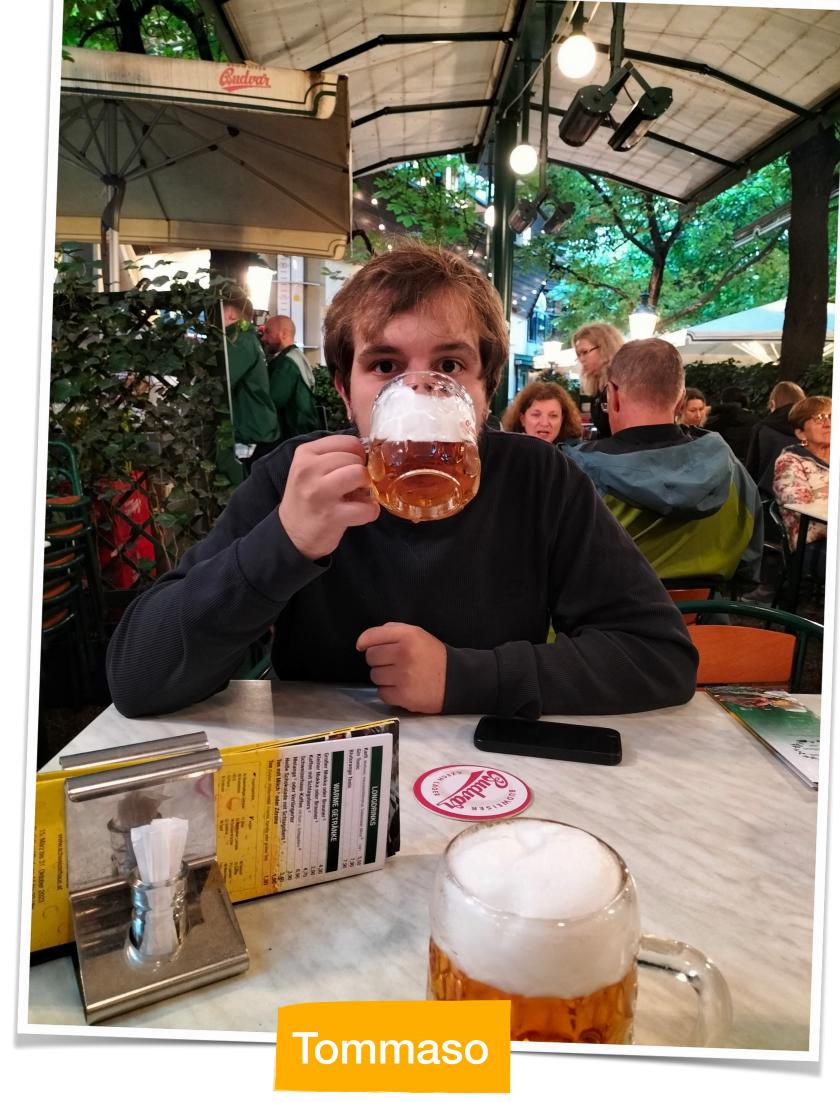
RICCI FLOWS, MASSIVE GRAVITY AND TT

TOMMASO MORONE









2017 – 2023 BSc & MSc in physics

2023 – today PhD in physics

Tommaso



2017 – 2023 BSc & MSc in physics







Statistical Field Theory group – Torino

Integrable models

Andrea Cavaglià Roberto Tateo

Nicolò Brizio
Tommaso Morone
Michelangelo Preti
Nicolò Primi

Lattice field theory

Michele Caselle Marco Panero

Andrea Bulgarelli
Tommaso Canneti
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Alessandro Mariani
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A non-exhausting list of cool things born in Turin

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TT-deformed 2D Quantum Field Theories

Andrea Cavaglià 1 , Stefano Negro 2 , István M. Szécsényi 3 , Roberto Tateo 1

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integrable deformations of QFTs



Introduced as an irrelevant deformation of 2D QFTs

[Cavaglià, Negro, Szécsényi, Tateo 2016] [Smirnov, Zamolodchikov 2016]

$$\partial_{\lambda} S_{\lambda} = \frac{1}{2} \int d^2 x \sqrt{g} \left(T_{ab} T^{ab} - T_a^a \right)$$

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- Well defined at the quantum level
- Preserves integrability of the theory
- Allows exact computation of S-matrix elements and energy spectrum

What is the TT?

Standard example: deforming the free boson

$$S_0 = \int d^2x \sqrt{g} (\partial \varphi)^2$$

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Nambu-Goto action in static gauge (+ vacuum energy term)

$$S_{\lambda} \propto \frac{1}{\lambda} \int d^2x \sqrt{g} \left(\sqrt{1 + 2\lambda(\partial\varphi)^2} - 1 \right)$$

Since 2016:

 $\mbox{\ensuremath{\ensuremath{\wp}}}$ Growing interest in $T\overline{T}$ and $T\overline{T}$ -like deformations

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- $\cline{\mathbb{P}}$ Growing interest in $T\overline{T}$ and $T\overline{T}$ -like deformations
- Higher-dimensional extensions (at the classical level)

[Taylor 2018]

[Conti, Romano, Tateo 2022]

$$\partial_{\lambda} S_{\lambda} = \int d^d x \sqrt{g} f\left(T_{ab}\right)$$

Since 2016:

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Stress-tensor deformations as couping to gravity

[Dubovsky, Gorbenko, Mirbabayi 2017] [Tolley 2019]

[Morone, Negro, Tateo 2024]

Today's goal

Look at higher-dimensional $T\overline{T}\mbox{-like}$ deformations from a different persective \ldots

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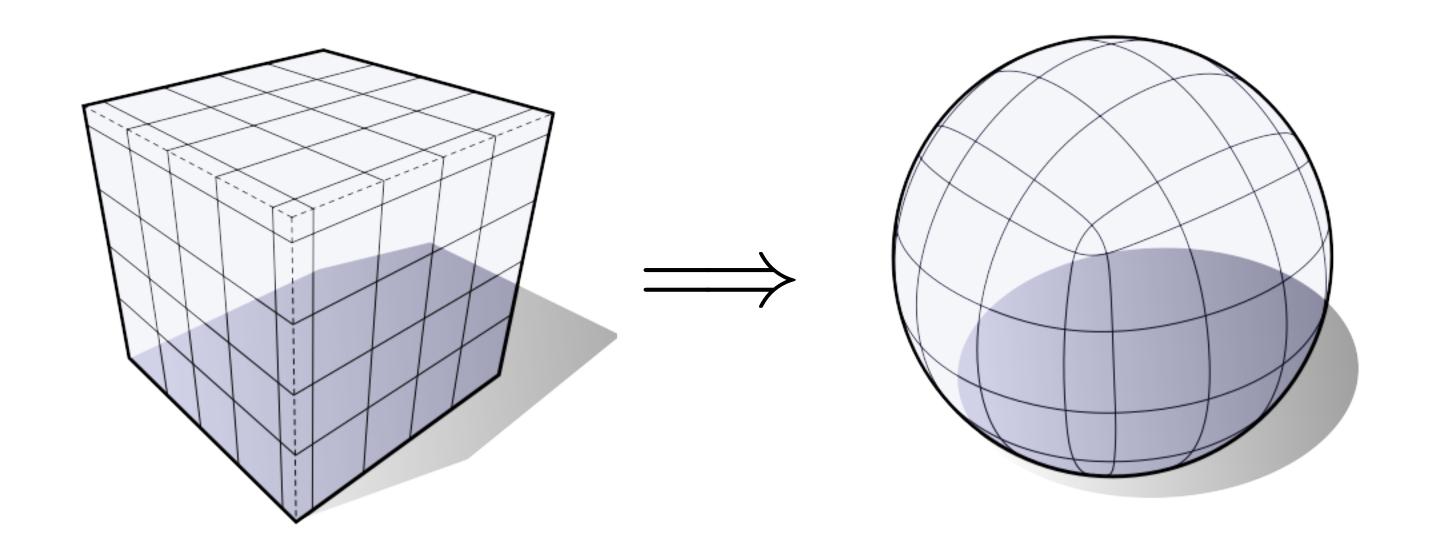
... going back to 1904

The Poincarè conjecture

The Poincarè conjecture

Every simply connected, closed 3-manifold can be continuously deformed into a 3-sphere

[Poincaré, 1904]



100 years later

The entropy formula for the Ricci flow and its geometric applications

Grisha Perelman*

100 years later



Mathematics genius leaves people stunned, rejects \$1 million prize and awards for his work

He ref Grigory Perelman, the maths genius New who said no to \$1m

words igazine

News Featt Perelman cracks a century-old conundrum, refuses the Health Spac reward, and barricades himself in his flat

Society

Grigori Perelman: The genius in hidi

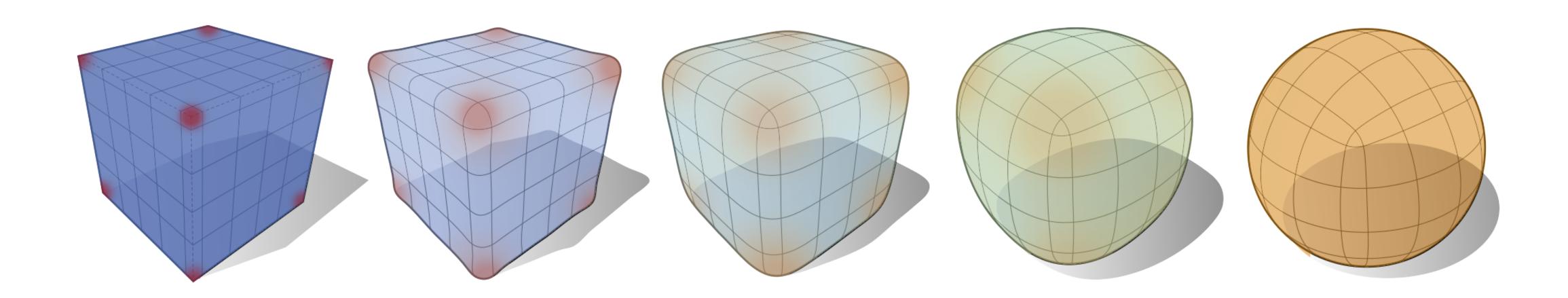
By Jennifer Ouellette

Let geometry do (almost) all the dirty work.

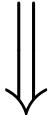
Let geometry do (almost) all the dirty work.



Use geometric flows to "smooth out" high curvature regions and ultimately the manifold reduces to a sphere.



Let geometry do (almost) all the dirty work.

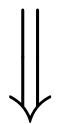


$$\partial_{\lambda}g_{ab}(\lambda)\propto R_{ab}(\lambda)$$
 RICCI FLOW

Let geometry do (almost) all the dirty work.



Let geometry do (almost) all the dirty work.



$$\partial_{\lambda}g_{ab}(\lambda) \propto R_{ab}(\lambda)$$



A COMPLETE PROOF OF THE POINCARÉ AND GEOMETRIZATION CONJECTURES – APPLICATION OF THE HAMILTON-PERELMAN THEORY OF THE RICCI FLOW*

HUAI-DONG CAO† AND XI-PING ZHU‡

The full proof is 320+ pages long!

100 years later

The entropy formula for the Ricci flow and its geometric applications

Grisha Perelman*

100 years later

The entropy formula for the Ricci flow and its geometric applications

5.3* An entropy formula for the Ricci flow in dimension two was found by Chow [C]; there seems to be no relation between his formula and ours.

The interplay of statistical physics and (pseudo)-riemannian geometry occurs in the subject of Black Hole Thermodynamics, developed by Hawking et al. Unfortunately, this subject is beyond my understanding at the moment.

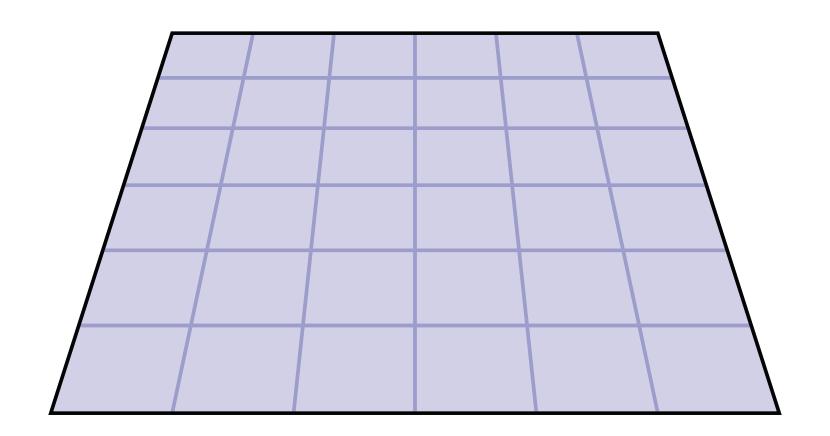
Question

Do Ricci flows arise in physical space-time as well?

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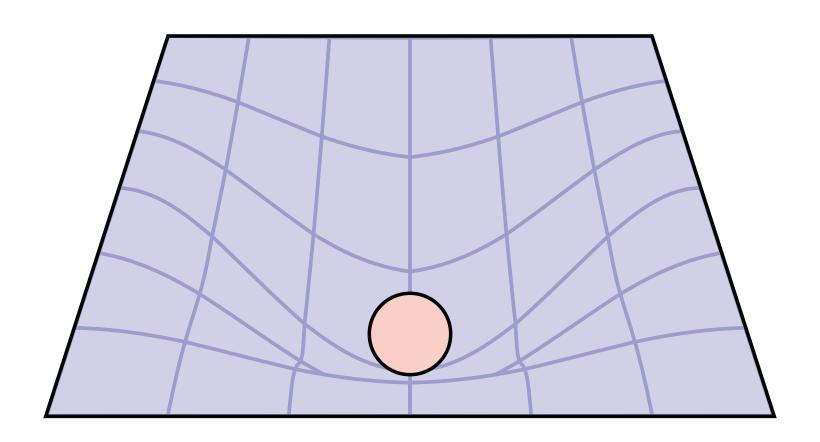
Take a lesson from General Relativity:



Question

Do Ricci flows arise in physical space-time as well?

Take a lesson from General Relativity:



Matter fields influence space-time geometries

$$R_{ab} - \frac{1}{2}Rg_{ab} = T_{ab} \implies \text{solve Einstein field equations to see}$$
 what space-time looks like



Are there solutions to the Einstein field equations which satisfy a Ricci flow?



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Short answer

Yes! If the matter action satisfies the PDE

$$\partial_{\lambda} S_{\lambda} \propto \int d^d x \sqrt{g} \left(T_{ab} T^{ab} - \frac{1}{2} T_a^a \right)$$



Are there solutions to the Einstein field equations which satisfy a Ricci flow?

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[Conti, Negro, Tateo 2018]

$$\partial_{\lambda} S_{\lambda} \propto \int d^d x \sqrt{g} \left(T_{ab} T^{ab} - \frac{1}{2} T_a^a \right)$$

[Brizio, Morone, Tateo 2024]

Ricci solitons = self-similar solutions to the Ricci flow.

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Easy physical realization: vacuum energy term.

$$S(0) = -\int d^4x \sqrt{g} \Lambda \mapsto S(\lambda) = -\int d^4x \sqrt{g} \left(\frac{\Lambda}{1 + \lambda \Lambda}\right)$$

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Deformed field equations:

$$ds^{2} = -f(r)dt^{2} + \frac{dr^{2}}{f(r)} + r^{2}d\Omega^{2}, \quad f(r) = 1 - \frac{2}{3}\frac{\Lambda}{1 + \lambda\Lambda}r^{2}$$

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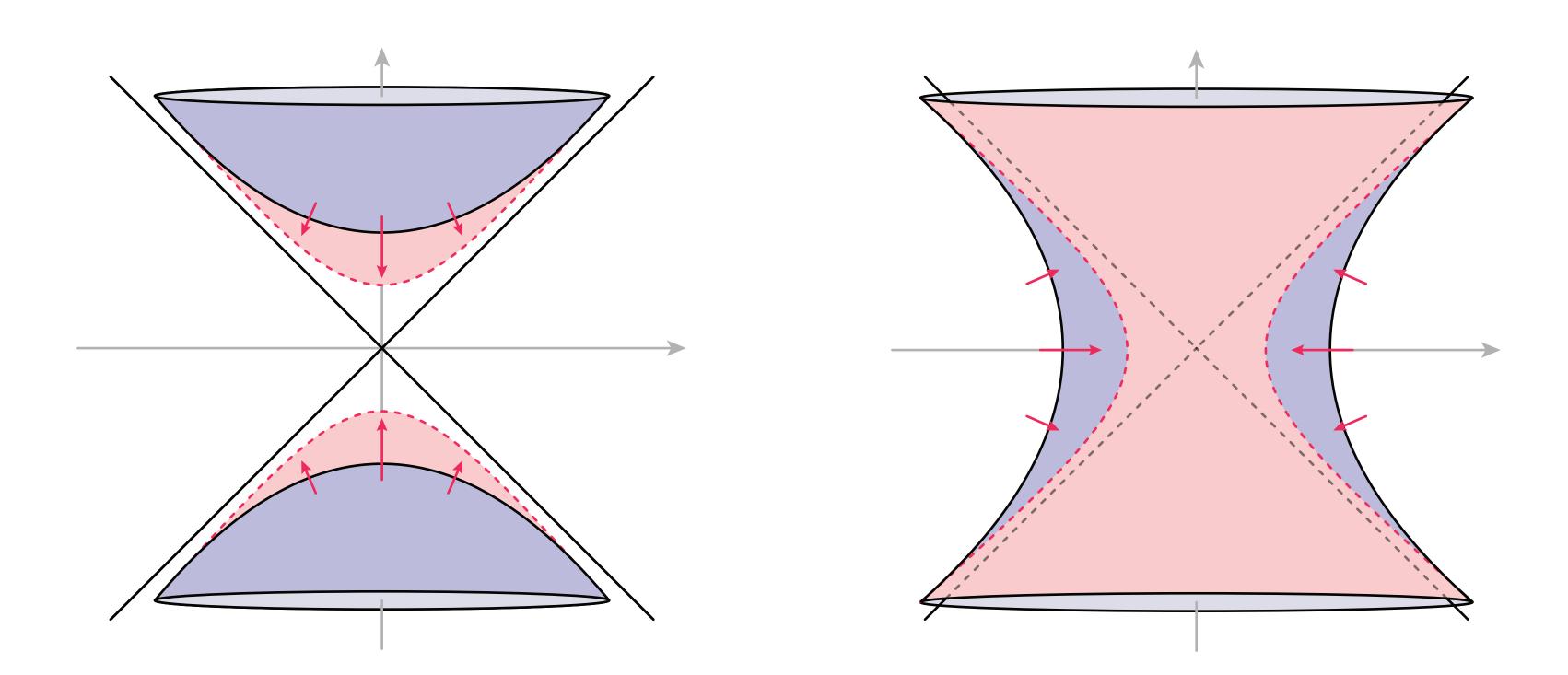
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$$\partial_{\lambda}g_{ab} = R_{ab}$$

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No, you don't!

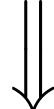
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Dressing mechanisms allow you to work directly on the EoMs.

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Start from known solutions to Einstein's equations



Deform them to obtain Ricci flows in space-time

Start from Reissner–Nordstrom solution = static, spherically symmetric electromagnetic black hole

$$f(r) = 1 - \frac{2m}{r} + \frac{Q^2}{r^2}$$

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$$f(r) = 1 - \frac{2m_0}{r} - \frac{r^2}{3\lambda} \left(1 - \sqrt{1 - \frac{2\lambda Q^2}{r^4}} \right) + \frac{4Q^2}{3r^2} \left[{}_2F_1\left(\frac{1}{4}, \frac{1}{2}; \frac{5}{4}; \frac{2\lambda Q^2}{r^4}\right) \right]$$

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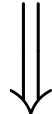
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(Born-Infeld black hole)

$$S_0 = \frac{1}{4} \int d^4x \sqrt{g} F_{ab} F^{ab}$$



$$S_{\lambda} = \frac{1}{\lambda} \int d^4x \left(\sqrt{\det(|g_{ab} + \lambda F_{ab})|} - \sqrt{g} \right)$$

[Conti, Iannella, Negro, Tateo 2018]

Back to 2D:

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 $T\overline{T}$ -deformed matter = undeformed matter + massive gravity theory

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[Tolley 2019]

Ghost-free massive gravity theories \implies bi-metric theories

$$S_{\text{grav}} = S_{\text{grav}}[g, h]$$

2D $T\overline{T}$ deformations:

$$S_{\lambda}[h] = S_0[g^*] + S_{\text{grav}}[g^*, h]$$

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Higher dimensions: [Babaei-Aghbolagh, He, Morone, Ouyang, Tateo 2024]

$$\frac{1}{2} \int d^d x \sqrt{h} R[h] + S_{\lambda}[h] = S_0[g^*] + S_{\text{grav}}[g^*, h]$$

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E.g., in 4D: [Morone, Negro, Tateo 2024]

$$S_{\text{grav}} = \frac{1}{\lambda} \int d^4x \left(\sqrt{|\det(g_{ab} + \lambda R_{ab}(\Gamma))|} - \sqrt{g} \right)$$

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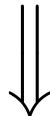
Generally true for any stress-energy tensor deformation.

E.g. root- $T\overline{T}$ deformations: [Babaei-Aghbolagh, He, Morone, Ouyang, Tateo 2024] [Morone, Brizio, Tateo 2024]

$$S_{\text{grav}} = \frac{1}{2} \int d^4x \sqrt{g} \left[R(\Gamma) \cosh \frac{\gamma}{2} + \sinh \frac{\gamma}{2} \sqrt{4R^{ab}(\Gamma)R_{ab}(\Gamma) - (R(\Gamma))^2} \right]$$

Final remarks

 $T\overline{T}$, Ricci flows and massive gravity theories



Look like very different things, but ultimately are the same

- Quantum (or, at least, semi-classical) description?
- Connections with self-dual, non-linear electrodynamics [Aschieri, Ferrara 2013]
- ${\mathbb F}$ $T\overline{T}$ and Bekenstein-Hawking entropy [Morone, Brizio, Tateo 2024]

Thanks!