Spindles, Toric Orbifolds and Holography

Alessio Fontanarossa, University and Politecnico of Turin

based on [2210.16128] and [2402.08724] with D.Martelli, F.Faedo

Seminari di Fisica, Genova 9/10/2024



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- Motivation: Evolution of Physics and Black Holes
- Pramework (Supersymmetry, Superstring, Supergravity, AdS/CFT)
 - The Spindle
- 4 Actual Research (Branes Wrapped on Spindles, Orbifolds and Holography)
- 6 Recap and Comments

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- Gravity \rightarrow classical field theory \rightarrow General Relativity;

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• Effective theory valid until $M_p = \sqrt{\frac{\hbar c}{G_N}} \sim 10^{19}$ GeV.

$$S_{\rm GR} \simeq M_p^2 \int {\rm d}^4 x \sqrt{g} R \rightarrow R_{\mu\nu} - \frac{R}{2} g_{\mu\nu} = 0 \,. \label{eq:GR}$$

- What is gravity above M_p ?
- What is the meaning of a quantum-fluctuating geometry at very small scales

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• BH are giant microwaves in the universe:

$$\begin{split} T_H &= \frac{\hbar}{k_B} \frac{\kappa}{2\pi} \,, \quad -2\kappa^2 \equiv \nabla^{\mu} \xi^{\nu} \nabla_{\mu} \xi_{\nu} \big|_{\text{hor}} \,, \\ \delta M_{BH} &= \frac{\kappa}{8\pi} \delta A_{\text{hor}} + \phi Q + \dots \underset{\text{Hawking radiation}}{\rightarrow} S_{BH} = \frac{A_{\text{hor}}}{4G_N} \,. \end{split}$$

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• Since " $S_{BH} \gg 1 +$ no-hair theorem", where are all the microstates?

String theory in a nutshell

- A change of perspective is necessary:
 - Special Relativity + Quantum Mechanics = QFT; $x^{\mu} = (\vec{x}, t)$ $\hat{x}^{(t)}, \hat{p}^{(t)}$ $\hat{\phi}_{(\vec{x}, t)}$

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String theory in a nutshell



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$$\underbrace{\Box_{\vec{x},y}\phi(x^{\mu},y)=0}_{\text{scalar in }d+1=d(x^{\mu})+1(y)}, \underbrace{\phi(x^{\mu},y)=\sum_{n}\varphi_{n}(x^{\mu})e^{\mathrm{i}ny/R}}_{y \text{ is compact on }S_{R}^{1}} \Longrightarrow \underbrace{\sum_{n}\left[\Box_{x^{\mu}}-\frac{n^{2}}{R^{2}}\right]\varphi_{n}(x^{\mu})=0}_{\infty \text{ massive scalars in }d}.$$

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- Anti-de-Sitter_d:
 - maximally symmetric space time: $R_{\mu\nu\rho\sigma} = L^{-2}(g_{\mu\nu}g_{\rho\sigma} g_{\mu\rho}g_{\nu\sigma});$
 - (maximally supersymmetric s.t.);
 - Einstein s.t.: $R_{\mu\nu} \propto -L^{-2}(d-1)g_{\mu\nu} \implies R < 0;$
 - from the embedding in $\mathbb{R}^{2,d-1}$: $-x_0^2 + \sum_{i=1}^{d-1} x_i^2 x_d^2 = -L^2$; $(ds_{\mathbb{R}^2}^2 = dx^2 + dy^2, x^2 + y^2 = L^2 \implies ds_{S^1}^2 = L^2 d\theta^2)$;
 - invariant under SO(2, d-1);

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 - invariant under SO(2, d-1);
- Conformal Field Theory_d
 - QFT with Poincarrè invariance (Lorentz+translations);
 - CFT=QFT+dilatations+special conformal transformations $\implies SO(2, d)$;

$$(x')^{\mu} = \Lambda^{\mu}_{\nu} x^{\nu} + x^{\mu} + a^{\mu} + \lambda x^{\mu} + \frac{x^{\mu} + b^{\mu} x^{2}}{1 + 2b \cdot x + b^{2} x^{2}},$$

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AdS/CFT correspondence proposal by J. Maldacena. Holography: "there is a correspondence between gravity in AdS_{d+1} and a CFT_d on the boundary of AdS_{d+1} ";





Maldacena motivation:

• The near horizon limit of a stack of N D3-branes, which is the maximally supersymmetric background $AdS_5 \times S^5$;

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$$\mathcal{N} = 4, d = 4$$
 SYM with $G = SU(\mathbf{N})$;

Have the same symmetries!

gravity: $32 \epsilon \times SO(2,4) \times SO(6)$,

CFT: (16 + 16) supercharges \times conf. symmetry \times R-sym. .

Given a particular gravity solution, how to identify the dual (quantum/conformal) theory?

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Branes are the key: whenever a solution can be interpreted as the NHL of some configuration of branes, we have (at least..) a hint of the parent theory. How? We start by the well known dualities

M-theory, M2	$\mathrm{AdS}_4 \times S^7$	$\mathcal{N}=6, d=3$ ABJM
Type IIB, D3	$\mathrm{AdS}_5 \times S^5$	$\mathcal{N}=4, d=4$ SYM
Type IIA, D4	$AdS_6 \times S^4$	$\mathcal{N} = 1, d = 5 \text{ SCFT}$
M-theory, M5	$\mathrm{AdS}_7 \times S^4$	$\mathcal{N} = (2,0), d = 6 \text{ SCFT}$

$$\begin{array}{c} \overbrace{\operatorname{AdS}_{6} \times S^{4}}^{\text{(m)-TypeIIA}} & \longrightarrow \mathcal{N} = 1, d = 5 \\ \\ \text{BH in } \operatorname{AdS}_{6} & \swarrow & (\mathcal{N} = 1, d = 5)/S^{2} \\ & \downarrow & \downarrow \\ \\ \underbrace{\operatorname{AdS}_{4} \times S^{2}}_{\text{Near-horizon geom}} & \times S^{4}_{\longleftarrow} \\ \end{array} \xrightarrow{\mathcal{N} = 2, d = 3} \end{array}$$

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Recap and Comments

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• A spindle Σ is the orbifold $\mathbb{WCP}^1_{[n_-,n_+]} \equiv S^3/U(1)_{n_{\pm}}$ where $\lambda(z_1, z_2) = (\lambda^{n_-} z_1, \lambda^{n_+} z_2), \ \lambda \in U(1);$



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- A spindle \mathbb{Z} is the orbifold $\mathbb{WCP}^1_{[n_-,n_+]} \equiv S^3/U(1)_{n_{\pm}}$ where $\lambda(z_1, z_2) = (\lambda^{n_-} z_1, \lambda^{n_+} z_2), \ \lambda \in U(1);$
- Locally modelled by $\mathbb{C}/\mathbb{Z}_{n_{\pm}}$ singularities: $dS^2 \underset{\theta \sim \theta_{\pm}}{\simeq} d\theta_{\pm}^2 + \frac{\theta_{\pm}^2}{n_{\pm}^2} d\phi^2 \Delta \phi = 2\pi;$



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- Physical perspective: BH can accelerate!
- Spindle solutions (meaning $AdS_D \times \mathbb{Z}$ spacetimes) have been built for D = 2, 3, 4, 5 In the last few years.



Spindles: arXiv:2011.10579

 $\bullet\,$ Minimal gauged supergravity in D=5 has eom and susy variation

$$\begin{split} R_{\mu\nu} &= -4g_{\mu\nu} - \frac{2}{3}F_{\mu\rho}F^{\rho}_{\ \nu} - \frac{1}{9}g_{\mu\nu}F_{\rho\sigma}F^{\rho\sigma} ,\\ d\star F &= -\frac{2}{3}F\wedge F \,, \quad F_{\mu\nu} = (\mathrm{d}A)_{\mu\nu} \,,\\ \Big[\nabla_{\mu} - \frac{\mathrm{i}}{12}(\Gamma^{\nu}_{\ \rho\mu} - 4\delta^{\nu}_{\mu}\Gamma^{\rho})F_{\nu\rho} - \frac{1}{2}\Gamma_{\mu} - \mathrm{i}A_{\mu}\Big]\epsilon = 0 \end{split}$$

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Near-horizon-like geometry

$$ds_{5}^{2} = \frac{4y}{9} ds_{AdS_{3}}^{2} + ds_{\Sigma}^{2}, \quad A = \frac{1}{4} \left(1 - \frac{a}{y} \right) dz,$$
$$ds_{\Sigma}^{2} = \frac{y}{q(y)} dy^{2} + \frac{q(y)}{36y^{2}} dz^{2} \implies \Delta z =?, \quad q(y) = 4y^{3} - 9y^{2} + 6ay - a^{2}.$$

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if

$$a = \frac{n_-^2 - n_+^2}{n_-^2 + n_+^2}, \quad \Delta z = \frac{\sqrt{2}\sqrt{n_-^2 + n_+^2}}{n_- n_+}\pi,$$

then

$$\frac{1}{2\pi}\int_{\mathbb{Z}}F=\frac{n_--n_+}{2n_-n_+}\,,\quad \chi(\mathbb{Z})=\frac{1}{4\pi}\int_{\mathbb{Z}}R_{\mathbb{Z}}\operatorname{vol}_{\mathbb{Z}}=\frac{n_-+n_+}{n_-n_+}.$$

N D3 at the sing. of the CY_3 cone over SE_5 : $AdS_5 \times SE^5 \longleftrightarrow \mathcal{N} = 1, d = 4$

$$\begin{array}{c} \mathsf{B}\mathsf{H} \text{ in } \mathrm{AdS}_5 \hspace{0.5cm} \bigcup \hspace{0.5cm} (\mathcal{N}=1,d=4)/\mathbb{Z} \\ \hspace{0.5cm} \bigcup \hspace{0.5cm} \bigcup \hspace{0.5cm} \mathsf{AdS}_3 \times \mathbb{Z} \times SE^5 {\longleftrightarrow} \mathcal{N}=(2,0), d=2 \end{array}$$

$$S_{BH}\big|_{\rm grav} = \log(\Omega)\big|_{\rm CFT}$$

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 In D = 6 gauged Supegravity we found (arXiv: 2210.16128) AdS₂ × Σ_g κ Σ₂ and more surprisingly AdS₂ × Σ₁ κ Σ₂;

$$\begin{split} \mathrm{d}s^2 = &(y^2 h_1 h_2)^{1/4} \left[\frac{x^2}{4} \, \mathrm{d}s^2_{\mathrm{AdS}_2} + \overbrace{\frac{x^2}{q(x)}}^{\overline{\Sigma}_{[m_-,m_+]}} \mathrm{d}x^2 + \frac{q(x)}{4x^2} \, \mathrm{d}\psi^2 \right. \\ &+ \underbrace{\frac{y^2}{f(y)}}_{\overline{\Sigma}_{[n_-,n_+]}} \mathrm{d}y^2 + \underbrace{\frac{f(y)}{h_1 h_2}}_{\overline{\Sigma}_{[n_-,n_+]}} - \underbrace{\left[\frac{1}{2m} \left(1 - \frac{\mathbf{a}}{x} \right) \mathrm{d}\psi} \right)^2 \right], \\ A_i &= -\frac{y^3}{h_i} \left(\mathrm{d}z - \frac{1}{2m} \left(1 - \frac{\mathbf{a}}{x} \right) \mathrm{d}\psi \right), \qquad X_i = (y^2 h_1 h_2)^{3/8} h_i^{-1}, \\ B &= \frac{\mathbf{a}y}{2m} \operatorname{vol}(\mathrm{AdS}_2), \\ h_i(y) &= y^3 + q_i, \ f(y) = m^2 h_1 h_2 - y^4, \ q(x) = x^4 - 4x^2 + 4\mathbf{a}x - \mathbf{a}^2. \end{split}$$

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- In D = 6 gauged Supegravity we found (arXiv: 2210.16128) AdS₂ × Σ_g κ Σ₂ and more surprisingly AdS₂ × Σ₁ κ Σ₂;
- These can be uplifted to D = 10 solutions of massive type IIA sugra: no singularities, thus allowed AdS/CFT computations!

$$\begin{split} \mathrm{d}s^2 = & (y^2 h_1 h_2)^{1/4} \left[\frac{x^2}{4} \, \mathrm{d}s^2_{\mathrm{AdS}_2} + \overbrace{\frac{x^2}{q(x)} \, \mathrm{d}x^2 + \frac{q(x)}{4x^2} \, \mathrm{d}\psi^2}^{\mathbb{E}_{[m_-,m_+]}} \right. \\ & \left. + \underbrace{\frac{y^2}{f(y)} \, \mathrm{d}y^2 + \frac{f(y)}{h_1 h_2} \Big(\mathrm{d}z - \underbrace{\frac{1}{2m} \Big(1 - \frac{\mathbf{a}}{x} \Big) \, \mathrm{d}\psi} \Big)^2 \Big] \,, \\ & \left. A_i = -\frac{y^3}{h_i} \Big(\mathrm{d}z - \frac{1}{2m} \Big(1 - \frac{\mathbf{a}}{x} \Big) \, \mathrm{d}\psi \Big) \,, \qquad X_i = (y^2 h_1 h_2)^{3/8} h_i^{-1} \,, \\ & B = \frac{\mathbf{a}y}{2m} \operatorname{vol}(\mathrm{AdS}_2) \,, \\ & h_i(y) = y^3 + q_i \,, \ f(y) = m^2 h_1 h_2 - y^4 \,, \ q(x) = x^4 - 4x^2 + 4\mathbf{a}x - \mathbf{a}^2 \,. \end{split}$$

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 After uplifting to massive (m)type IIA sugra → near-horizon limit of a system of N D4-branes and N_f D8-branes wrapped on M = ∑₁ × ∑₂;

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- The dual CFT will (should..) be again obtained as $(\mathcal{N} = 1, d = 5)/\mathbb{M}$;
- Difficult in practice! (Sometimes nearly impossible...)
- We take advantage from the toric properties of our solution:





$$S = \frac{\sqrt{2}\sqrt{m_{+}^{2} + m_{-}^{2}} - (m_{+} + m_{-})}{2m_{+}m_{-}} \frac{n_{+} + n_{-}}{n_{+}n_{-}} \frac{\sqrt{3}\pi N^{5/2}}{\sqrt{8 - N_{f}}} \frac{[3\mu(x^{2} + 1) - x(x^{2} + 5)]^{3/2}}{x(x^{2} + 3)(\mu - x)^{1/2}},$$

$$t_{i} \equiv \frac{1}{2\pi} \int_{\Sigma_{2}} dA_{R} = \frac{+n_{+} + n_{-}}{n_{+}n_{-}}, \quad s_{i}^{\pm} \equiv \frac{1}{2\pi} \int_{\Sigma_{1}^{\pm}} dA_{R} = \frac{+m_{+} - m_{-}}{m_{+}m_{-}} \pm \frac{t/n_{\pm}}{m_{+}m_{-}},$$

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 Idea: Use a "geometrical" approach to find the entropy functions from the toric data! For "'toric-symplectic'" orbifolds we conjectured an off-shell free energy.

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$$\begin{split} &\overbrace{F(\varphi_i,\varepsilon_I;\mathfrak{p}^a_i)\propto\sum_a\frac{\eta^a_d}{d_{a,a+1}}\frac{\mathcal{F}_d(\Phi^a_i)}{\varepsilon^a_1\varepsilon^a_2}}, \quad \mathcal{F}_d(\Phi^a_i)\propto N^{(d-1)/2}(\Phi^a_1\Phi^a_2)^{(d-3)/2}, \\ &\Phi^a_i=\varphi_i-\mathfrak{p}^a_i\varepsilon^a_1-\mathfrak{p}^{a+1}_i\varepsilon^a_2, \quad \varepsilon^a_1=-\frac{\det(\vec{w}_{a+1},\vec{\varepsilon})}{d_{a,a+1}}, \quad \varepsilon^a_2=\frac{\det(\vec{w}_a,\vec{\varepsilon})}{d_{a,a+1}}, \\ &\varphi_1+\varphi_2-\det(\vec{W},\vec{\varepsilon})=2, \quad \mathfrak{p}^a_1+\mathfrak{p}^a_2=\sigma^a+\det(\vec{W},\vec{w}_a), \\ &\mathfrak{q}^a_i=\sum_a D_{ab}\mathfrak{p}^b_i, \quad \sum_a\mathfrak{q}^a_i\vec{w}_a=0 \implies \mathfrak{q}^a_R=\sum_b D_{ab}\sigma^b, \end{split}$$

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From 2402.08724:

$$ds^{2} = (H_{1}H_{2})^{1/(D-2)} \left(\frac{x^{2}y^{2}}{a^{2}} ds^{2}_{AdS_{D-4}} + ds^{2}_{\mathbb{M}_{4}}\right),$$

where

$$ds_{\mathbb{M}_{4}}^{2} = \frac{(H_{1}H_{2})^{-1}}{\Xi^{2}(x^{2} - y^{2})} \left[(V_{y}^{2}\Delta_{x} - V_{x}^{2}\Delta_{y}) d\psi^{2} + (\widetilde{V}_{y}^{2}\Delta_{x} - \widetilde{V}_{x}^{2}\Delta_{y}) d\phi^{2} - \frac{4c_{1}c_{2}\widetilde{c}_{1}\widetilde{c}_{2}}{a^{2}} \left(\frac{N_{y}\Delta_{x}}{y^{D-5}} - \frac{N_{x}\Delta_{y}}{x^{D-5}} \right) d\psi d\phi \right] + \frac{y^{2} - x^{2}}{\Delta_{y}} dy^{2} + \frac{x^{2} - y^{2}}{\Delta_{x}} dx^{2}.$$

The scalar fields are

$$X_i = (H_1 H_2)^{(D-3)/2(D-2)} H_i^{-1},$$

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the gauge potentials are given by

$$\begin{split} A_1 &= \frac{2N_y s_1}{\Xi (x^2 - y^2) H_1 y^{D-5}} \left(c_1 \widetilde{c}_2 V_x^{(1)} \, \mathrm{d}\psi - c_2 \widetilde{c}_1 \widetilde{V}_x^{(1)} \, \mathrm{d}\phi \right) \\ &- \frac{2N_x s_1}{\Xi (x^2 - y^2) H_1 x^{D-5}} \left(c_1 \widetilde{c}_2 V_y^{(1)} \, \mathrm{d}\psi - c_2 \widetilde{c}_1 \widetilde{V}_y^{(1)} \, \mathrm{d}\phi \right), \\ A_2 &= \frac{2N_y s_2}{\Xi (x^2 - y^2) H_2 y^{D-5}} \left(c_2 \widetilde{c}_1 V_x^{(2)} \, \mathrm{d}\psi - c_1 \widetilde{c}_2 \widetilde{V}_x^{(2)} \, \mathrm{d}\phi \right) \\ &- \frac{2N_x s_2}{\Xi (x^2 - y^2) H_2 x^{D-5}} \left(c_2 \widetilde{c}_1 V_y^{(2)} \, \mathrm{d}\psi - c_1 \widetilde{c}_2 \widetilde{V}_y^{(2)} \, \mathrm{d}\phi \right), \end{split}$$

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and the $\left(D-4\right)\text{-}\mathsf{form}$ is

$$B = -\frac{4s_1s_2(N_yx^3 - N_xy^3)}{a^2(x^2 - y^2)} \operatorname{vol}(\operatorname{AdS}_2) \qquad (D = 6),$$

$$B = -\frac{2s_1s_2(N_yx^4 - N_xy^4)}{a^3(x^2 - y^2)} \operatorname{vol}(\operatorname{AdS}_3) - \frac{2N_ygs_1s_2x}{\Xi a^2y^2} \, \mathrm{d}\psi \wedge \mathrm{d}\phi \wedge \mathrm{d}x \qquad (D = 7).$$

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The functions are

$$\begin{split} \Delta_y &= -y^2 + a^2 - \frac{2N_y}{y^{D-5}} + g^2 \left(y^2 - \frac{2N_y s_1^2}{y^{D-5}} \right) \left(y^2 - \frac{2N_y s_2^2}{y^{D-5}} \right) - a^2 g^2 y^2 - \frac{2N_y a^2 g^2 s_1^2 s_2^2}{y^{D-5}} \,, \\ \Delta_x &= -x^2 + a^2 - \frac{2N_x}{x^{D-5}} + g^2 \left(x^2 - \frac{2N_x s_1^2}{x^{D-5}} \right) \left(x^2 - \frac{2N_x s_2^2}{x^{D-5}} \right) - a^2 g^2 x^2 - \frac{2N_x a^2 g^2 s_1^2 s_2^2}{x^{D-5}} \,, \\ V_y^2 &= V_y^{(1)} V_y^{(2)} \,, \qquad \widetilde{V}_y^2 = \widetilde{V}_y^{(1)} \widetilde{V}_y^{(2)} \,, \qquad V_R^2 = V_R^{(1)} V_R^{(2)} \,, \qquad \widetilde{V}_R^2 = \widetilde{V}_R^{(1)} \widetilde{V}_R^{(2)} \,, \\ V_y^{(i)} &= 1 - g^2 \left(y^2 - \frac{2N_y s_i^2}{y^{D-5}} \right) \,, \qquad \widetilde{V}_y^{(i)} = 1 - \frac{1}{a^2} \left(y^2 - \frac{2N_y s_i^2}{y^{D-5}} \right) \,, \\ V_x^{(i)} &= 1 - g^2 \left(x^2 - \frac{2N_x s_i^2}{x^{D-5}} \right) \,, \qquad \widetilde{V}_x^{(i)} = 1 - \frac{1}{a^2} \left(x^2 - \frac{2N_x s_i^2}{x^{D-5}} \right) \,, \\ H_i &= 1 + \frac{2s_i^2}{x^2 - y^2} \left(\frac{N_y}{y^{D-5}} - \frac{N_x}{x^{D-5}} \right) \,, \\ \Xi &= 1 - a^2 g^2 \,, \qquad s_i = \sinh \delta_i \,, \qquad c_i = \cosh \delta_i \,, \qquad \widetilde{c}_i = \sqrt{1 + a^2 g^2 s_i^2} \,. \end{split}$$

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Table of Contents

- Motivation: Evolution of Physics and Black Holes
- Pramework (Supersymmetry, Superstring, Supergravity, AdS/CFT)
- 3 The Spindle
- 4 Actual Research (Branes Wrapped on Spindles, Orbifolds and Holography)

6 Recap and Comments

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Comments:

- Our conjecture is valid also for M5-branes wrapped over \mathbb{M}_4 with minor tweaks and for an arbitrary number of fixed points;
- It works also for the $AdS_{2,3} \times \mathbb{M}_4$ solutions of arXiv:2402.08724;
- sum over fixed points \implies equivariant integration at work \implies equivariant volume ([Colombo-Faedo-Martelli-Zaffaroni])/ localization of the action (Genolini-Gauntlett-Sparks);
- Proved: M2, D3 and M5 on $\mathbb{\Sigma},$ M5 on $\mathbb{M}_4;$

Some open problems:

- D4 branes?
- Are there other interesting equivariant quantities?
- Black objects with \mathbb{M}_4 horizon in D = 6,7 gauged supergravity?
- Existence of solutions?

$$\begin{aligned} \mathcal{S}_{\pm}(\varphi_i,\varepsilon;p_i,n_{\pm}) &\equiv \frac{1}{\varepsilon} \left[F_d(\varphi_i + p_i\varepsilon) \pm F_d(\varphi_i - p_i\varepsilon) \right], \\ \sum_i \varphi_i - \frac{n_- - \sigma n_+}{n_- n_+} \varepsilon &= 2, \quad \sum_i p_i = \frac{n_- + \sigma n_+}{n_- n_+}, \end{aligned}$$

Thanks for your attention!

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Back-up slides

$$\mathbb{S}^{3}(a,b;m_{\pm};t) \equiv \mathbb{S}^{3}(a,b;m_{\pm})/\mathbb{Z}_{t} = (\mathbb{C}_{1}/\mathbb{Z}_{m_{-}} \times \mathbb{C}_{2}/\mathbb{Z}_{m_{+}})/\mathbb{Z}_{t}, \quad \frac{|z_{1}|^{2}}{a^{2}} + \frac{|z_{2}|^{2}}{b^{2}} = 1$$

with the identification

$$(z_1, z_2) \sim e^{\frac{2\pi i}{t}} \circ \left(e^{\frac{2\pi i}{m_-}} z_1, e^{\frac{2\pi i}{m_+}} z_2 \right) \sim \left(e^{\frac{2\pi i}{t}} z_1, e^{\frac{2\pi i}{t}} z_1, e^{\frac{2\pi i}{t}} z_2 \right) \sim \left(e^{\frac{2\pi i}{t}} z_1, e^{\frac{2\pi i}{t}} z_2 \right).$$

Consider $t = k_+ k_- \overline{t}$ and $m_{\pm} = k_{\pm} \overline{m}_{\pm}$:

$$(e^{\frac{2\pi i m_{+}}{t}}z_{1}, e^{\frac{2\pi i m_{-}}{t}}z_{2}) = (e^{\frac{2\pi i \overline{m}_{+}}{\overline{t}k_{-}}}z_{1}, e^{\frac{2\pi i \overline{m}_{-}}{\overline{t}k_{+}}}z_{2}) \implies (e^{\frac{2\pi i \overline{m}_{+}k_{+}}{k_{-}}}z_{1}, e^{2\pi i \overline{m}_{-}}z_{2}).$$

Alternatively

$$\vec{\hat{w}}_1 = (n_-, 0), \quad \vec{\hat{w}}_2 = (ta_+, m_-), \quad \vec{\hat{w}}_3 = (-n_+, 0), \quad \vec{\hat{w}}_4 = (ta_-, -m_+),$$

with $a_+m_+ + a_-m_- = 1$ (or $r_+m_+ + r_-m_- = t$ with $r_{\pm} = ta_{\pm}$).

Recap and Comments

Back-up slides

$$\begin{split} & Z_{\mathsf{GC}}^{\mathsf{CFT}}(\varphi_a, \varepsilon_i; p_a) = \sum_{Q_a, J_i} c(Q_a, J_i) \mathrm{e}^{\mathrm{i}(\varphi_a Q_a + \varepsilon_i J_i)} \,, \\ & \mathrm{e}^{S_{\mathsf{BH}}(Q_a, J_i)} = c(Q_a, J_i) = \int \frac{\mathrm{d}\varphi_a}{2\pi} \frac{\mathrm{d}\varepsilon_i}{2\pi} Z_{\mathsf{GC}}^{\mathsf{CFT}}(\varphi_a, \varepsilon_i) \mathrm{e}^{-\mathrm{i}(\varphi_a Q_a + \varepsilon_i J_i)} \,, \\ & S_{\mathsf{BH}}(Q_a, J_i) = \left[\log Z_{\mathsf{GC}}^{\mathsf{CFT}}(\varphi_a, \varepsilon_i) - \mathrm{i}(\varphi_a Q_a + \varepsilon_i J_i) \right] \Big|_{(\varphi_a^*, \varepsilon_i^*)} \,, \\ & \partial_{(\varphi, \varepsilon)} \Big[\log Z_{\mathsf{GC}}^{\mathsf{CFT}}(\varphi_a, \varepsilon_i) - \mathrm{i}(\varphi_a Q_a + \varepsilon_i J_i) \Big] \Big|_{(\varphi_a^*, \varepsilon_i^*)} \equiv 0 \,. \end{split}$$

Moreover

$$Z_{\mathsf{CFT}} = Z_{\mathsf{grav}} = \int \mathcal{D}[\Phi] \mathrm{e}^{-S[\Phi]} \longrightarrow \mathrm{e}^{-S^{\mathsf{OS}}[\Phi]} \,.$$

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Back-up slides

Squashed-branched lens space:

$$\begin{split} \mathbb{S}^{3}(a,b;m_{\pm};t) &\equiv \mathbb{S}^{3}(a,b;m_{\pm})/\mathbb{Z}_{t} = (\mathbb{C}_{1}/\mathbb{Z}_{m_{-}} \times \mathbb{C}_{2}/\mathbb{Z}_{m_{+}})/\mathbb{Z}_{t} \,,\\ &\frac{|z_{1}|^{2}}{a^{2}} + \frac{|z_{2}|^{2}}{b^{2}} = 1 \,, \end{split}$$

with the identification

$$(z_1, z_2) \sim e^{\frac{2\pi i}{t}} \circ \left(e^{\frac{2\pi i}{m_-}} z_1, e^{\frac{2\pi i}{m_+}} z_2 \right) \sim \left(e^{\frac{2\pi i}{t m_-}} z_1, e^{\frac{2\pi i}{t m_+}} z_2 \right) \sim \left(e^{\frac{2\pi i m_+}{t}} z_1, e^{\frac{2\pi i m_-}{t}} z_2 \right).$$

Consider $t = k_+ k_- \overline{t}$ and $m_{\pm} = k_{\pm} \overline{m}_{\pm}$:

$$(e^{\frac{2\pi i m_{+}}{t}}z_{1}, e^{\frac{2\pi i m_{-}}{t}}z_{2}) = (e^{\frac{2\pi i \overline{m}_{+}}{\overline{t}k_{-}}}z_{1}, e^{\frac{2\pi i \overline{m}_{-}}{\overline{t}k_{+}}}z_{2}) \implies (e^{\frac{2\pi i \overline{m}_{+}k_{+}}{k_{-}}}z_{1}, e^{2\pi i \overline{m}_{-}}z_{2}).$$

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$$\begin{aligned} x^2 + y^2 + z^2 + t^2 &= 1 \,, \quad z_1 = x + \mathrm{i}y \,, \quad z_2 = z + \mathrm{i}t \quad |z_1|^2 + |z_2|^2 &= 1 \,, \end{aligned}$$

Choosing $(z_1, z_2) = (\cos \frac{\theta}{2} e^{\mathrm{i}(-\phi/2 + \psi)}, \sin \frac{\theta}{2} e^{\mathrm{i}(\phi/2 + \psi)})$ the metric appears as the

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