

Plasma Wakefield Acceleration – Advantages and Challenges

Livio Verra

Fundamental Research and Applications with the EuPRAXIA Facility at LNF

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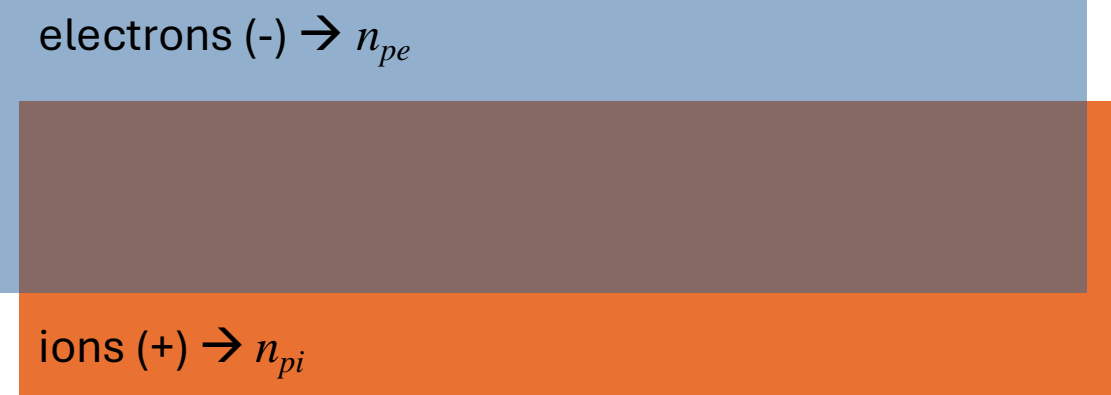


Istituto Nazionale di Fisica Nucleare
Laboratori Nazionali di Frascati

Plasma

Plasma:

- Ionized gas
- Collisions can be (most of time) neglected
→ Electromagnetic interaction dominates
- Large number of particles → **collective** behavior
- Quasi-neutral ($n_{pe} \sim n_{pi}$)



electrons (-) → n_{pe}

The diagram consists of a horizontal rectangle divided into three colored regions. The top region is light blue and contains the text 'electrons (-) → n_{pe}'. The middle region is a darker brownish-grey and is empty. The bottom region is orange and contains the text 'ions (+) → n_{pi}'. The blue and orange regions are of equal height, while the middle region is taller, making the total height of the rectangle greater than the sum of the blue and orange regions.

ions (+) → n_{pi}

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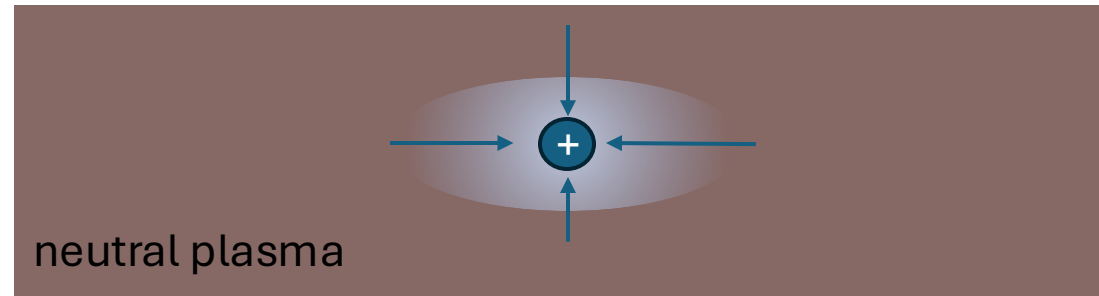


neutral plasma

Plasma

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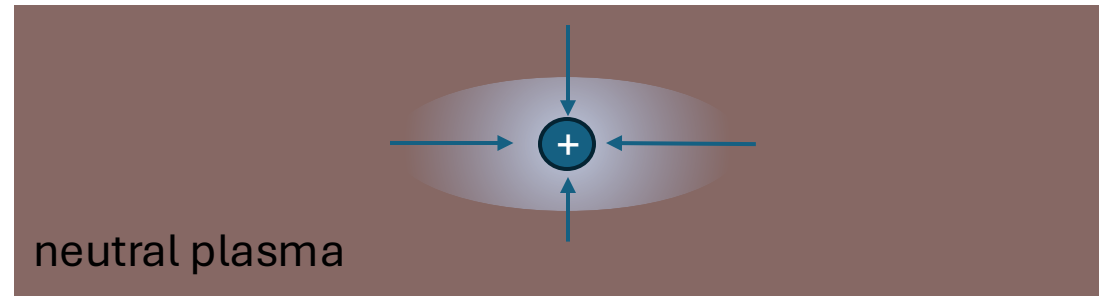
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Plasma electrons ($m_{ip} \gg m_{pe}$) move to compensate for the disturbance
 - Plasma screens electromagnetic fields



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When the equilibrium is perturbed:

- Electrons oscillate with angular frequency $\omega_{pe} = \sqrt{\frac{n_{pe}e^2}{m_e\epsilon_0}}$
- Ions with $\omega_{pi} = \sqrt{\frac{n_{pi}e^2}{m_i\epsilon_0}} \ll \omega_{pe}$ (ions considered immobile for short time-scales)

Plasma Wakefield Acceleration

- Let's take a plasma with density n_{pe}
- Let's take a relativistic charged bunch (e.g. e^-) or a laser pulse

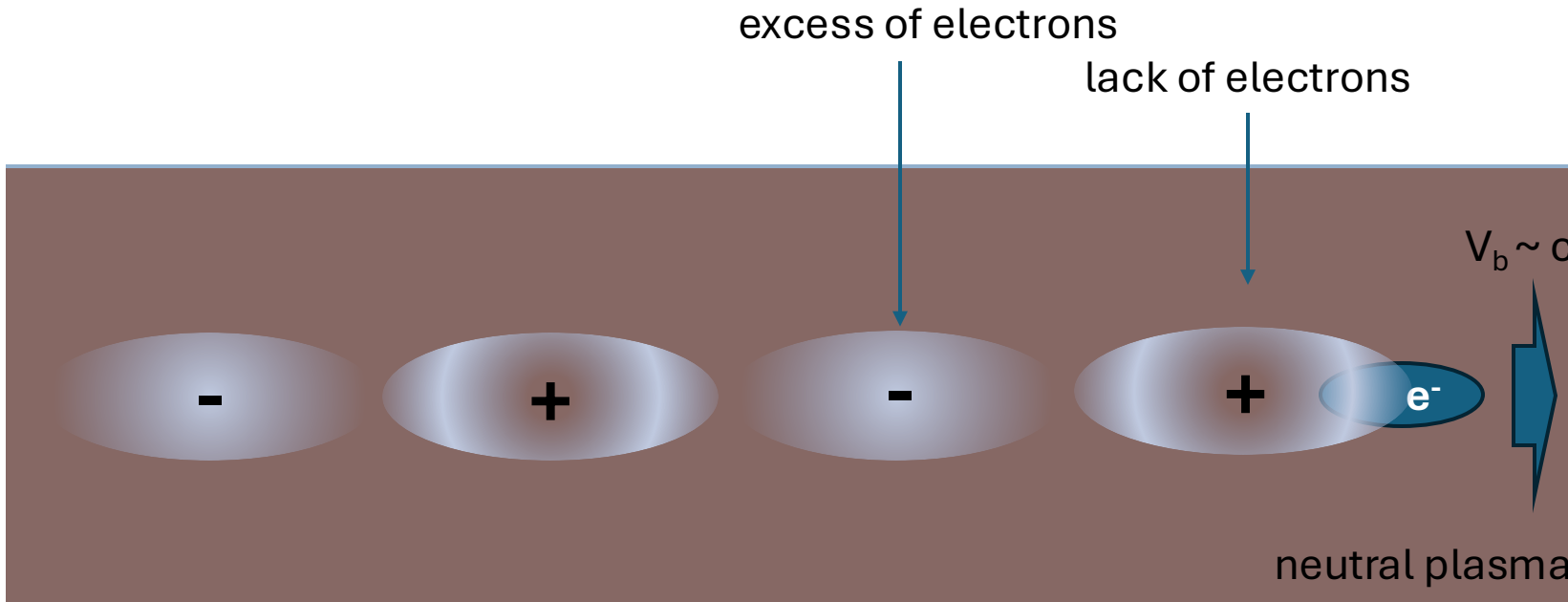
1. Transverse E field expels plasma electrons
2. Positively charged region behind the bunch head
→ restoring force



(inspired by P. Muggli's CAS lecture)

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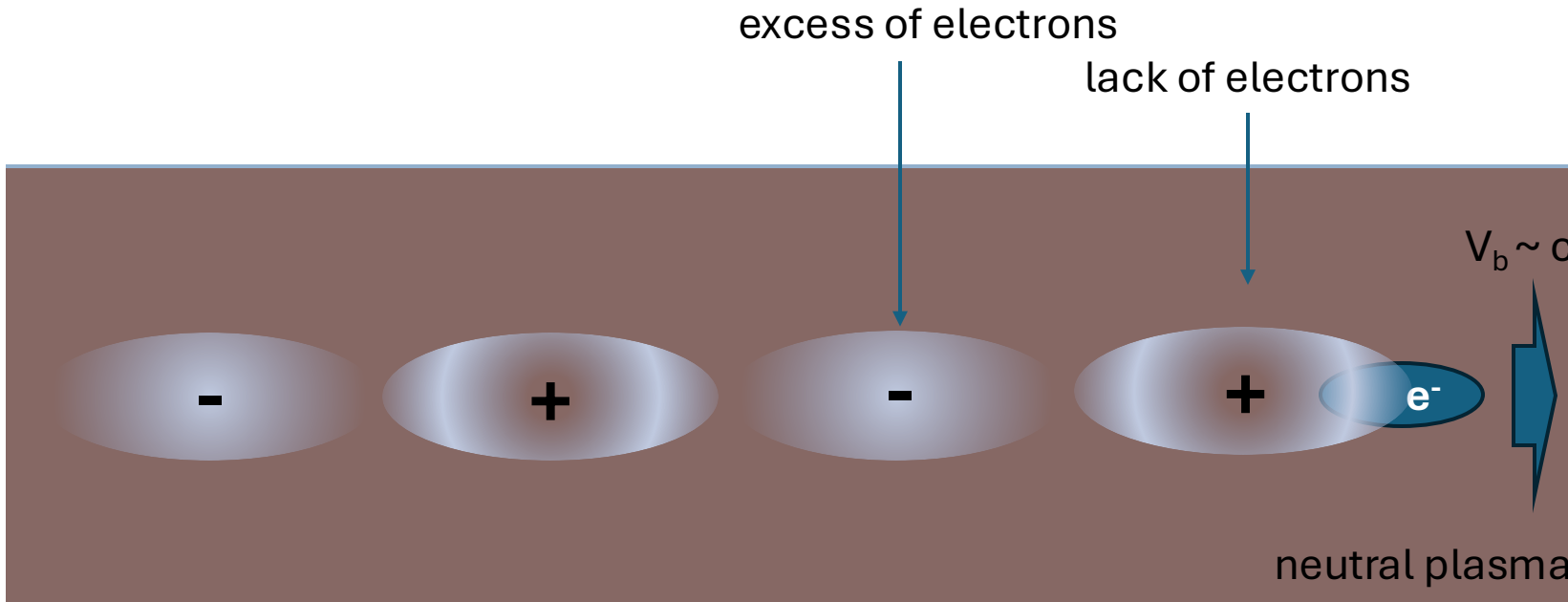


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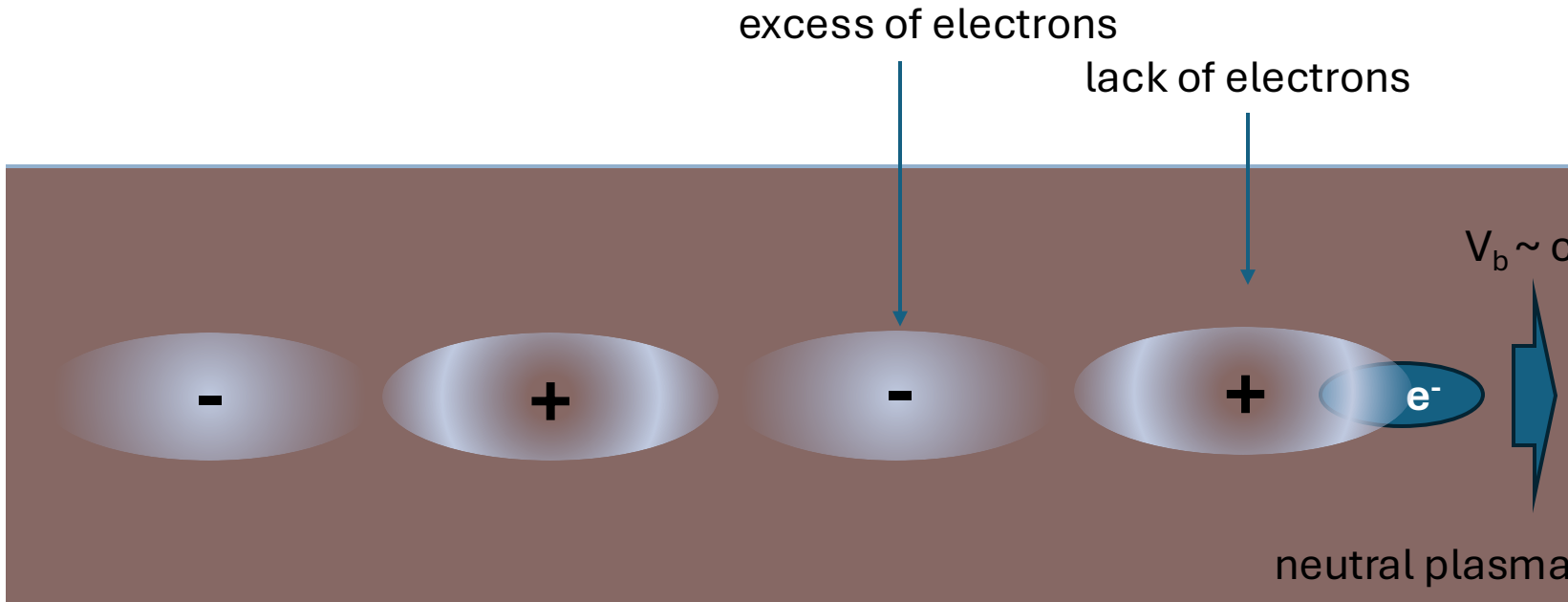
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3. Oscillation of plasma e^- with ω_{pe}
→ periodic density variation

(inspired by P. Muggli's CAS lecture)

$$\lambda_{pe} = \frac{2\pi c}{\omega_{pe}}$$

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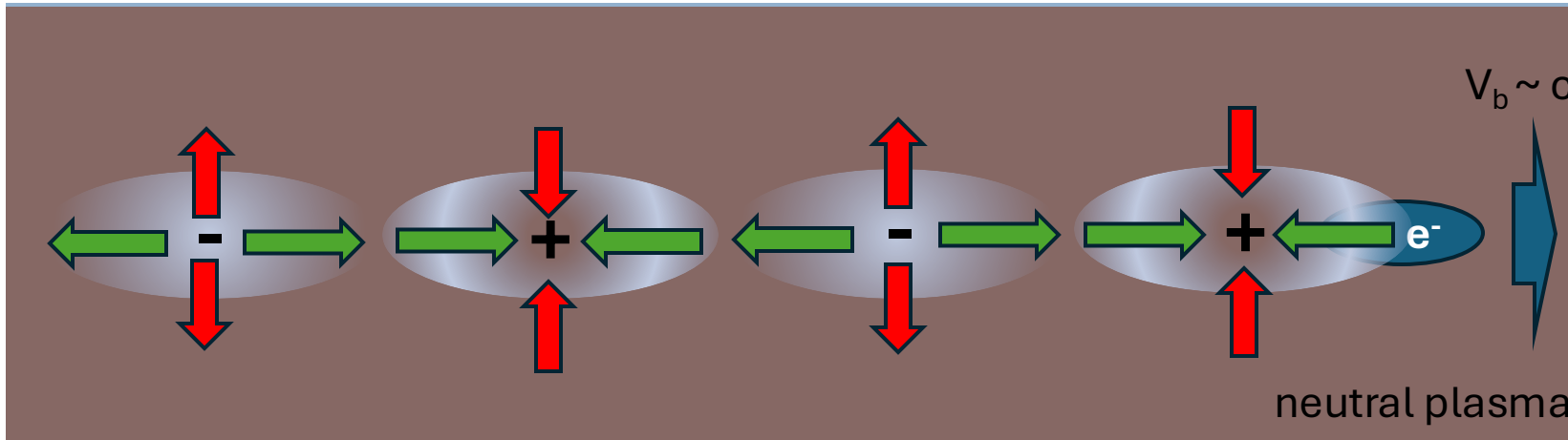
$$\nabla \cdot \vec{E} = \frac{\rho}{\epsilon_0}$$

→ Wakefields ←

Plasma Wakefield Acceleration

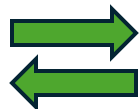
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 Longitudinal (accelerating – decelerating) wakefields

 Transverse (focusing – defocusing) wakefields

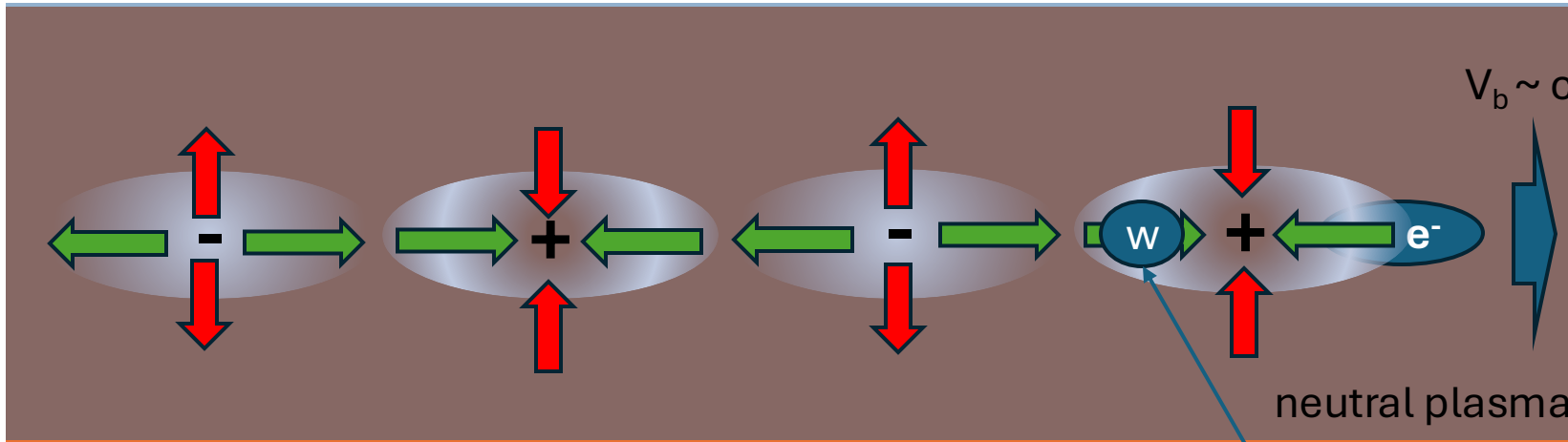
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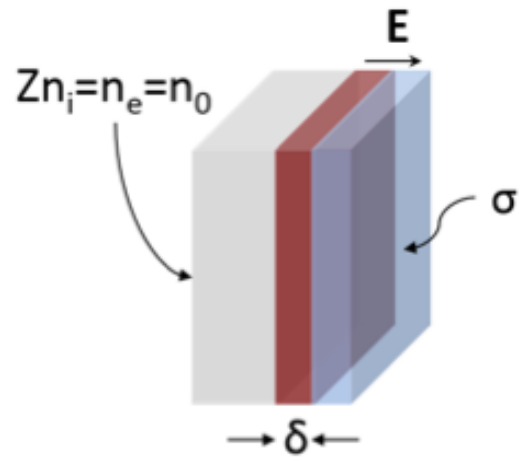
$$\nabla \cdot \vec{E} = \frac{\rho}{\epsilon_0}$$

→ Wakefields ←

A witness bunch can be accelerated to high energies if injected in the right phase!

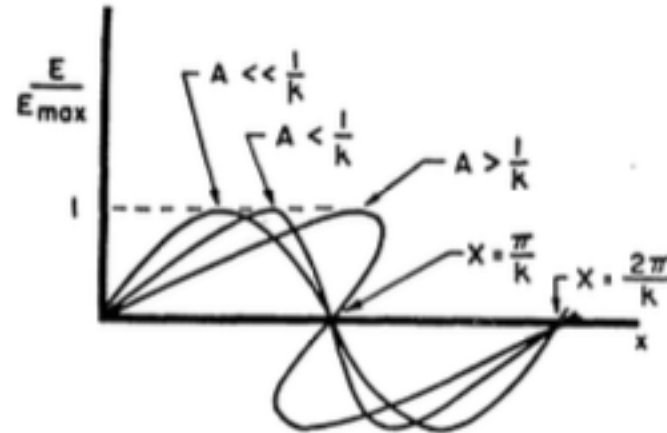
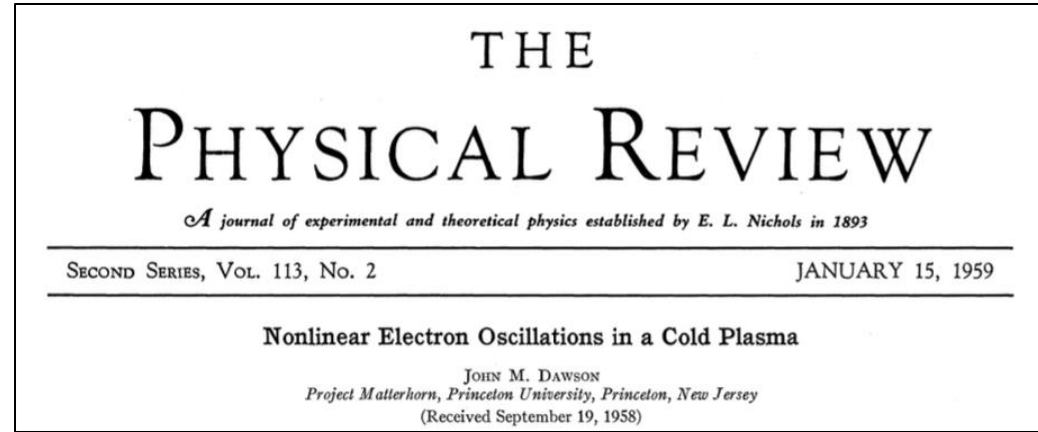
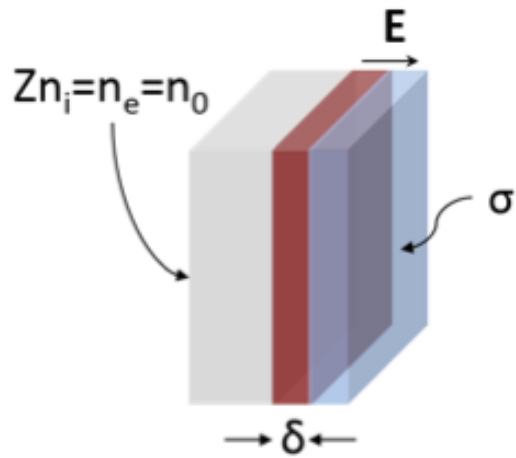
Accelerating Gradient

- Fields in plasmas are sustained by the charge separation



Accelerating Gradient

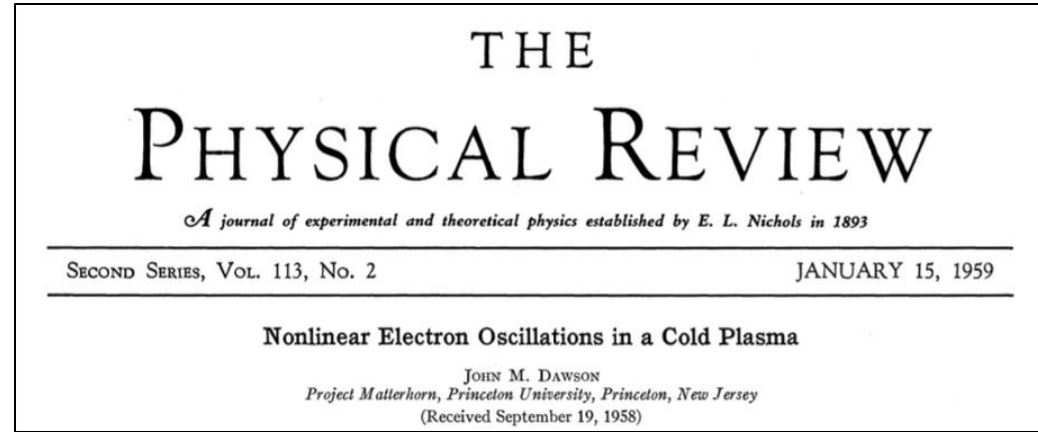
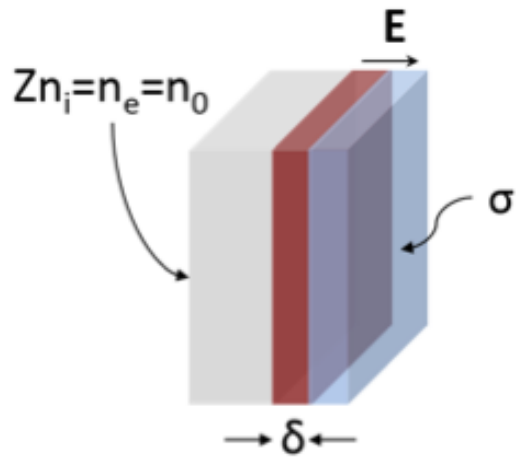
- Fields in plasmas are sustained by the charge separation
 - As high as the cold wave-breaking field: $E_{WB} = \frac{m_e c \omega_{pe}}{q} \rightarrow$ oscillation length cannot exceed plasma wavelength
 - E.g. for $n_{pe} = (10^{14} - 10^{18}) \text{ cm}^{-3}$, $E_{WB} \sim 100 \frac{V}{m} \sqrt{n_{pe} [\text{cm}^{-3}]} = (1 - 100 \text{ GV/m})$



Wave «breaks» when the maximum amplitude is reached

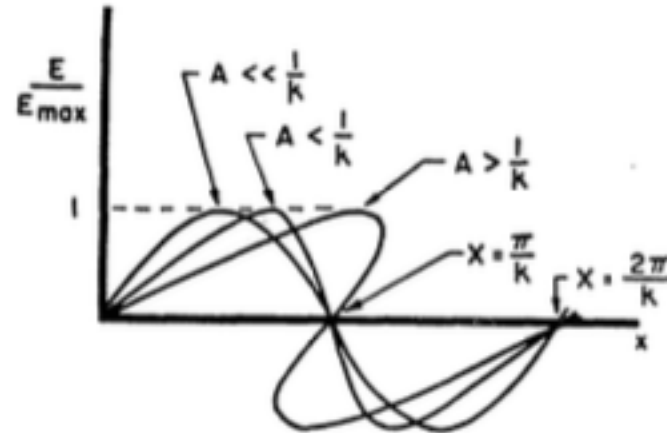
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- RF cavities limited to 100MV/m by breakdown, caused e.g. by fatigue, pulse heating, etc..

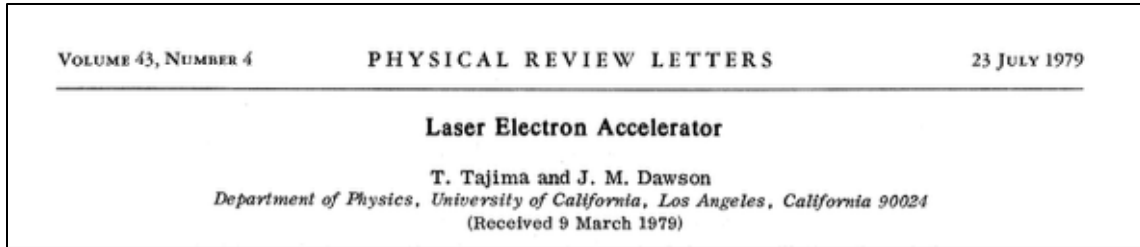
→ one could dream of shrinking down the size of accelerators by orders of magnitude!



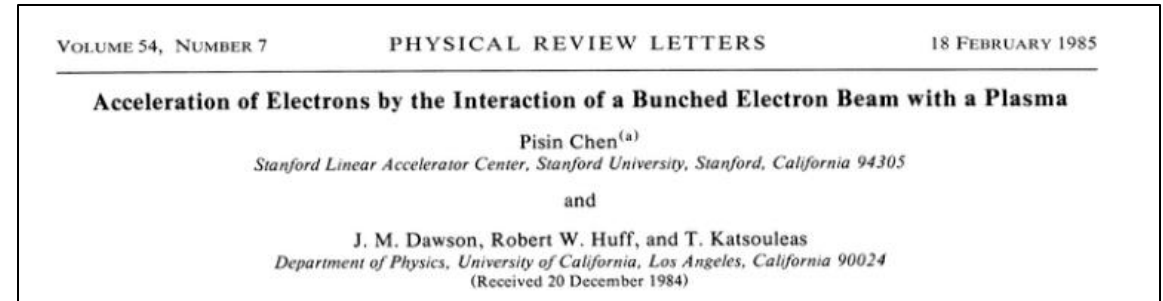
Wave «breaks» when the maximum amplitude is reached

Accelerating Gradient – Experimental Results

Laser Wakefield Acceleration (LWFA)

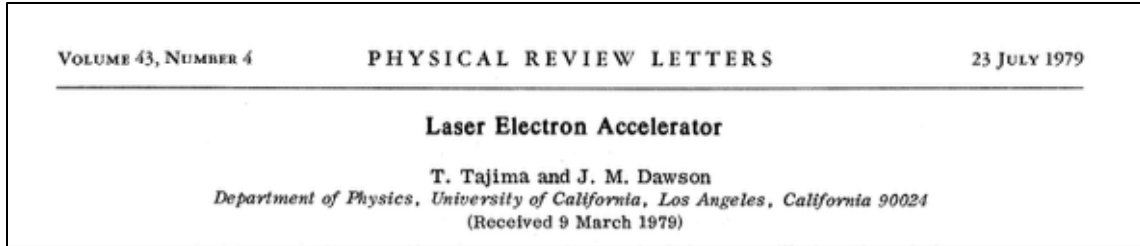


Beam-Driven Plasma Wakefield Acceleration (PWFA)

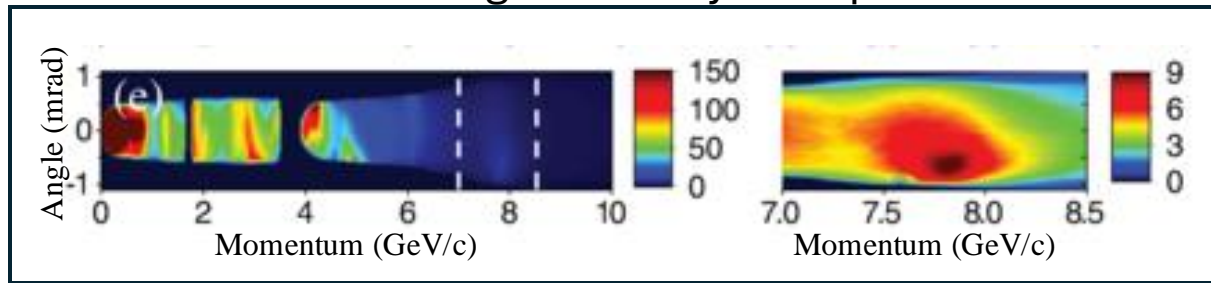


Accelerating Gradient – Experimental Results

Laser Wakefield Acceleration (LWFA)



Driver: high-intensity laser pulse

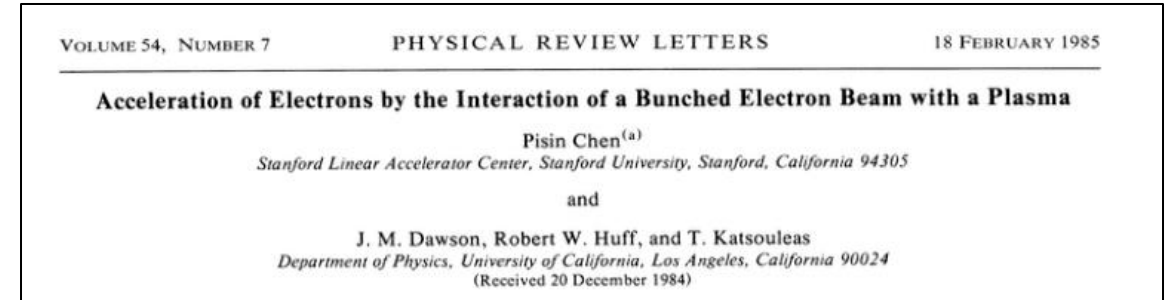


~ 8 GeV in 20 cm → ~40GV/m

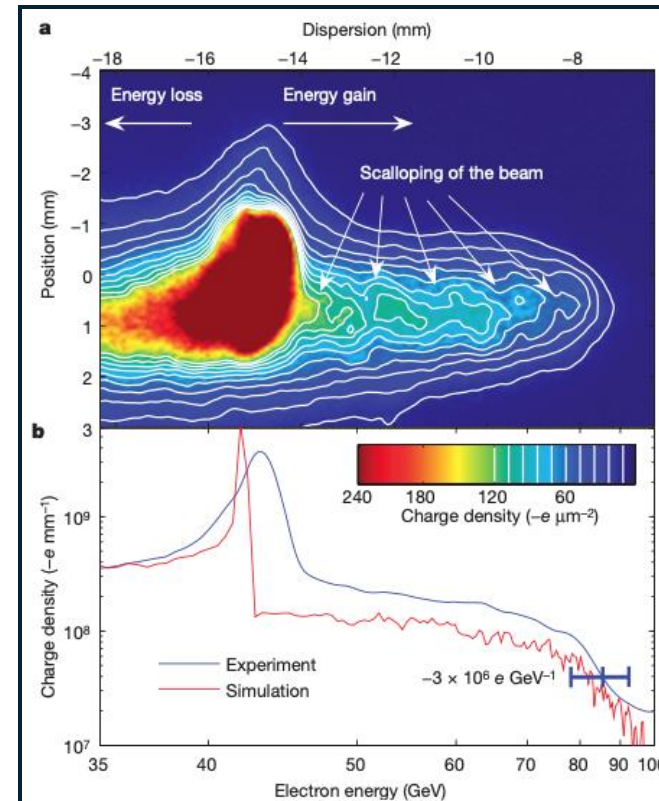
$$n_{pe} = 3.4 \times 10^{17} \text{ cm}^{-3}$$

A. J. Gonsalves et al., PRL 122, 084801 (2019)

Beam-Driven Plasma Wakefield Acceleration (PWFA)



Driver: relativistic charged particle bunch

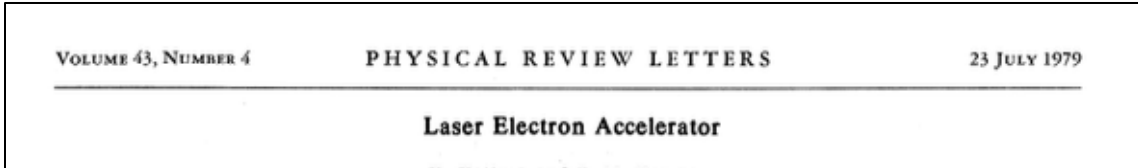


~ 42 GeV in 85 cm → ~50GV/m
 $n_{pe} = 2.8 \times 10^{17} \text{ cm}^{-3}$

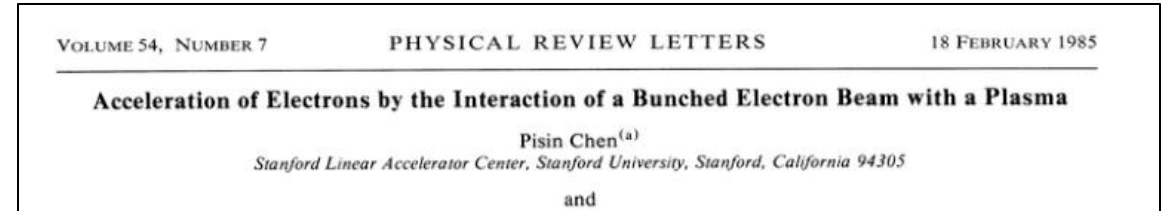
I. Blumenfeld et al.,
 Nature 455, 741-744 (2007)

Accelerating Gradient – Experimental Results

Laser Wakefield Acceleration (LWFA)



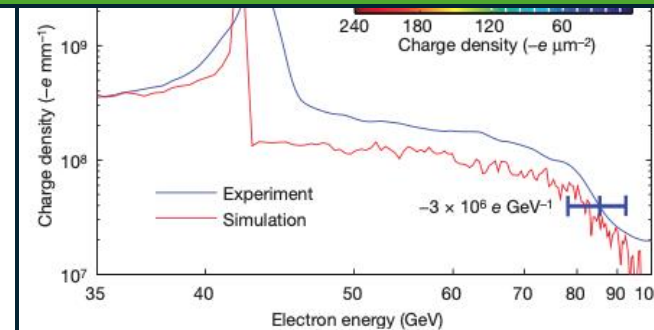
Beam-Driven Plasma Wakefield Acceleration (PWFA)



Plasma's main advantage:
Same (or higher) beam energy in considerably smaller space

- Cheaper accelerators
- Cheaper light sources, etc..
- More of them

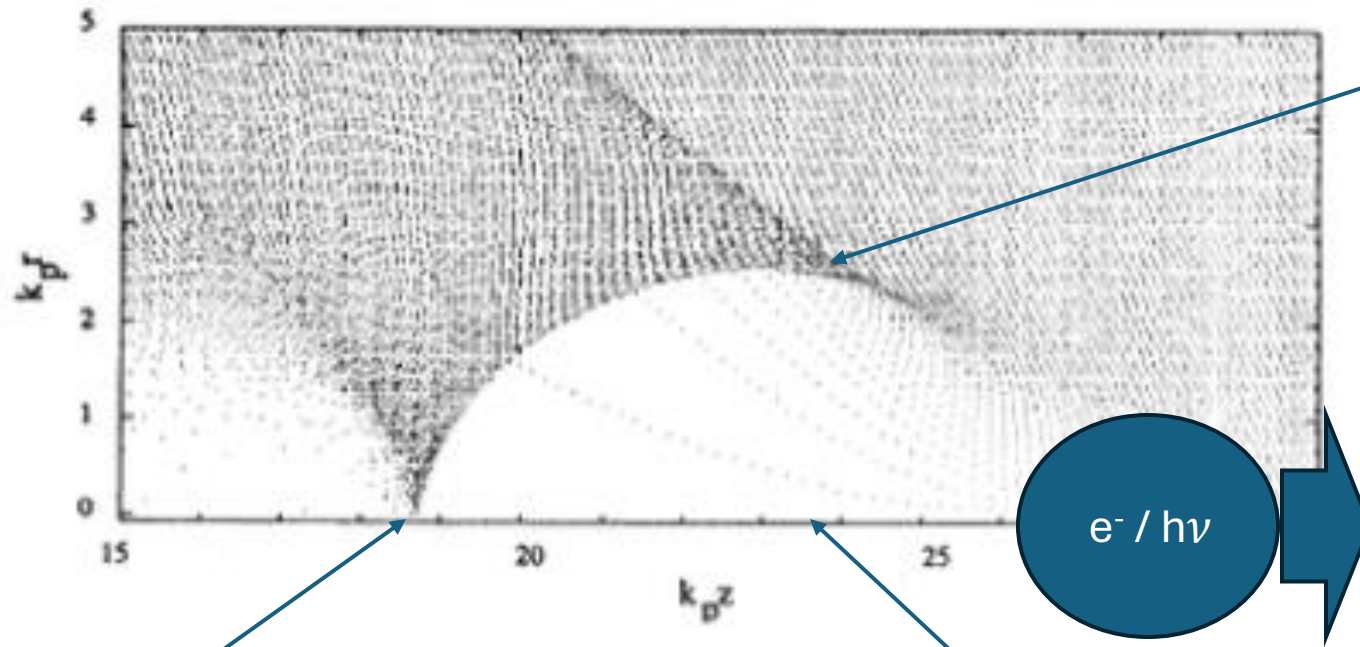
A. J. Gonsalves et al., PRL 122, 084801 (2019)



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Main challenge – Beam Quality Preservation

Most of PWFA's work in the non-linear *blowout* regime:



Plasma electrons are expelled outwards forming a “bubble”

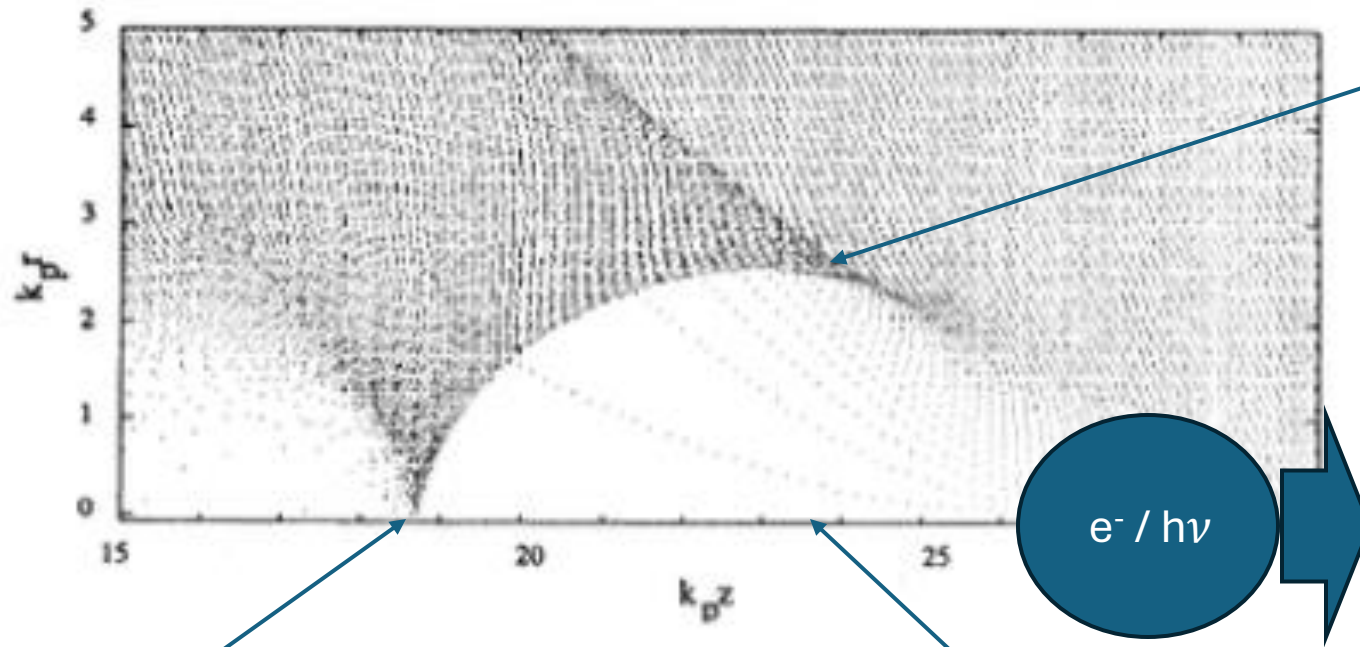
They cross the axis after $\sim \lambda_{pe}$
→ Singularity
→ Non-linearities

No electrons on axis
(blowout)

Main challenge – Beam Quality Preservation

Most of PWFA's work in the non-linear *blowout* regime:

- High gradient
- Linear focusing force



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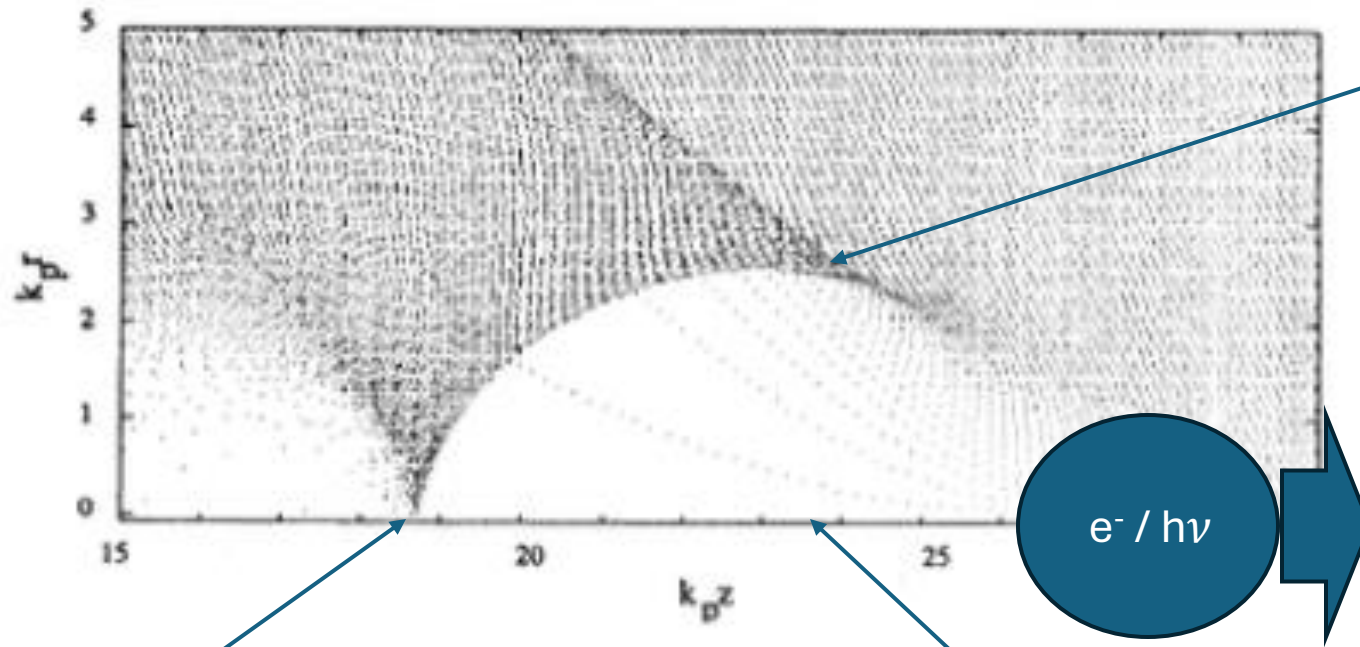
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→ Causing the main challenge: beam quality preservation (vital for applications)

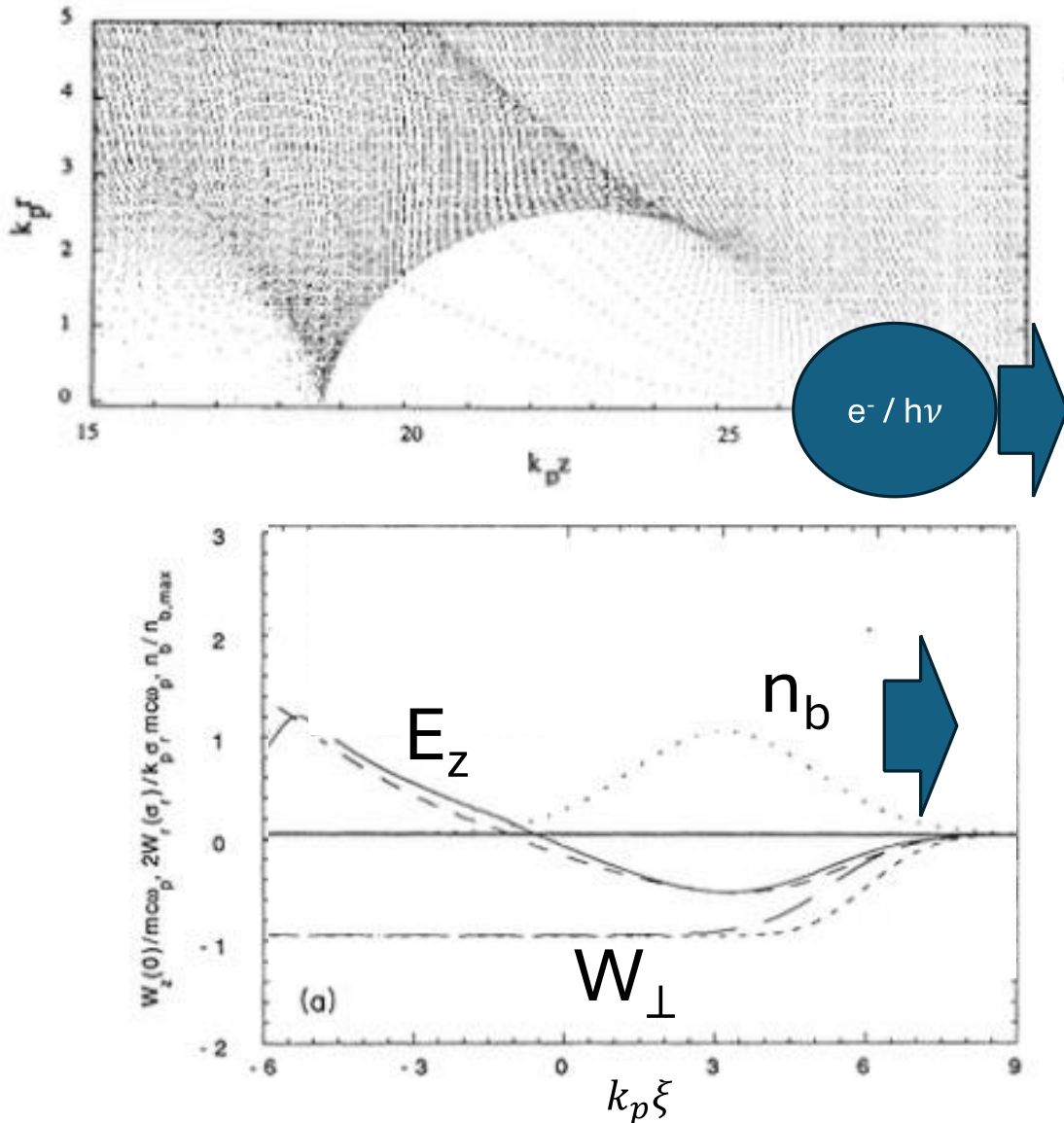


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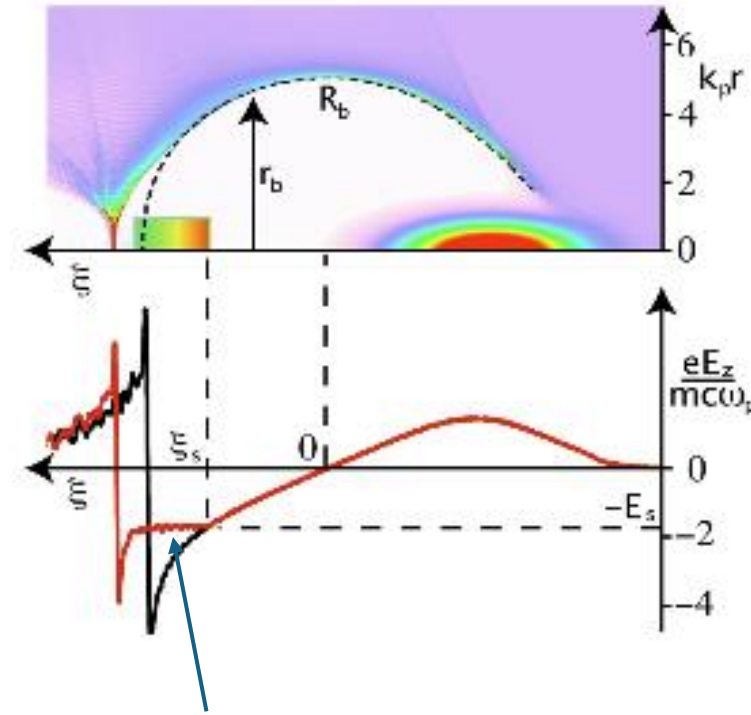
They cross the axis after $\sim \lambda_{pe}$
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No electrons on axis
(blowout)

Main challenge – Energy spread (*longitudinal quality*)



- Accelerating field not uniform along the bubble
 \rightarrow Energy spread increasing while acceleration happens
- Solution: “loading” the wake with the presence of the witness bunch itself



Constant accelerating field within the witness bunch

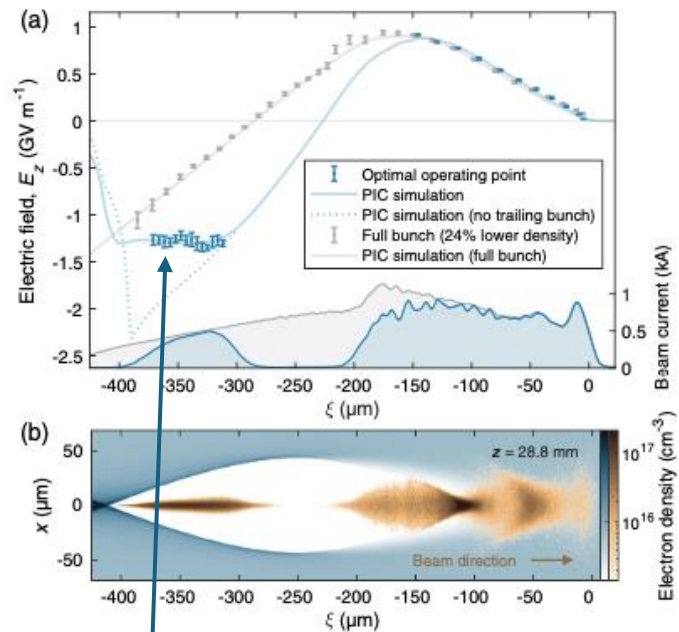
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➔ Experimental Demonstrations:

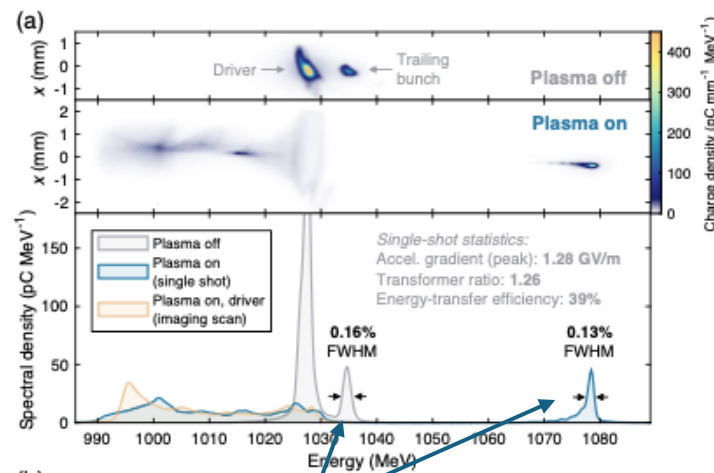
PHYSICAL REVIEW LETTERS **126**, 014801 (2021)

Energy-Spread Preservation and High Efficiency in a Plasma-Wakefield Accelerator

C. A. Lindström^{1,*}, J. M. Garland¹, S. Schröder^{1,2}, L. Boulton^{1,3,4}, G. Boyle¹, J. Chappell⁵, R. D'Arcy¹, P. Gonzalez^{1,2}, A. Knetsch^{1,9}, V. Libov¹, G. Loisch¹, A. Martinez de la Ossa¹, P. Niknejadi¹, K. Pöder¹, L. Schaper¹, B. Schmidt¹, B. Sheeran^{1,2}, S. Wesch¹, J. Wood¹ and J. Osterhoff¹



Flattened, loaded wake



~GV/m acceleration
Preserved relative energy spread

nature physics

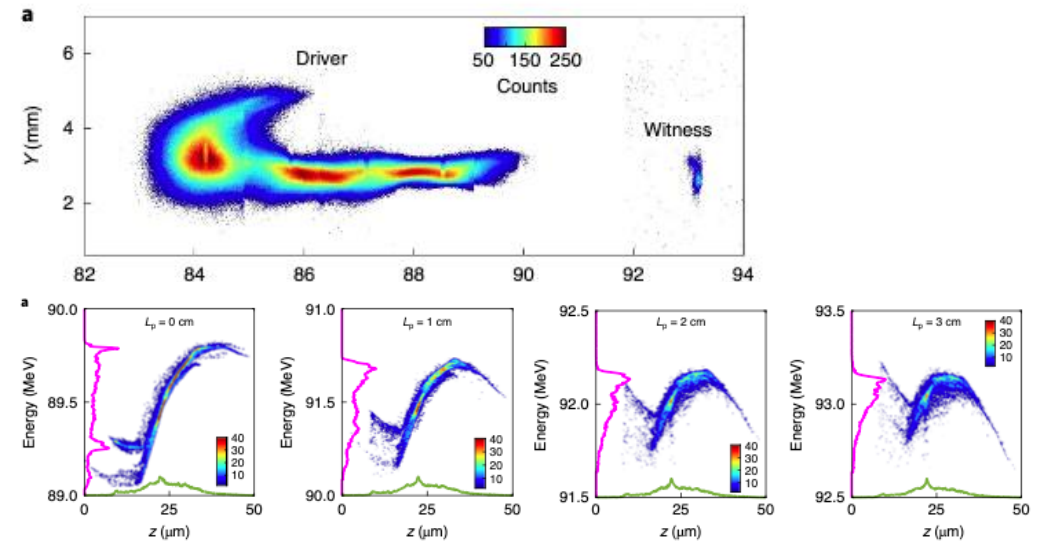
LETTERS

<https://doi.org/10.1038/s41567-020-01116-9>

Check for updates

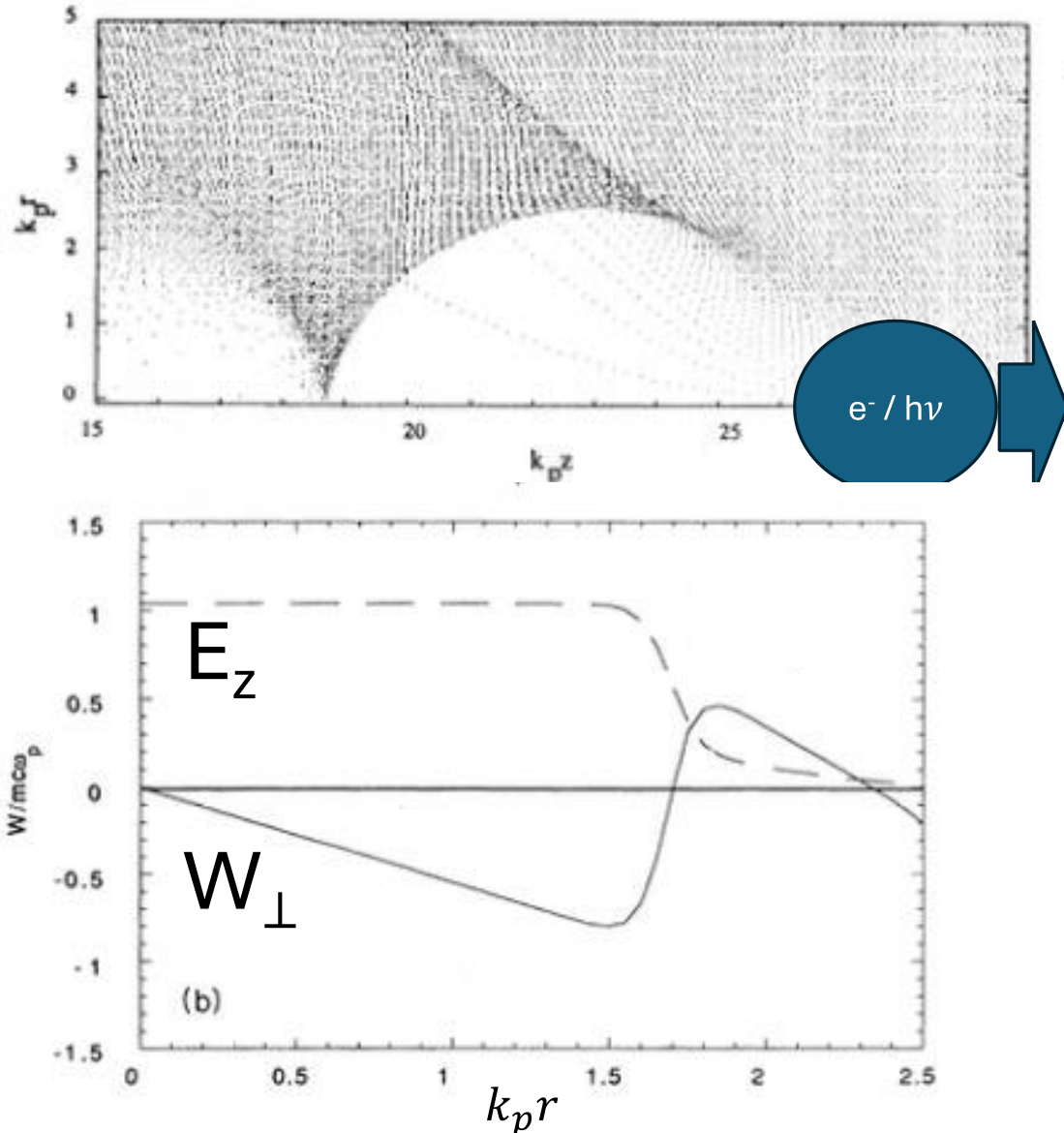
Energy spread minimization in a beam-driven plasma wakefield accelerator

R. Pompili¹, D. Alesini¹, M. P. Anania¹, M. Behtouei¹, M. Bellaveglia¹, A. Biagioni¹, F. G. Bisesto¹, M. Cesarini^{1,2}, E. Chiadroni¹, A. Cianchi³, G. Costa¹, M. Croia¹, A. Del Dotto¹, D. Di Giovenale¹, M. Diomedè¹, F. Dipace¹, M. Ferrario¹, A. Giribono¹, V. Lollo¹, L. Magnisi¹, M. Marongiu¹, A. Mostacci², L. Piersanti¹, G. Di Pirro¹, S. Romeo¹, A. R. Rossi⁴, J. Scifo¹, V. Shpakov¹, C. Vaccarezza¹, F. Villa¹ and A. Zigler^{1,5}



Combination of beam loading and initial chirp to obtain final small energy spread

Main challenge – Emittance Preservation (*transverse quality*)



- Focusing force is transversely linear

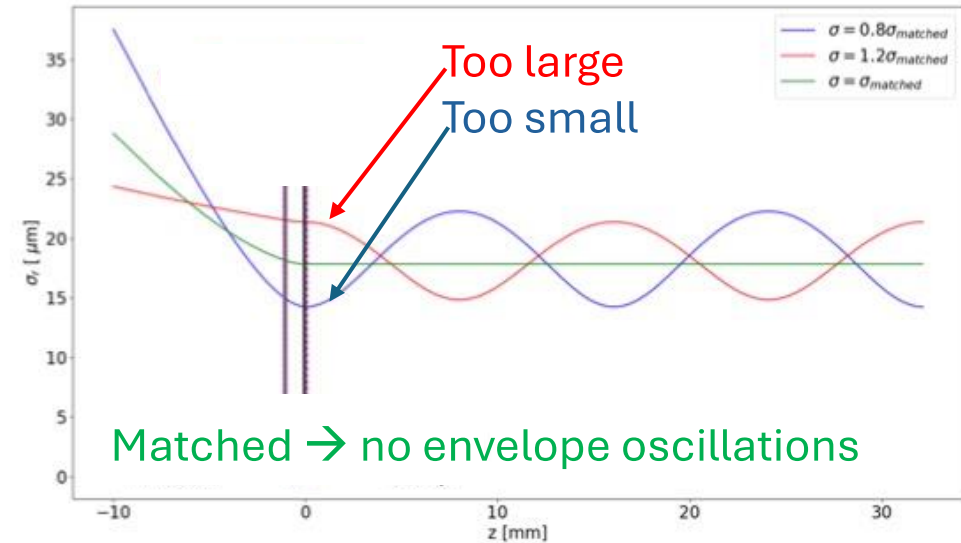
if: beam envelope is *matched* to the focusing force
(which is extremely strong!)

→ Possible emittance preservation

else: different energy slices rotate at different rates in transverse phase space

→ slice emittance preserved (linear focusing)

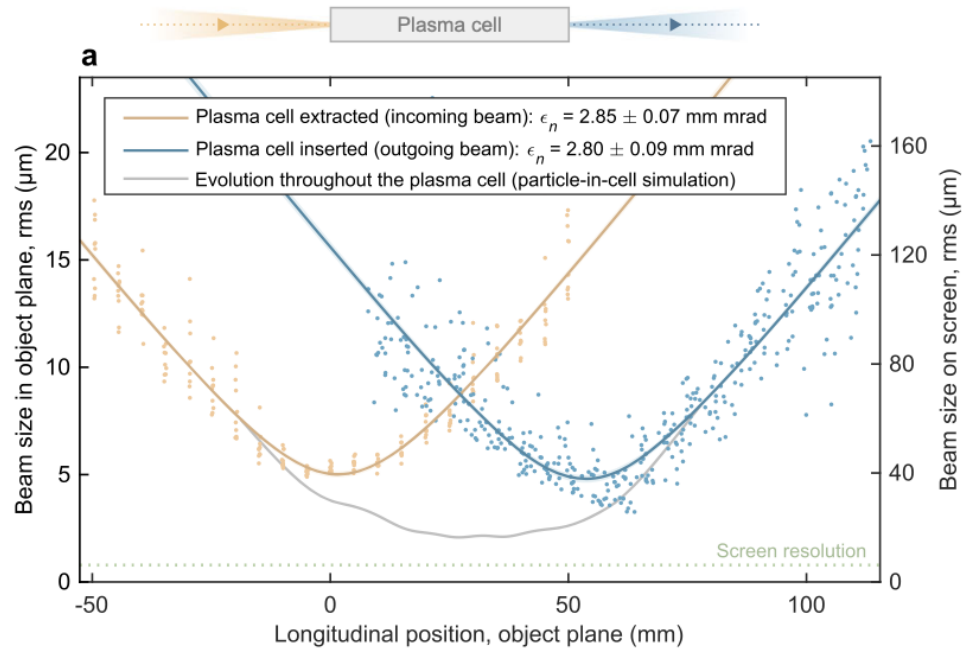
→ projected (i.e., overall) emittance grows!



(L. Verra *et al* 2020 *J. Phys.: Conf. Ser.* **1596** 012007)

Main challenge – Emittance Preservation (*transverse quality*)

→ Experimental Demonstration:



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nature communications

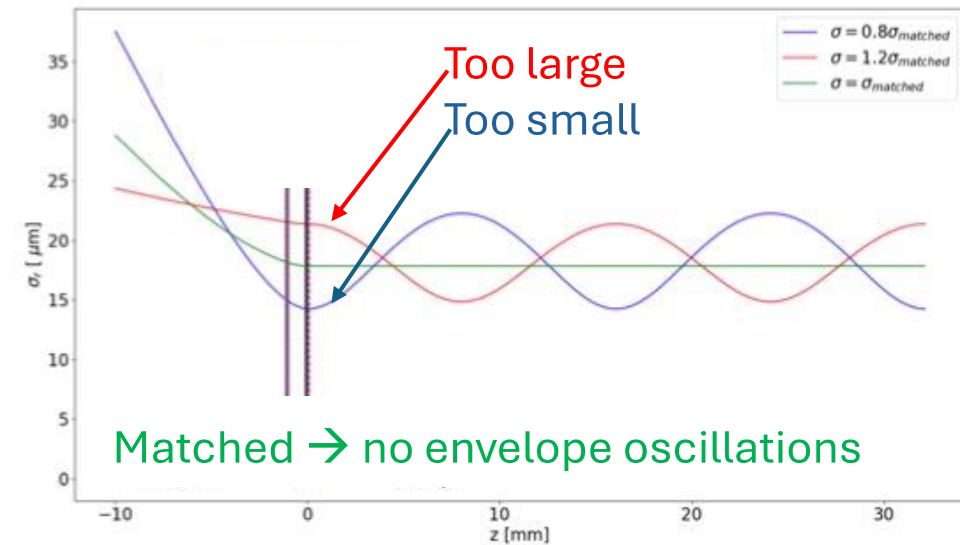
Article <https://doi.org/10.1038/s41467-024-50320-1>

Emittance preservation in a plasma-wakefield accelerator

Received: 17 July 2023 | Accepted: 4 July 2024 | Published online: 19 July 2024

C. A. Lindström^{1,2}, J. Beinortaitė^{1,3}, J. Björklund Svensson¹, L. Boulton^{1,4,5}, J. Chappell³, S. Diederichs^{1,6}, B. Foster⁷, J. M. Garland¹, P. González Caminal^{1,6}, G. Loisch¹, F. Peña^{1,6}, S. Schröder¹, M. Thévenet¹, S. Wesch¹, M. Wing^{1,3}, J. C. Wood^{1,3}, R. D’Arcy¹ & J. Osterhoff¹

Check for updates



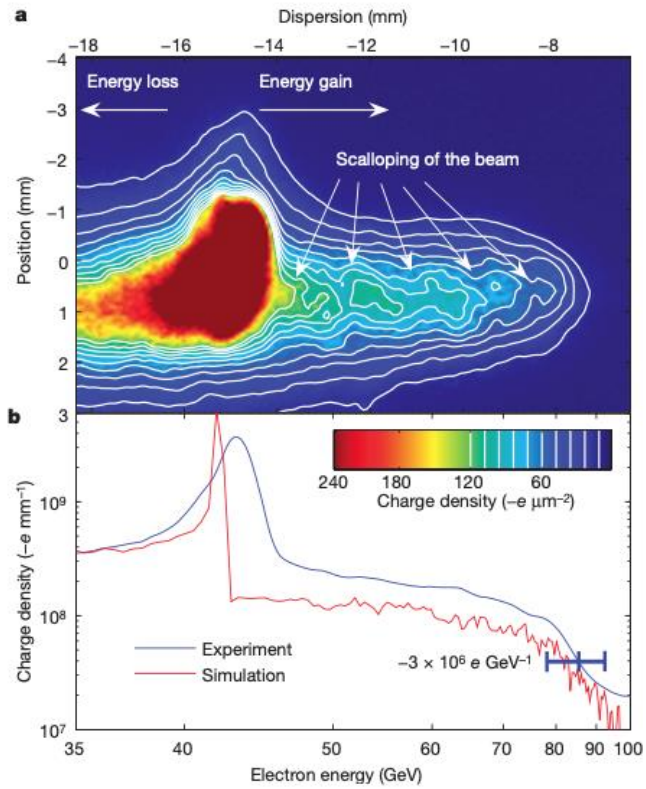
(L. Verra *et al* 2020 *J. Phys.: Conf. Ser.* **1596** 012007)

Main Challenges

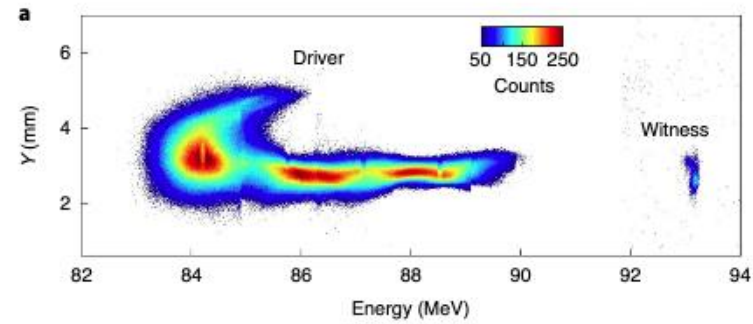
The main challenge remains:

- Do everything at the same time

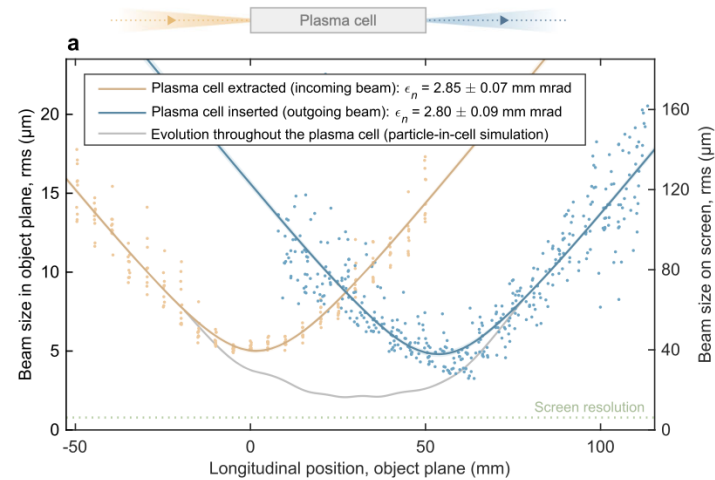
Extremely high gradient



Energy spread minimization



Emittance preservation

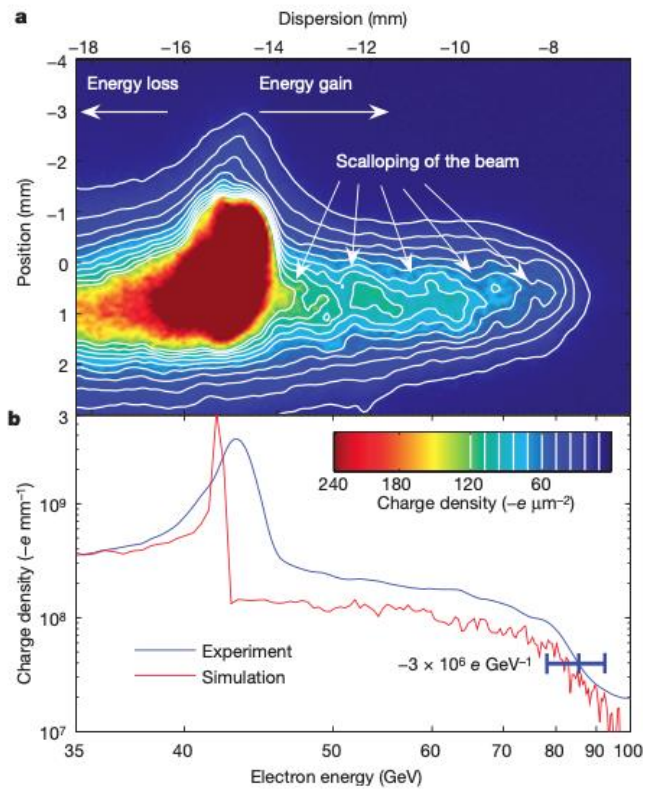


Main Challenges

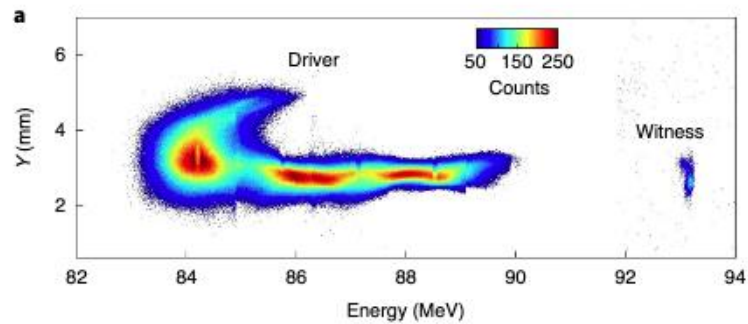
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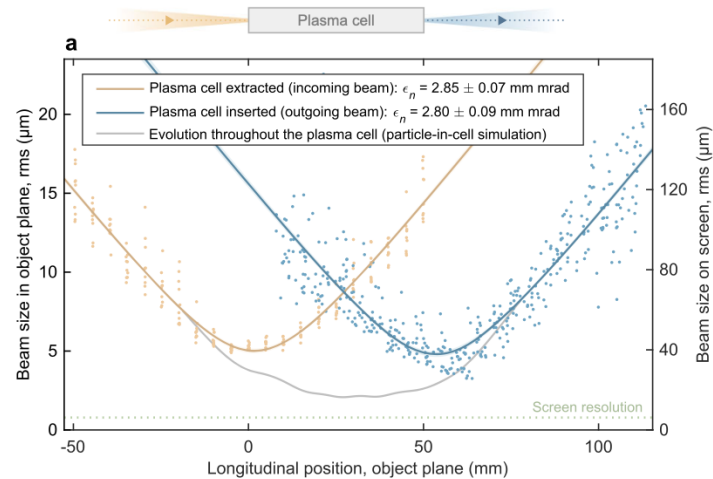
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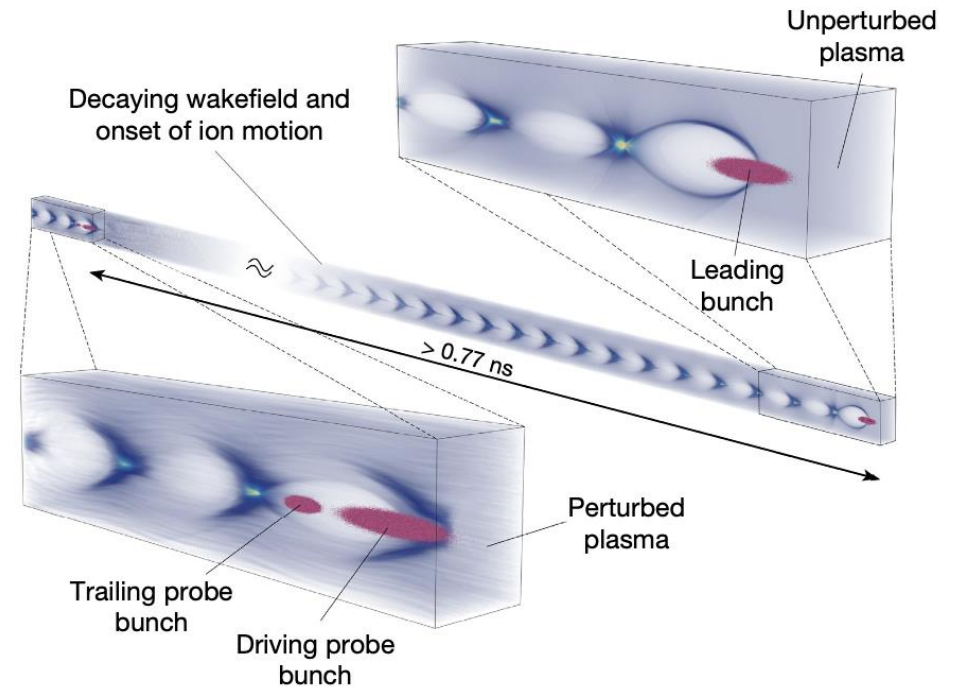
Energy spread minimization



Emittance preservation



- Do it many times \rightarrow high repetition rate



Plasma was shown to recover at ns time-scale

R. D'Arcy et al., Nature **603**, 58–62 (2022)

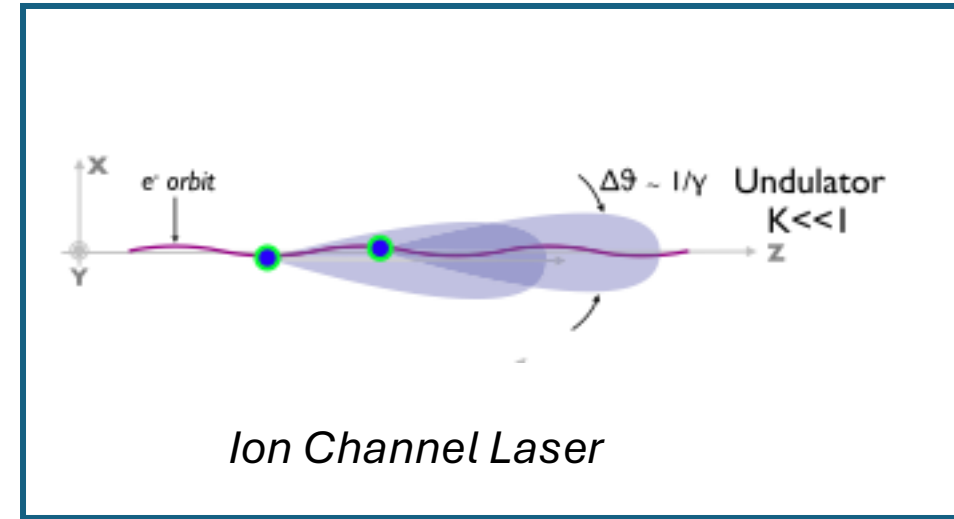
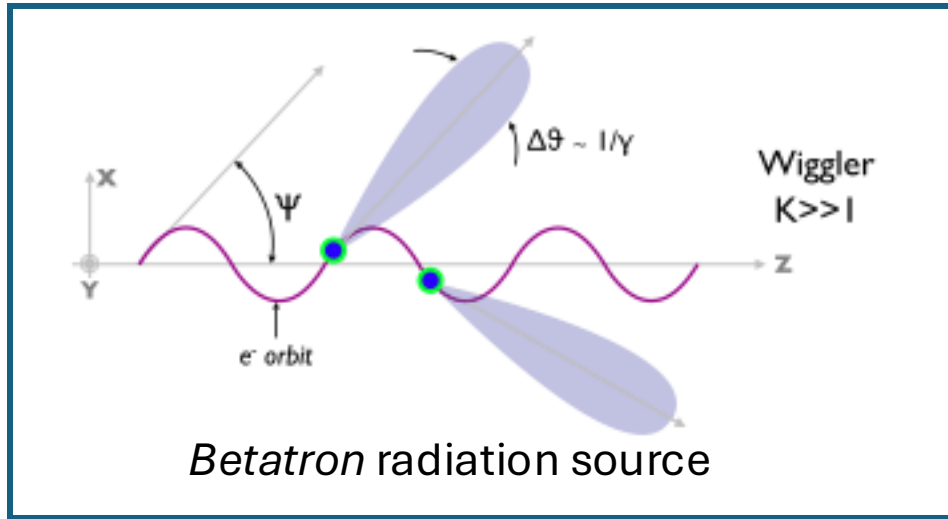
R. Pompili et al., Comm. Phys. **7**, 241 (2024)

Still a technological challenge:

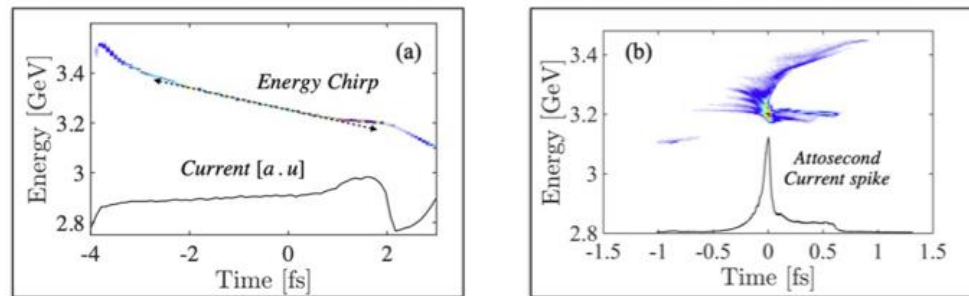
(Angelo Biagioni's talk)

Challenges → Opportunities

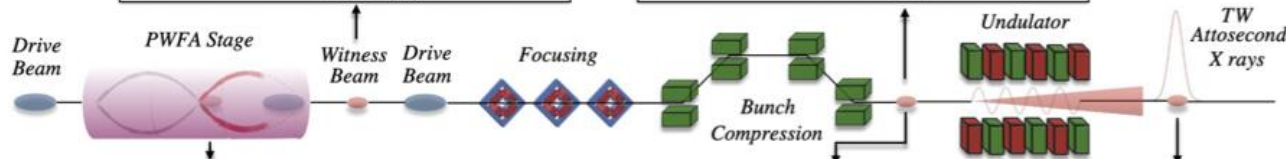
- Transverse (betatron) oscillation induces radiation emission in plasma (A. Frazzitta's talk)



- Large energy chirp may be used to compress the witness bunch



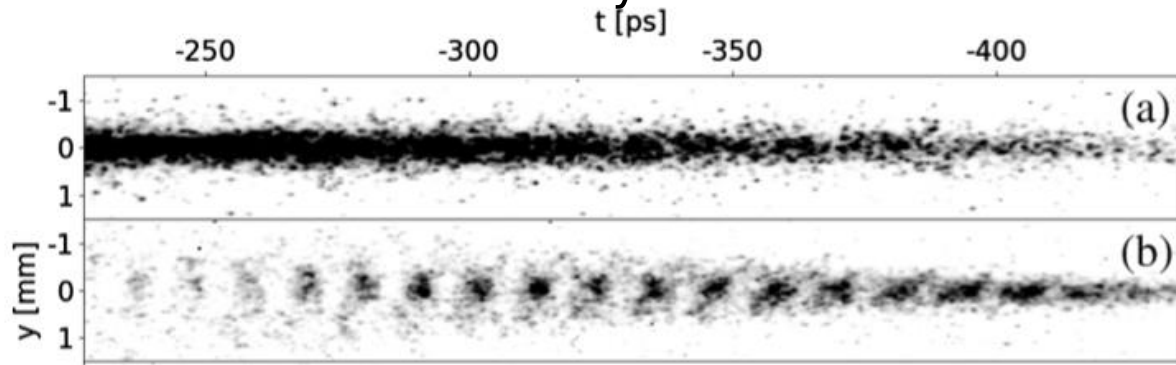
→ Extremely short bunches and light pulses (<fs)



Challenges → Opportunities

- Study of Beam-Plasma interactions and instabilities is an active field (e.g., laboratory astrophysics)

Self-Modulation Instability

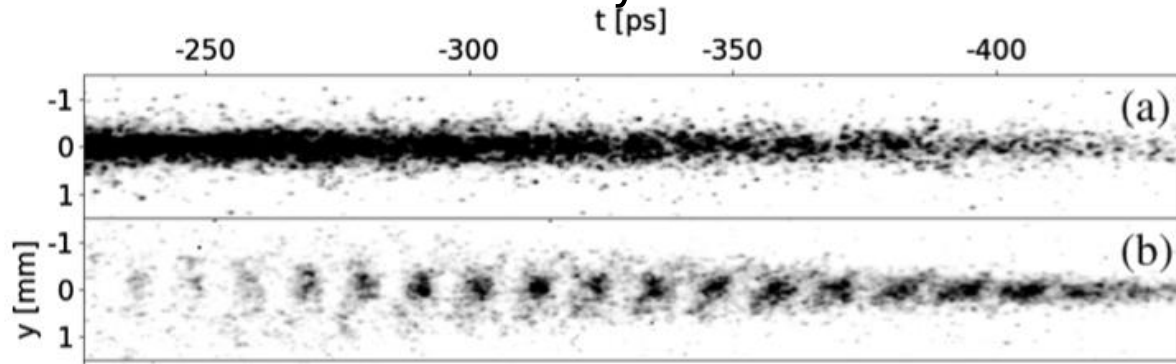


L. Verra et al. (AWAKE Collaboration) Phys. Rev. Lett. **129**, 024802 (2022)

Challenges → Opportunities

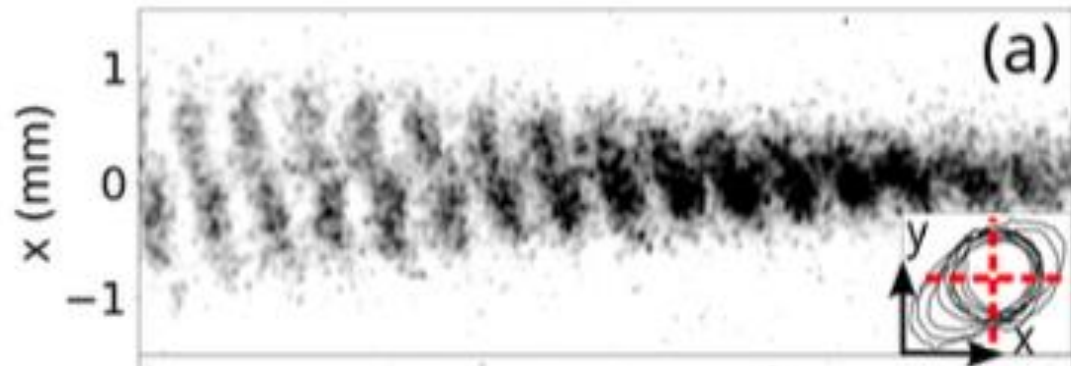
- Study of Beam-Plasma interactions and instabilities is an active field (e.g., laboratory astrophysics)

Self-Modulation Instability



L. Verra et al. (AWAKE Collaboration) Phys. Rev. Lett. **129**, 024802 (2022)

Hosing Instability

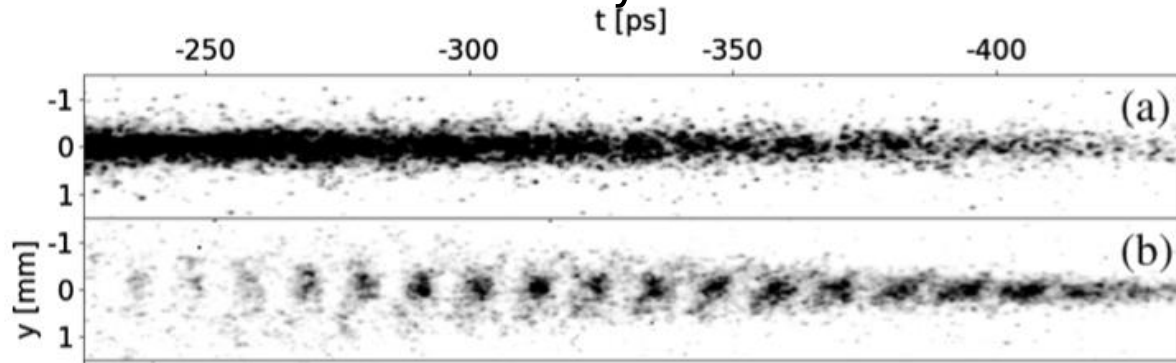


T. Nechaeva, L. Verra et al. (AWAKE Coll.) Phys. Rev. Lett. **132**, 075001 (2024)

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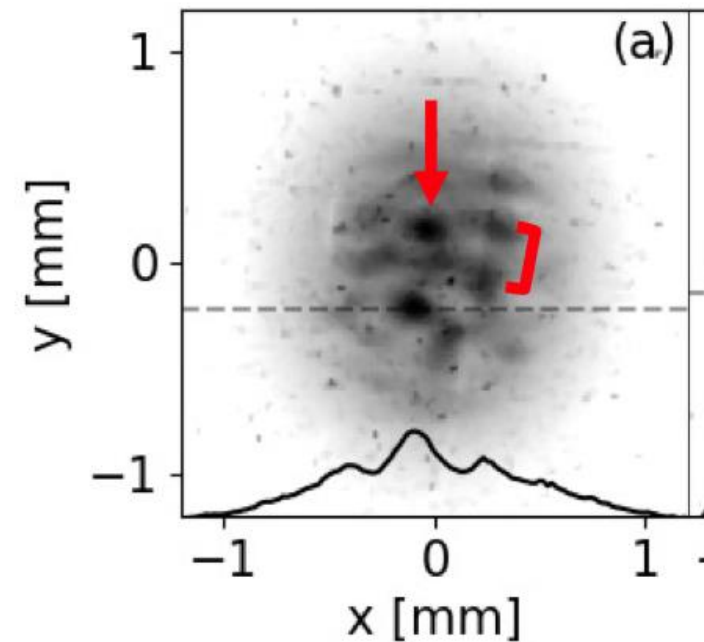
L. Verra et al. (AWAKE Collaboration) Phys. Rev. Lett. **129**, 024802 (2022)

Hosing Instability



T. Nechaeva, L. Verra et al. (AWAKE Coll.) Phys. Rev. Lett. **132**, 075001 (2024)

Filamentation Instability



L. Verra et al. (AWAKE Collaboration), Phys. Rev. E **109**, 055203

- We will exploit the advantages (large energy gain), while handling the challenges to deliver high-quality bunches
 - Compact Free Electron Laser
- Opportunities for “fun physics”, where the challenges are used
 - Radiation generation
 - Compression to ultra-short bunches
 - Beam-plasma instabilities
 -

Thank You For Listening!