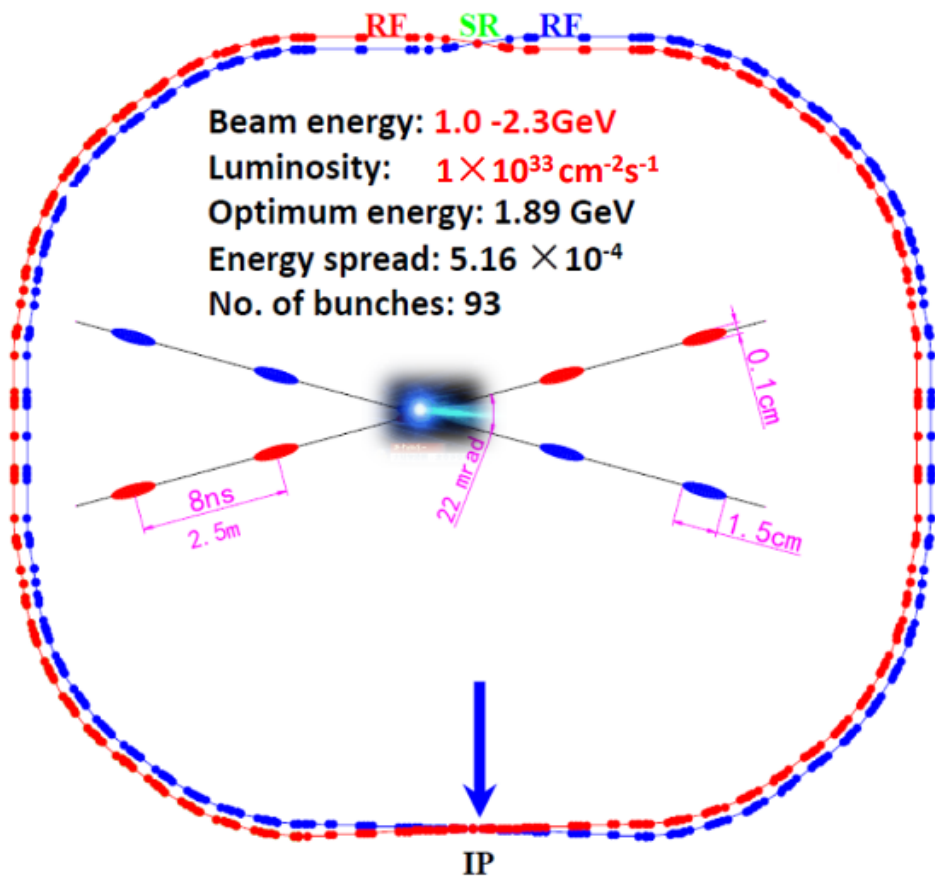
✓ The 95% CL. upper limits of $\mathcal{B}(J/\psi \rightarrow \gamma a)$ is set based on 10B data

Accepted by Phys. Rev. D (Letter)



Most stringent constraints to date for $0.18 < m_a < 2.85$ GeV





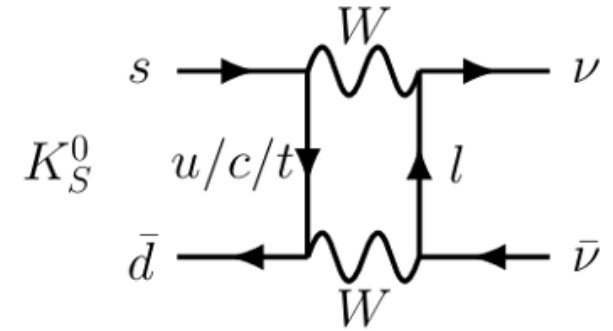
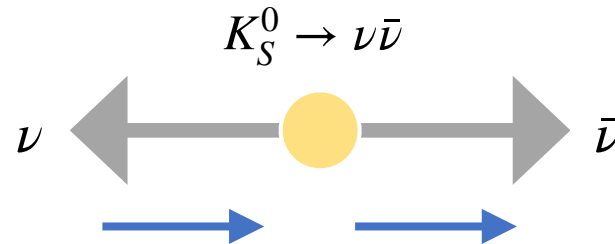
Invisible decay of K_S^0

Phys.Rev.D 91 (2015) 1, 015004

SM decay

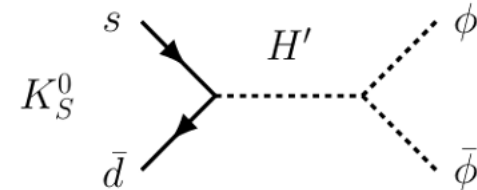
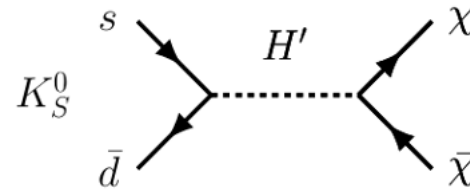
$$K_S^0 \rightarrow \nu \bar{\nu}$$

FCNC && helicity suppression, **BF** < 10^{-16}



Decay to DM

$$K_S^0 \rightarrow \chi \bar{\chi}$$



2HDM model
BF ~ $\mathcal{O}(10^{-6})$

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Invisible