Report from SAND Calibration WG

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SAND Calibration WG

- Calibration: from detector signals to physical variables
 - ECAL: energy, time and positions of the particles
 - GRAIN: tracks, time, energy,
 - Tracker : r-t relations, track momentum, dE/dx for PID,
 - Timing alignment among the subdetectors
- Define a strategy for each subdetector:
 - Sources: cosmics, particles from beam, ...
 - Choose suitable processes (given the expected fluxes of particles in the detector, e.g. for the ECAL: cosmic μ's as MIPs, MIPs from the beam, electrons and photons)
 - Set a calibration procedure (Which level of precision ? How much time expected ?)
 - Reference people: ECAL P.Gauzzi, GRAIN: A.Surdo, Tracker:
- WG meetings generally every three weeks, on Thursday at 3 p.m. CET
- WG mailing list: <u>dune-nd-sand-calibration@fnal.gov</u>

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Studies on ECAL Calibration

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- Calibration constants C_i determined with cosmic rays, Data-taking without circulating beams: muons = MIPs
- 2.5 kHz of cosmics \Rightarrow "golden" MIPs, ~ 100 Hz



1 day data-taking $\Rightarrow \sim 10^3$ evts/cell C_i = peak of the MIP distribution $\Rightarrow \sim 1 - 2$ % accuracy Repeated every few months



- Average energy scale 38 MeV / MIP crossing a cell at the center (measured at test beams)
- Absolute energy scale set with Bhabha scattering events (e⁺e⁻→e⁺e⁻) and e⁺e⁻→ γγ: showers of 510 MeV
- Repeated every run (every 1 or 2 hours)
- $4 5 \times 10^4$ Bhabha evts in the Barrel $O(10^5)$ in the Endcaps $10^3 - 10^4$ yy events in one run



Sapienza

ECAL calibration in SAND

MIPs from cosmic rays: muon flux at surface ~ 0.02 μ /(s cm²)

 $\Rightarrow \sim 10^4 \; \mu/s$ on ECAL (\Rightarrow 100 Hz of "golden mips" in KLOE)

- Underground reduction of a factor of about 100 $\Rightarrow \sim 100 \mu/s$ on ECAL (no selection)
- Rough estimate by rescaling the KLOE numbers \Rightarrow 1 day (24 hrs): ~ 10 evts/cell
- Relaxing the "golden mip" selection: in few days ~ 10³ evts/cell

MIPs from beam (rock, magnet and Fe yoke,

upstream ECAL modules)

- We need also muons from beam for the modules around the median plane and for the endcaps
- Started MC study of the rate of muons from beam events reaching the ECAL





- Generation of 100000 \textit{v}_{μ} events with vertices in the hall and in the rock surrounding the hall



• Selecting events with at least a muon in the ECAL \Rightarrow ~ 2000 evts.







- All events
- Events with at least one muon in the ECAL
- We can restrict the generation window to DUNE_ND_HALL (X and Y in ~ -6.0 – 6.0 m) and to cut at Z > -10 m



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- Test: generation of 25000 \textit{v}_{μ} in that window \Rightarrow 797 events with at least 1 cluster from μ
- This sample corresponds to ~ 30 spills





MIPs from beam Events / (3.33333

- Golden mips: all the cluster cells in the same column
- Low statistics
- Clean distribution
- Good peak fit

- Less stringent selection: at least 3 cells in the same column
- Peak still clear





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Events / (3.33333

1400

1200

1000

800

600

400 200

> 10 20

- Occupancy:
 - No conditions on muon clusters

- At least 3 cells in one column

- Golden mips



Energy scale calibration in SAND

- γ 's from π^0 decays, invariant mass reconstruction (need a vertex from the tracker)
- γ + electrons: ~ 30% of photons from π^0 convert in the tracker
 - $\Rightarrow \sim 50\%$ of π^0 have at least one $\gamma \rightarrow e^+e^-$ (from DUNE-doc-13262 A Near Detector for DUNE)
- High energy electrons from v_e interactions \Rightarrow need the momentum measurement in the tracker
- Possibility to exploit $K^0 \rightarrow \pi^0 \pi^0 \rightarrow 4\gamma$
- From a naive rescaling of K⁰→π⁺π⁻ ⇒ O(10⁵) evts in 5 years of FHC data-taking
- Reconstruct a vertex with the ECAL only, back-propagating each of the 4 photons, but the times of the ECAL cells must be very well aligned







Time calibration

 $-t_{G}^{0}$

 \mathcal{X}

 MIPs: uniform illumination of the ECAL for calibration of time and coordinate along the fibers

$$t = \frac{1}{2}(t_A + t_B) - \frac{L}{2v} - t_0$$

$$A \xleftarrow{x} B$$

$$= \frac{1}{2}v(t_A - t_B) - \Delta t_0$$

$$\Delta t_0 = \frac{1}{2}(t_A^0 + t_B^0)$$

$$\Delta t_0 = \frac{1}{2}(t_A^0 - t_B^0)$$



Fit function: sum of two Error functions Width \Rightarrow 2L/v , L =430 cm fixed, v free parameter \approx 17 cm/ns

 t₀'s from fit of straight tracks (p > 6 GeV): cosmic muons and beam muons



Conclusions

- We need both cosmic and muon beams for the ECAL calibration
- Evaluating the rate of good muons from beam from MC
- Next steps:
 - Produce few x 10⁶ events to increase the statistics
 - Define a strategy to calibrate the Endcaps and the modules with low statistics
 - Start the study of the energy scale calibration
- Strategy of time alignment with the other subdetectors has to be studied









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14 P.Gauzzi DUNE Italia meeting

Time calibration in KLO



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Time calibration in KLOE



 $E_i^{(A,B)}[\text{MeV}] = \frac{(Q_i^{(A,B)} - P_i^{(A,B)})[\text{ADC counts}]}{C_i[\text{ADC counts}/\text{MIP}]} K \times f_{MIP2MeV}[\text{MeV}/\text{MIP}]$

- C_i = peak of the MIP distribution
- Corrections to the C_i with the Bhabha scattering events $(e^+e^- \rightarrow e^+e^-)$: showers of 510 MeV
- Absolute energy scale K fixed at cluster level with the $e^+e^- \rightarrow \gamma\gamma$ events

$$\Rightarrow$$
 Calib. Const. $= \frac{K}{C_i}$



- Calibration constants C_i determined with cosmic rays, Data-taking without circulating beams: muons = MIPs
- 2.5 kHz of cosmics



f "golden" MIPs, ~ 100 Hz

 μ crossing one column (almost orthogonal to the module, within 10°) at the module center (± 20 cm 200 in the longitudinal coordinate) 150

- 1 day data-taking $\Rightarrow \sim 10^3$ evts/cell
- C_i = peak of the MIP distribution $\Rightarrow \sim 1 - 2 \%$ accuracy
- Repeated every few months
- Used to equalize HVs to have uniform trigger thresholds



- Average energy scale 38 MeV / MIP crossing a cell at the center (measured at test beams)
- Absolute energy scale set with Bhabha scattering events (e⁺e⁻→e⁺e⁻) and e⁺e⁻→ γγ: showers of 510 MeV
- Repeated every run (every 1 or 2 hours) (~ 100 nb⁻¹ in KLOE, ~ 1 pb⁻¹ in KLOE-2)
- $4 5 \times 10^4$ Bhabha evts in the Barrel $O(10^5)$ in the Endcaps

 $10^3 - 10^4 \gamma\gamma$ events in one run







400 450 500 550 600 650 700

300 350

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- Linearity of the response and energy resolution measured with radiative Bhabha scattering $(e^+e^- \rightarrow e^+e^-\gamma)$ by detecting the charged tracks in the drift chamber
- Linearity within 1% for E > 70 MeV $\frac{E_{cl} E_{\gamma}}{E_{\gamma}}$

$$E_{\gamma} = \sqrt{s} - E_{+} - E_{-}$$

 E_+ and E_- from p_+ and p_- measured in the Drift chamber (much better resolution for charged tracks)

• For $E = 100 \text{ MeV} \Rightarrow \sigma_E = 18 \text{ MeV}$





• Measured with different processes: $\phi \rightarrow \pi^0 \gamma \ (\pi^0 \rightarrow \gamma \gamma)$, $\phi \rightarrow \eta \gamma \ (\eta \rightarrow \gamma \gamma), \ \phi \rightarrow \pi^+ \pi^- \pi^0, \ e^+ e^- \rightarrow e^+ e^- \gamma$

$$\sigma_t = \frac{57 \text{ ps}}{\sqrt{E \text{ [GeV]}}} \oplus 140 \text{ ps}$$

 The constant term has two contribution: a term common to all the cells, due to the spread of the DAΦNE Interaction Point position, and a proper constant term, uncorrelated among cells, due to a residual miscalibration

140 $\mathrm{ps}=92~\mathrm{ps}\oplus105~\mathrm{ps}$

