

The history of cosmic ray studies after Hess

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History

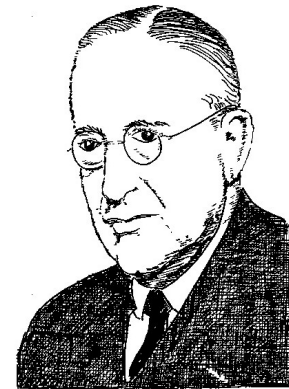
The golden years

A few words on accelerators

Renaissance of cosmic rays

Accelerators in the sky

A start into a new era



History

1895 Wilhelm Conrad Röntgen

Discovery of X rays

1896 Henri Antoine Becquerel

Discovery of radioactivity

1912 Victor Franz Hess

Discovery of cosmic rays

Particles or rays?

The New York Times

VOL. LXXXII...No. 27,370.

December 31, 1932

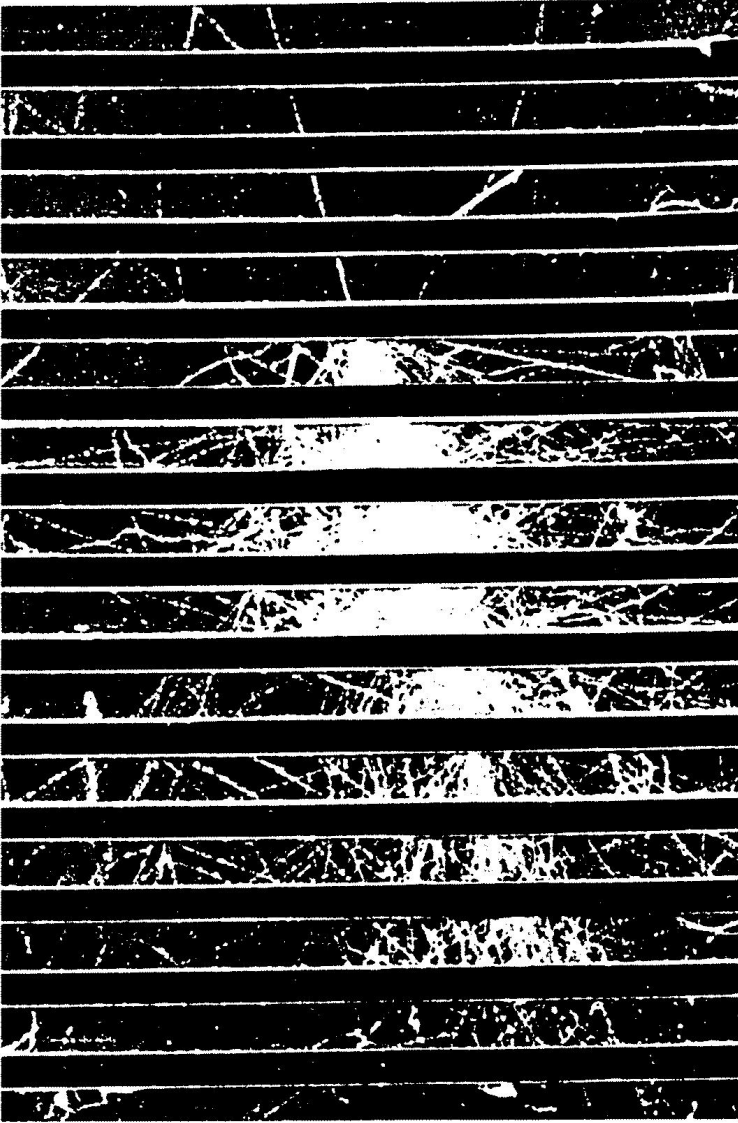
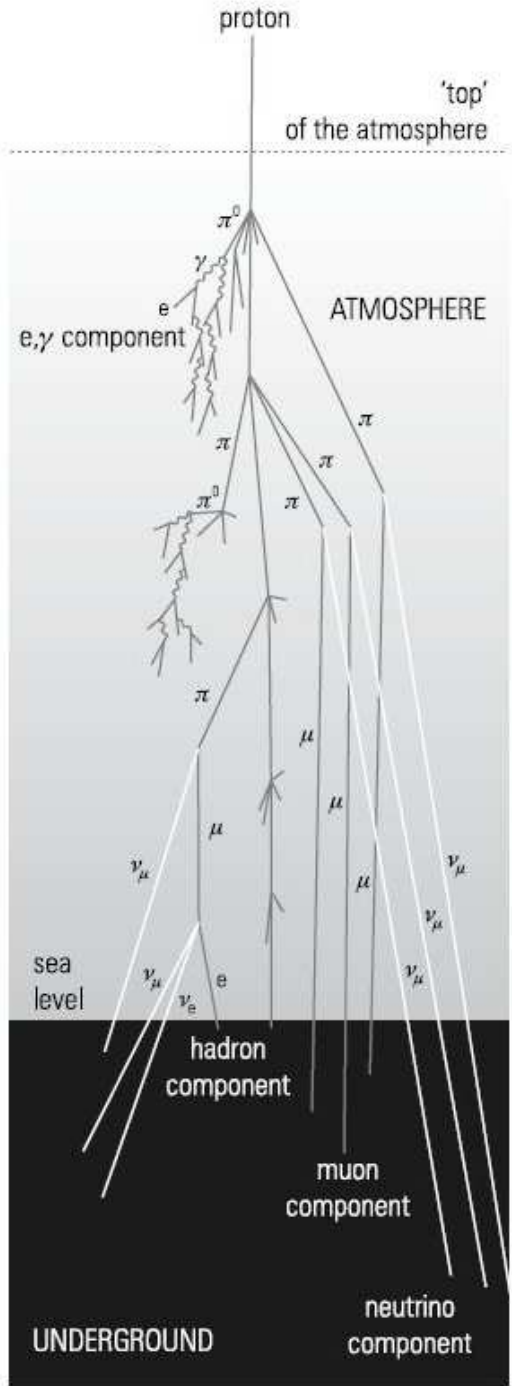
MILLIKAN RETORTS HOTLY TO COMPTON IN COSMIC RAY CLASH

Debate of Rival Theorists
Brings Drama to Session
of Nation's Scientists.

THEIR DATA AT VARIANCE

New Findings of His Ex-Pupil
Lead to Thrust by Millikan
at 'Less Cautious' Work.

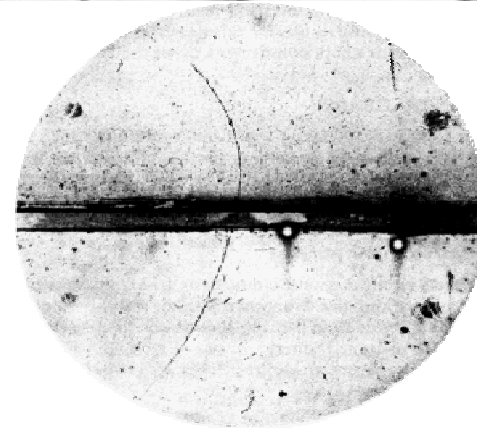
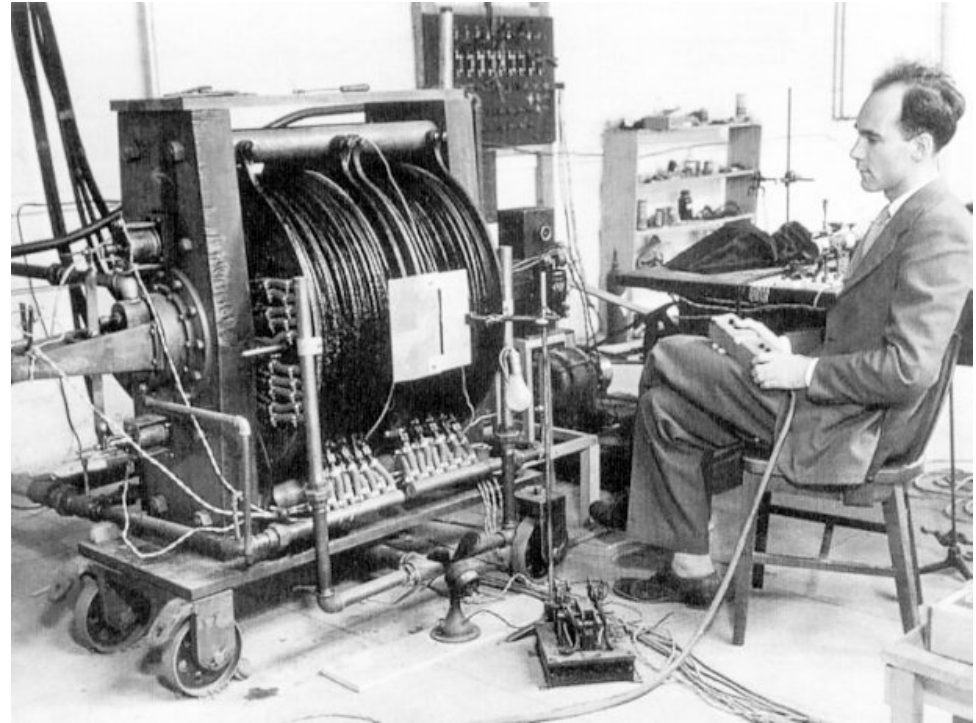
Cosmic ray induced cascade



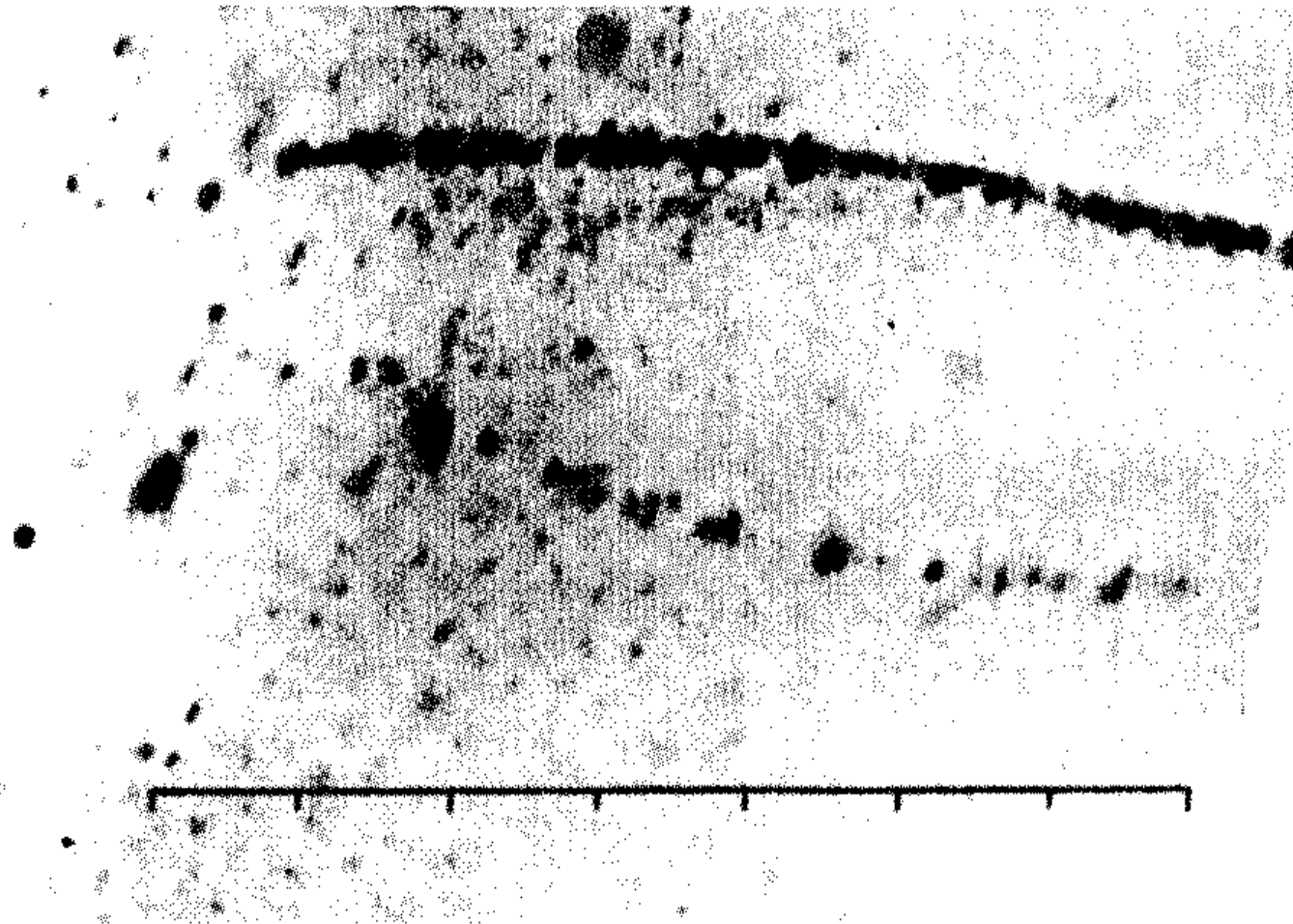
3027 m
altitude;
~10 GeV
proton

Fretter
1949

The golden years: discovery of the positron 1932

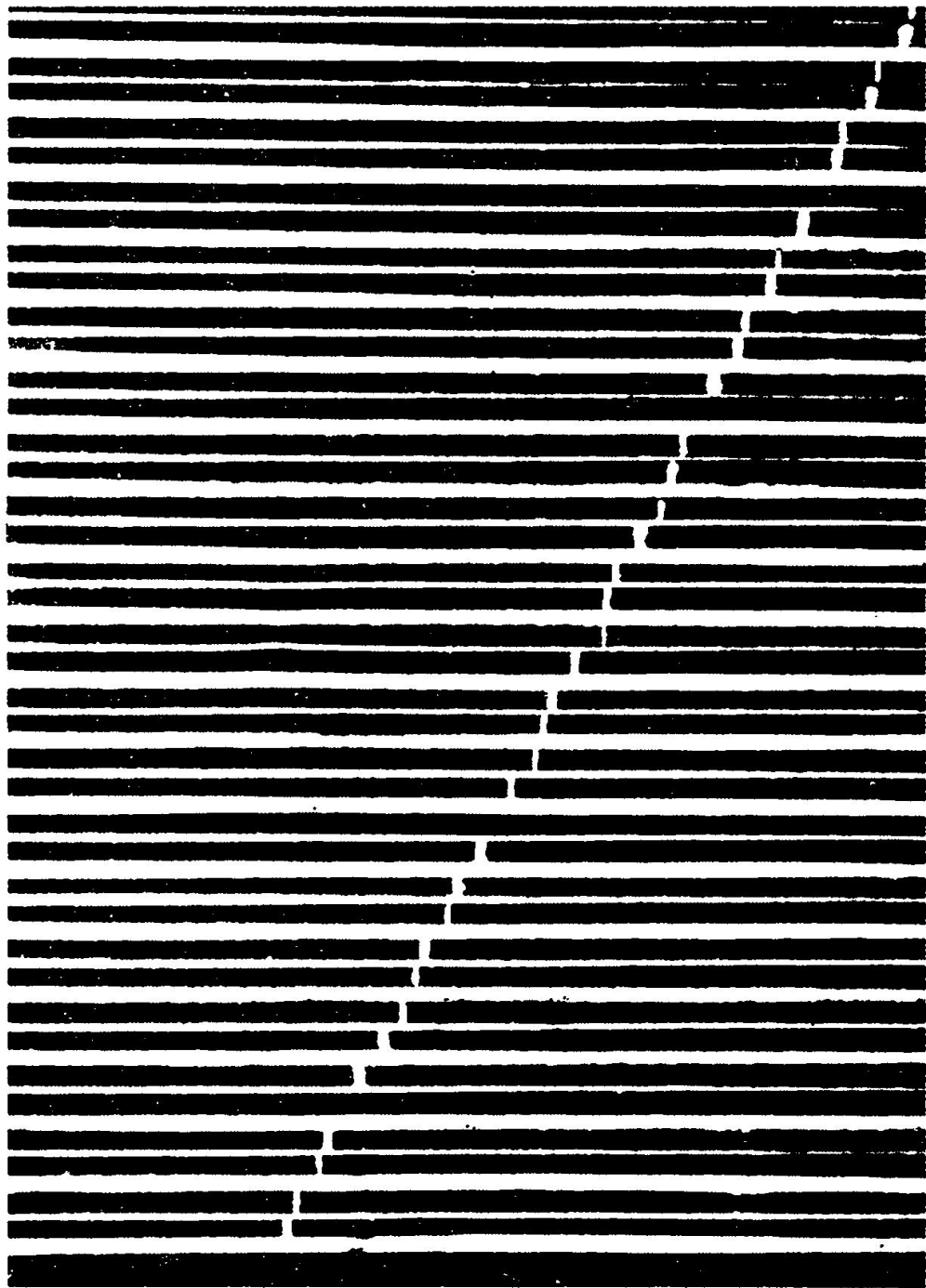


First indication of the muon by
Paul Kunze
in Rostock in 1932



muon

electron

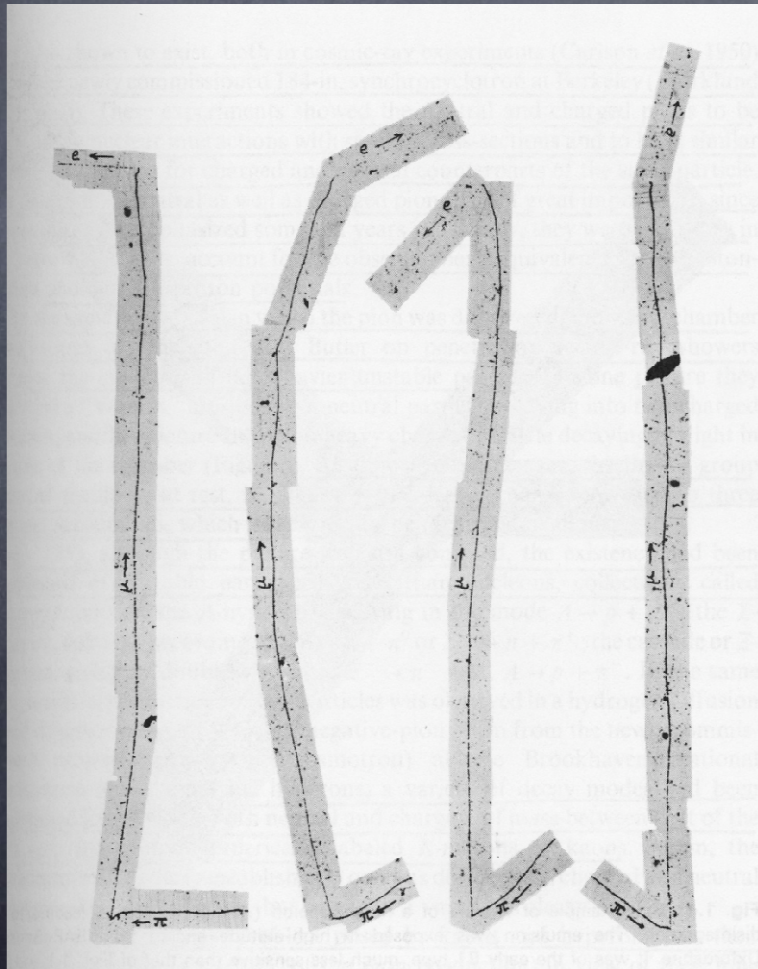


Discovery of the muon
by Anderson and
Neddermeyer 1937

Muon in a multiplate
spark chamber
(V.S. Kaftanov 1963)

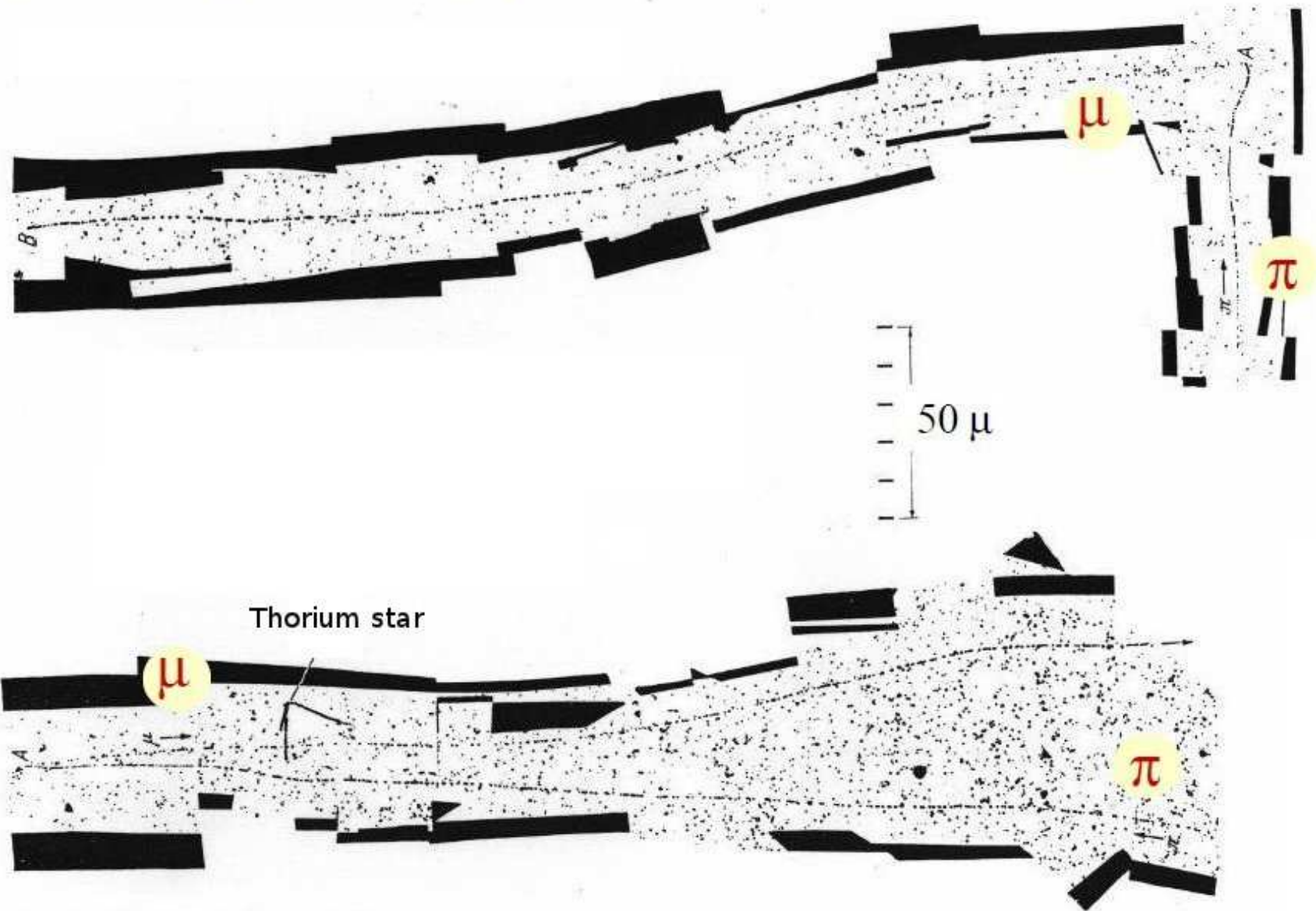
Pion discovery 1947: Perkins, Powell, Occhialini, Lattes and Muirhed

Nobel Prize 1950

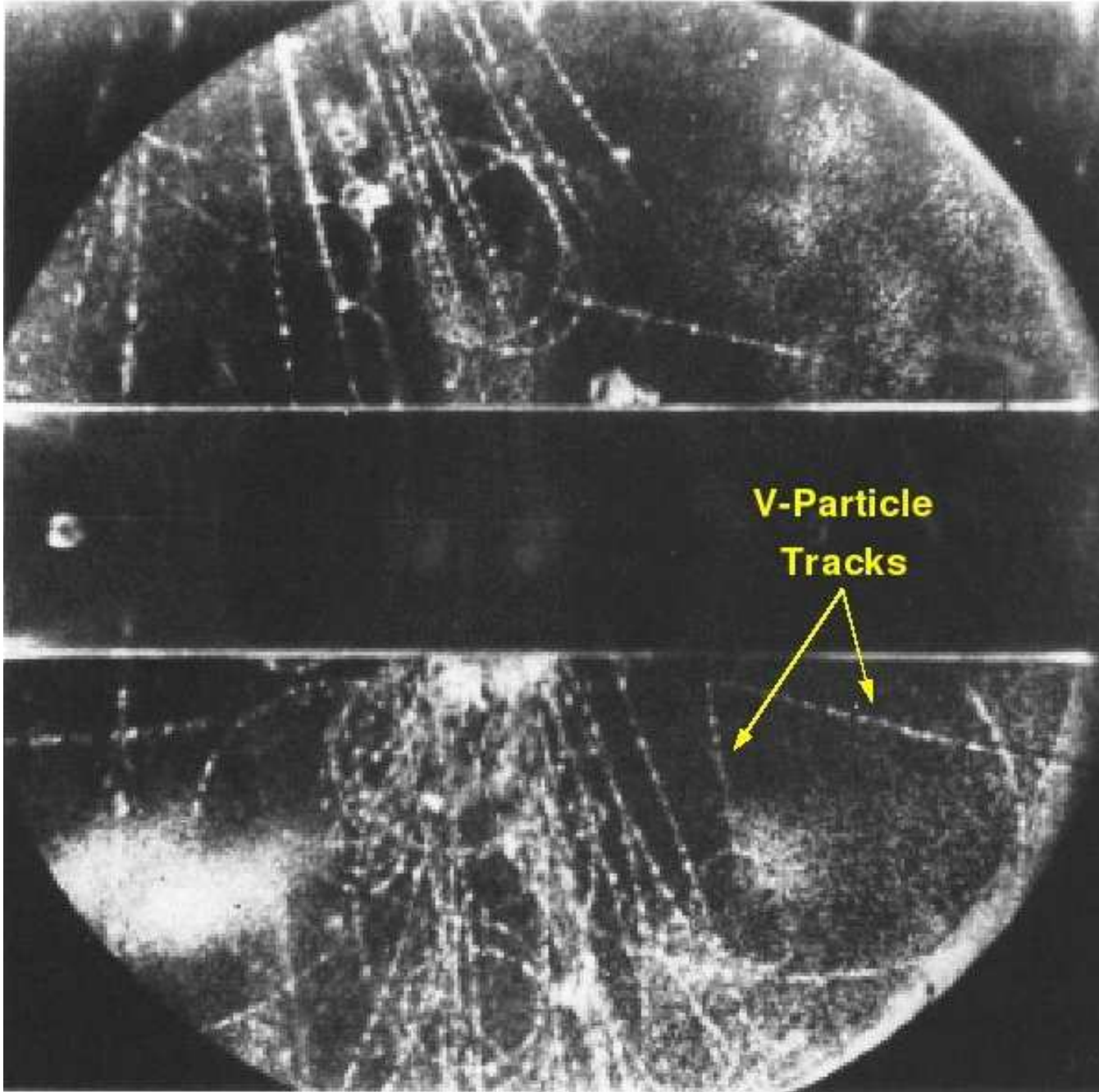


Powell

Discovery of $\pi \rightarrow \mu + \nu_\mu$ (C.F.Powell : 1947)

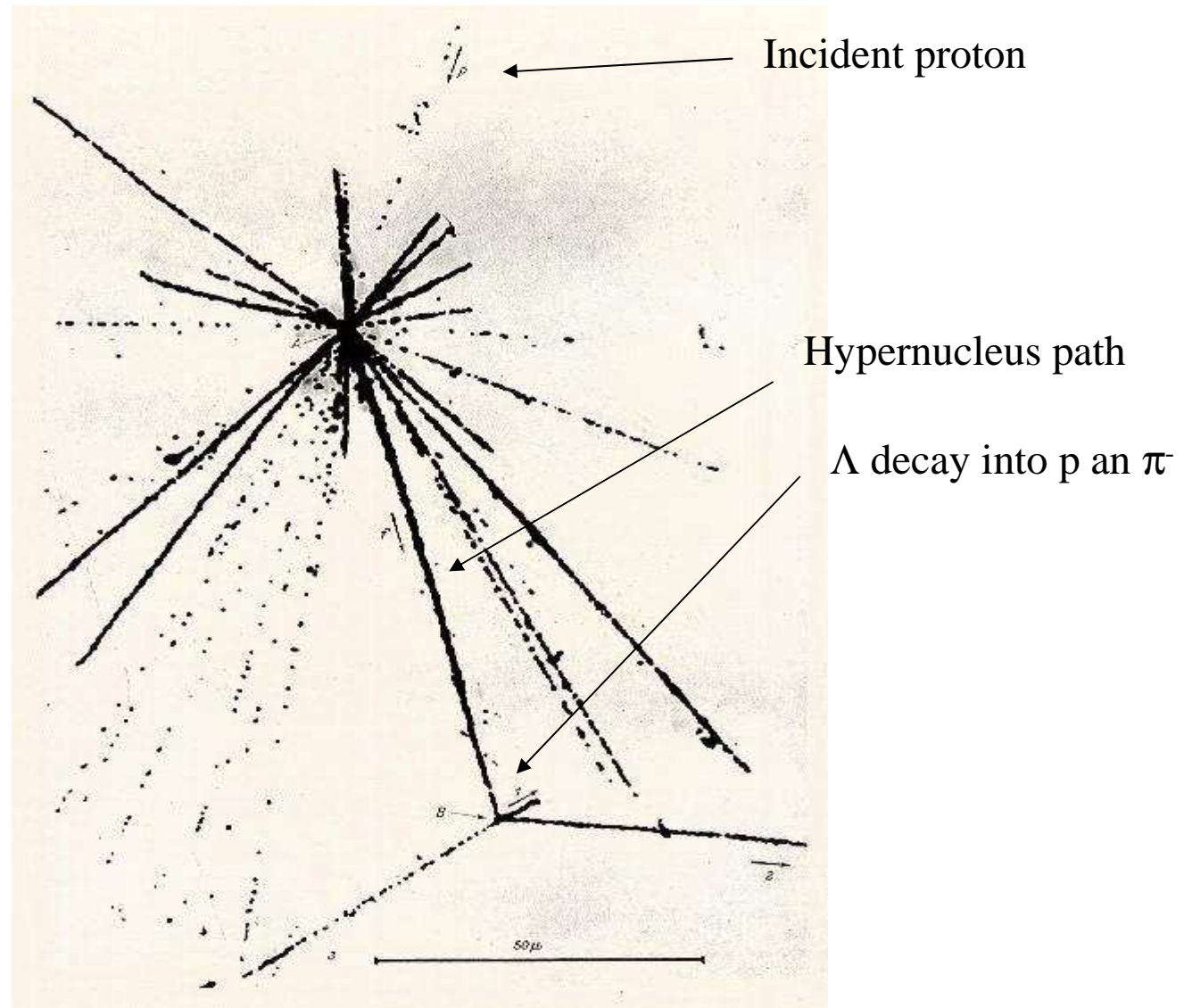


LATTES, MUIRHEAD, OCCHIALINI
and POWELL; Nature 159, 694 (1947).



Neutral Kaon
Discovery
1947
Rochester and
Butler

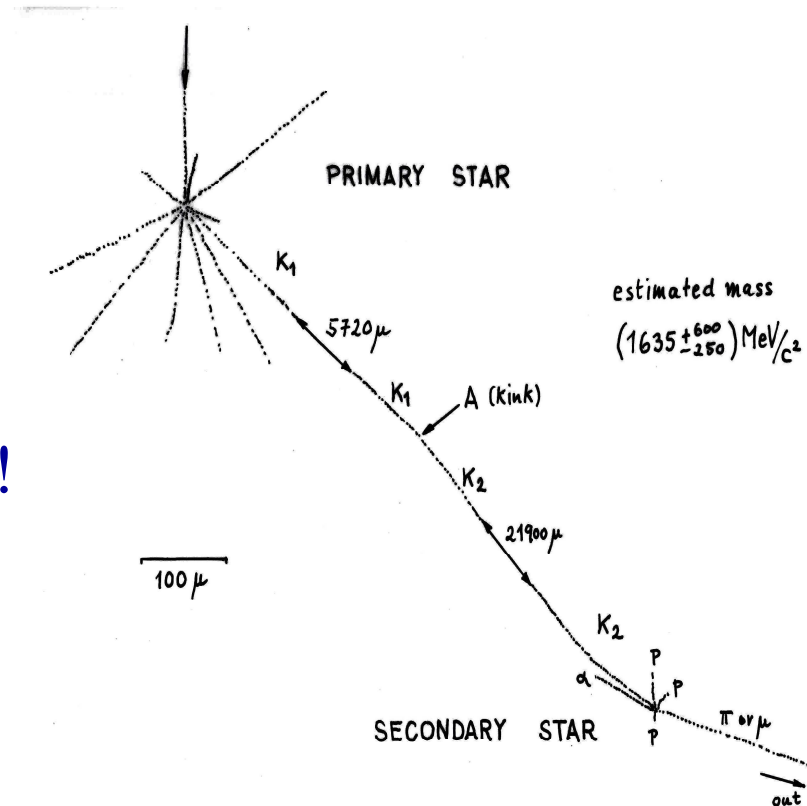
Observation of a Λ in cosmic rays



M. Danysz and J. Pniewski Phil. Mag. 44 (1953) 348

The missing link in the quark model

It is believed that an unidentified track found in 1954 in a stack of nuclear emulsions exposed to cosmic rays at 100 000 ft altitude by Yehuda Eisenberg was the path of an Omega minus! M. Gell-Mann: “Perhaps it (the Omega minus) could explain the old Eisenberg event ...”



Alvarez:

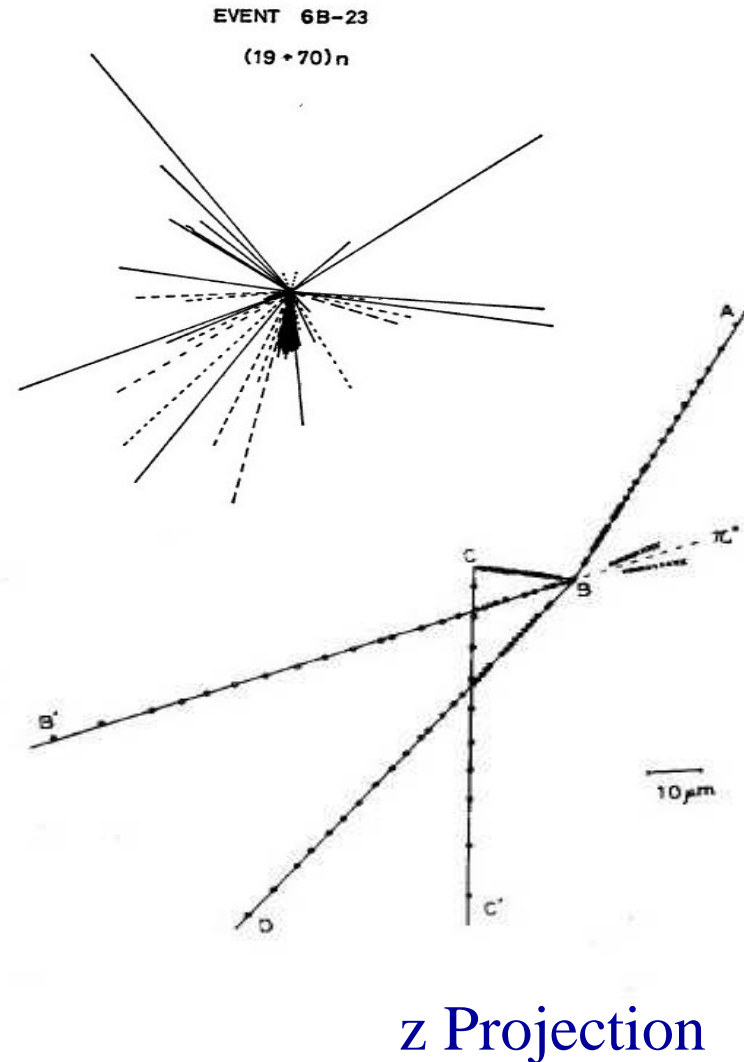
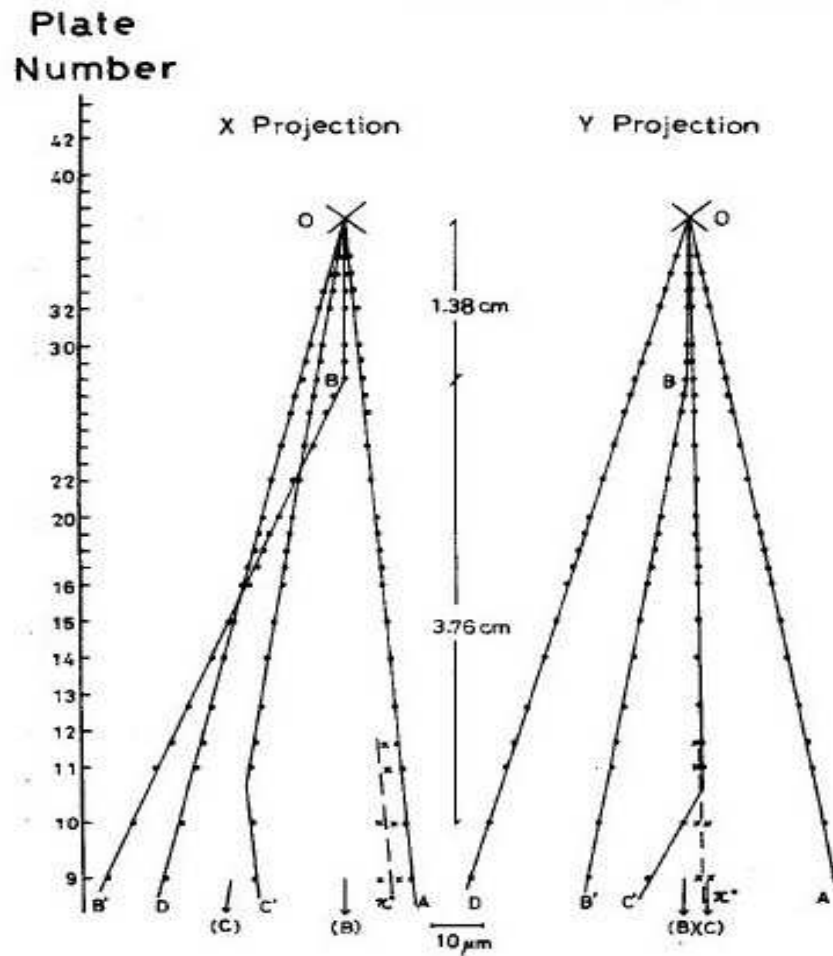
Y. Eisenberg; Phys. Rev. 96 (1954) 541

the Ω interacted with an Ag nucleus to give $K^- \Xi \text{ Ag}$.

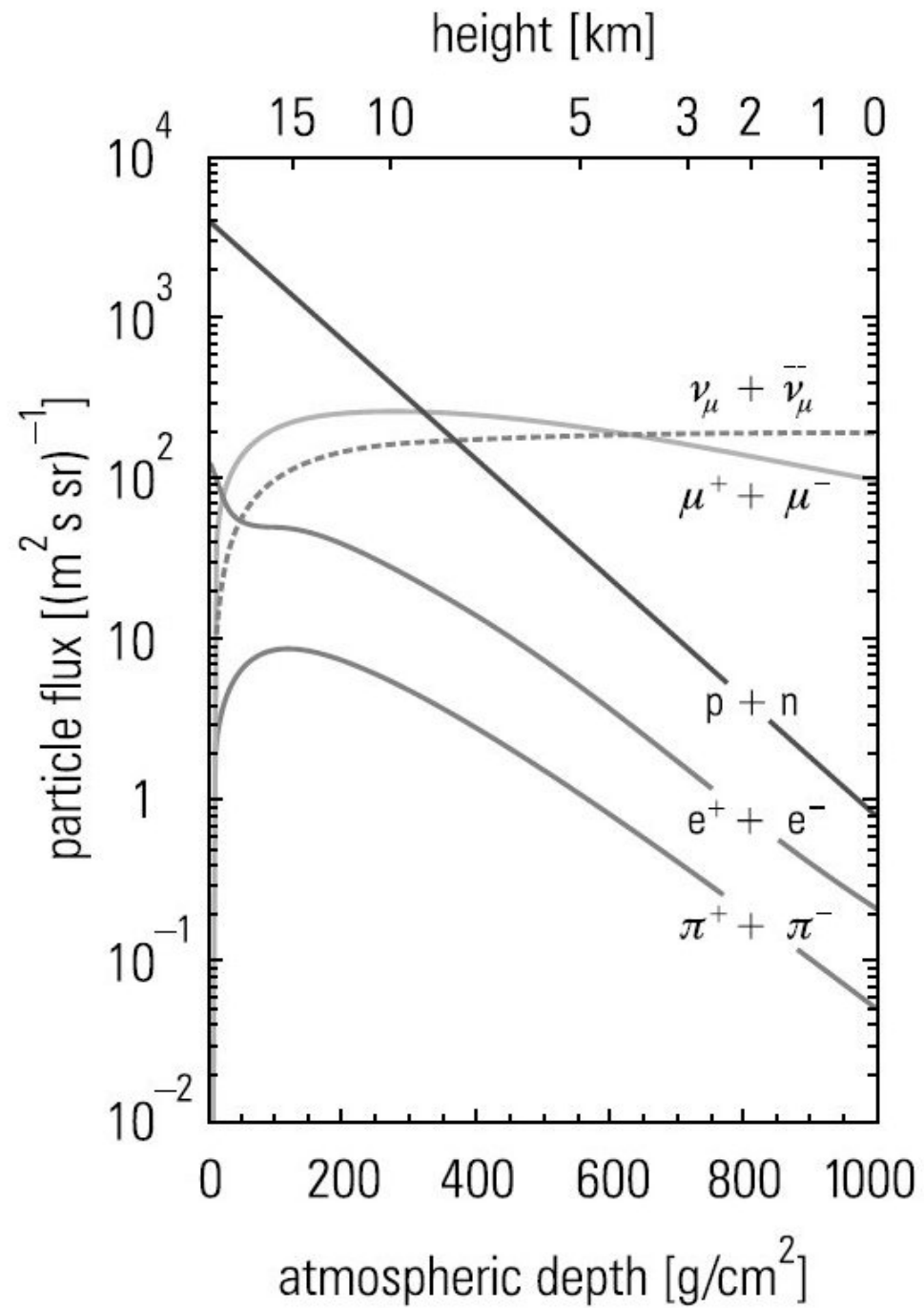
Discovery of X-particle (K.Niu : 1971)

Emulsion stack on a jet cargo air plane;
Estimated mass ~ 2 GeV; short-lived

The charm ?!!



Transformation in the atmosphere



Accelerators and Storage Rings take over

Electron neutrino:

Cowan and Reines 1956

Muon neutrino :

Lederman, Schwartz, Steinberger 1962

Tau and Tau neutrino

Perl 1975

Gluon, DESY 1979 to name a few

Three Generations of Matter (Fermions)

	I	II	III	
mass →	2.4 MeV/c ²	1.27 GeV/c ²	171.2 GeV/c ²	0
charge →	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$	0
spin →	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1
name →	u up	c charm	t top	γ photon
Quarks	4.0 MeV/c ²	104 MeV/c ²	4.2 GeV/c ²	0
	$-\frac{1}{3}$	$-\frac{1}{3}$	$-\frac{1}{3}$	0
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1
	d down	s strange	b bottom	g gluon
Leptons	<2.2 eV/c ²	<0.17 MeV/c ²	<15.5 MeV/c ²	91.2 GeV/c ²
	0	0	0	0
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1
	ν_e electron neutrino	ν_μ muon neutrino	ν_τ tau neutrino	Z⁰ Z boson
	0.511 MeV/c ²	105.7 MeV/c ²	1.777 GeV/c ²	80.4 GeV/c ²
	-1	-1	-1	± 1
	$\frac{1}{2}$	$\frac{1}{2}$	$\frac{1}{2}$	1
	e electron	μ muon	τ tau	W[±] W boson

Gauge Bosons

Renaissance of Cosmic Rays

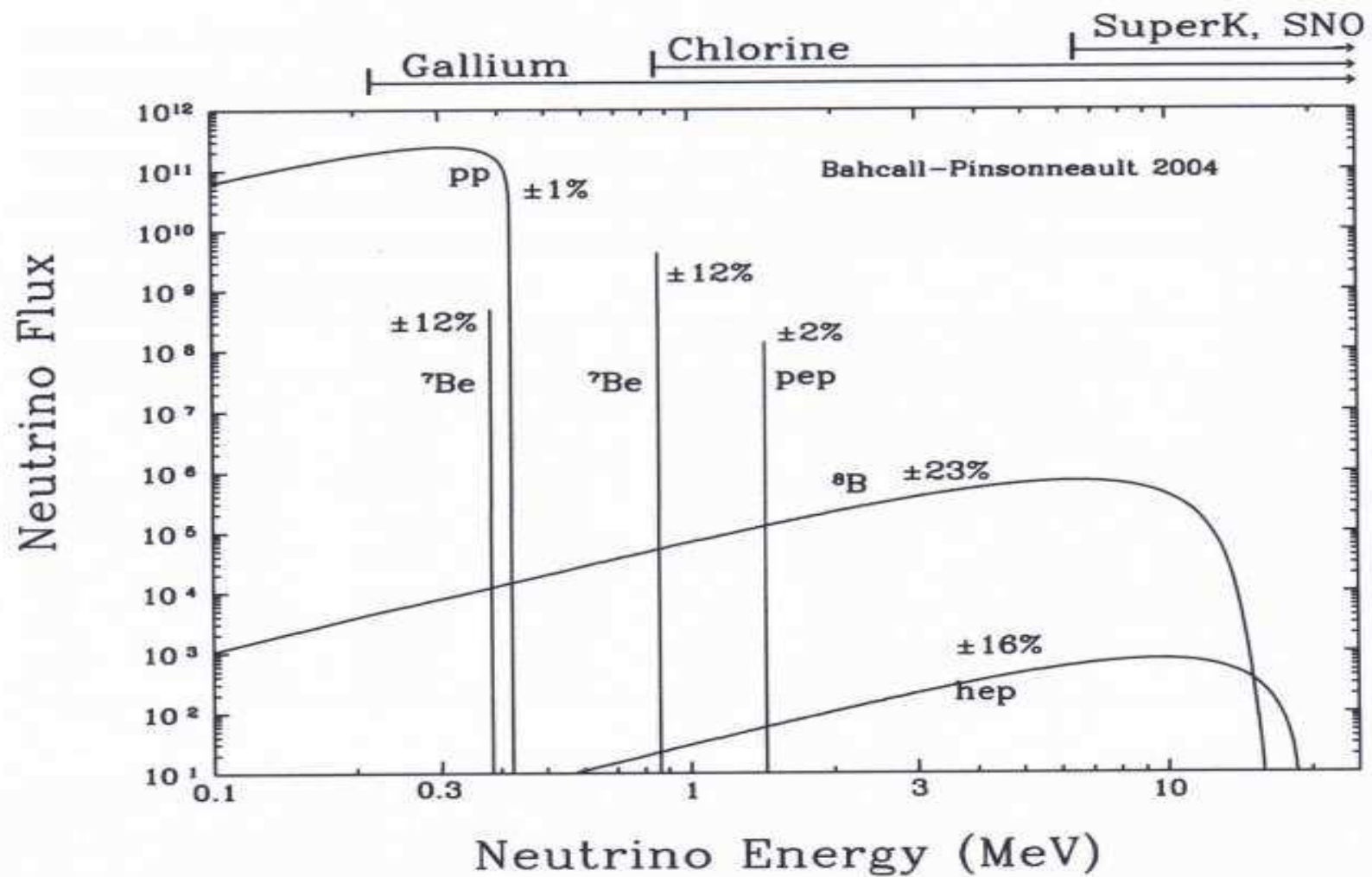
Ray Davis jun.



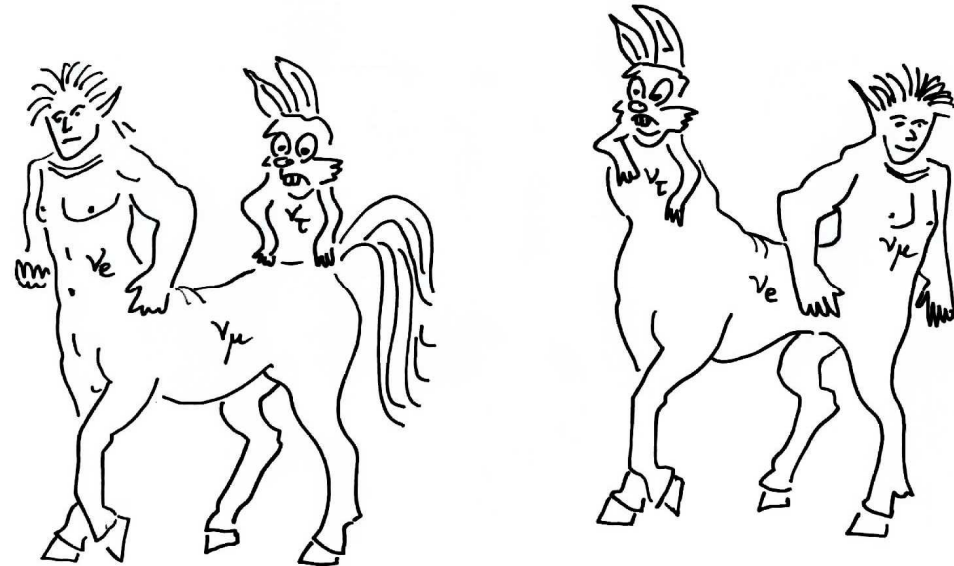
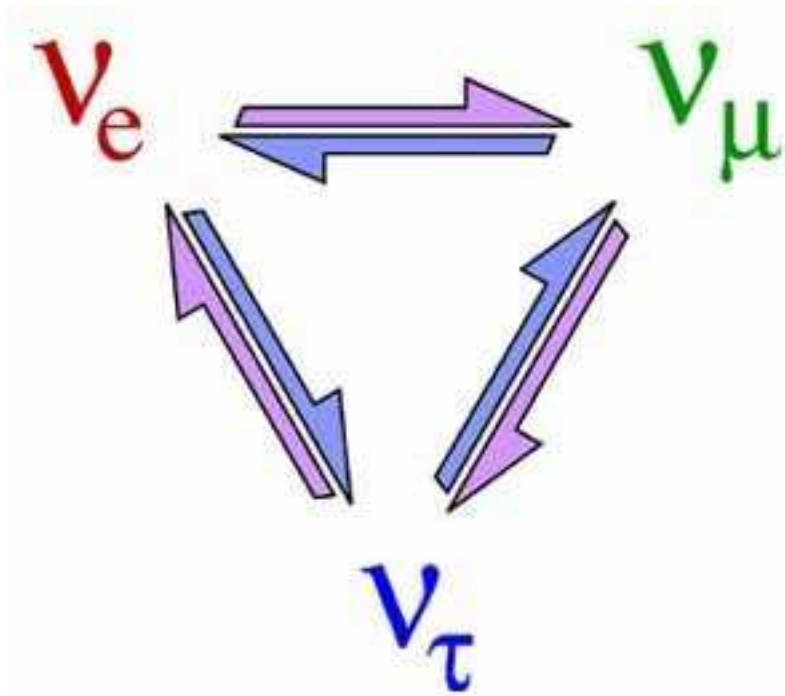
from 1967



Solar Neutrinos



Neutrino-Oscillations



animalistic oscillations

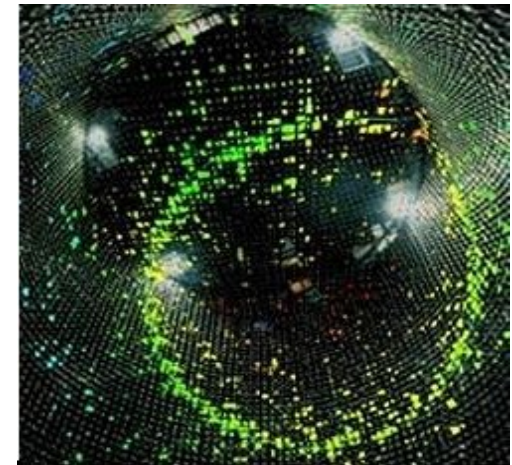
Dependent on the mixing angle and the difference of the masses squared

Super-Kamiokande Experiment

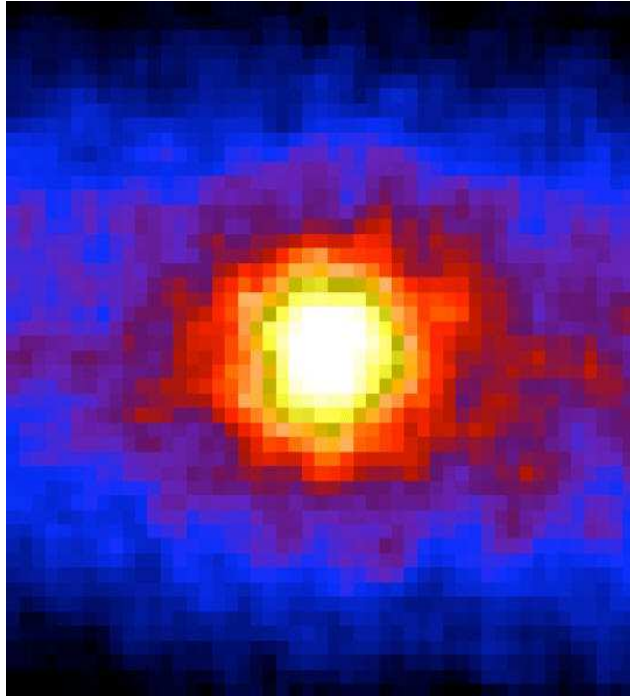


Masatoshi Koshiba

Cherenkov-Ring



Sun in the light of neutrinos



Supernova 1987 A



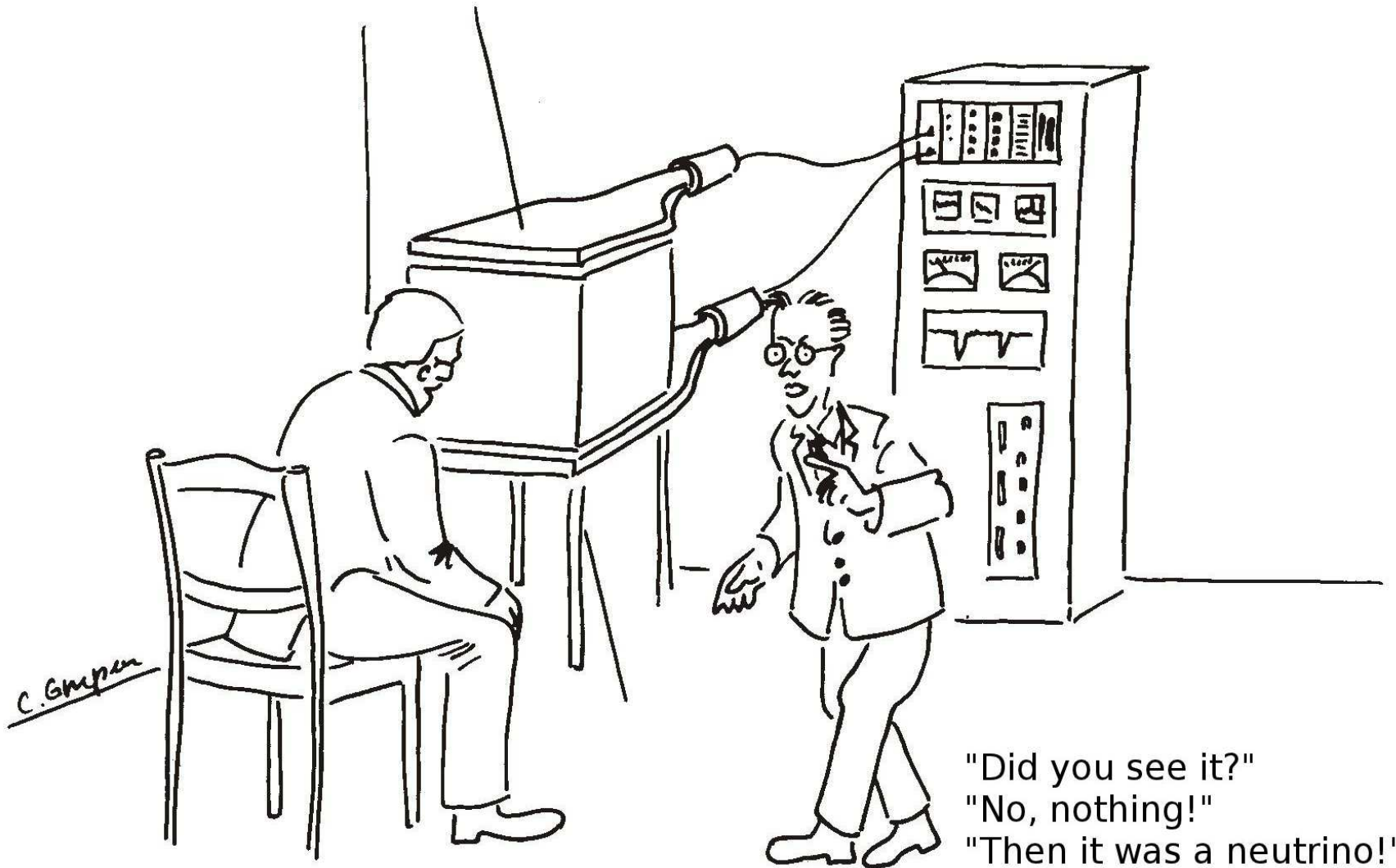
Supernova 1987 A

Supernova-Explosion
in the Large Magellanic Cloud
Distance 170 000 light years

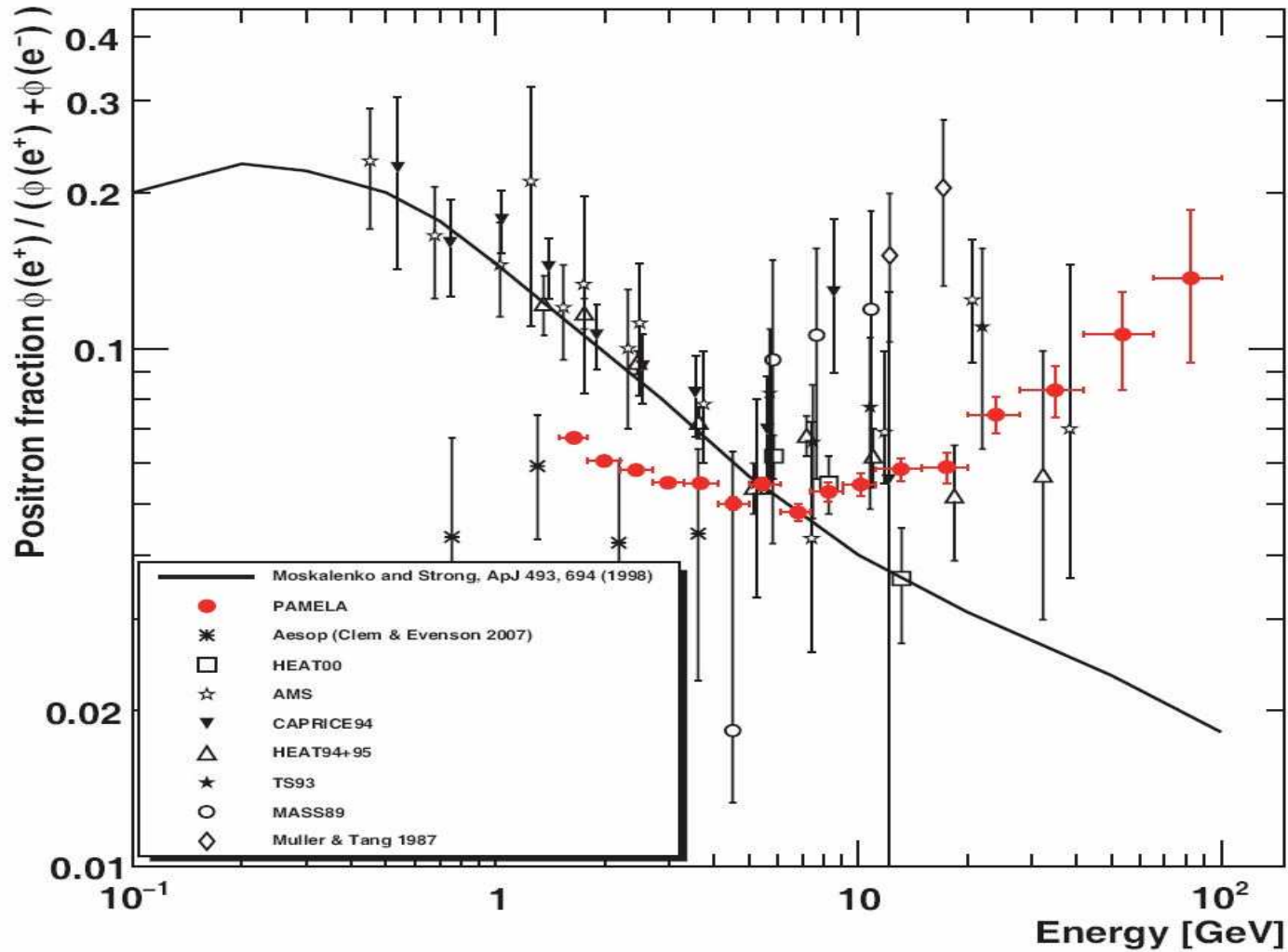
energy output 6×10^{46} Joule
(energy consumption 10^{21} Joule per year
for the whole world)

10^{58} neutrinos, only 19 measured in IMB and
Kamiokande

Problems to measure neutrinos



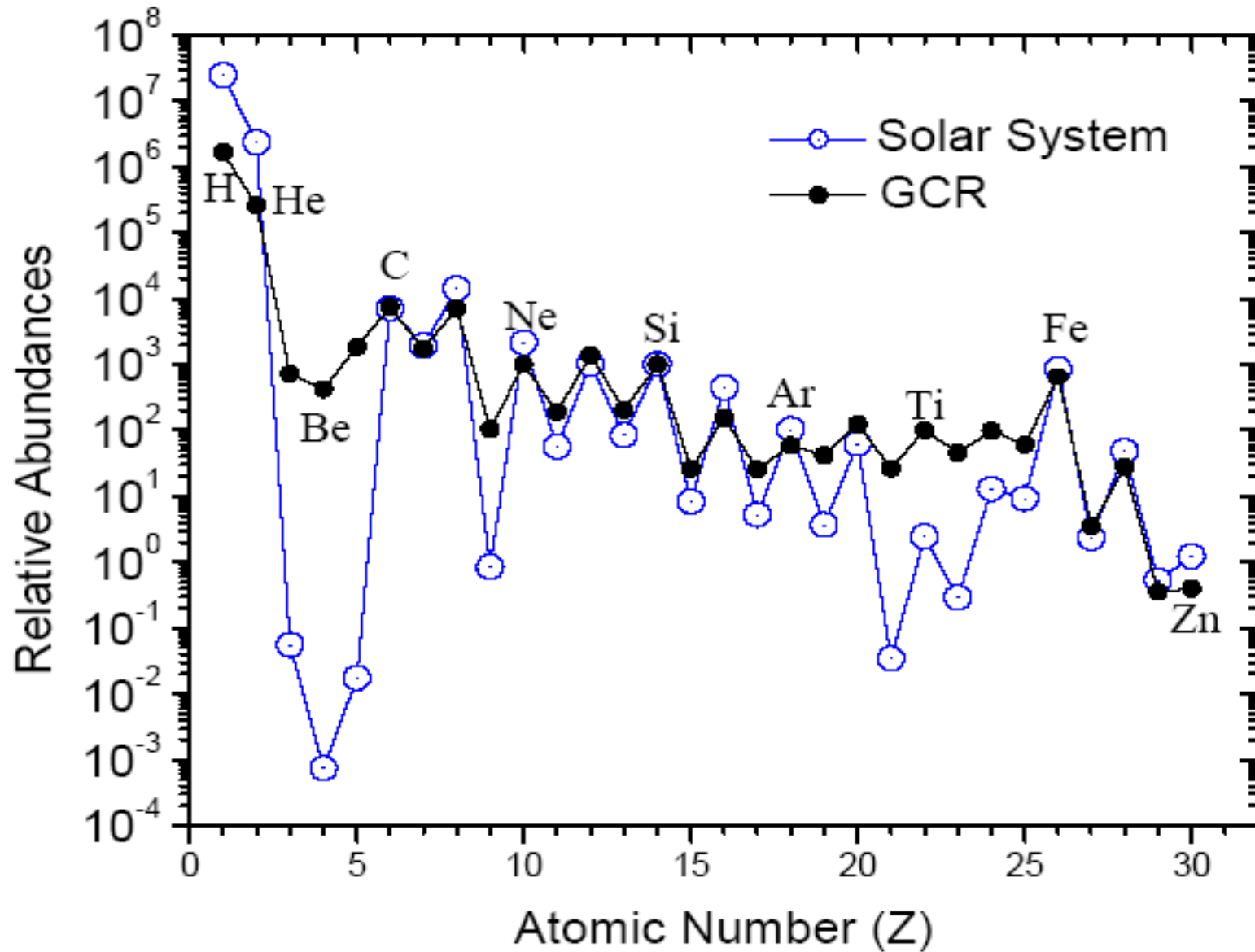
Positron spectra from PAMELA



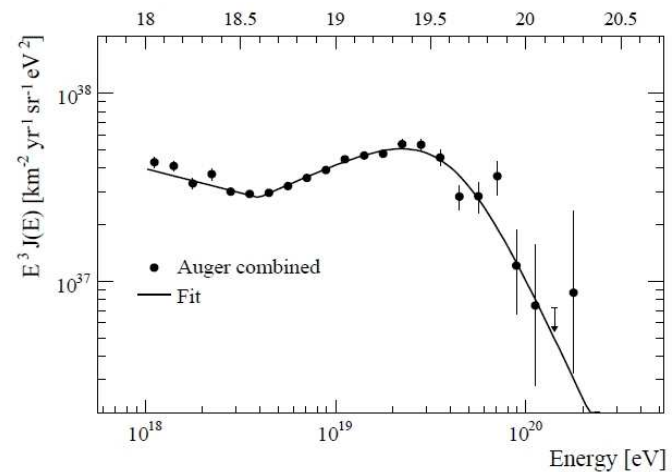
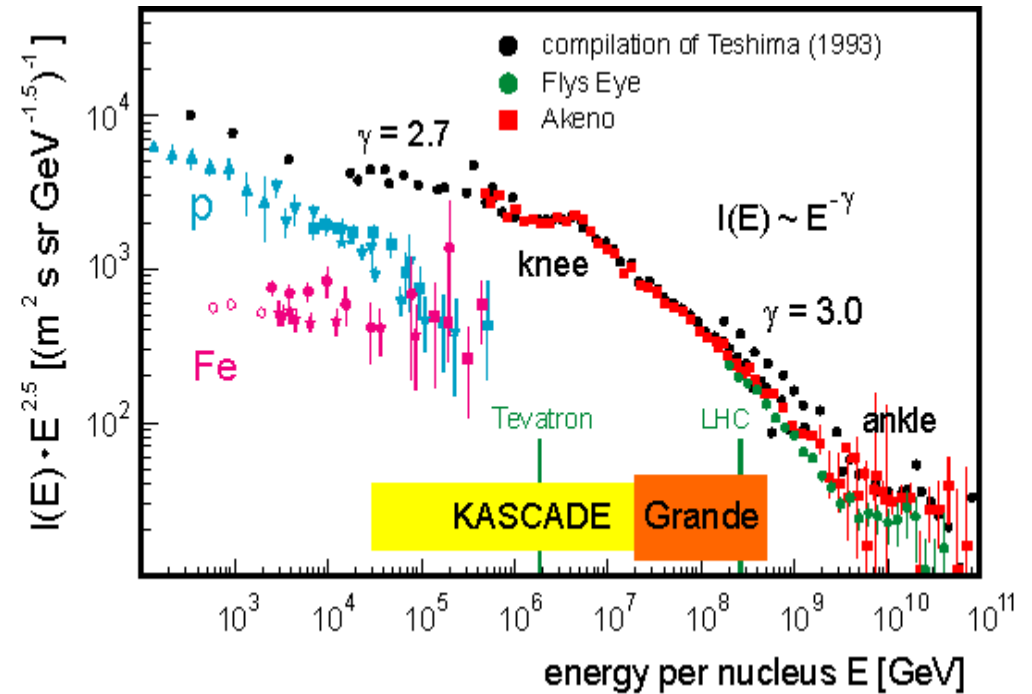
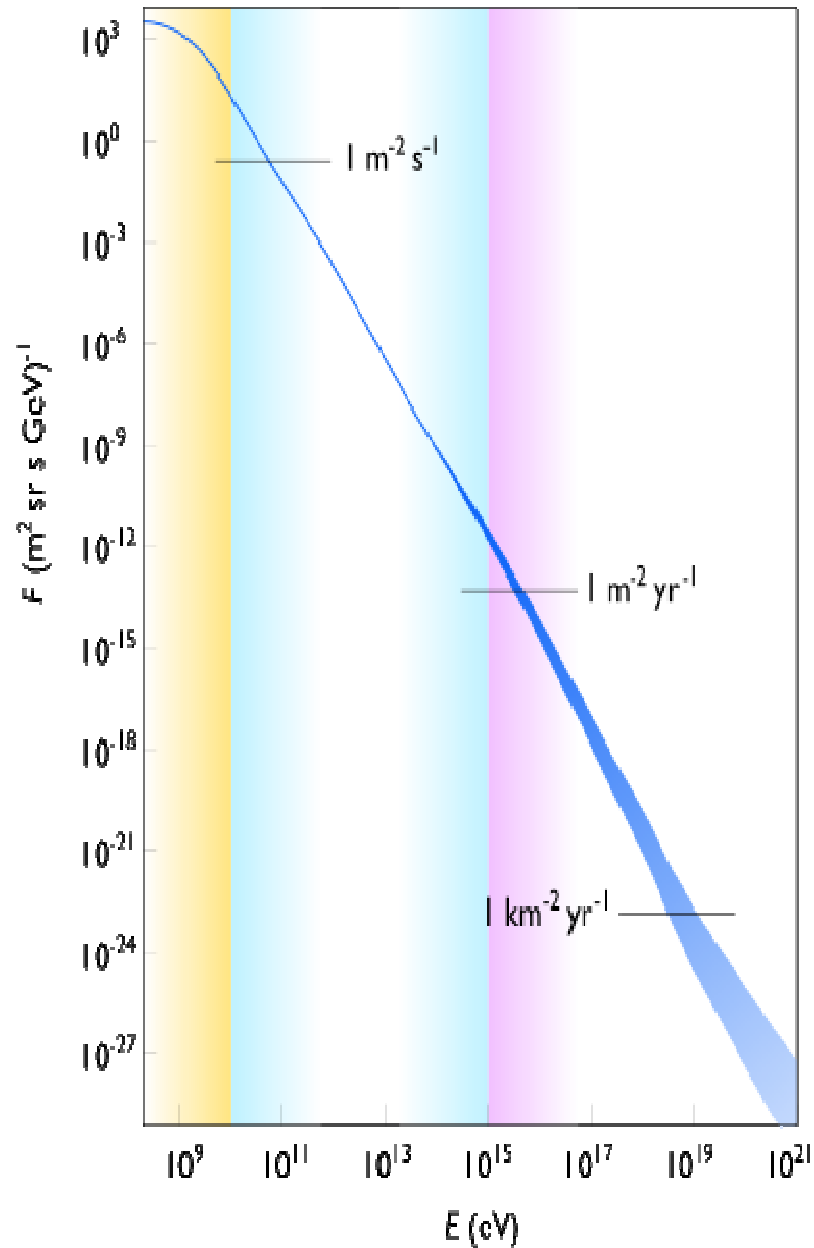
Origin of Cosmic rays?



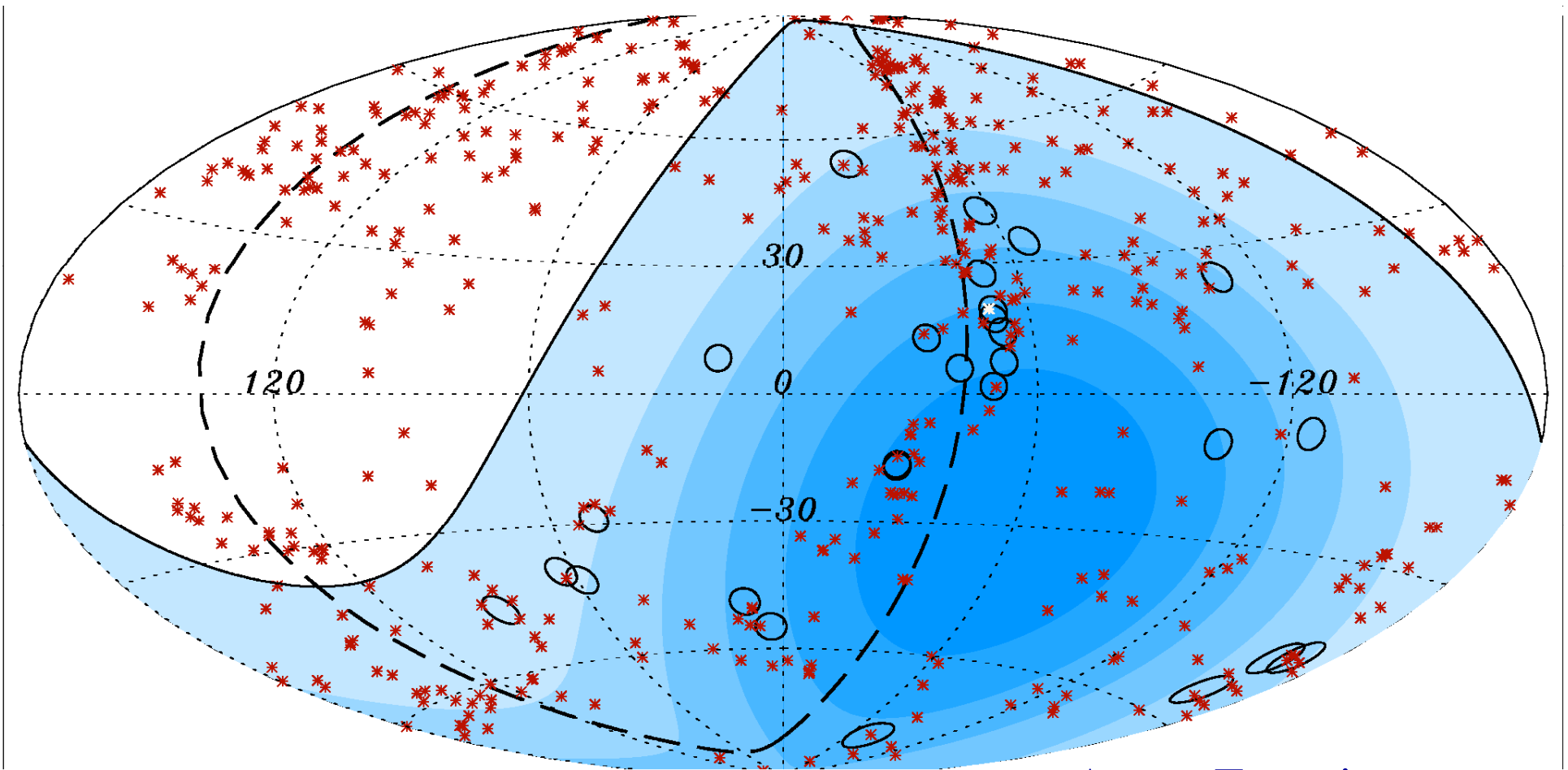
Elemental abundance



Extensive air showers

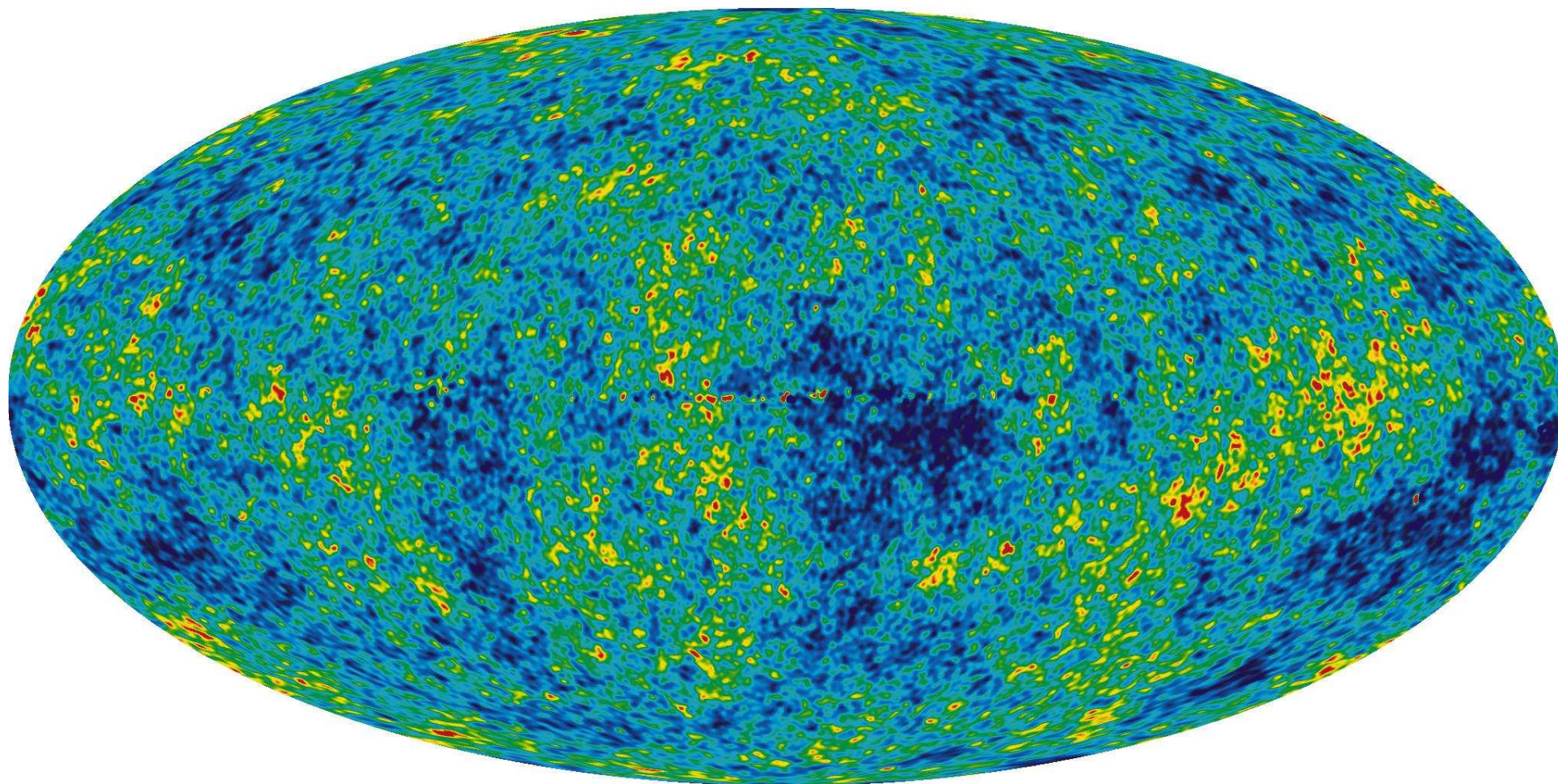


Where are the sources?



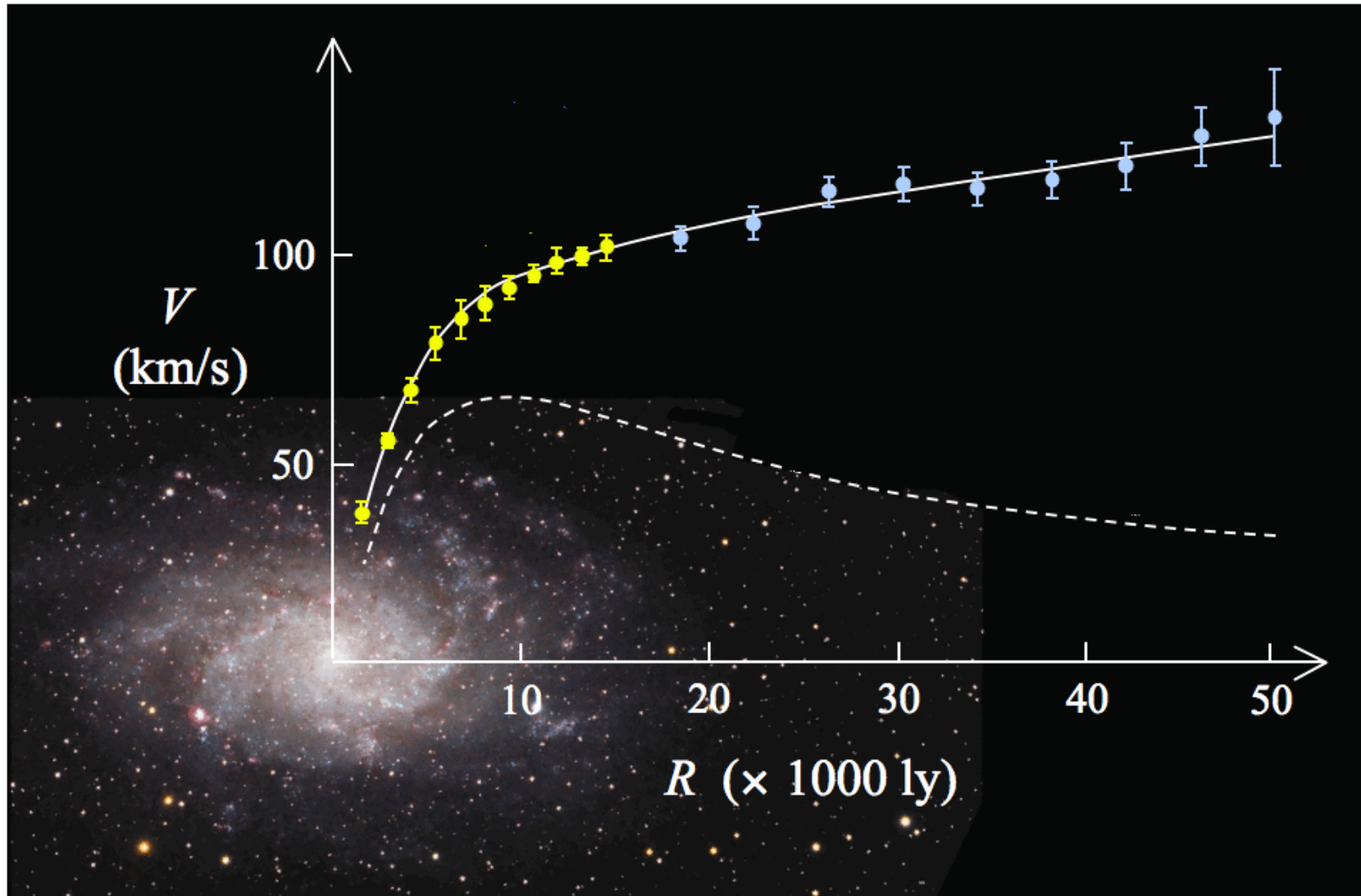
Auger Experiment

Echo of the Big Bang



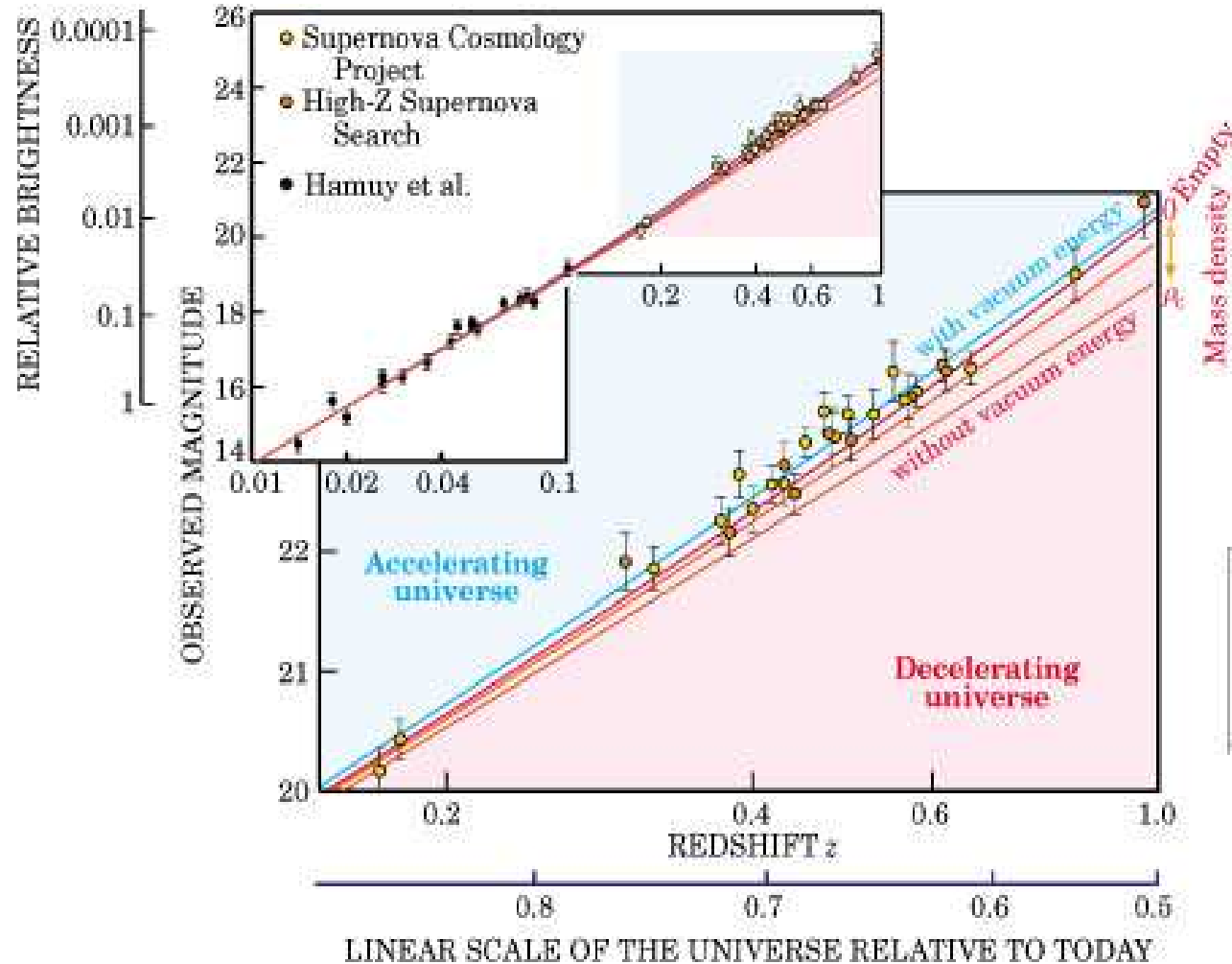
WMAP 2010

Evidence for dark matter



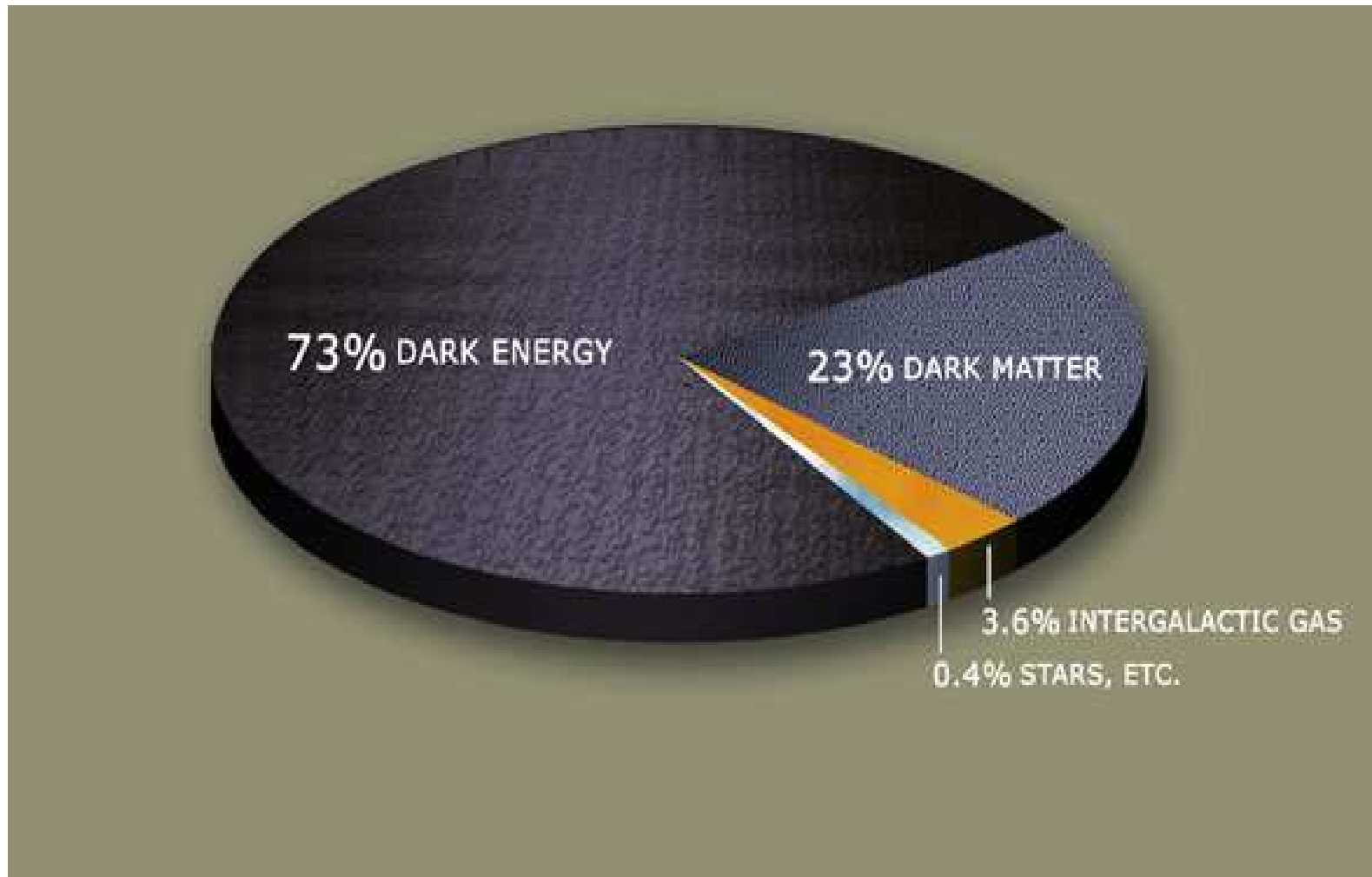
M33

Dark matter, dark energy

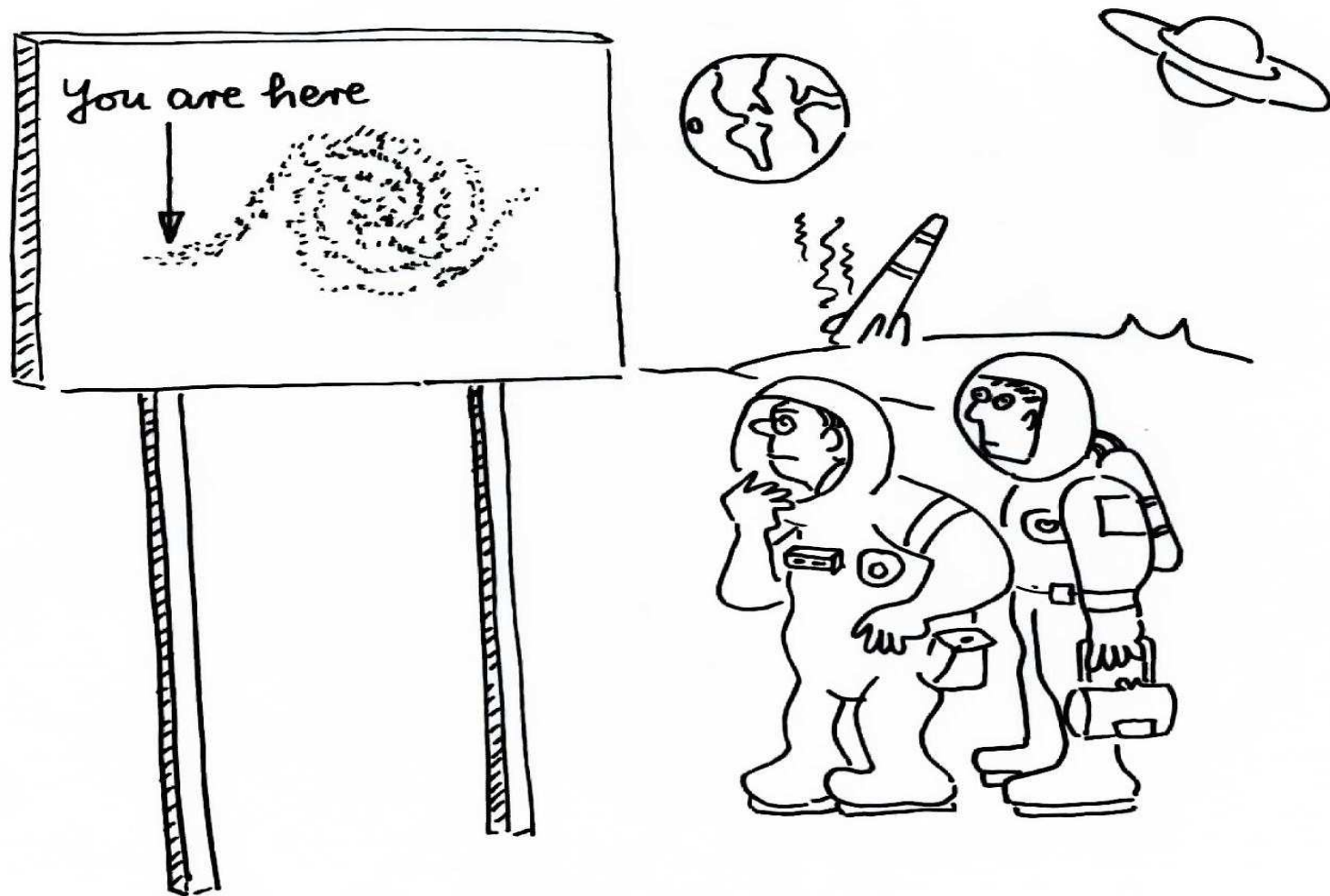


$$z = \frac{\lambda_{\text{obsv}} - \lambda_{\text{emit}}}{\lambda_{\text{emit}}}$$

What makes up the Universe?



Manned space travel?



Positive effects of cosmic rays?



"They call it, I believe, 'Cosmetic Rays'.
They guarantee that there will be no
negative side effects!"

Outlook

Cosmic rays is the birthplace of elementary particles

Since 1987 we experience a Renaissance of cosmic rays:

Neutrino astronomy

Gamma-ray astronomy

Particle astronomy at the highest energies

Gravitational wave astronomy (to come)

The highest energies provided by Nature
will never be reached at accelerators

*Cosmic rays will continue to be a lab without
competition*