

# HENSA collaboration at LNGS: characterization of the underground neutron flux

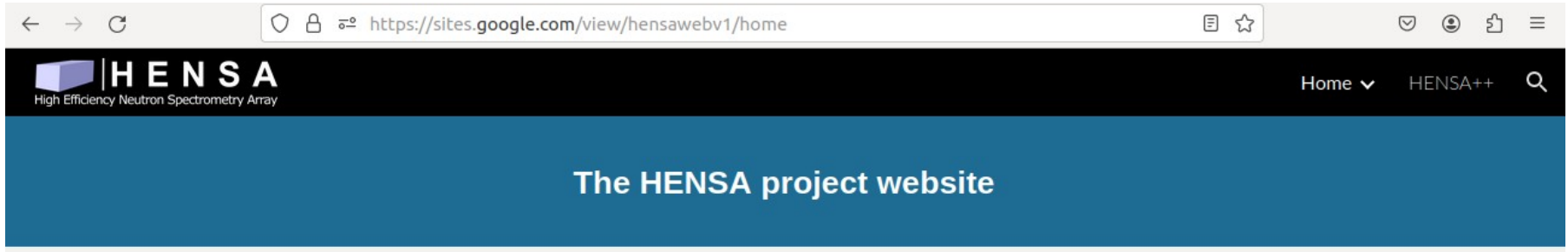
Nil Mont-Geli<sup>1</sup> on behalf of the HENSA collaboration<sup>2</sup>

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<sup>2</sup>[www.hensaproject.org](http://www.hensaproject.org)



# The HENSA collaboration



[www.hensaproject.org](http://www.hensaproject.org)

# The HENSA collaboration



## *HENSA collaboration at Laboratori Nazionali del Gran Sasso (LNGS)*



UNIVERSITAT POLITÈCNICA DE CATALUNYA  
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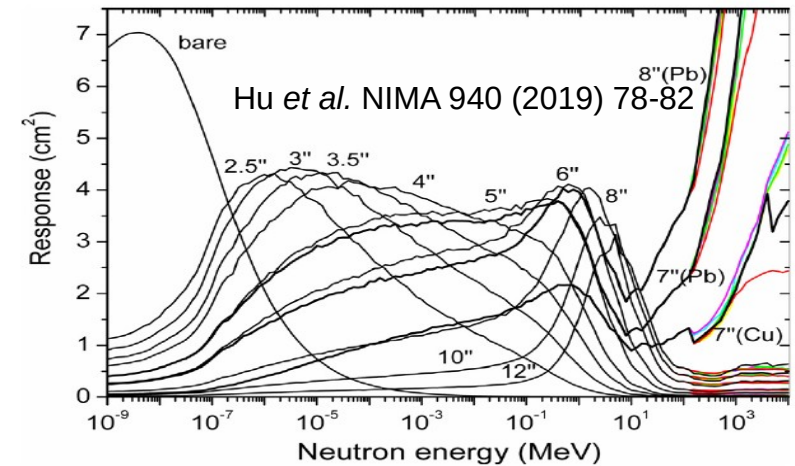
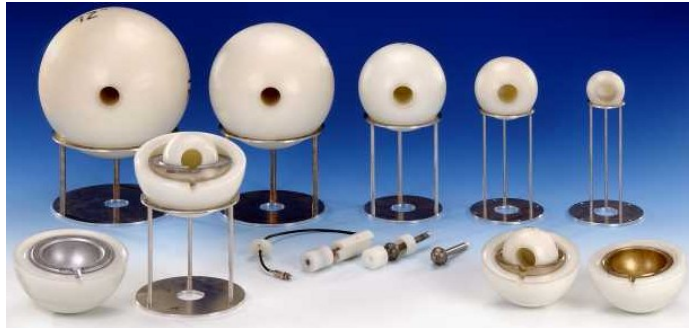
## Characterization of the underground neutron flux at LNGS

- Absolute magnitude of the neutron flux.
- Energy spectrum.
- Local variations.
- Temporal/seasonal fluctuations.

# What is HENSA?

It is a high efficiency neutron detection system based on the same principles than Bonner Spheres.

- Array of moderated proportional neutron counters with variable sizes in order to provide sensitivity at different neutron energies.



$$M_i = \int R_i(E)\phi(E) dE. \quad \rightarrow \quad M_i = \sum_{j=1}^n R_{ij}\phi_j$$

**From M and R, we estimate  
the real spectrum  $\Phi$   
(unfolding)**

# What is HENSA?



**It is a high efficiency neutron detection system based on the same principles than Bonner Spheres.**

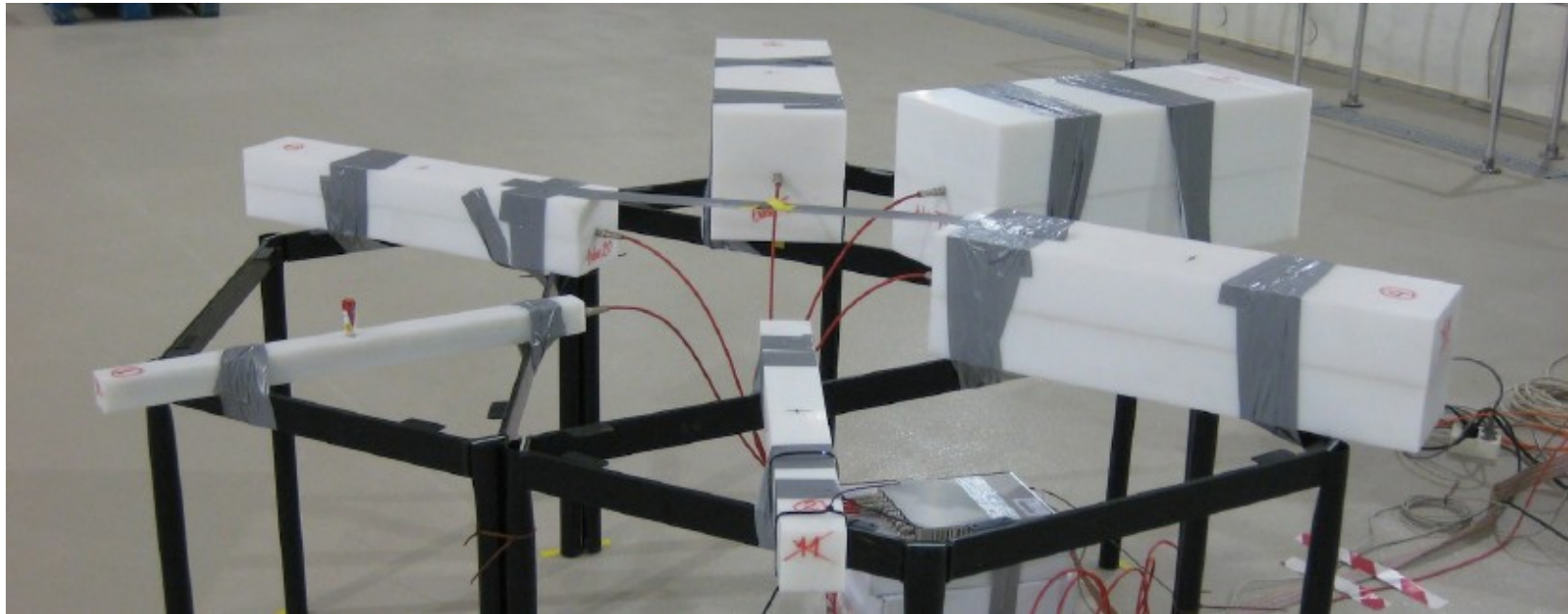
- **Original idea by J.L. Tain (IFIC) in 2010**: high efficiency spectrometer with digital acquisition system for CUNA project (Canfranc Underground Nuclear Astrophysics).
- HENSA is achieved by a topological change in Bonner Spheres in order to benefit from high detection efficiency in cylindrical proportional neutron counters (60 cm active length).
- HENSA project is a scientific collaboration for the exploitation of the spectrometer. Focus on measurements in **underground laboratories** and **secondary neutrons produced by cosmic-rays**.

# What is HENSA?

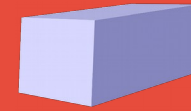


## Above thermal energies: moderation (HDPE)

First version of the HENSA setup (6 detectors) in 2011



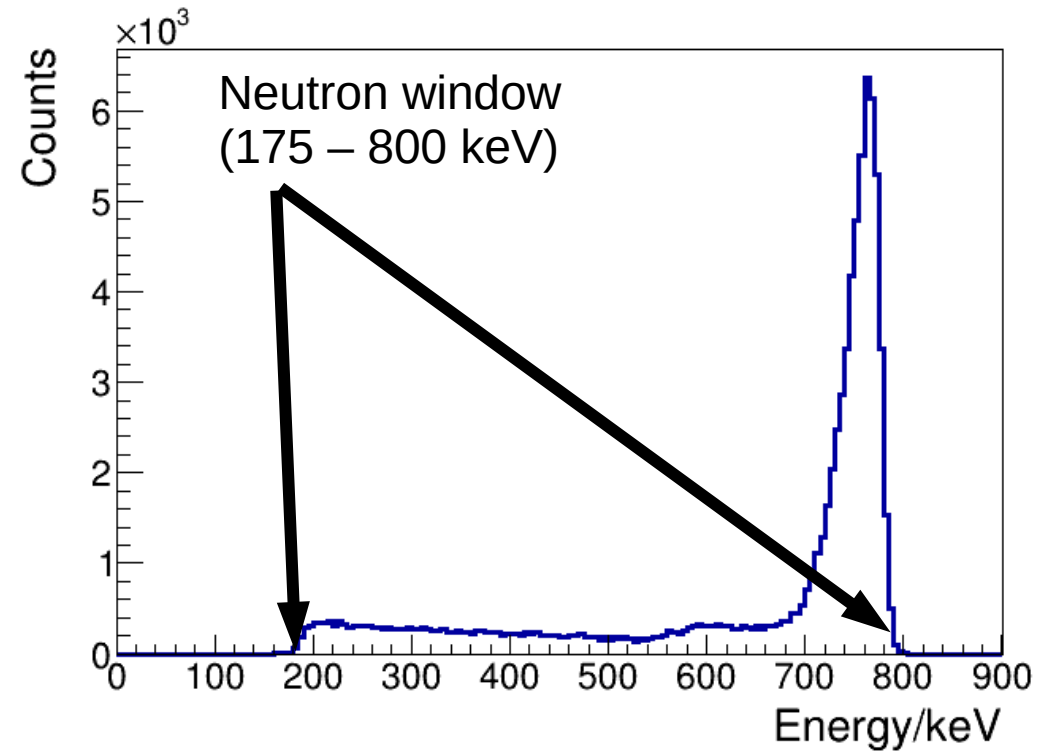
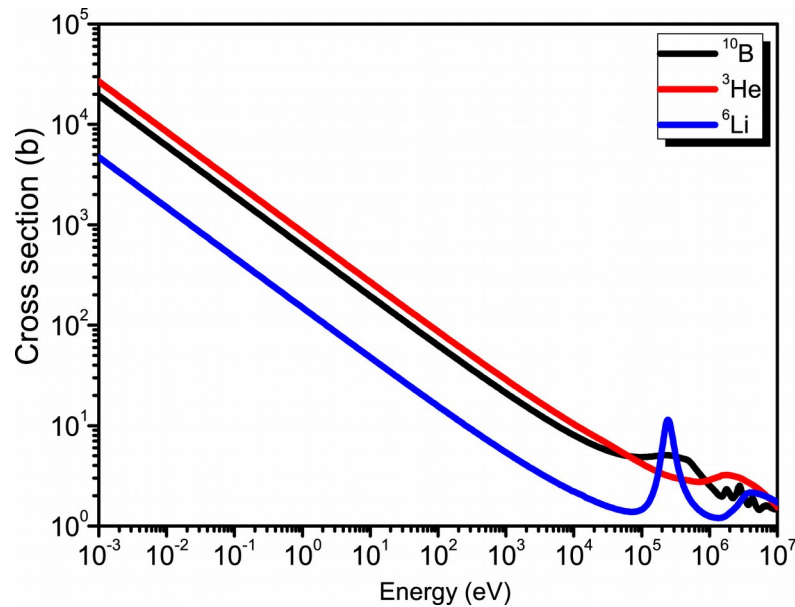
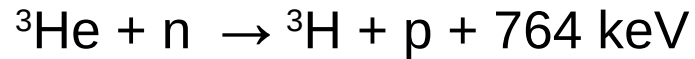
# What is HENSA?



# HENSA

High Efficiency Neutron Spectrometry Array

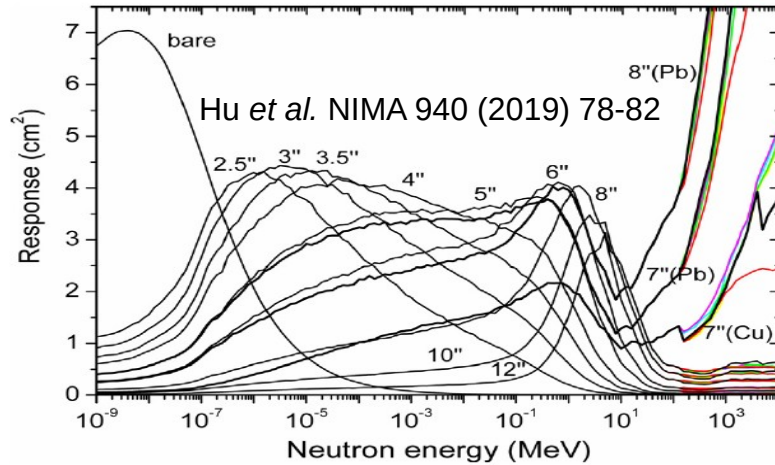
## Neutron detection in He-3 proportional counters



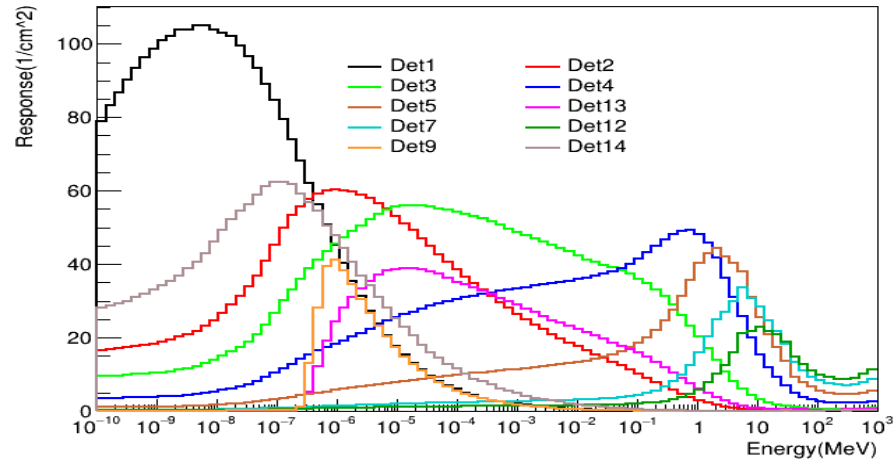


# What is HENSA?

## Standard extended Bonner Spheres



## HENSA version 2022



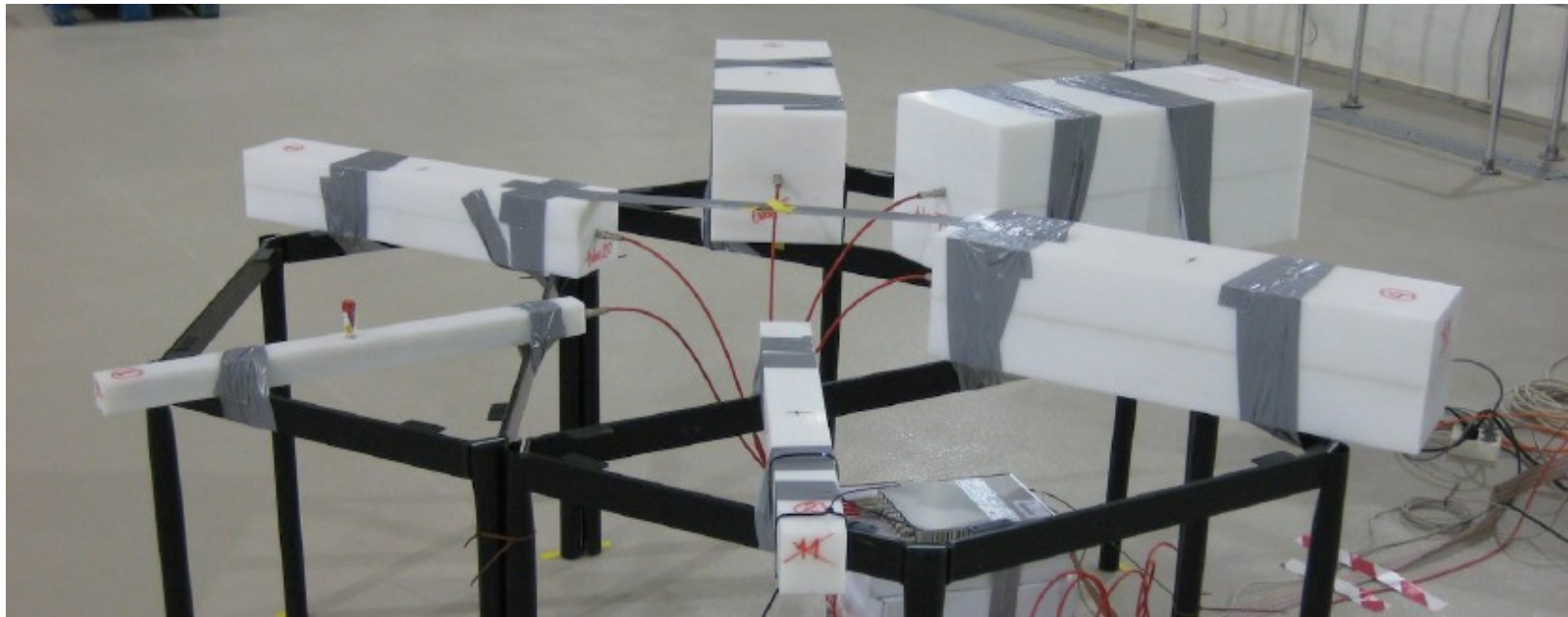
**HENSA** neutron response is **~5-15 times larger** than standard Bonner Spheres systems in the energy range **from thermal up to 20 MeV**.

Use of **larger counters** (3He) and of moderators (HDPE) with the shape of a rectangular prisms.

## HENSA at LSC (first measurement in hall A)

D. Jordan *et al* 2013, *Astroparticle Physics* 42, 1 – 6

*Corrigendum*: D. Jordan *et al* 2020, *Astroparticle Physics* 118, 102372



## HENSA at Felsenkeller

M. Grieger *et al* 2020, Phys. Rev. D 101, 123027



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M. Grieger *et al* 2020, Phys. Rev. D 101, 123027



## Long-term measurement at LSC (hall A)

S. E. A. Orrigo *et al* 2021, J. Phys.: Conf. Ser. 2156, 012169

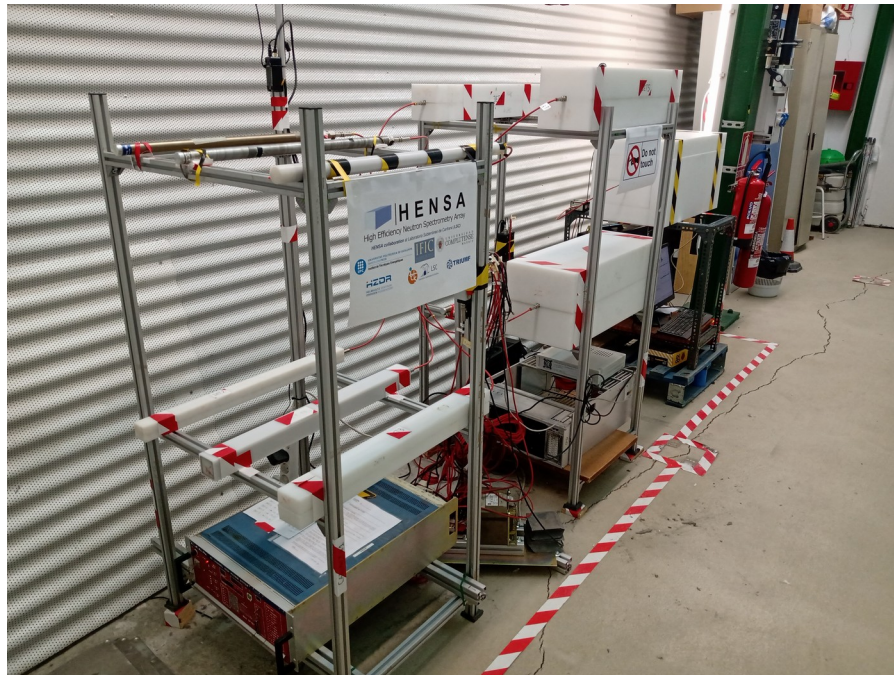
S. E. A. Orrigo *et al* 2022, European Physics Journal C 82, 814



## Long-term measurement at LSC (hall B)

N. Mont-Geli *et al* 2021, J. Phys.: Conf. Ser. 2156, 01223

N. Mont-Geli *et al* 2023, Proceedings of Science 441, 312



# HENSA before LNGS



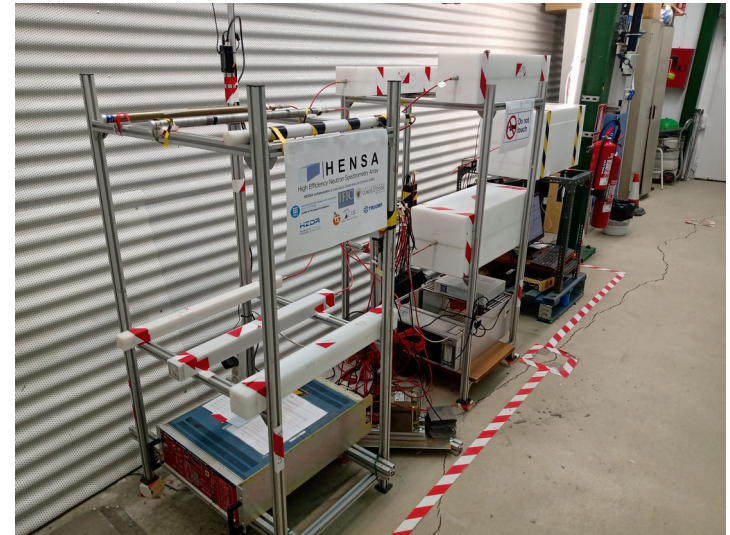
## HENSA for CR (in 2020 and since 2023/2024)



## Long-term measurements in halls A and B

Two setups:

- HENSA-v2019 (hall A, October 2019 – March 2021).
- HENSA-v2019 (hall B, February 2022 – August 2022).
- HENSA-v2022 (hall B, since August 2022) – **Better spectral resolution**

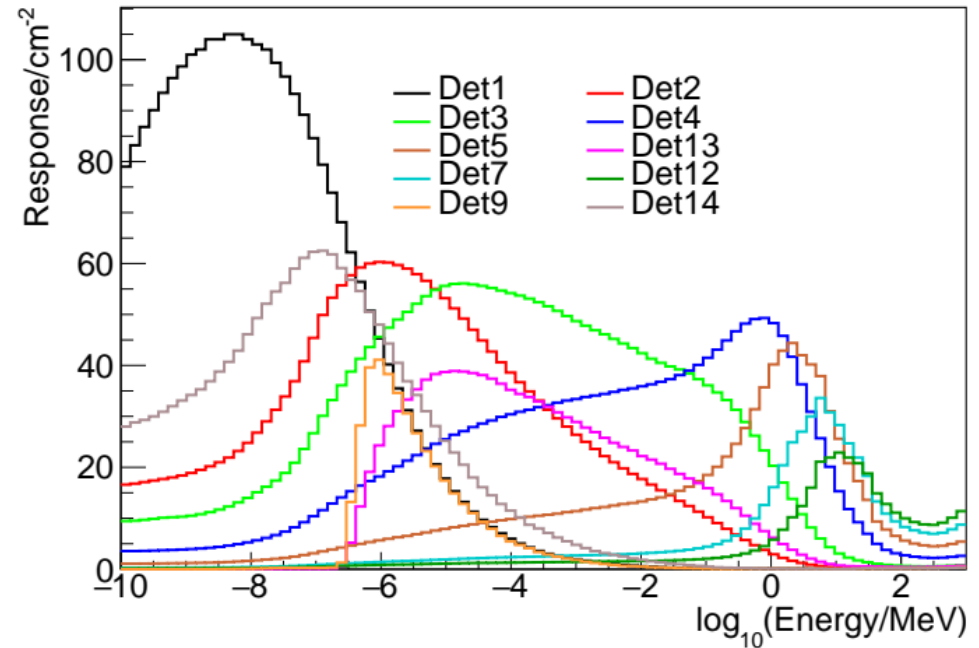
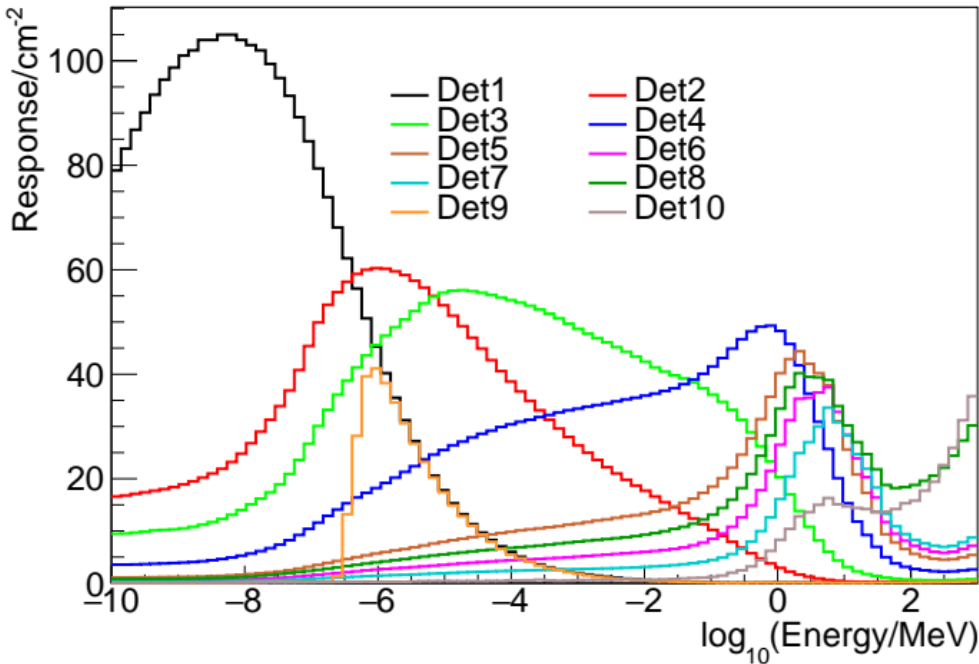




## Responses



[www.particlecounter.net](http://www.particlecounter.net)



## Neutron source in hall A

Monte Carlo FLUKA (v. 4.3.2) calculations used to estimate the neutron flux **in Canfranc**.

- **( $\alpha$ ,n) reactions.**
- **Spontaneous fission.**
- **Muon-induced neutrons**

Thermal peak  $\sim 4.5 - 6.5 \cdot 10^{-8}$  MeV

Isolethargic intermediate region

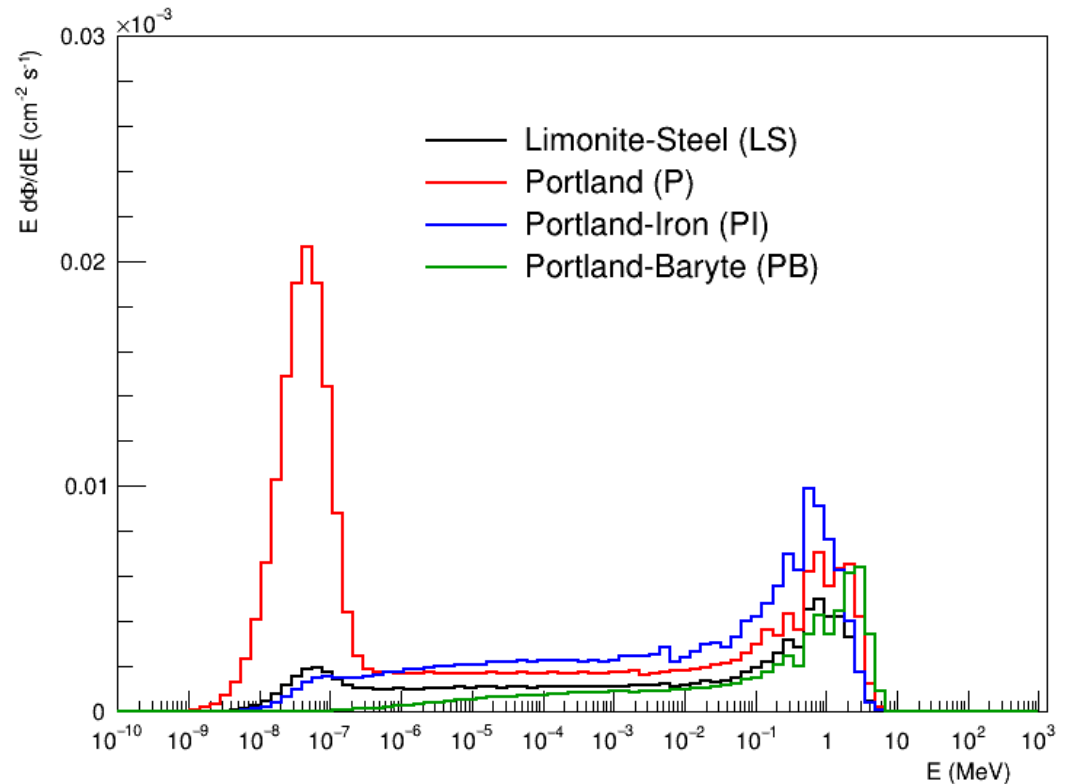
Fast peak  $\sim 0.5 - 3$  MeV

HE peak  $\sim 50$  MeV

**Concrete contribution: 94%**

**Rock contribution: 6%**

**Muons contribution: 0.03%**



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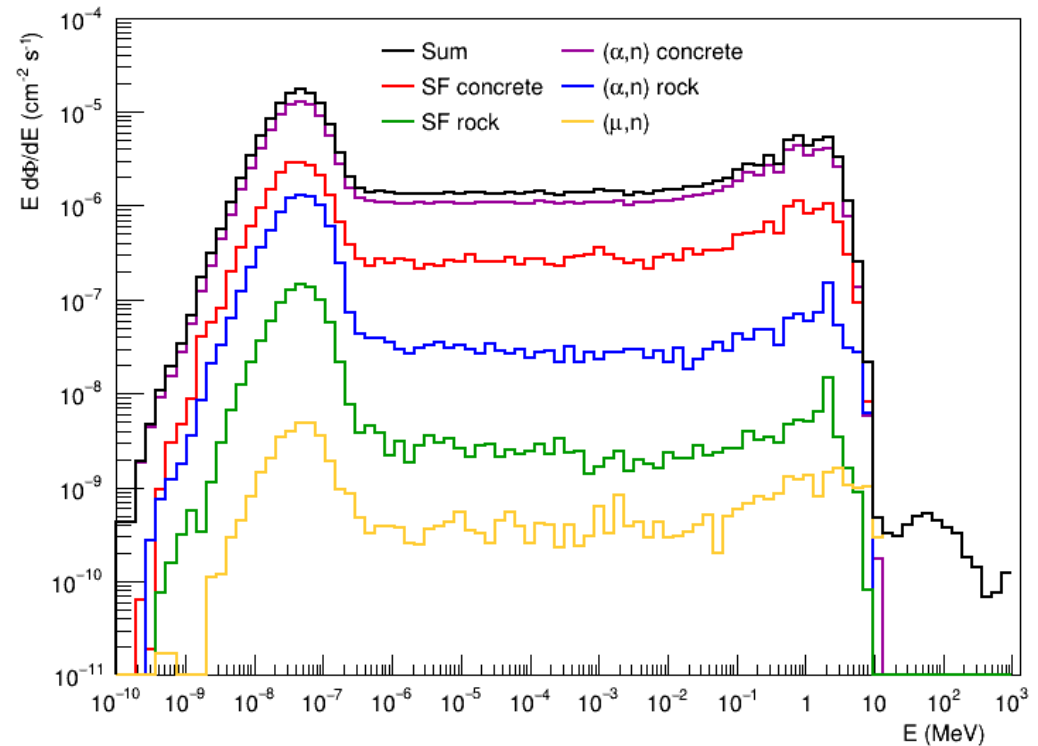
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## Energy spectrum

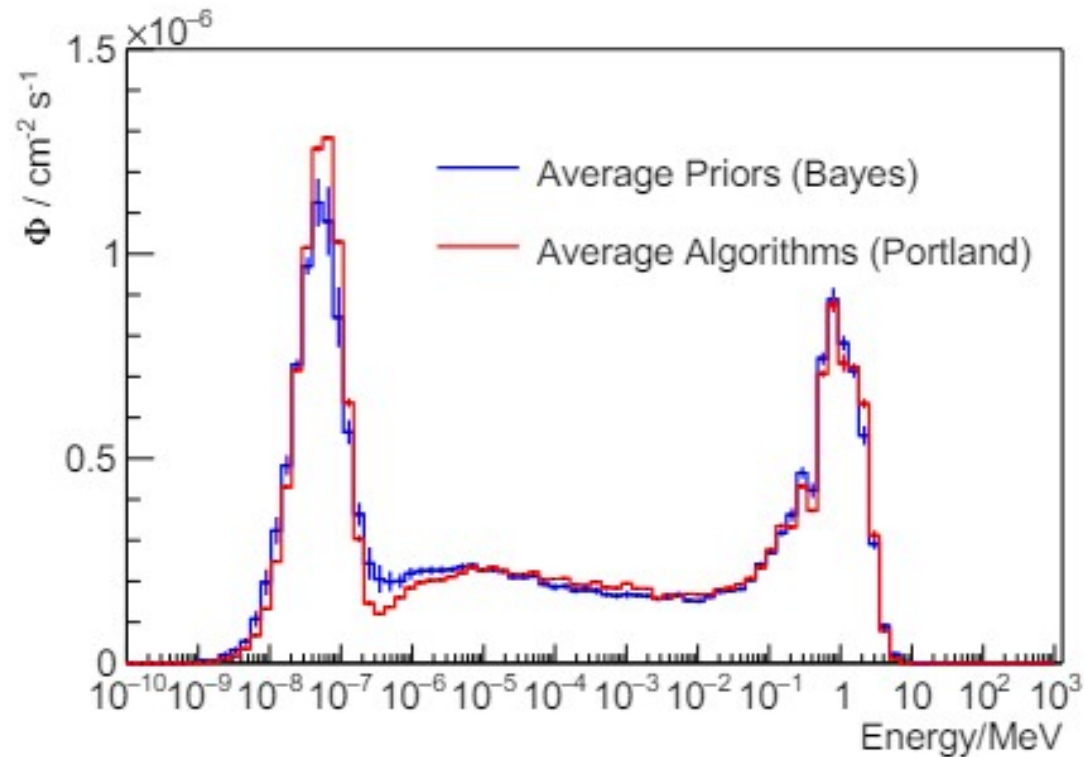
Unfolding algorithms:

$$M_i = \int R_i(E)\phi(E) dE. \quad \rightarrow \quad M_i = \sum_{j=1}^n R_{ij}\phi_j$$

- Bayesian method: **S. Agostini, NIM A 362 (1995) 487**
  - Computational implementation by J.L Tain (IFIC) and D. Cano-Ott (CIEMAT).
- GRAVEL and MAXED implemented in UMG package version 3.3 (free-distributed by PTB).
  - MAXED: **M. Reginatto et al., NIM A 476 (2002) 242**
  - GRAVEL: **M. Matzke, Report PTB-N-19 (1994)**

Unfolding typically relies on some preliminary knowledge about the spectrum (prior spectrum)

## Energy spectrum (hall B)

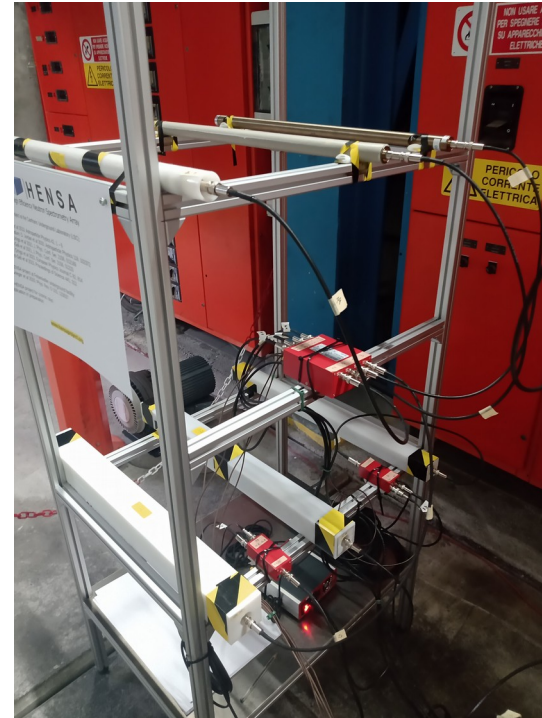
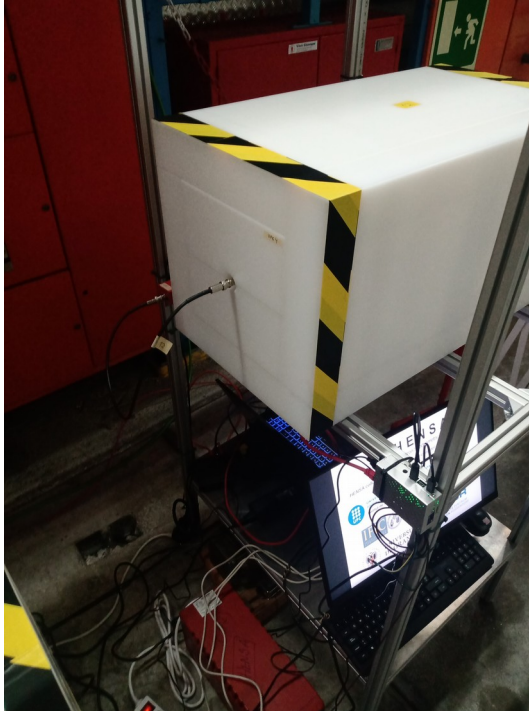


N. Mont-Geli *et al* 2023, Proceedings of Science 441, 312

- Setup assembled in hall A from 17 to 20 April 2024.
- Acquisition started on 20 April 2024.



- 10 He-3 thermal neutron counters (2.54 cm diameter, 60 cm length)
- HDPE shielding with variable sizes (HDPE thickness from 32 to 380 mm)



## Current HENSA plans at LNGS

- Short measurements (2 – 4 months) at different locations:
  - Hall A, since late April 2024.
  - Inside the new STELLA facility (begin in mid-July until late September/October).
  - Other locations: hall B, hall C.
- Long-term monitorization of the neutron flux (minimum 1 year).
  - Location to be decided.



## Status of the neutron data at LNGS

Work	Hall	Technique
Belli (1989)	A	BF3 counters + variable size moderators (Spec.)
Aleskan (1989)	A	Li-6 scintillator (> 3 MeV)
Arneodo (1999)	C	Proton recoil scintillation detector (Spec.)
Belloti (1985)	B	$^3\text{He}$ counters, bare + 1 paraffin (thermal and fast)
Debicki (2009)	?	$^3\text{He}$ bare counters (thermal)
Cribier (1995)	A	$\text{CaNO}_3$ radiochemical detector (> 2.5 MeV)
Rindi (1888)	?	$^3\text{He}$ counters, bare + paraffin + bare and cd
Best (2016)	No hall	$^3\text{He}$ bare counters (thermal)
Debicki (2018)	(same 2009)	$^3\text{He}$ counters + long counter for fast neutrons
Bruno (2019)	A (LVD)	Liquid scintillators (above 10 MeV)

## Status of the neutron data at LNGS

I have compiled ten different previous works related with neutrons (1985 – 2019).

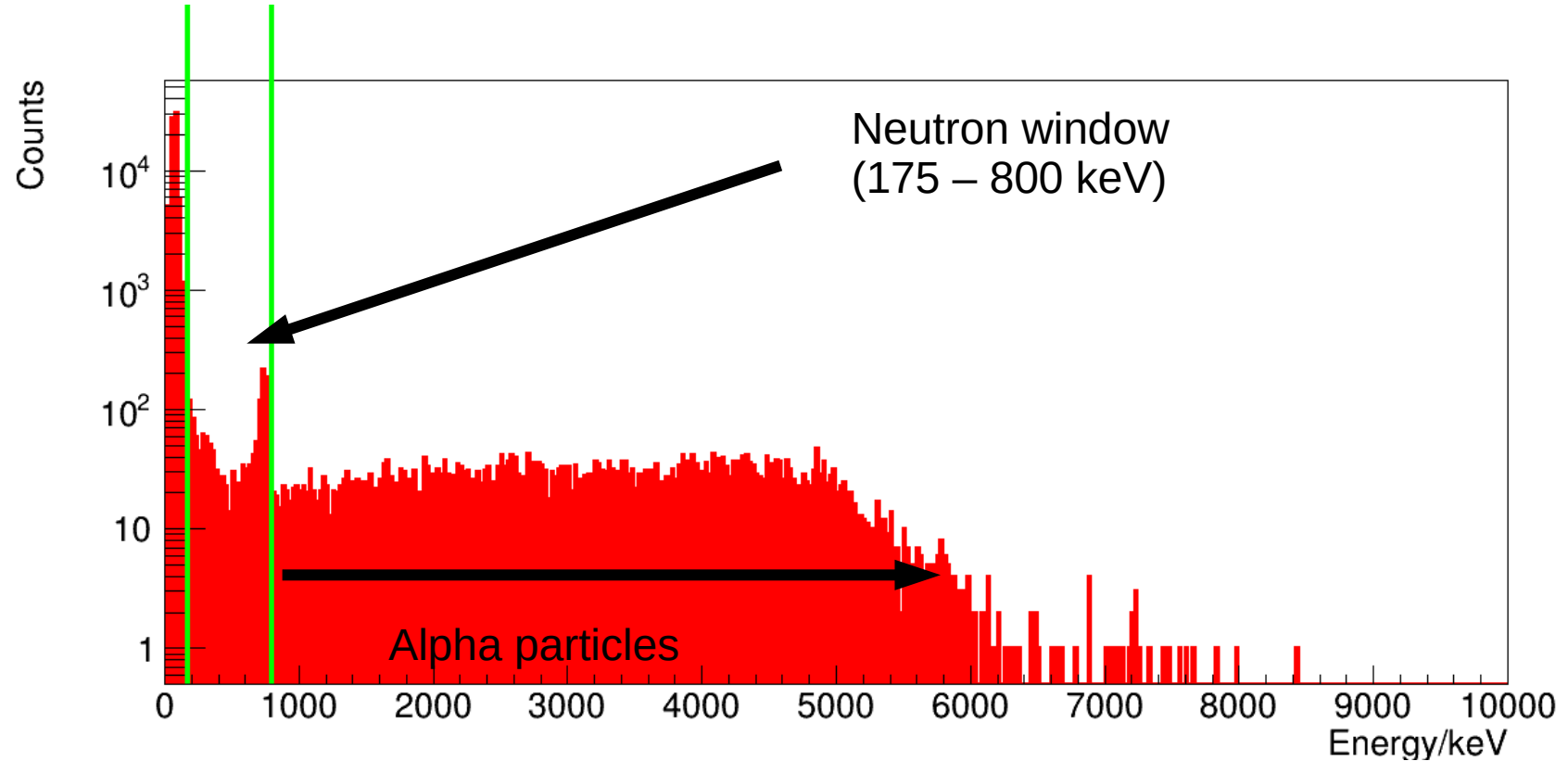
Only two works provides a “complete” spectrum (from thermal to fast energies):

- Belli (1989) using a HENSA-like setup (but no unfolding) in hall A.
- Arneodo (1999) direct energy measurement (proton recoil) in hall C.

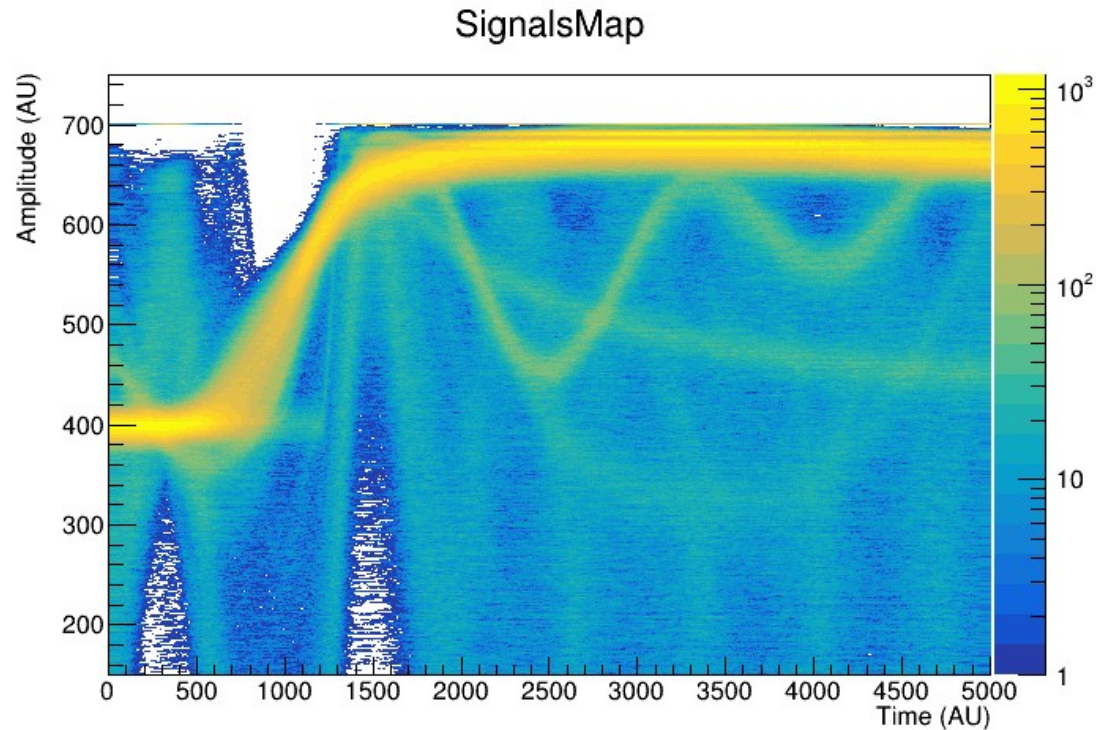
Moreover, discrepancies between measurements. Example in hall A:

- Fast flux ( $\text{cm}^{-2} \text{s}^{-1}$ ):
  - Belli (1989):  $0.32(14) \text{E-6}$  (contribution above 2.5 MeV)
  - Aleskan (1989):  $0.78(4)\text{E-6}$  (only above 3 MeV)
  - Cribier (1995):  $0.09(6)\text{E-6}$  (only above 2.5 MeV)

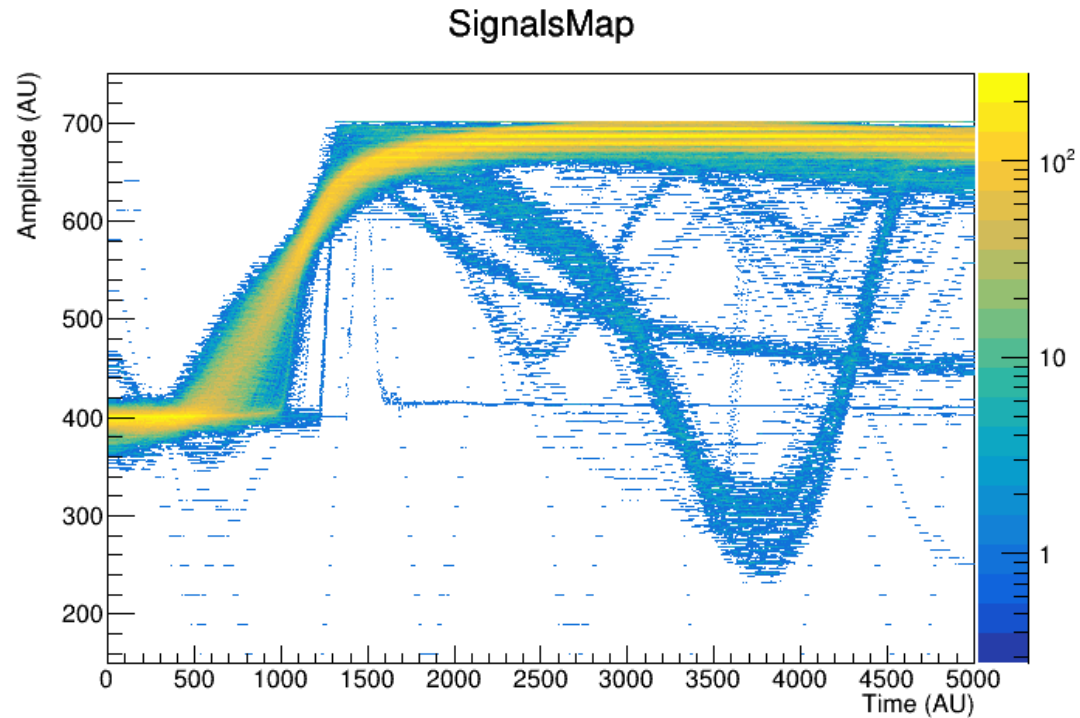
## Pulse-height spectrum (neutron events overlapped with other contributions)



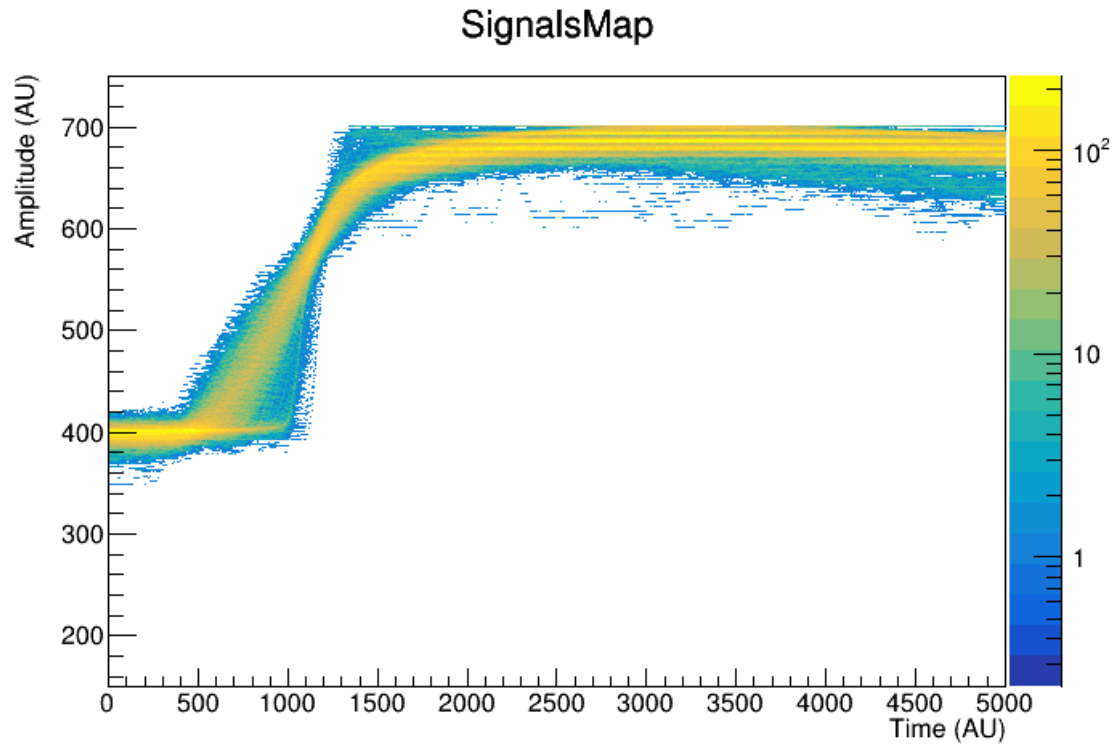
## All signals (from 0 to 1000 keV)



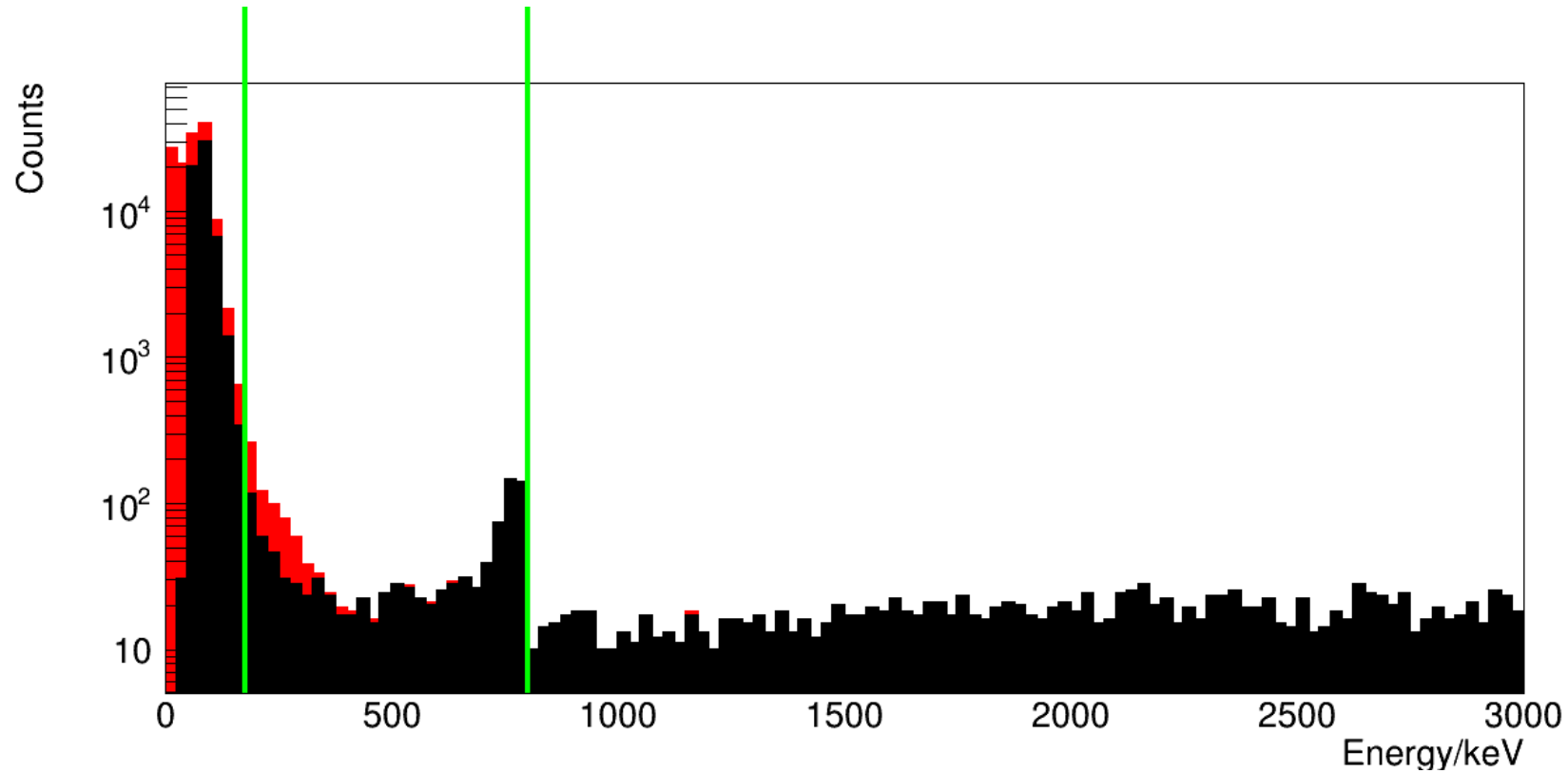
## All signals (from 100 to 1000 keV)



## Pulse Shape Discrimination (noise filter)

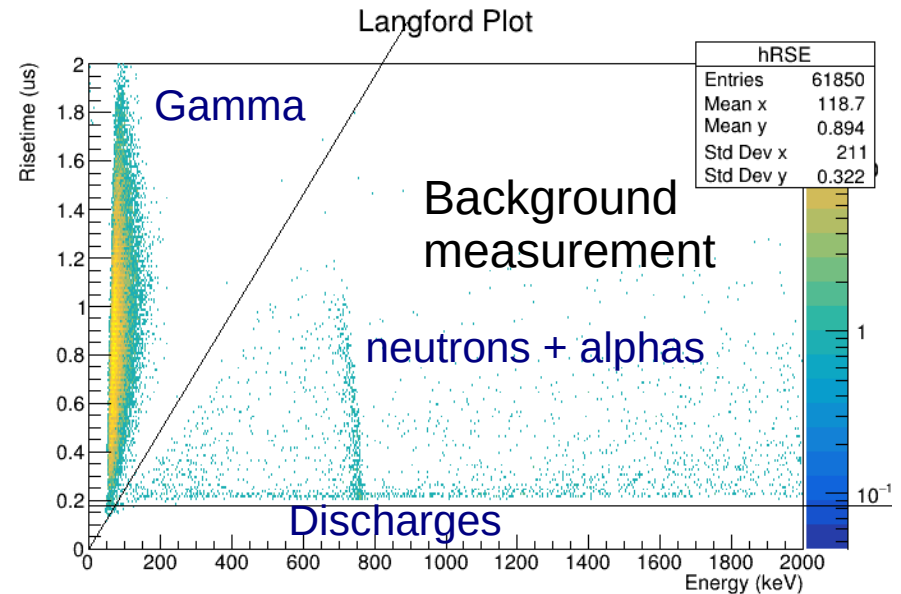
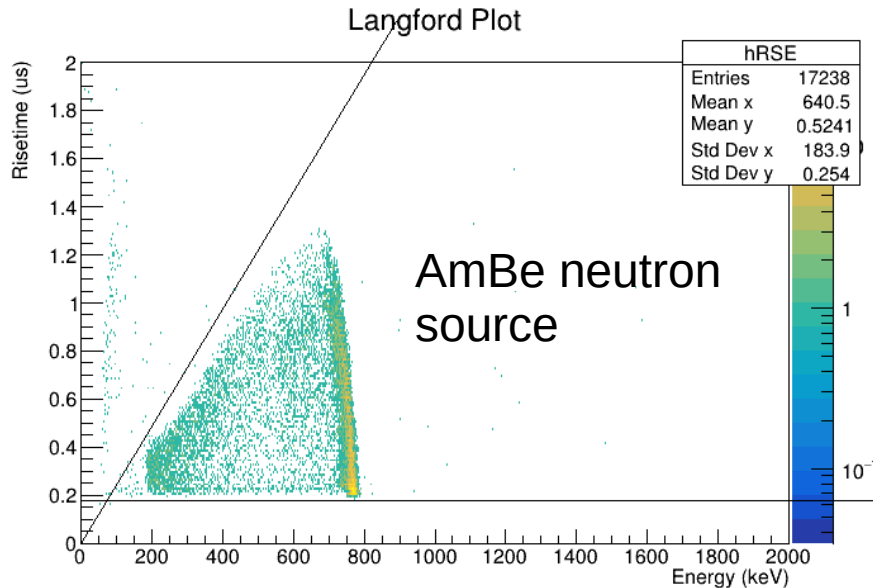


## Pulse Shape Discrimination (noise filter)



## Pulse Shape Discrimination (gamma rays and “fast” discharges filter)

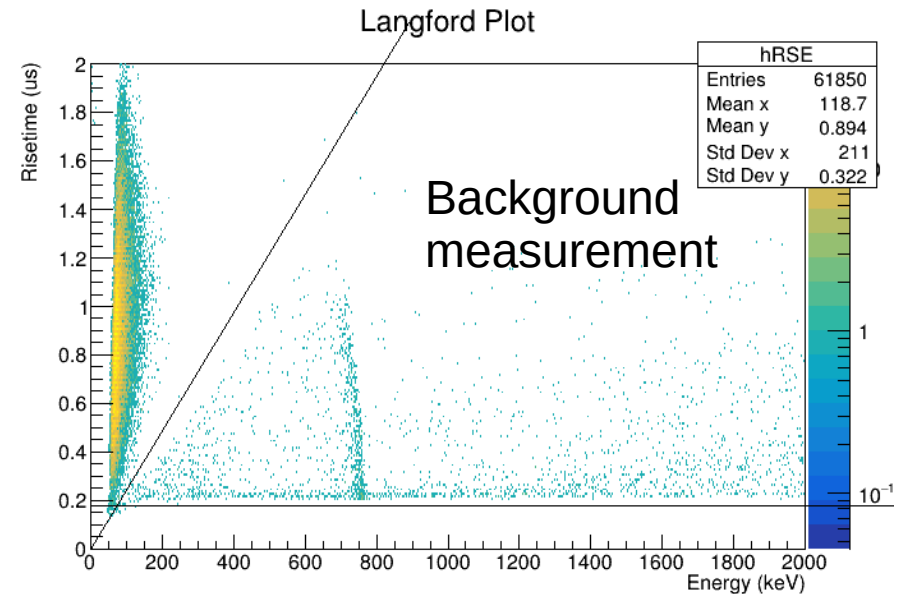
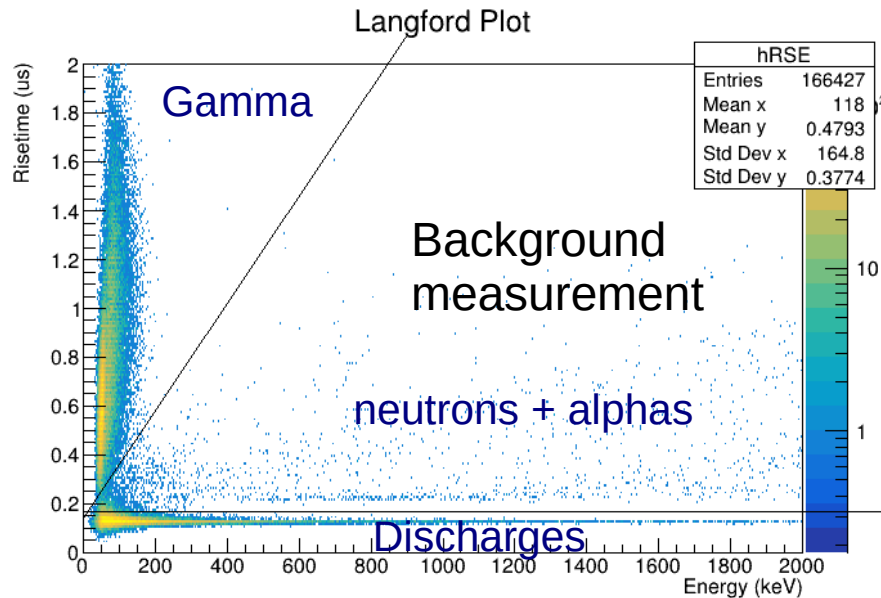
- A filter based on the the risetime differences: Langford *et al* NIM A 717 (2013)



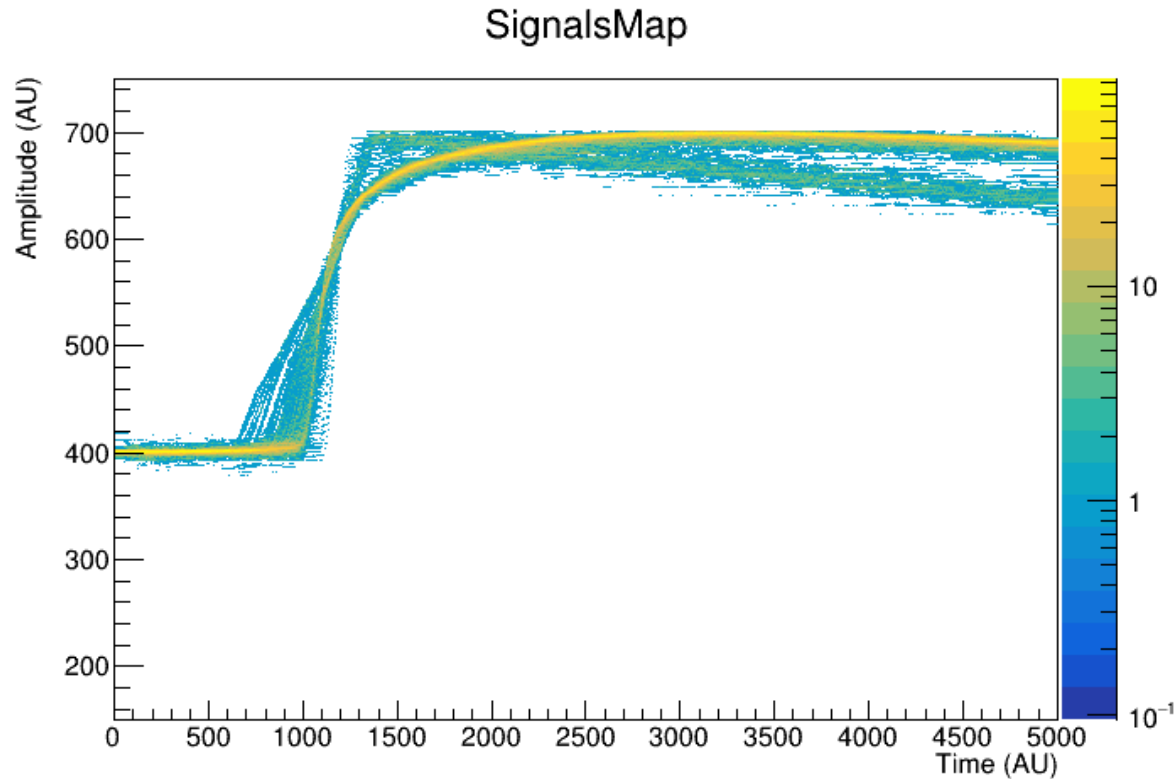


## Pulse Shape Discrimination (gamma rays and “fast” discharges filter)

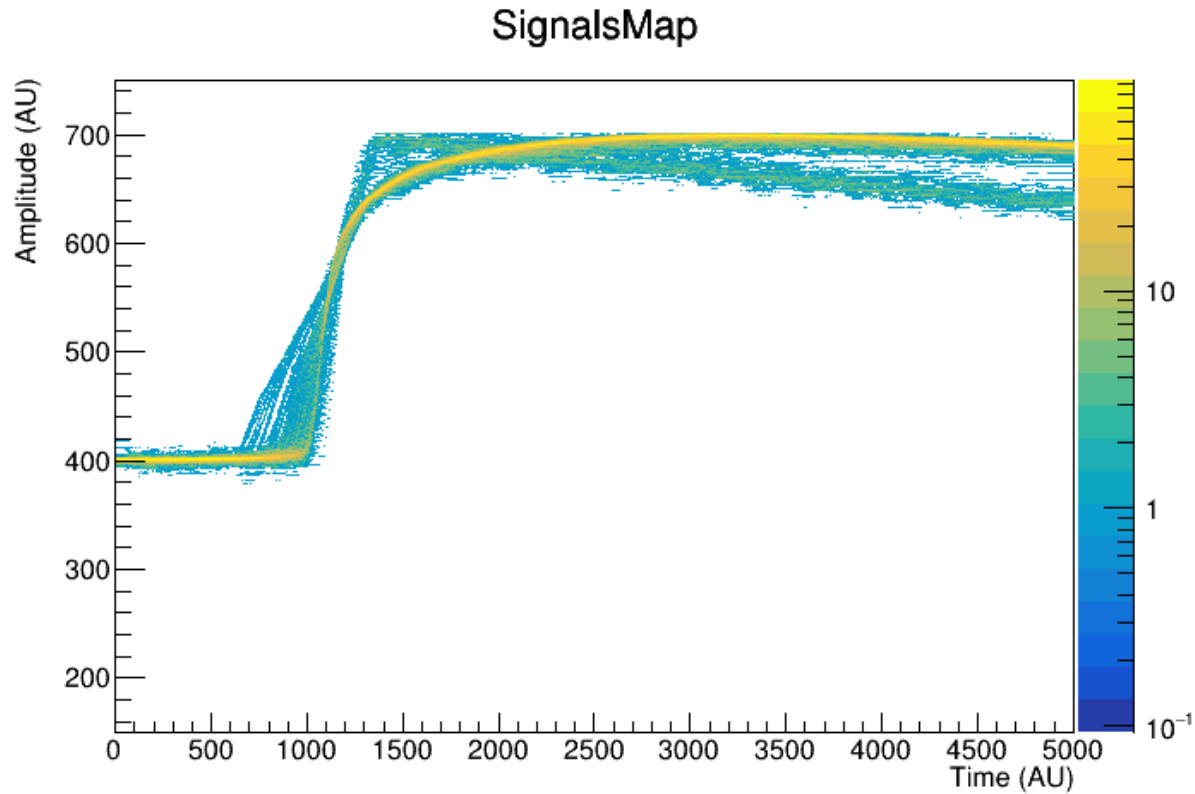
- A filter based on the the risetime differences: Langford *et al* NIM A 717 (2013)



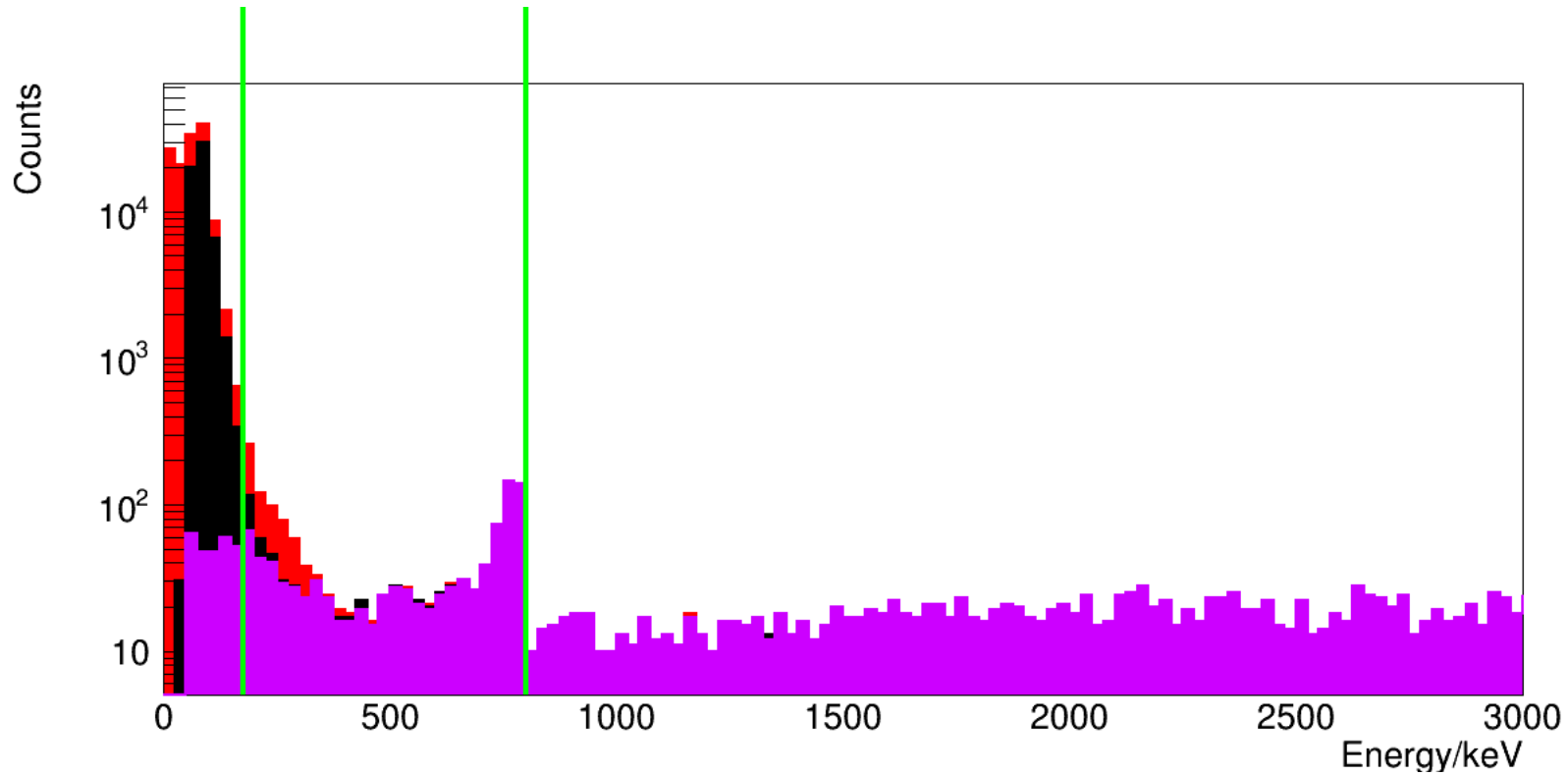
## Pulse Shape Discrimination (Gamma filter)



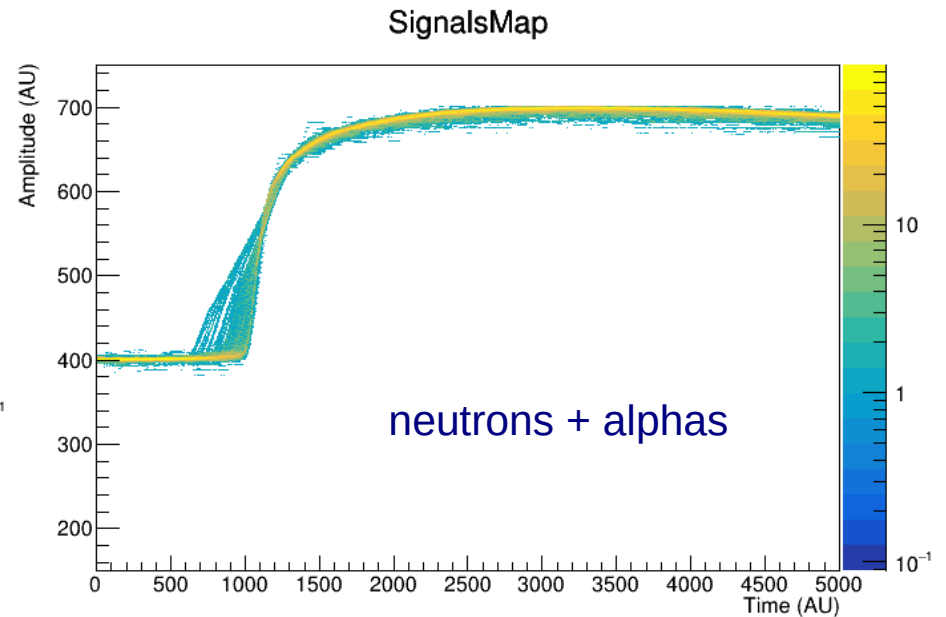
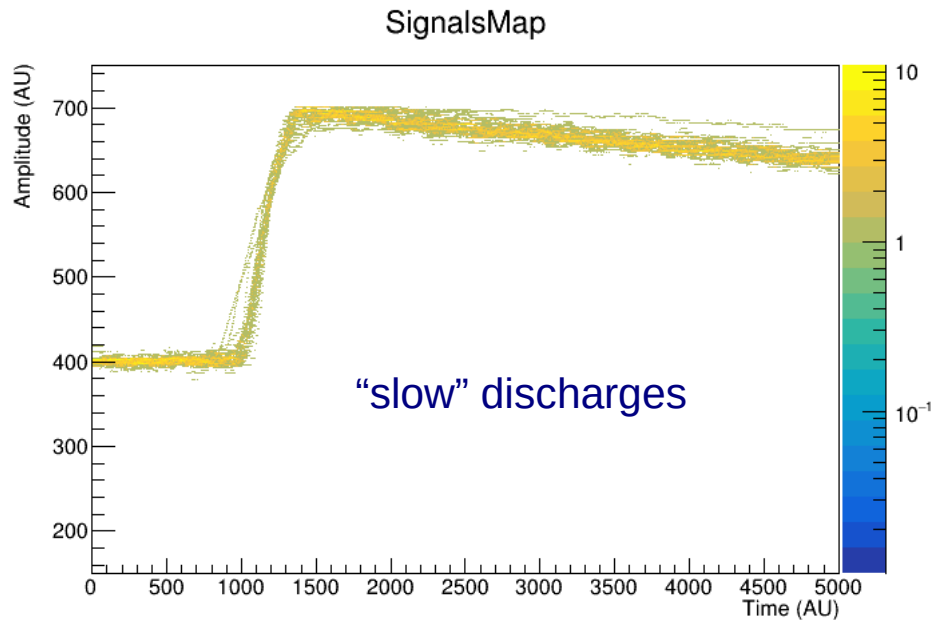
## Pulse Shape Discrimination (Langford filter)



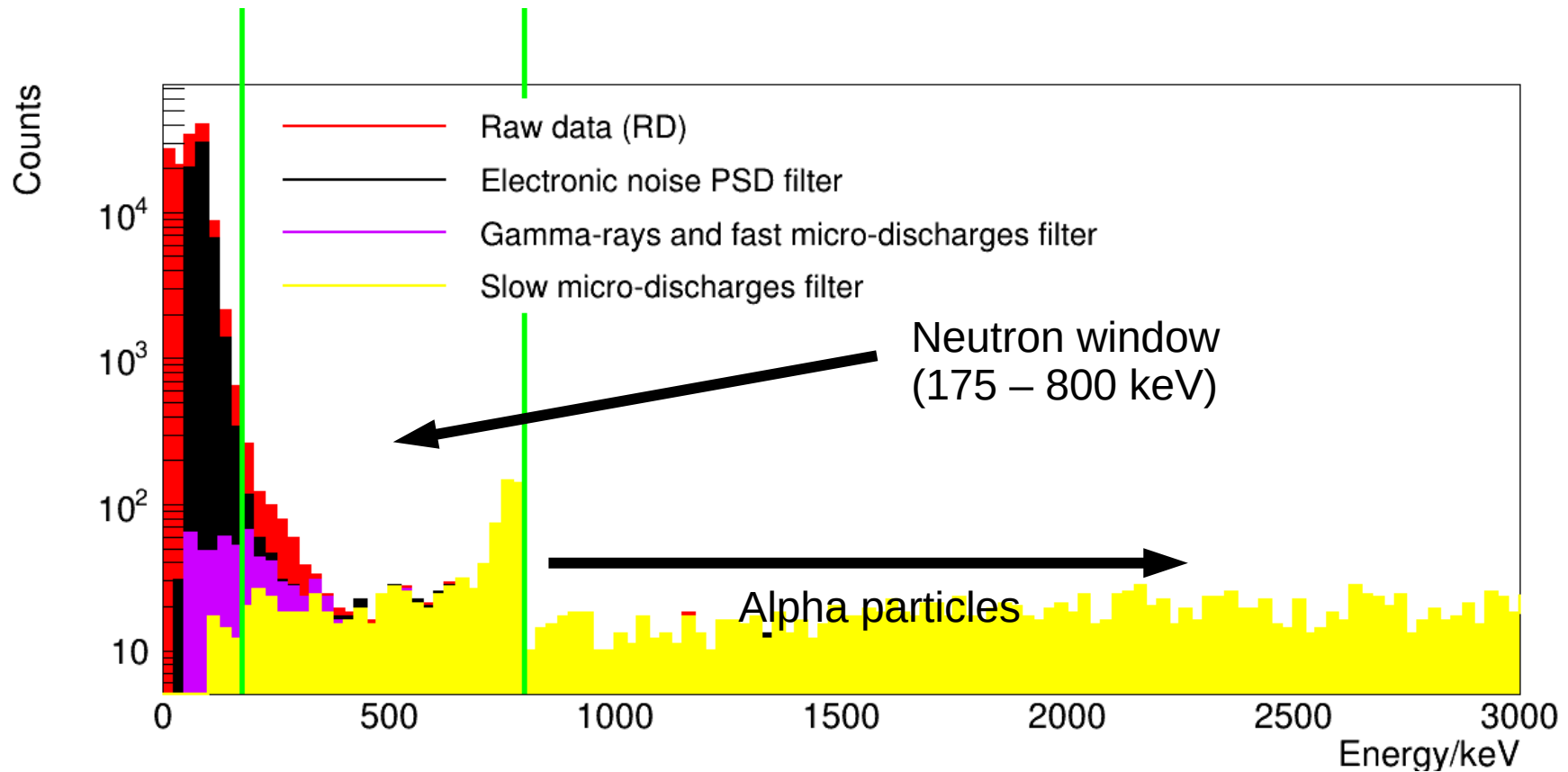
## Pulse Shape Discrimination (noise + Langford filter)



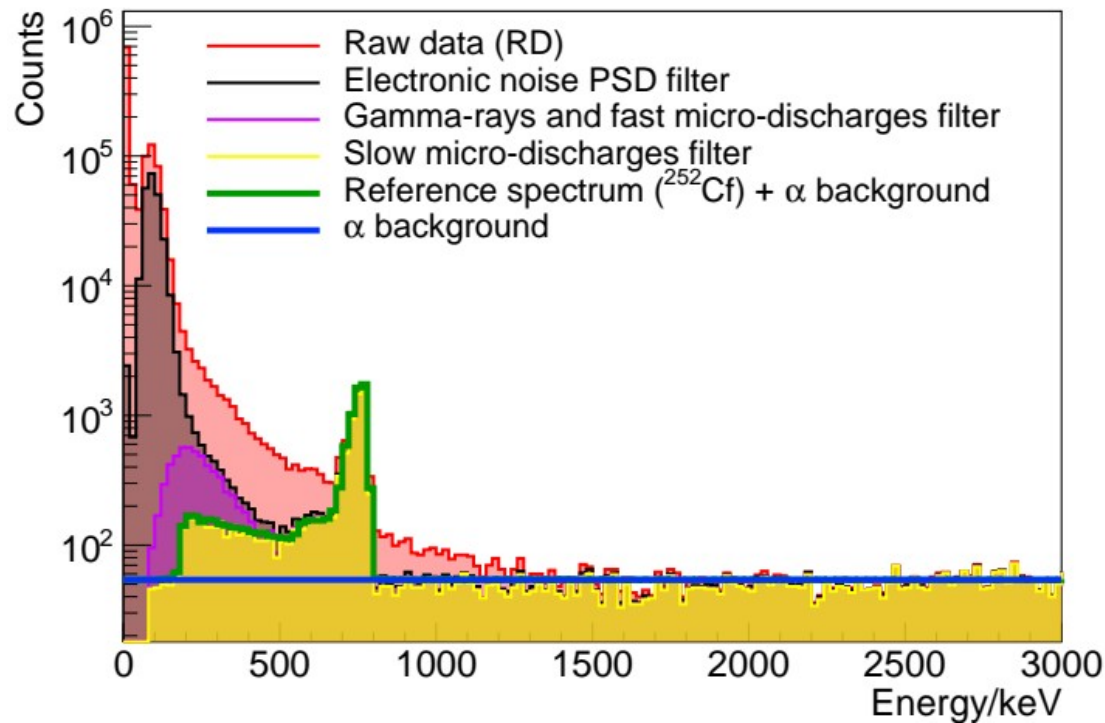
## Pulse Shape Discrimination (“slow” discharges filter)



## Pulse-height spectrum (+ PSD)



## Pulse-height spectrum at LSC (comparison with a pure $^{252}\text{Cf}$ neutron spectrum)



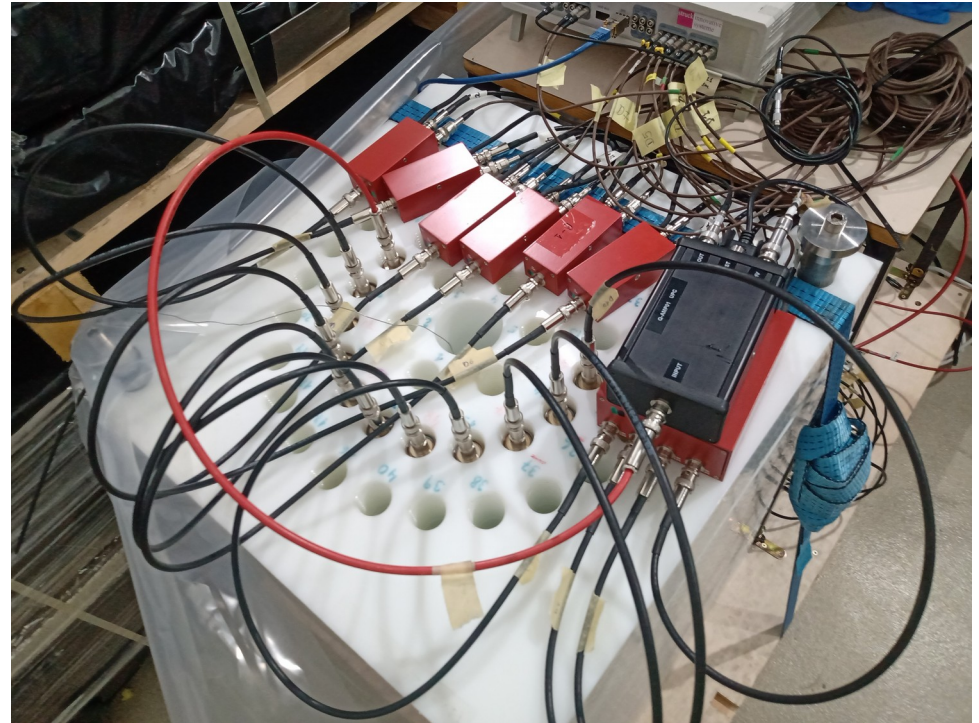
## Characterization of the alpha background

Characterization of the intrinsic background at Canfranc (LSC, Aug 23 – Feb 24)

- Counters completely covered with cadmium.
- Embedded in a HDPE shielding.

The result is that all neutrons are absorbed before reaching the  $^3\text{He}$ .

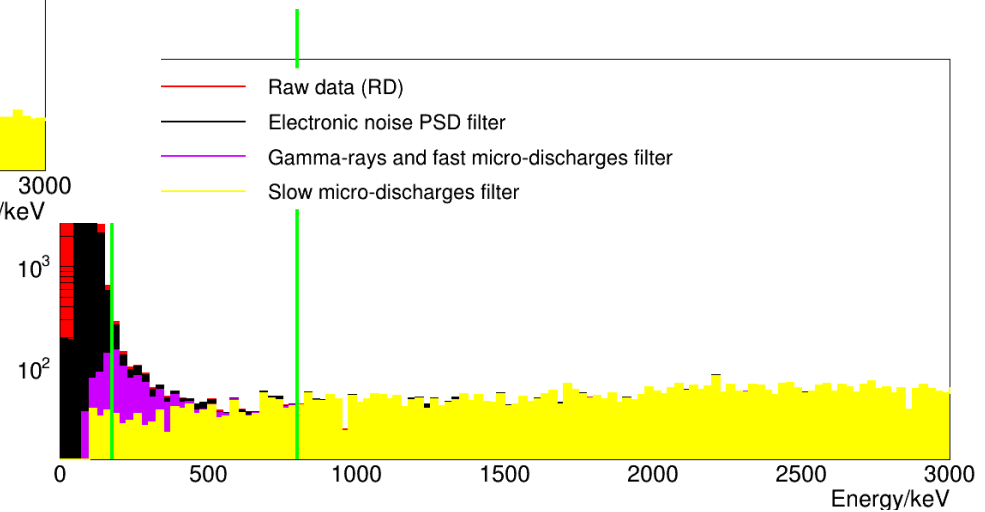
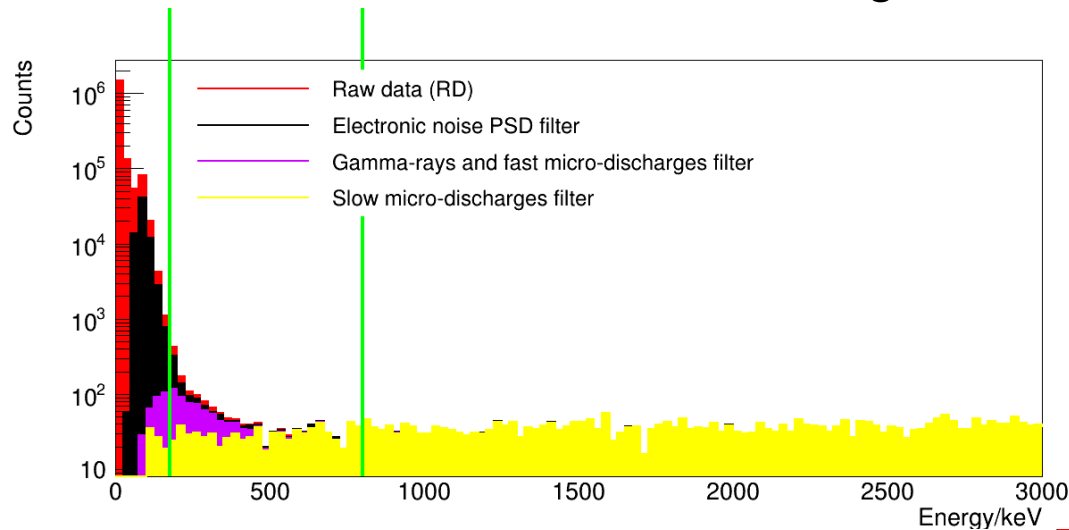
No neutrons are detected.





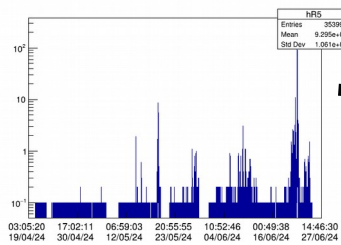
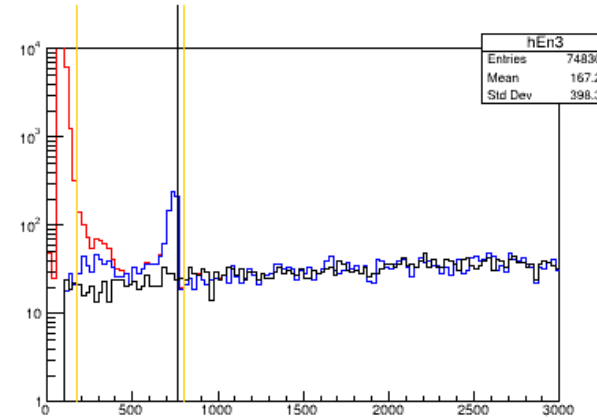
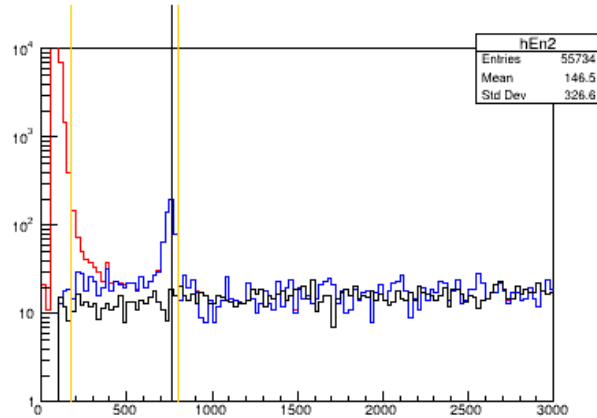
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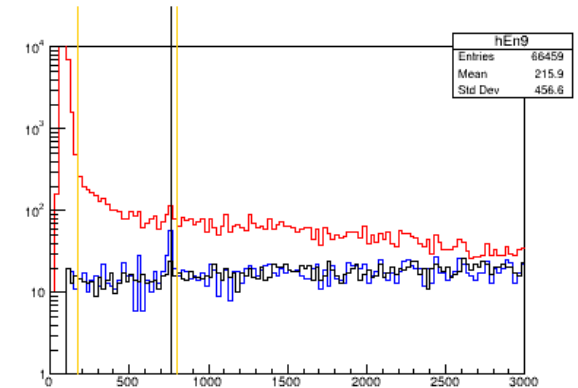
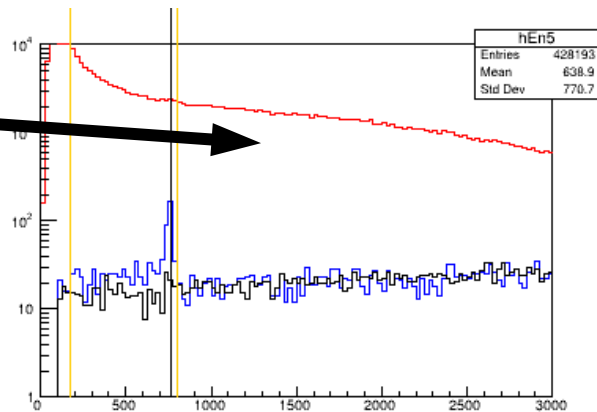


## Subtraction of the alpha background

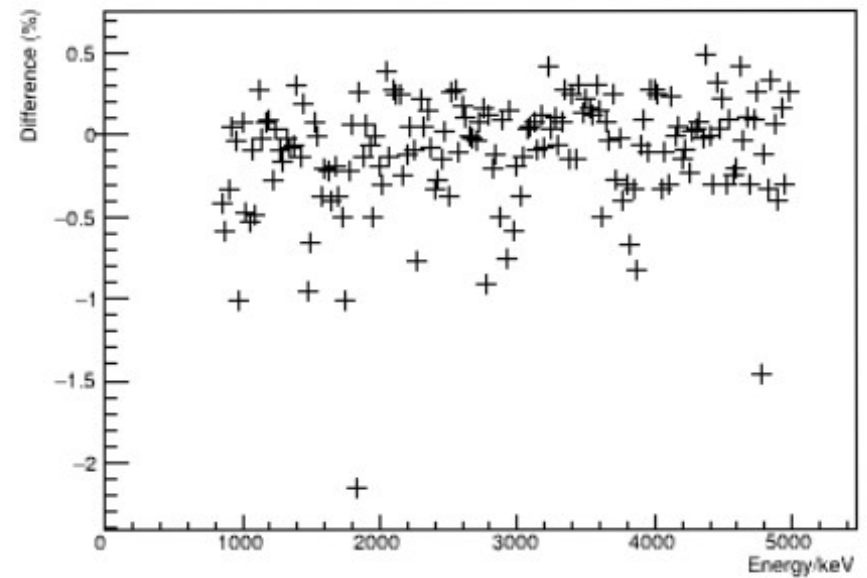
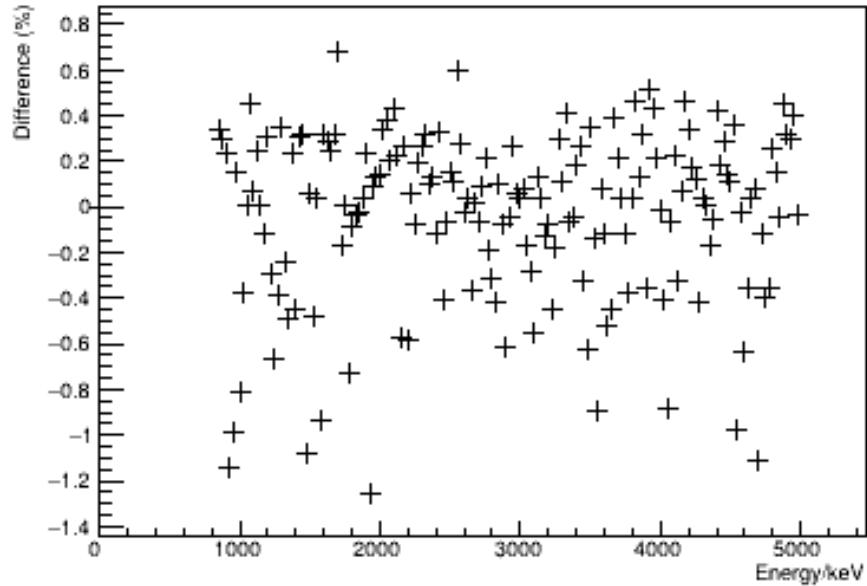
The alpha background is our main limitation!



High micro-discharges rate periods (max: ~ 200 cpm)



## Subtraction of the alpha background (residuals)



## Subtraction of the alpha background (counting time: 59 days)

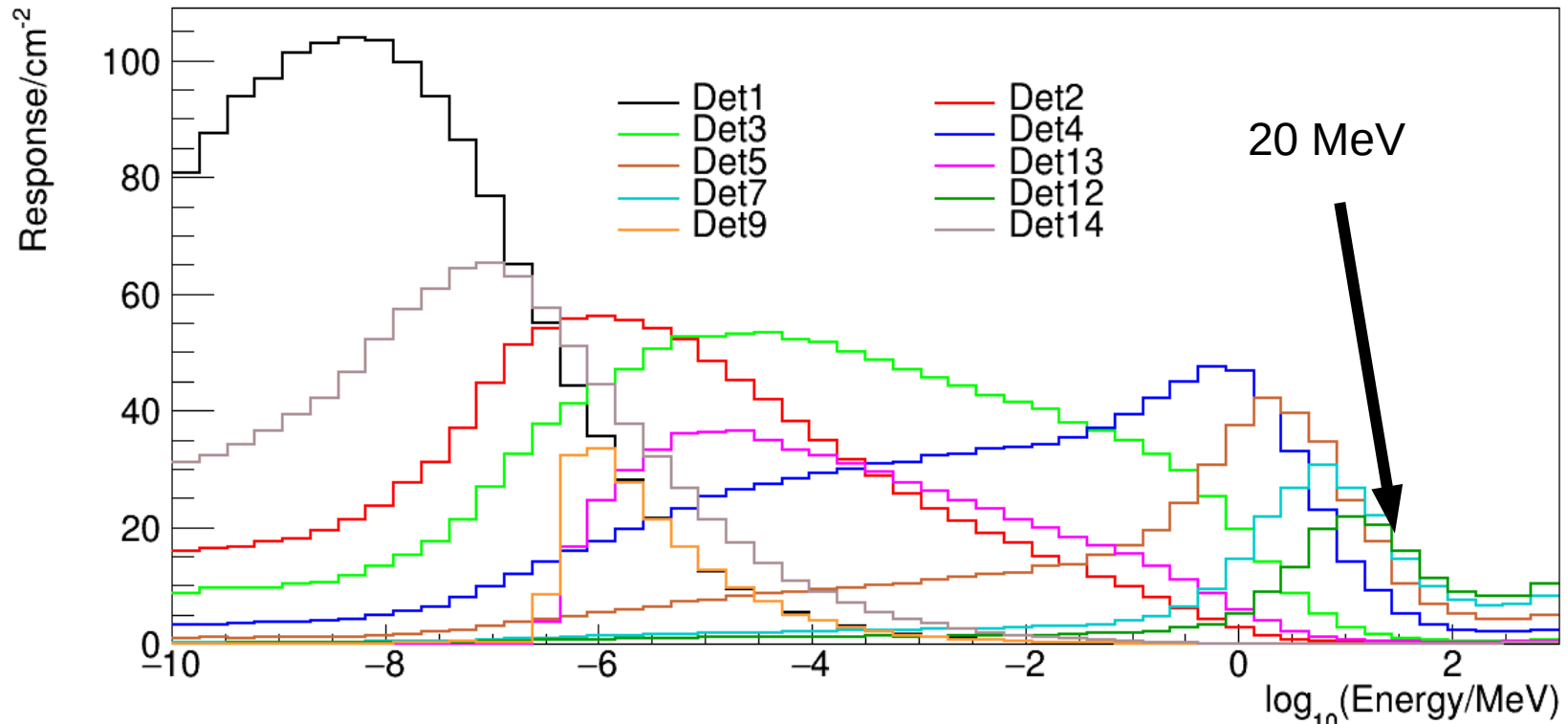
Detector	Rate (cps)	Error (cps)	Error
1	1.38E-04	7.31E-06	5.3%
2	1.48E-04	7.37E-06	5.0%
3	1.76E-04	8.55E-06	4.9%
4	1.42E-04	7.47E-06	5.3%
5	1.12E-04	6.82E-06	6.1%
7	4.51E-05	5.87E-06	13.0%
9	1.33E-05	5.30E-06	39.9%
12	1.67E-05	5.38E-06	32.3%
13	6.97E-05	6.15E-06	8.8%
14	1.25E-04	7.13E-06	5.7%

## Substraction of the alpha background (counting time: 59 days)

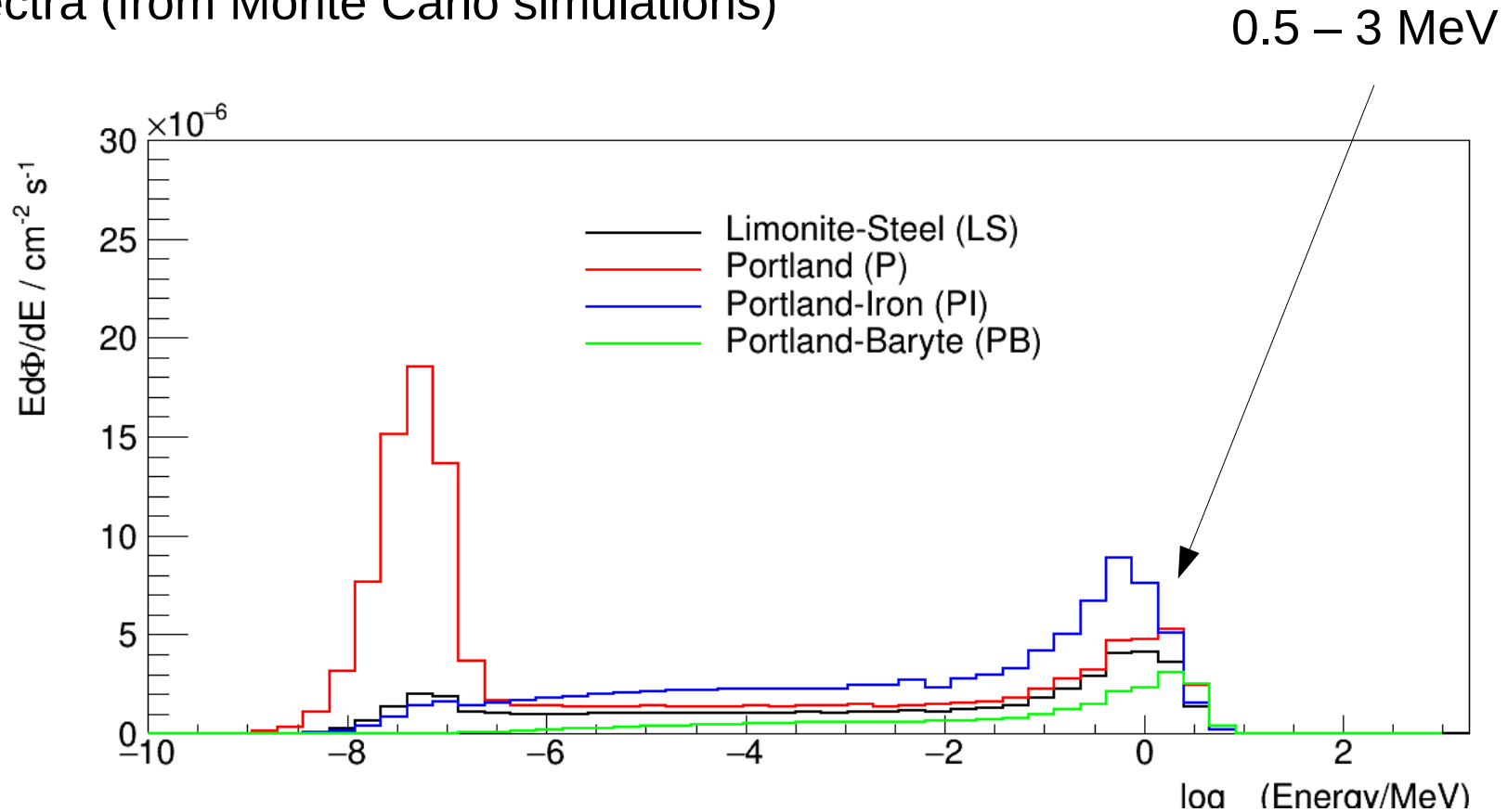
Detector	Rate (cps)	Error (cps)	Error
1	1.38E-04	3.28E-06	2.4%
2	1.48E-04	2.77E-06	1.9%
3	1.76E-04	3.96E-06	2.2%
4	1.42E-04	3.11E-06	2.2%
5	1.12E-04	2.92E-06	2.6%
7	4.51E-05	2.97E-06	6.6%
9	1.33E-05	3.20E-06	24.1%
12	1.67E-05	2.90E-06	17.4%
13	6.97E-05	2.89E-06	4.1%
14	1.25E-04	3.28E-06	2.6%

## Energy spectrum

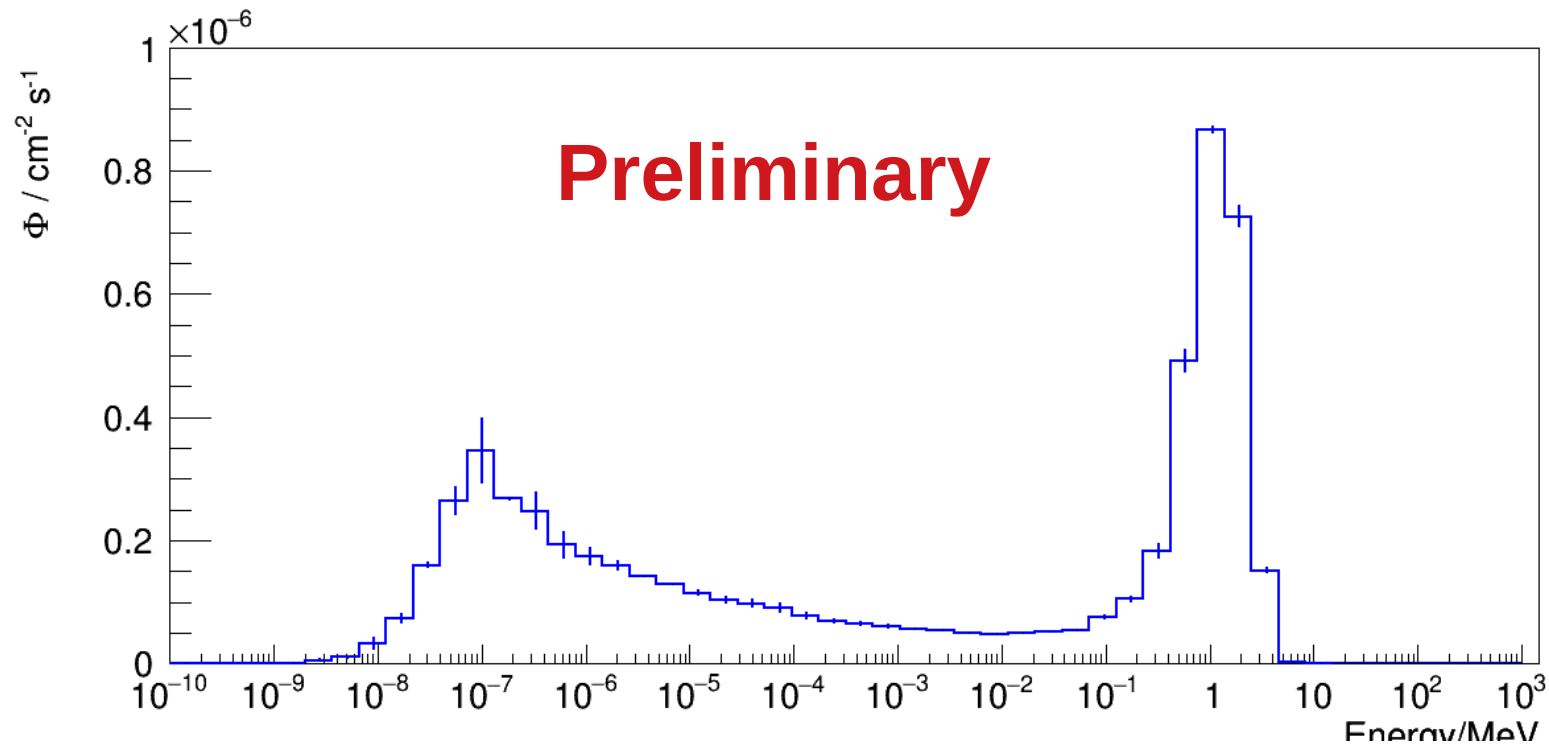
HENSA sensitivity from thermal up to 20 MeV



## Energy spectrum (up to ~ 20 MeV) Prior spectra (from Monte Carlo simulations)



## Energy spectrum: (average from different priors)





## Neutron flux in hall A (preliminary)

	Thermal	Intermediate	Fast flux	Total flux
LNGS (hall A)	<b>1.40(6)</b>	<b>2.16(5)</b>	<b>2.60(3)</b>	<b>5.85(8)</b>
LSC (hall B)	7.45(1)	7.425(7)	5.57(2)	20.45(2)
Belli 1989*	1.08(2)	-	0.86(14)	3.78(26)
<i>Best (2016)**</i>	<i>0.320(130)</i>	-	-	-

\*P. Belli et al. (1989): a measurement of the neutron flux in hall A of the LNGS using an array of 5 large BF<sub>3</sub> counters shielded with polyethylene and cadmium. The upper limits for thermal region they used is 0.5e-7 eV.

\*\*Not in hall A, but in the the “interferometer tunnel”

Thermal:  $0 < E < 3.2E-7$  MeV

Intermediate:  $3.2E-7 < E < 0.1$  MeV

Fast:  $0.1 < E < 20$  MeV

About Belli (1989):

Maximum polyethylene shielding was 140 mm. In HENSA is 380 mm.

- **HENSA is a high efficiency neutron detection system based on the same principle than Bonner Spheres providing sensitivity to neutrons from thermal energies up to MeV.**
- **HENSA has been successfully used for many years at the Canfranc Underground Laboratory (LSC) in order to characterize the underground neutron flux.**
- **A replica of HENSA has been assembled at LNGS. First results have been presented.**

We acknowledge Spanish Grant PID2022-138297NB-C22 funded by MCIN/AEI and “ERDF a way of making Europe” and the MRR (Recuperation and Resilience Mechanism) co-financed by the Generalitat de Catalunya, the Spanish MCIN and the EU.