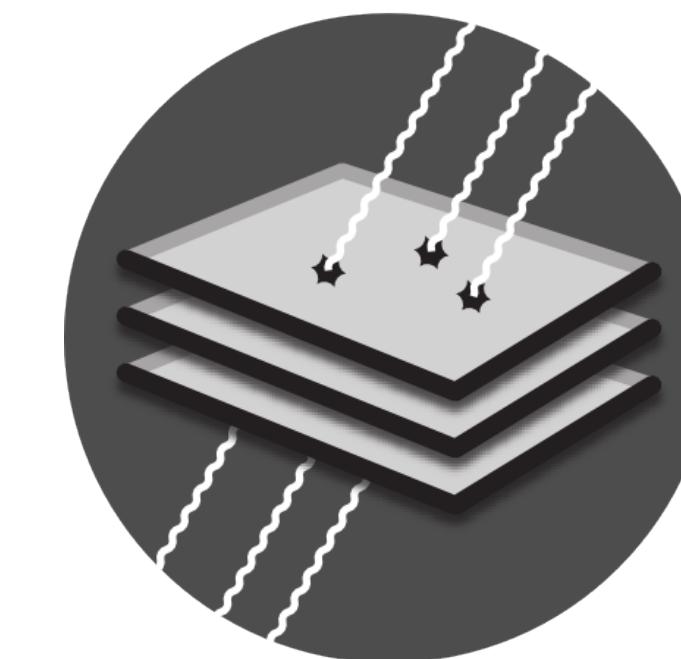
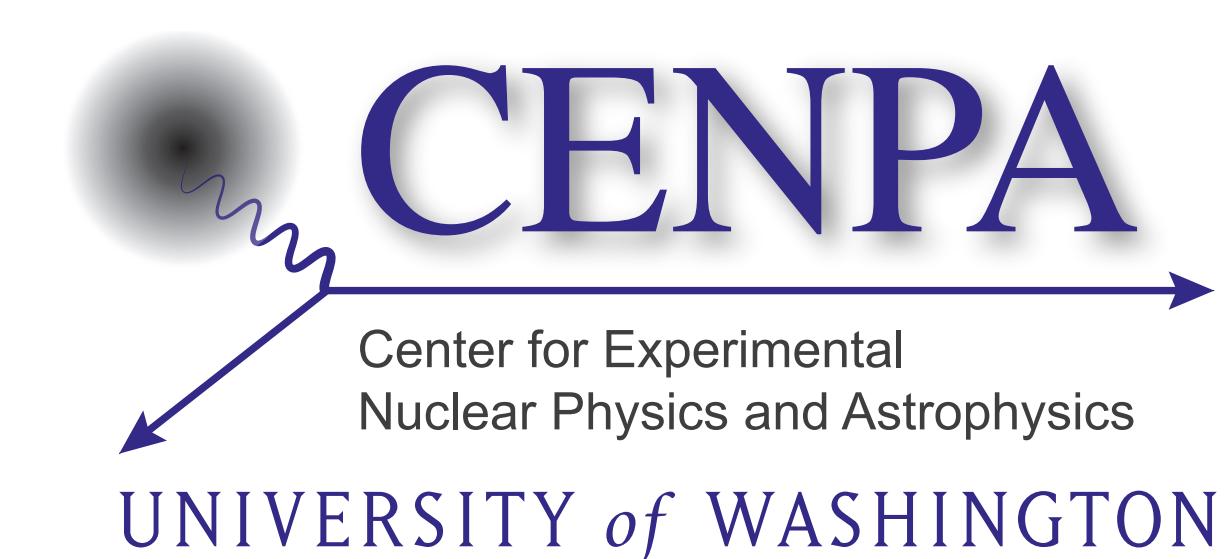
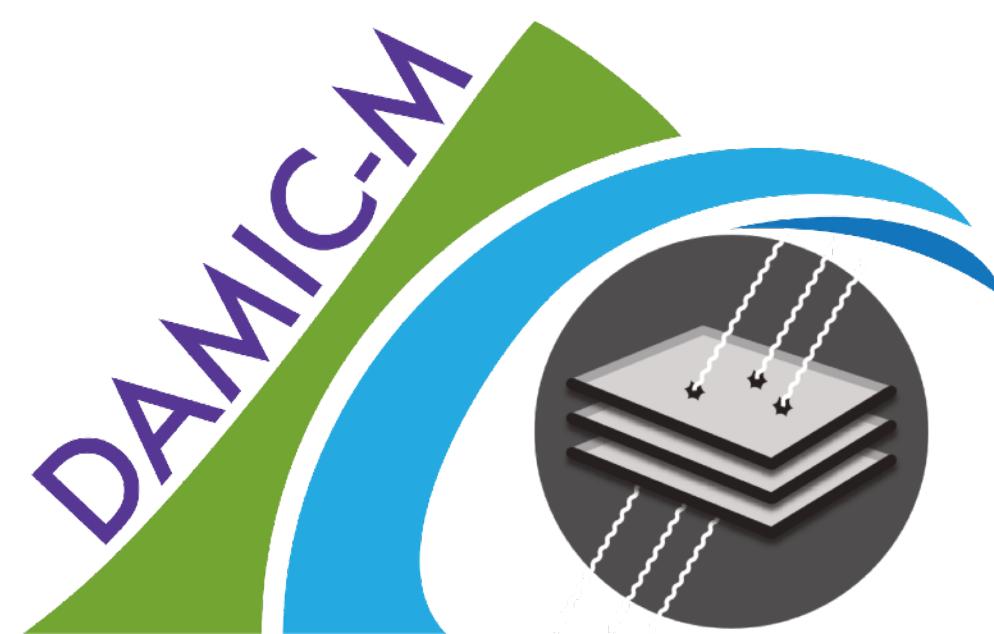


The search for dark matter with CCDs

Alvaro E. Chavarria
University of Washington



Outline

- Dark matter (DM) direct detection.
- DM-e⁻ scattering searches.
- Charge-coupled devices (CCDs).
- DAMIC-M at Modane Underground Lab.
- Results from the LBC test stand.
- DAMIC-M status and prospects.

Dark matter particles

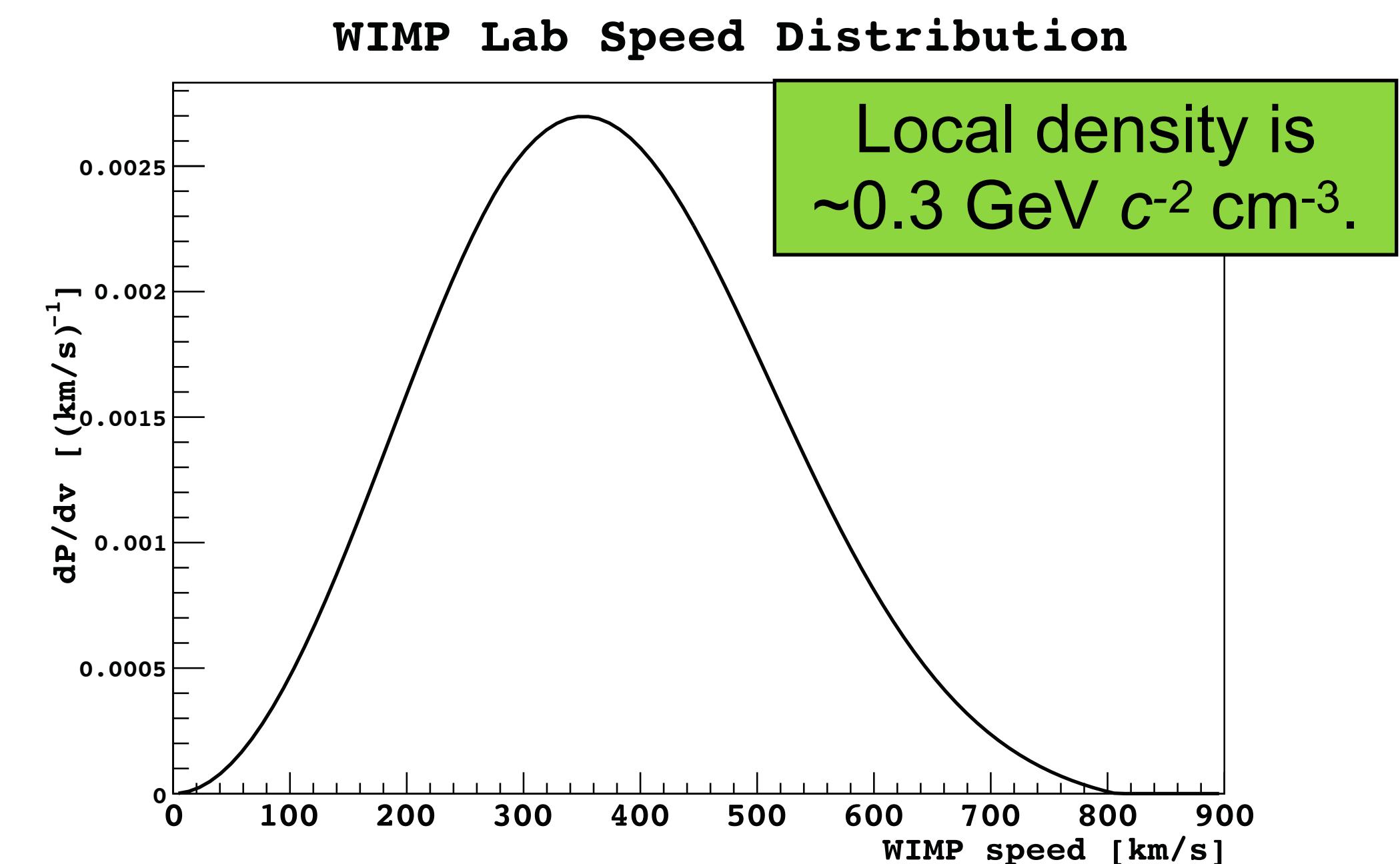
Dark matter is *cold*, i.e., it is bound to the galaxy.

Hence, the dark matter particle speed is ~the same as stars: 100s km/s.

$$E_\chi = \frac{1}{2} M_\chi v^2$$

$$E_\chi = \frac{1}{2} M_\chi c^2 \beta^2 \quad \beta \approx 10^{-3}$$

$$E_\chi \approx \left(\frac{M_\chi c^2}{\text{GeV}} \right) \text{keV}$$



We do not know the particle mass (M_χ)

A 1 GeV (proton-mass) particle has 1 keV of kinetic energy (very little).

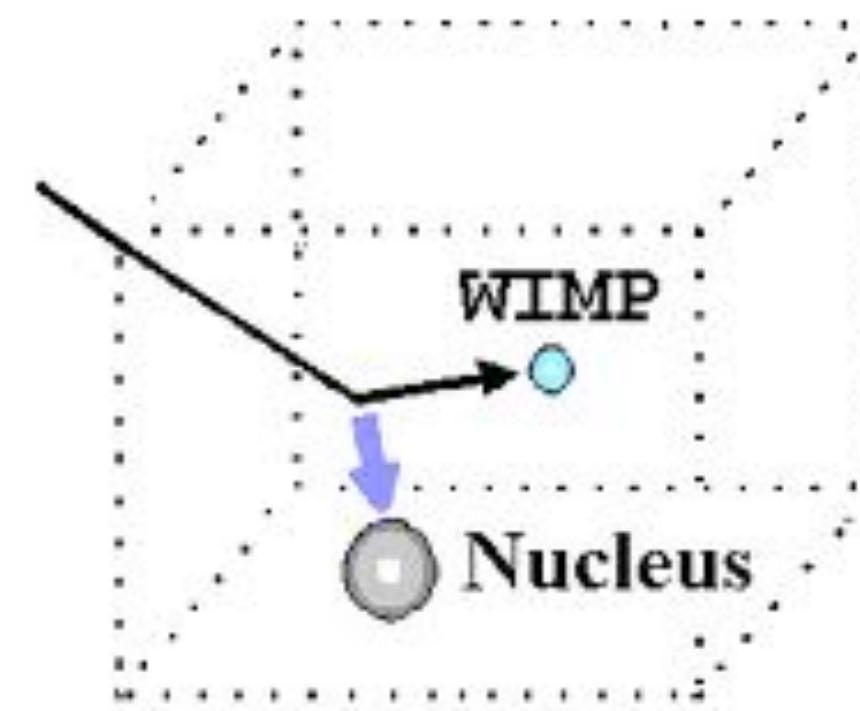
Dark matter signal

- Small interaction cross-section.
- Need detector with low energy threshold, largest possible exposure and correspondingly low backgrounds.

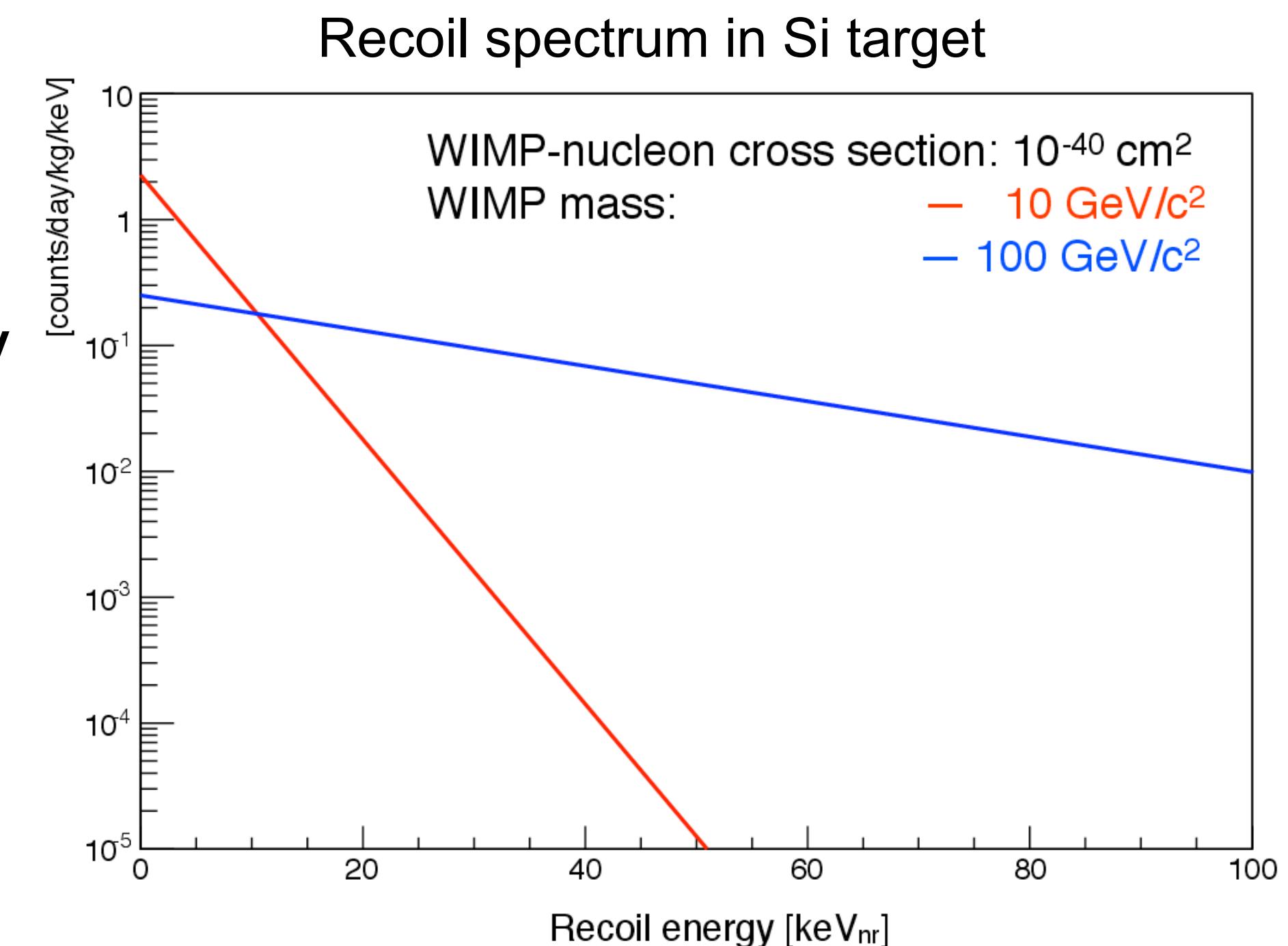
Traditional mechanism
for WIMP searches:

Coherent
enhancement:

$$\sigma_N \propto A^2$$



Maximum energy transfer when $M \sim A$

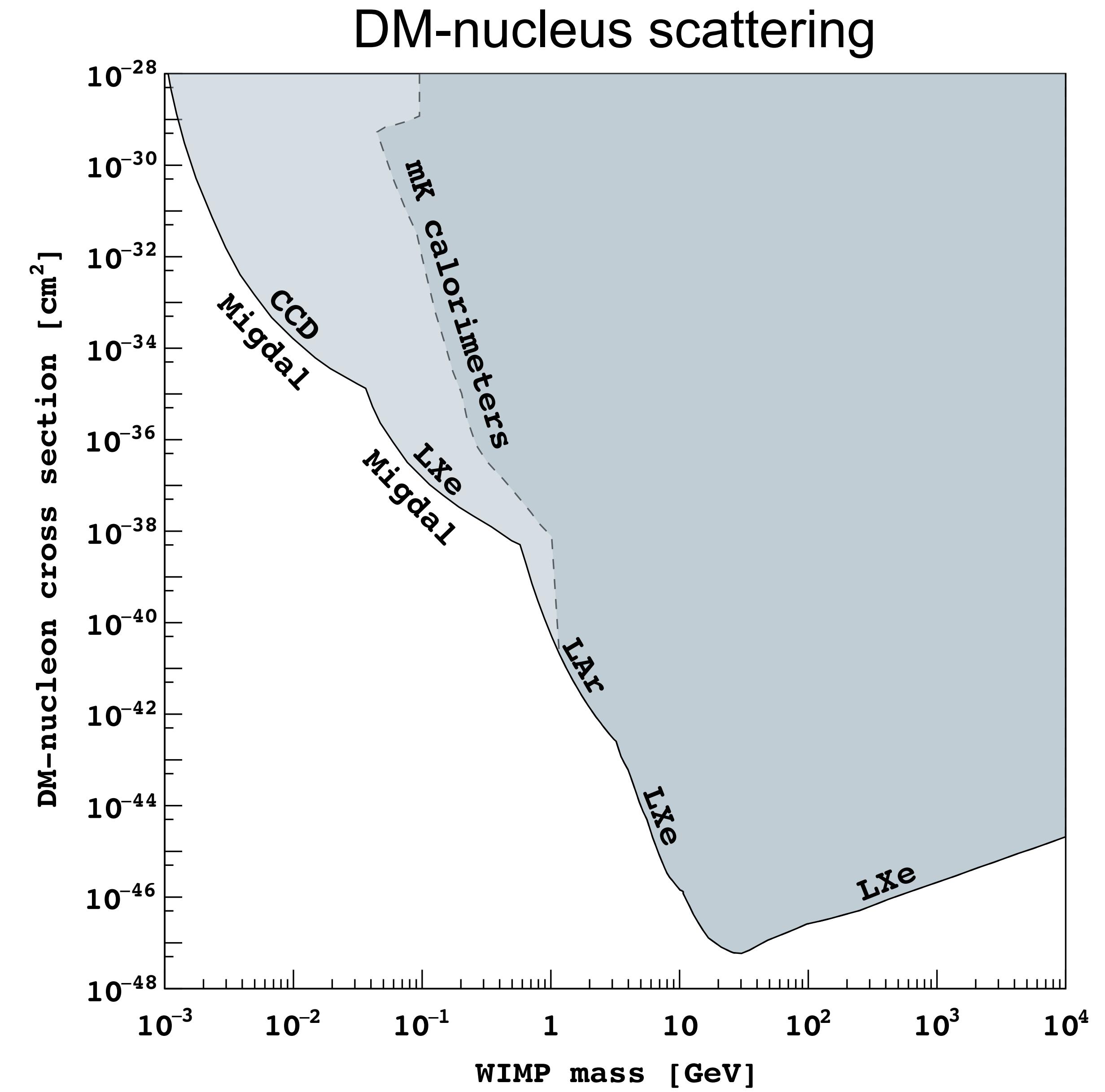


For low-mass WIMP: $M_T \gg M_\chi$

$$E_T < 4 \frac{M_\chi}{M_T} E_\chi$$

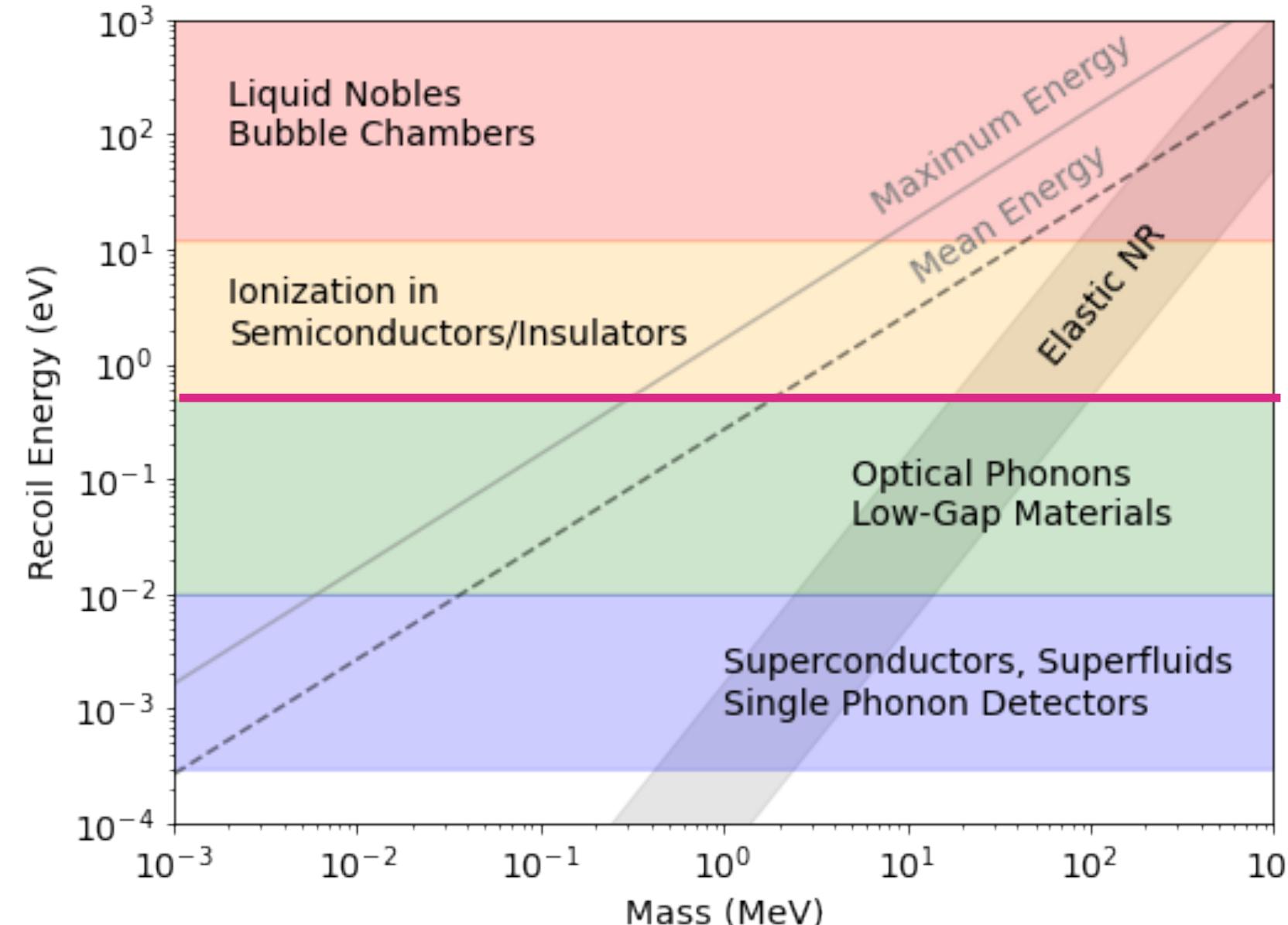
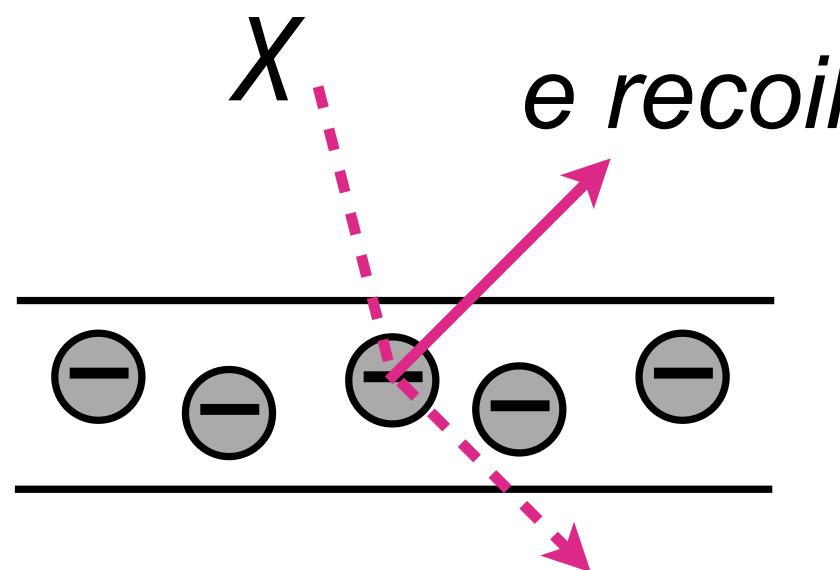
World Status

- World-wide effort to directly detect DM signals.
- For “particle” DM the search currently spans from ~ 1 MeV to the Planck mass.
- Different technologies target different mass ranges.
- CCDs have greatest sensitivity below 35 MeV*
- *Depends on the Migdal effect (10^{-6} probability of ionization, see McCabe’s talk)



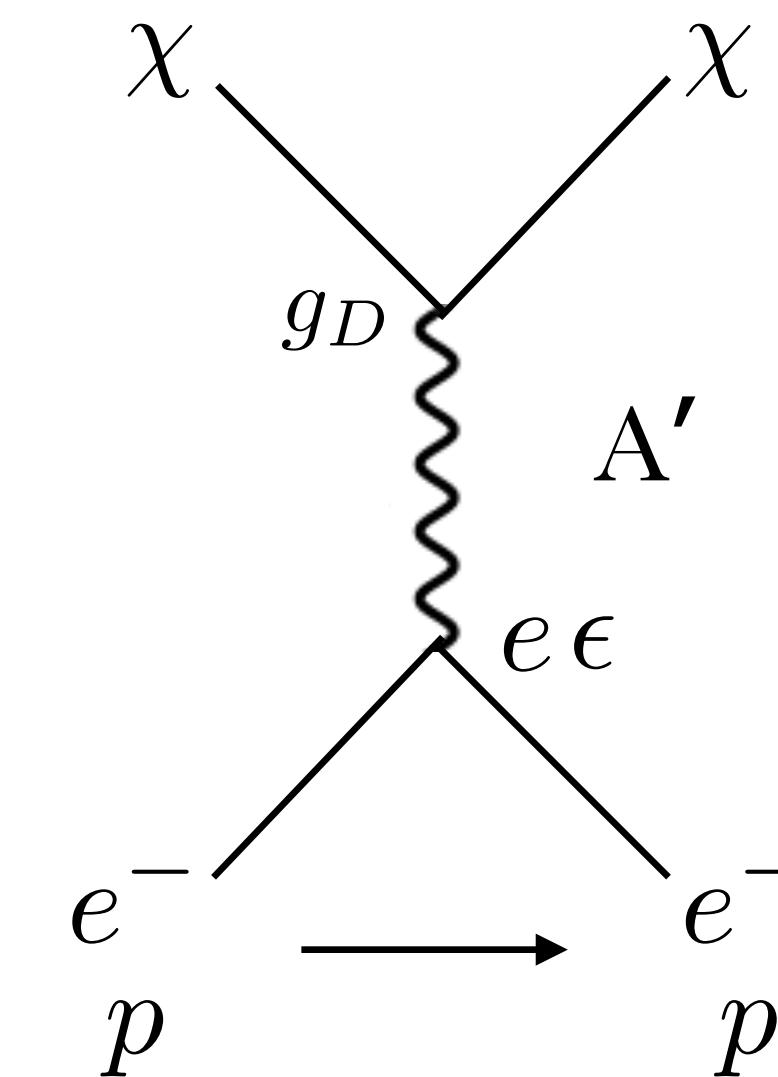
DM-e scattering

- ▶ Electrons are a lighter target and *ER visible as ionization*.
- ▶ Electrons bound with some momentum; there is a region of phase-space where the electron carries most of the WIMP kinetic energy.



- ▶ Momentum distributions in some targets better “kinematically matched” to the DM than others.
- ▶ Phase-space ‘penalty,’ no coherent enhancement and probing DM-e interaction cross-section.

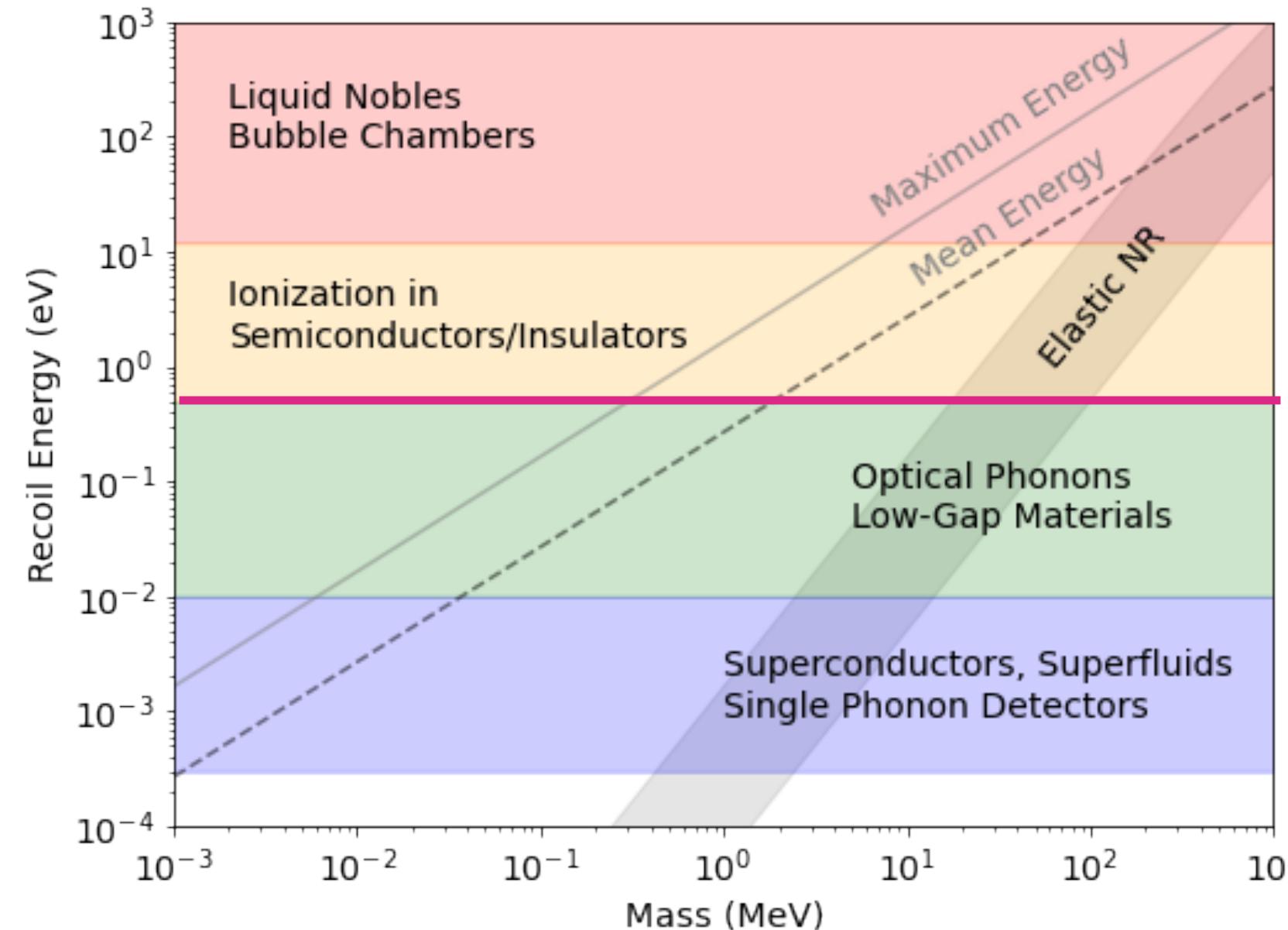
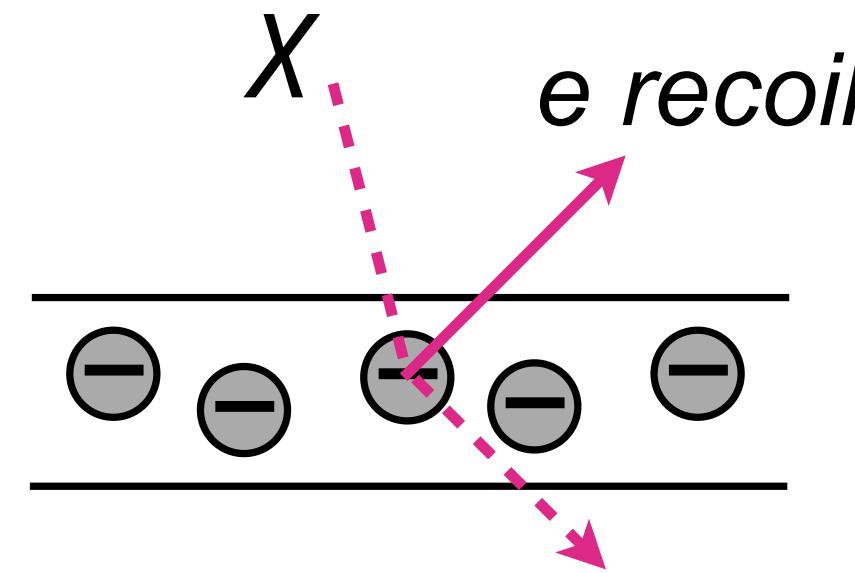
- ▶ Mediator A' mixes with SM photon.



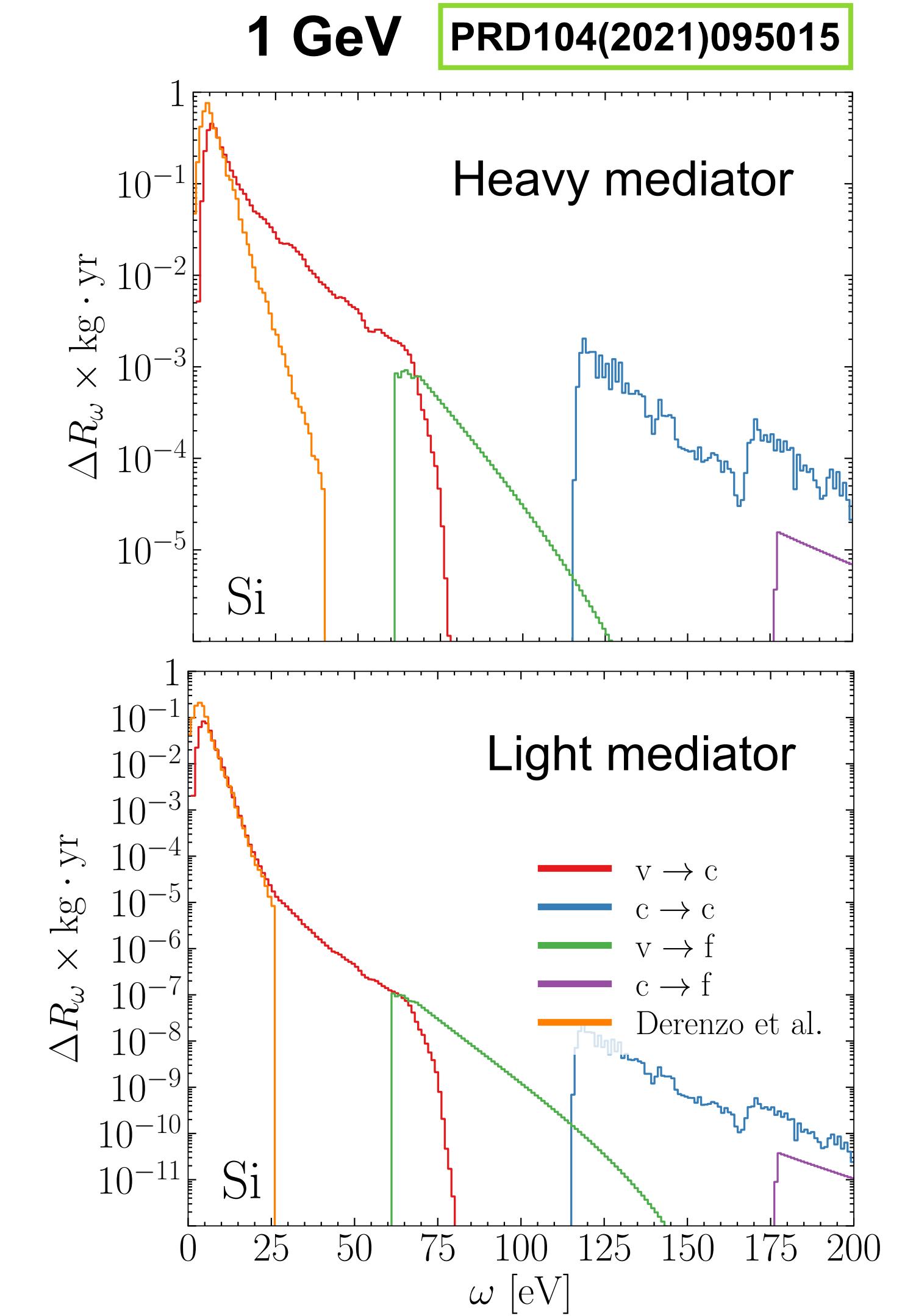
- ▶ Most sensitive direct-detection probe for sub-GeV hidden-sector DM!

DM-e scattering

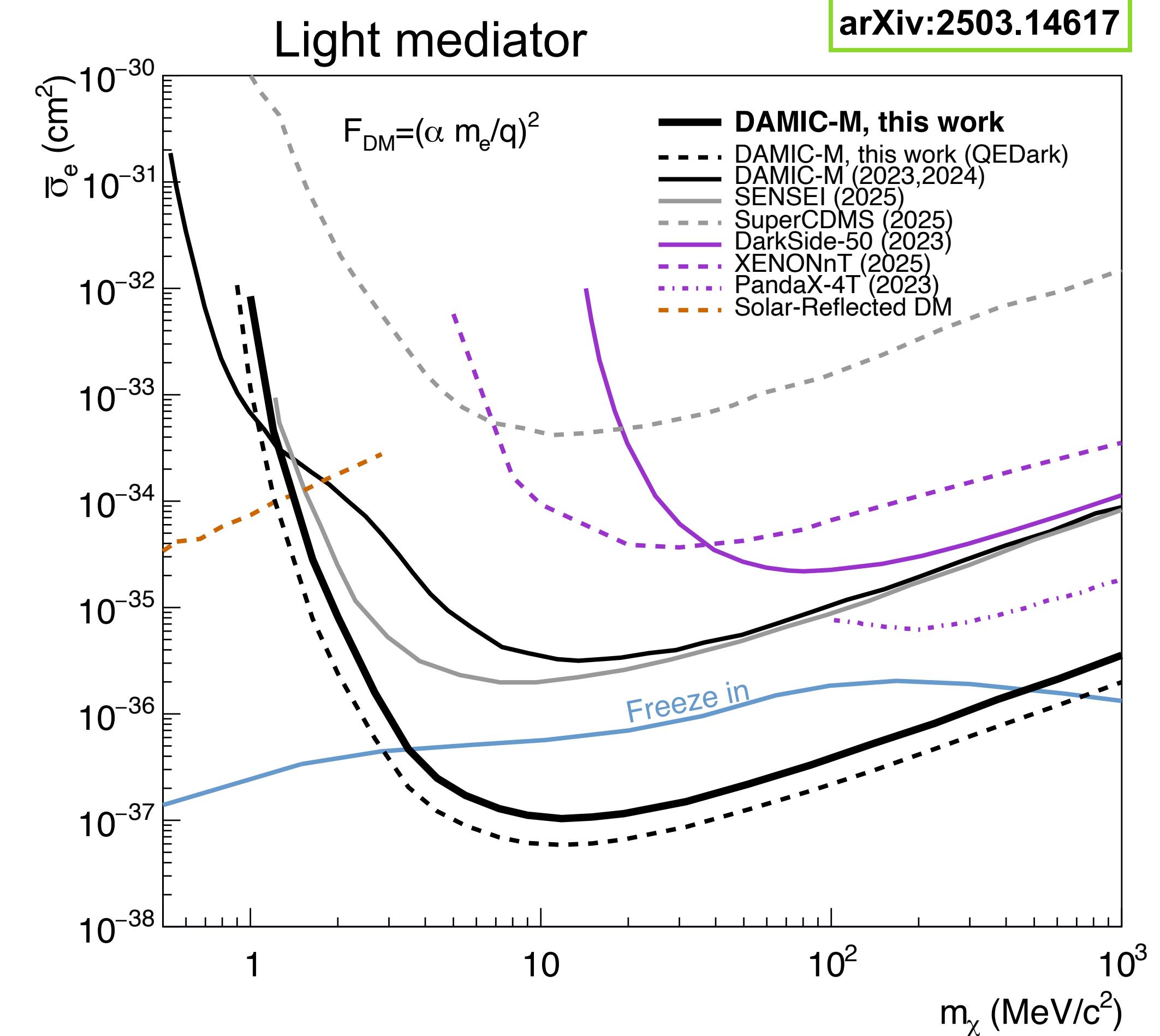
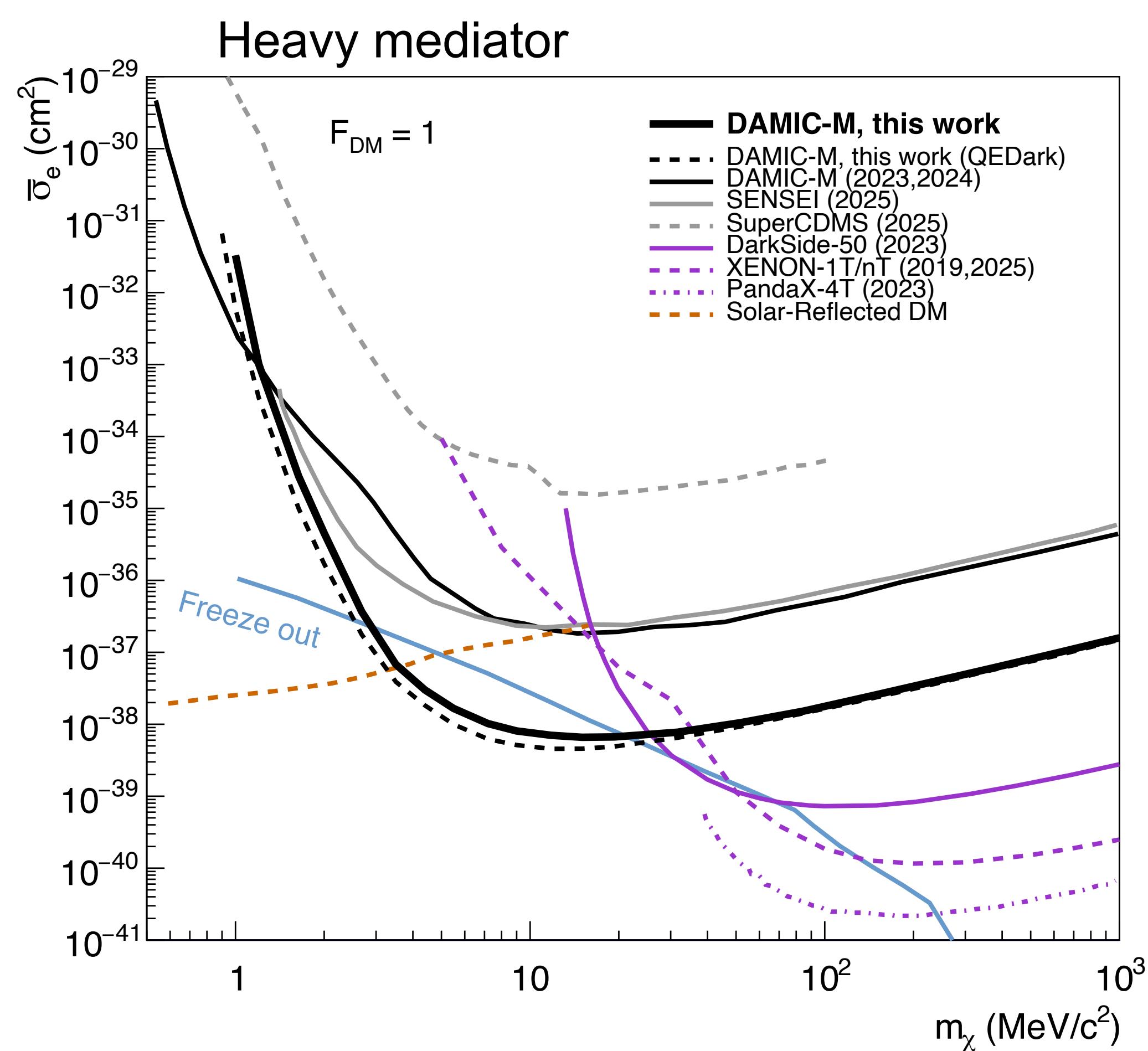
- ▶ Electrons are a lighter target and *ER visible as ionization*.
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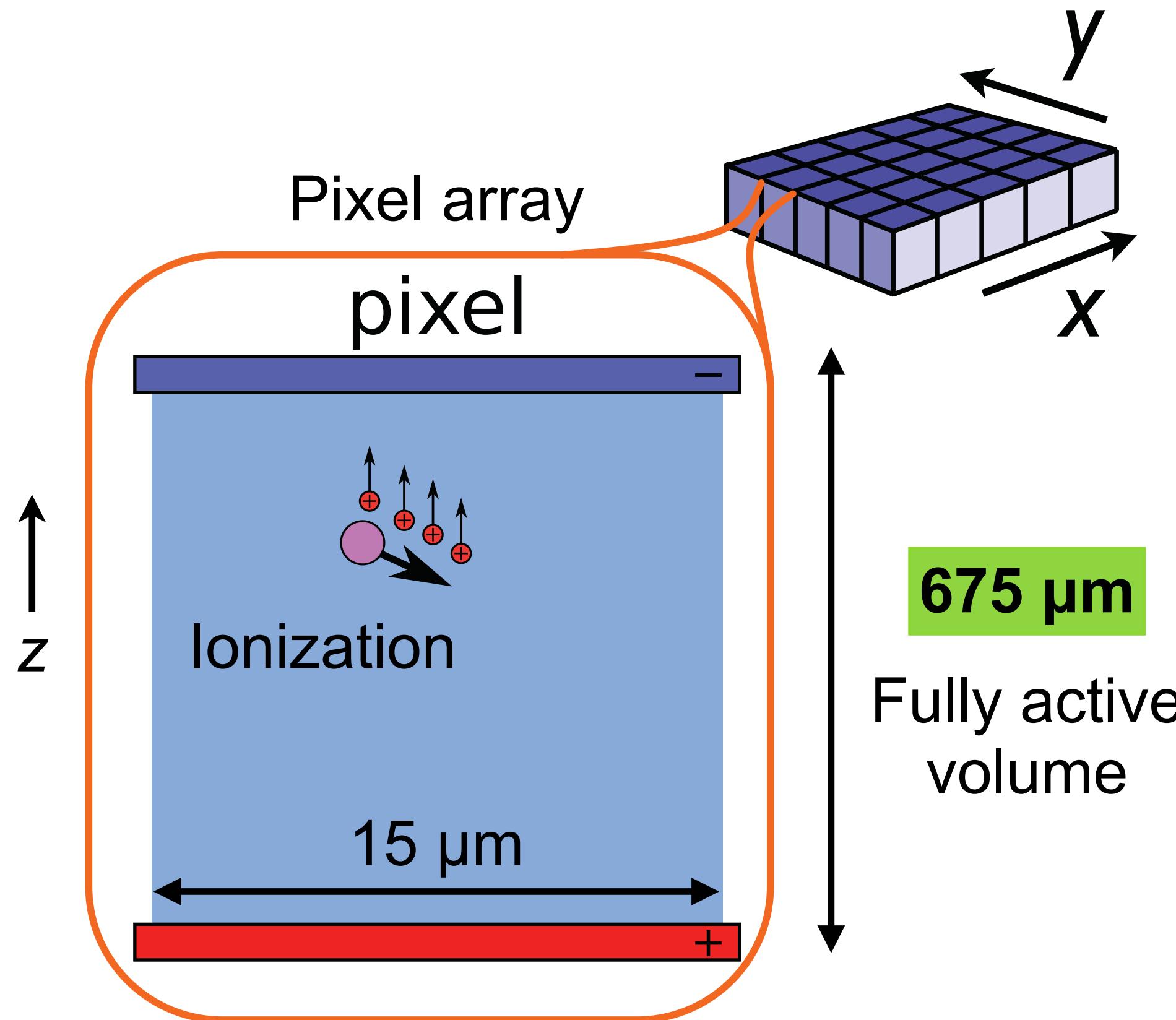
DM-e exclusion limits



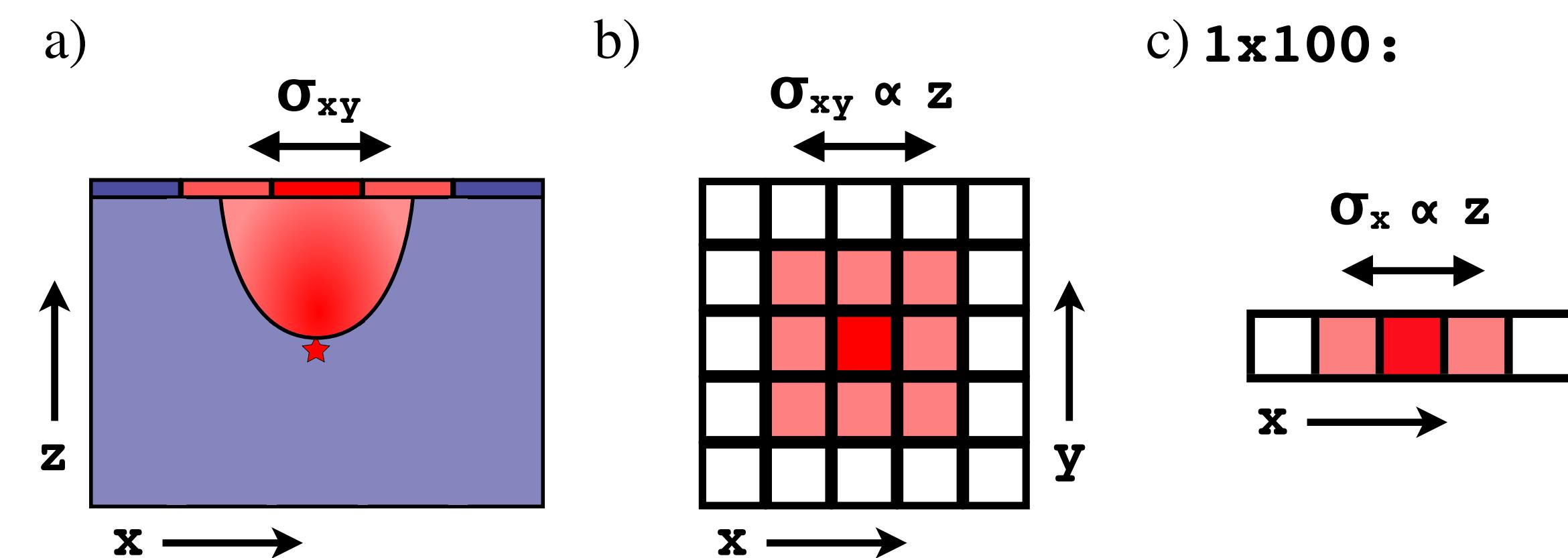
arXiv:2503.14617

► DAMIC-M has world-leading exclusion limits for sub-GeV hidden-sector DM!

Charge-coupled devices



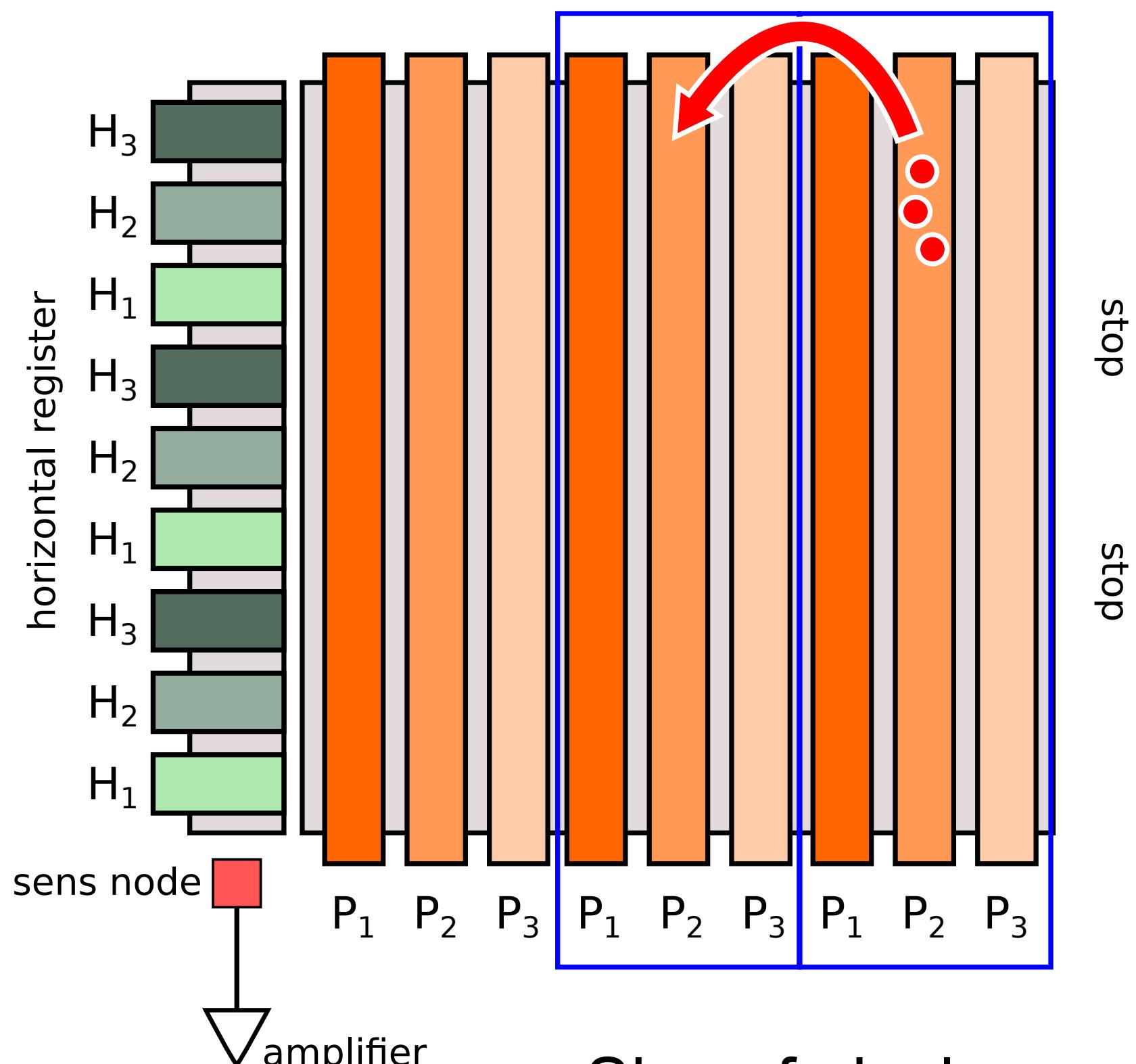
Silicon band-gap: 1.2 eV.
Mean energy for 1 e-h pair: 3.8 eV.



- ▶ Depth (z) reconstructed from distribution of charge on pixel array.
- ▶ Device is “exposed,” collecting charge until user commands readout.
- ▶ Readout can be slow: **very low noise**.
- ▶ Standard fabrication in semiconductor industry and easy cryogenics (~ 100 K).

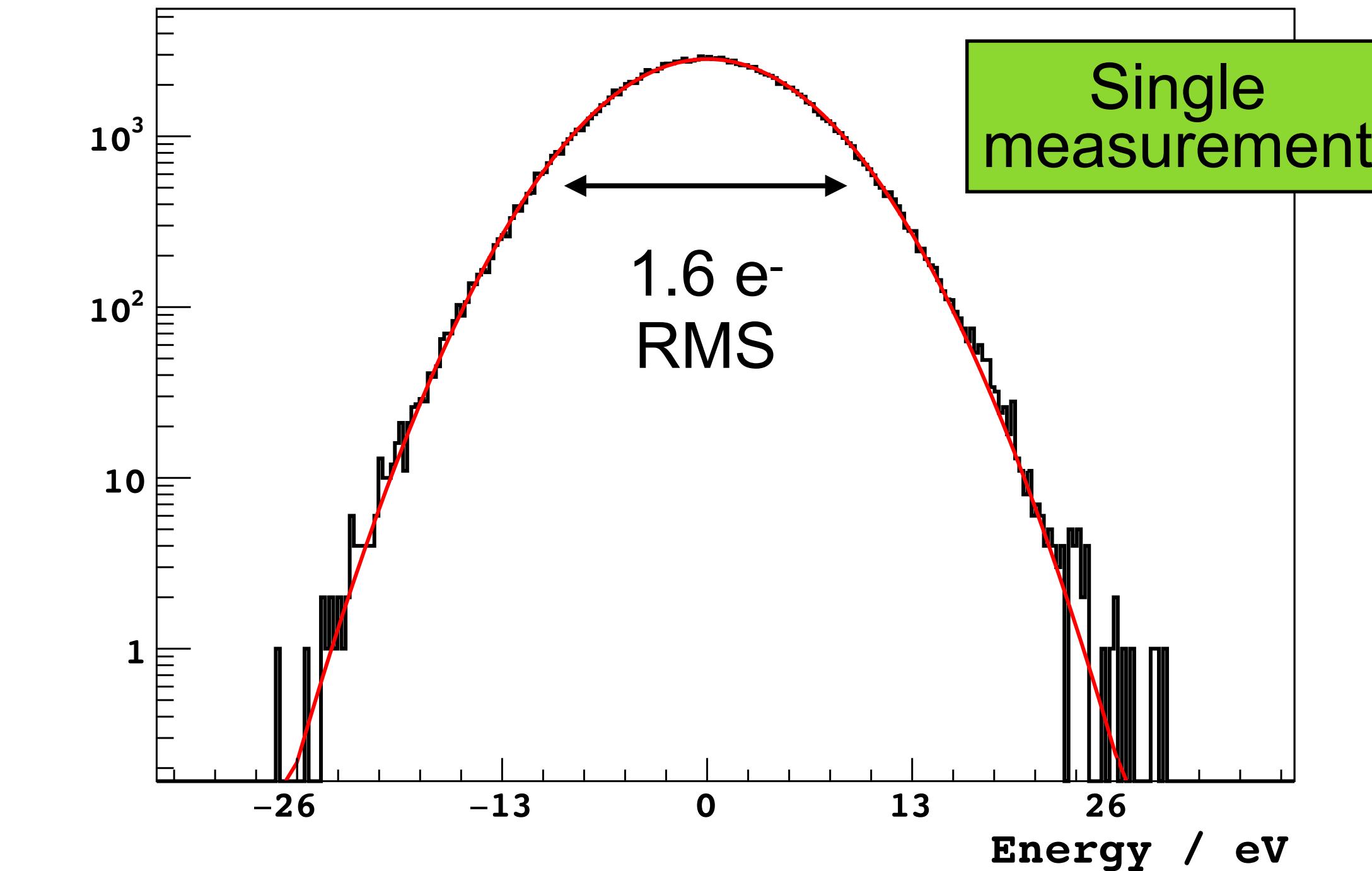
CCD readout

3x3 pixels CCD



$$C \sim 10 \text{ fF}$$

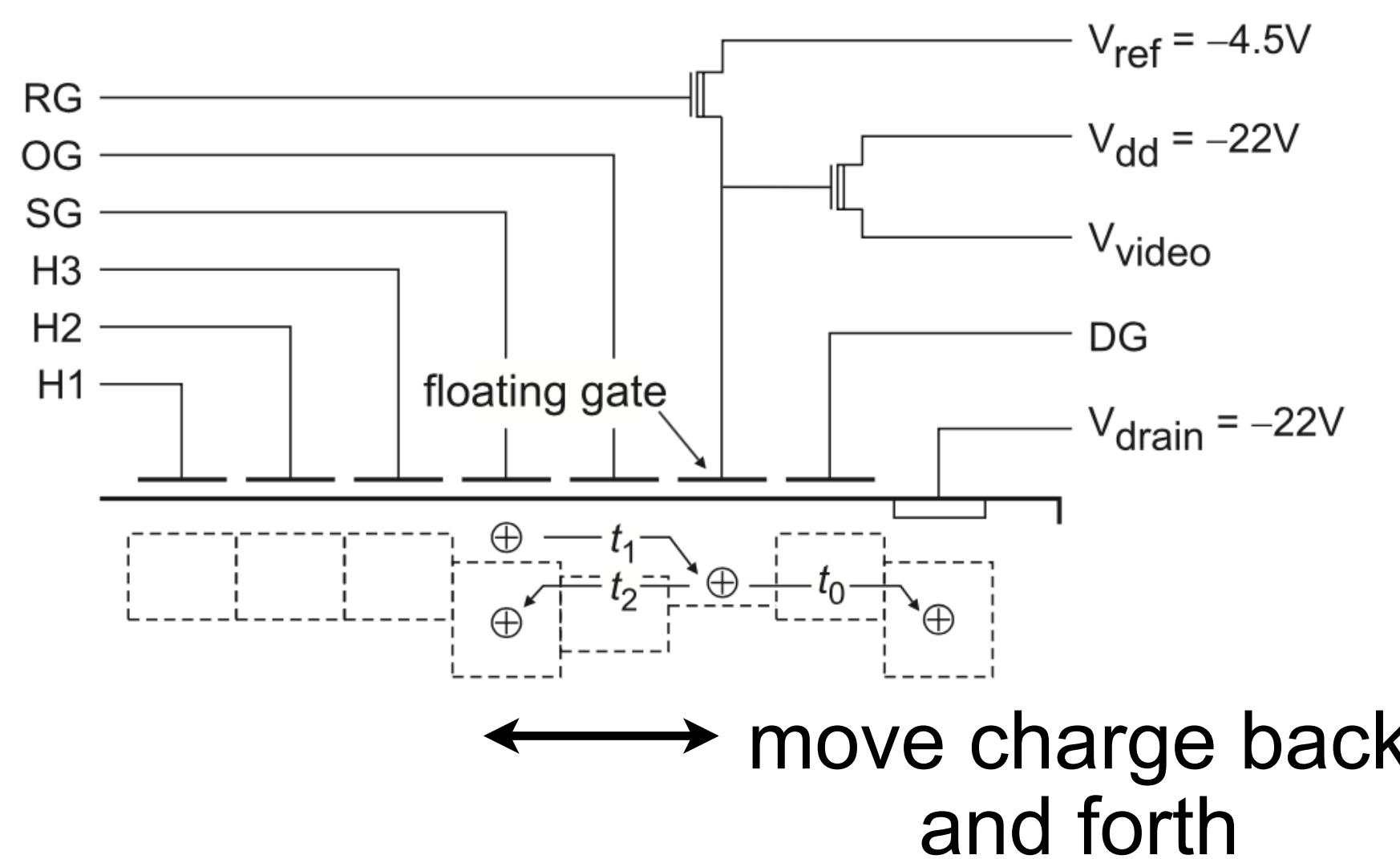
- Size of pixel array decoupled from capacitance of output node (device noise).



- Extremely low $\sim 10^{-6}$ inefficiency in charge transfer.
- Extremely low leakage current $\sim 7 \text{ e-}/\text{cm}^2/\text{day}$. [arXiv:2410.18716](#)

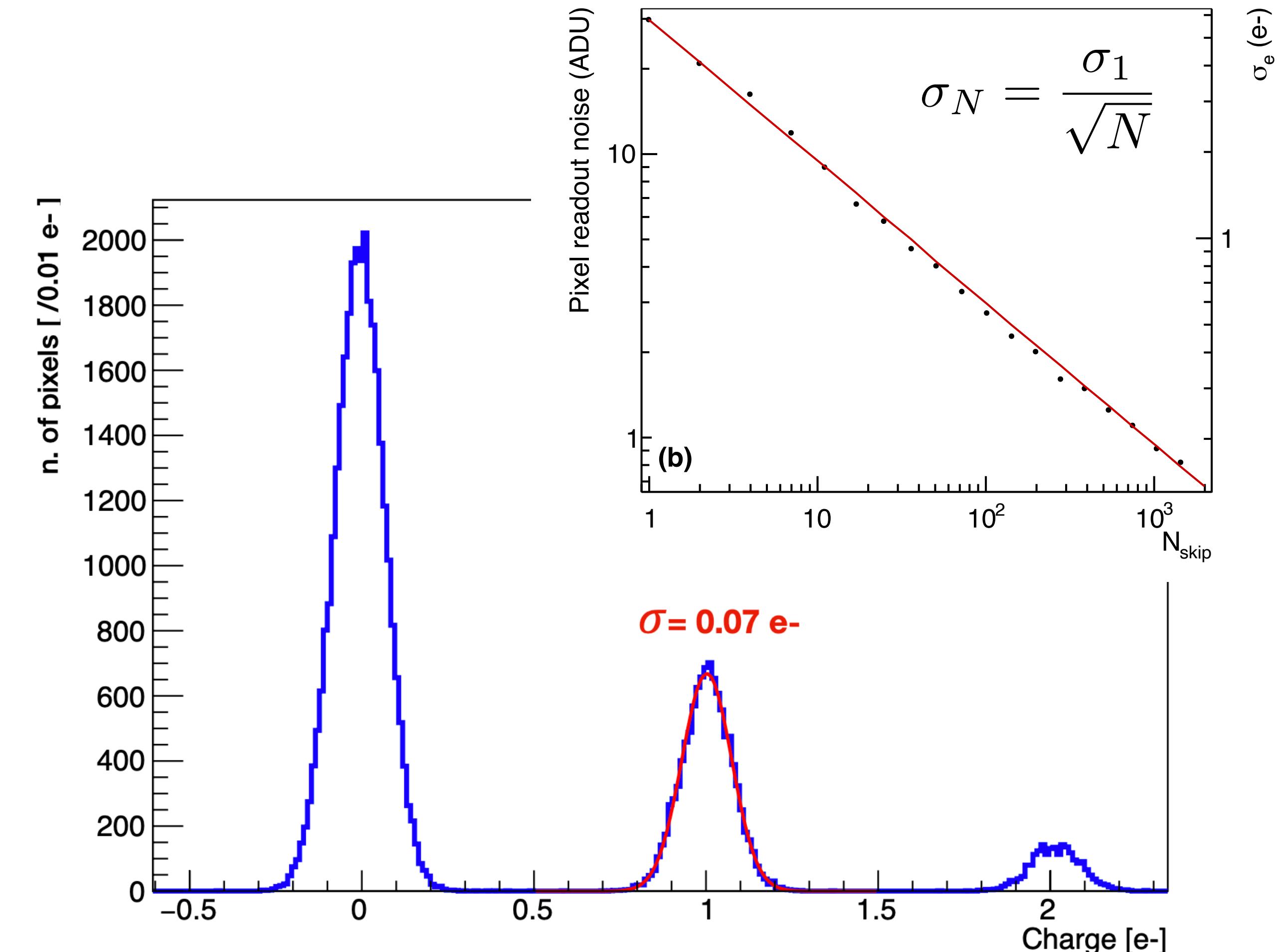
Skipper readout

“Skipper” readout:
Perform N uncorrelated
measurements of the same pixel.

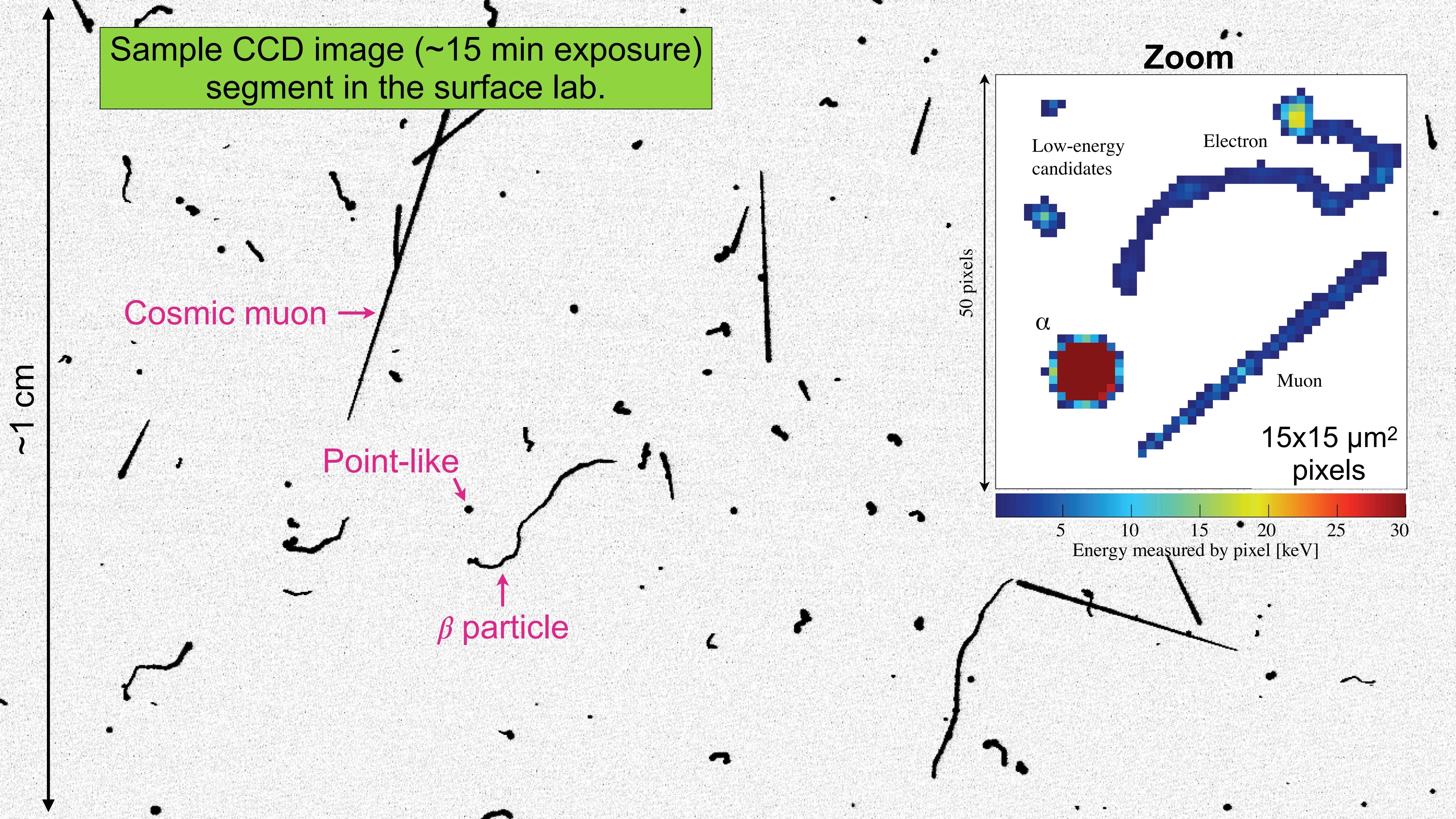


Introduced to particle physics in 2017

PRL119(2017)131802



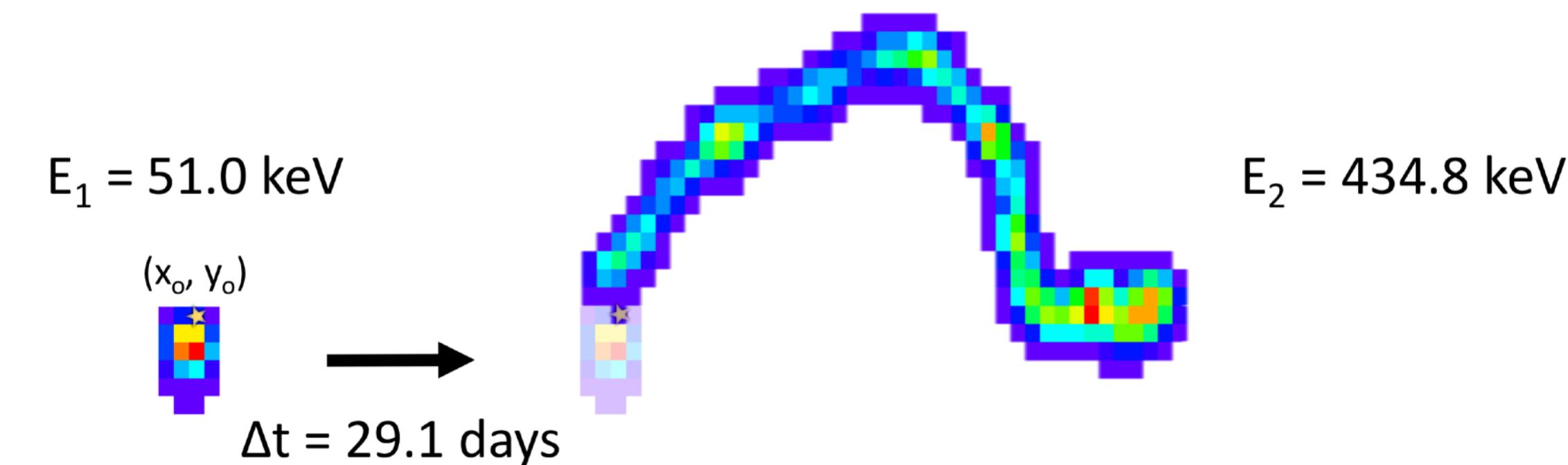
Sample CCD image (~15 min exposure)
segment in the surface lab.



Spatial resolution

- Surface background rejection by depth (z) reconstruction, and classification (α , β , NR) by track topology (at high $E > 80 \text{ keV}_{\text{ee}}$).
- Spatial coincidence searches to identify decay sequences: [JINST16\(2021\)P06019](#)
- **Cosmogenic ^{32}Si :** $^{32}\text{Si} (\text{T}_{1/2} = 150 \text{ y}, \beta) \rightarrow ^{32}\text{P} (\text{T}_{1/2} = 14 \text{ days}, \beta)$

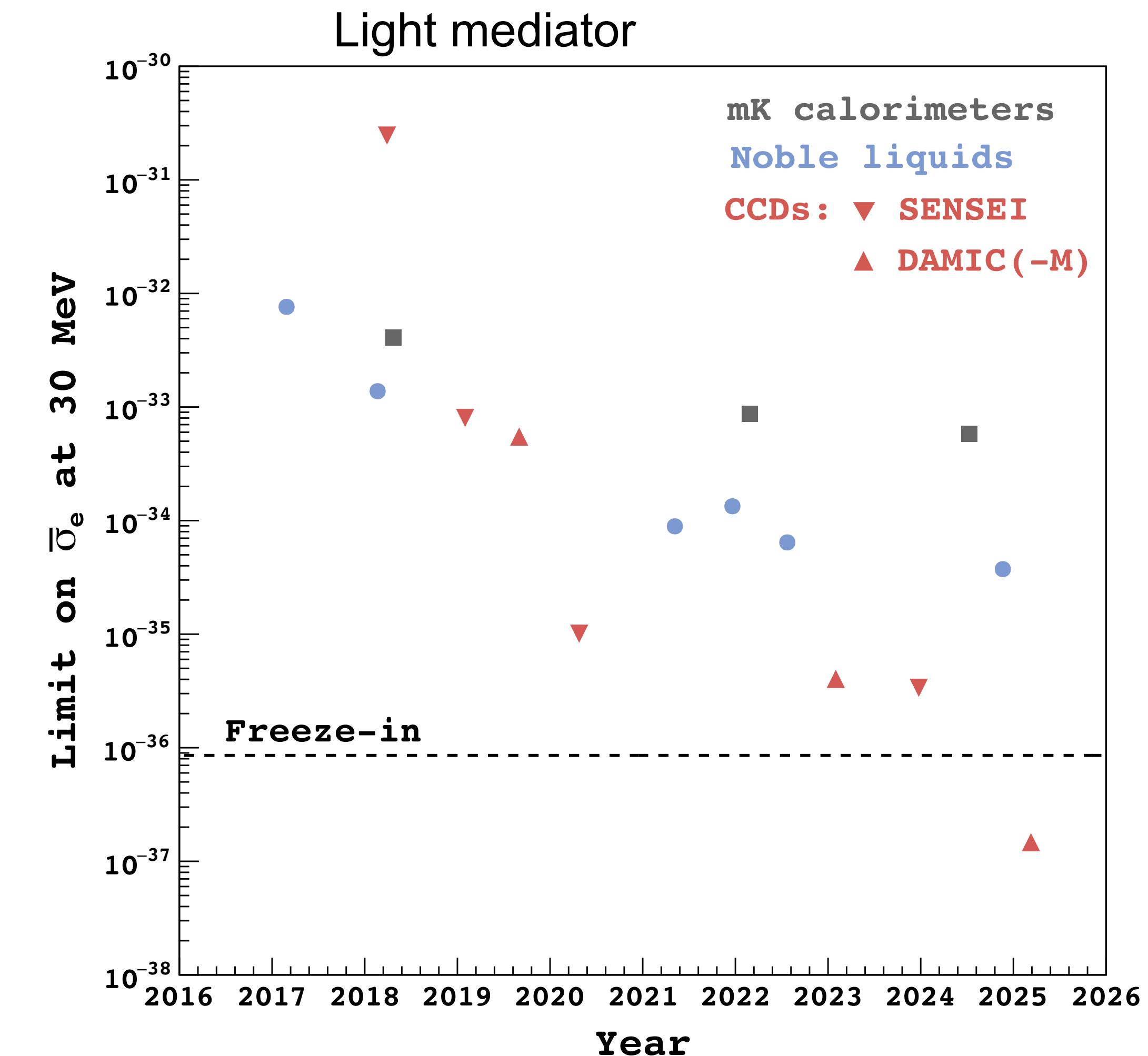
$140 \pm 30 \mu\text{Bq} / \text{kg}$



- Also upper limits on every β emitter in the U/Th chain.
- Reject crystal defects “hot spots” that dominate device leakage current.
- NR identification by spatial correlation between ionization event and defect left behind in the crystal (R&D): [PRD110\(2024\)043008](#)

CCD timeline

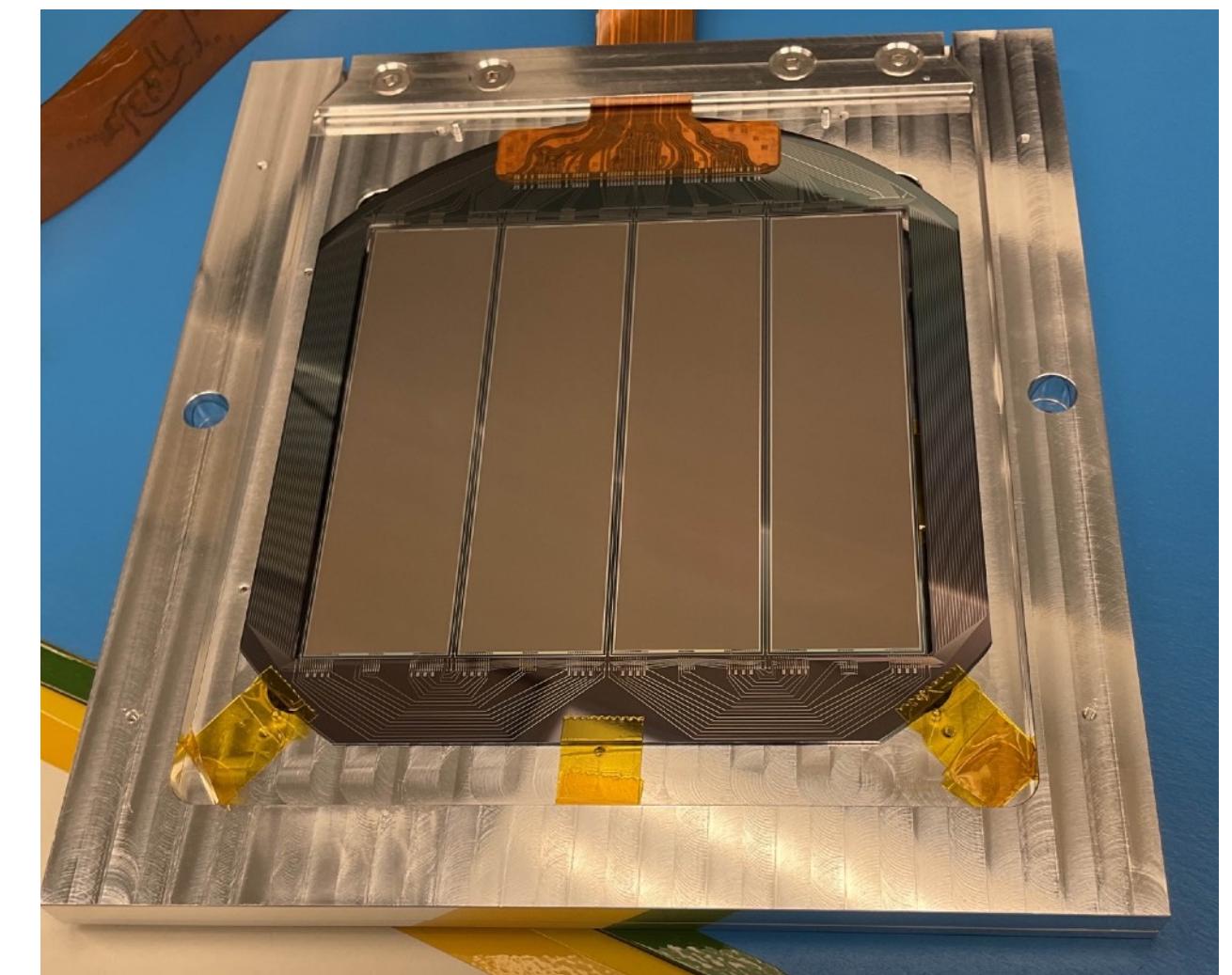
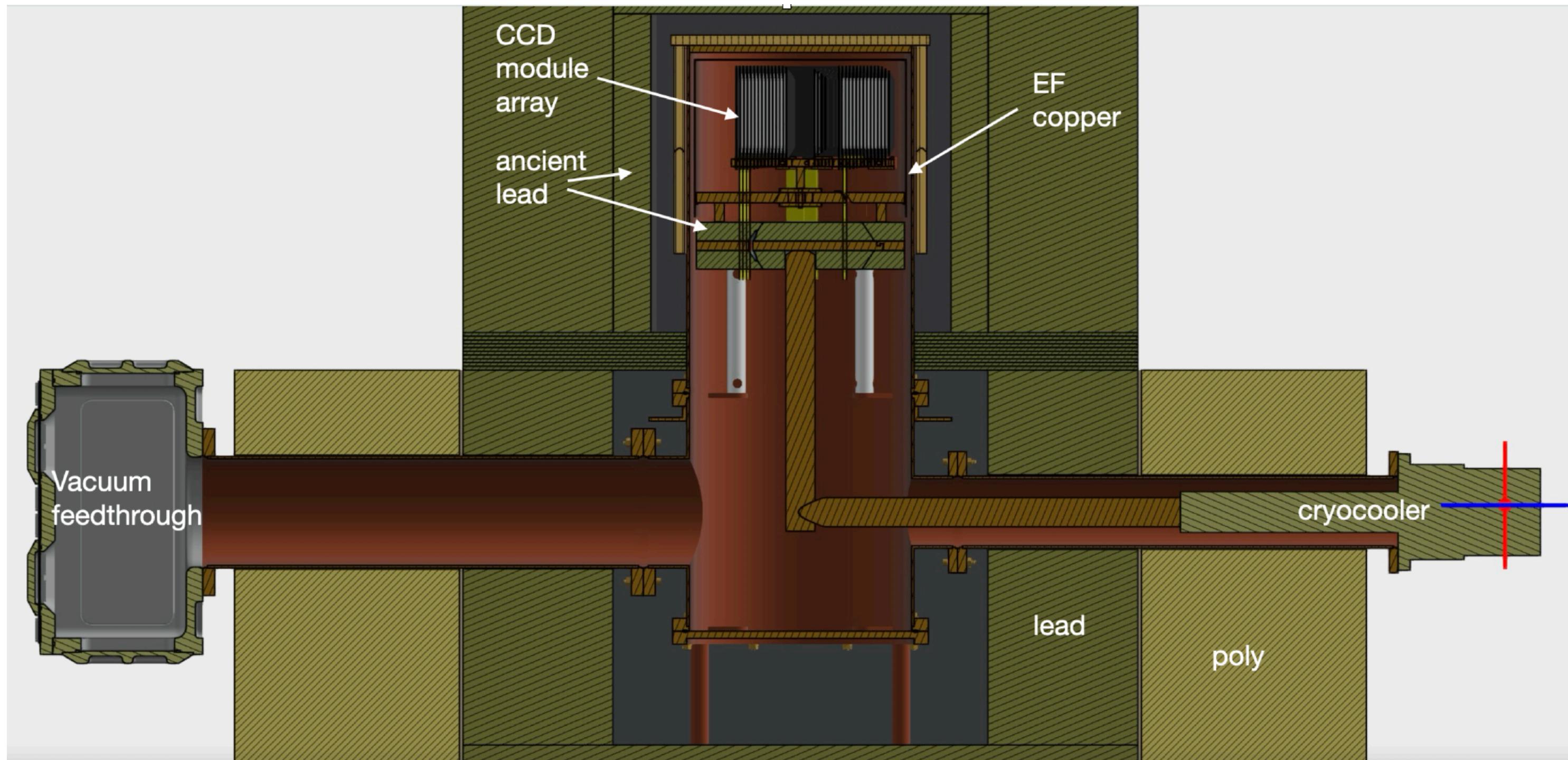
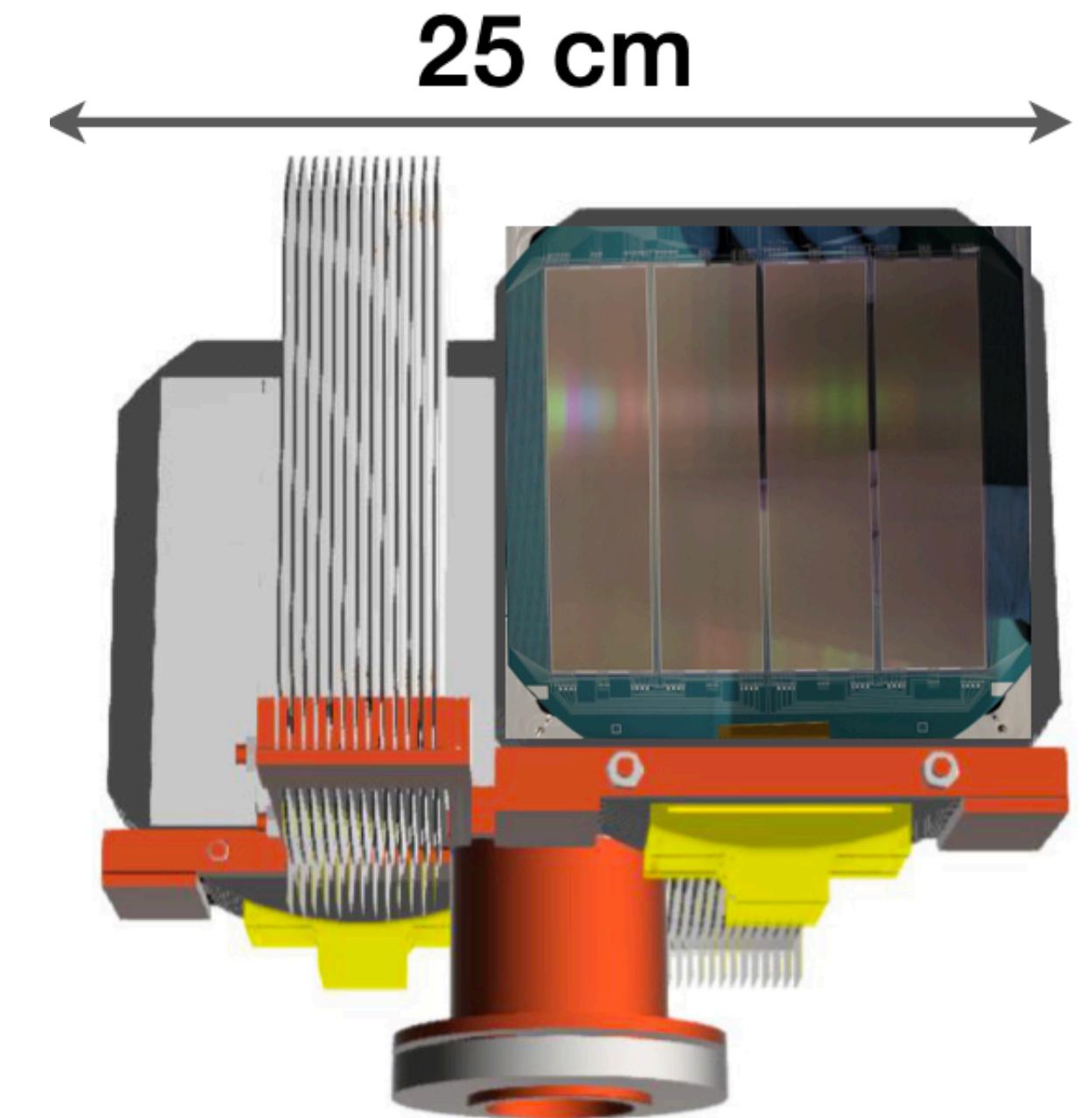
- ▶ **2012–2017:** we built DAMIC at SNOLAB, the first low background CCD array for DM searches.
- ▶ **2017:** DAMIC releases first DM search results from \sim eV ionization signals.
- ▶ **2017:** “Skipper” CCDs are introduced for 10x improvement in noise to provide increased sensitivity for DM-e searches.
- ▶ Two skipper-CCD experiments started: **SENSEI** and **DAMIC-M**.
- ▶ Multiple detector iterations, fast progress from both collaborations until now.
- ▶ **2025:** DAMIC-M’s LBC probes benchmark hidden-sector models: **Highlight of today**
- ▶ **2026+:** DAMIC-M coming online with >100 CCDs!



DAMIC-M

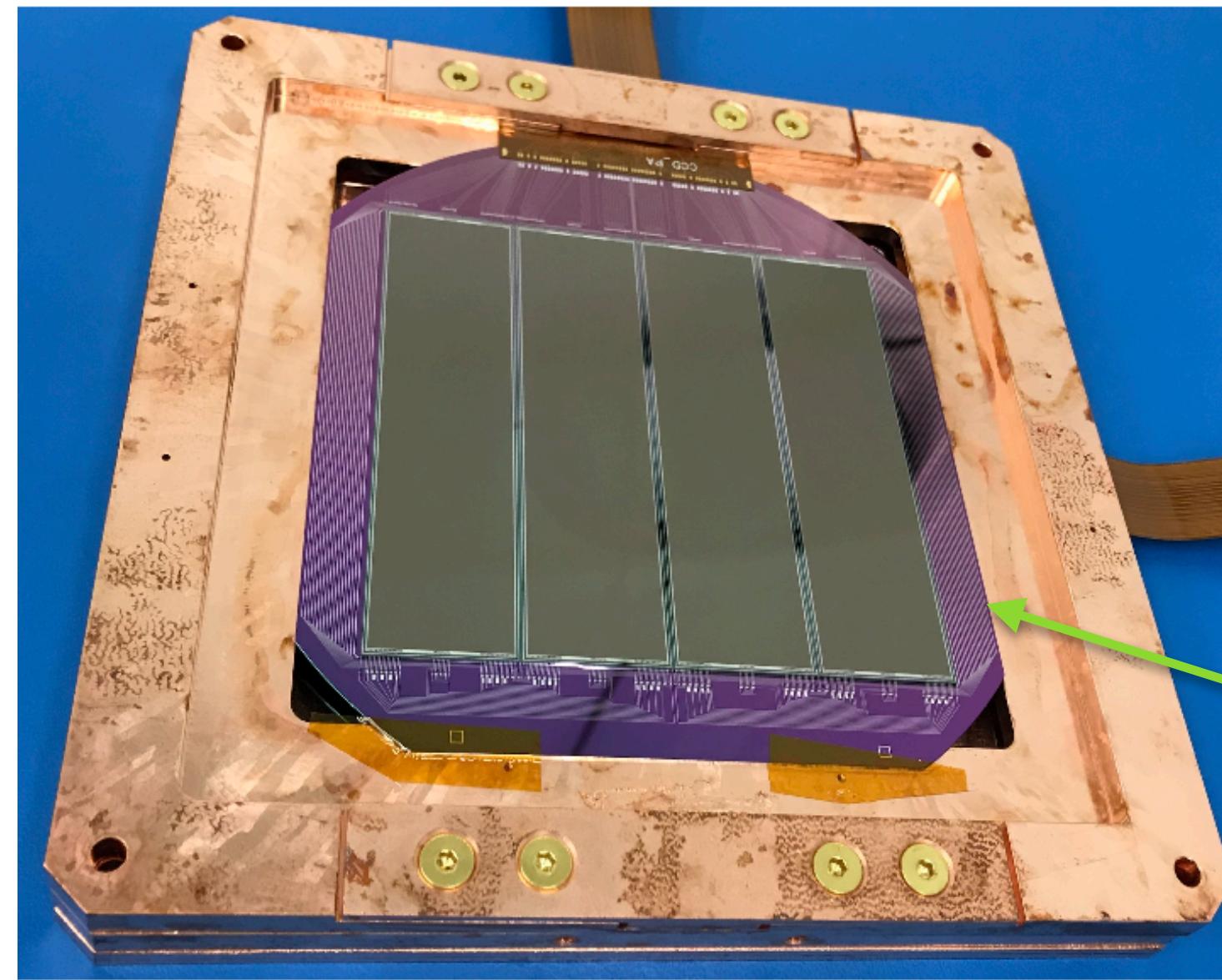
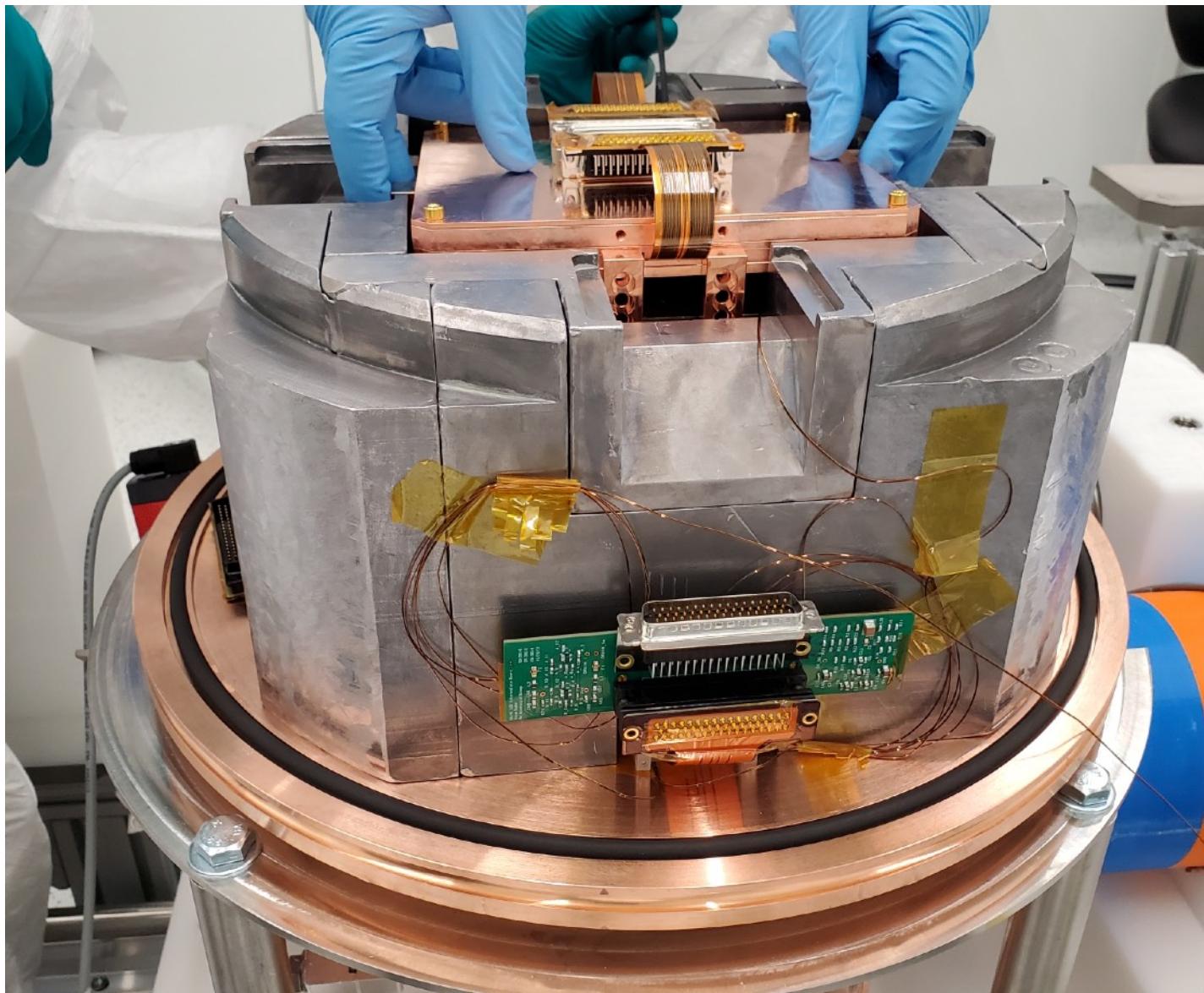
DAMIC-M

- ▶ 52 CCD modules in LSM (France) for kg-year target exposures.
- ▶ Skipper readout for 2 or 3 e- threshold.
- ▶ Background reduction to a fraction of d.r.u. (events per kg-day).
- ▶ Under construction. Commissioning by end of 2025!



Law Background Chamber

- ▶ Low Background Chamber (LBC) test setup for DAMIC-M at LSM for performance and background studies.
- ▶ Operating in LSM clean room since 2022.
- ▶ Several detector iterations. Details in [JINST19\(2024\)T11010](#)
- ▶ First science results. Spectral analysis: [PRL130\(2023\)171003](#)
- ▶ Daily modulation: [PRL132\(2024\)101006](#)

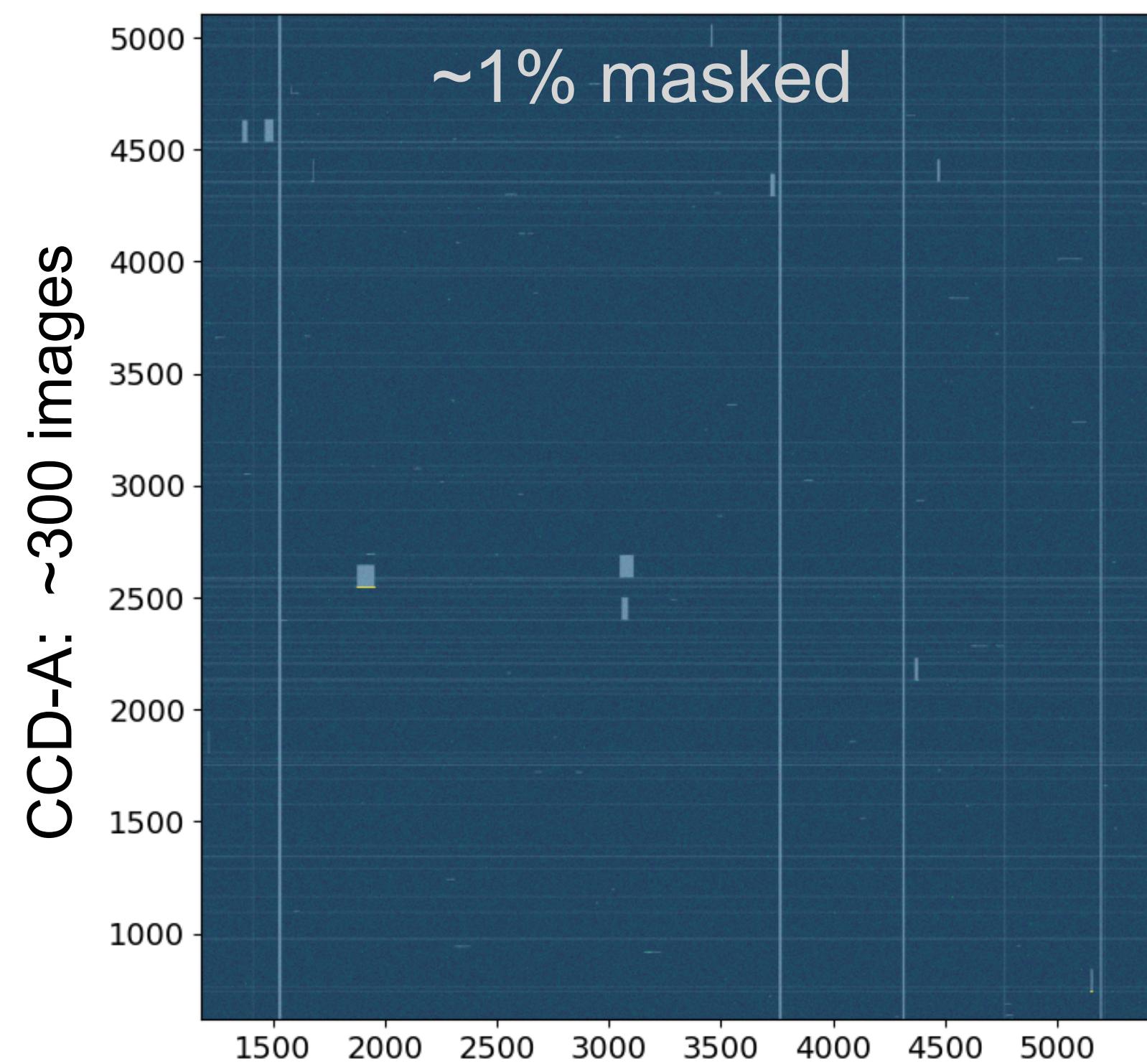


Prototype CCD module
packaged and tested at UW

Science Run 2

Detector upgrades

- Two DAMIC-M modules (8 CCDs for 26g)
- electroformed copper box lids
- DAMIC-M low-noise electronics



Parameters

- read out 1 amplifier per CCD
- binning: 1 pixel x 100 pixel (col x row)
- temperature: ~130 K

Performance

- reduced dark current: $\sim 10^{-4}$ e-/pixel/day (previous 50x)
- background: ~15 dru with shield partly open
- readout noise = 0.16e- with 500 skips
- data set exposure: 1.3 kg-day (previous 85 g-day)

Image masking – 95% of data are kept

- hot regions in CCDs (large 1e- rate)
- clusters of high-charge pixels ($\geq 6e^-$)
- clusters in CCDs of same module (cross-talk)
- charge-correlated pixels in CCDs of same module
- 100 pixels above + row of pixel with >100 e- (charge traps)

Pattern analysis

Blind analysis

- Data set 1 (D1): selection sample (130 g-day)
- Data set 2 (D2): blinded analysis set (1.3 kg-day)

Candidate selection

- look for horizontal or consecutive pixels with 2, 3, or 4 e-:
 $\{11\}$, $\{21\}$, $\{111\}$, $\{31\}$, $\{22\}$, $\{211\}$
- exclude isolated pixels with $\geq 2e^-$

Efficiency

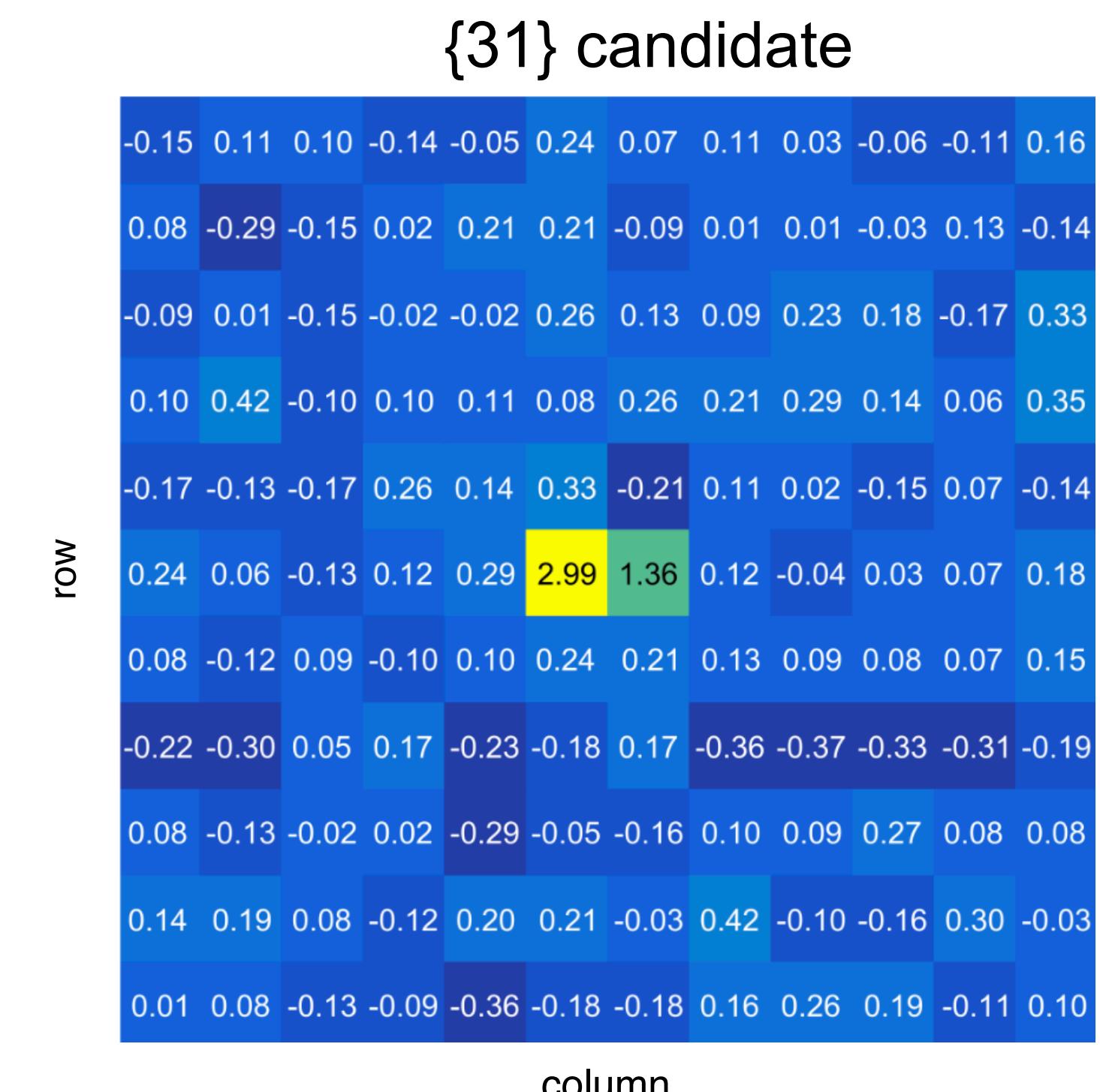
- calculate probability to obtain pattern from ionization events with initial charge N_e (includes charge diffusion and noise)

Backgrounds

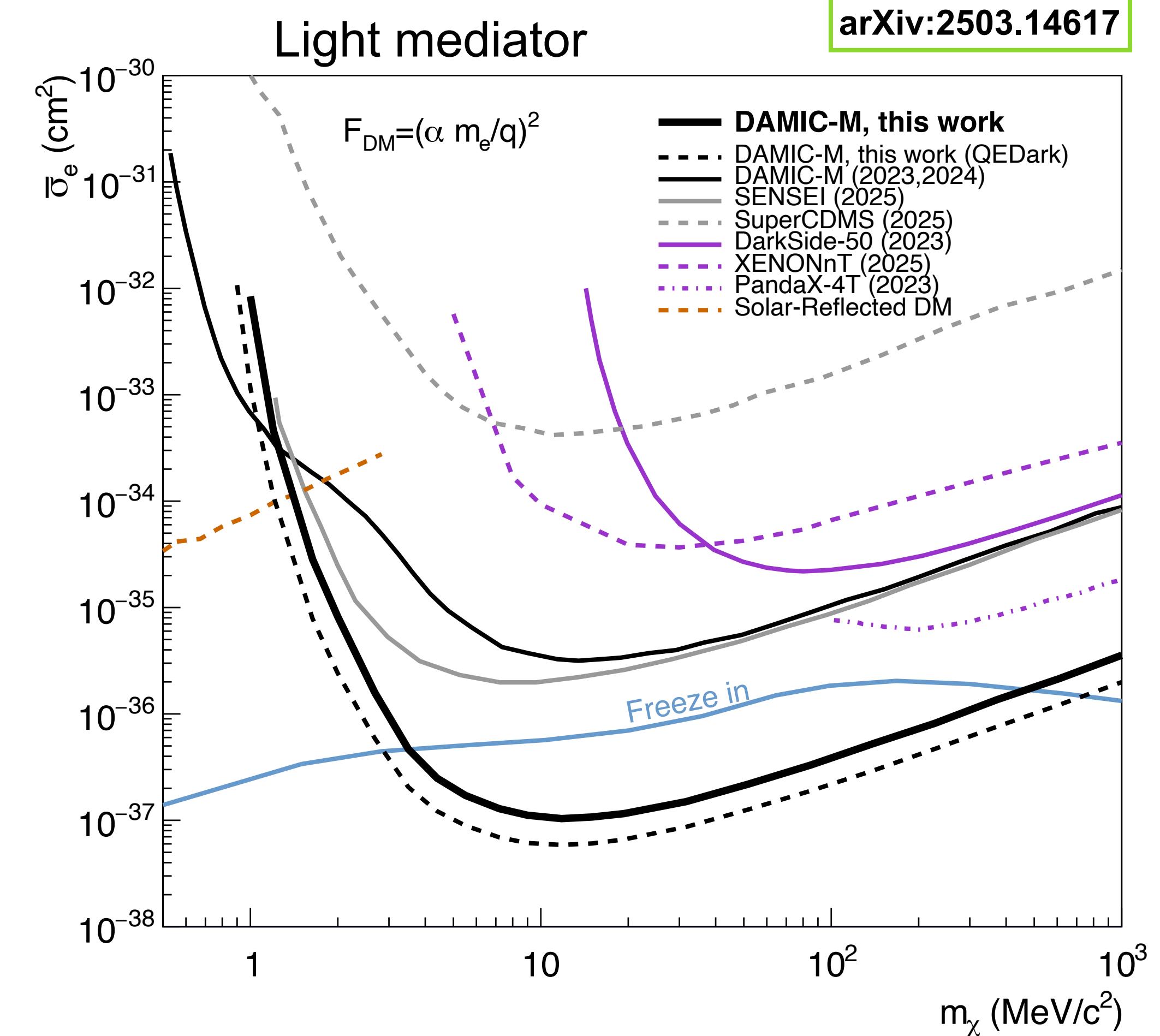
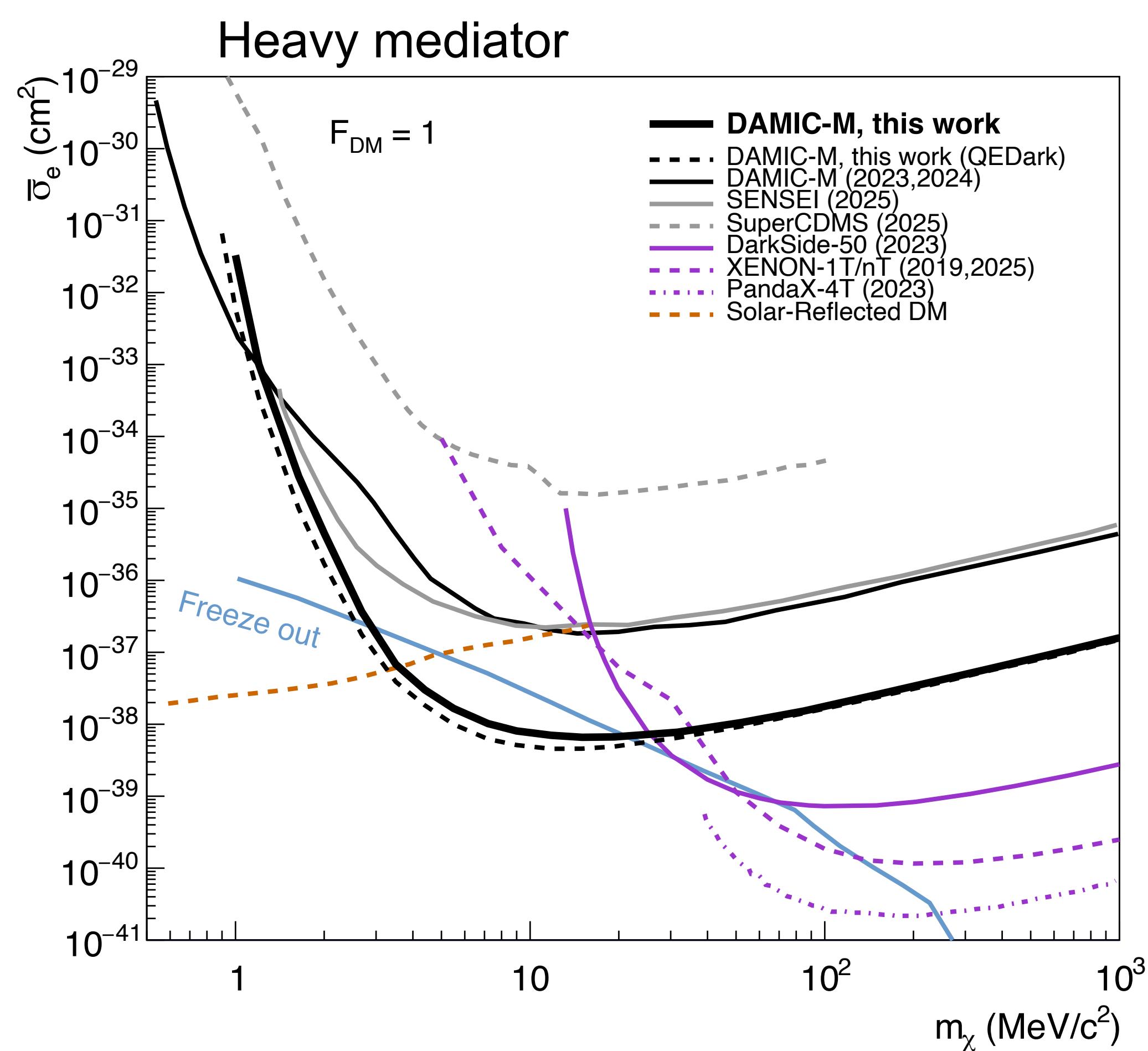
- estimate radiogenic background by scaling measured high energy events (2.5 to 7.5 keV) with Geant4
- random coincidences of uncorrelated pixels next to each other evaluated toy MC

	Pattern p		
	$\{11\}$	$\{21\}$	$\{111\}$
D_p	144	0	0
B_p^{rc}	141.4	0.111	0.042
B_p^{rad}	0.039	0.039	0.016
	$\{31\}$	$\{22\}$	$\{211\}$
D_p	1	0	0
B_p^{rc}	0.019	$2.5 \cdot 10^{-5}$	$5.8 \cdot 10^{-5}$
B_p^{rad}	0.052	0.011	0.035

TABLE I. The number of candidates D_p in the D2 data set, and the number expected from backgrounds due to random coincidences, B_p^{rc} , and to radioactive decays, B_p^{rad} .



DM-e exclusion limits

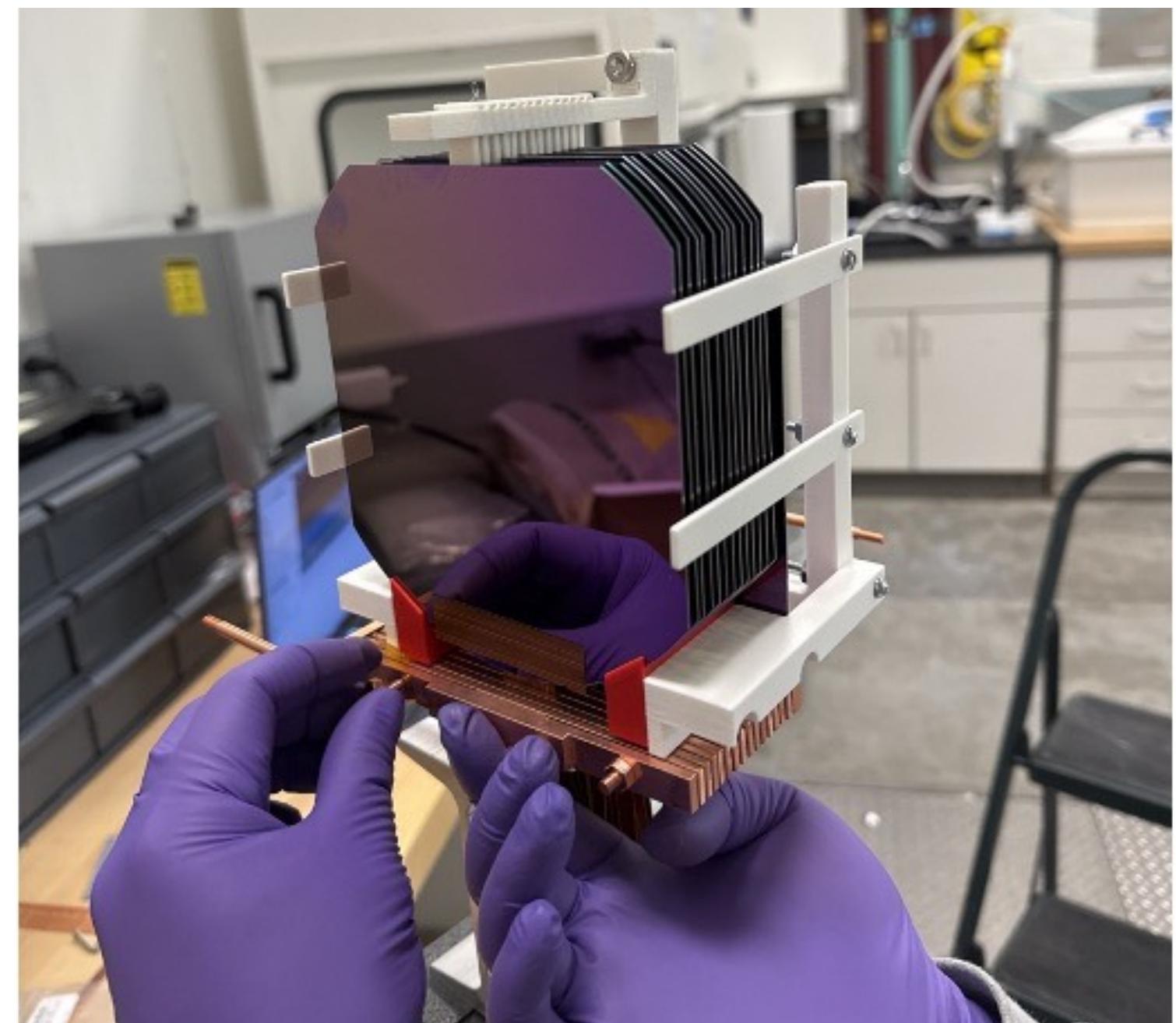
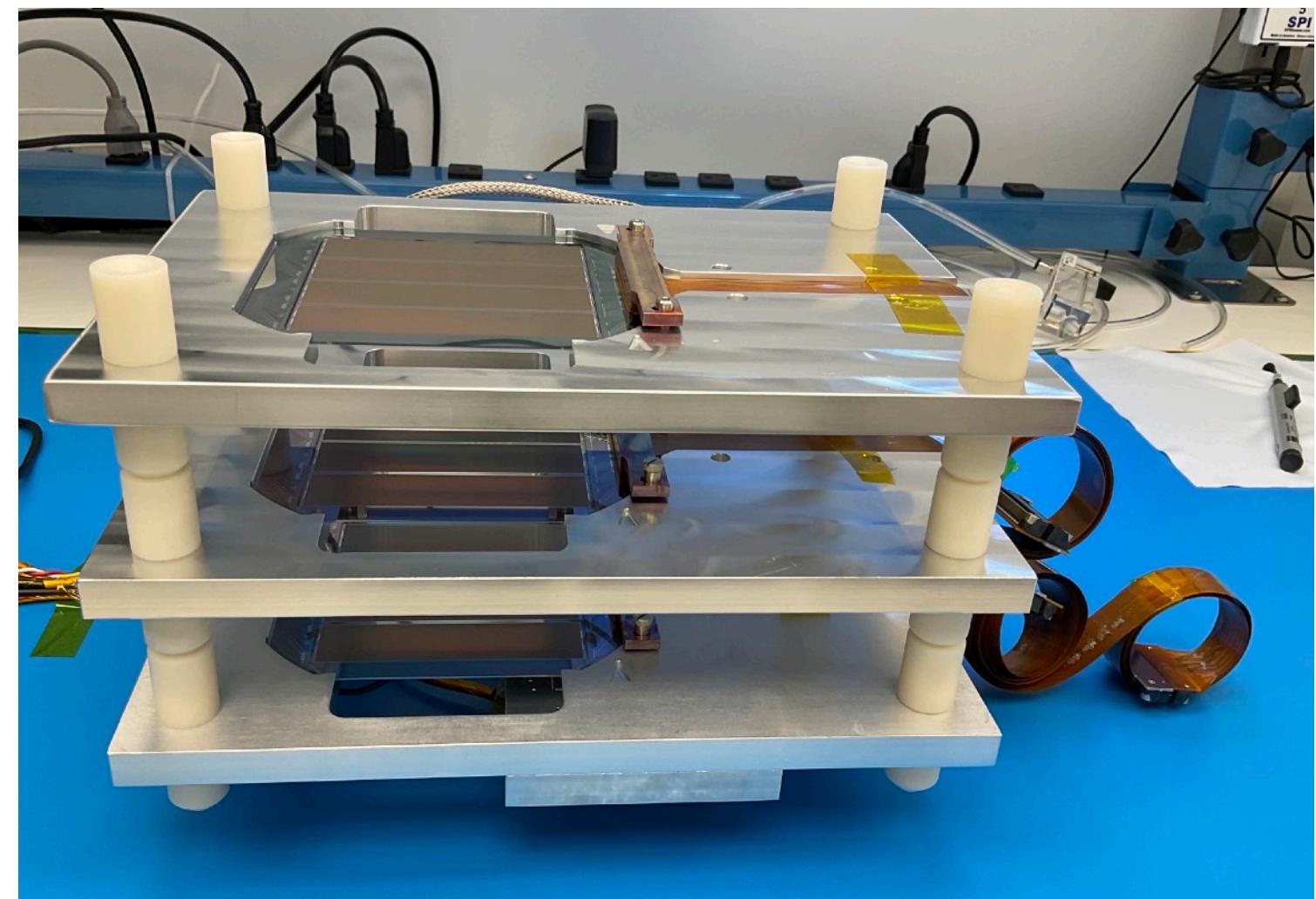


arXiv:2503.14617

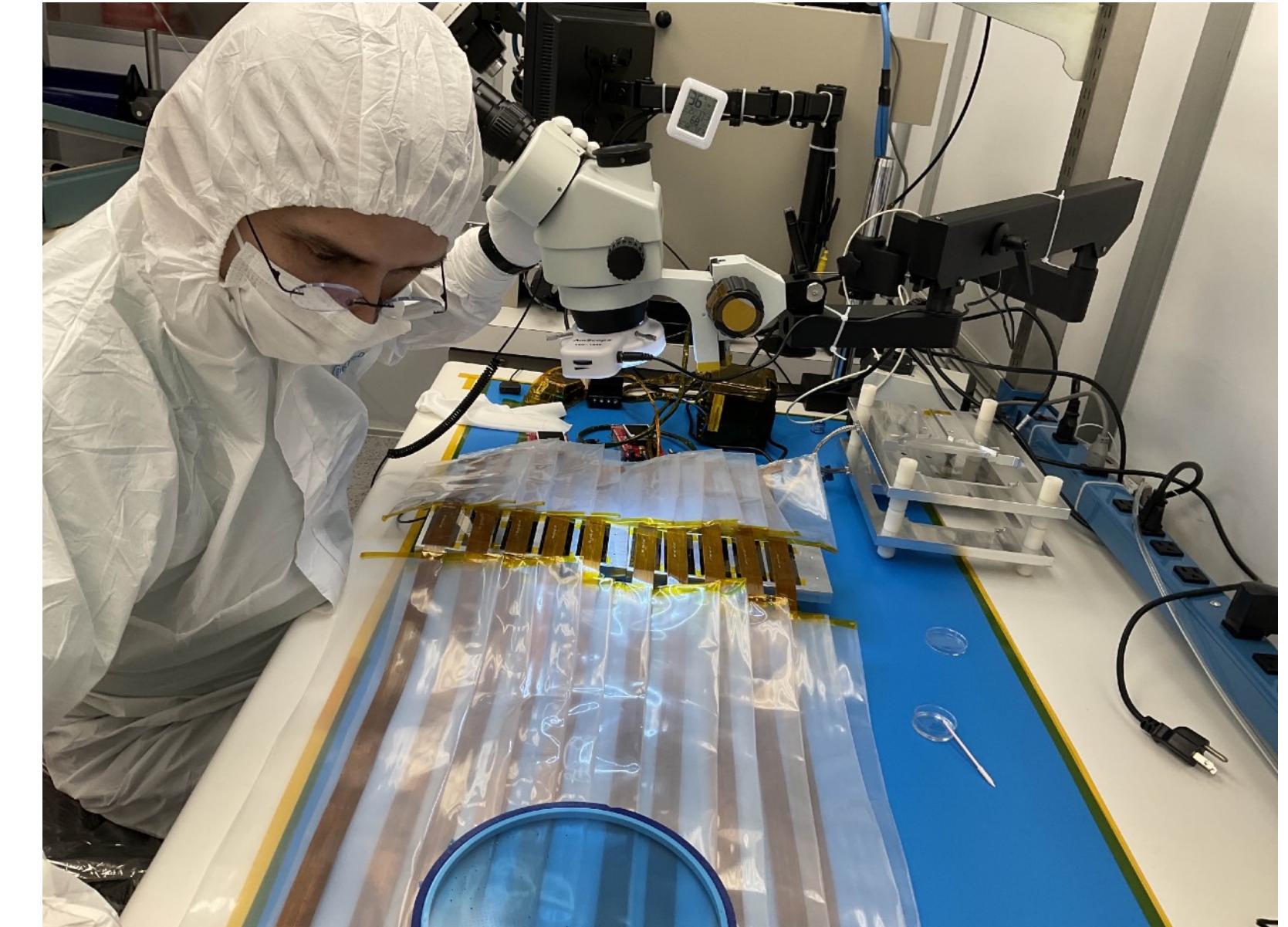
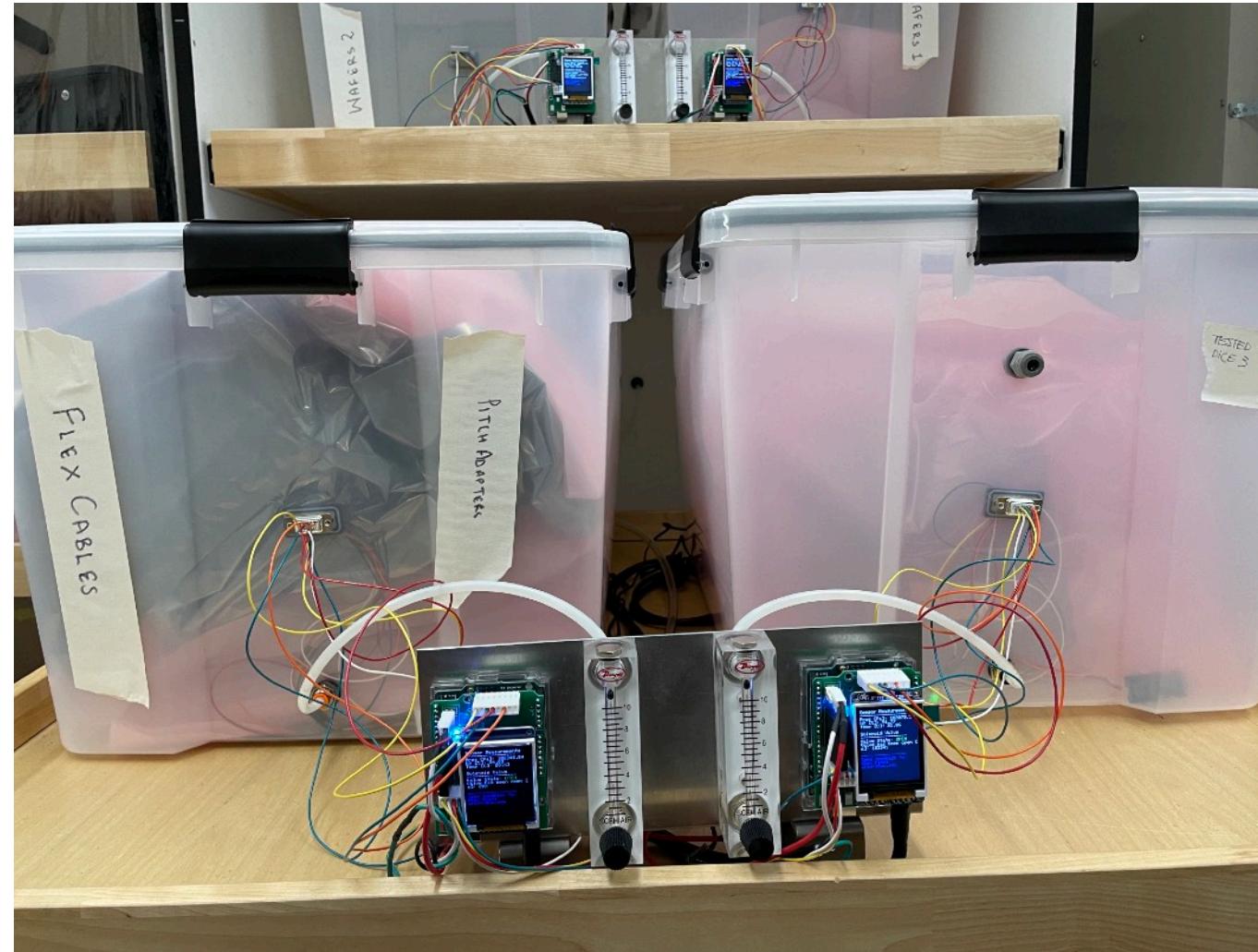
► DAMIC-M probes benchmark hidden-sector dark-matter models!

DAMIC-M Status

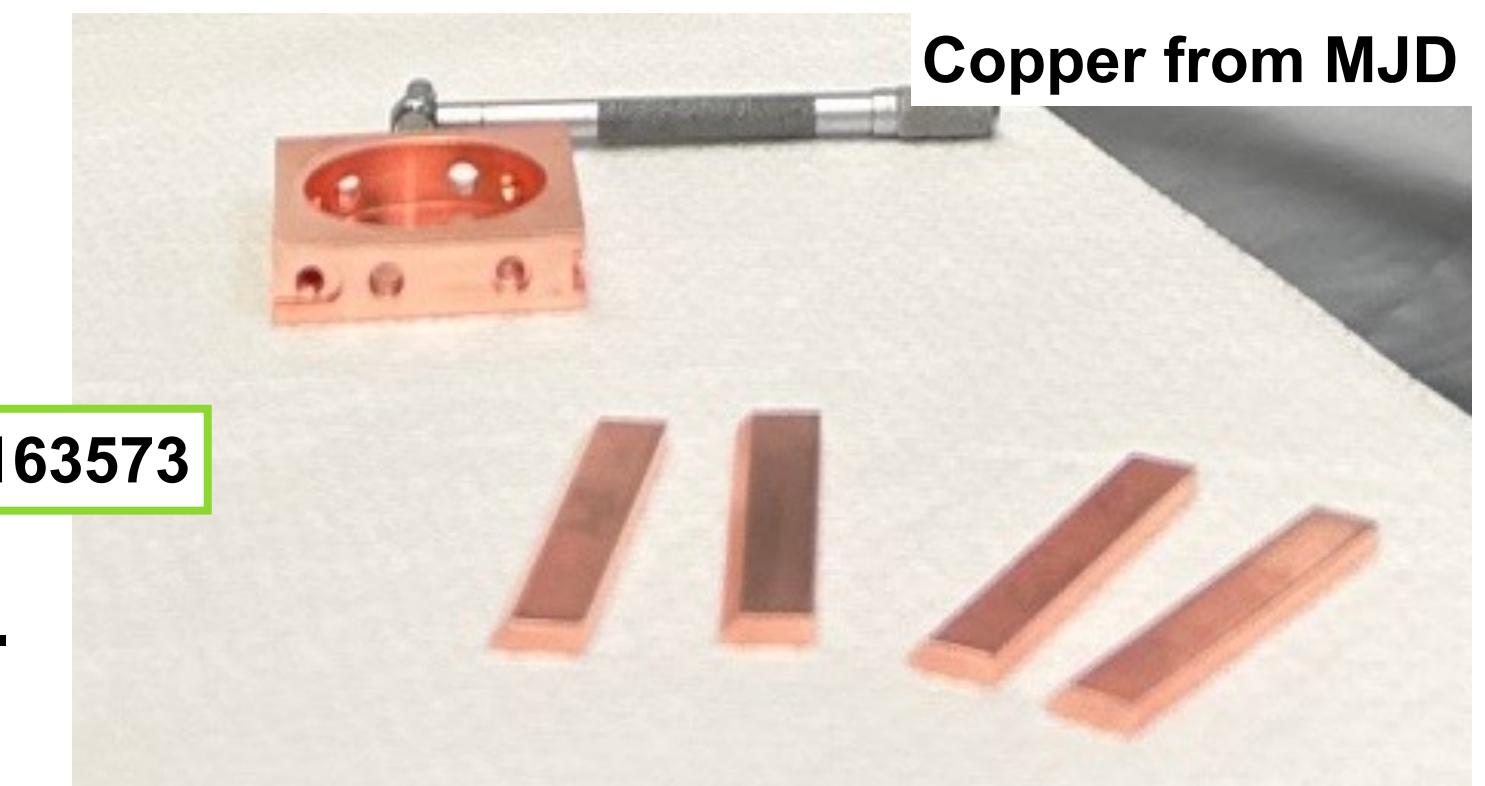
- Tested 188 CCDs at UW in Q3 2024.
- 28 CCD modules fabricated, shipped to LSM in their shielded container.
- LBC relocated to second clean room at LSM.
- Reconditioning clean rooms for CCD array assembly and detector installation.
- Copper machining ongoing, lead and poly being prepared for shipment.
- Custom electronics under procurement and testing.
- Detector installation in second-half 2025, commissioning by end of 2025.



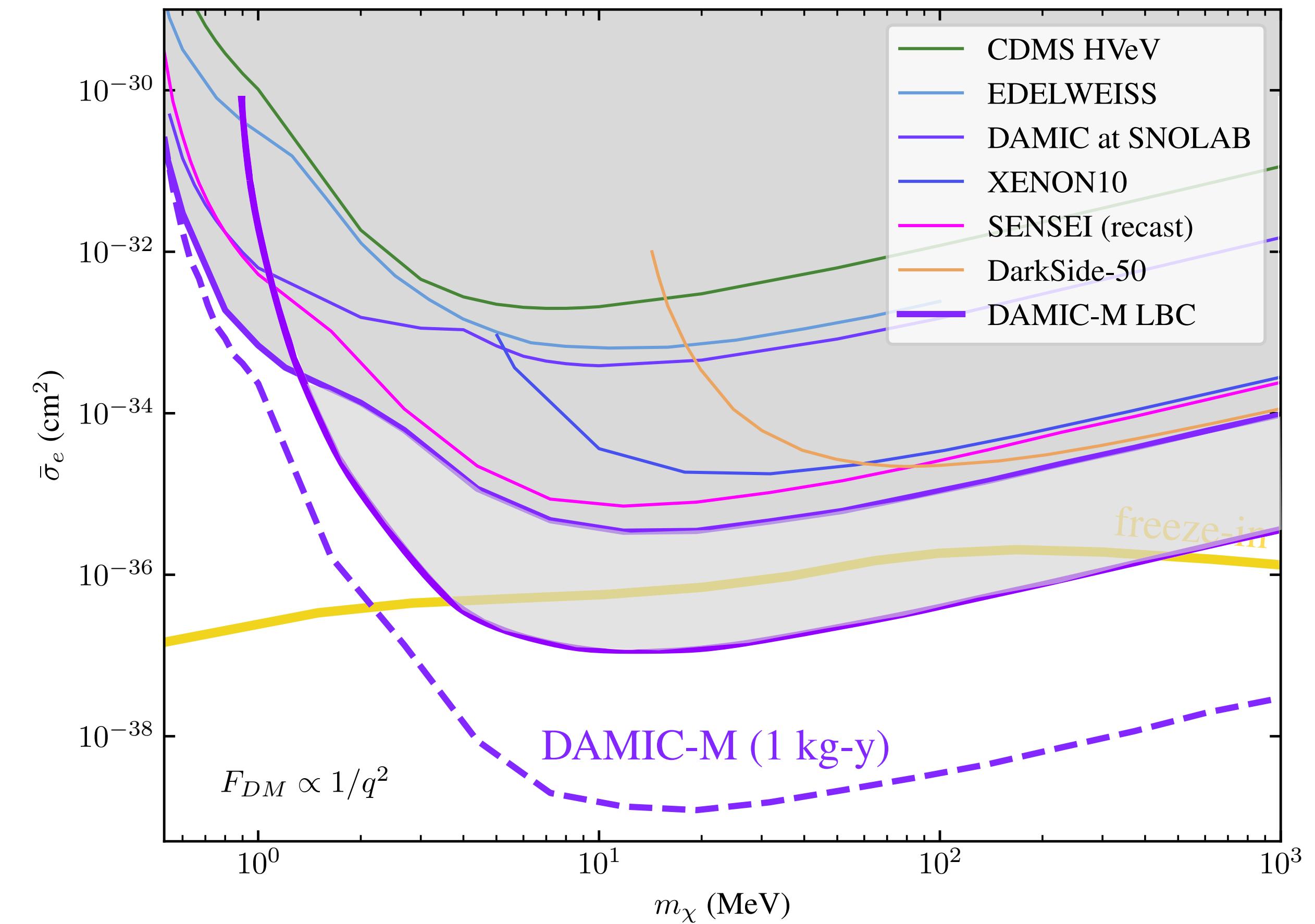
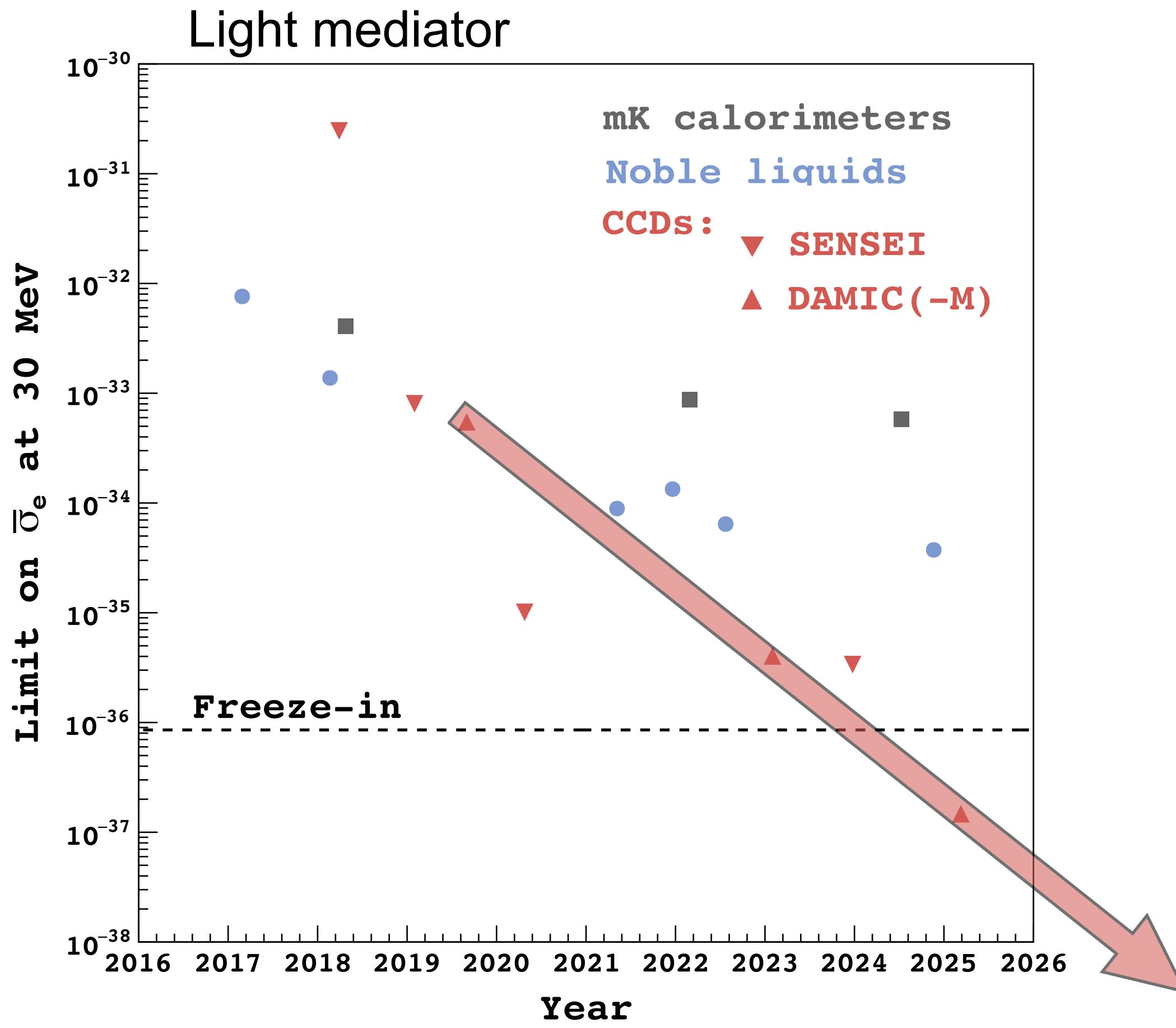
Low-background



- ▶ Transport in shielded container.
- ▶ Shielded, Rn-free storage.
- ▶ Clean room operations.
- ▶ Low-radioactivity flexes. NIMA959(2020)163573
- ▶ Copper electroformed underground.
- ▶ Light-tight infrared shield.



DAMIC-M Forecast



Conclusions

- The range of DM particle masses searched for by direct detection has expanded greatly in recent years.
- DM-e⁻ scattering is a powerful probe for sub-GeV DM particles.
- Charge-coupled device (CCD) experiments lead the sub-GeV mass window.
- DAMIC pioneered the use of CCDs to search for dark matter.
- Steadfast progress by skipper-CCD experiments SENSEI and DAMIC-M in the last 8 years.
- DAMIC-M's LBC now excludes several hidden-sector benchmark models.
- Progress should continue for at least one more detector generation with DAMIC-M.

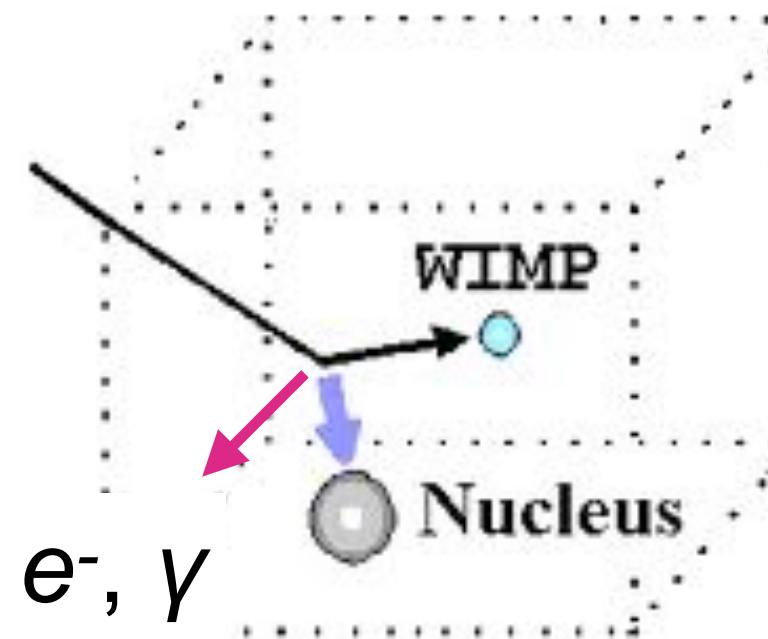
Thank you!



Backup Slides

Other e-recoils

Three-body final state:



- ▶ An additional e^- or γ in the final state.
- ▶ Migdal effect (atomic e^-) or Bremsstrahlung (γ).
- ▶ E and p can be conserved even when e^- or γ take most of the WIMP kinetic energy.
- ▶ Probability of e^- or γ emission $< 10^{-6}$. Rare.
- ▶ Never observed for recoils with keV energies. Uncalibrated.

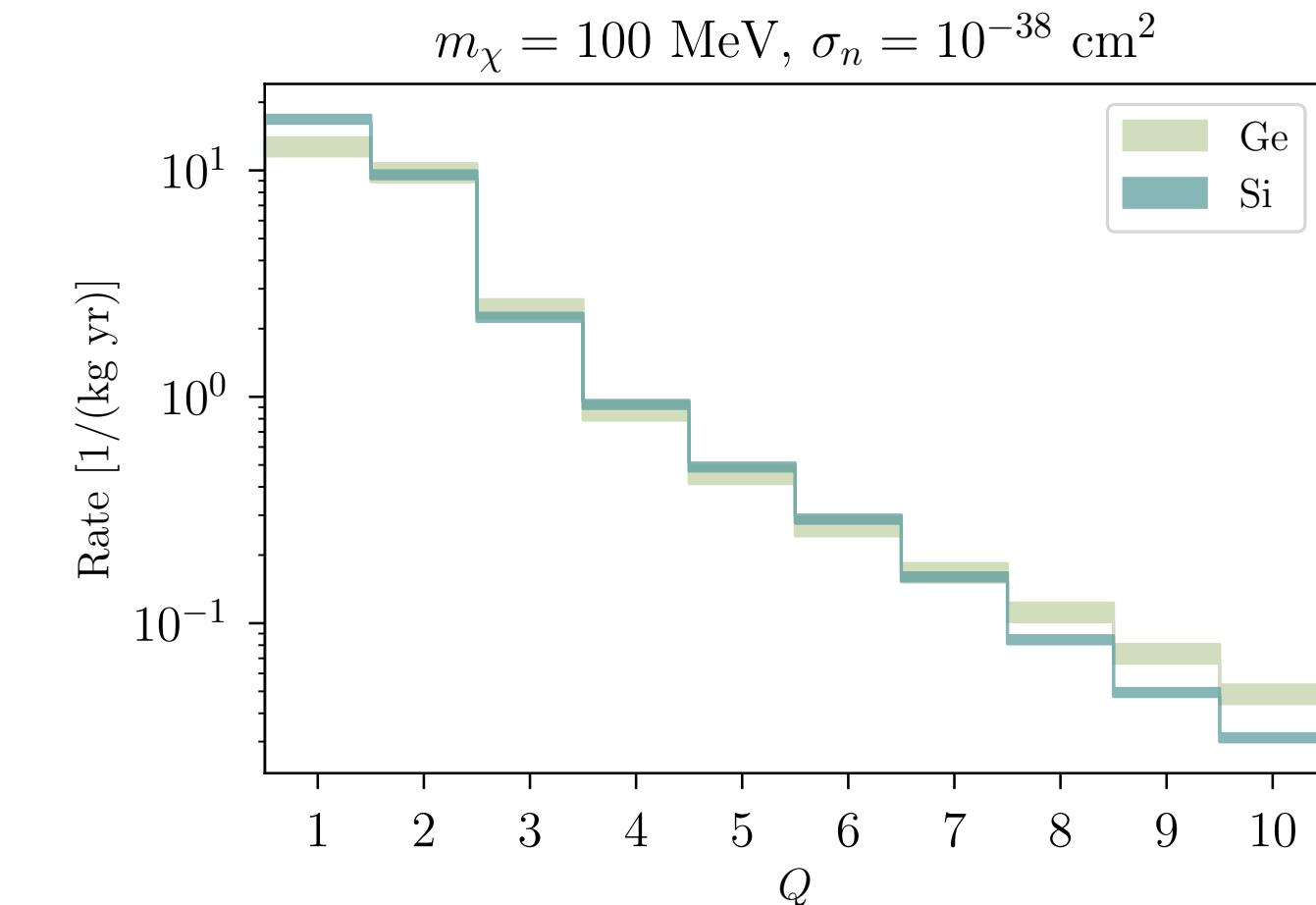
Bosonic DM absorption:

- ▶ DM particle is a boson that couples to the electron, e.g., a “dark” or “hidden” photon.
- ▶ DM is absorbed by the target electron and its rest energy released as electronic recoil K.E.

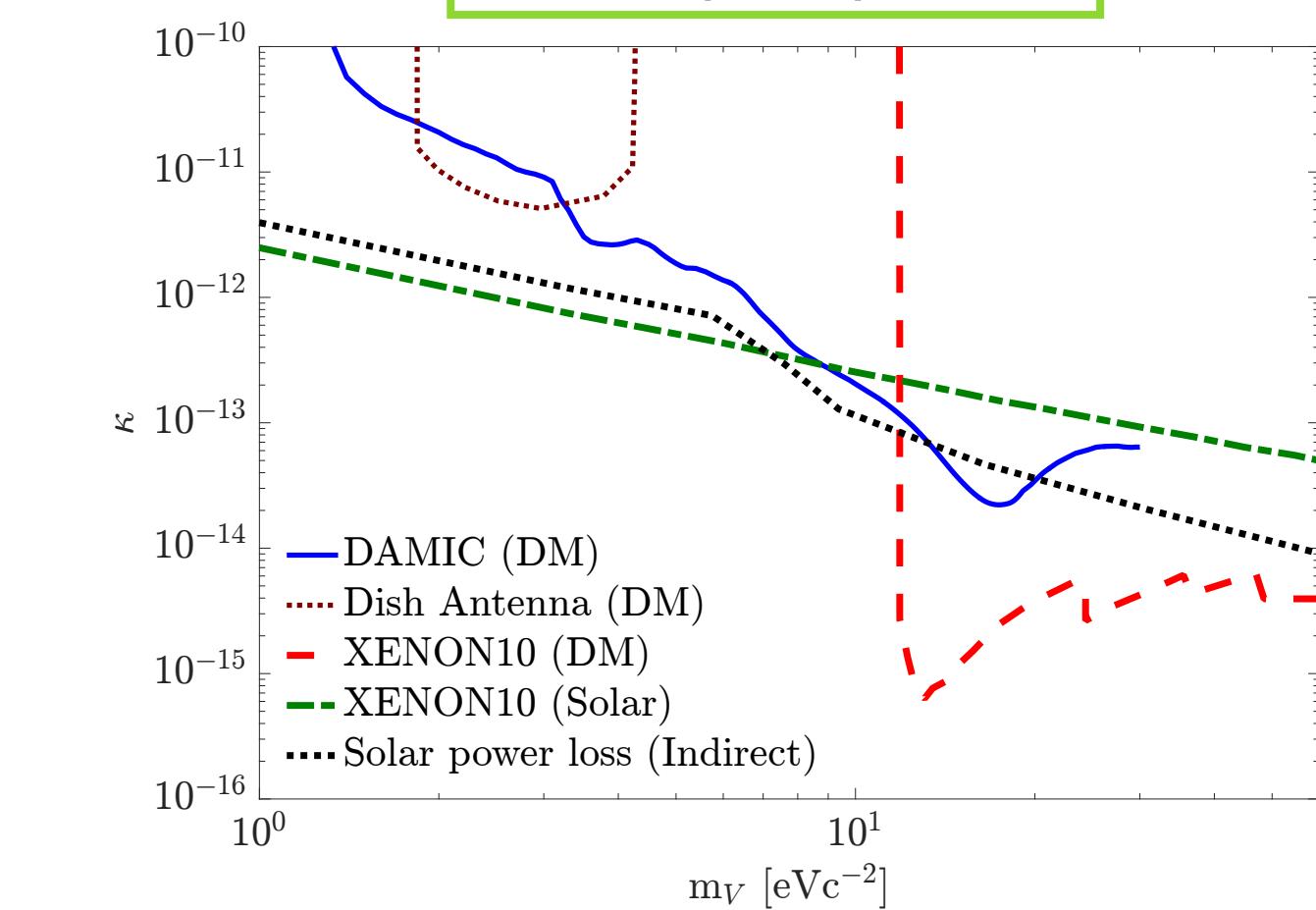
Electronic recoil result could also be interpreted as limit on DM-N scattering (Migdal) or DM absorption

I will use DM-e scattering parameter space as benchmark

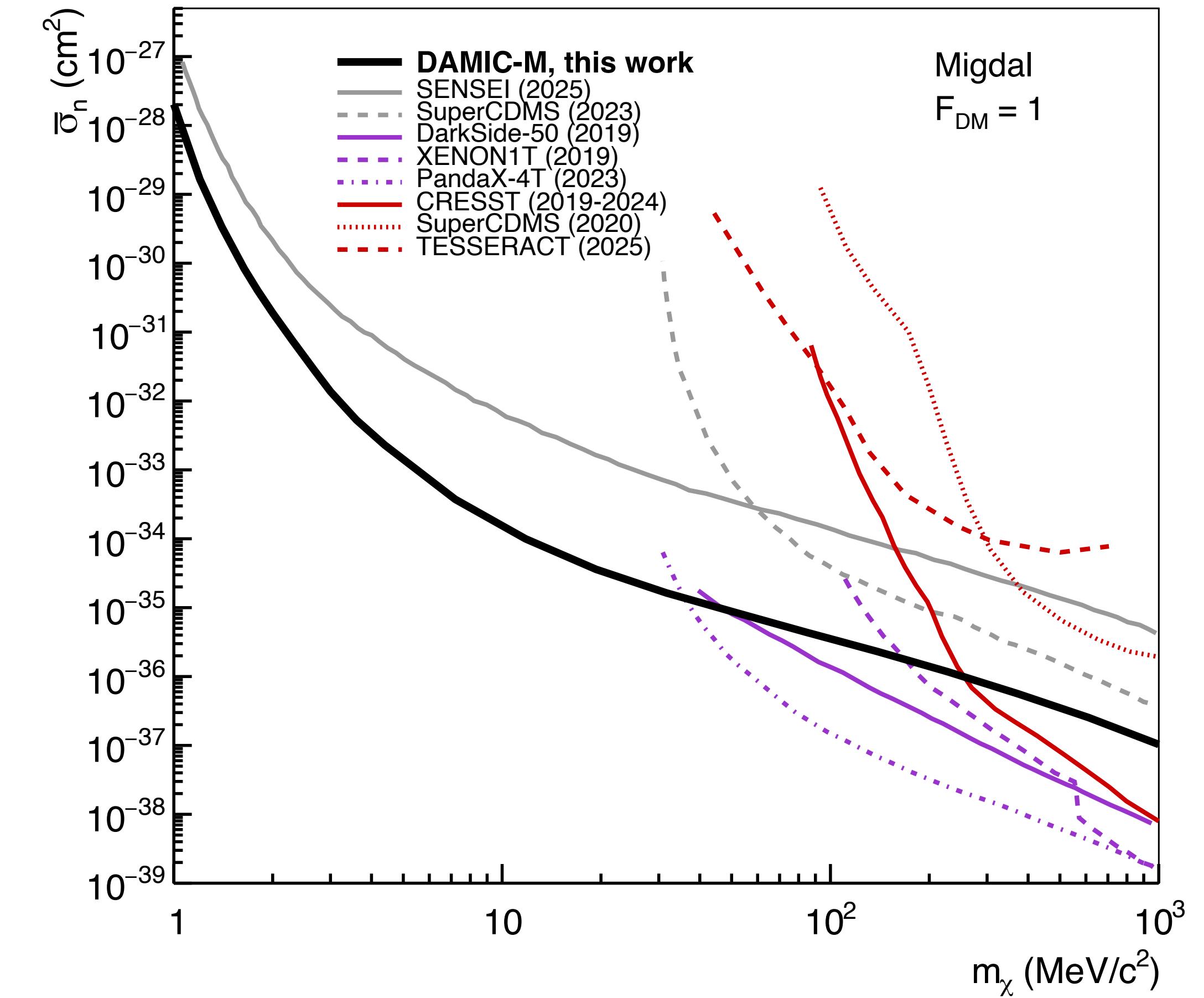
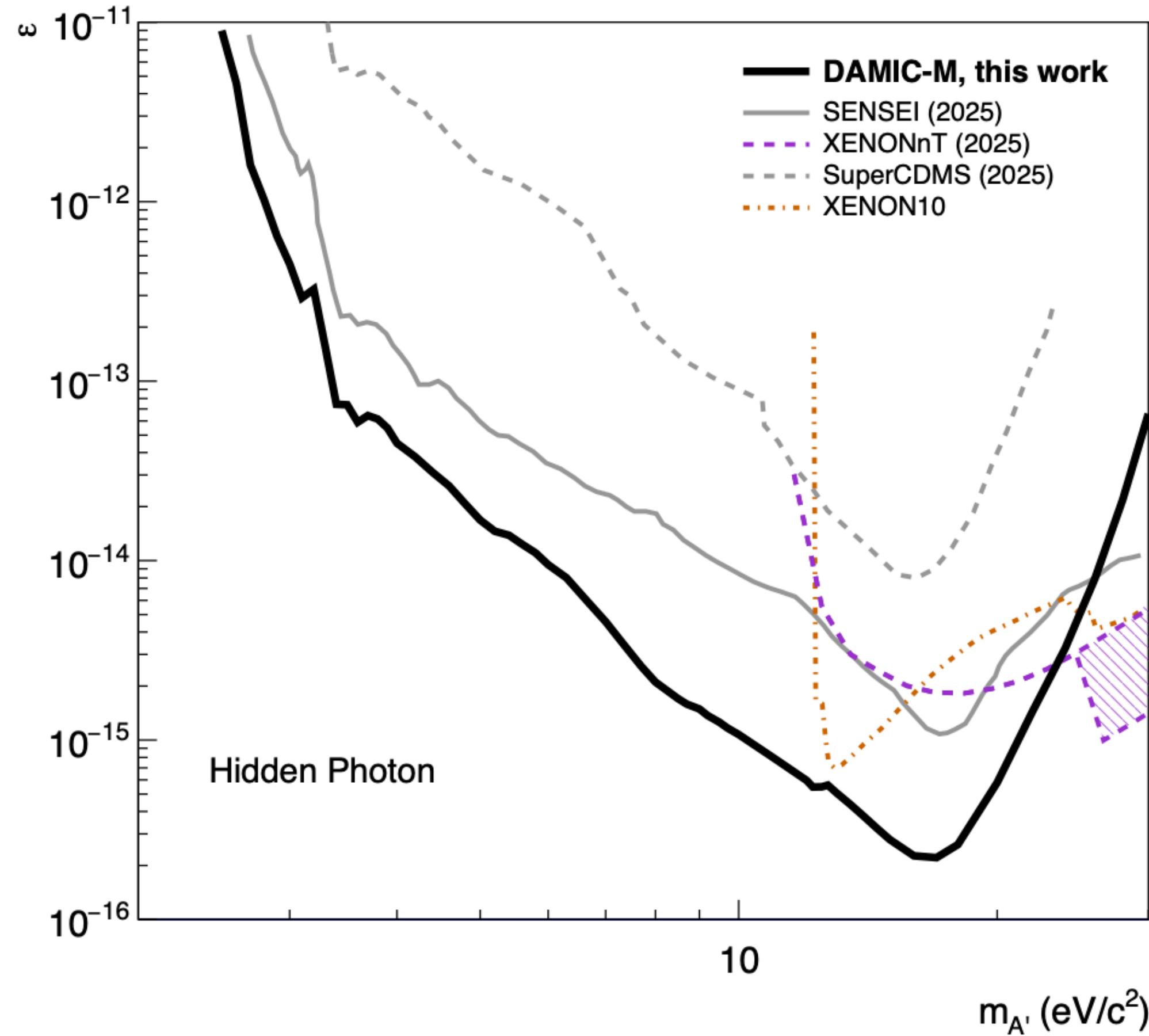
PRL127(2021)081805



PRL118(2017)141803

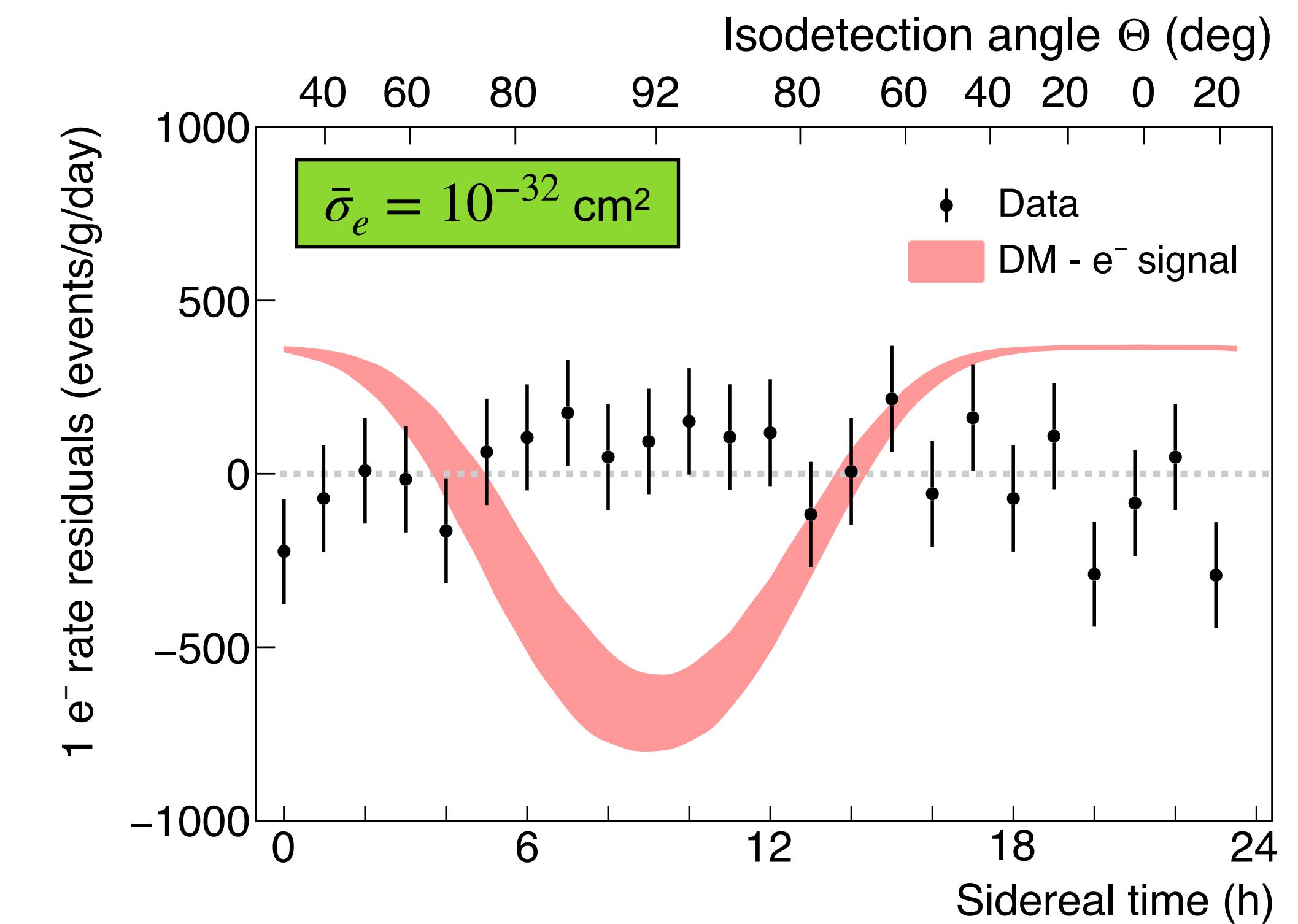
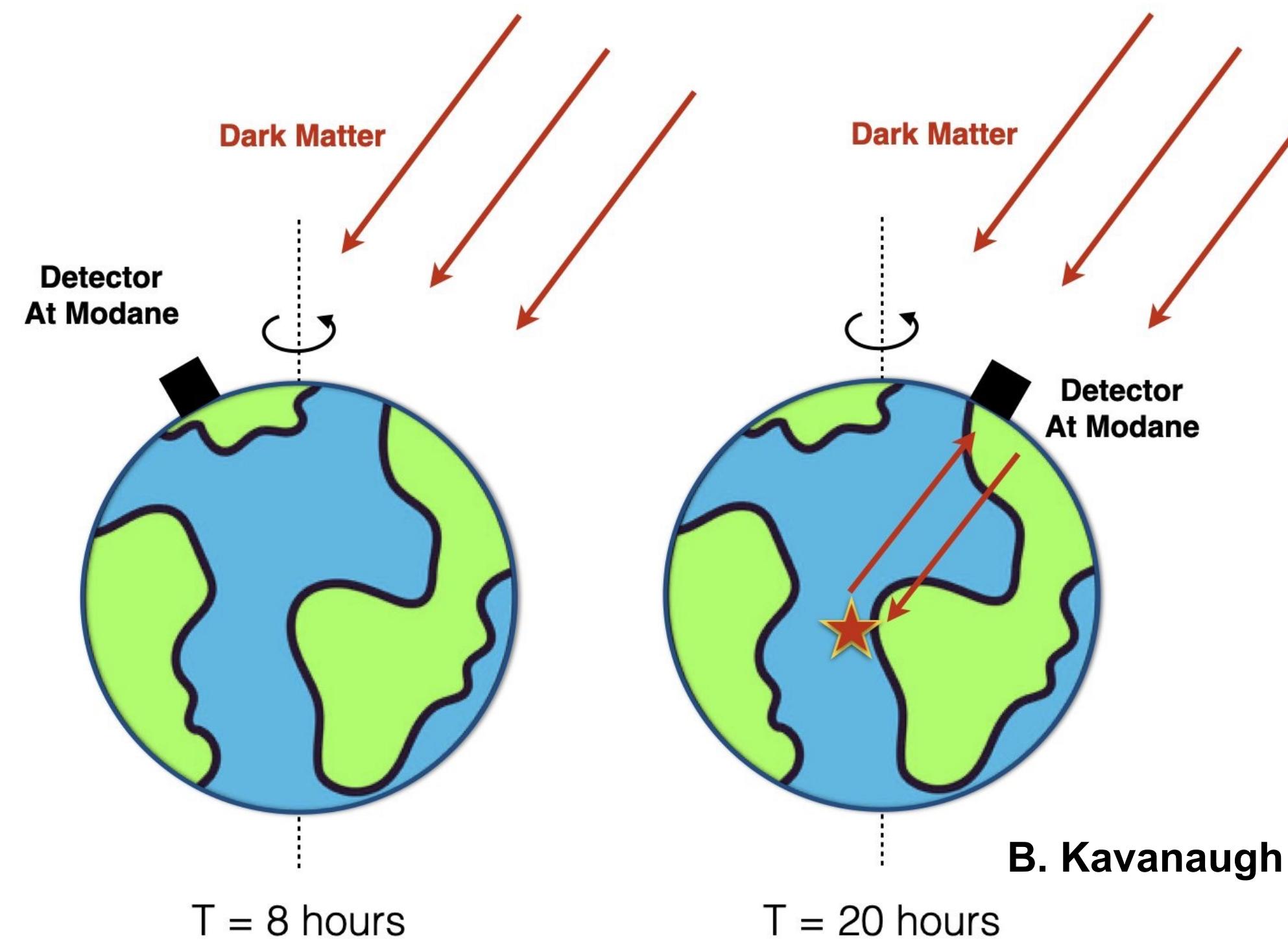


Other exclusion limits



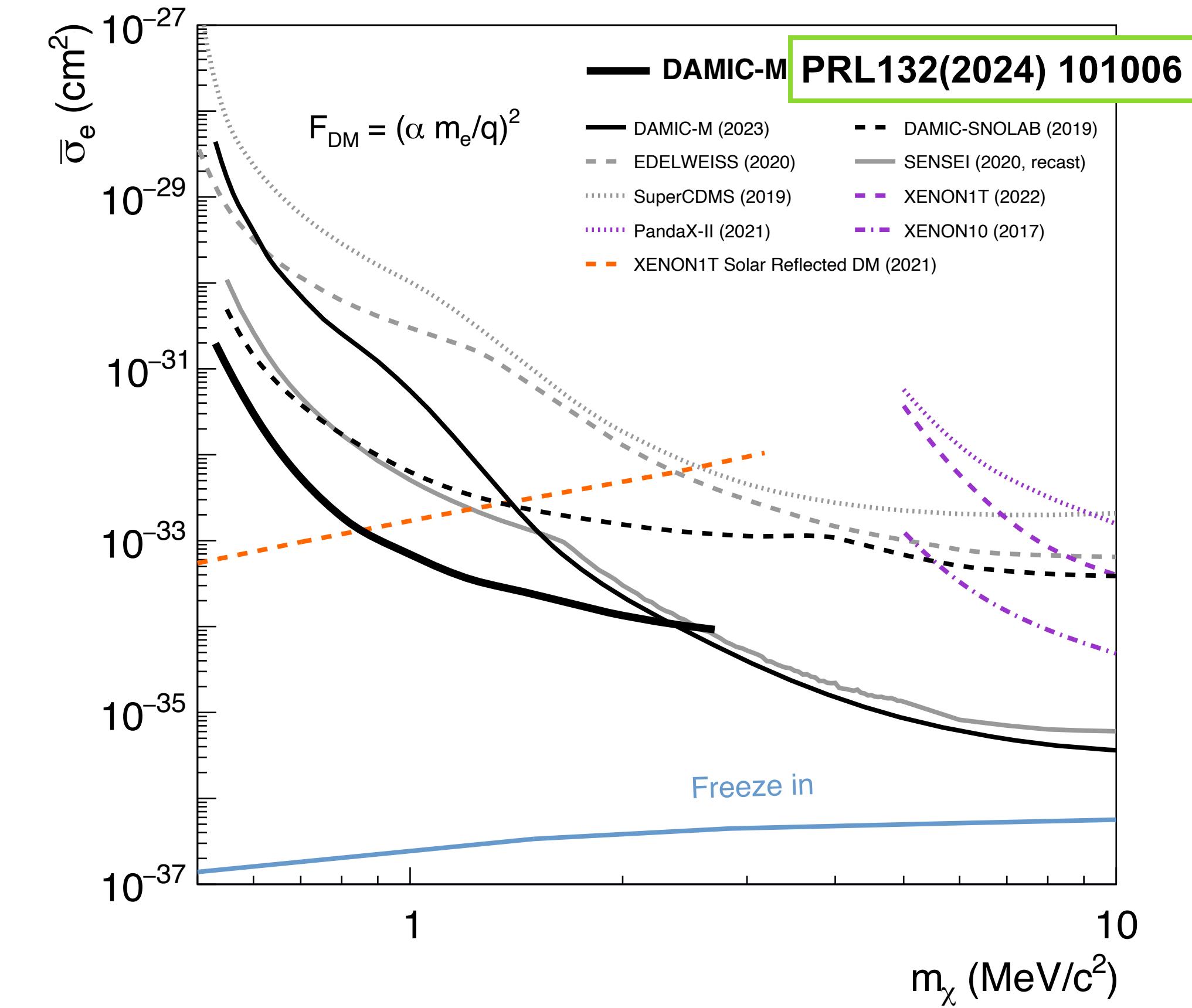
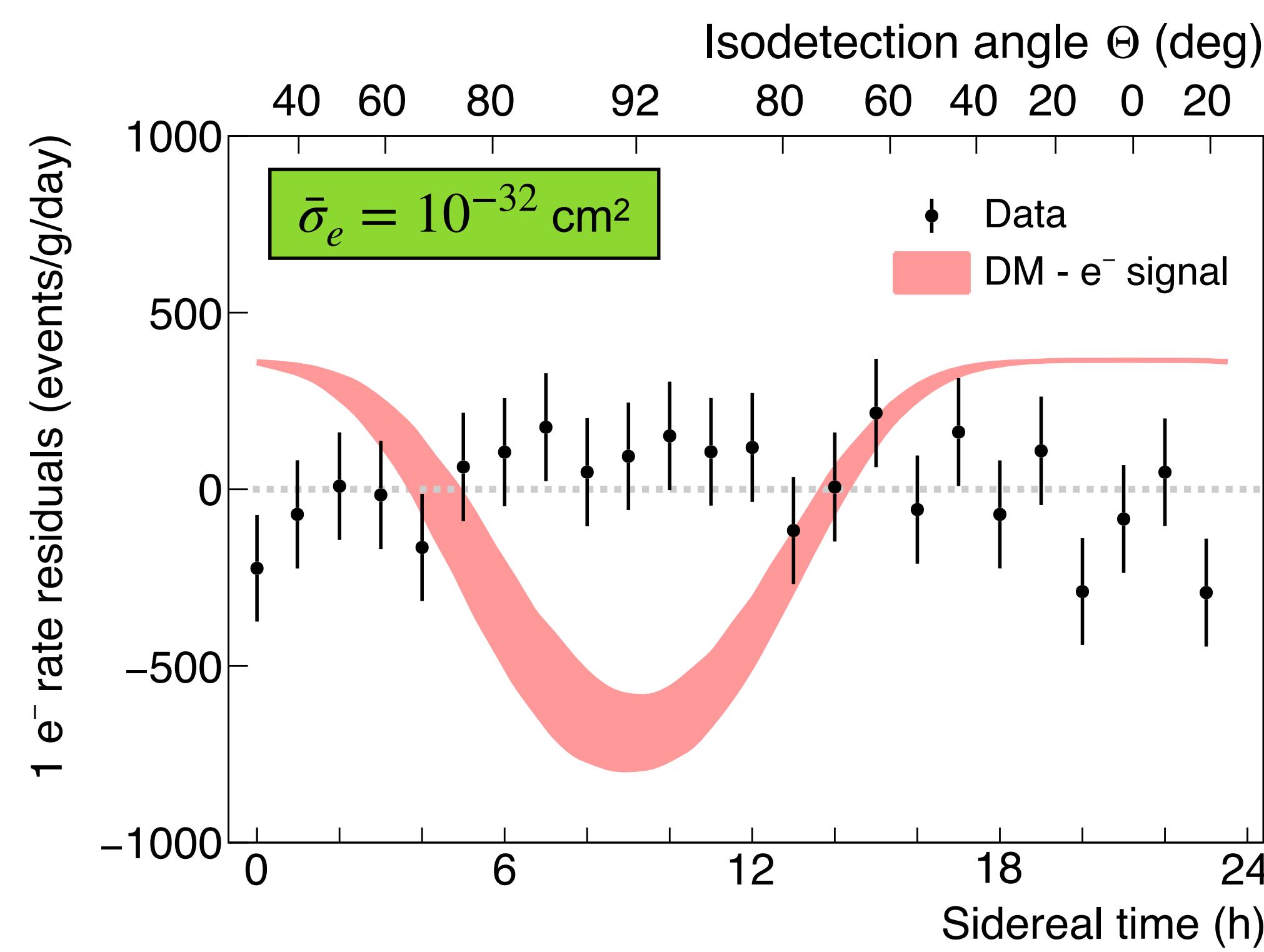
- DAMIC-M also sets stringent constraints on **absorption of relic hidden photon** by atomic electron on silicon and **DM-nuclear scattering via Migdal effect** (see End Matter arXiv:2503.14617). All limits available as text on linked Github.

Daily modulation



- ▶ At large enough σ , DM scatterings in the Earth cause a daily modulation of the flux in the lab.
- ▶ Although “large,” these σ had never been probed for ~ 1 MeV DM.

Daily modulation



- ▶ Previous daily modulation analysis x100 more sensitive than spectral analysis for $m_\chi \sim 1$ MeV.
- ▶ Currently updating daily modulation analysis with latest data set.

Spectroscopy

- Calibrations to confirm CCDs can measure spectral features near threshold.
- Precision measurement with a skipper CCD of electronic recoil spectrum from Compton scattering: [PRD106\(2022\)092001](#)

- “Steps” at the binding energies of the atomic shells in silicon.
- Spectral features not well-reproduced by Geant4.
- Spectrum down to 23 eV_{ee} reproduced with *FEFF* code, which performs full atomic physics treatment.

