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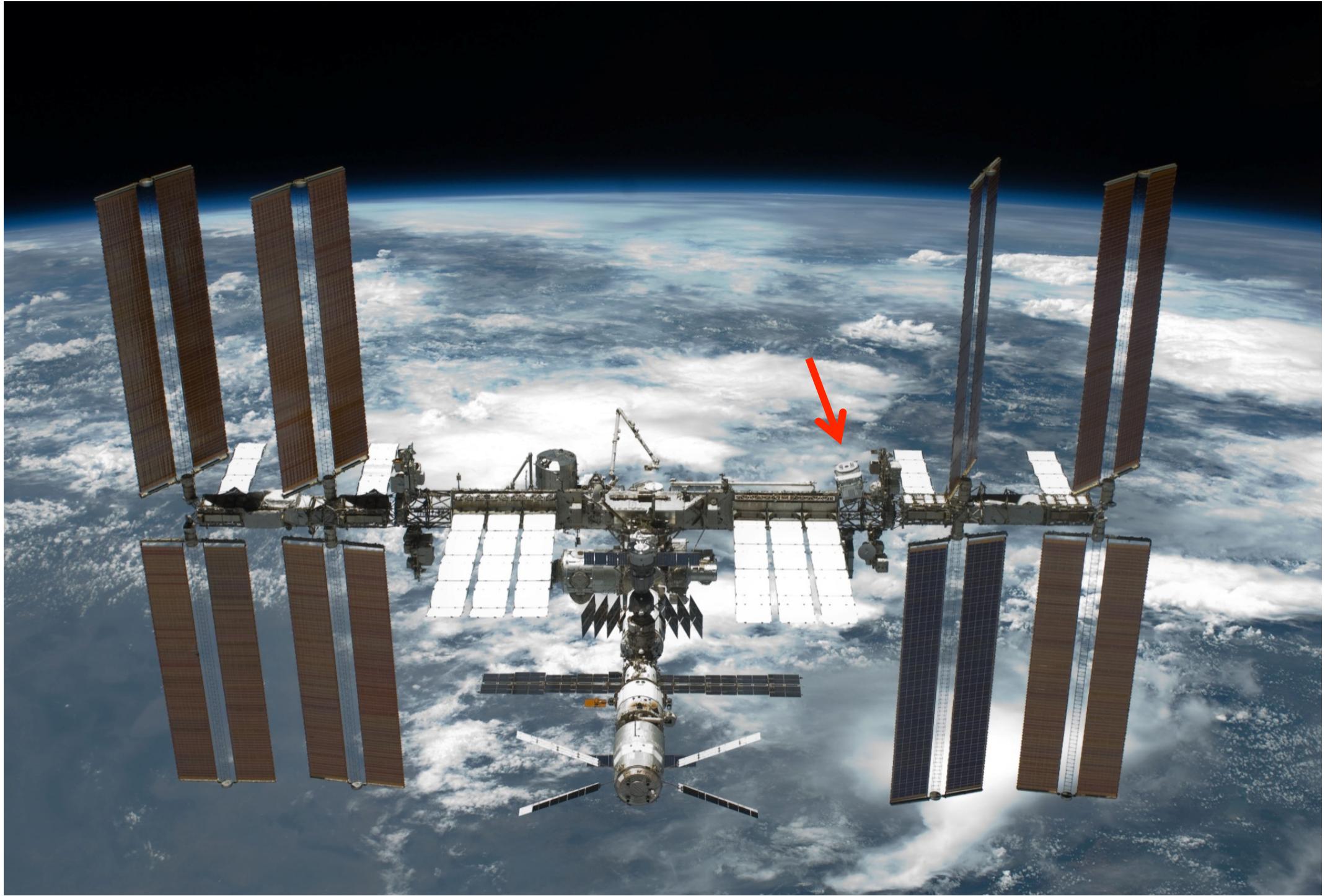
# AMS-02 one year in space

Cecilia Pizzolotto

on behalf of the AMS-02 Collaboration  
INFN Perugia and ASDC Frascati



SciNeGHE 20-22 June 2012, Lecce





# AMS-02 goals

The experimental challenge: perform accurate, high statistics, long term measurements of charged cosmic rays (0.5 GV – O(TV)) and  $\gamma$  rays ( $E > 1\text{GeV}$ )

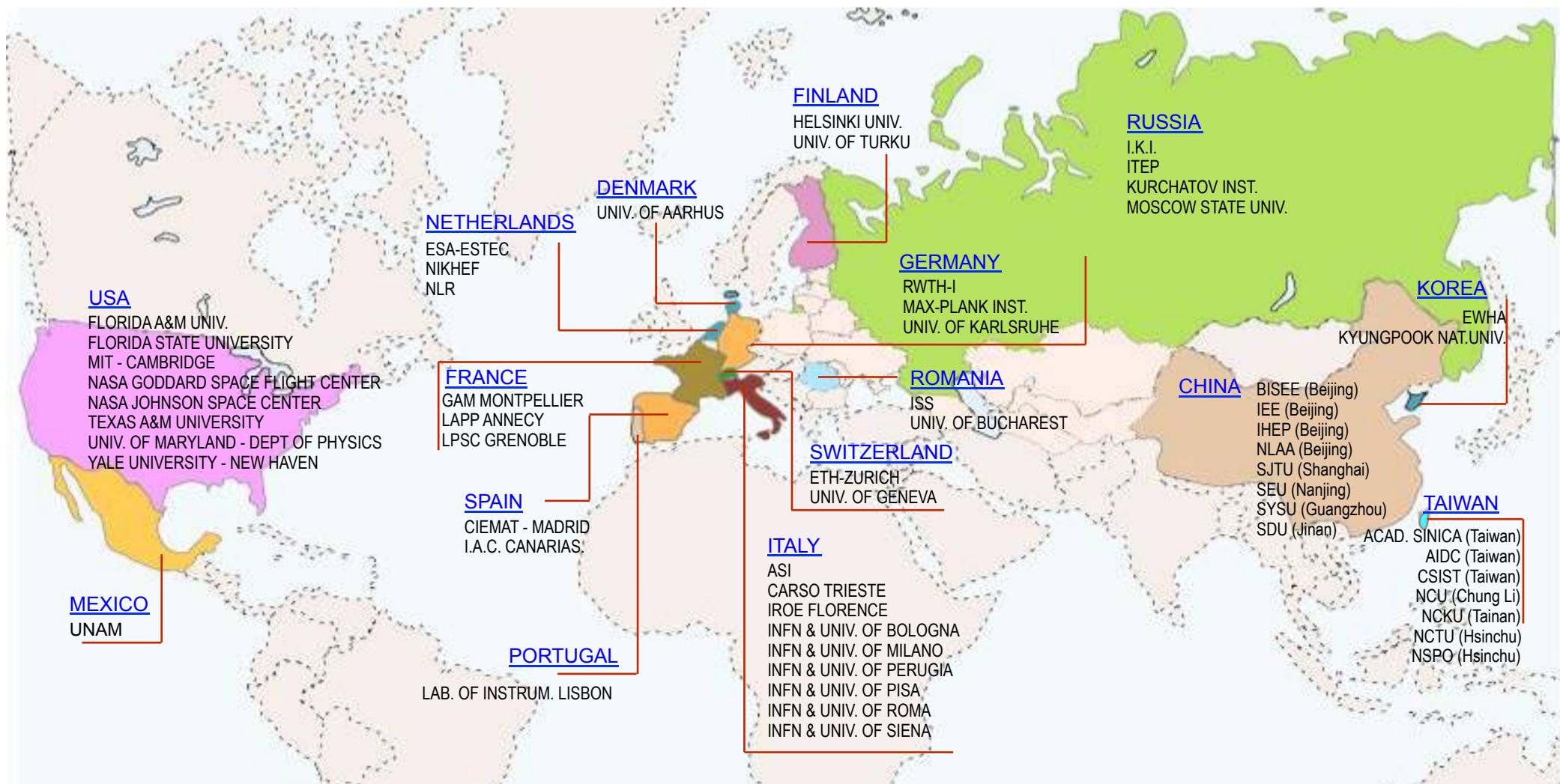
- Primordial Antimatter search with  $10^{-9}$  sensitivity
- Indirect Dark Matter search ( $e^+, \dots, \gamma$ )
- Relative abundance of nuclei and isotopes in primary cosmic rays
- $\gamma$  ray astrophysics



AMS in the shuttle cargo bay

# AMS international collaboration

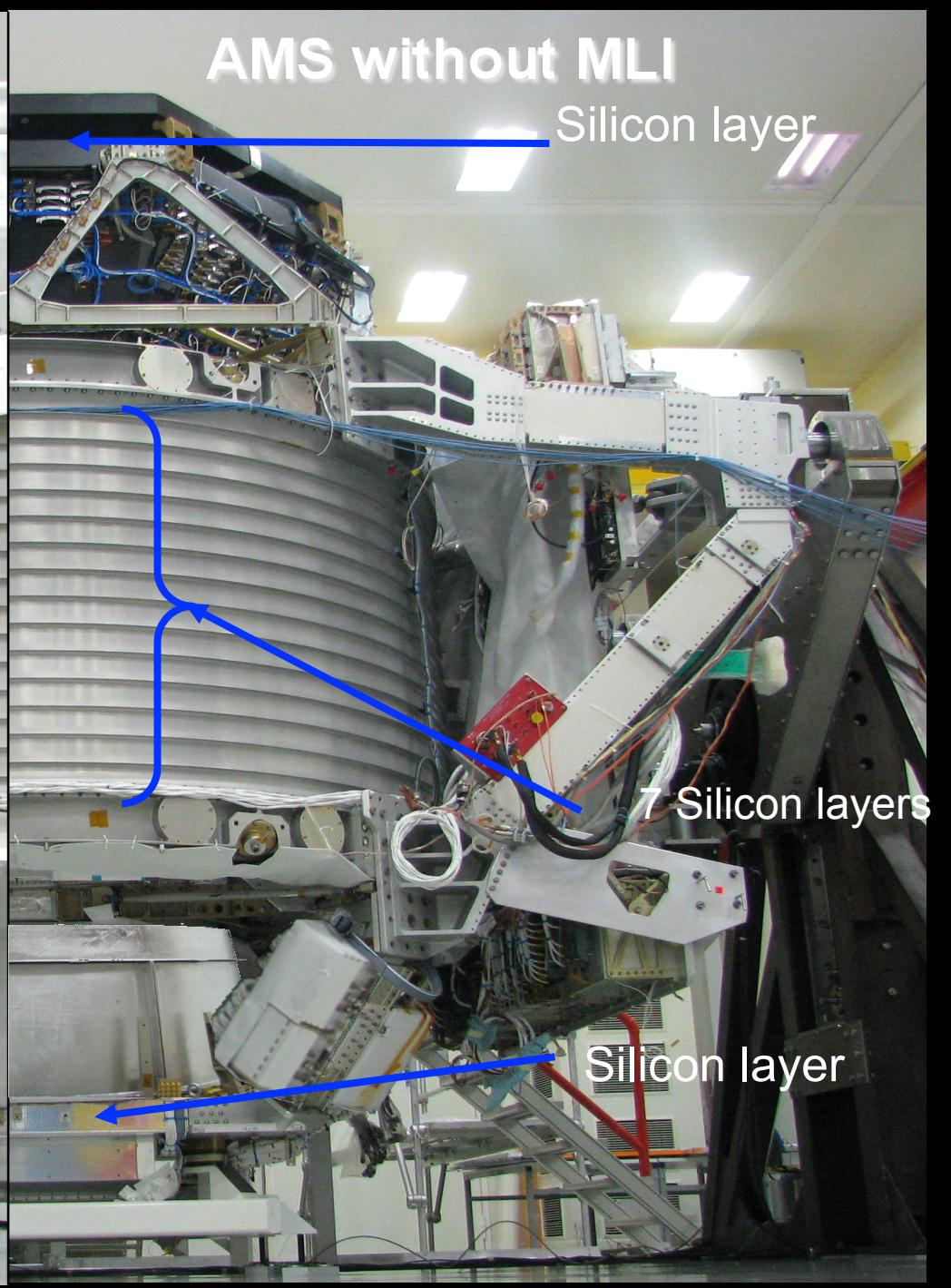
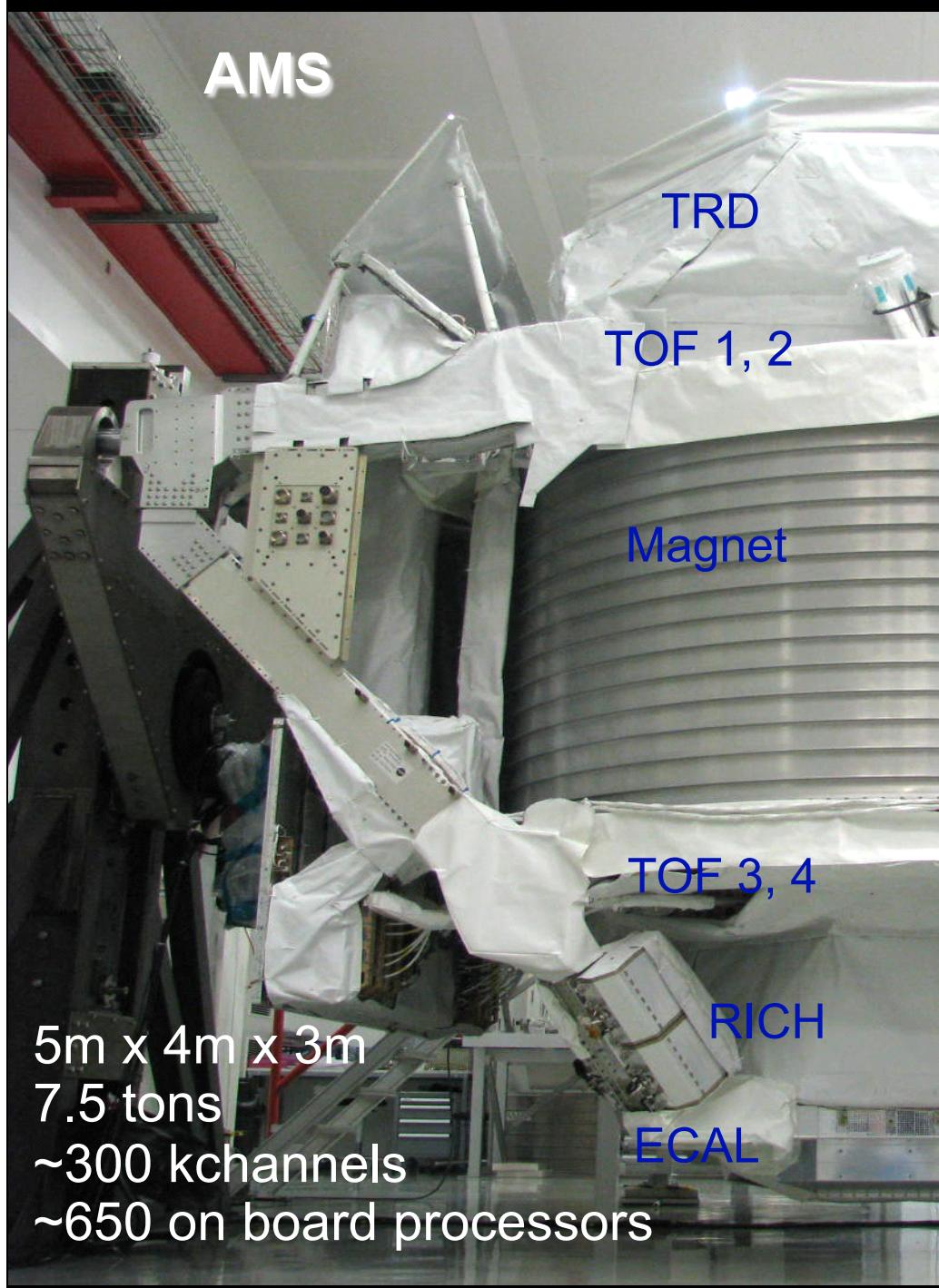
16 Countries, 60 Institutes and 600 Physicists, 17 years



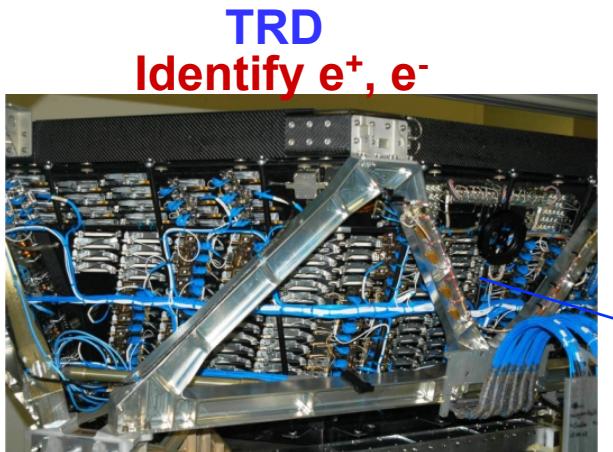
The detectors and electronics were built all over the world and assembled at CERN, Switzerland.



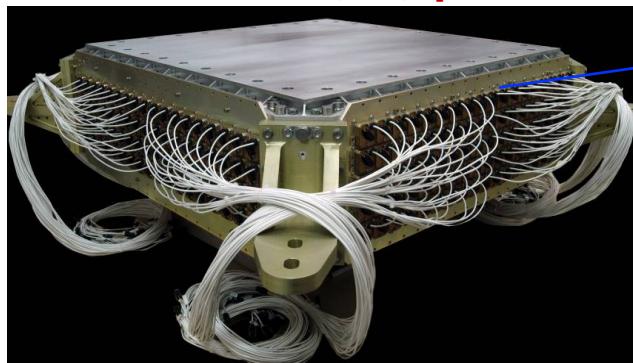
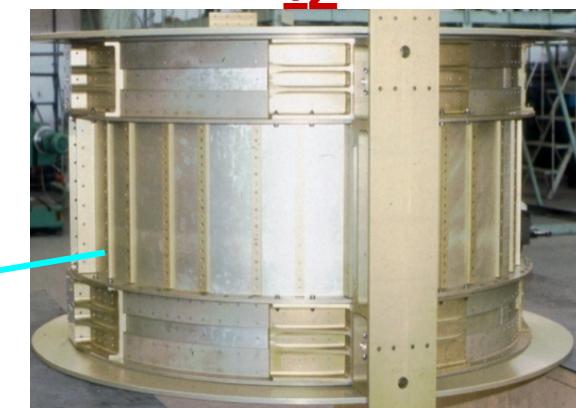
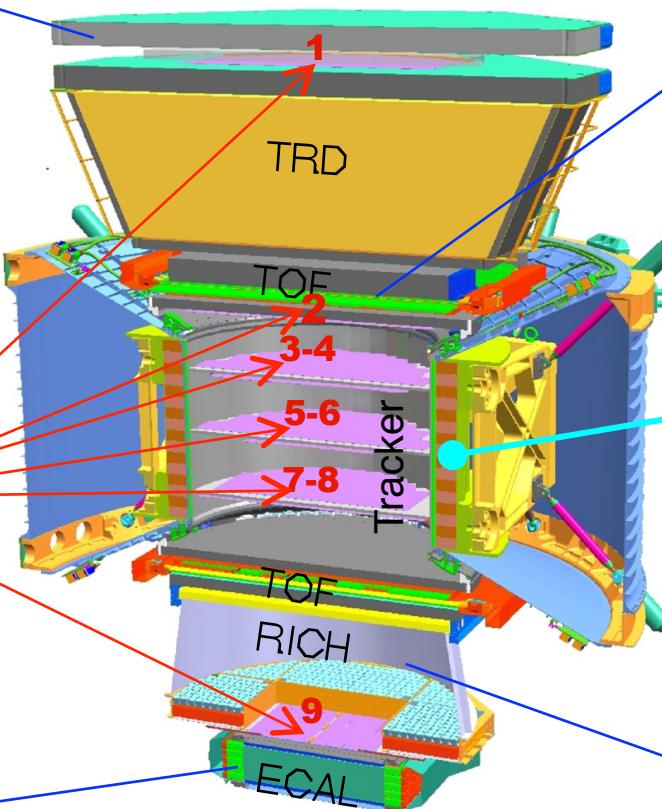
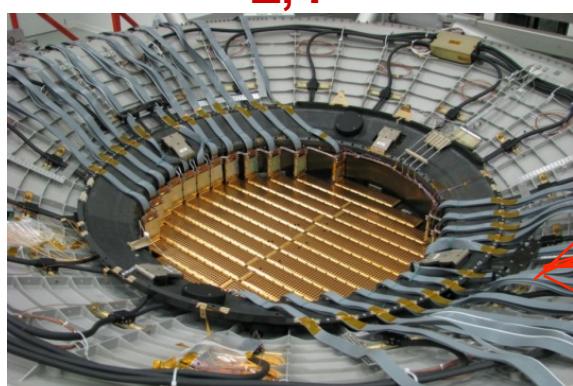
JSC Control Room, May 2011



# AMS-02: A TeV precision, multipurpose spectrometer



Particles and nuclei are defined by their charge ( $Z$ ) and energy ( $E \sim P$ )



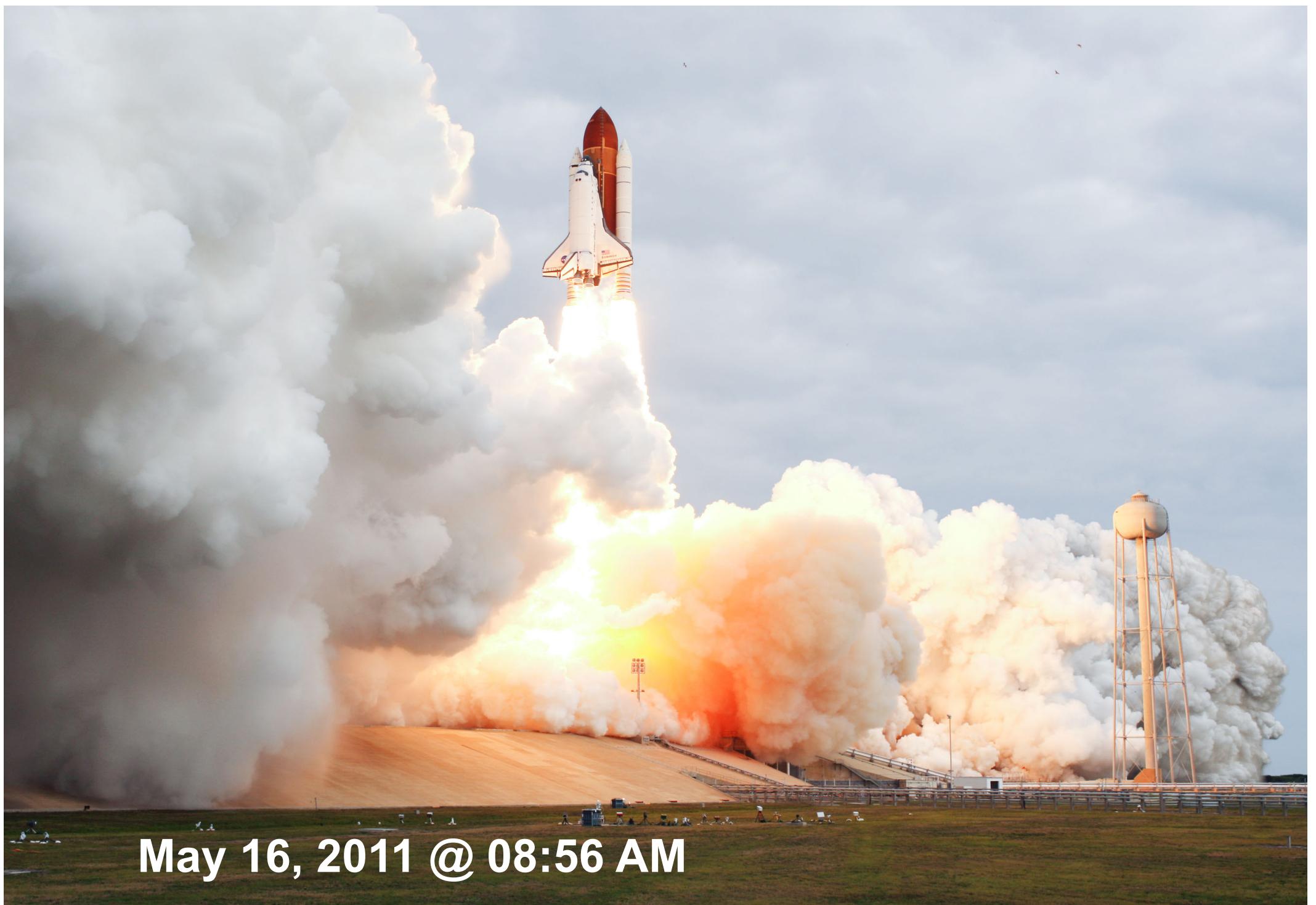
$Z, P$  are measured independently by the Tracker, RICH, TOF and ECAL

MDR = 2 TeV



# Particle Identification with AMS

	$e^-$	P	He,Li,Be,..Fe	$\gamma$	$e^+$	$\bar{P}, \bar{D}$	$\bar{He}, \bar{C}$
TRD							
TOF							
Tracker							
RICH							
ECAL							
Physics example	Cosmic Ray Physics Strangelets				Dark matter		Antimatter

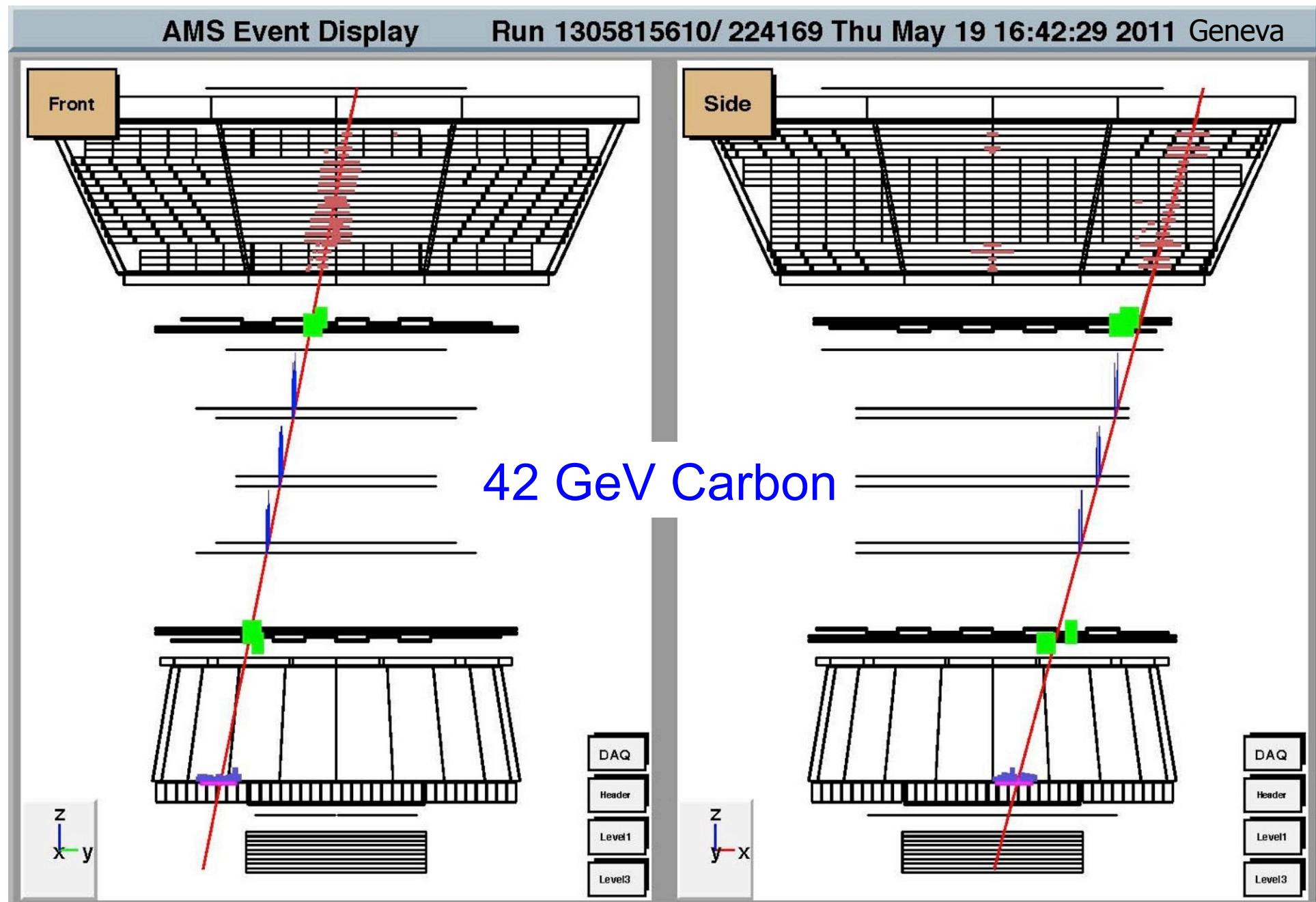


May 16, 2011 @ 08:56 AM

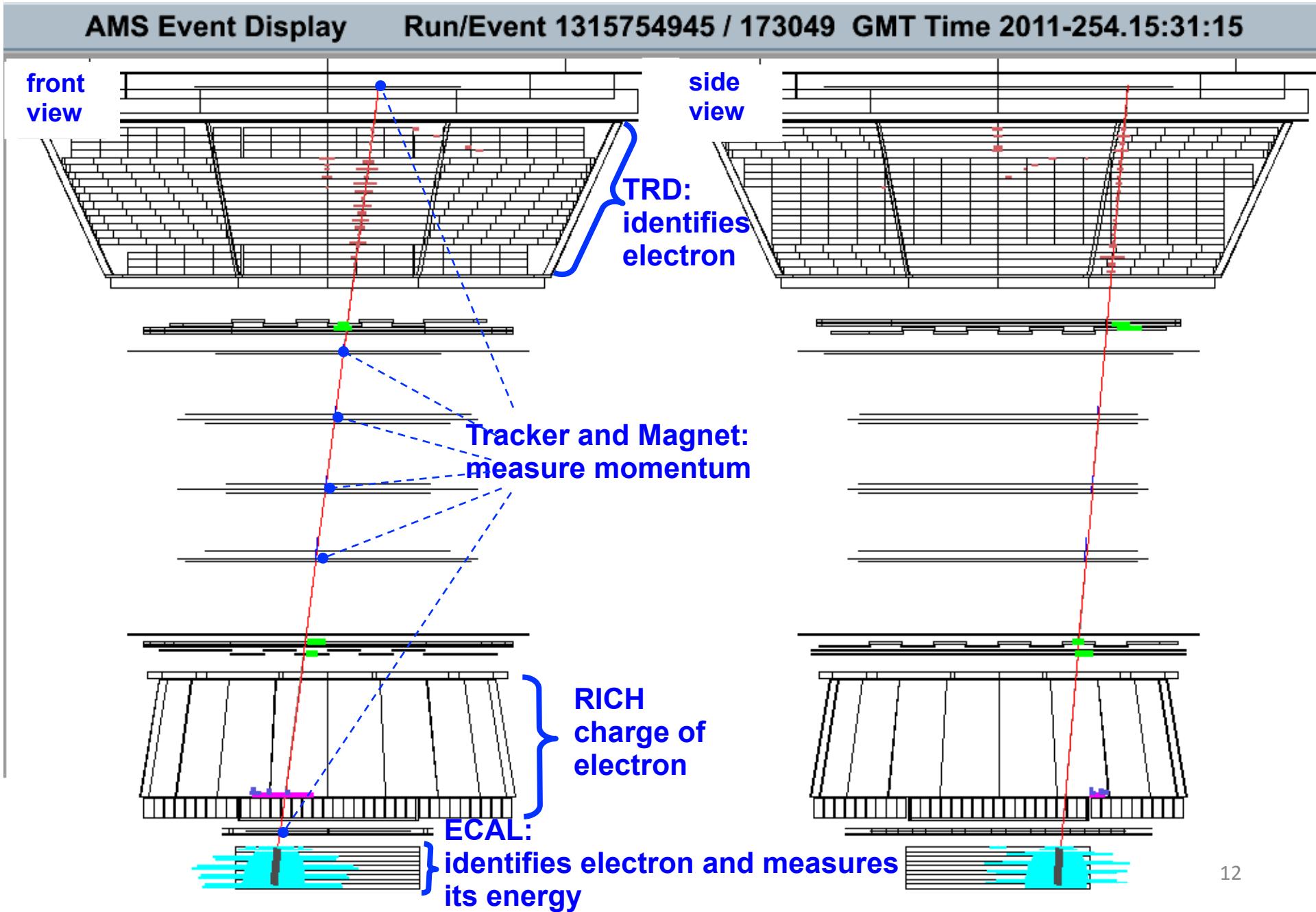


**May 19, 2011: AMS installation completed at 5:15 CDT,  
data taking started at 9:35 CDT**

**May 19: AMS installed on ISS 5:15 CDT, start taking data 9:35 CDT**



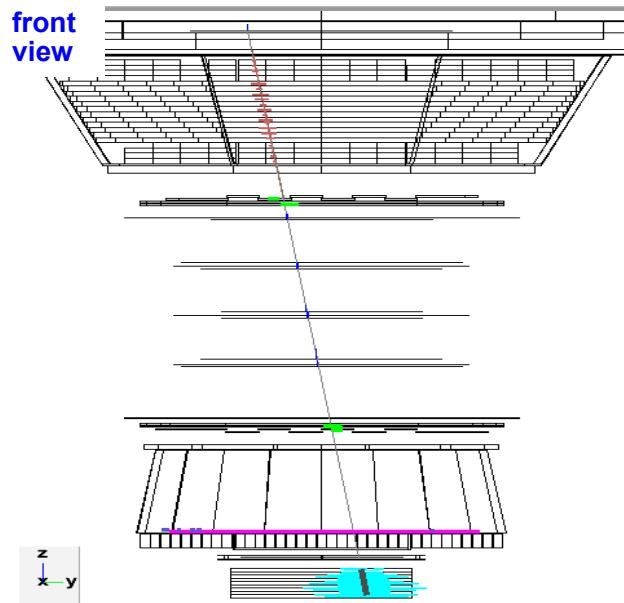
# 1.03 TeV electron



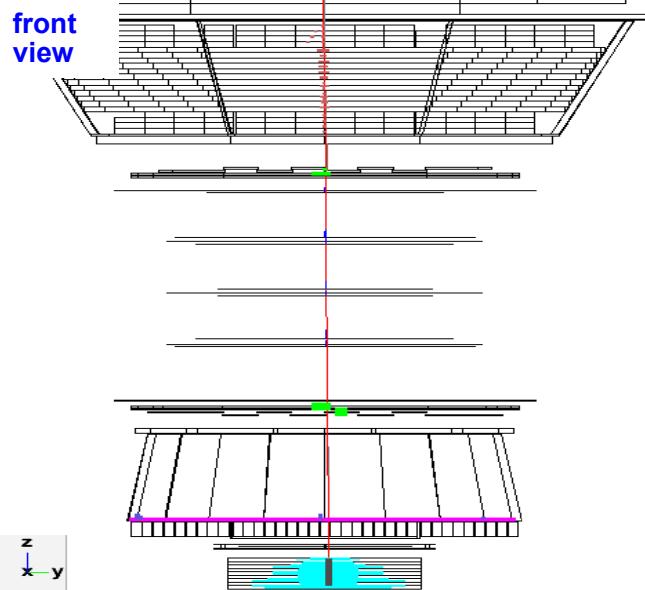
12

# High energy e+

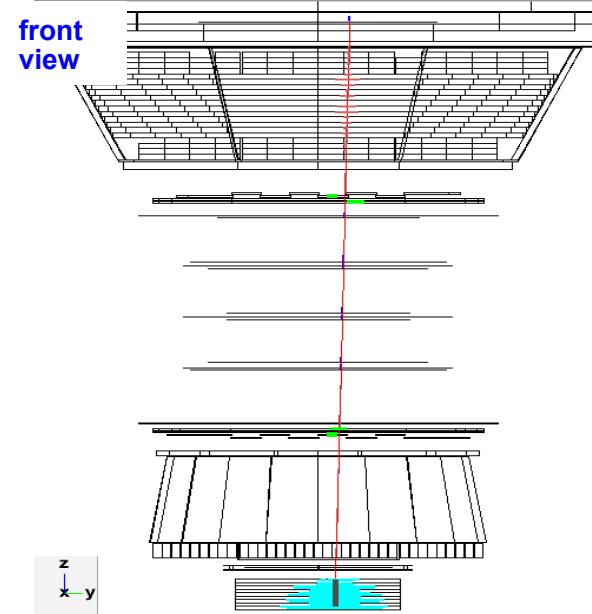
**205 GeV Positron**



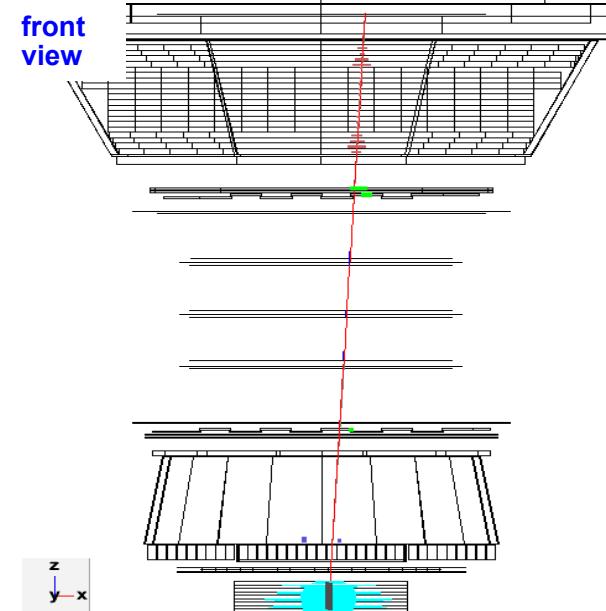
**388 GeV Positron**



**369 GeV Positron**



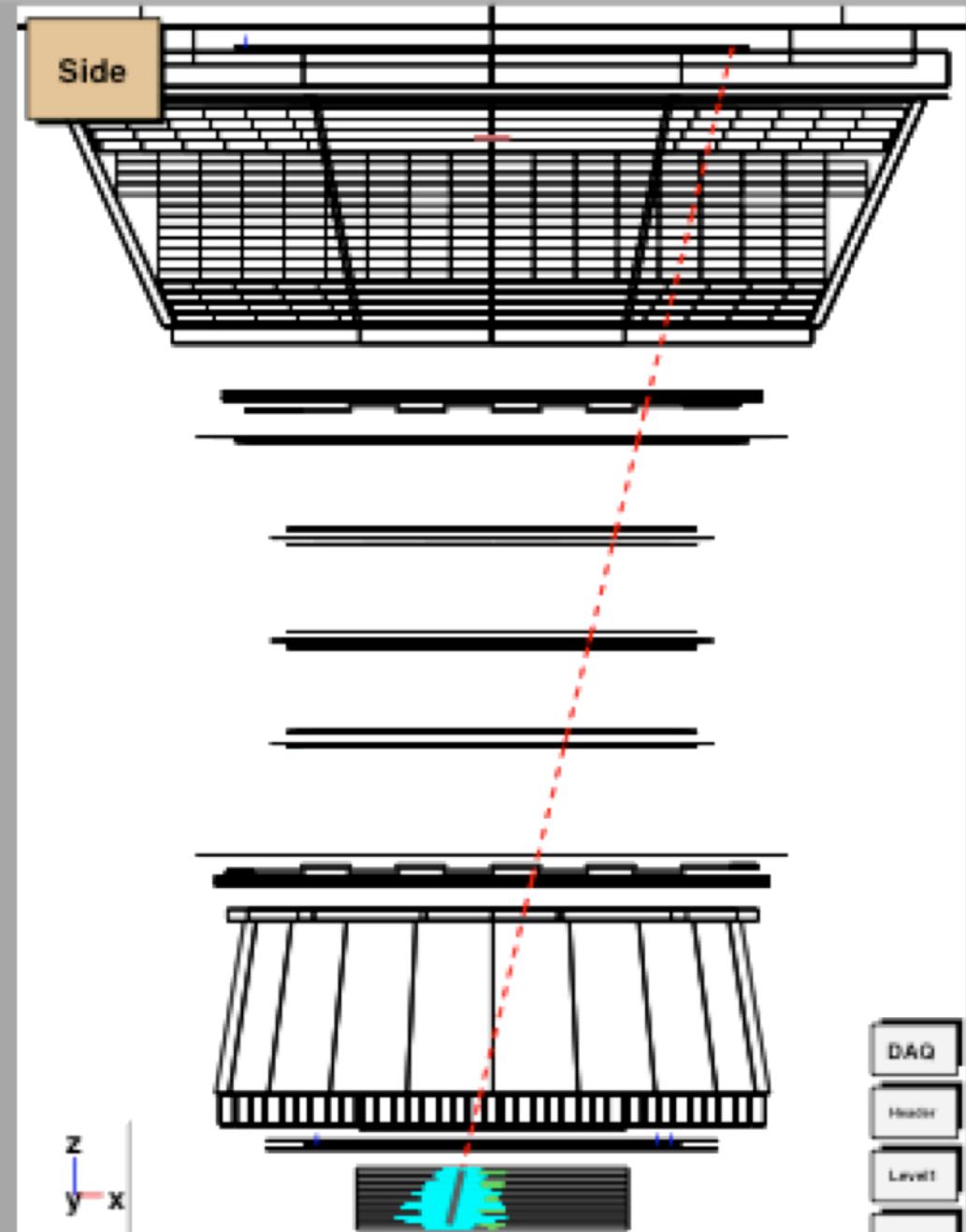
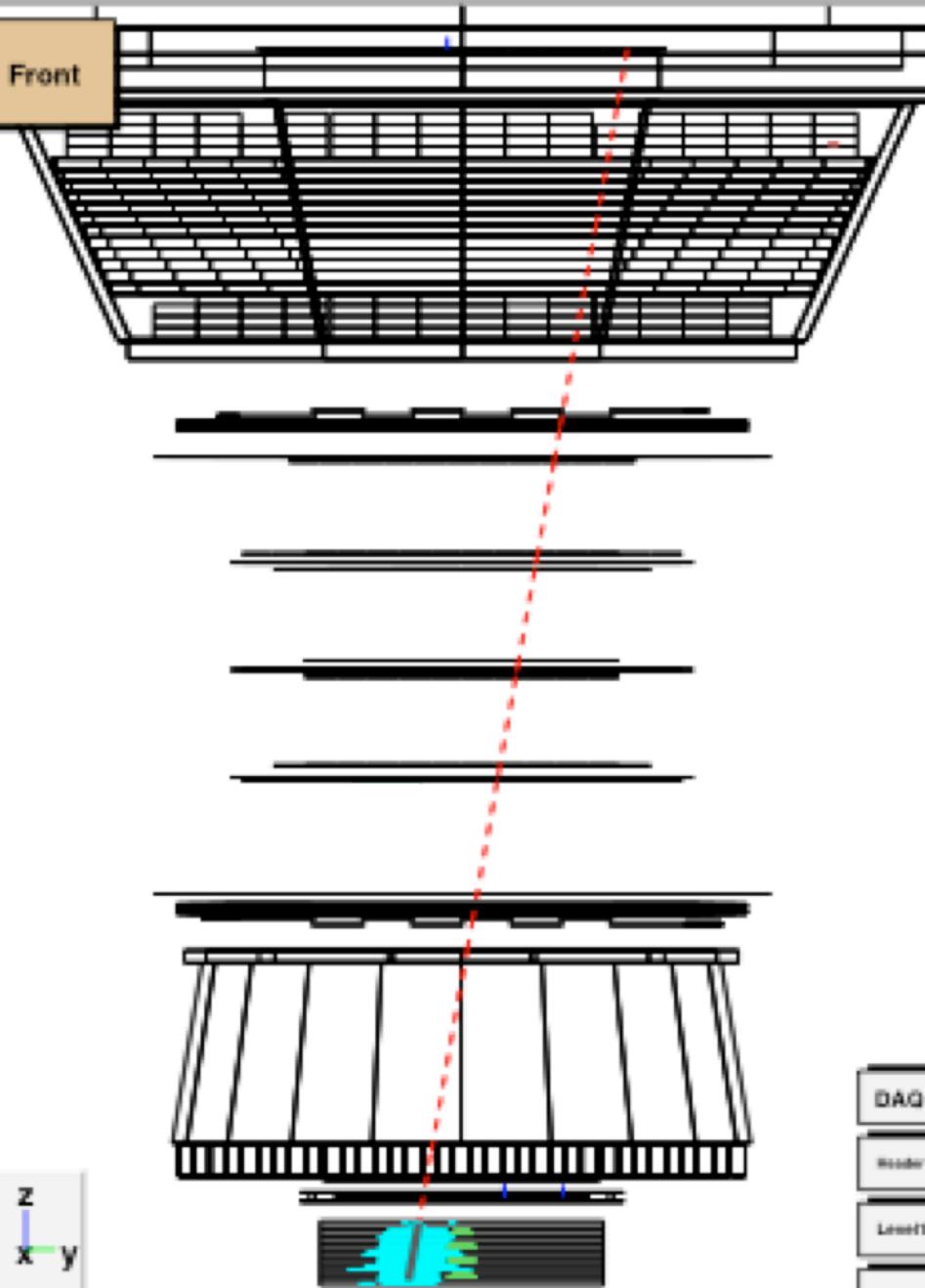
**424 GeV Positron**



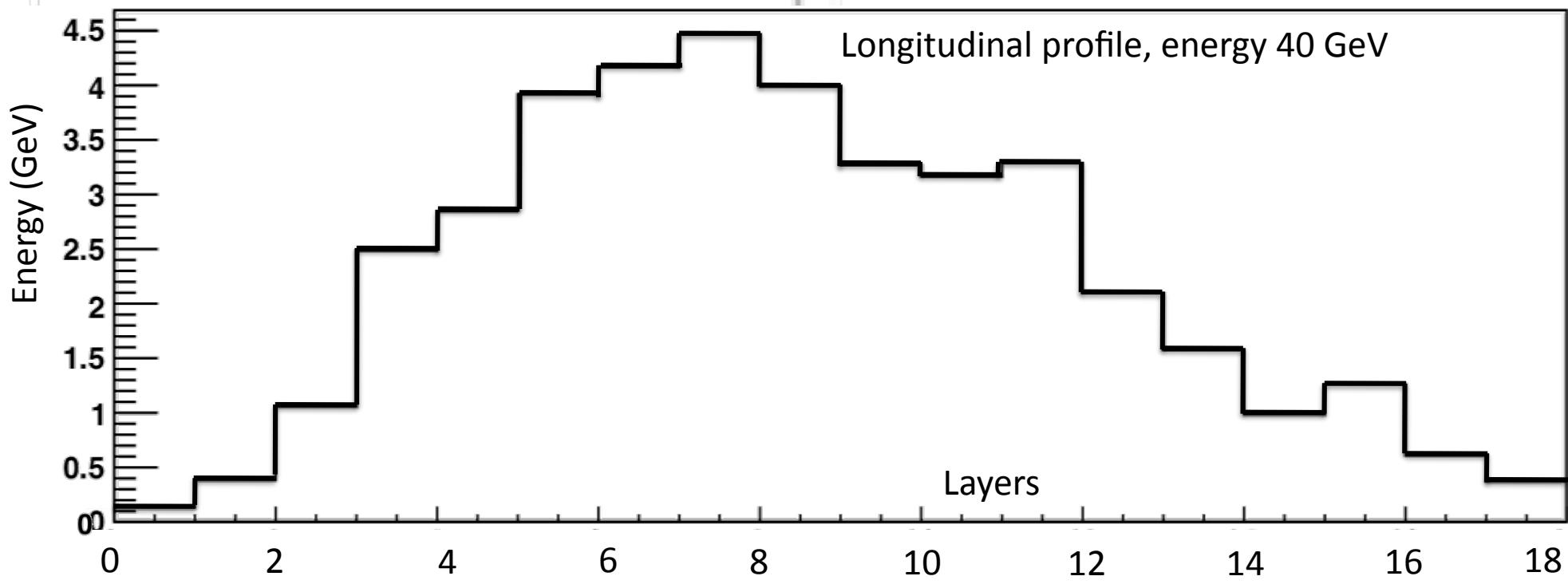
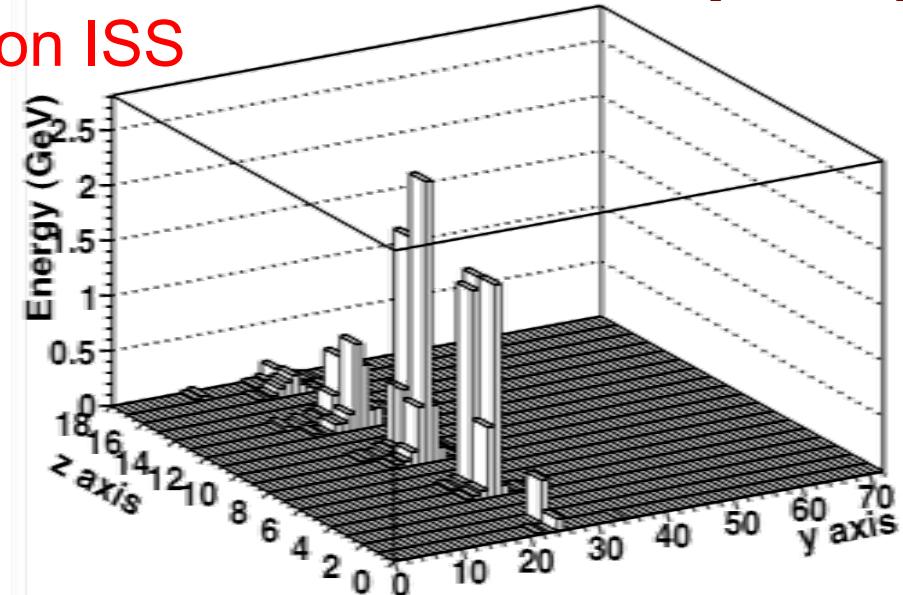
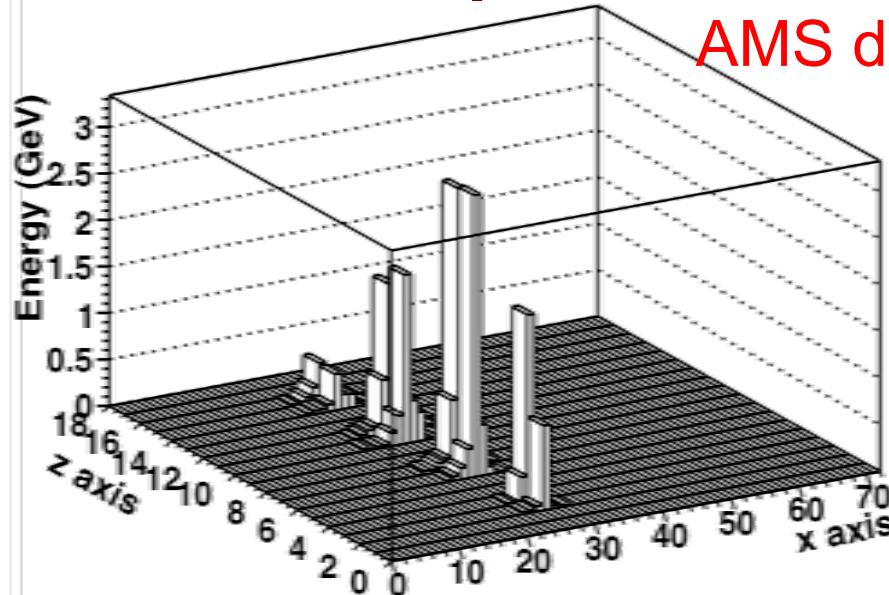
# Photon 40 GeV

AMS Event Display

Run/Event 1306127850 / 159966 GMT Time 2011-143.05:26:24



# 40 GeV photon, 3D shower sampling

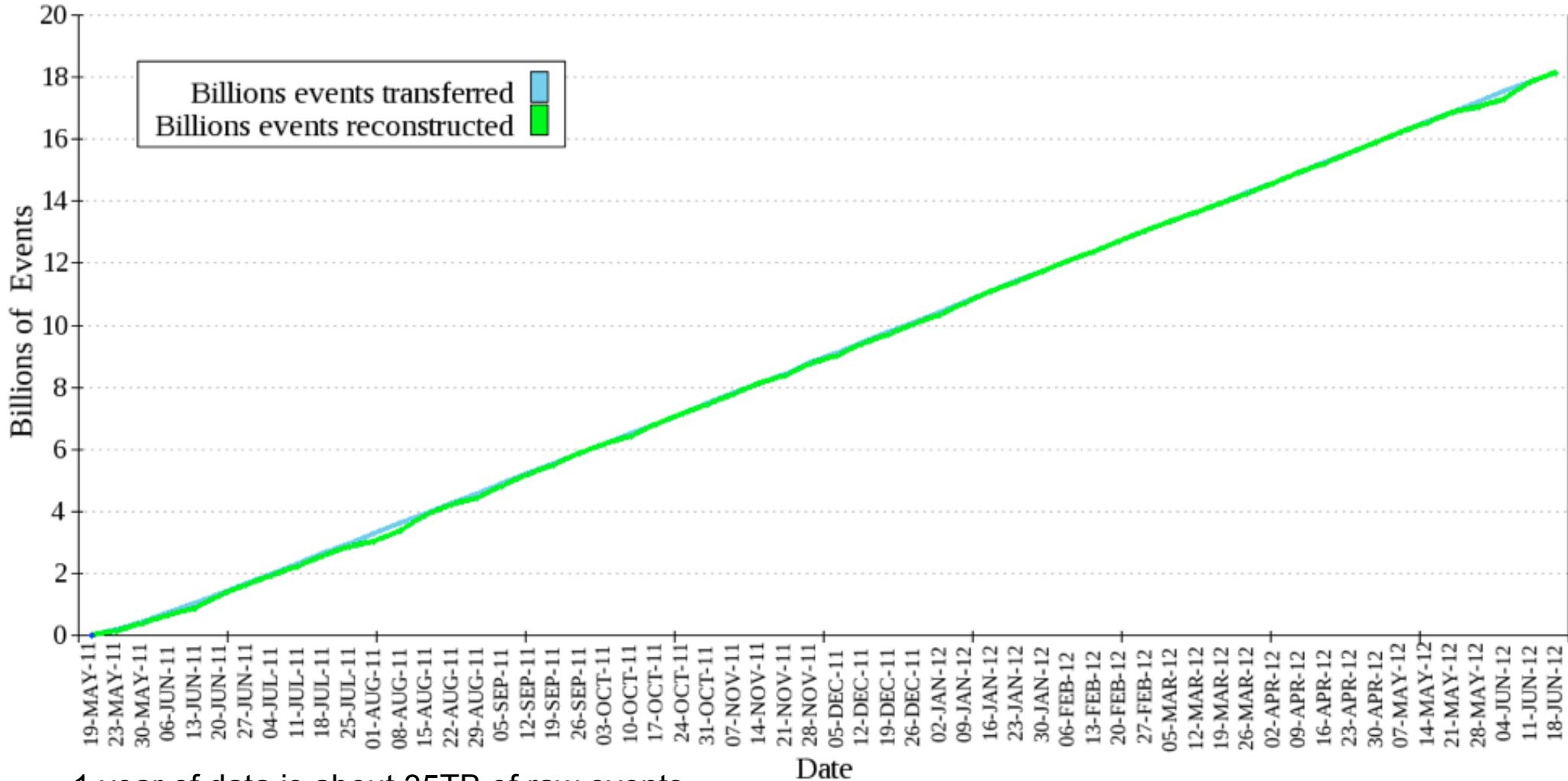




# AMS collected over 18 billion of events

Every year, we will collect  $16 \times 10^9$  events

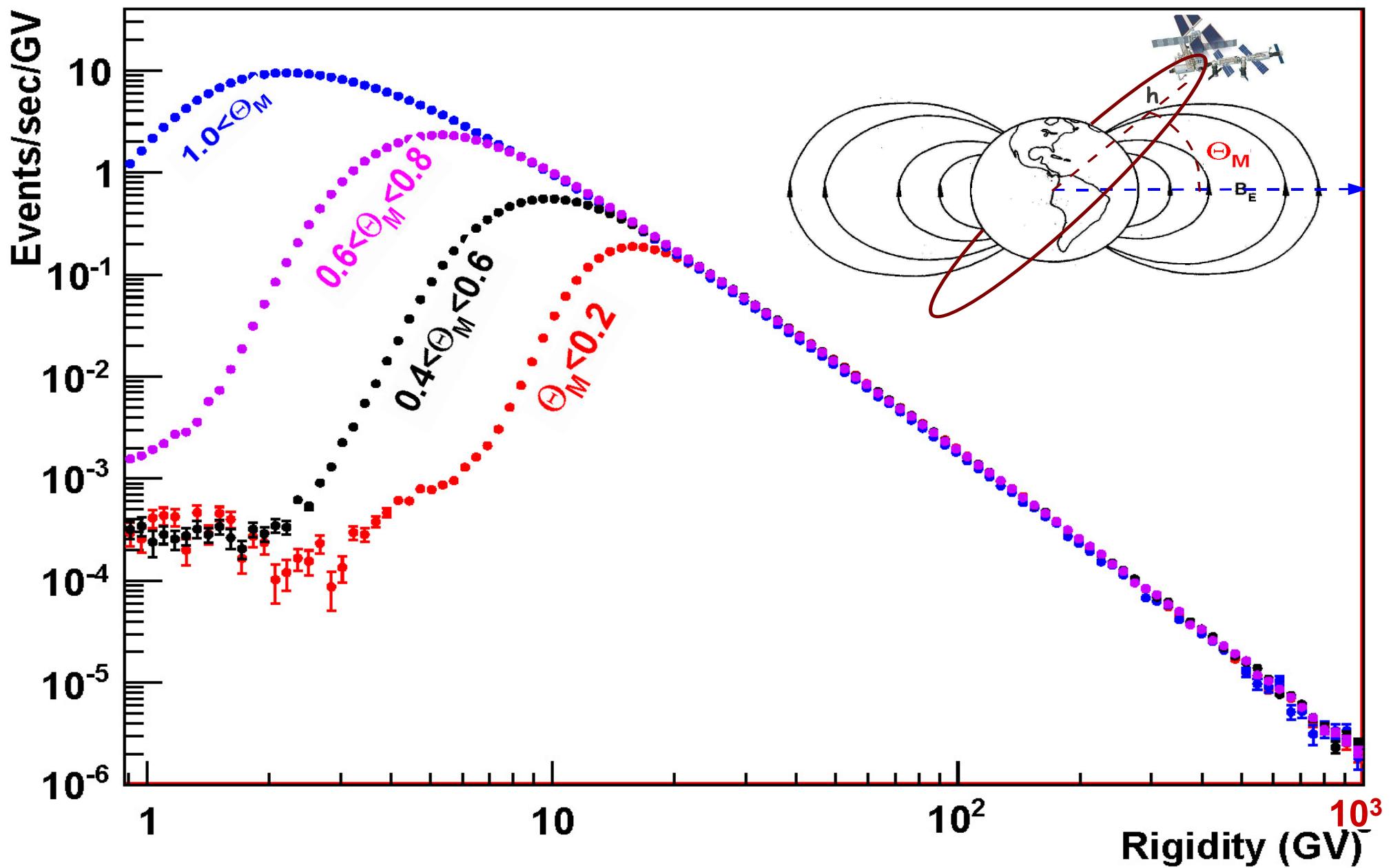
This will provide unprecedented sensitivity to search for new physics.



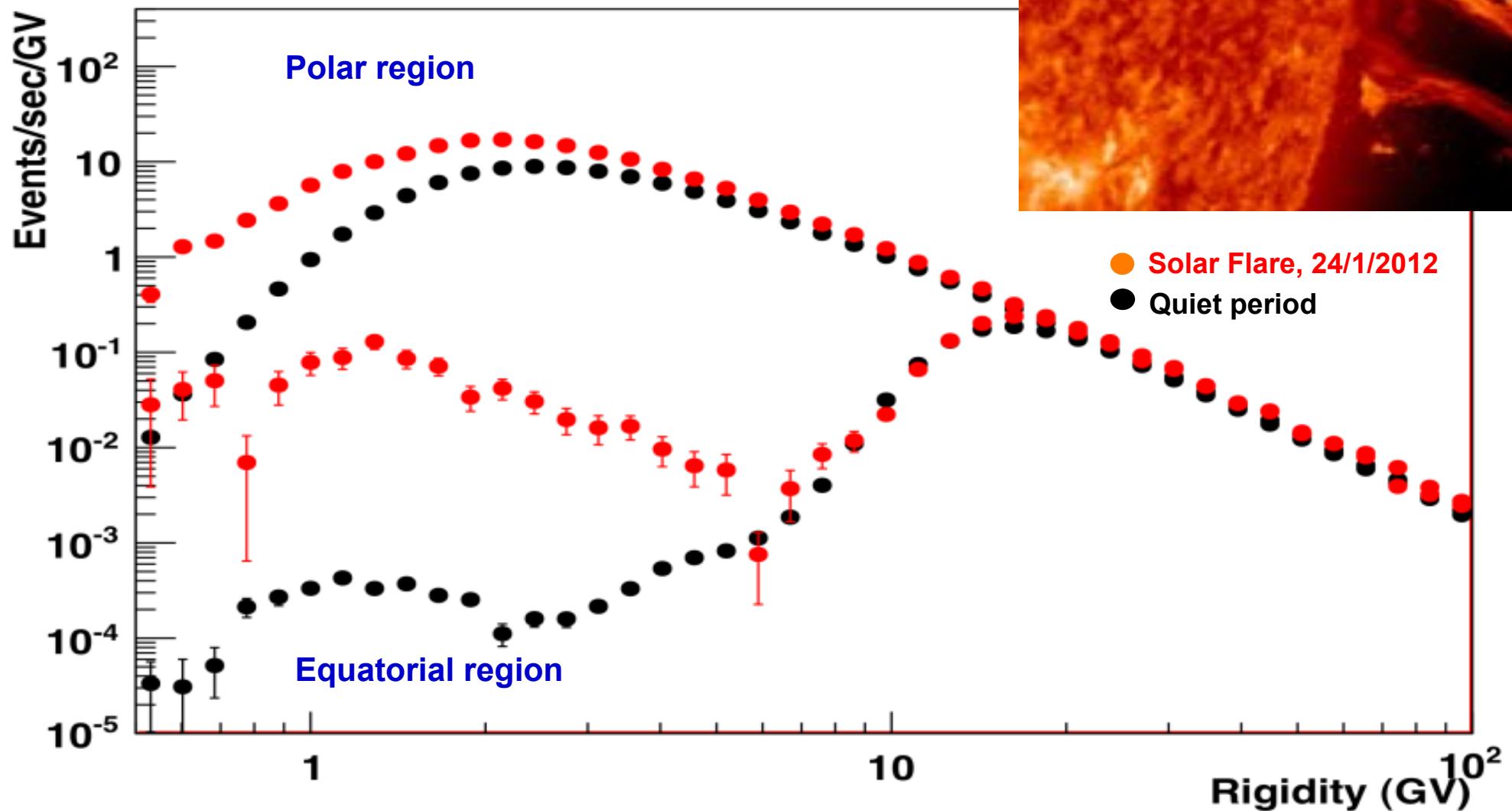
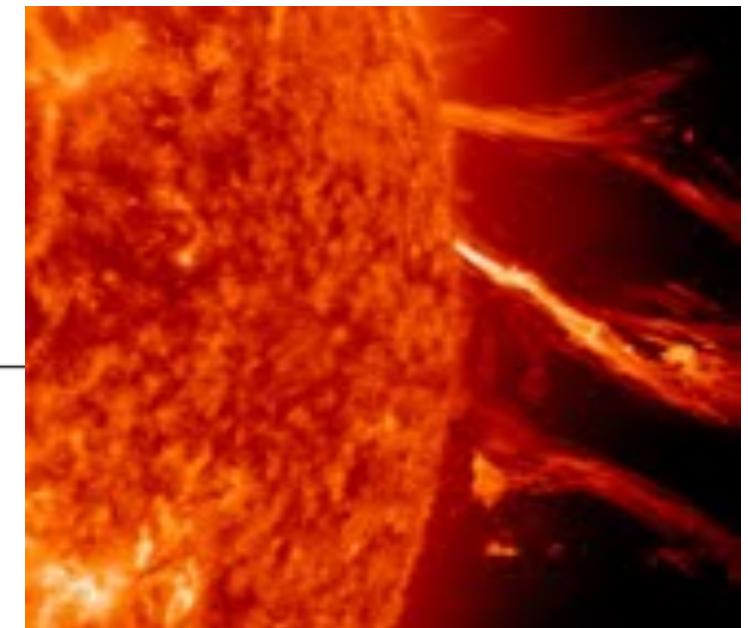
1 year of data is about 35TB of raw events

First year of data reproduction is completed after detector calibration and ready for analysis

# AMS data on the ISS: He rate



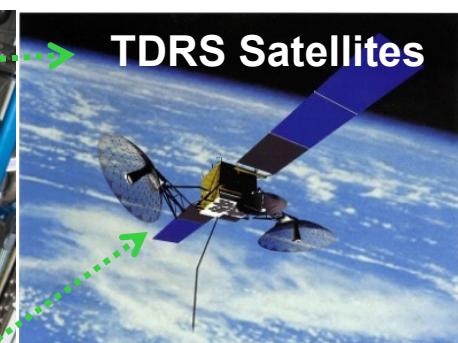
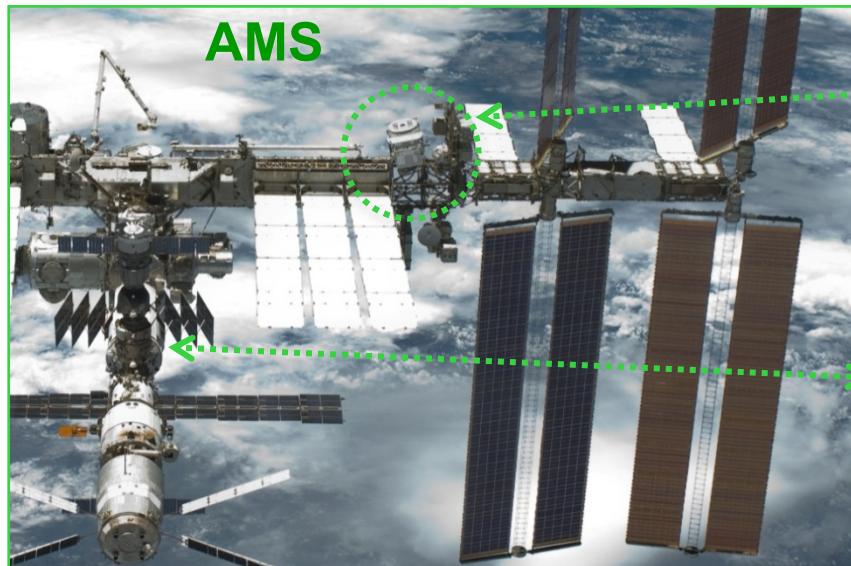
# AMS data: He rate and Solar flare



The first year:  
→ Operation  
→ Calibration in Space



# AMS Operations



AMS Payload Operations Control and  
Science Operations Centers  
(POCC, SOC) at CERN

Flight Operations  
Ground Operations



AMS Computers  
at MSFC, AL



White Sands Ground  
Terminal, NM

Ku-Band  
High Rate (down):  
Events <10Mbit/s>

S-Band  
Low Rate (up & down):  
Commanding: 1 Kbit/s  
Monitoring: 30 Kbit/s

# Payload Operations Control Center (POCC)

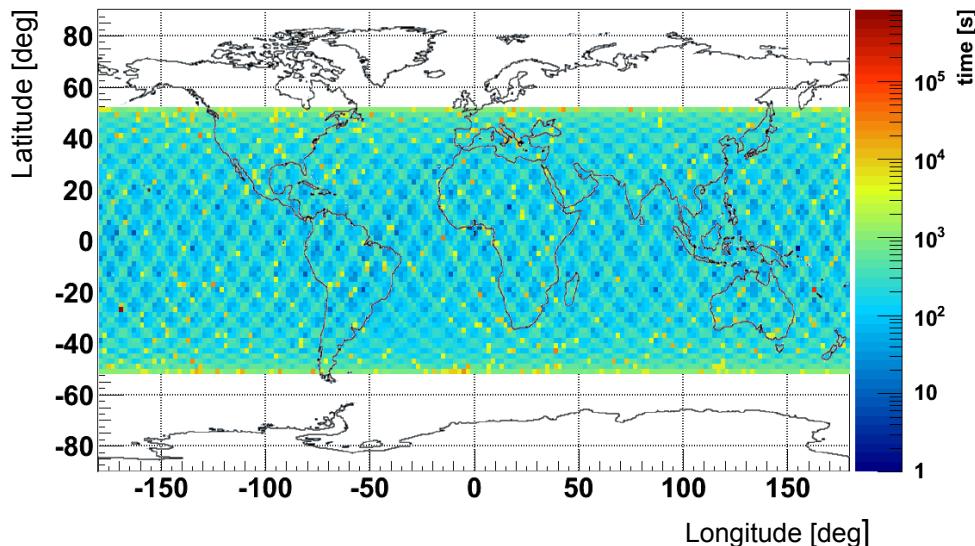
## at CERN since 19th June 2011



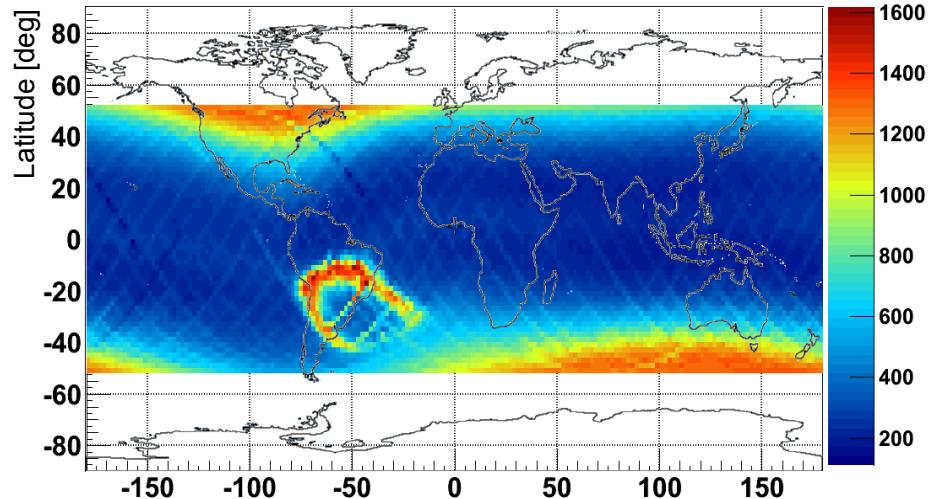
**SOC** (Science Operation Center): production of *root* files for science analysis

# Orbital DAQ parameters

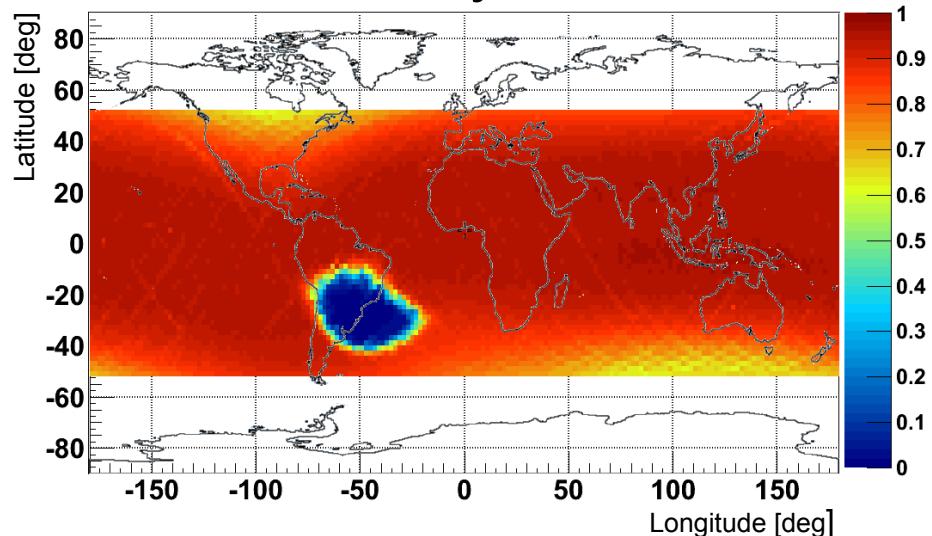
Time at location [s]



Acquisition rate [Hz]



DAQ efficiency

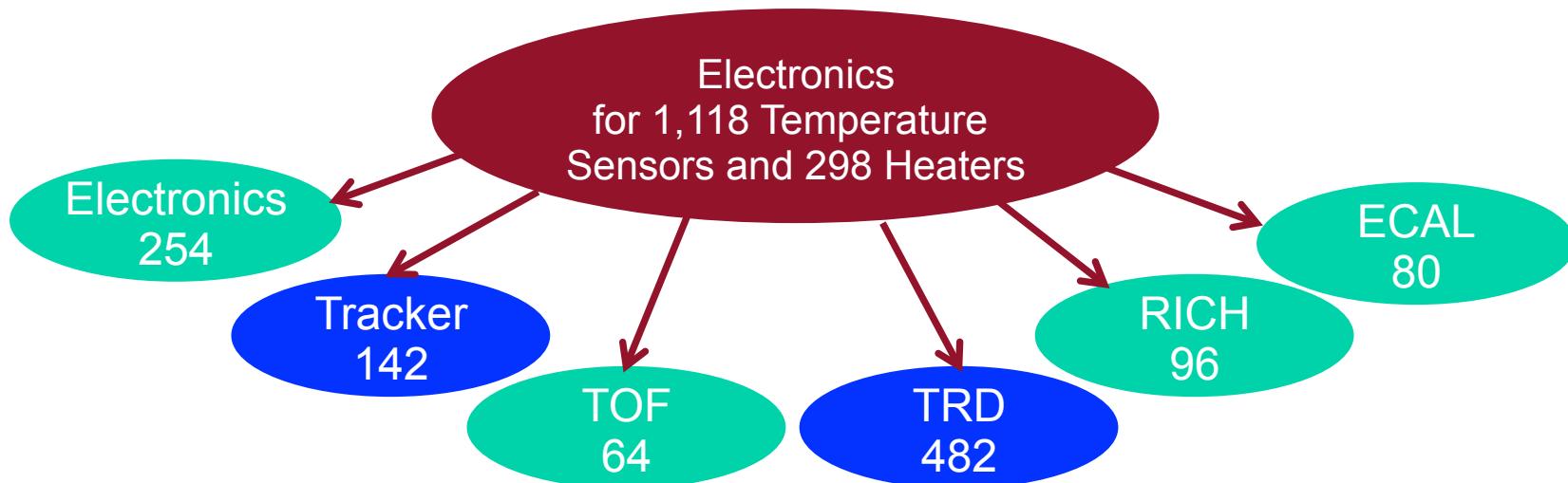


ISS orbits at about 400km  
with an inclination of 51.7 degrees

Particle rates: 200 to 2000 Hz per orbit

Orbit average: DAQ efficiency 85%  
DAQ rate ~530Hz

# AMS Flight Electronics for Thermal Control



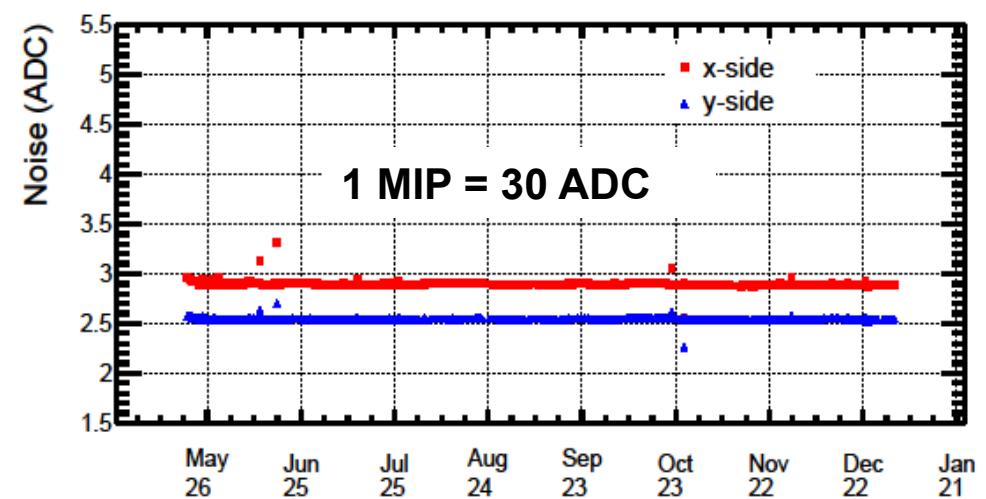
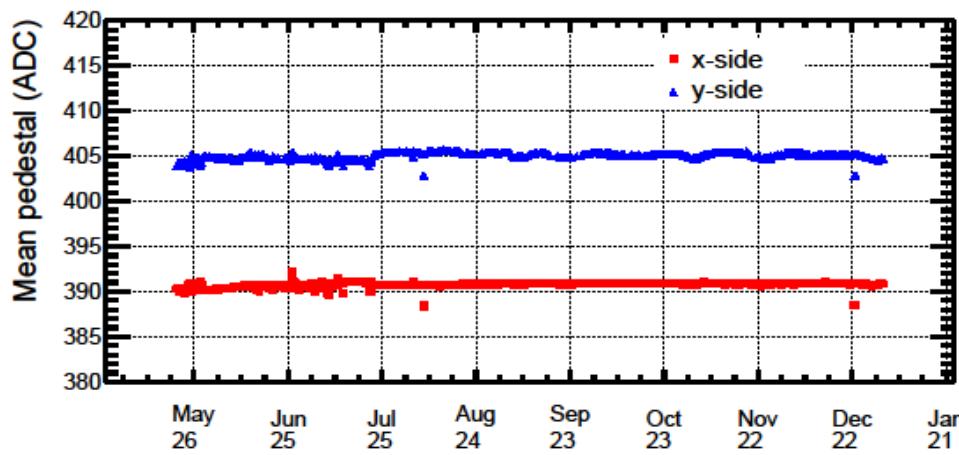
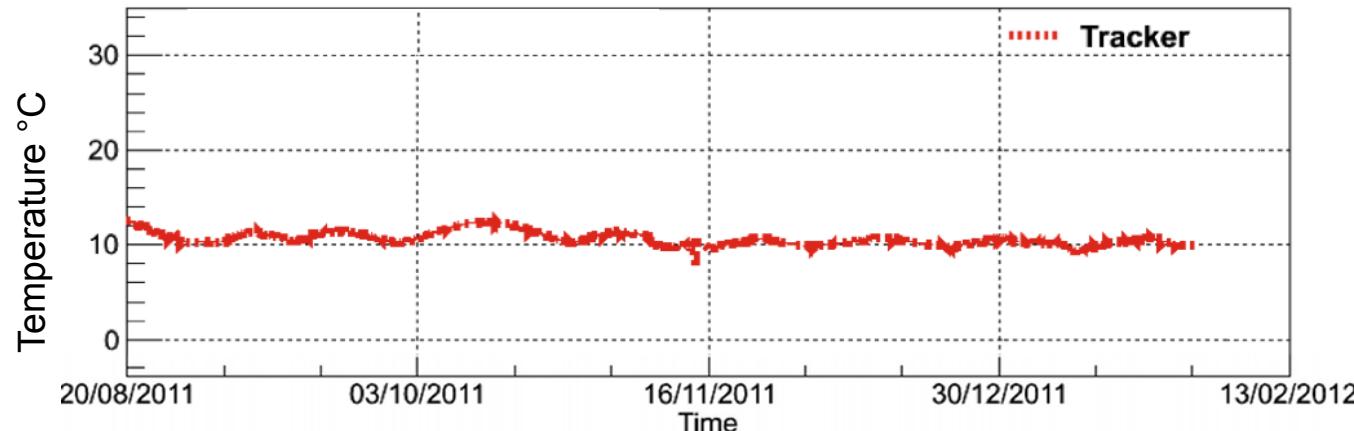
One of the major challenges of operating in space is the extreme thermal environments to which the experiment is exposed.

AMS developed **computers which are programmable from the POCC** for all the monitoring and control. They readout the **1118 temperature sensors** and control **298 heaters**.

The Tracker and TRD have dedicated Thermal Control Systems. Thermal control of the other detectors is done by the Global Thermal Control System.

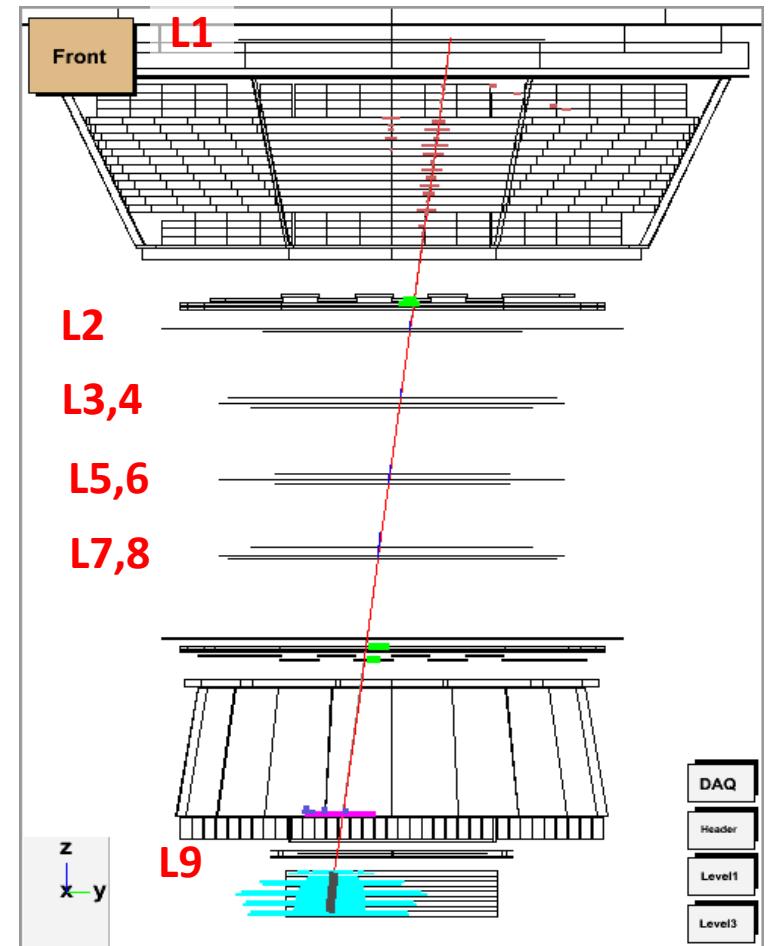
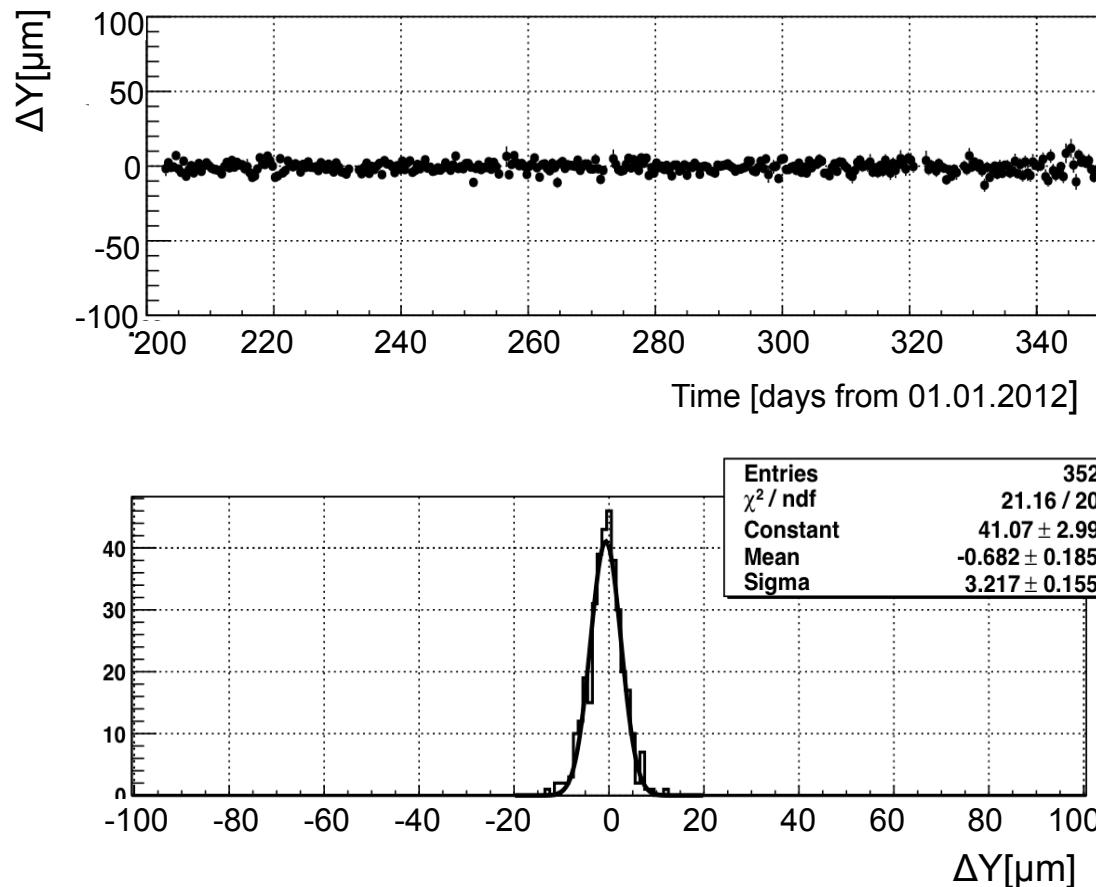
# Tracker on orbit performance

## Temperature effects

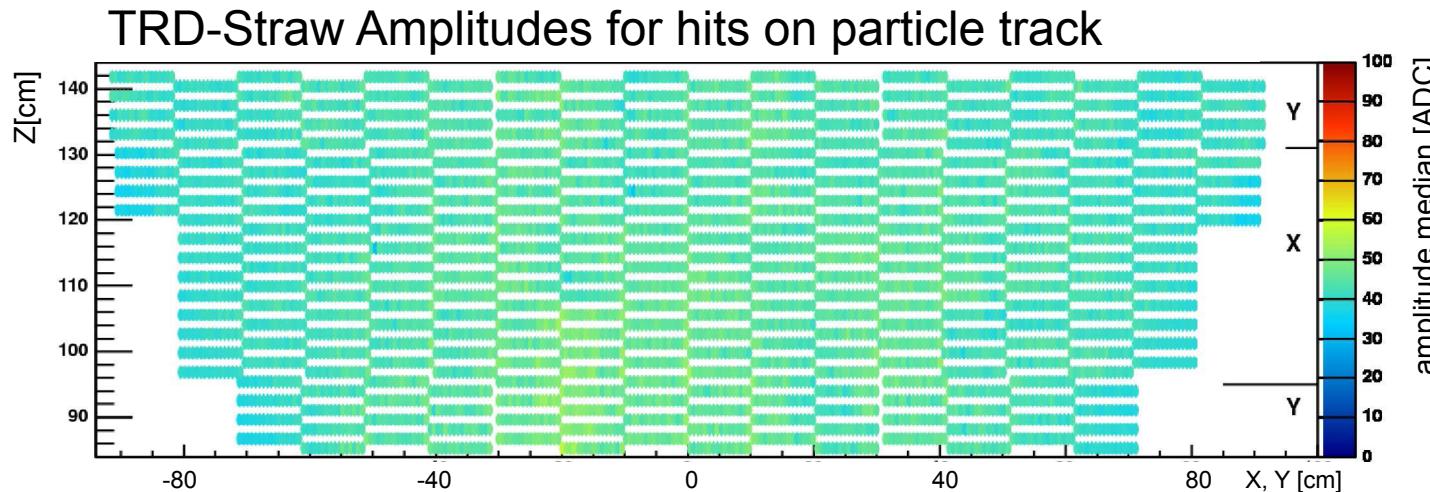


# Time stability of Tracker alignment

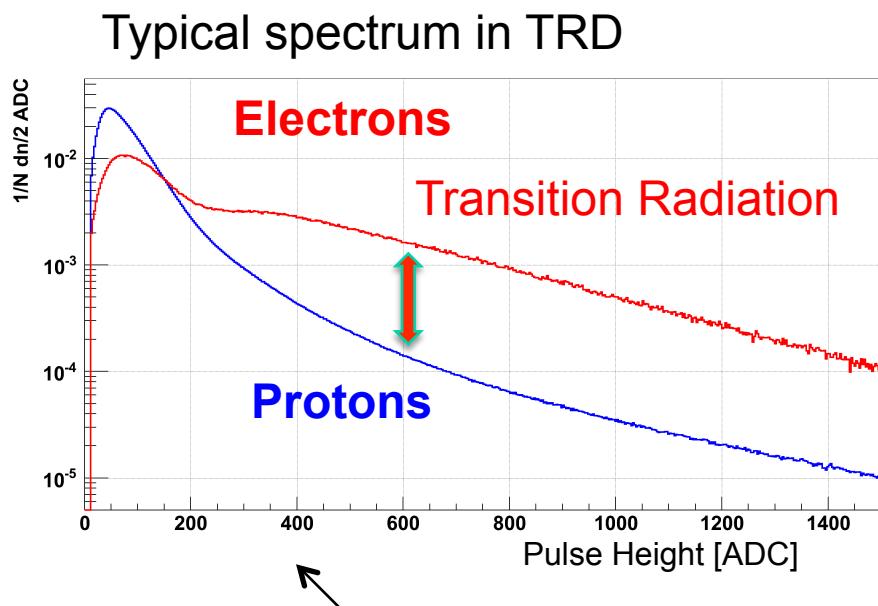
Stability of Layer 1 alignment tracker alignment  
with respect to the inner tracker



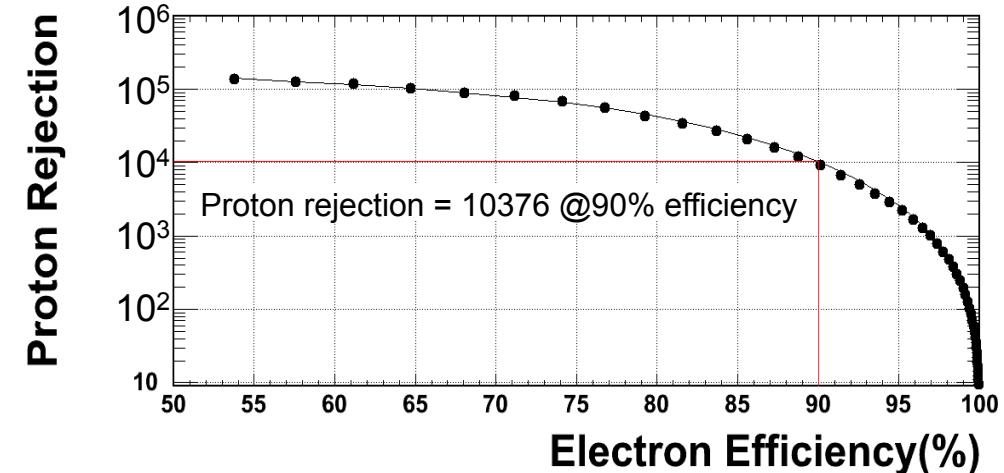
# TRD on orbit performance



All 5248  
channels  
operational

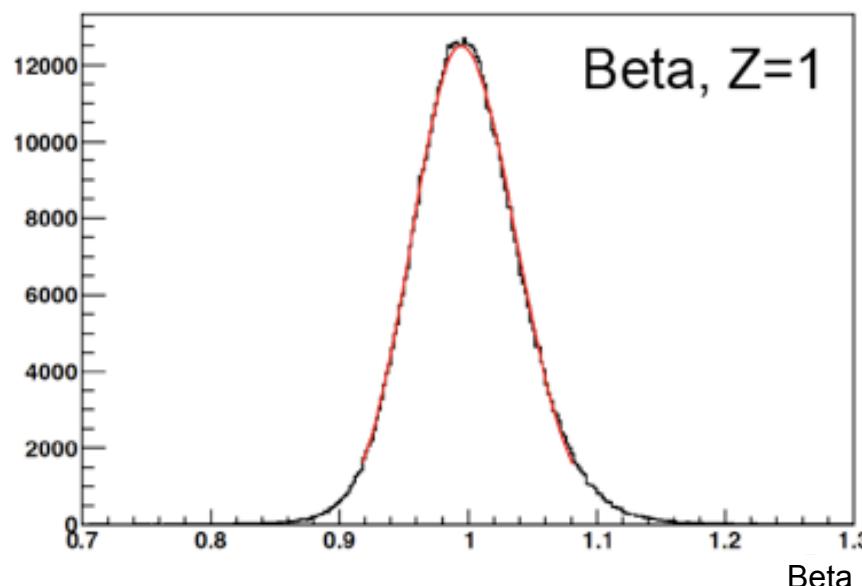
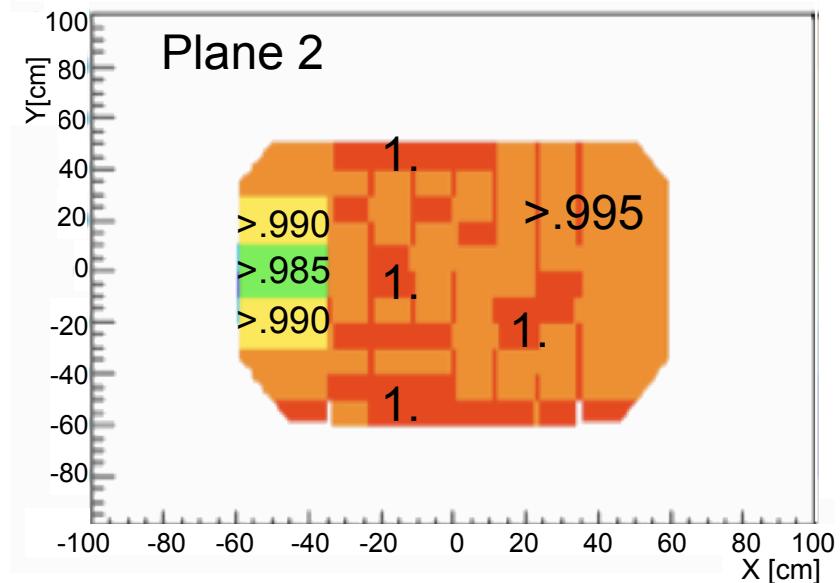
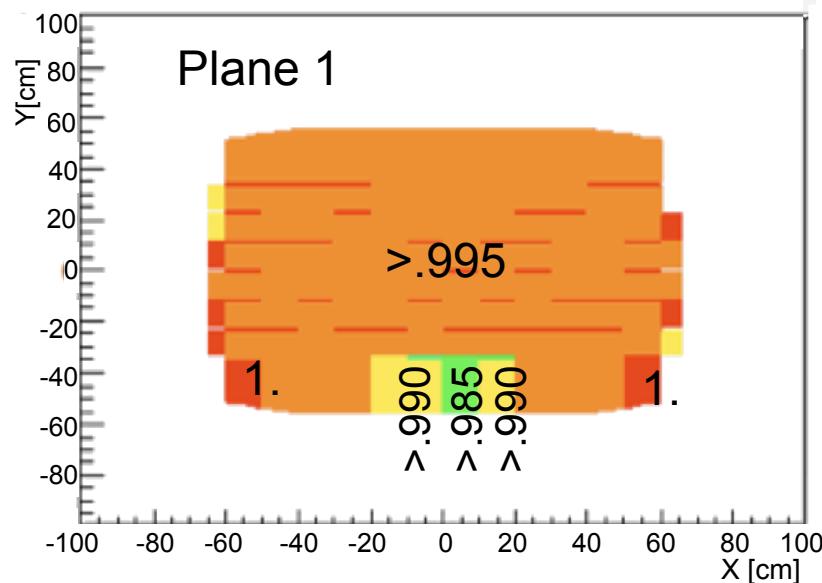


Measurement with 1 of the 20 TRD Layers



# Time of Flight: Trigger & Beta

Single plane efficiencies > 99.5 %

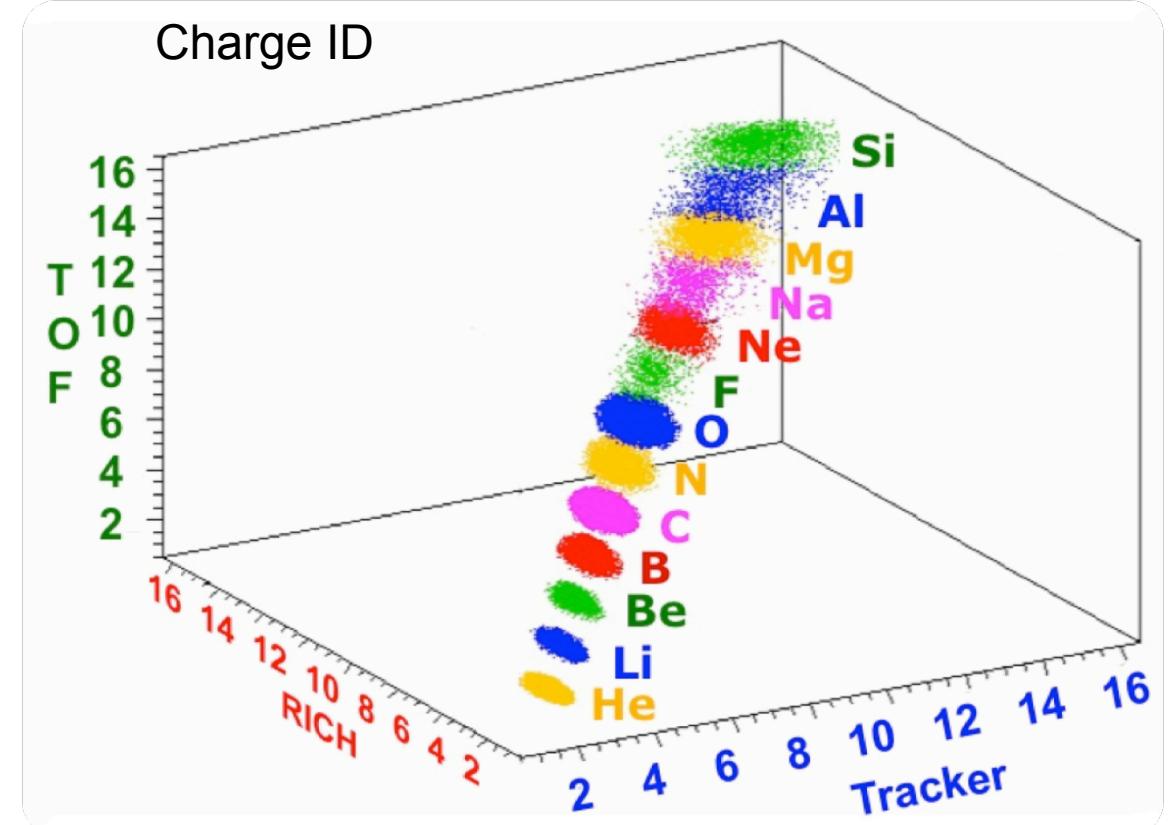
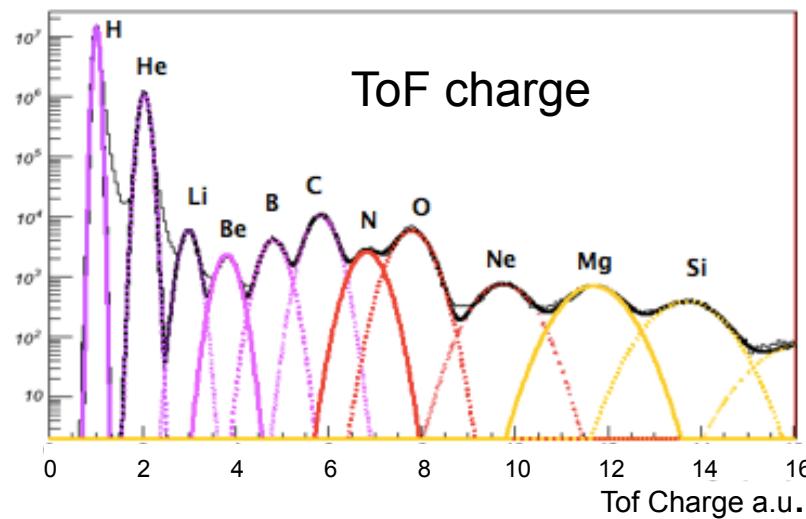
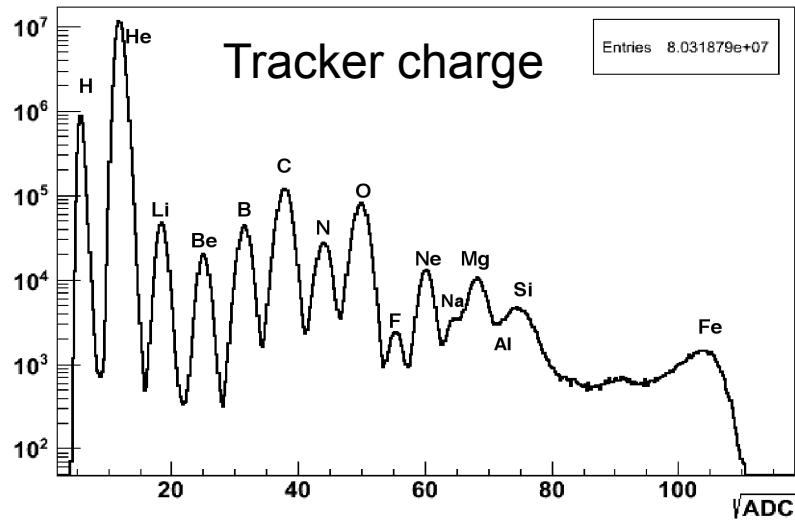


Resolution 4% for protons  
and improves with higher Z

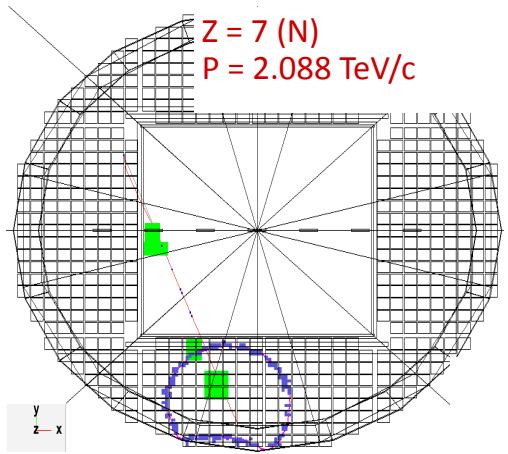
# Charge identification

ToF, Tracker, TRD, ECAL:  
 $dE/dx \approx Z^2$

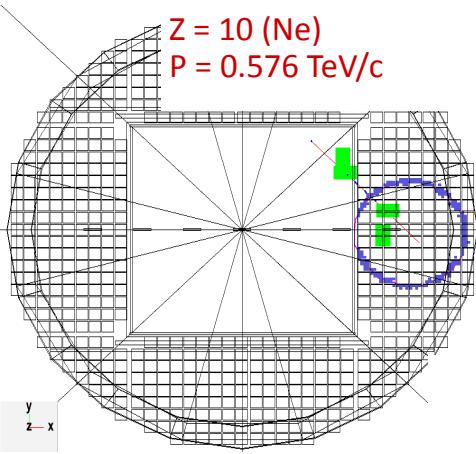
RICH:  
 $Z \approx \#photoelectrons$



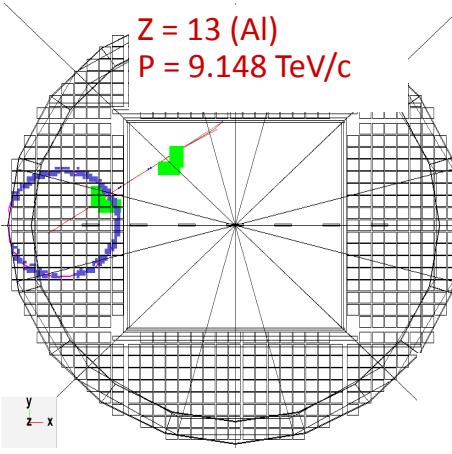
# RICH: Nuclei in the TeV range



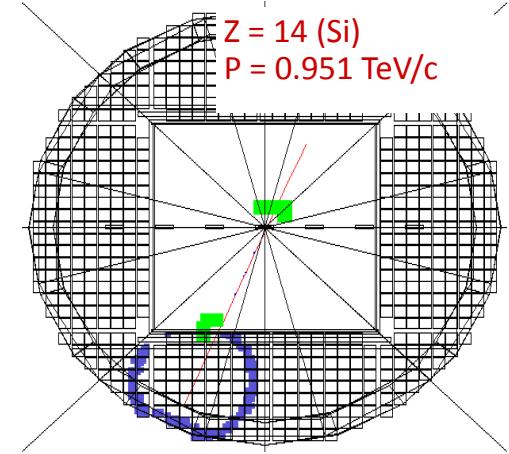
$$\begin{array}{l} Z = 15 \text{ (P)} \\ P = 1.497 \text{ TeV/c} \end{array}$$



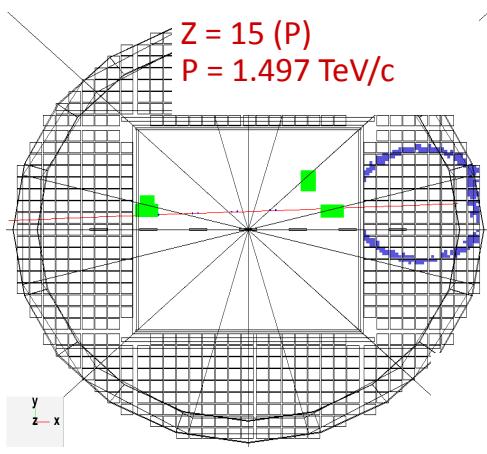
$$E = 16 \text{ (S)}$$



$$\begin{aligned}Z &= 13 \text{ (Al)} \\P &= 9.148 \text{ TeV/c}\end{aligned}$$

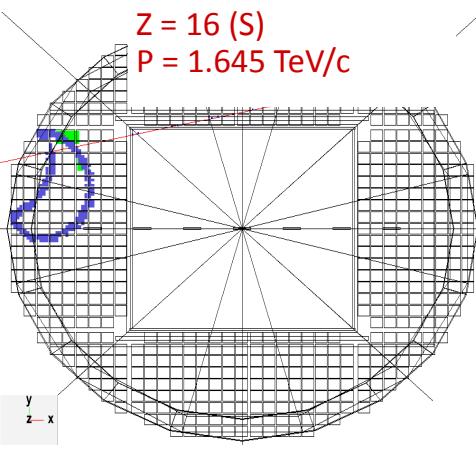


$Z = 14 \text{ (Si)}$   
 $P = 0.951 \text{ TeV/c}$

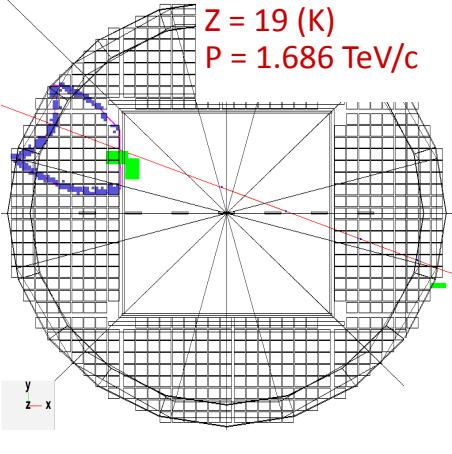


$$Z = 21 \text{ (Sc)}$$

$$P = 0.390 \text{ TeV/c}$$

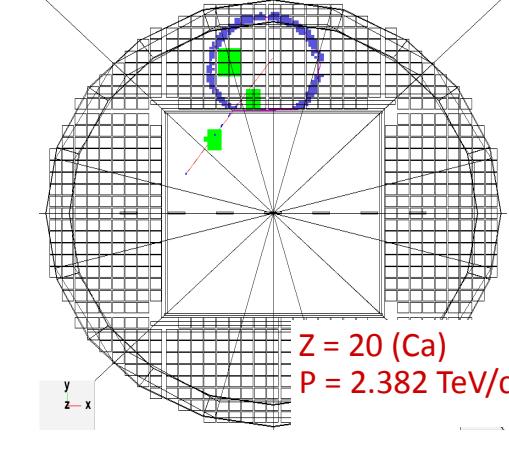


$Z = 22$  (Ti)  
 $P = 1.288 \text{ TeV}/c$

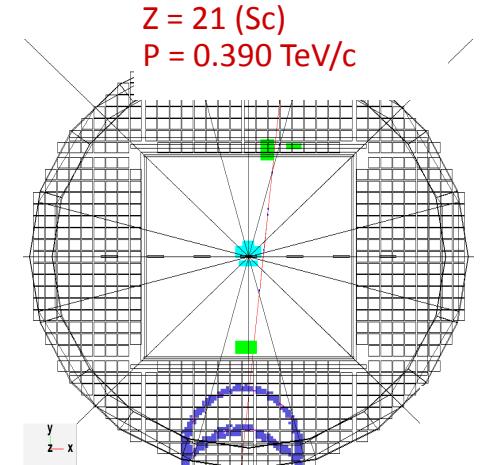




$Z = 19 \text{ (K)}$   
 $P = 1.686 \text{ TeV/c}$

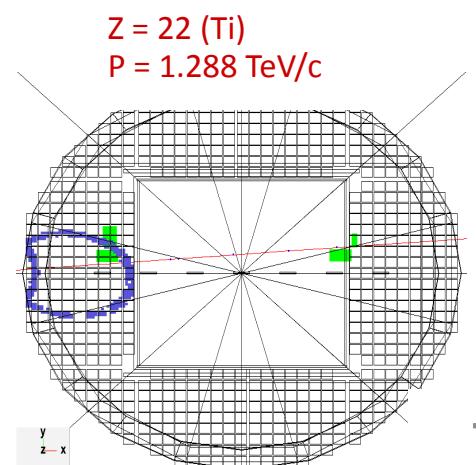


$$Z = 20 \text{ (Ca)} \\ P = 2.382 \text{ TeV/c}$$



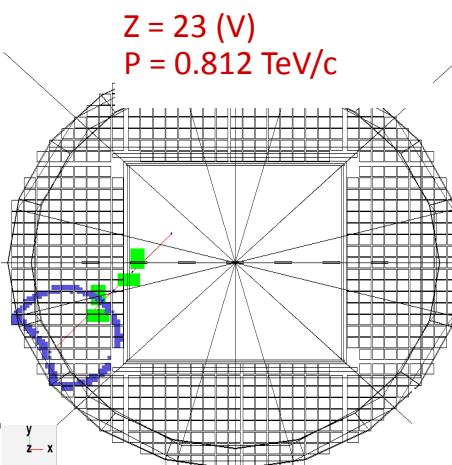
$$Z = 21 \text{ (Sc)}$$

$$P = 0.390 \text{ TeV/c}$$



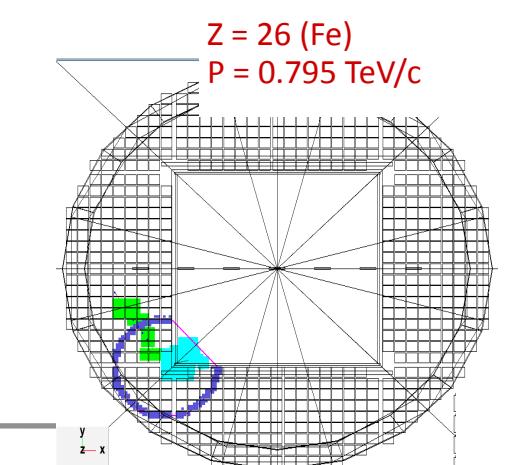
$$Z = 22 \text{ (Ti)}$$

$$P = 1.288 \text{ TeV/c}$$

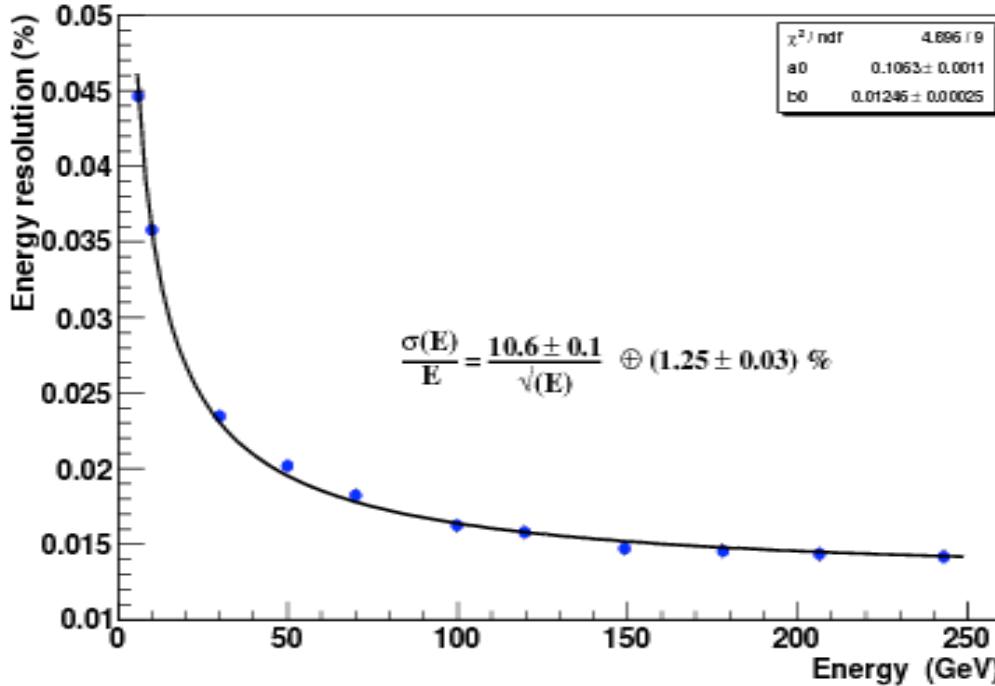


$$Z = 23 \text{ (V)}$$

$$P = 0.812 \text{ TeV/c}$$

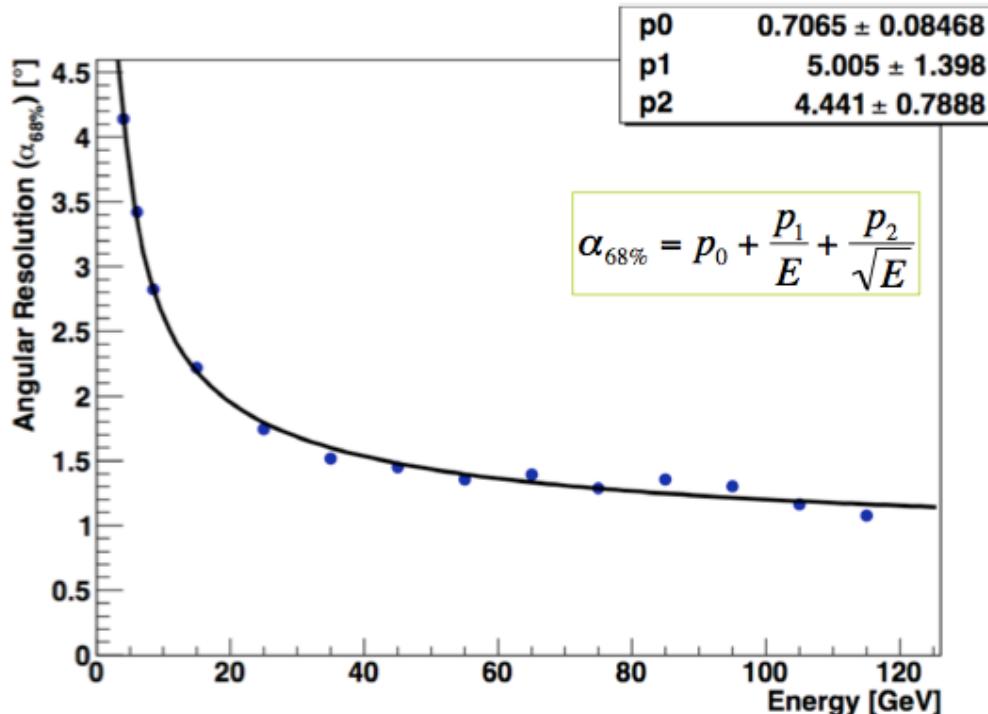


$$\begin{aligned}Z &= 26 \text{ (Fe)} \\P &= 0.795 \text{ TeV/c}\end{aligned}$$



Energy Resolution:  
Test Beam

Better than 2% @ 100GeV



Angular Resolution:  
Flight Data

evaluated with electrons, comparing  
tracker track with ecal shower axis

Better than 1.5 deg for E>40GeV

# Calibration & analysis

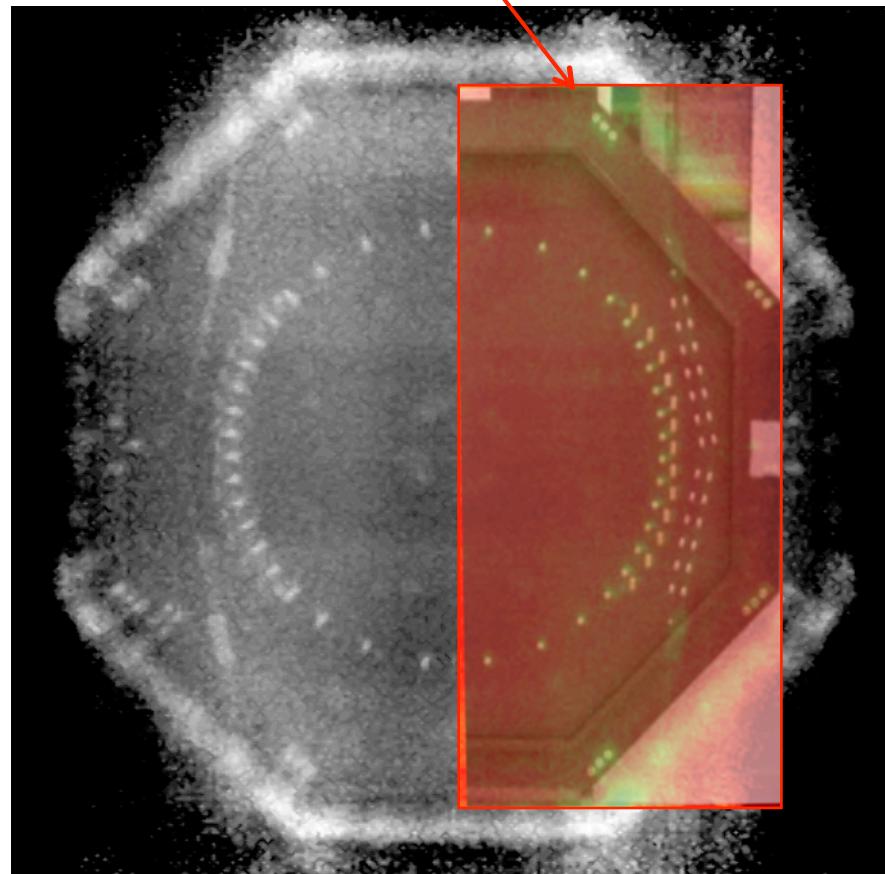
- An intense effort on detector calibration has been developed during the past year
- Important results on stability, alignment, energy calibration have been achieved: the detectors operate smoothly with excellent performances, sometimes better than measured on earth

# Tomography of support plane

Dishomogeneity of the He/p ratio

He “missing” particles extrapolated to the first mechanical Tracker support plane

Photo of the support plane



# Conclusions

- AMS02 is in orbit since May 16<sup>th</sup> 2011
- No damage due to the launch stress or to the space environment, all the systems are working in both the primary and redundant part
- All the detectors are properly functioning with DAQ in nominal conditions since May 19<sup>th</sup> 2011
- Ground operations (POCC and SOC) run smoothly
- Detector calibration procedures are completed and constant monitoring will continue
- 10-20 years on board the ISS: great discovery potential

Thank you

