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Data-Driven Dark Energy: Probing $w(a)$ with Flexknots

Recent cosmological surveys have opened a new window onto the nature of dark energy. In our work we reconstruct the dark energy equation of state using a “flexknot” parameterisation that represents $w(a)$ as a linear spline with free-moving nodes. By combining the latest DESI Baryonic Acoustic Oscillation measurements with Pantheon+ supernovae data—and cross-checking our results with an independent Cobaya-based pipeline—we obtain posterior distributions for $w(a)$ that reveal an unexpected W-shaped structure. Although the Bayesian evidence does not ultimately favour dynamical dark energy over the standard Λ CDM model, our analysis shows that even non-CMB datasets can indicate deviations from a constant $w = -1$.

We also generalise dataset tension statistics to marginalise over multiple models, ensuring our unexpected results are not driven by inter-dataset disagreement.

We finish with a brief discussion of the pedagogical advantages and computational benefits of developing an analysis pipeline in-house, which, in addition to increasing efficiency, allows us to analytically marginalise nuisance parameters and provides confidence that these features are genuinely driven by the data.

AI keywords

Bayesian inference; dataset tension; model selection

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