



EROMA

TRE

SAPIENZA NIVERSITÀ DI ROMA

PMT Analysis for <u>3D tracking</u>



David J. G. Marques
<u>CYGNO</u> collaboration





The University Of Sheffield.

OIMBRA

PMT Reco & Analysis - Previously





Full framework retrieves and saves all the relevant information automatically.

Previous episodes:

- 1.[Reconstruction and Analysis meeting (16 maggio 2024) · Agenda (Indico)]Initial look at alpha tracks for directional & head-tail determination
- 2. [Technical Reconstruction meeting (5 June 2024)]

<u>1 Update on ... - 3D reconstructed alpha tracks</u>

3. [Reconstruction and Analysis meeting]

2 Update on ... - 3D reconstructed alpha tracks

4. [Technical Reconstruction meeting]

3 Update on ... Techical_update_Coordinates_&_Shadows

5. [<u>Two Days of Reconstruction and Analysis</u>]

4 Update ... - 3D reconstructed alpha tracks - Real Analysis

- <u>[Reconstruction and Analysis meeting]</u>
 <u>Update ... 3D alphas analysis</u>
- 7. This one [CYGNO Collaboration Meeting 2024]

<u> 3D reconstruction - David</u>

\star 🐘 My thesis 🏇

LIME - Detector and Data





- → 50 L & 50 cm drift gaseous TPC
- → He:CF4, 60:40, 1 Atm (910 mbar), 293 K
- → 4 PMTs + 1 sCMOS camera

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- → <u>One event =</u>
 - 1 CMOS pic: R_{AF} = 300 ms
 - **X PMT WFs** = N_{triggers} * N_{PMTs} * N_{digitiz}
 - R_{\Deltat} = **1.3 ns & 4 ns**



1000

1500

2000

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 - X PMT WFs = N_{triggers} * N_{PMTs} * N_{digitiz}
 - R_{\Delta t} = **1.3 ns & 4 ns**
- → <u>The information needs to be matched!</u>





3D Event Reconstruction



- CZGNO G S
- → To fully reconstruct the information of one event we need to merge the CMOS and PMTs information.
 - We developed a <u>1-to-1 association</u> to merge the *CMOS clusters* to *PMT triggers*.
 - 1. Light seen by each PMT depends

on their relative positions.



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 - We developed a <u>1-to-1 association</u> to merge the *CMOS clusters* to *PMT triggers*.
- Light seen by each PMT depends on their relative positions.
- We apply a multi-variable Bayesian fit in a time window by integrating
 <u>the charge</u> ⇒ We retrieve (x, y, L) information



GS

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 The final goal is to implement this technique at the detector front-end level. The <u>efficiencies</u> are promising! (The reference is the 33x33cm² GEM plane)



→ **Optimization undergoing** concerning effects like

saturation, lens barreling, gain inter-calibration, etc.

GS



3D Event Reconstruction





3D Events - CMOS analysis

- The analysis of the CMOS images starts with a directional iDBSCAN algorithm which clusters groups of pixels belonging to the same ionization event.
 - For <u>PID</u>, each <u>cluster</u> can be selected through its: *light integral, length, slimness, photon density, dE/dx, etc.*





Directional iDBSCAN to detect cosmic-ray tracks for the CYGNO experiment - IOPscience

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GS

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GS

3D Event















CZGNO G S Experiment S I

3D Events - PMT analysis

→ The <u>PMTs</u> gives us information regarding the <u>longitudinal coordinate Z</u>, and allows us to close the <u>3D geometry</u>



CZGNO G S

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CZGNO G S

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Experiment S I

3D Events - PMT analysis

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- 1. Time-over-Threshold ⇒ Traveled Z
- **2.** Highest integral \Rightarrow XY Quadrant
- 3. Bragg peak position \Rightarrow Moving towards GEMs or cathode



Experiment S

3D Events - PMT analysis

- → The **PMTs** gives us information regarding the **longitudinal coordinate Z**, and allows us to close the **3D geometry**
- **1.** Time-over-Threshold \Rightarrow Traveled Z
- **2.** Highest integral \Rightarrow XY Quadrant
- 3. Bragg peak position ⇒ Moving towards GEMs or cathode

*Additionaly, the alpha PID is performed at this stage, using these and others variables. Non-alphas are discarded. Optimized?



Experiment S I

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Additional reconstructed info:

- ΔΖ
- Sign of θ



CZGNO G S

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3D Event Reconstruction





3D Events - Absolute Z

- → When studying the distribution of the supposedly Rn alphas, two variables are important:
 - Absolute Z position + Angular distribution



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CZGNO G S Experiment S I

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the transverse (Y) light distribution of each segment of the track: the σ_Y obtained from the Gaussian fit is in fact increasing with \sqrt{Z} . Therefore, the segment's original Z can be deduced by measuring σ (Fig. 8). The observed values of σ are related to the track's Z coordinate by $\sigma = \sqrt{\sigma_0^2 + B^2 Z}$ The transverse diffusion coefficient B in the gas is measured to be $129.7 \pm 3.1 \frac{\mu m}{m}$, well in agreement with the expected value obtained with Garfield simulation (130 $\frac{\mu m}{\sqrt{m}}$) [47] The intercept at zero $\sigma_0 = 292 \pm 12 \ \mu m$ is due to the contribution of the electron avalanche propagation in the GEM stack and confirms results obtained with a radioactive source [47]. https://www.sciencedirect.com/science/article/pii/S0168900221001935 of the diffusion parameters. Now we can use these parameters to predict absolute *z* for any segment of ionization charge detected: $z = \frac{\sigma^2 - A^2}{2}$ (3)The absolute z accuracy is measured by the horizontal standard deviation of the distribution from the fit curve in Fig. 4. For all segments, without any selections on fit quality besides requiring that the fit not fail (efficiency 0.994), the absolute z measurement accuracy is 0.5-1.5 cm throughout the full z range. For nuclear https://www.sciencedirect.com/science/article/pii/S0168900215003320

for production points farther in Z from the GEWIS. This is reflected in

David Marques

(a)

(c)

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orthogonal direction \Rightarrow <u>"Transverse profile"</u>

Mean x Mean y Std Dev x

Transversal Profi



The transverse diffusion coefficient *B* in the gas is measured to be $\frac{129.7 \pm 3.1}{\sqrt{cm}}$ well in agreement with the expected value obtained with Garfield simulation (130 $\frac{\mu m}{\sqrt{cm}}$) [47].

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$$z = \frac{\sigma^2 - A^2}{B^2}.$$

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- ➔ I used LEMOn values. Clear offset since LIME is 50 cm.
- → I also tried Roque's result ⇒ Huge uncertainty, and similar result
- → Usable to hint absolute Z position









- 6. Improvement of 3D Reco
 - a. Sensor-cut alphas removal ⇒ <u>Cleans low energy tail</u>
 - b. This can affect the effectiveness of the directionality code, thus decreasing our **head-tail capabilities**.



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Noisy band cut on the reconstruction code



• Accounts for 15% of events

C/GNO G S Experiment S I

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3D Event Reconstruction

Dual sensor 3D analysis

CZGNO G S

3D Events - 3D analysis

- → With the sensors' individual and merged information, we can now fully analyze the events and perform <u>particle</u>
 <u>ID</u>, <u>reject backgrounds</u> from known sources, and fully <u>characterize the 3D direction</u> ⇒ <u>Directionality</u>
 - The first studies were focused on <u>Alpha Particles</u>





Alpha studies



What I have been showing...

3D Events - The Rn study

→ When comparing our data vs. simulation, we found discrepancies, *possibility attributed to*²²²*Rn*.



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Alpha studies



Alphas - Alpha range in LIME



- But where do these ranges come from? Flaminia's thesis? Geant4? SRIM?
- We have seen previously that the **alpha length** could be slightly overestimated due to the *"minimum signal possible"*. "Consistent" with the fact Geant4 shows smaller lengths...

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- We have seen previously that the **alpha length** could be slightly overestimated due to the *"minimum signal possible"*. "Consistent" with the fact Geant4 shows smaller lengths...
 - I tried to understand the origin of this... I checked and compared:
 - Davide Pinci's values ⇒ Compatible with data, but no source (?)
 - Flaminia's thesis ⇒ No final values in the thesis ⇒ "range tables by SRIM for our gas mixture"
 - **CYGNO Simulation** ⇒ Incompatible with data, consistent between simulators (Fiorina & Melba)
 - Theoretical calculation using NIST calculators ⇒ Approximated for He:CF4 starting from air ⇒ Inconsistent with everything.
 - SRIM ⇒ Supposedly the source of evertyhing
 - My 3D analysis (= data)



- I tried to make the new alpha digitization work to see if the difference in MC and data is recovered there.
 - (Thanks Melba for the alphas, and Giorgio & Piacentini for digi details)



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Outcomes...

- Lengths for the 3 Rn alphas are, on average:
 - 61.6 ; 69.4; 97.9 mm
- They are bigger than simulation 🔽
- But also much bigger than data* 🗙

*Actually Pietro found encouraging results on these!







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- → Profile of digitized track not really Gaussian
 - Neither data, but hint for improvement!
 - Saturation UNDERSTANDING ?!
- ★ Digi not there to help me, but almost
 - Empirical explanation incoming





<u>SRIM</u>

• SRIM needs two important inputs: stoichiometry of gas mixture and *gas density. What's the gas density of LIME?*

Alphas - SRIM

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 - Theoretical: **d** = **1.4 kg/m**³
 - Ranges calculated with this density, consistent with Geant4
 - Now we know SRIM works.

Ic Ene	ergy E	HE/dx Elec.	dE/dx Nuclear	Proj Ra	ected
5.50 Me	V 6.790E	-01 4.29	9E-04 39).51 m	ım
6.00 Me	V 6.382E	-01 3.98	3E-04 44	1.93 m	ım
6.50 Me	V 6.023E	-01 3.71	2E-04 50	0.68 m	ım
7.00 Me	V 5.707E	-01 3.47	8E-04 56	5.77 m	ım
7.69 Me	V 5.337E	-01 3.20	1E-04 65	5.69 m	ım



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 - density = [1.46 1.50 kg/m3]

	Ion Energy	dE/dx Elec.	dE/dx Nuclear	Projected Range
5.	50 MeV	6.790E-01	4.299E-04 3	39.51 mm
6.	00 MeV	6.382E-01	3.983E-04 4	14.93 mm
6.	50 MeV	6.023E-01	3.712E-04 5	50.68 mm
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CYGNO Collaboration Meeting - Nov 27 - 29, 2024

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Energy	Range (avg) ± Error (1/2*max diff)	
5.49	37.37 ± 0.51	
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*claminia wrote in	Energy	Range (avg) ± Error (1/2*max diff)	
her thesis the CYGNO	5.49	37.37 ± 0.51	
gas mixture density is	6.0	42.50 ± 0.58	
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Let's re-start the discussion...

With correct energies & ranges

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- → There's a missing component not being account for: the *minimum XYZ signal*
- \rightarrow I calculated the transverse and longitudinal diffusion from ⁵⁵Fe signals, and add the average to the expected ranges.
- → Since I don't know the position of the event in Z à priori, I will take the *mean value (step 3)* to correct the 3D lengths.



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 - Minimum 3D length = 8.39 mm ± half of the maximum variation = 1.04mm



- Final result for the Minimum XY / Z: XY: 5.7mm Z: 3.95 mm
- Longitudinal diffusion smaller than transverse 🔽



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Table 1.7: Summary of the expected ranges for the energies of the 3 major ²²²Rn alphas along with the 3 major peaks visible in the distribution of measured 3D lengths of alphas in LIME.

Energy [MeV]	Range in LIME [mm]	Range + 3D min length [mm]	Measured 3D length [mm]
5.49	37.37 ± 1.01	$45.76 \pm 0.51 \pm 1.04$	$44.3 \pm 1.8 (1\sigma)$
6.00	42.50 ± 1.15	$50.89 \pm 0.58 \pm 1.04$	$51.2 \pm 1.8 (1\sigma)$
7.69	62.14 ± 1.68	$70.53 \pm 0.84 \pm 1.04$	$72.9 \pm 2.6 (1\sigma)$

Now the results are consistent!

- Final result for the Minimum XY / Z: XY: 5.7mm Z: 3.95 mm
- Longitudinal diffusion smaller than transverse 🔽

No shady business!

But is it really *Radon*?

David Marques



- → If I remove the contribution from the field rings...
 - Clear selection of events in the center (likely Radon)



(a) Geometrical cut - central square



- → If I remove the contribution from the field rings...
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(b) 3D length of alphas

CZGNO G S Experiment S I

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CZGNO G S Experiment S I

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- → Alphas emitted at higher Z at each step of the chain because the Rn daughters are emitted charged (mostly)
 - Thus why 7.69 MeV alpha is emitted mostly towards the GEMS (from the cathode). Seen also by DRIFT, MIMAC, and other
- → Alphas at ~4 cm length (²²²Rn) see a preferential emission towards the cathode instead, although it's **not clear** why.
- → Remaining alphas at lower (< 4 cm) length are thought to come from the U/Th chains...





loole #	Measured 3D	Estimated real range =	Estimated energy
eak #	length (mm)	3D length - 3D min length (mm)	from range (MeV)
1	$33.68 \pm 4.56 (1\sigma)$	$25.29 \pm 4.56 \pm 1.04$	$4.170\substack{+0.645\\-0.725}$
4	$60.98 \pm 3.41 \ (1\sigma)$	$52.59 \pm 3.41 \pm 1.04$	$6.905\substack{+0.375\\-0.385}$





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			*	Lev	els	
Energy (keV)	Intensity (%)	Туре	Origin	Start*	End*	Parent
4 038 (5)	0.13 (3)	α	U-238	0	2	U-238
4 151 (5)	22.3 (5)	α	U-238	0	1	U-238
4 198 (3)	77.5 (5)	α	U-238	0	0	U-238
5 748.46 (11)	0.118 (15)	α	Rn-220	0	1	Rn-220
6 288.22 (10)	99.882 (15)	α	Rn-220	0	0	Rn-220
5 988.4 (7)	0.0019 (3)	α	Po-216	0	1	Po-216
6 778.4 (5)	99.9981 (3)	α	Po-216	0	0	Po-216





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*The tails could be U						
alphas that lose part						
of energy in material.						
*We also kind of see						
Po210 shoulder.						

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		Туре	Origin	Start*	End*	Parent
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CXGNO G S Experiment S I

→ If we look at the other peaks, at ~3.4 cm and ~6.1 cm, we can estimate their energies



→ And who is emitting all this Uranium?

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→ And who is emitting all this Uranium? G S

Entries

→ If we look at the other peaks, at ~3.4 cm and ~6.1 cm, we can estimate their energies

12450



- → Mostly at the borders
 - Not surprising!



(b)

30

35

Begin X [cm]

GS

But what about the energy? Measurable?



→ First we need to select alphas with the same saturation:
S G

First we need to select alphas with the same saturation: ->

Cuts on: absolute Z; Z angle (!); central region (only Rn alphas)

2398

278 / 124

 120.4 ± 5.7

 4.62 ± 0.01

 171.2 ± 6.8

 5.273 ± 0.004 0.1011 ± 0.0025

> 21.59 ± 1.38 7.504 ± 0.012

 0.164 ± 0.008

3D length [cm]

 0.1398 ± 0.0050

Entries

 γ^2 / ndf

0a

p1

n2

p7

6 7 8 9



(b) Alpha particle - Reconstructed event.



→ First we need to select alphas with the same saturation:

Cuts on: absolute Z; Z angle (!); central region (only Rn alphas)



→ Peaks clearly suggested a priori...

C/GNO G S

→ First we need to select alphas with the same saturation:

Cuts on: absolute Z; Z angle (!); central region (only Rn alphas)



→ But if I look at the singular contributions...



→ First we need to select alphas with the same saturation:

Cuts on: absolute Z;

Z; Z angle (!);

e (!); central region (only Rn alphas)



Yesterday

analysis...

Alphas - The profile shape



→ The alphas are not really Gaussian... they seem *more triangular* (!?) ... also different from in NID...



- → Gaussian fit not really appropriate
- → Alphas strongly saturated
 - The study of the profile could help the study or understanding / modelling of the saturation with the help of digitization
- → Samuele suggested me to use the Transverse Profile distribution (histo) RMS

Samuele suggested me to use the alpha profile RMS isntead of the (poorly calculated) sigma...



→ Samuele suggested me to use the alpha profile RMS isntead of the (poorly calculated) sigma...



Profile RMS

David Marques

G S

→ Samuele suggested me to use the alpha profile RMS isntead of the (poorly calculated) sigma...



→ Samuele suggested me to use the alpha profile RMS isntead of the (poorly calculated) sigma...



- Gained much resolution!
 Rn alpha confirm!
- Rn progeny drift
 towards cathode even
 more clear!
- Is RMS (from PROFILE) a good measure of diffusion and subsequently absolute Z?
- Maybe LIME is really 80 cm long (did you measure it?)

GS

Conclusions - What has been done



- → The merging of CMOS (X-Y) with PMT (Z) information is performed with a Bayesian fit.
 - <u>Reconstructing</u> ionization events in <u>3D</u> greatly improves <u>spatial resolution</u> and our <u>PID capabilities</u>.
- → From the suspicion of the presence of Rn, the <u>3D analysis</u> allows us to almost confirm its presence and, through its emission direction, its origin became more clear.
 - Other decay chains identified with this method
 - 3D analysis is more efficient than energy analysis (better resolution)
 - Shape of the alpha cloud profile interesting \Rightarrow Relation to saturation and absolute Z
- → Currently **paper** on **CMOS+PMT association**, **3D reco** and **Rn lengths** undergoing.

Conclusions - What should be done



- Studies on the application of the 3D analysis for NR and others. \rightarrow
 - The framework is there \Rightarrow Just need to PID different particles and analyze them
- Improvements in the code \rightarrow
 - I'm lacking optimization of parameters, and there are a lot!
 - **PMT simulation is paramount!**
- Implement the new version of BAT \rightarrow
 - And probably implement it at DAQ level 0
- Alphas are cool, but what now?
 - Additional following paper on 3D results with discussion of the origin of the background? With possible comparison with MC?
 - We can still work together after I'm gone! 🧡 🏴 •



Thank you for

your attention!

The CYGNO Project counts with the collaboration of several international researchers, coming from:



Backups & more details

Alphas - Improvement of 3D reco

- C/GNO G S Experiment S I
- The CMOS sensor active row cut is checked using Piacentini's formula. I use a 2 mm window for small variations.

```
bool cutted frame = false;
double activeRow = 0, active row cm = 0;
double active range low = std::min(c begin Y, c end Y), active range high = std::max(c begin Y, c end Y);
double window check = 0.2;
if (ttt time > 184.4 && ttt time < 300) {
   active row cm = 36.; // inside Global exposure
   std::cout << "--> This track is *NOT* cut by the sensor." << std::endl;</pre>
} else {
    std::cout << "--> This track *COULD* be cut, let's check..." << std::endl;</pre>
    if
            (ttt time < 184.4) activeRow = 2304.0 - (2304.0 * (ttt time / 184.4));
    else if (ttt time > 300.0) activeRow = 2304.0 - (2304.0 * ((ttt time - 300.0) / 184.4));
    active row cm = activeRow * granu;
    if (active row cm > (active range low - window check) & active row cm < (active range high + window check)) {
        std::cout << "--> This track *IS* cut by the sensor!" << std::endl;</pre>
        cutted frame = true;
```

Alphas - Improvement of 3D reco



- A cross-check is now performed in order to identify alphas where this happens.
- The noisy cut band it's easy to check, and I check if the track starts or end within 1 mm from the cuts:

```
// To ensure we consider the correct limits independently of the track direction
double active_range_low = std::min(c_begin_Y, c_end_Y);
double active range high = std::max(c begin Y, c end Y);
```

// 1 mm window to mitigate mismatch between pixel rows and cm position due to granularity
double window = 0.1; //millimiter

double noisy_Y_band_high = (2304. - 304.) * granu - window; double noisy_Y_band_low = (0. + 250.) * granu + window;

if (active_range_high > noisy_Y_band_high || active_range_low < noisy_Y_band_low) {</pre>

```
std::cout << "--> This track *IS* cut by the noisy band cut in the reco." << std::endl;
cutted track = true;
```

Alphas - Trans/Long Diffusion



- I took about 20 calibration runs * 5 steps from batch6 of the alpha dataset.
- I measure the total time and space length of the ⁵⁵Fe signal seen by the PMT and CMOS, respectively.
- For the PMT ⇒ PMT_Avg_Events-> Draw("pmt_peak_FullWidth>>hist_tw", "pmt_wf_sampling == 1024 && pmt_peak_Number == 1", "goff");
- For the CMOS ⇒ Events->Draw("sc_length>>hist_sc_length," "sc_length < 200 && sc_rms > 6 && sc_tgausssigma > 0.5/0.152 && sc_tgausssigma/sc_lgausssigma > 0.7 &&
 sc_width/sc_length > 0.7 && sqrt((sc_xmean-2304/2)**2 + (sc_ymean-2304/2)**2) < \$00"yoff");
 </p>
- I perform a Gaussian fit for each **step**
- I convert the mean to cm:

d. $\mu_{new}[cm] = \mu[ADC]$ * time_to_cm = μ * (4./3.) * 5.471E-3 **e.** $\mu_{new}[cm] = \mu[pixels]$ * granularity = μ * 0.0155

- I plot the mean of the Gaussian (μ_{new}) as a function of the step position in cm
 - I fit these points with sqrt([0]*[0] + [1]*[1]*x)







 ${\rm Minimum}~{\rm 3D}~{\rm length}~=~$



v2 / ndl

Constant

Fime window fit 10801

1162.72/52

1166 41 +/- 15 38

49.06 +/- 0.05

4 10 +/- 0.03

Pixels [#]

Cluster length: step 5

(1.12)



Alphas - LIME Gas density

s (He: CFy) T= 21°C Atte = 91000 . 0,004 = 5,000 8,314 . 29.34 kg/m3 MAC = 49/wol 11 cfy = 88 g/wol PU= ART (=) D= PM RT Acfy = 91000 · 0,088 8,314 · 2934 R= 8,314 J/mol 7:1 = 0.6 × AHL + 0.4 . Act 1 mbor = 100 Pa SHe, 14 T= 20°C = 293 k 60:40 SINMOON 1,40 Kg/m =