

# OPERA HIGHLIGHTS

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On behalf of

The OPERA collaboration





# Outline

- OPERA
  - physics goal
  - detector
  - physics results
    - Oscillation
    - Neutrino speed
  - conclusion



# The OPERA Collaboration

## 160 physicists, 30 institutions, 11 countries

**Belgium**  
IIHE-ULB Brussels



**Italy**  
LNGS Assergi  
Bari



**Korea**  
Jinju



**Croatia**  
IRB Zagreb



**Bologna**  
LNF Frascati  
L'Aquila  
Naples  
Padova  
Rome  
Salerno

**Russia**  
INR RAS Moscow  
LPI RAS Moscow  
ITEP Moscow  
SINP MSU Moscow  
JINR Dubna



**France**  
LAPP Annecy  
IPNL Lyon  
IPHC Strasbourg



**Germany**  
Hamburg



**Japan**  
Aichi  
Toho  
Kobe  
Nagoya  
Utsunomiya



**Switzerland**  
Bern  
ETH Zurich



**Israel**  
Technion Haifa



**Turkey**  
METU Ankara



<http://operaweb.lngs.infn.it/scientists/?lang=en>

## Neutrino propagation

- Neutrinos are created as weak eigenstate, but propagate as mass eigenstate

$$|\nu\rangle = \begin{pmatrix} |\nu_e\rangle \\ |\nu_\mu\rangle \\ |\nu_\tau\rangle \end{pmatrix} \qquad |\nu'\rangle = \begin{pmatrix} |\nu_1\rangle \\ |\nu_2\rangle \\ |\nu_3\rangle \end{pmatrix}$$

- When we detect them after their travel, we can detect their current weak eigenstate

$$|\nu'(t)\rangle = e^{-iH't} |\nu'(0)\rangle \qquad |\nu(t)\rangle ?$$

- This leads to a possible flavour change of neutrinos during their travel, which was observed for solar (Davis, SNO) and atmospheric neutrinos (Super KAMIOKANDE)

# Neutrino oscillation – theoretical description

$$\begin{pmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{pmatrix} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \\ \nu_3 \end{pmatrix}$$

3 angles  $\theta_{ij}$ ; the fraction of  $\nu_1, \nu_2$  or  $\nu_3$  in a flavor eigenstate  
 1 phase  $\delta$ ; CP violation (for Dirac neutrinos)  
 atmospheric (23) and solar (12)

$$\begin{matrix} c_{ij} = \cos\theta_{ij} \\ s_{ij} = \sin\theta_{ij} \end{matrix}$$

Oscillation probability in the three neutrinos scenario

$$P(\nu_\ell \rightarrow \nu_{\ell'}) = \left| \sum_i U_{\ell i} U_{\ell' i}^* e^{-i(m_i^2/2E)L} \right|^2$$

$$= \sum_i |U_{\ell i} U_{\ell' i}^*|^2 + \Re \sum_i \sum_{j \neq i} U_{\ell i} U_{\ell' i}^* U_{\ell j}^* U_{\ell' j} e^{i \frac{|m_i^2 - m_j^2| L}{2E}}$$

$\frac{|m_i^2 - m_j^2| L}{2E}$

$\Delta m^2$

$L$

$E$

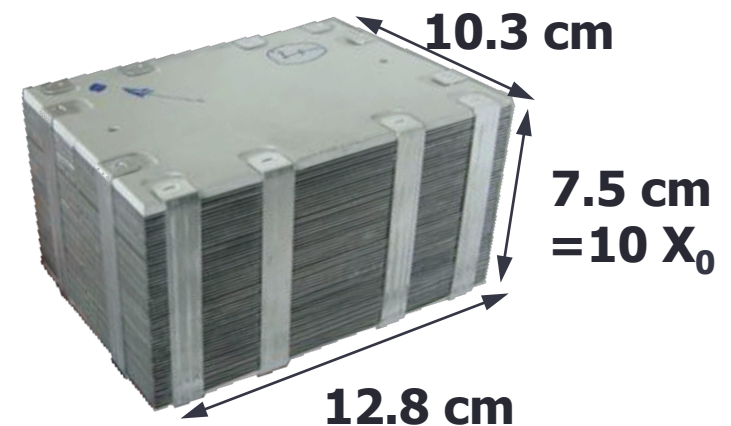
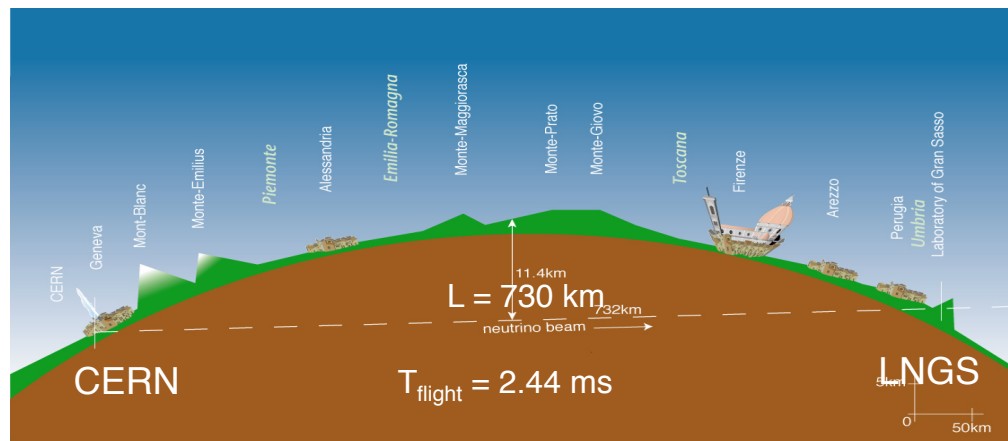
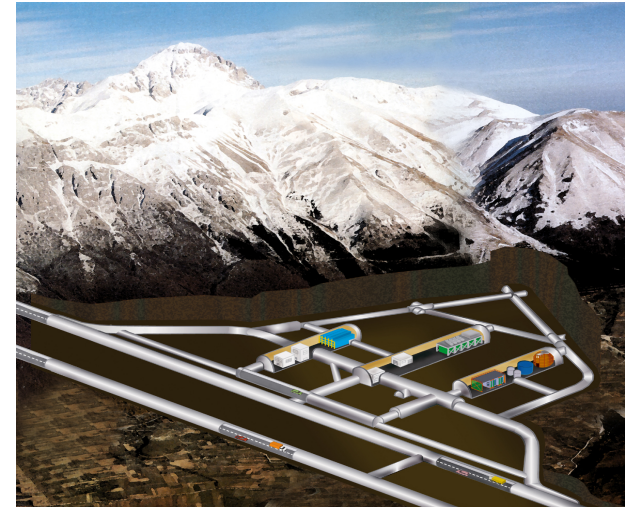
2 mass differences  $\Delta m^2_{12} \ll \Delta m^2_{23}$

2 neutrino approximation:

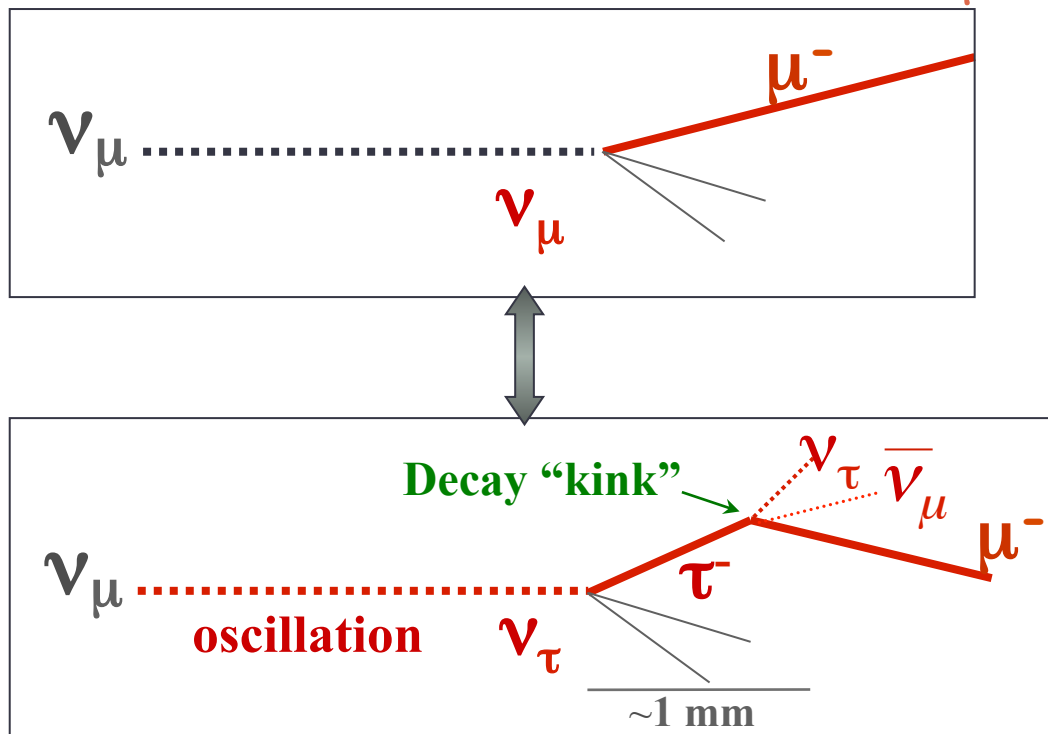
$$P(\nu_\mu \rightarrow \nu_\tau) \sim \sin^2 2\theta_{23} \underbrace{\cos^4 \theta_{13}}_{\sim 1} \sin^2 (\Delta m^2_{23} L / 4E)$$

# Physics goal of Opera

- First detection of neutrino oscillation in direct appearance mode via the  $\nu_\mu$  to  $\nu_\tau$  channel.
- Detection of the  $\nu_\tau$  by the  $\tau$  decay topology
- Low cross section leads to a large mass
  - Emulsion Cloud Chamber: Lead brick, interleaved with emulsion
- Low background requirement: LNGS
- Neutrino beam optimized for detection



# The direct probe for $\nu_\mu \rightarrow \nu_\tau$ appearance



leptonic, 1 + 3-prong decay modes

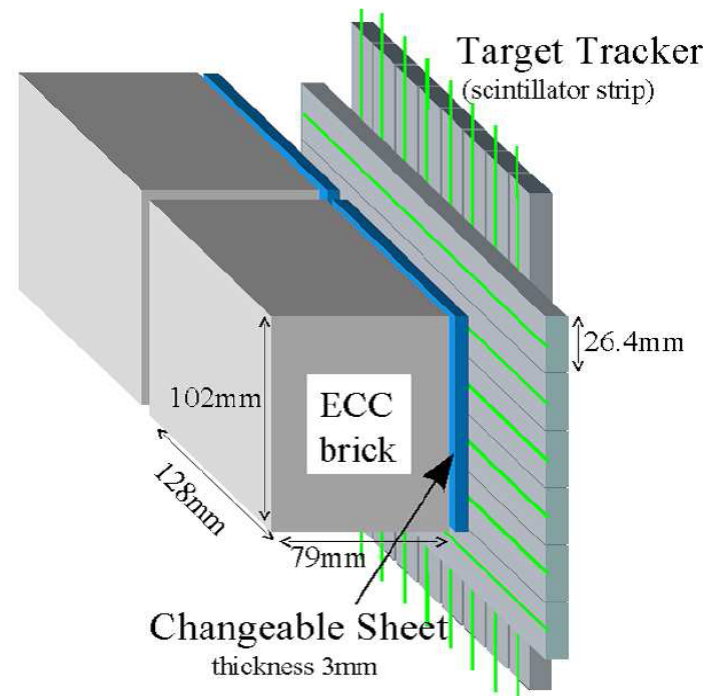
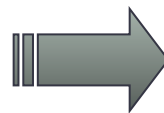
Two conflicting requirements:

➤ **Large mass (low Xsection)**

➤ **High spatial resolution**

➡ signal selection

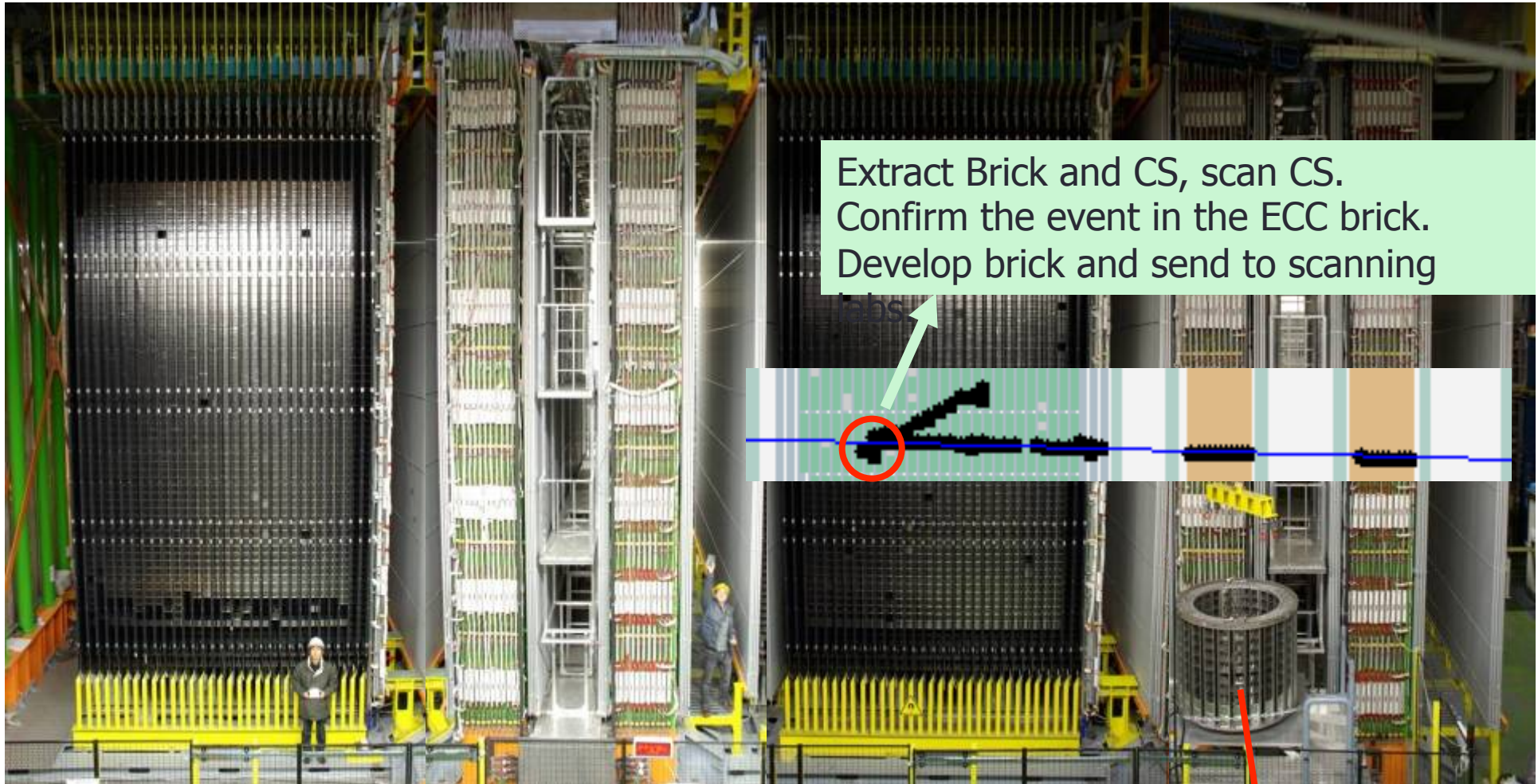
➡ background rejection



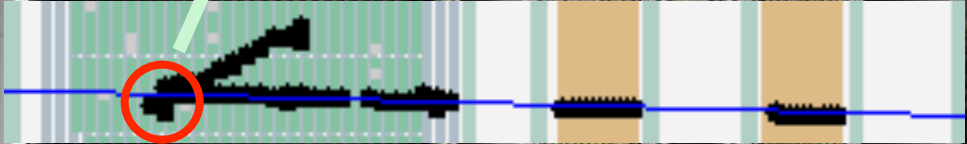
Target:  $\sim 1250$  tons,  
 $22.5E^{19}$  pot during 5 years

- $>20000$  neutrino interactions
- $\sim 100$   $\nu_\tau$  interactions
- $\sim 10$   $\nu_\tau$  identified
- $<1$  background event

# The OPERA detector



Extract Brick and CS, scan CS.  
Confirm the event in the ECC brick.  
Develop brick and send to scanning  
labs



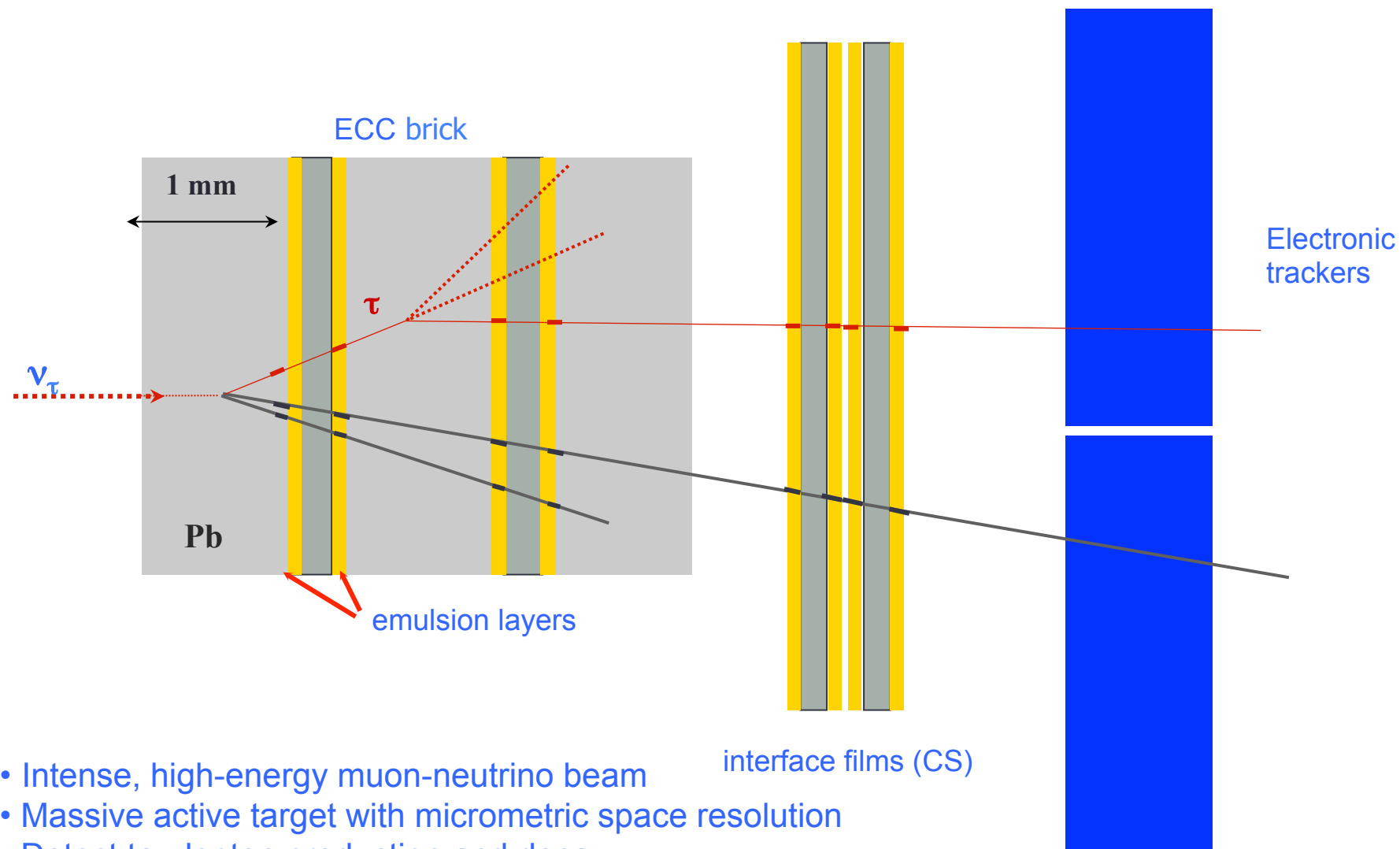
Target area  
(ECC + CS + TT)

Muon spectrometer  
(Magnet+RPC+PT)

Brick Manipulator System



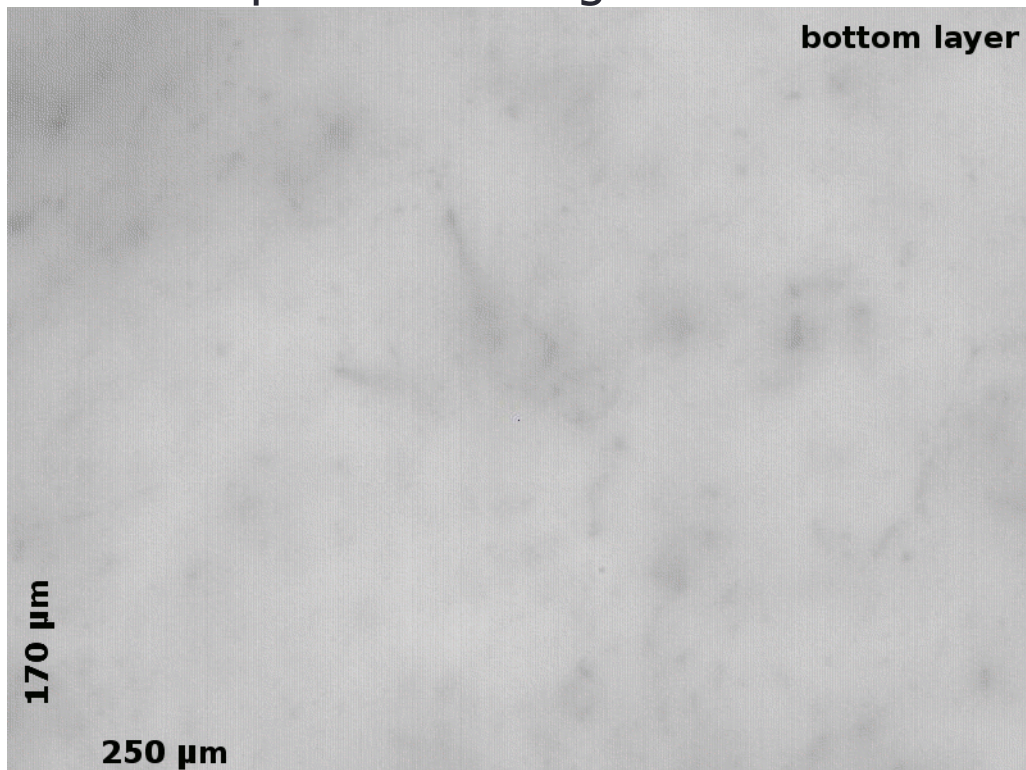
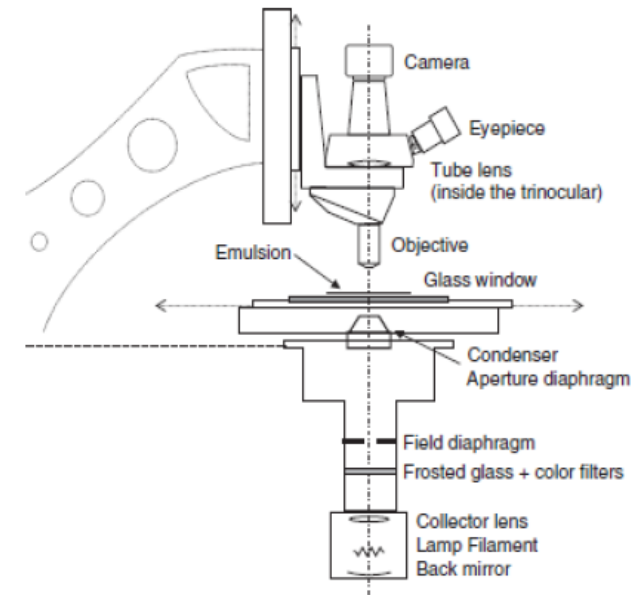
# ECC combined with electronic detector



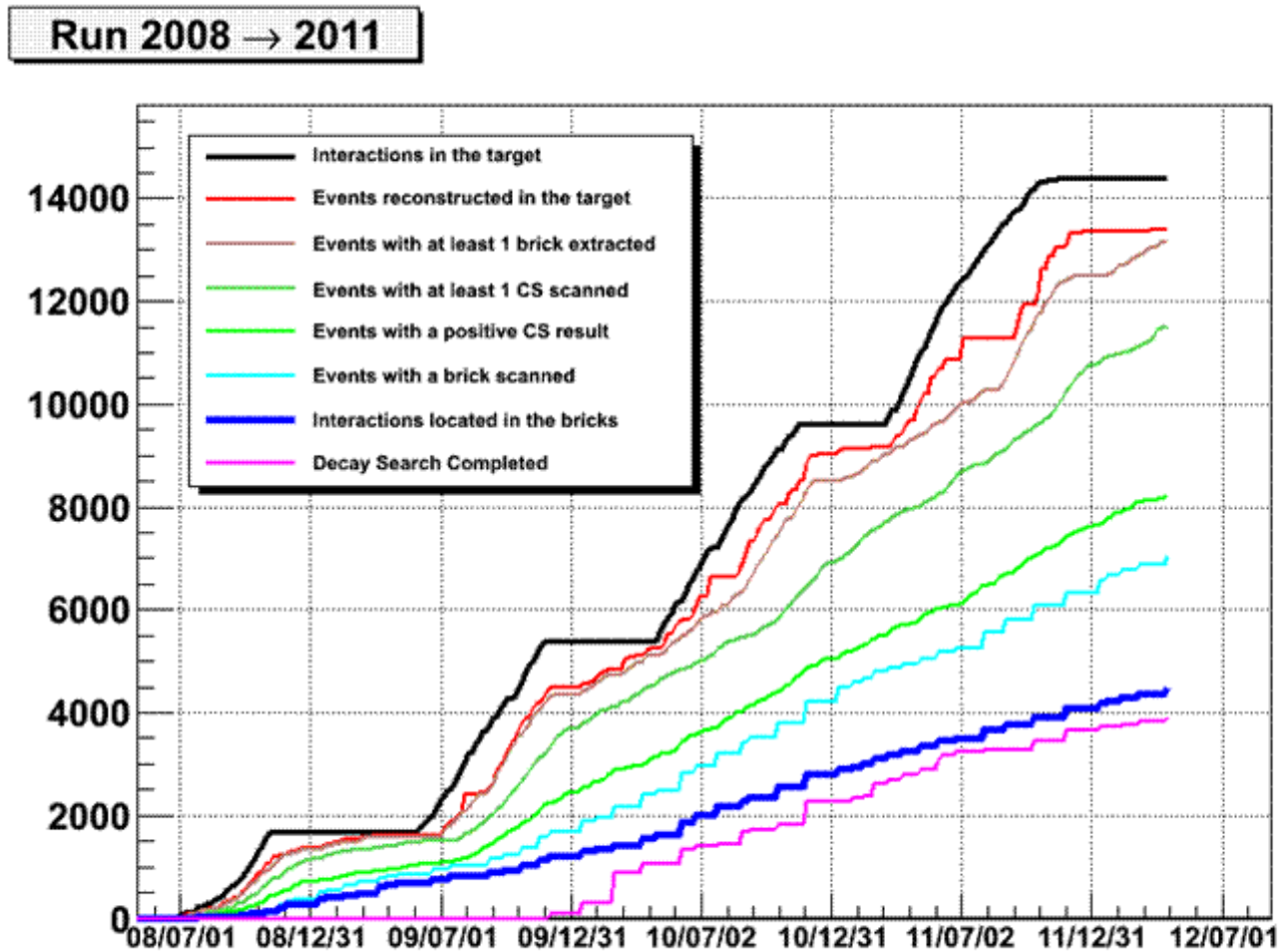
- Intense, high-energy muon-neutrino beam
- Massive active target with micrometric space resolution
- Detect tau-lepton production and decay
- Use electronic detectors to provide “time resolution” to the emulsions and preselect the interaction region

# Emulsion Scanning

- Scanning Labs in Europe and Japan
  - 50% sharing
  - European CS scanning at LNGS
  - Data storage of scanned data in DB
  - Example of neutrino vertex
  - Example of scanning station at Bern

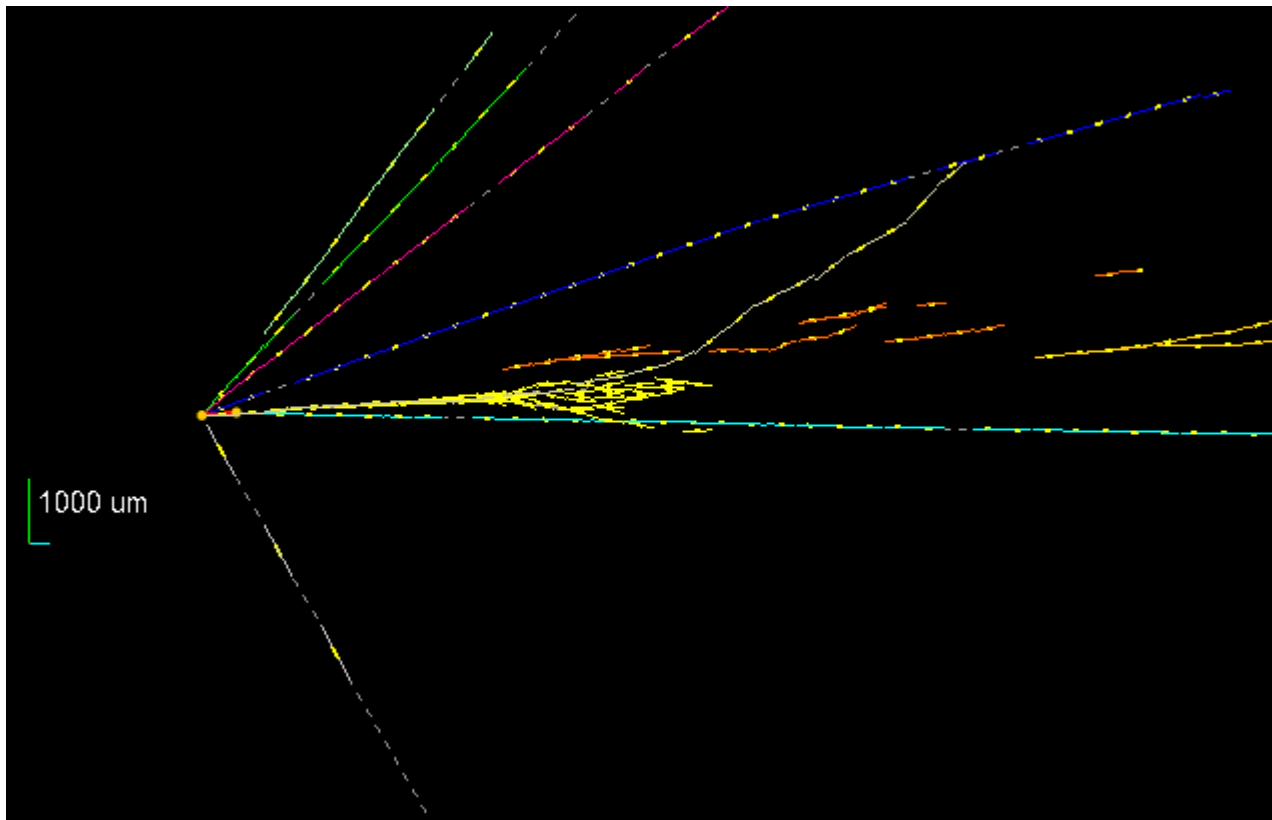


# Event collection

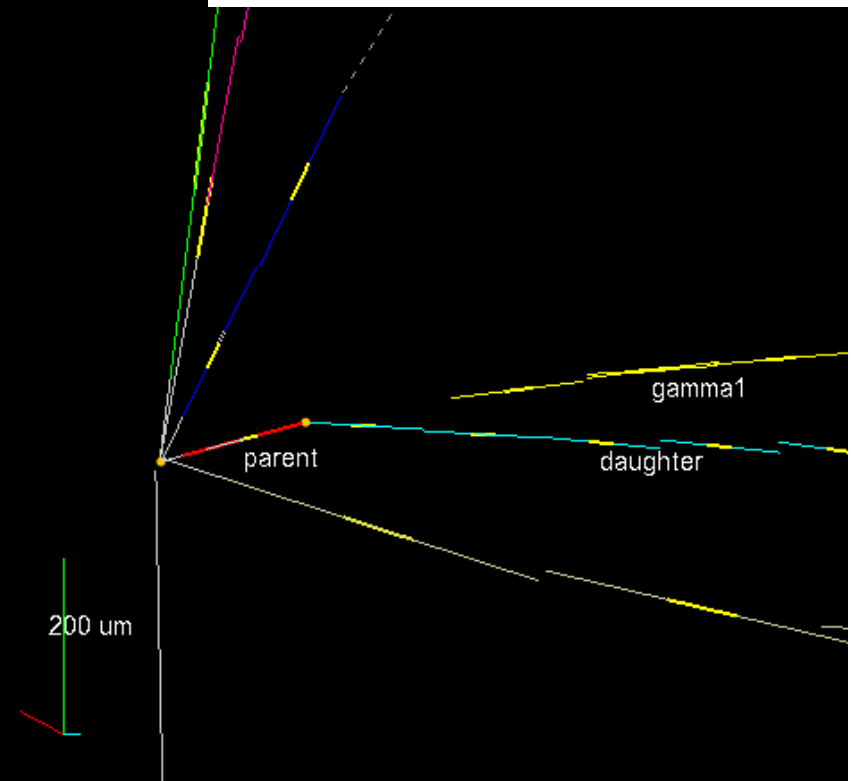


Run 2012 is started, currently “bunched beam” mode for  $v$  velocity measurement

# First $\nu_\tau$ candidate



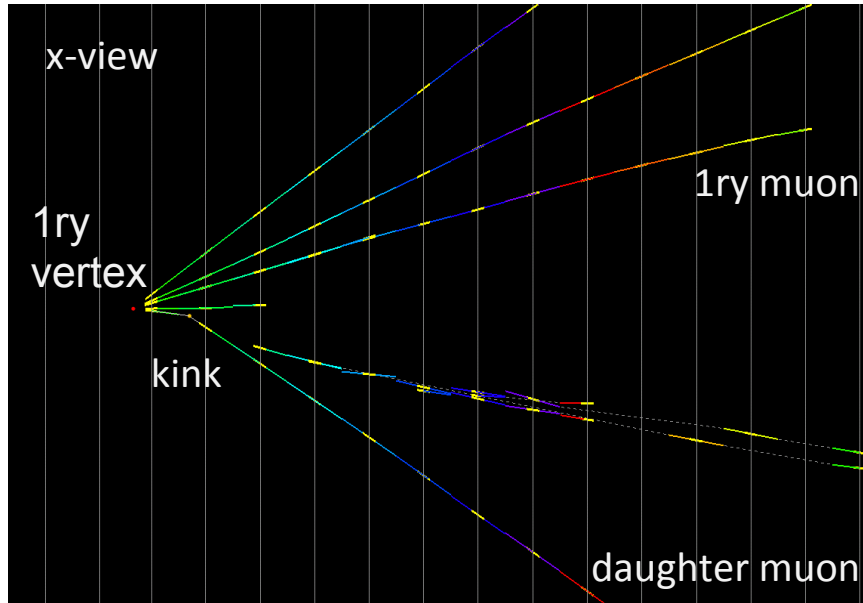
- PLB 691 (2010) 138
- No  $\mu$  found in spectrometer
- All tracks followed till stopping point or reinteraction (no  $\mu$  found)
- In agreement with expectation
- Compatible with a hadronic  $\tau$  decay:  $\tau^- \rightarrow h^- (n\pi^0) \nu_\tau$
- Full update for Neutrino 2012
  - New Candidates
  - New Background estimation



## About the first $\nu_\tau$ event

- All cuts defined at the proposal time applied, event passed the selection.
- Current expectation 1.7 events in all tau channels (0.5 in singleprong)
- Propability for a background event 5%, expected background so far only 0.16 (0.05) for  $4.9 \times 10^{19}$  pot
- What else can we do except for  $\tau$  decay
  - Other rare events: charm
  - electron neutrinos
  - Timing of neutrino events

# Charm event

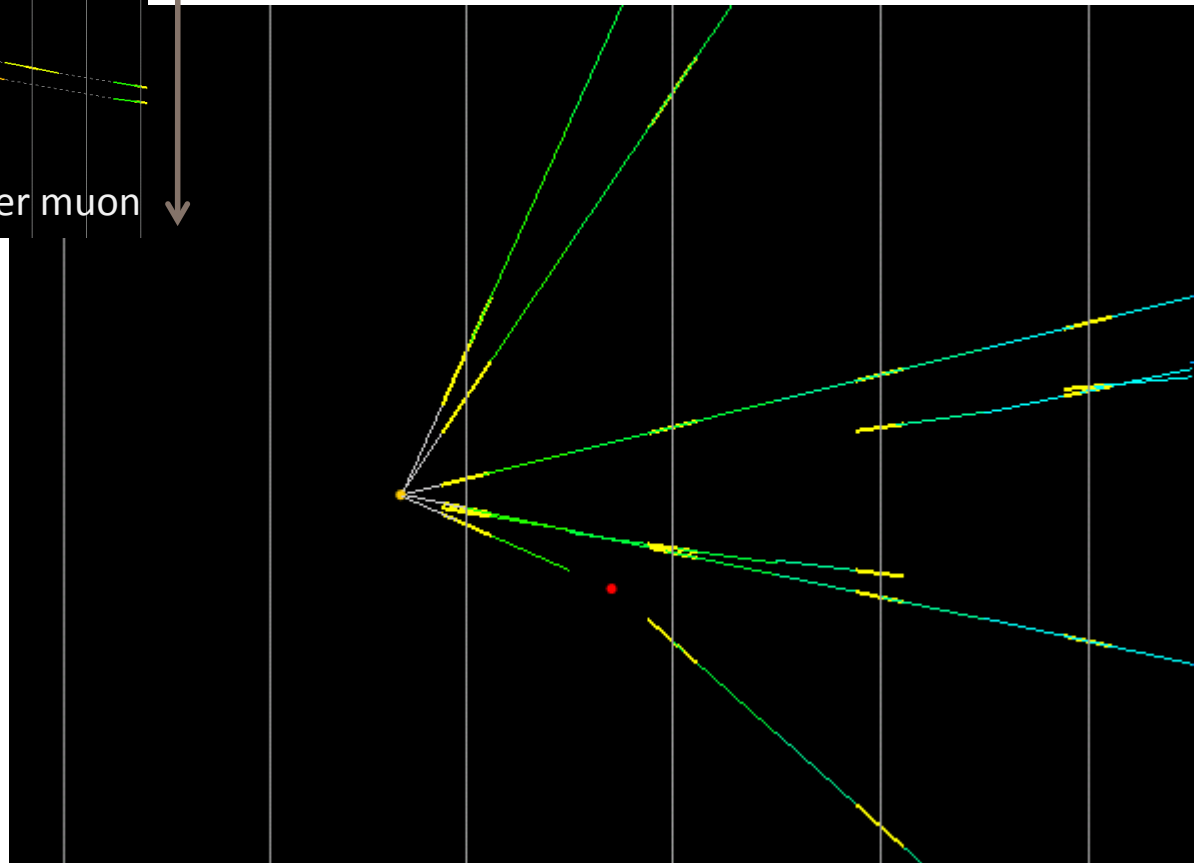


4 mm

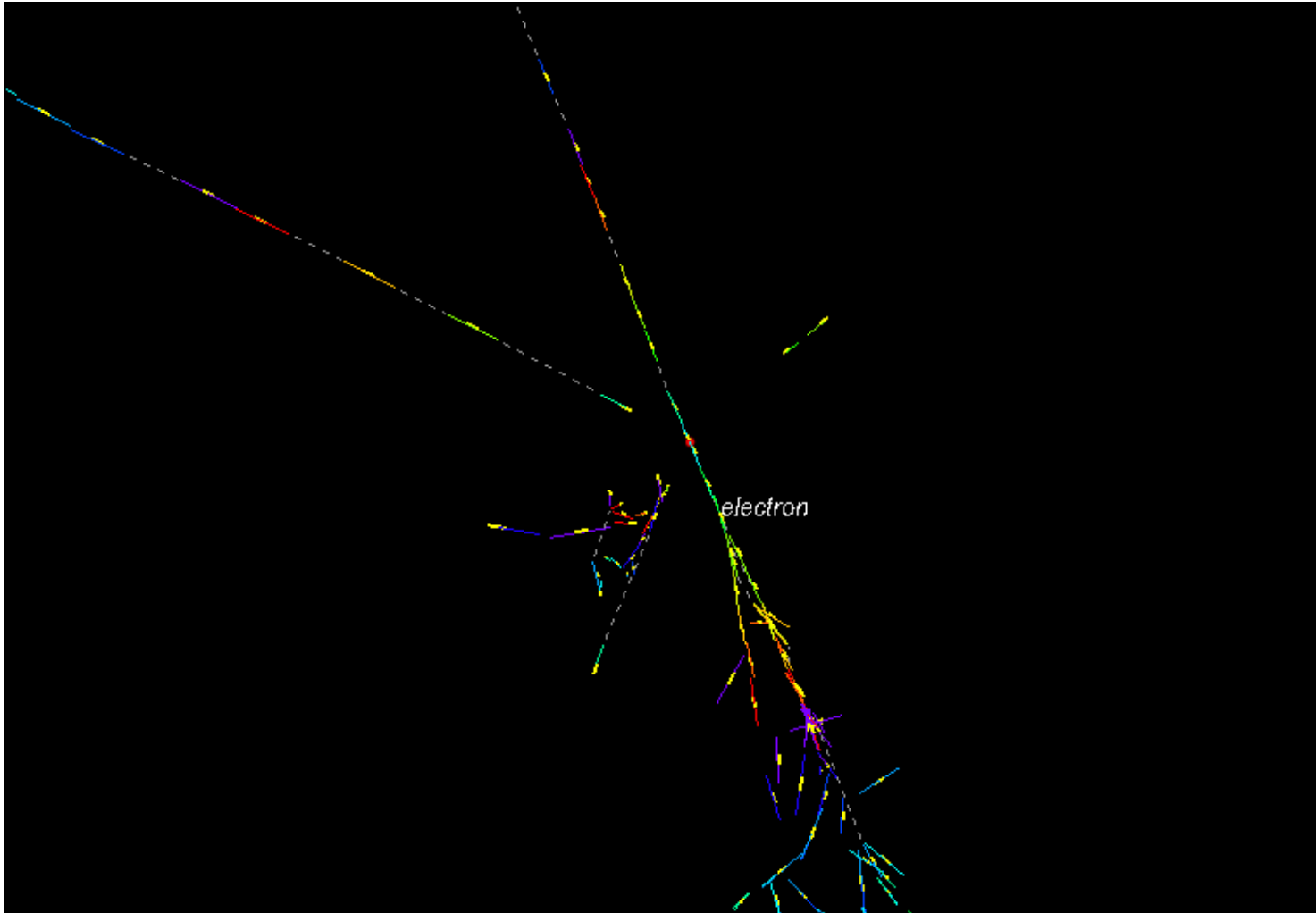
flight length: 1330 microns  
kink angle: 209 mrad  
IP of daughter: 262 microns  
daughter muon: 2.2 GeV/c  
decay  $P_t$ : 0.46 GeV/c

1.3 mm

Di muon candidate  
 $\mu^+$  and  $\mu^-$  in spectrometer



# $\nu_e$ candidate



# Neutrino velocity measurement

- Just a brief description of technics and concept, more details and new numbers will come at Neutrino 2012
- TOF: Time of flight  $\text{TOF}_\nu = t_B - t_A - t_{\text{delay}}$ 
  - $t_A$  – production time
  - $t_B$  - detection time
- Distance: between source (CNGS) and detector (Opera)
- Velocity: distance /  $\text{TOF}_\nu$
- Statistical analysis on neutrino events collected 2009-2011
- Special bunched beam mode end of 2011 and in may 2012 for direct comparison event by event



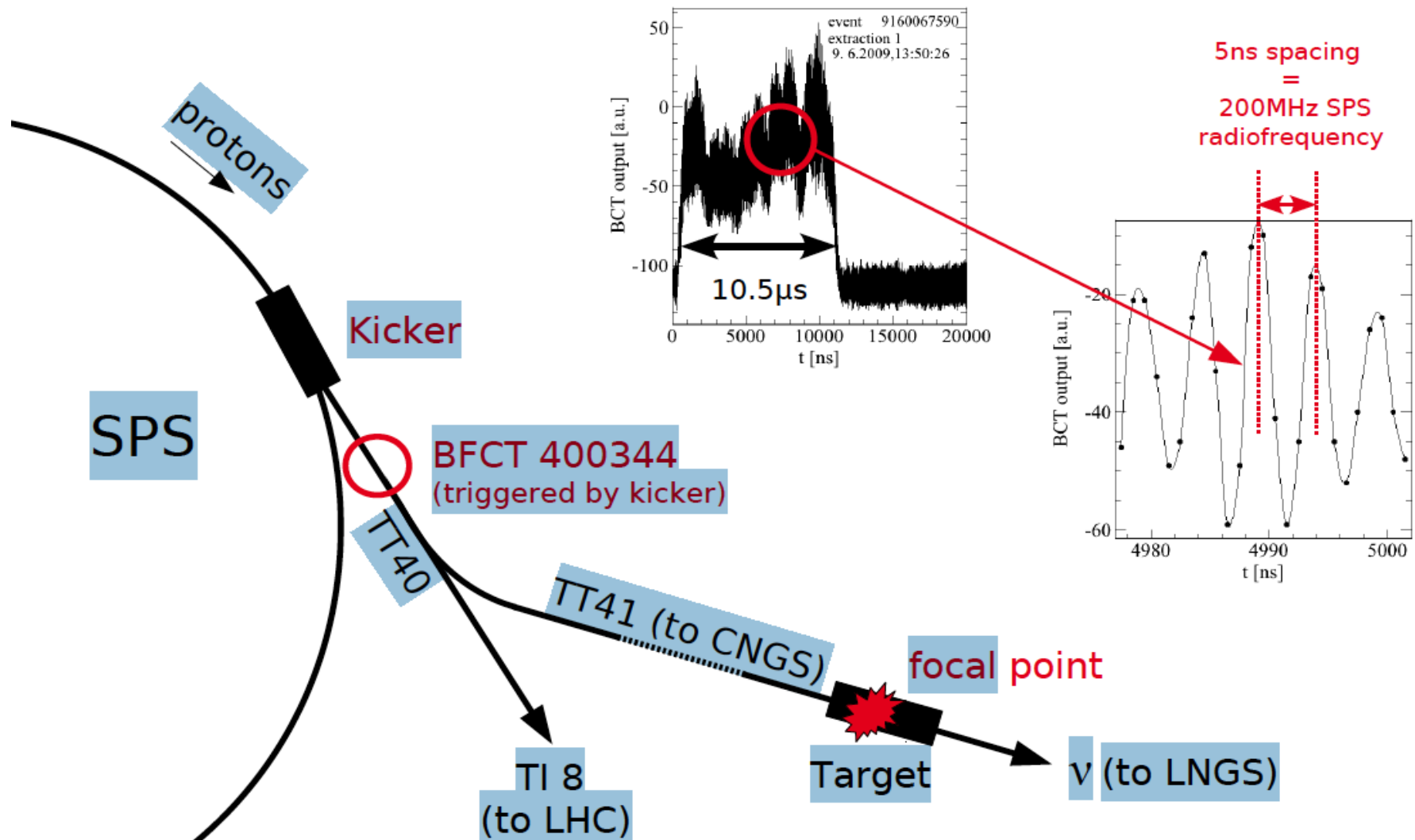
# Previous measurements

- 1979: FNAL (**Phys. Rev. Lett. 43 (1979) 1361**)
- short distance, 30 GeV  $\nu_\mu$ , comparison of  $\nu_\mu$  and  $\mu$  TOF
- $|v-c|/c \leq 4 \times 10^{-5}$
- 1988: SN1987A (**Phys. Lett. B 201 (1988) 353**)
- very long distance (168.000 light years), 10 MeV anti- $\nu_e$ ,
- comparison of  $\nu$  and photon arrival time (not SN mod.-dep.)
- $|v-c|/c \leq 2 \times 10^{-9}$
- 2007: MINOS (**Phys. Rev. D 76 (2007) 072005**)
- 730km distance,  $\sim 3$  GeV  $\nu_\mu$ , near detector comparison
- $(v-c)/c = (5.1 \pm 2.9) \times 10^{-5}$
- 2011: OPERA
- 730km distance,  $\sim 17$  GeV  $\nu_\mu$ , proton – BCT comparison

## CERN production timestamp creation $t_A$

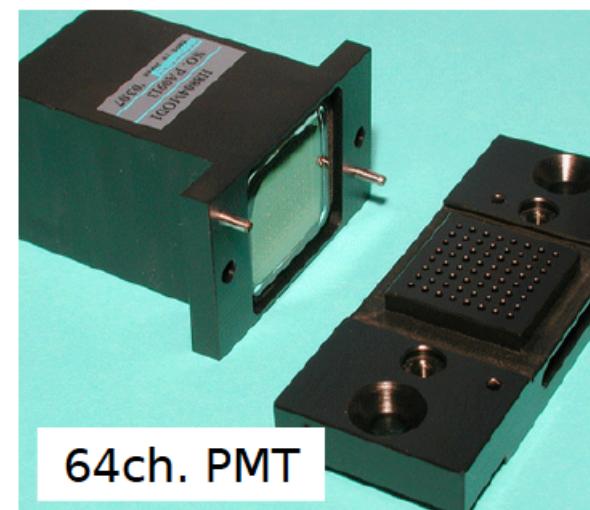
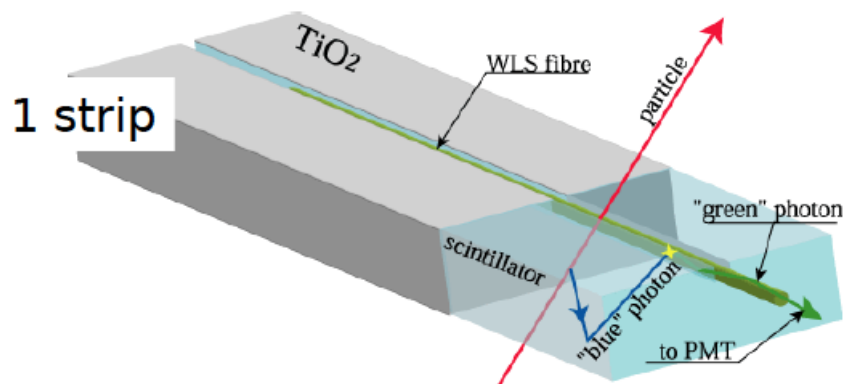
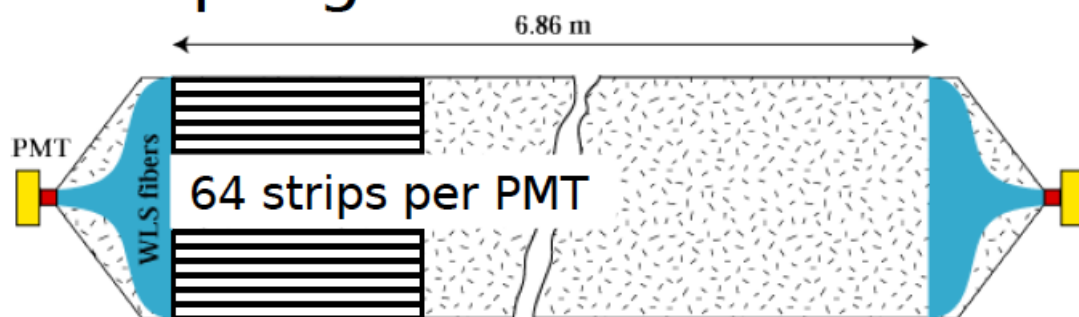
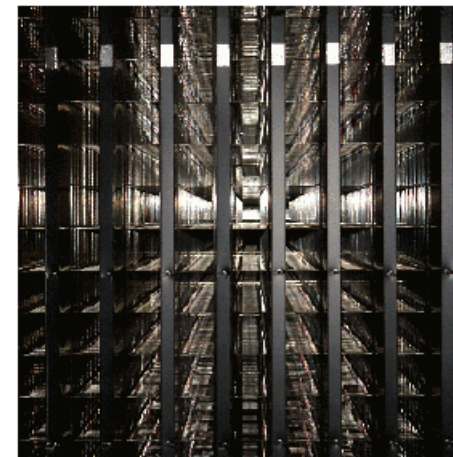
- protons accelerated in SPS@CERN
- protons extracted to TT40 transfer tunnel by kicker magnets (“global  $t_0$ ”)
- two extractions per SPS filling separated by 50ms, each 10.5 $\mu$ s long, consist of thousands of 2ns-long “bunches”
- proton distribution after kicker measured by fast beam current transformer (BFCT) and read-out by a waveform digitizer (WFD)
- protons focused on graphite target
- Proton interaction: subsequent pion (kaon) decay into neutrinos

# CERN production timestamp creation $t_A$



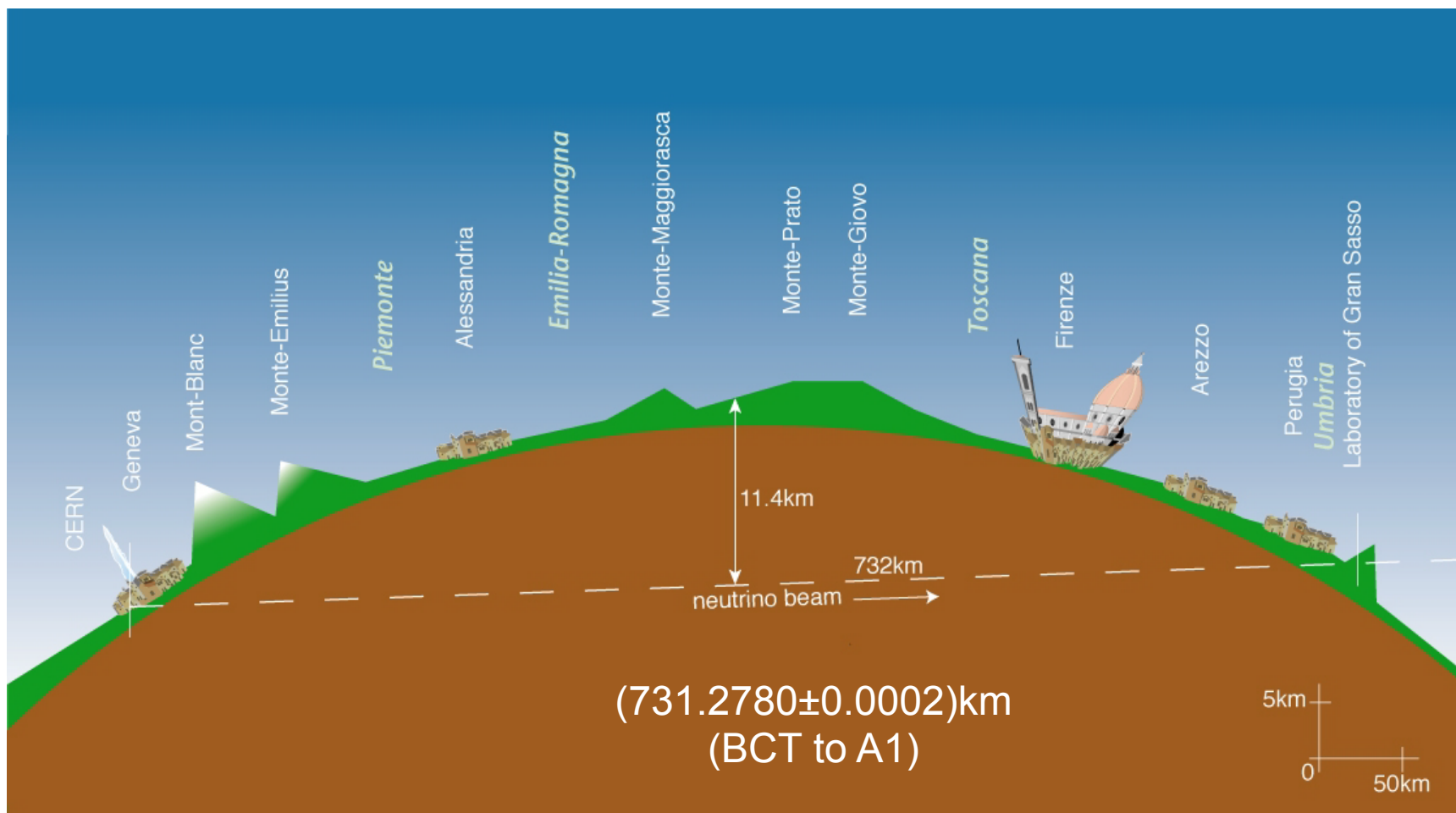
# OPERA detection time $t_B$ (LNGS)

- use plastic scintillators only
- first hit in target trackers is the stop signal

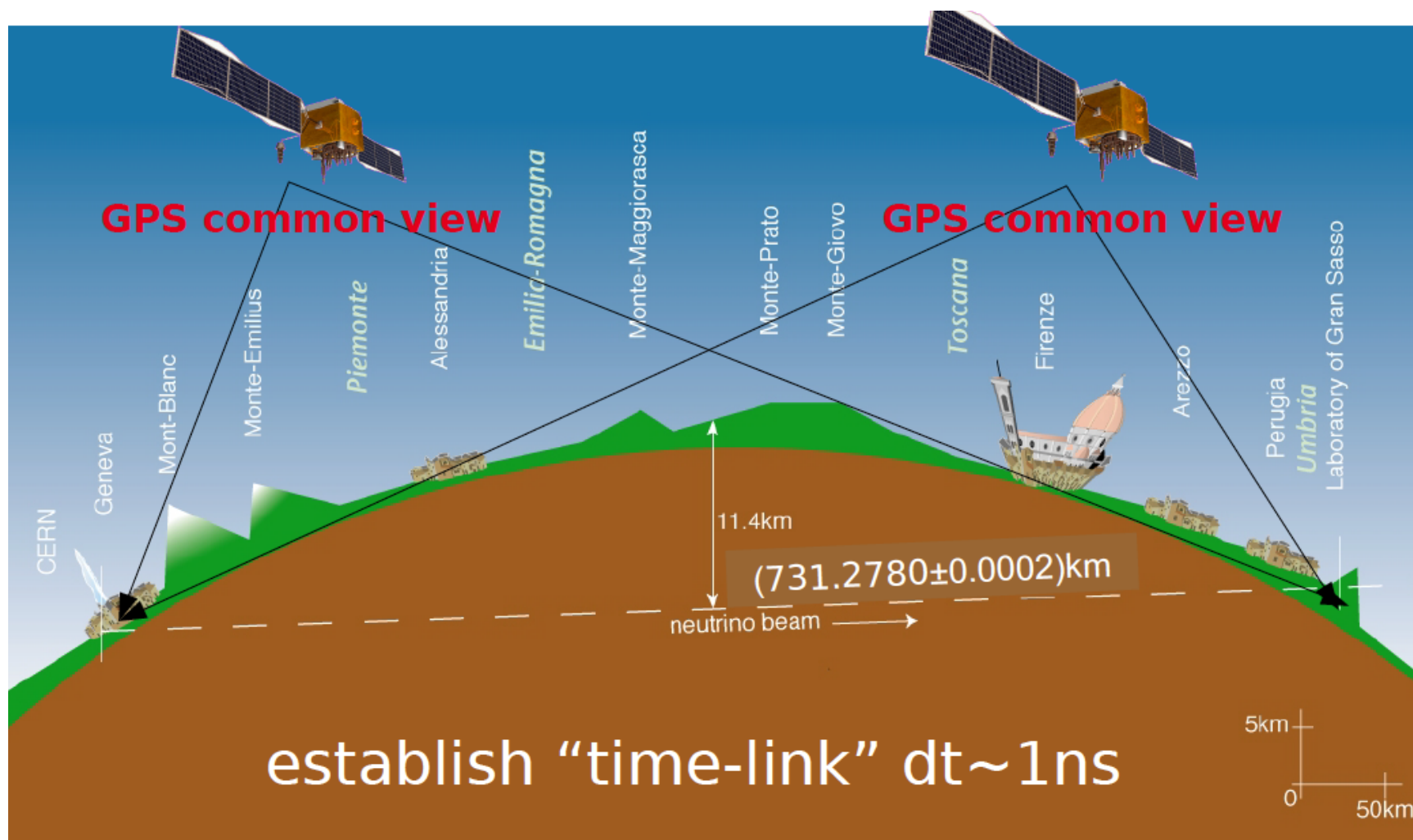


64ch. PMT

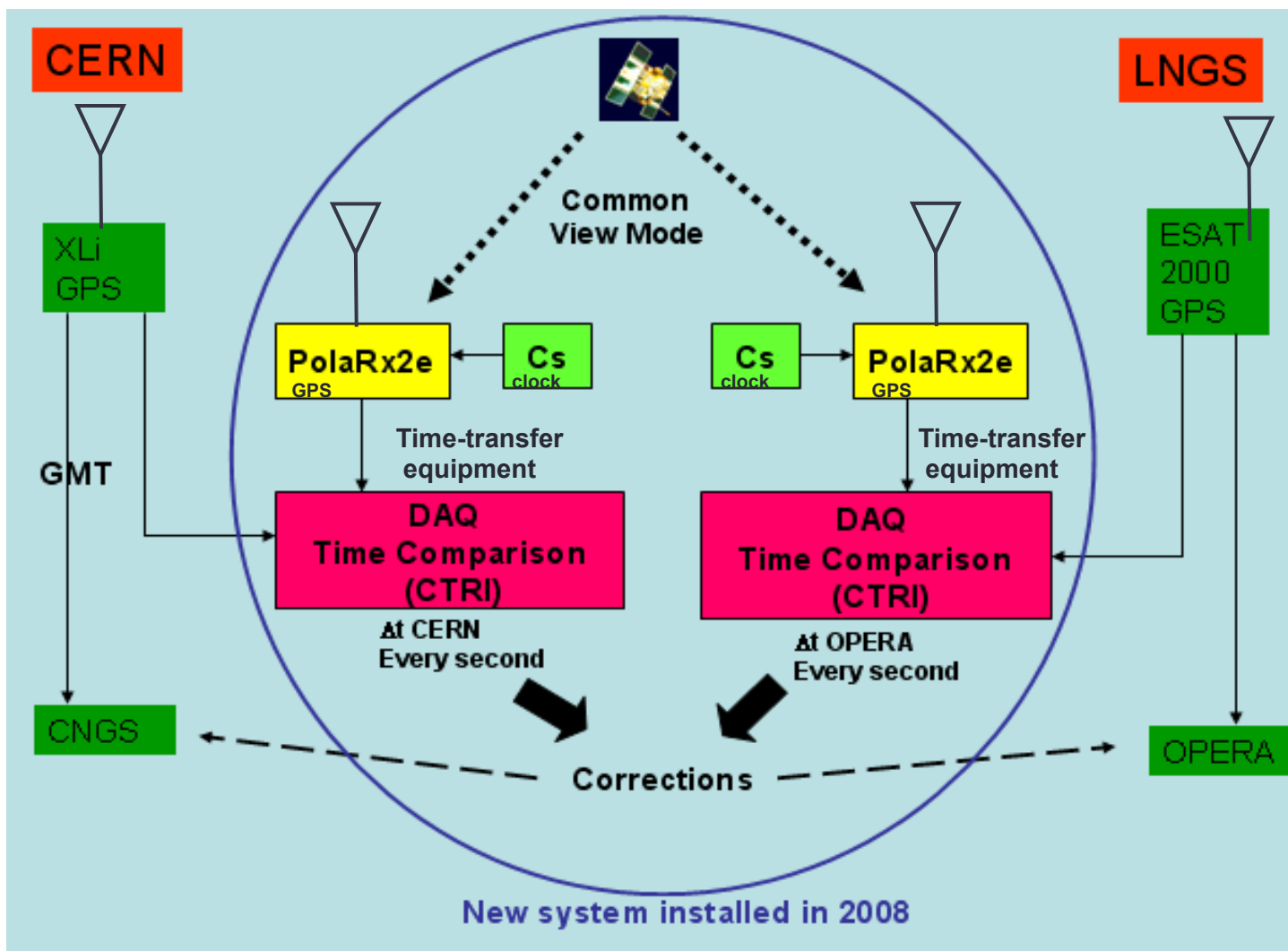
# Distance measurement CNGS – LNGS



# Time link between LNGS and CERN



# Time link between LNGS and CERN



## Data selection for statistical analysis Sep. 2011

- selection of neutrino events
- internal events (within fiducial volume, same as for the oscillation search): 7586
- external events (interactions in rock) with reconstructed 3D muon track: 8525 ( $\pm 2\text{ns}$  additional uncertainty)
- at least 4 GPS satellites in common view
- first hit not isolated in time or space
- if neutrino event passes selection:
  - select the corresponding BCT waveform
- **7235 internal and 7988 external events**

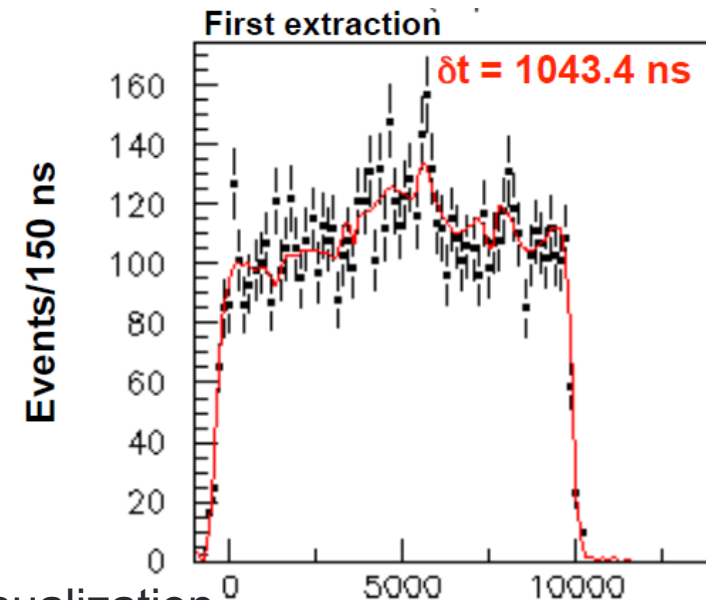
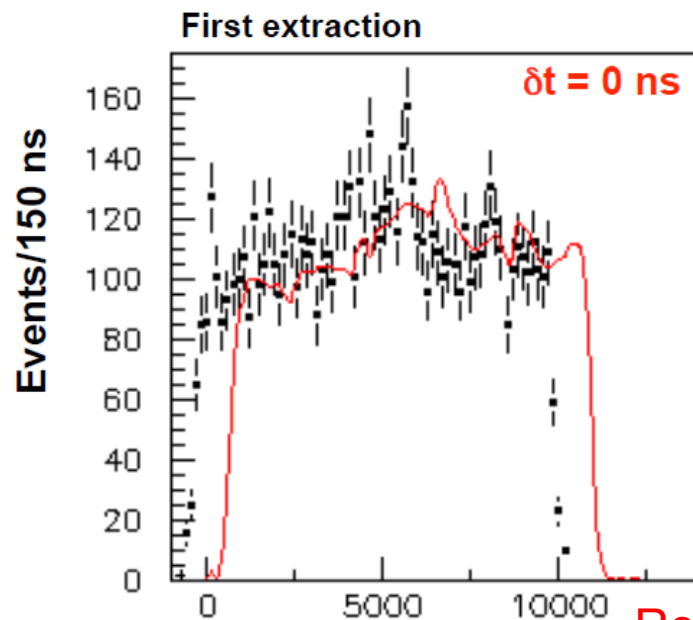


# Statistical analysis

- Build a likelihood from the sum of the waveforms
  - All waveforms/ only waveform of events which were selected

$$L_k(\delta t_k) = \prod_j w_k(t_j + \delta t_k) \quad k = 1, 2 \text{ extractions}$$

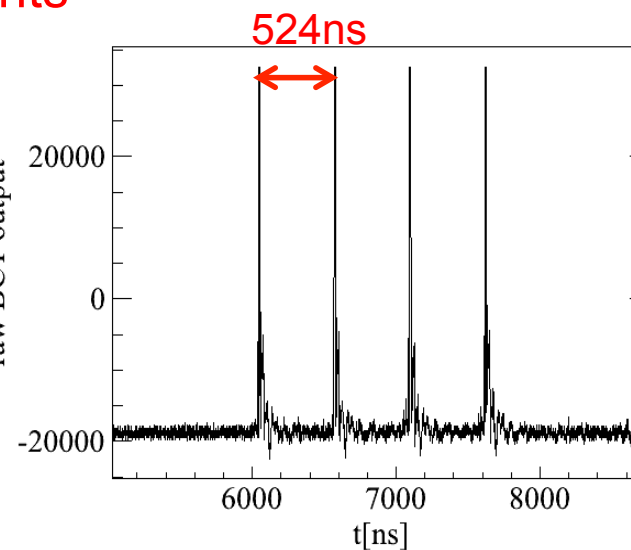
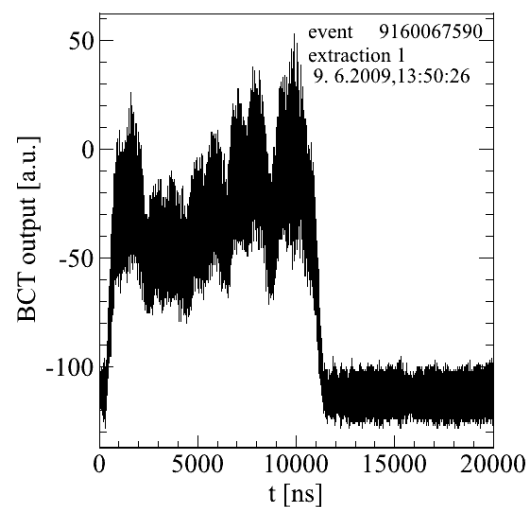
- measurement of delay  $\delta t$  for Opera



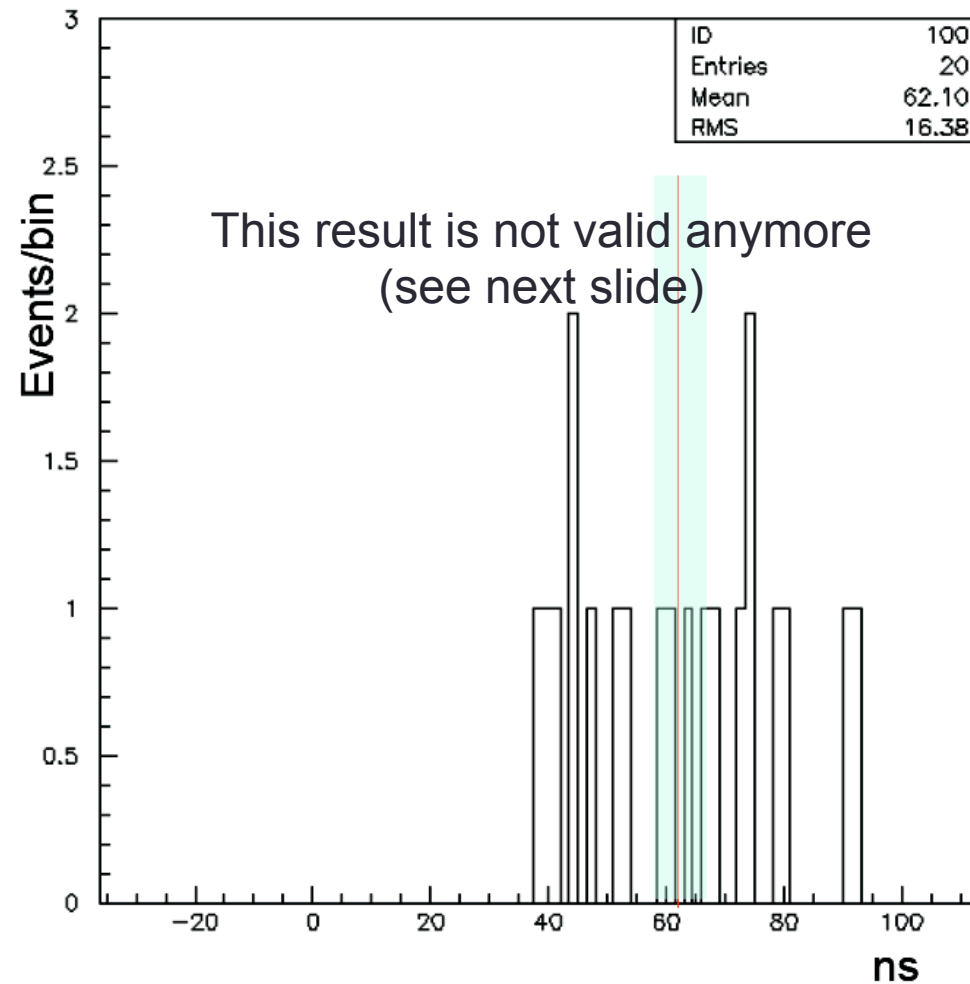
Red curve only visualization

# Bunched beam 2011

- bunched beam: instead of  $10.5\mu\text{s}$  extractions:
- 4 single, 3ns-wide bunches, separated by 524ns  
→ single-event TOF measurement!
- October 22 to November 6, 2011
- beam intensity lower than nominal ( $\sim 1/60$ )
- collected 35 events, same selection criteria,
- same delay corrections  
→ 14 external and 6 internal events



# Bunched beam 2011



# The story of $\delta t$

- “The OPERA Collaboration [...] has identified two issues that could significantly affect the reported result.
  - [...] the oscillator used to produce the events time-stamps [...]
  - [...] the connection of the optical fiber [...]” (Feb. 23<sup>rd</sup> 2012)
- Taken into account these correction, the time of flight could be in agreement with the speed of light.
- Currently a new bunched beam mode (different settings, to increase statistics, all LNGS experiments taking data)
- New preliminary results to be expected at Neutrino 2012

# Conclusions

- Opera is the first experiment that observed muon to tau neutrino oscillation appearance by identifying the event candidates by the topological characterization of the tau decay
- Opera is detecting other rare decays as charm decays and is detecting electron neutrino interactions
- Opera is able to measure the time of flight for neutrinos with a higher precision than any previous earth based experiment
- A lot of more exiting news will be released in 2 weeks time, stay tuned

Questions?