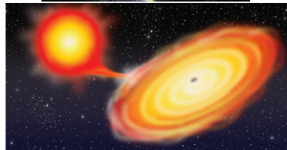
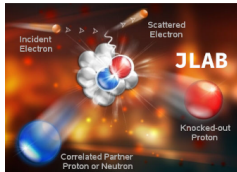
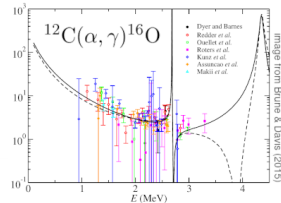
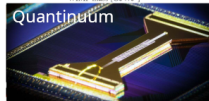
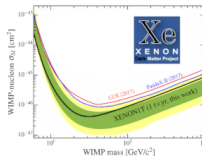
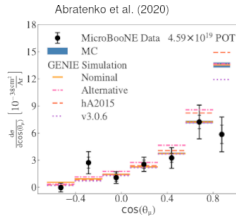


Ab-initio nuclear response functions with quantum (inspired) algorithms

Alessandro Roggero



Ischia – 22 May, 2025



The need for ab-initio many-body dynamics in NP

- ν scattering for supernovae explosion and NS cooling
- capture reactions for crust heating and nucleosynthesis
- cross sections for dark-matter discovery and neutrino physics
- transport properties of neutron star matter for X-ray emission

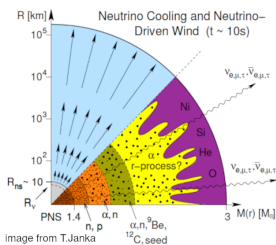


image from T.Janka

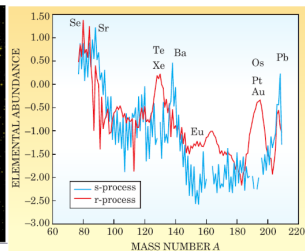
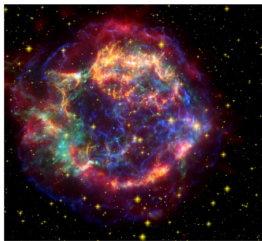


Image from Cowan & Thielemann (2004)



© Piro 2005

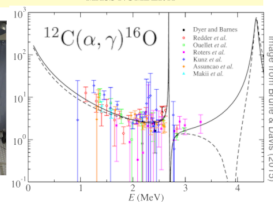
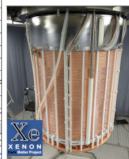
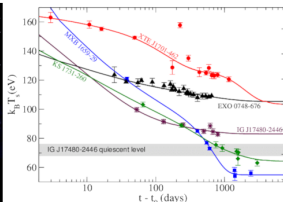
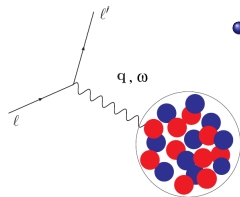


Image from Brune & Davis (2013)

Inclusive cross section and the response function

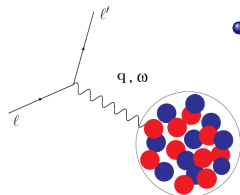


- cross section determined by the response function

$$R_O(\omega) = \sum_f \left| \langle f | \hat{O} | \Psi_0 \rangle \right|^2 \delta(\omega - E_f + E_0)$$

- excitation operator $\hat{O}$ specifies the vertex

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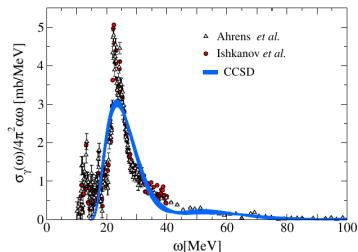
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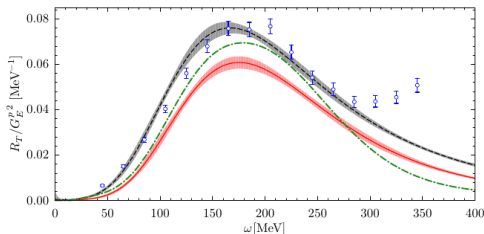
Extremely challenging classically for strongly correlated quantum systems

- dipole response of ^{16}O



Bacca et al. PRL(2013) **LIT+CC**

- quasi-elastic EM response of ^{12}C



Lovato et al. PRL(2016) **GFMC+Laplace**

(see also F.Marino's and G.King's talks)

Many body dynamics with Integral Transforms

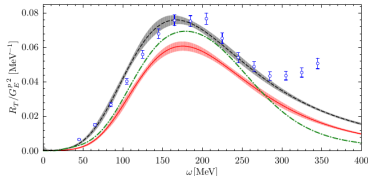
A possible way out with integral transform techniques

Efros (1989), Carlson & Schiavilla (1992), Efros, Leidemann & Orlandini (1994)

$$T(\sigma) = \int d\omega K(\sigma, \omega) R_O(\omega) = \langle 0 | \hat{O}^\dagger K(\sigma, \hat{H} - E_0) \hat{O} | 0 \rangle$$

Laplace

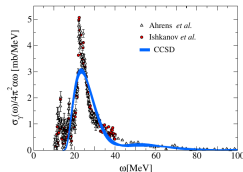
$$K(\sigma, \omega) = e^{-\sigma\omega}$$



Lovato et al. PRL(2016) **GFMC**

Lorentz

$$K(\sigma, \omega; \Gamma) = \frac{\Gamma}{\Gamma^2 + (\sigma - \omega)^2}$$



Bacca et al. PRL(2013) **LIT+CC**

PROBLEM: the inversion procedure is often ill-posed, difficult to assign error bars on the reconstructed response function

Many body dynamics with Integral Transforms II

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ADVANTAGE: if we did, we could do more than linear response!

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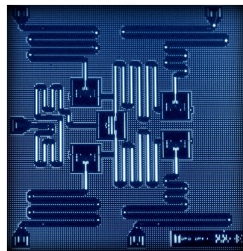
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Simulations of nuclear dynamics

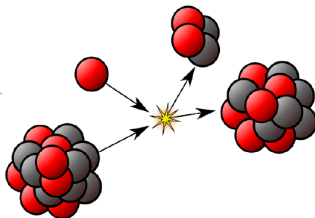
Classical simulations



Quantum simulations



IBM



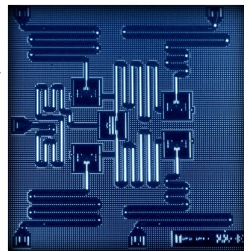
several talks in Ischia: D. Lacroix, J. Menendez, E. Costa

Simulations of nuclear dynamics

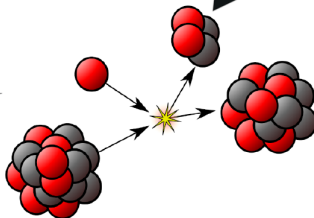
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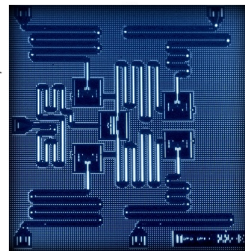
Simulations of nuclear dynamics

"Quantum inspired"

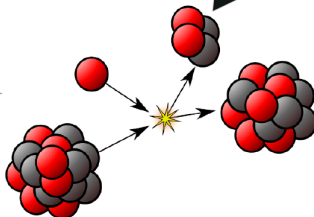
**Classical
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**Quantum
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IBM



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Inclusive cross section from QC

For inclusive scattering seems reasonable to get real-time correlators

$$R_O(\omega) = \int dt e^{i\omega t} C(t) \quad \text{with} \quad C(t) = \langle \Psi_0 | O(t) O(0) | \Psi_0 \rangle$$

- Can be done “easily” using one additional qubit (Somma et al. (2001))

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Turns out it is much more convenient to compute moments

Somma(2019), AR et al.(2020), AR(2020), Rall(2020), Baroni et al.(2021), AR&Sobczyk(2022), Kiss et al.(2023)

$$M_F(t) = \langle \Psi_0 | O e^{-itH} O | \Psi_0 \rangle \qquad M_C(n) = \langle \Psi_0 | O T_n(H) O | \Psi_0 \rangle$$

- Chebyshev Polynomials T_n appear naturally (see also S. Wang's talk)

$$f(H) |\Phi\rangle = \sum_{n=0}^{\infty} c_k T_n(H) |\Phi\rangle \approx \sum_{n=0}^M c_k T_n(H) |\Phi\rangle$$

- Very popular recently for early fault-tolerant ground state energy estimation (and preparation) [Lin & Tong (2022), Dong et al. (2022), Wan et al. (2022)]

Quantum inspired simulation of reactions

Since we want Chebyshev moments, why not get them classically instead?

Orthogonal polynomials satisfy recurrence relations, for Chebyshev

$$T_0(H) = 1 \quad T_1(H) = H \quad \Rightarrow \quad T_{n+1}(H) = 2HT_n(H) - T_{n-1}(H)$$

To get Chebyshev moments we need a many-body method such that

- we can prepare a good approximation to the ground state $|\Psi_0\rangle$
- we can apply the Hamiltonian efficiently

$$\begin{aligned} |\phi_0\rangle &= |\Psi_0\rangle & |\phi_1\rangle &= H |\Psi_0\rangle \\ & & |\phi_n\rangle &\rightarrow |\phi_{n+1}\rangle = 2H |\phi_n\rangle - |\phi_{n-1}\rangle \end{aligned}$$

- we can take overlaps efficiently $m_k = \langle \phi_0 | \phi_k \rangle = \langle \phi_k | \phi_0 \rangle$

Once we have the moments, all the post processing is carried out as if we obtained them from a quantum computer

Quantum inspired simulation of reactions with CC-theory

Sobczyk & Roggero (2022), Sobczyk, Jiang, Roggero (2025)

Coupled-cluster theory allows for accurate nuclear ground states to be prepared efficiently. We can use EOM-CC to study excited-states/moments

$$|\Psi_0\rangle = e^T |HF\rangle \quad \langle\tilde{\Psi}_0| = \langle HF|(1 + \Lambda)e^{-T}$$

the natural construction uses a similarity transformed Hamiltonian

$$\overline{H} = e^{-T} H e^T \quad \text{in CCSD operator } T \text{ contains 1p1h and 2p2h}$$

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Within EOM the excited states are parametrized as

$$|\phi_{n+1}\rangle = 2\overline{H} |\phi_n\rangle - |\phi_{n-1}\rangle = \mathcal{R}_n |\Psi_0\rangle$$

Once we collected the parameters of $\mathcal{R}_n$ we can get moments as

$$m_k = \langle\tilde{\Psi}_0|\phi_k\rangle$$

Reasonably similar to the LIT-CC method (see F.Marino's talk)

Spin response of bulk neutron matter



Dynamic spin structure factor

$$S_{\sigma}(\vec{q}, \omega) \propto \int dt e^{i\omega t} \langle \vec{s}(t, \vec{q}) \cdot \vec{s}(0, \vec{q}) \rangle$$

νN scattering and ν pair-production emissivity dominated by $S_{\sigma}(\vec{q}, \omega)$ for small wave-lengths $|\vec{q}| \rightarrow 0$

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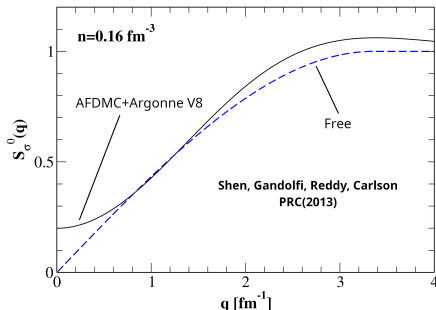
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Total strength given by sum rule

$$S_{\sigma}^0(\vec{q}) = \int d\omega S_{\sigma}(\vec{q}, \omega)$$

Tensor & Spin-orbit terms lead to

$$S_{\sigma}^0(0) = \frac{4}{3N} \langle S^2 \rangle \neq 0$$



Spin response of bulk neutron matter from CC

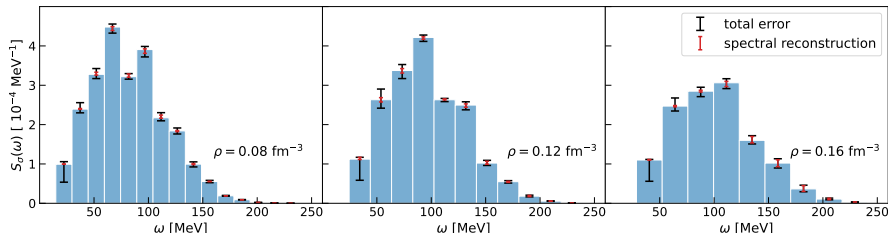
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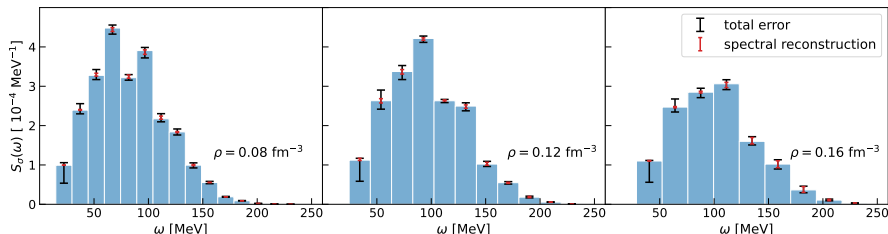


First ab-initio calculation of the frequency dependent spin response of neutron matter with **controllable errors**

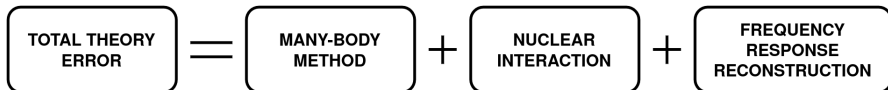
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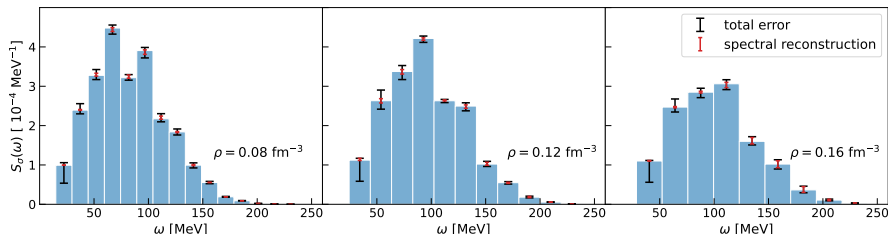
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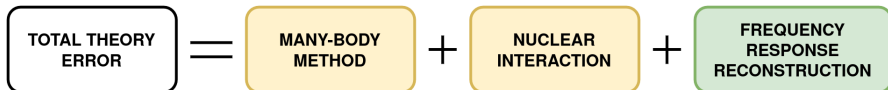
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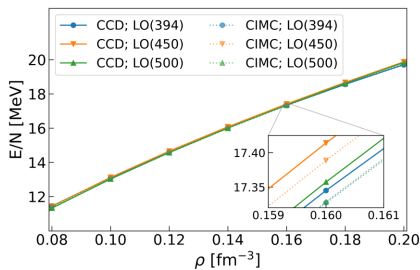


Spin response of bulk neutron matter from CC

Despite being a low energy observable, important interaction dependence

We tested three different Chiral potentials at LO using CC and CIMC

- energy per particle has negligible dependence on method/model

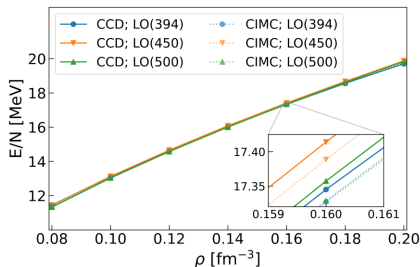
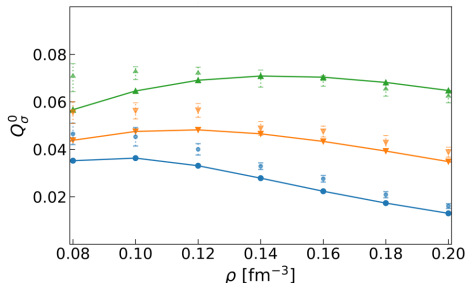


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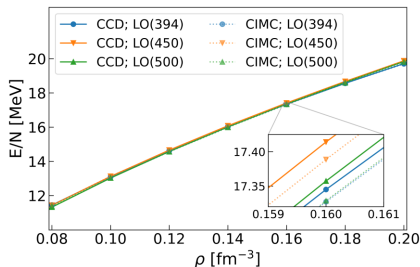
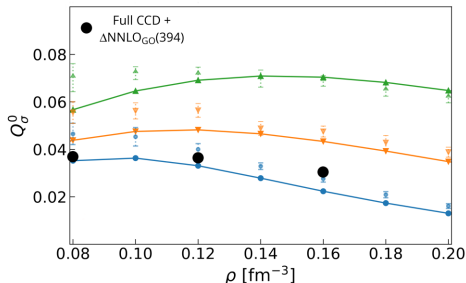


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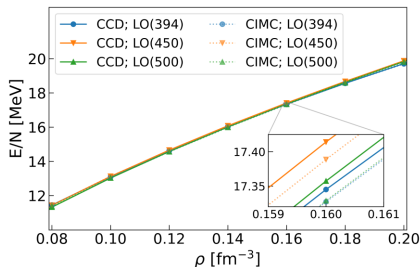
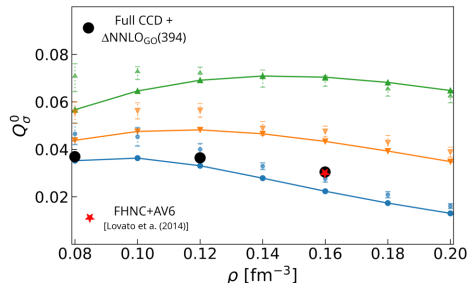
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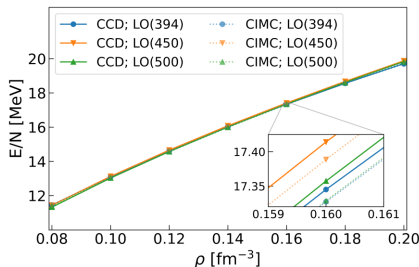
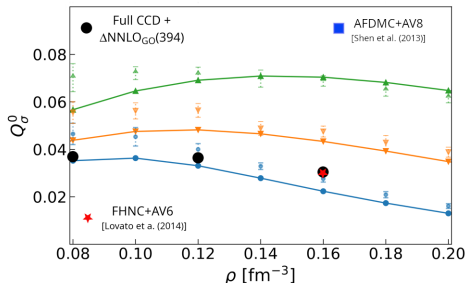
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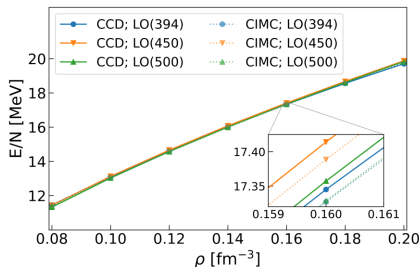
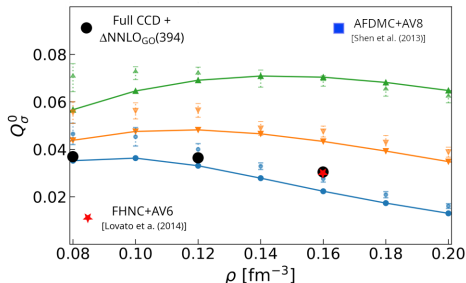
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What is the value of $\langle S^2 \rangle$ in zero temperature neutron matter?

Summary and perspective

- ab-initio treatment of nuclear dynamics is important for both terrestrial experiments and extreme astrophysical sites
- incredible recent progress especially for static nuclear properties but new tools might be needed for some dynamical processes
- quantum computing is a strong candidate to considerably improve our simulations of nuclear physics, especially dynamical properties
- some of the techniques developed for use on future large scale quantum computers can be employed now on classical HPC!
- calculation of spin response shows important interaction/method dependence: **neutron matter around ρ_0 not boring anymore!**

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Thanks to my collaborators

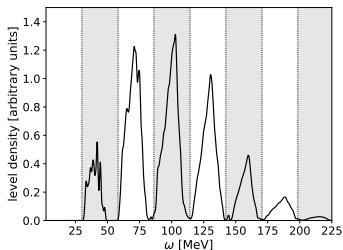
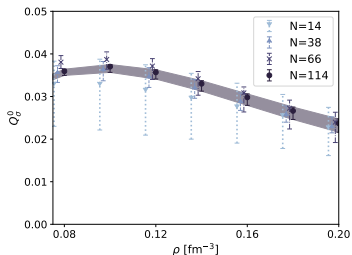
- J. Sobczyk (Mainz → Chalmers)
- W. Jiang (Mainz)



Mainz Institute for
Theoretical Physics

Finite size systematics

We use TABC to minimize finite size effects. This works well for sum rules but residual N dependence in the density of states (and thus the response)



Clustering can be explained as a shell effect: at fixed density $\rho = N/L^3$ so the free single particle energies are $E_n \propto n(2\pi/L)^2$

