nuT: Neutrino Reconstruction with Transformers

Transfer Learning Across KM3NeT/ORCA Telescopes for Neutrino Oscillations Physics

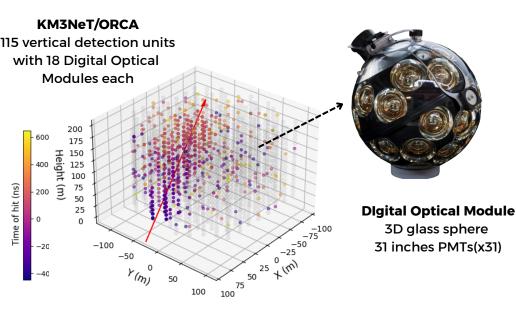
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KM3NeT in search of neutrinos

Deep-sea neutrino telescopes in the Mediterranean sea [1]

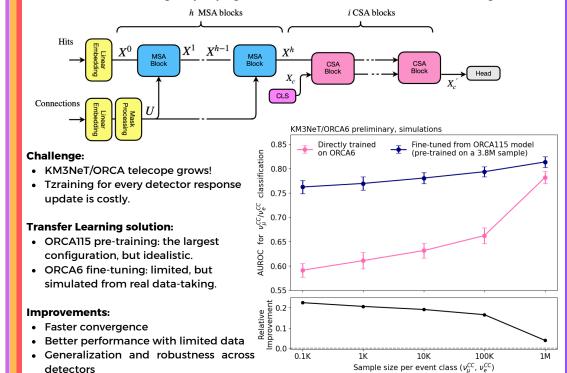
- KM3NeT/ARCA: identify high-energy neutrino sources in the universe
- KM3NeT/ORCA: study of neutrino properties
- Optimized for 1 TeV 10 PeV and 1-100 GeV energies, respectively.



A **muon** from a neutrino interaction near the sea bottom travels through the telescope radiating Cherenkov photons collected by the PMTs with high time and position resolution

One Model, Many detectors

A transformer-based model [2] tailored for neutrino event reconstruction Reconstructs from sparse light patterns detected in the telescope Uses attention masks inspired by physics principles and detector geometry [3, 4] Uses transfer learning to propagate information across detector configurations



Challenges of neutrino physics reconstruction in KM3NeT/ORCA

Reconstruct neutrinos from light

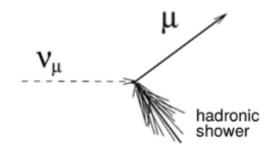
Maximum Likelihood Fit (MLF) algorithms:

- Reconstruct under track or shower hypothesis
- · Do not reconstruct the neutrino itself

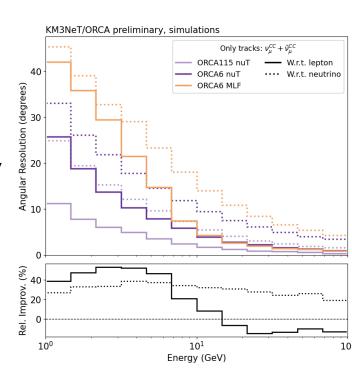
nuT model beyond MLF:

- · Directly reconstructs neutrinos from
- Simultaneously reconstruct all event types

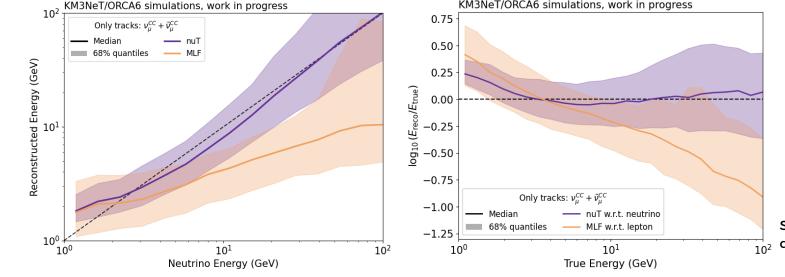
Altough not compared here, an MLF algorithm simultaneously reconstructing the track and shower is in development [5]



10 Only tracks: $v_{ii}^{CC} + \bar{v}_{ii}^{CC}$ Median nuT 68% quantiles _ MLF



KM3NeT/ORCA6 simulations, work in progress



Training large models

KM3NeT generates realistic Monte Carlo samples based on actual data-taking runs, capturing the complexities of deep-sea conditions.

The efficiency of the nuT model and the use of data and computing resources are maximized by leveraging pre-trained models.

Scaling up with precision

KM3NeT/ORCA6, with already promising physics results [6], limits the reconstruction. Nevertheless, with increased detector and statistic, the understanding of neutrino physics improves.

Angular resolution

- MLF: <10° resolution above 10 GeV
- nuT:
 - +30% improvement below 10 GeV
 - Improved resolution w.r.t. neutrino

Energy reconstruction

- MLF: underestimation from missing hadronic component.
- nuT: accounts for both visible and nonvisible energy.

Saturation effect caused by to limited detector size (6 DUs).

Further improvements

Shower events (u_e^{CC}) reconstruction with nuT model Angular resolution: relative improvement of 40% w.r.t. MLF (<10° above 10 GeV)

Energy reconstruction: better accuracy w.r.t. to tracks because of its containment

Background signal suppresion

Atmospheric muons: <1% reconstructed as upgoing (< 8% of contamination) Optical background: effectively removed from quality cuts in reconstruction

References

[1] KM3NeT Collaboration, Letter of intent for KM3NeT 2.0, J. Phys. G 43 (2016) 084001. [2] A. Vaswani et al. Attention is all you need. NIPS (2017). [3] H. Bukhari, D. Chakraborty, P. Eller, T. Ito, M. V. Shugaev and R. Ørsøe, IceCube – Neutrinos in Deep Ice The Top 3 Solutions from the Public Kaggle Competition, (arxiv.org/abs/2310.15674, 2023). [4] Huilin Qu, Congqiao Li and Sitian Qian. Particle Transformer for Jet Tagging, (arxiv.org/abs/2202.03772.2023). [5] B. O Fearraigh. Following the light - Novel event reconstruction techniques for neutrino oscillation analyses in KM3NeT/ORCA, (2024). [6] V. Carretero. Measurement of neutrino oscillation parameters with the first six detection units of KM3NeT/ORCA (arxiv.org/abs/2408.07015, 2024)