The Zoo of AGN: Experimental Tests of the General Relativity Theory and its pseudo-complex extension

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Outline

1. Present tests of the theory of General Relativity

2. The Pseudo-Complex Theory

3. Future tests of GR theories

1. Present tests of the theory of General Relativity dynamical tests in the low mass regime





2. The Pseudo-Complex Theory (Hess, Greiner 2007-2012)

$$X = X_R + I X_I$$
 with $I^2 = +1$

X: pseudo-complex number

$$R^{\mu\nu} - \frac{1}{2} g^{\mu\nu} R = -\frac{8\pi\kappa}{c^2} T^{\mu\nu} \sigma_{-}$$

$$\sigma_{-} = \frac{1}{2} (1 - I) \qquad \sigma_{-} \sigma_{+} = 0 \qquad \sigma_{-}^2 = \sigma_{-}$$

new Einstein equation energy represents repulsion

$$g_{00} = \frac{r^2 - 2mr + a^2 \cos^2 \theta + \frac{B}{2r}}{r^2 + a^2 \cos^2 \theta}$$

- g₀₀: metric tensor
- B: new pseudo-complex variable
- a: spin parameter

no coordinate singularity at r = 2m for a = 0

Effective potential in Pseudo-Complex theory



The Pseudo-Complex Theory vs. General Relativity BH spins becomes different



maxima: infall into BH

ISCO predictions: last stable orbits different



3. Future tests of the GR theories

VLTI interferometry

Eisenhauer et al. 2007-2012



X-ray spectroscopy in the strong limit

Boller2012, World Scientific Publishing



GRAVITY on VLT: R = 3 mas, $\Delta x = 10 \mu as$ the inner accretion flows around black holes emit significant amounts of X-rays

the radiation originates so close to the black hole that it allows for probing general relativity models

3. Future tests of GR theories

3.1 VLTI in the GC

Eisenhauer et al. 2007-2012

3.2 X-ray spectroscopy

Boller, Müller:

"Astronomical tests of General Relativity and its pseudo-complex extension" Springer 2012, special issue of the Symposium on "Astrophysics, Quark-Gluon-Plasmas, Biology", Makutsi, South Africa, in press

Schönenbach, Caspar, Hess, Boller, Müller, Greiner: "Possible experimental tests of the pseudo-complex General Relativity" "Journal of Astronomy and Astrophysics (IJAA)", 2012, submitted

3.3 X-ray and NIR timing analysis

papers as above

3.4 Tracing infall of matter

Boller, Müller, Räth, Dovciak, Svoboda GRAVITAS Yellow Book, 2011

3.1 VLTI in GC: GRAVITY

Sub-mm shadowing



3.1 VLTI in AGN: GRAVITY

sub-mm shadowing in M87



image different for PC and standard GR due to different z_G

3.2 X-ray spectroscopy

gravitational redshift as a function of distance



3.2 X-ray spectroscopy

gravitational redshift as a function of distance



significant differences in z_G between standard and PC theory



GR and PC similarities

Keplerian frequencies



local maximum in Keplerian frequencies below last stable orbit in ART

local maximum in Keplerian frequencies before last stable orbit in PC theory



The orbital velocity $v^{(\Phi)}$ described in the *Zero Angular Momentum Observer*'s frame reaches a minimum and from thereon increases again (Müller&Camenzind2004)

3.4 Tracing the infall of matter The model in standard GR



matter is falling on a Schwarzschild BH, disappearing at event horizon at 2 R_G

T. Boller, A. Müller, C. Räth, M. Dovciak, J. Svoboda GRAVITAS, YELLOW BOOK, 2010

3.4 Tracing the infall of matter

Feasibility study for a 1m² class X-ray telescope

first infall segment: the relativistic line is clearly visible

third infall segment:

the infalling matter is receding from the observer and strongly Doppler-deboosted

fourth infall segment:

the relativistic line profile becomes visible again at around 4.5 keV due to strong Doppler-boosting

Significant differences between GR and Pseudo-Complex theory are expected for

- the Fe K α line profiles
- the infall frequencies

Summary

Gravitation is well described by Einstein 's General Relativity, however several predictions like the existence of an event horizon are under debate

we have contrasted predictions of the GR with the pseudo-complex field theory

among them we have studied

- gravitational redshift effects
- orbital motions
- perihelion shifts
- timing measurements and gravitationally distorted spectral lines

we consider supermassive black holes as ideal laboratories to test theoretical predictions in the regime of strong gravity