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# Analysis of low energy Nuclear Recoils' from AmBe neutron source

## Data taking setup





12

 $\overline{2}$ 

#### AmBe excess selection







### AmBe excess selection (zoom)



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#### AmBe excess selection - Some samples



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5

## Selected clusters Energy/Density spectra





## Directionality evaluation

- Principal Component Analysis (PCA) with 2 parameters on the most intense part of the clusters to extract the clusters' axes.
- Use always the **biggest eigenvector** to compute the angle with respect to the  $\hat{x}$ direction.
- Impose the head-tail, since we know this excess comes from the AmBe source.
- Do the same on the Background dataset and compare to see if there are differences.





#### Directionality evaluation - Examples



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#### Directionality evaluation - Examples



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### Map of AmBe Nuclear Recoils



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Angle (degrees)

 $\Omega$ 

**Flat Distribution Uncertainty** 

 $\frac{1}{1}$ 

50

100

150

AmBe

 $-50$ 

- Flat Distribution

AmBe Uncertainty

## Directionality evaluation - AmBe vs. Bkg



#### • Observations:

- Excess of vertical clusters in Bkg sample. Compatible with flat distribution.
- Excess of horizontal clusters in AmBe sample. Not compatible with flat distribution.
- Is this expected?



20

 $-150$ 

 $-100$ 

30

25

20

 $rac{15}{2}$  15

## Monte Carlo validation

Strategy:

- · Simulate a fake nuclear recoil inside the detector frame.
- Model the interaction as a simple elastic scattering.
- Project the angle on the GEM plane and compare with the observed distribution.

#### **LAB** frame  $\mathbf U$  $\theta_{U}$  $\mathbf v$ m  $\mathbf b$ M  $\theta_{W}$  $\mathbf R$ W  $S1n \gamma$  $\theta_W$  $=$  arctan  $\cos \nu$









#### • Vertical region is not perfectly matched by the AmBe sample, but **Bkg is for sure flatter.**

## MC validation - Simulated angle



# $E_i(\sigma) = E_i + error$

#### where  $error \in Gauss(0, \sigma)$

- The differences in the distributions could be due to our angular resolution, which is absent in the simulation.
- We can simulate it by means of a gaussian smearing.
- In order to statistically compare the distributions, we can use the reduced  $\chi^2$ , indicating the measurements with  $O_i$  and the simulation with  $E_i(\sigma)$  .

## MC validation - Gaussian Smearing



$$
\chi^2 = \sum_{i} \frac{\left[O_i - E_i(\sigma)\right]^2}{E_i(\sigma) \times \nu} \qquad \nu = \text{\# of bins}
$$



#### Angular resolution = 5°:



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#### Angular resolution = 25°:



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num\_events=384

17

#### Angular resolution = 40°:



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num\_events=384



#### Angular resolution = 55°:



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num\_events=384



#### Angular resolution = 85°:



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MC validation -  $\chi^2$  vs. Resolution

- Two claims can be extracted from  $10<sup>2</sup>$ these tests:
	- Our measurement resolution with this method is ground 40-45°.
	- There is a preferential direction in the AmBe dataset.

 $10^0$ 

 $\chi^2$ 





#### We can reconstruct the source position



• The AmBe source was placed at half height (Y ~ 1150pixels)



First evidence of the directionality of LIME for Nuclear Recoils



MC validation -  $\chi^2$  vs. Resolution

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 $10^0$ 

 $\chi^2$ 



## Clusters 3D range reconstruction

- . 3D range  $= \sqrt{\text{sc\_length}^2 + L_z^2}$  , with  $L_z = v_{\text{drift}} \times \text{Tor}^{\text{max}}$
- 55Fe source half-way in the drift direction.

• Both lengths should be preprocessed removing the diffusion, evaluated from data taken with







## Clusters 3D range reconstruction

- A 5.9 keV  $e^-$  travel ~0.5 mm in He:CF<sub>4</sub>.
- From the previous slide we obtain:

#### $8.63 \pm 0.9$  mm

- spot size mainly due to diffusion only.
- This measurements can be interpreted as offsets to be subtracted to their relative physical quantities.



- Since the effect of the diffusion increases with the distance, the length offset does it too;
- Diffusion of ionisation electrons scales with the square root of the distance in drift chambers.
- Transverse profile *σ* gives α measure of the position of small clusters in the drift direction.





- Fit Energy vs Range simulation with a 2nd order polynomial function.
- With this we can extrapolate energies outside the simulated range domain and compute the "expected energy".





3D range distribution for AmBe NRs sample

- Combining camera and PMTs we can obtain 3D range for each cluster.
- Most of the clusters are shorter than 10 mm.



Expected energy spectrum for AmBe NRs sample

## True energy spectrum from 3D range

- Known non-linearity response for very dense tracks.
- Using previous range vs energy simulations, the true energy spectrum is extracted.
- Maximum bin for NR with reconstructed energies between 200 and 300 keV



Energy saturation factor distribution







#### Conclusions



First RUN with AmBe lasted unfortunately less than 48 hours;

With a very simple selection, 1461 NR were identified, to be compared with 71 in a same data-taking without source;

From an evaluation based on their length, their energy was reconstructed to be mainly below 1 MeV;

The distribution of their angles reconstructed with a PCA performed on the saved clusters is:

- different for the AmBe and bkg neutron, indicating a clear sensitivity to the NR preferred direction

- compatible with a direction resolution of about 40-45O