

Solar neutrinos in DUNE and SoLAr

Activities in Milano-Bicocca

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Opportunities at low-energy

Dominant interaction channel:

CC interaction $\nu_e(^{40}\text{Ar}, ^{40}\text{K}^*)e^-$ (eff. thrs. ≈ 5 MeV)

\leftrightarrow final state: $e^- + \text{few } \gamma$

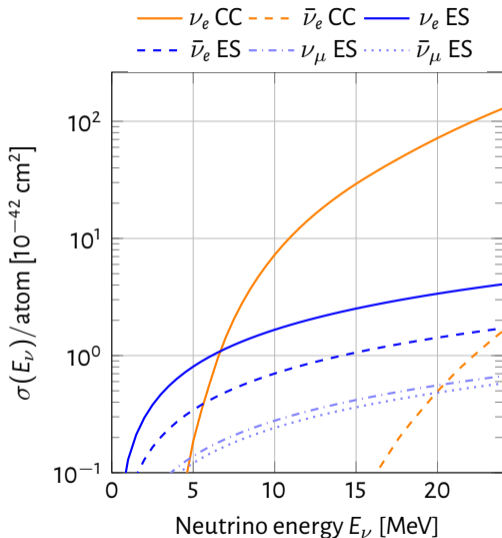
- Supernova neutrinos
- ^8B + *hep* solar neutrinos
 - ^8B : Complementary measurement of θ_{12} and Δm_{12}^2 + Solar core temperature
 - *hep*: Only component of the solar neutrino flux still undetected

LArTPC can probe this energy region
without charge multiplication

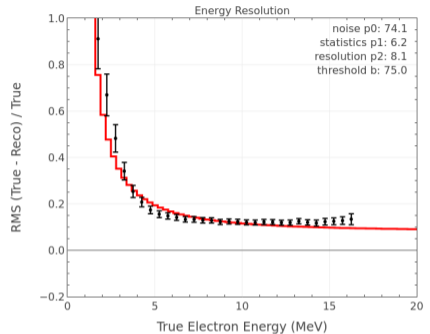
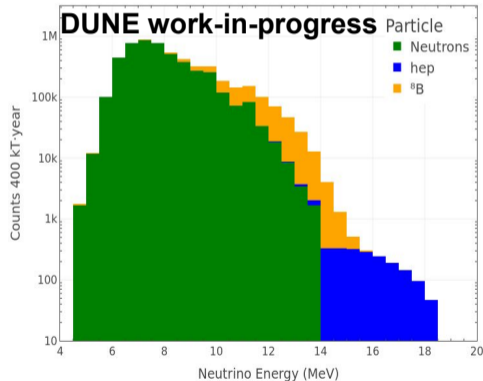
Complementary to LS/Cherenkov experiment

MicroBooNE, arXiv:2410.18419

MicroBooNE, *Phys.Rev.D* 109 (2024), arXiv:2307.03102



- Reconstruction for low energy events using a hit-clustering scheme based on channel and time proximity (3 channels, 12.5 μ s window).
- Electron energy resolution $\approx 10\%$ (ideal drift reconstruction).



Challenges:

- High background level
- Difficult matching between PDS and TPC clusters
 - ➔ Flash-matching proposal
- Energy resolution can be improved
 - ➔ Charge+Light calorimetry at low-energy

PDS simulation...

➤ **Production** phenomenological model modifying the Birks' charge recombination model and reproducing light/charge anticorrelation and its dependence with dE/dx and E_{field} [F. Marinho et al., 2022 JINST17 C07009]

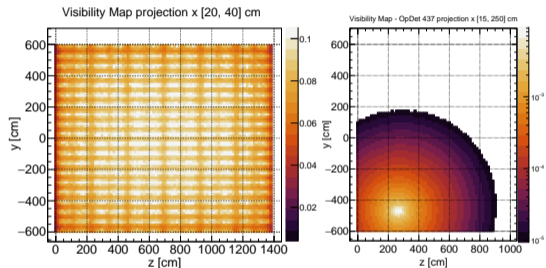
➤ **Propagation** photons ray-tracing unbearable in G4

↪ **Semi-analytical** light propagation model

$$N_{\text{ph hit}}^{\text{opdet } i} = N_{\text{ph}} \cdot Q_{\text{abs}} \cdot Q_{\text{trans}} \cdot G(d, \theta) \cdot T(d, \theta) \cdot QE_{\text{eff}}$$

[D. Garcia-Gamez et al, Eur.Phys.J.C 81 (2021) 4, 349]

PhotonVisibilityExport_module for producing 3D visibility maps (full detector and single opdets)
[duneopet PR#64]

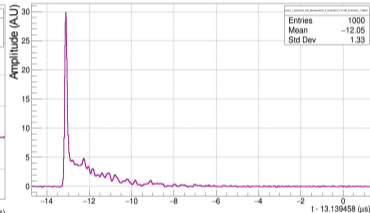
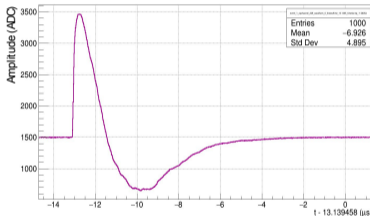


➤ **Digitization** produce a digitized waveform for each PDS channel

OpDetDigitizerDUNE_module uses a **testbench model** for the SiPM's single p.e. response

- **Waveform filtering** for restoring the “true” photon hit’s time profile removing the SiPM impulse response

Deconvolution_module implements two digital filters (Wiener, SBND) for removing the possible undershoot in the SiPM response
Now used in ProdoDUNE-HD LArSoft analysis



PDS simulation and reconstruction - automatic validation

Motivation

Automatize the production of high statistics datasets and the comparison of selected distributions with a reference sample

Originally focused on PANDORA reconstruction, now including much more (gen, sim, hit level)
+ **PDS**

General structure

- Process run using the `lar_ci` framework
- Define a **workflow** specifying the nr of jobs + number of events + progress of simulation & reconstruction stages [gen, g4, sim, reco]
- Produce distributions for different variables and compare them with a reference

Workflow

Simulation workflow follows the “standard” adopted for PDS studies within the low-energy WG

- 100k events simulated by MARLEY with flat energy spectrum
- sim and reco stages for light only
- \approx 5 h run time

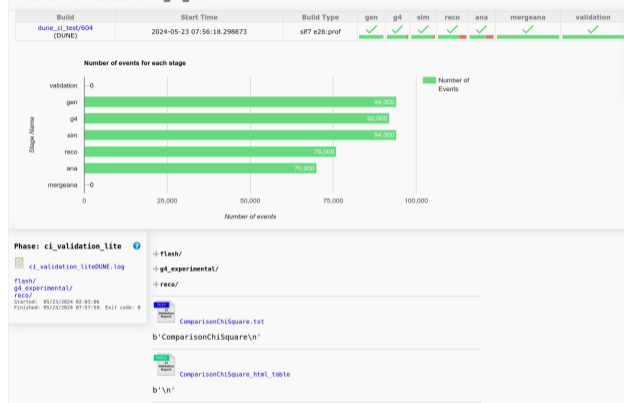
Validation

Prepare histograms and compare with reference

- FD-specific list of histograms keys to be compared
- Compatibility quantified using a simple χ^2 as “test-statistics”
- Compute residuals between validation and reference distributions

PDS simulation and reconstruction - automatic validation

DUNE HD PDS Validation v09_88_00d00



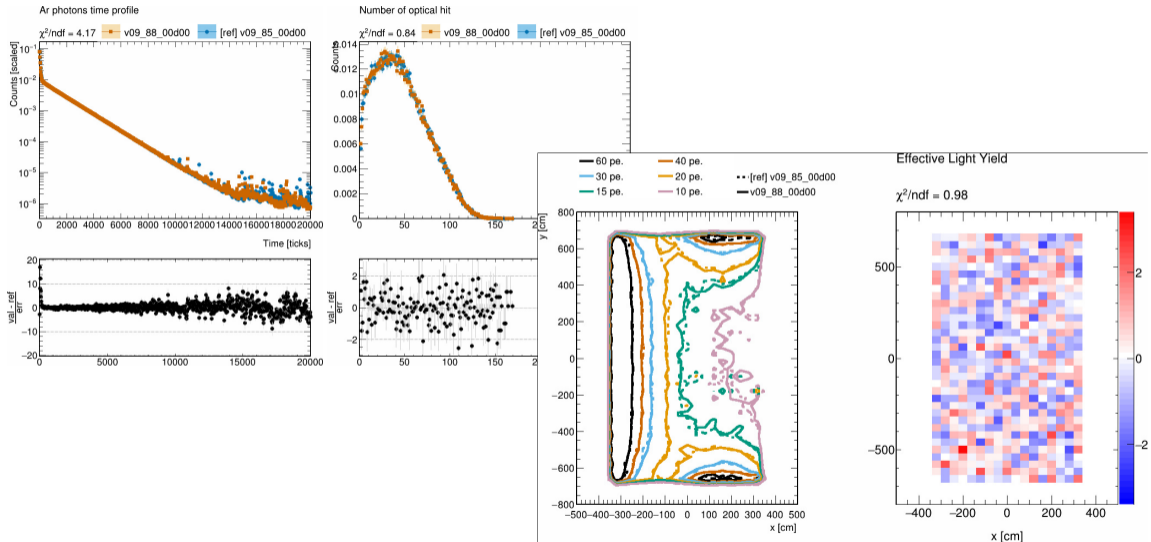
g4
Monte Carlo Truth

Reco
Reconstructed variables

Flash
Flash variables

DUNE Release	Parameter	Chi2
v09_85_00d00 vs v09_88_00d00	reco/effly_map	1.036374
v09_85_00d00 vs v09_88_00d00	reco/ophit_distr	0.875366
v09_85_00d00 vs v09_88_00d00	reco/recope_vs_distance	1.413047
v09_85_00d00 vs v09_88_00d00	reco/recope_vs_opDet	0.874868
v09_85_00d00 vs v09_88_00d00	reco/opdet_avg_pe	0.874868
v09_85_00d00 vs v09_88_00d00	flash/n_opflash	1.616197
v09_85_00d00 vs v09_88_00d00	flash/flash_pe	0.835508
v09_85_00d00 vs v09_88_00d00	flash/flash_pe_event	0.846874
v09_85_00d00 vs v09_88_00d00	flash/flash_purity	454.048895
v09_85_00d00 vs v09_88_00d00	g4_experimental/opdet_avg_gg	1.010757
v09_85_00d00 vs v09_88_00d00	g4_experimental/opdet_avg_trueph_direct	0.876648
v09_85_00d00 vs v09_88_00d00	g4_experimental/opdet_avg_trueph_reflected	0.779359
v09_85_00d00 vs v09_88_00d00	g4_experimental/trueph_time_profile_direct	2.099404
v09_85_00d00 vs v09_88_00d00	g4_experimental/trueph_time_profile_reflected	3.649947
v09_85_00d00 vs v09_88_00d00	g4_experimental/direct_trueph_vs_distance	1.663674
v09_85_00d00 vs v09_88_00d00	g4_experimental/reflected_trueph_vs_distance	2.407212
v09_85_00d00 vs v09_88_00d00	g4_experimental/total_recope_vs_trueph_prf	0.965064
v09_85_00d00 vs v09_88_00d00	g4_experimental/opdet_recope_vs_trueph_prf	0.959080

PDS simulation and reconstruction - automatic validation



Next steps at low-energy

- Combined charge+light calorimetry requires correct matching of PDS and TPC signals
 - Low energy challenge: $\mathcal{O}(100)/(1000)$ flash candidates per TPC cluster in FD HD/VD
 - Idea: make use of the light propagation model to select the most compatible match
- Low-energy events \sim point-like \rightarrow described by 5 parameters $\theta = \{\mathbf{x} = \{x, y, z\}, t, E\}$

TPC hits constrain E +
position on the anode (y, z)

PDS flash constrains E +
time \rightarrow position along drift (x)

For each flash candidate evaluate a likelihood function like

[Patterson et al, *Nucl.Instrum.Meth.A* 608,2009]

$$\mathcal{L}(\{\text{opHit}\}|\theta) = \prod_i^{\text{unhit XA}} (1 - P(i \text{ hit}|\theta)) \cdot \prod_i^{\text{hit XA}} P(i \text{ hit}|\theta) f_q(q_i|\theta) f_t(t_i|\theta)$$

where each term can be evaluated from the detector response and light propagation model (e.g., opDet visibility maps built with PhotonVisibilityExport_module)

Towards DUNE Phase II: SoLAR

Improved background suppression

Passive shielding, material selection, pulse-shape discrimination, event topology

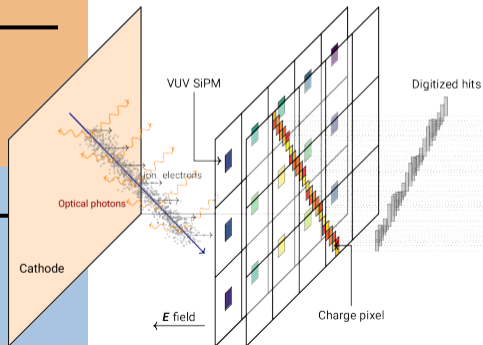


Pixelated readout:

Pixel readout plane will enhance event reconstruction, while replacing TPC wires is expected to simplify construction and installation

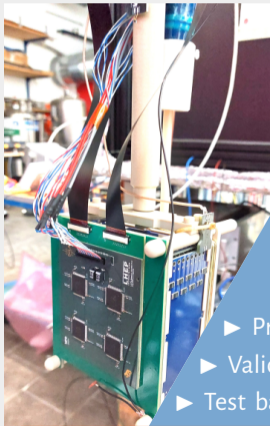
Improved light sensors:

Arapuca-style modules + **VUV SiPMs integrated on the anode**
Exploit the light signal in LAr to perform **combined Q + L calorimetry**: Target $\Delta E/E \approx 7\%$



Phase I: Prototyping

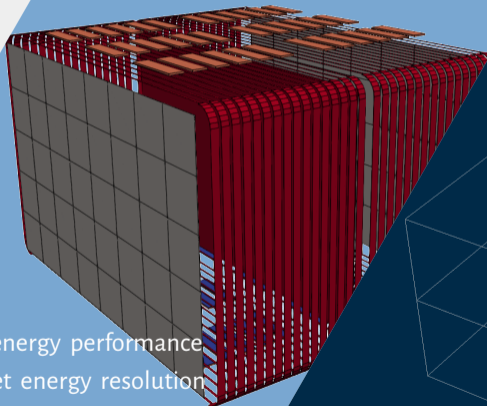
Development of integrated charge+light readout



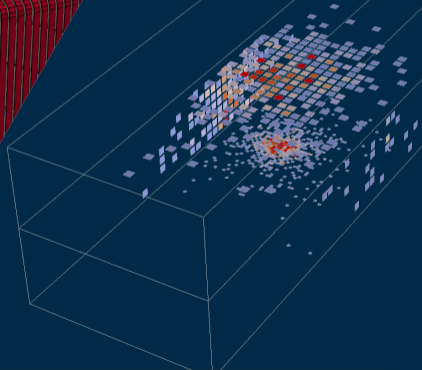
- ▶ Prove low-energy performance
- ▶ Validate target energy resolution
- ▶ Test background suppression methods

Phase II: Medium Scale Experiment

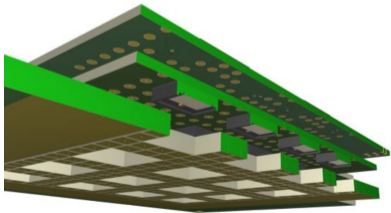
First detection of solar neutrinos in LAr



Phase III: DUNE MoO

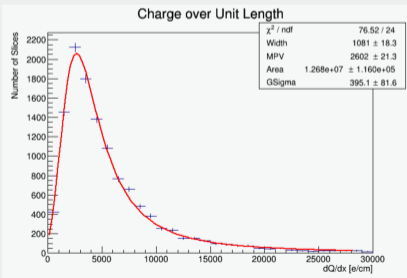
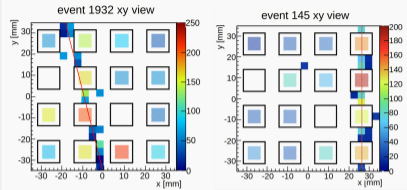


- **Charge readout:**
256, 3×3 mm² pads, readout by 4 LArPix v2a ASIC + PACMAN
- **Light readout:**
16 Hamamatsu S13370-6050CN (6×6 mm²) + Cold pre-amp + Warm amp + 62.5 MS/s digitizer
- **Anode assembly:**
Three stacked PCB layers to accommodate SiPMs packaging SiPM floating bias to enhance charge collection



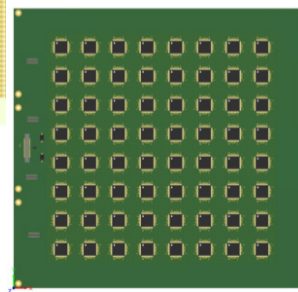
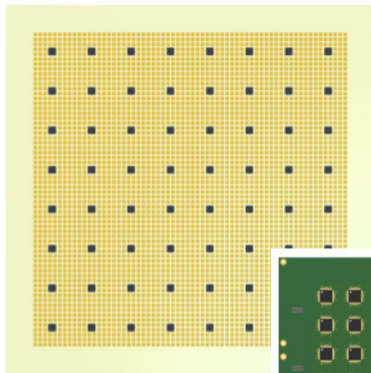
Test outcome

- combined operation of charge+light sensors
- calorimetric response ok



SoLAr prototyping activities: SoLAr_v2

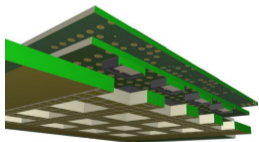
- Tile dimension: $32 \times 32 \text{ cm}^2$
(active area $25.6 \times 25.6 \text{ cm}^2$)
- Divided into 8×8 regions (64 — 4 pixel, 1 SiPM)
- 20 LArPix (room for 64)
- 64 Hamamatsu VUV MPPC with independent readout
- Complete re-design of the PCB



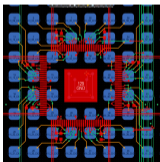
SoLAr tile V1



SoLAr tile V2



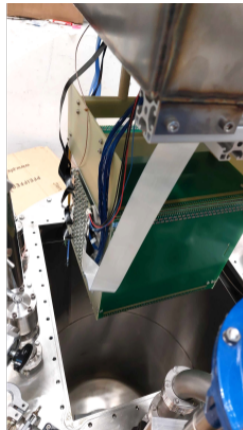
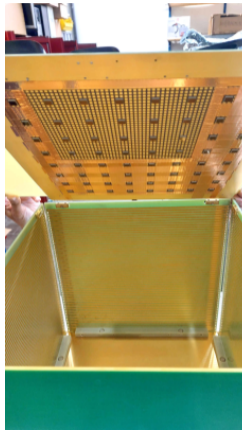
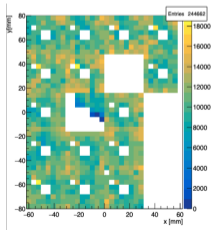
Three PCB stacked



Single multilayer PCB

SoLAr_v2: Operations

- 30 cm drift length
- Un-routed pads grounded with copper tape
- 2 days of cosmic data taking with nominal HV 15 kV + special runs at 7.5 and 3.75 kV
- Additional run with ^{60}Co source
- Good LAr purity
- Low charge hit threshold $\approx 3.8 \text{ ke}$
- few dead areas on the anode

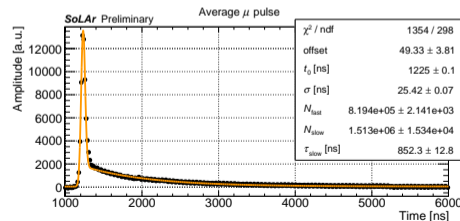
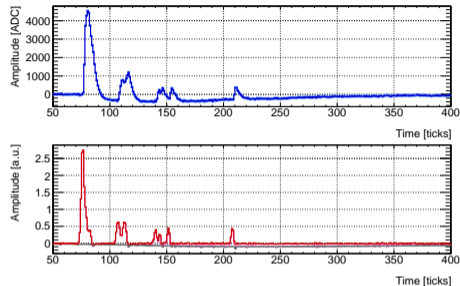


SiPM waveform analysis

- SiPM waveform show a characteristic undershoot due to electronics coupling
- SiPM impulse response studies with dedicated LED runs
- Various waveform filtering strategies (Wiener filter, ...)
- Residual baseline modulation corrected applying a SNIP algorithm

Preliminary results

- Scintillation time profile obtained averaging filtered waveforms from ≈ 80 min cosmic run
- Slow time constant and relative weight of fast/slow component show a small discrepancy with expected values



SoLAr v2: Preliminary results

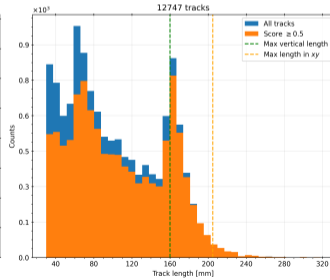
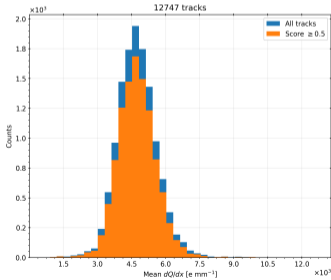
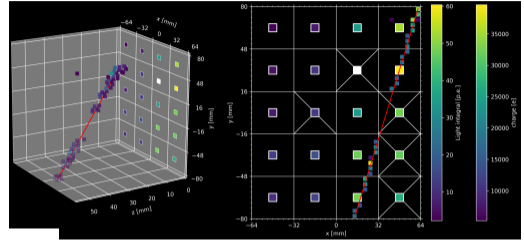
Charge hit analysis

➤ Clustering

Solving ambiguities due to dead areas using simulated data

➤ Track fit

Identify outliers and secondary tracks



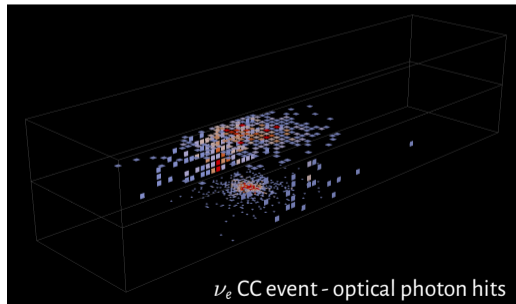
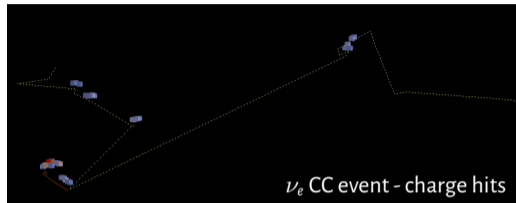
Preliminary results

- Track length distribution influenced by dead areas on the anode tile
- dQ/dx distribution compatible with similar experiments

SoLAr simulation

solar-sim Geant4-based simulation package

- **Geometry** description in external configuration file (including cavern, cryostat)
- **Physics**
 - Multiple generators integration: MARLEY, BxDecay0, CRY, RadSrc, Externals, ...
 - HP Physics for n and low-energy ions
 - Currently integrating G4Cascade for (n, γ) cascades
 - Includes charge-light anticorrelation (LARQL)
 - Geant4 ray-tracing for optical photons
- Include charge diffusion and pixel **readout effects** (in progress)
- **Ongoing studies:**
 - SoLAr_v2 run
 - end-to-end simulation of neutron background
 - oscillation parameters sensitivity



PDS Simulation and Reconstruction

- MiB activities in PDS simulation and reconstruction covers a significant part of the workflow (Digitizer, Deconvolution, Production of visibility maps)
- Develop automatic validation of PDS simulation and reconstruction based on Low-Energy WG studies
- Ready use our expertise on PDS for flash-matching with high background level
 - ↔ key step towards combined calorimetry at low-energy

SoLAr

- Two prototyping runs demonstrated the technological concept
Track reconstruction, matching light flash, analysis ongoing
- Developed a simulation package to establish the detector requirements and guide the design