PMT Reco & Analysis - Previously





Figure 1.18: Full example of the analysis pipeline for the alpha particles identified in a given picture (run 41525, picture 399): (a) Original picture and sCMOS analysis of both alphas; (b) Light transverse profile of the alpha tracks; (c) and (d) the set of 4 waveforms for each trigger identified as alpha tracks; (e) waveform - cluster association and final 3D projection in the real LIME framework.

Previous episodes:

1. [https://agenda.infn.it/event/41735/]

Initial look at alpha tracks for directional & head-tail determination

2. [https://agenda.infn.it/event/42030/]

<u>1 Update on ... - 3D reconstructed alpha tracks</u>

- 3. [https://agenda.infn.it/event/42653/]
 - <u>2 Update on ... 3D reconstructed alpha tracks</u>

Full framework retrieves and saves all the relevant information automatically, including plots .

These are *not* cherry-picked events



#3 Update on ... 3D alphas - Optimization of parameters

David Marques and PMT Working Group Technical / Analysis meeting 18-07-2024



<u>Fixing camera</u> <u>coordinates & rotations</u>



... There's still work to be done:



Some incoherence in the 3D projection and CMOS data...

... Maybe there are some coordinate / granularity / centering corrections to do...



Let's analyze the lengths of the tracks

Using **sc_length**



Track start = With Analyzer::Edges (first())* granularity

Track end =

- (sc_length[sc_i] * granularity * cos(cam.angle_XY * TMath::Pi()/180.));
- (sc_length[sc_i] * granularity * sin(cam.angle_XY * TMath::Pi()/180.));



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Using **edges** (from cmos:analyzer class)



Track start = With Analyzer::Edges (first())* granularity
Track end = With Analyzer::Edges (back())* granularity



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Track start = With *Analyzer::Edges* (first())* granularity

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Track end = With Analyzer::Edges (back())* granularity



→ It's because of the shadows 🗲

• These get merged into the main cluster, making some of the variables wrong...



Track and shadow merged in the final code ⇒ **Be careful**

with reco variables for alphas!

The alpha selection



The alpha selection with occupancy>0.7



- These are the first 10 selected events

D. Pinci saw the same issue



→ It's because of the shadows 🔚

- These get merged into the main cluster, making some of the variables wrong...
- With the <u>CMOS::analyzer</u> this doesn't happen because the "<u>removeNoise</u>" function removes all of this
 - (NB: If the cluster gets mixed with an alpha with a cosmic, still doesn't work. (CMOS reco could be improved))



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→ Track ends up outside the frame?! Length very big ?!

→ Also greatly impacts Angle Z and total length!



Some more examples to confirm it's working...





Some more examples to confirm it's working...



- I would say the variables matching are on point! \Rightarrow Solved \checkmark
- Additionally, if you look at the coordinates, it's not 33x33.
 - This also proves how much the **camera sampling** \Rightarrow 36x36 cm².

TH3F *axis = new TH3F(name, name,

1, <u>0, 36</u>, 1, 0, 50, 1, <u>0, 36</u>);



#4 Update on ...



<u>Statistical results +</u>

comparison with D. Pinci

David Marques

Technical / Analysis meeting 18-07-2024



Statistical Analysis

PMT 3D reco – Conclusions



→ But first...



To improve the **cluster-trigger association**

(1-to-1 association), we use the **BAT-fit to**

position the PMT signal in the GEM plane:





To improve the <u>cluster-trigger association</u> (1-to-1 association), we use the *BAT-fit to position the PMT signal in the GEM plane:*

- 1. Slice waveform
- BAT-fit the slice integrated charge ⇒
 (L,X,Y)
- 3. Place the point in the GEM plane.



$$V = R * I \Rightarrow V = R * Q/\Delta t \Rightarrow Q = \frac{V * t}{R}$$
 (1.4)

$$Q[nC] = A[ADU] * \frac{DGTZ \text{ dynamic range}[V]}{DGTZ \text{ resolution}[bits]} * \Delta t / R$$
$$= A[ADU] * \frac{1[V]}{12[bits]} * \frac{1}{DGTZ \text{ sampl. freq.}}[ns] / R[Omega] \qquad (1.5)$$
$$= A * \frac{1}{4096} * \frac{4}{3} * \frac{1}{50}$$



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- 4. Same for CMOS (using Analyzer::Edges)
- 5. Distance between points calculated
- 6. Cluster-trigger association done by smaller distances





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⇒ Works surprisingly well!









We can study the efficiency of this fit: 350 <u>ि</u> 2000 Y distance [pixe 300 250 200 150 -500 100 -1000 -1500 a hara hara hara hara hara hara -2000 2000-1500-1000 -500 0 500 1000 1500 2000 Counts [#] X distances X Distances Entries 264000 -53.04 Mean Y Distances Std Dev 147.1 10000 Y distances 264000 Entries Mean -0.4876 8000 Std Dev 149 6000 4000 2000 -1000 -800 -600 -400 -200 0 200 400 600 800 1000 Distance [pixels]

- We get **not very Gaussian distributions** (and it wasn't expected)
 - Standard deviations of <u>~150 pixels = 2,325 cm</u>



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- At the moment, is only necessary for matching *alpha* tracks.
 - Rarely 2 in one pic.
 - This resolution is more than enough.
 - At closer distances, also the <u>CMOS reco</u> starts *failing and* merging the tracks, which renders useless the 3D reco.





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I have many examples, there's a bit of everything...



David M.



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There are clear reasons to not have a perfect fit:

- 1. When the <u>waveforms are saturated</u> due to high gain, we lose proportionality \Rightarrow BAT-fit works better in the middle region.
- 2. Offsets due to improve this, but I haven't checked.
- 3. <u>Barreling effect</u> from lens towards the sides farther difficulties the fit ⇒ Giorgio working on it

While this is interesting, the optimization; test of other types of particles; implementation at front-end level is out-of-scope for my work.

GS



Statistical Results



- → Now we can actually do some statistical analysis on the results
 - Eventually compare with MC truth from Flaminia's simulations



The datasets used were:

- 1. Run 3
 - For optimization only
- 2. <u>Run 4</u>
 - Most of the long ranges of
 Bkg + calibs
- 3. Run 5
 - Different gain:
 - Interesting by will
 - *maybe* require
 - parameter tuning... in
 - the pipeline.





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TATUS	present in SSD?	Start	Stop	Numbers		Description	Data runs	Data pics Gas FI	ow Filter Line 1	Filter Line2	2 !
		2023-12-01 15:08	2023-12-04 9:39	40784 -	40917	Stability	133	53200	5 Blu	Not in use	1
ONE	YES	2023-12-04 10:23	2023-12-14 16:40	40919 -	42848	Bkg + Daily Calibrations	1929	771600	5 Blu	Not in use	Th
		2023-12-14 18:07:07	2023-12-16 10:17:27	42863 -	43185	Bkg + Daily Calibrations Low GAIN	322	128800	5 Blu	Not in use	the
		2023-12-15 11:54:46	2023-12-15 14:44:20	42985	43050	VGEM1 scan	65	26000	5¦Blu	Not in use	1
		2023-12-16 23:50:59	2023-12-17 21:53:14	43186 -	43231	Stability + Daily Calibrations- LOW Gas Flow : 2 I/h	45	18000	2 Blu	Not in use	1
		2023-12-17 22:45:16	2023-12-22 16:18:47	43232 -	43308	Stability + Daily Calibrations- LOW Gas Flow : 1 I/h	76	30400	1'Blu	Not in use	1
		2023-12-17 22:45:16	2023-12-22 16:18:47	43316 -	43486	Stability + Daily Calibrations- LOW Gas Flow : 1 l/h	170	68000	1 Blu + Rosso	Not in use	1
		2023-12-31	1	43502 -	43508	Daily Calibration			1 Blu + Rosso	Not in use	1
				43509 -	43515	Daily Calibration			1 Blu + Rosso	Not in use	1
		2024-01-04	1	43517 -	43522	Daily Calibration			1 Blu + Rosso	Not in use	-
		2024-01-06	5	43524 -	43529	Daily Calibration			1 Blu + Rosso	Not in use	1
		2024-01-08	3	43531 -	43536	Daily Calibration			2 Blu + Rosso	Not in use	1
		2024-01-10)	43636 -	43641	Daily Calibration			2 Blu + Rosso	Not in use	1
		2024-01-12	2	43732 -	43738	Daily Calibration			5 Blu + Rosso	Not in use	FIC
		2024-01-14	4	43849 -	43855	Daily Calibration			5 Blu + Rosso	Not in use	1
		2024-01-10	5	44047 -	44053	Daily Calibration			5 Blu + Rosso	Not in use	-
		2024-01-1	7	44203 -	44209	Daily Calibration			5 Blu + Rosso	Not in use	1
		2024-01-18	3	44367 -	44372	Daily Calibration			5 Blu + Rosso	Not in use	1
		2024-01-19	9	44553 -	44559	Daily Calibration			5 Blu + Rosso	Not in use	1
		2024-01-08 12:00:15	2024-01-08 18:38:15	43537 -	43701	Stability + Daily Calibrations- LOW Gas Flow : 2 I/h	164	65600	2,Blu + Rosso	Not in use	1
		2024-01-08 18:38:15	2024-01-15 9:00:00	43702 -	43885	Stability + Daily Calibrations- Gas Flow : 5 l/h	183	73200	5 Blu + Rosso	Not in use	1
0		2024-01-15 9:11:14	2024-01-23 12:31	43886 -	45213	Bkg + Daily Calibrations	1327	530800	5 Blu + Rosso	Not in use	1
		2024-01-23 15:44:30	2024-01-24 9:53:11	45214 -	45251	Stability + Daily Calibrations- Gas Flow : 5 I/h	37	14800	5 Blu + Rosso	Not in use	i.
ONE	YES	2024-01-24 10:27:00	2024-02-02 9:42	45259 -	46628	Bkg + Daily Calibrations	1369	547600	5 Blu + Rosso	Not in use	
		2024-02-02 9:47:28	2024-02-04 11:21:19	46636 -	46740	Stability + Daily Calibrations- Gas Flow : 4 I/h	104	41600	5 Blu + Rosso	Not in use	1
		2024-02-04 11:21:19	2024-02-05 14:03:49	46741 -	46802	Stability + Daily Calibrations- Gas Flow : 5 I/h	61	24400	5 Blu + Rosso	Not in use	1
		2024-02-05 14:13:49	2024-02-06 23:50	46803 -	47023	Bkg + Daily Calibrations	220	88000	5 Blu + Rosso	Not in use	1
		2024-02-06 23:59:42	2024-02-07 10:03:47	47024 -	47051	Stability + Daily Calibrations- LOW Gas Flow : 1 l/h	27	10800	1 Blu + Rosso	Not in use	1
		2024-02-07 10:03:47	1	47052 -	47108	Stability + Daily Calibrations- LOW Gas Flow : 0 I/h	56	22400	1 Blu + Rosso	Not in use	1
				47982 -	47985	DT test: trigger rate 36 Hz, PMT 590 V	4	1600	1 Blu + Rosso	Not in use	1
				47986 -	47989	DT test: trigger rate 26 Hz, PMT 580 V	4	1600	1 Blu + Rosso	Not in use	1
				47990 -	48014	DT test: trigger rate 4 Hz, PMT 560 V	25	10000	1 Blu + Rosso	Not in use	1
				48015 -	48054	DT test: trigger rate 2 Hz, PMT 555 V	40	16000	1 Blu + Rosso	Not in use	1
		2024-02-10 14:55:57	2024-02-15 13:07:13	47209 -	47981	Bkg + Daily Calibrations	772	308800	5 Blu + Rosso	Not in use	-
or sentine	el	2024-02-15 15:35:22	2024-03-05 9:33	48055 -	50891	Bkg + Daily Calibrations	2836	1134400	5 Blu + Rosso	Not in use	
		2024-03-17 16:20:14	2024-03-18 15:14	52664 -	52808	Bkg + Daily Calibrations	144	57600	5 Blu + Rosso	Not in use	
		2024-03-18 15:42:55	2024-03-19 15:19:04	52816 -	52874	Stability + Daily Calibrations- LOW Gas Flow : 1 I/h	58	23200	1 Blu + Rosso	Not in use	1
		2024-03-19 16:46:18	1	52882 -		Stability + Daily Calibrations- Gas Flow : 5 l/h	121	48400	5 Blu + Rosso	Not in use	Air
					53003	Stability + Daily Calibrations- Gas Flow : 5 I/h			5 Blu + Rosso	Not in use	Air
		2024-03-21 17:51:00		53004 -	53109	Stability + Daily Calibrations- LOW Gas Flow : 1 l/h	105	42000	1,Blu + Rosso	Not in use	Air
ONE	YES	2024-03-23 18:20:34	2024-03-26 9:41:19	53110 -	53502	Bkg + Daily Calibrations	392	156800 5+20	Blu + Rosso	Not in use	Air
ONE	YES	2024-03-29 10:01:40	2024-04-02 10:02:22	53707 -	54403	Bkg + Daily Calibrations	696	278400 5+20	Blu + Rosso + RADON	Not in use	Air
		2024-04-02 10:42:22	1	54411 -	54502	Stability + Daily Calibrations- HIGH recirculation 40 l/h	91	36400 5+40	Blu + Rosso + RADON	Not in use	Air
ONE	YES	2024-04-04 8:31:50	2024-04-08 8:26:06	54503 -	55093	Bkg + Daily Calibrations	590	236000 5+40	Blu + Rosso + RADON	Not in use	Air
0		2024-04-08 13:00:06	5	55101 -	56883	Bkg + Daily Calibrations - Low Gain - Low Drift	1782	712800 5+40	Blu + Rosso + RADON	Not in use	Air
									1		1



1. Alpha frequency in each batch

Batch	Runs	N runs	N matched alphas	N alphas per run	
1	40919-42848	1820	52800	29.01	
2	45259-46628	1370	8529	6.23	
3	53110-53502	390	2315	5.94	
4	54503-55093	591	3050	5.16	
5	53707-54403	531			
(6)	48055-50891	-	-	-	



High number of alphas in December, before oxygen and humidity filters were installed (correct?)



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High number of alphas in December, before oxygen and humidity filters were installed (correct?)

For all analysis, I don't have many cuts as I only saved <u>matched and alpha-PID</u> signals. \Rightarrow You

can assume near-perfect selection.

Given the statistics, I'll group the data as *before (batch 1) and after Christmas (batch 2-5)*.



2. 3D lengths distributions

2. **3D lengths distributions**

...comparing with Pinci's slides (2D alpha lengths)

The Radon Contamination



- So, a Rn contamination would produce:
 - 3 alphas:
 - ²²²Rn -> 5.590 MeV (about 43 mm)
 - ²¹⁸Po -> 6.115 MeV (about 50 mm)
 - ²¹⁴Po -> 7.833 MeV (about 73 mm)
 - 2 betas
 - a lot of gammas from 50 keV to 2200 keV

@Pinci, do we know the precision and source of these numbers. Iaminia has slightly bighes values for these alphas

Flaminia has slightly higher values for these alphas.





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C/GNO G S Experiment S I

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... Comparing batches:

- before (1) vs after (2-5) Rn filters
- Normalized to 1
- Statistics: 40k vs 10k samples





2. **3D lengths distributions**

... Comparing batches:

- before (1) vs after (2-5) Rn filters
- Normalized to 1
- Statistics: 40k vs 10k samples
- → The **Rn peaks** are in the same positions
 - 🕨 🛛 Good consistency 🔽
- → The relative quantity of Rn alphas in batch 2-5 reduced, highlighting other peaks.
 - Another proof of presence of Rn and posterior effectiveness of filters





David M.



2. **3D lengths distributions**

... Caveats:

 \rightarrow Actually, there is some discrepancy in my numbers and Pinci's \Rightarrow mine are systematically bigger.



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3D lengths distributions

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 - If I were to remove the ⁵⁵Fe length (= 63 samples = 4.6mm)...







2. 3D lengths distributions

* NB: This correction was also applied in the tilted cosmic (flux)

studies (https://agenda.infn.it/event/38654/contributions/217319/)

... Caveats:

- \rightarrow Actually, there is some discrepancy in my numbers and Pinci's \Rightarrow mine are systematically bigger.
 - Likely due to the fact that I'm not subtracting the <u>**"minimum temporal signal"**</u> from the measured ToTs.
 - If I were to remove the ⁵⁵Fe length (= 63 samples = 4.6mm)...
 - The results would be even closer.





- 222Rn -> 5.590 MeV (about 43 mm)
- 218Po -> 6.115 MeV (about 50 mm)
- 214Po -> 7.833 MeV (about 73 mm)

^{- 3} alphas:

Experiment S I

2. 3D lengths distributions - conclusion

... Conclusions:

2. 3D lengths distributions - conclusion

... Conclusions:

• Results are very interesting because, *unlike the energy that saturates* and spoils the spectrum, <u>the length of the</u> <u>tracks can be quite precise</u>, and indeed we can see, also with a **good resolution**!



- → We are <u>clearly in the presence of Rn</u> as we see the 3 contributions.
- → A simulation + digitization would help my analysis, mostly on the accuracy of the CMOS:analyzer class, since it was initially optimized only for ERs.
- → PMT simulation would be even better to test full 3D analysis!





3. dE/dx vs energy and vs length

... a quick look:

C/GNO G S

3. dE/dx vs energy and vs length

... a quick look:



- → Lines likely correspond to the **3 alphas observed**.
 - They are not points because of saturation!

-14

12 full_length 12

• Could saturation be studied from here?

CXGNO G S Experiment S I

4. Angles

4. Angles

→ Reference frame



Figure 1.5: System of coordinates and angles used in the analysis of LIME data.



4. Angles

→ Reference frame



Figure 1.5: System of coordinates and angles used in the analysis of LIME data.



4. Angles

- → Cuts based on the alpha 3D lengths:
 - > 4 cm: Rn alphas

<	4	cm:	Everyt	hing	else
---	---	-----	--------	------	------

		A	ZGNO	G	S
		full_length {full_length >	0 && full_length <	< 12}	
				hte	mp
	E			Entries	50808
	1800	6		Mean	4.583
	F	1		Std Dev	1.379
	1600	1			
1	1400				
	1400E	11			
	1200	11			
	1200				
	1000				
	E				
	800-				
	=	110			
	600	111			
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	, E	man	W. Newsman and a		
	0	2 4 6	8 10	12	
				tuli long	th

4. Angles

- → Cuts based on the alpha 3D lengths:
 - > 4 cm: Rn alphas







→ Before Christmas/Rn filters

- → Rn alphas have symmetric distribution of angles.
 - Makes sense since emission is random
- → Non-Rn alphas have preferential direction towards GEM
 - ♦ Perhaps coming from cathode? ⇒

To investigate with absolute Z



... Comparing before and after Rn filters:

 \rightarrow

4. Angles

- Cuts based on the alpha 3D lengths:
 - > 4 cm: Rn alphas

< 4 cm: Everything else</p>



... Comparing before and after Rn filters:



 \rightarrow

4. Angles

- Cuts based on the alpha 3D lengths:
 - > 4 cm: Rn alphas





... Comparing before and after Rn filters:



4. Angles

...We can also look at the <u>distribution of angles</u>





4. Angles

...We can also look at the <u>distribution of angles</u>

full_length:Z_angle {full_length > 4 && full_length < 12 && pmt_direction != 0}



full_length:Z_angle {full_length > 0 && full_length < 4 && pmt_direction != 0}

→ Alphas at 7 cm more towards GEMs. Po charged daughters that drift towards cathode?



4. Angles

..We can also look at XY angle, meaning <u>CMOS-only analysis*</u>

*Xenon nT style





4. Angles

..We can also look at XY angle, meaning <u>CMOS-only analysis*</u>



*Xenon nT style



- → This shows a much greater amount of track going downwards, which could be from the resistors?
 - To be confirmed with position dependent cuts...



G S