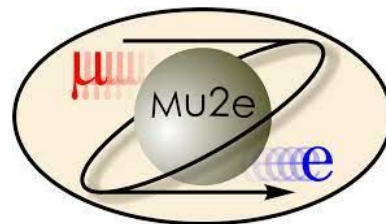




**Pulling out all the Stops:
Measuring the normalisation of signal events in the
Mu2e experiment**

George Sweetmore

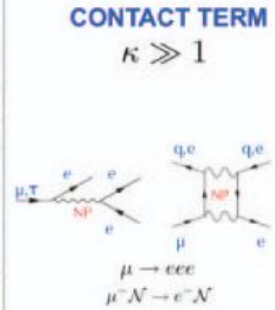
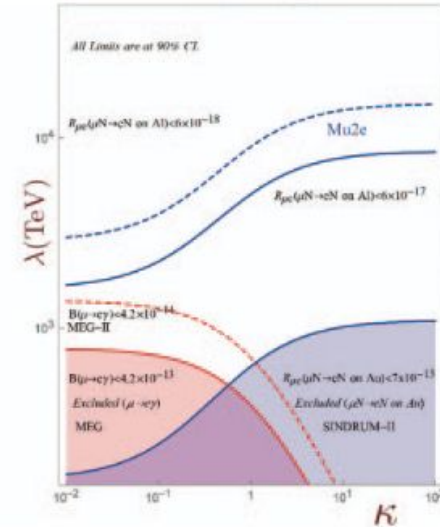
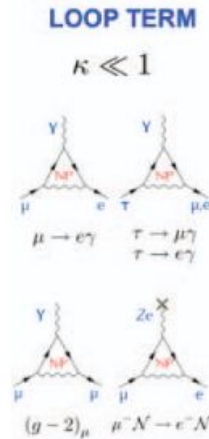
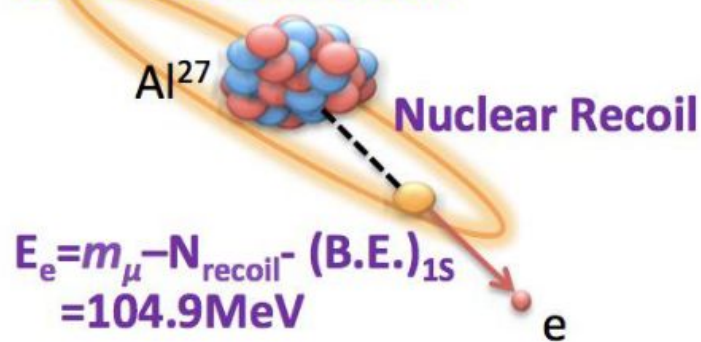
Muon4Future 2025



Mu2e experiment



Coherent Conversion



Measure neutrinoless conversion from muon to electron in proximity of a nucleus

Highly suppressed in SM
Allows us to probe large range of BSM models

Target Sensitivity:

$$R_{\mu e} = \frac{\Gamma[\mu^- + A(Z, N) \rightarrow e^- + A(Z, N)]}{\Gamma[\mu^- + A(Z, N) \rightarrow \nu_\mu + A(Z-1, N+1)]} < 6.7 \times 10^{-17} (90\% \text{ CL})$$

4 orders of magnitude better than current limits: SINDRUM II [W. Bertl et al., Eur. Phys. J. C 47, 337-346 (2006)]

Measure ratio between conversion events and stopped muons

Stopping Target Monitor



Target Sensitivity:

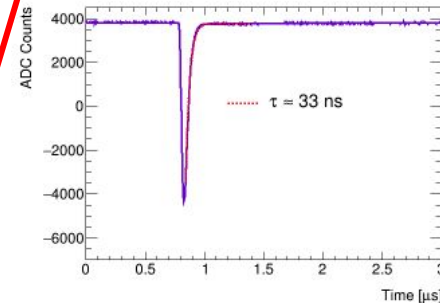
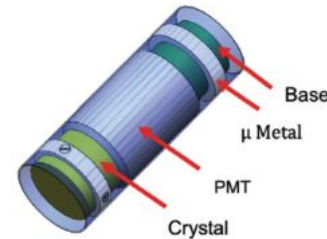
$$R_{\mu e} = \frac{\Gamma[\mu^- + A(Z, N) \rightarrow e^- + A(Z, N)]}{\Gamma[\mu^- + A(Z, N) \rightarrow \nu_\mu + A(Z-1, N+1)]} < 6.7 \times 10^{-17} \text{ (90\%CL)}$$

4 orders of magnitude better than current limits: SINDRUM II
[W. Bertl et al., Eur. Phys. J. C 47, 337-346 (2006)]

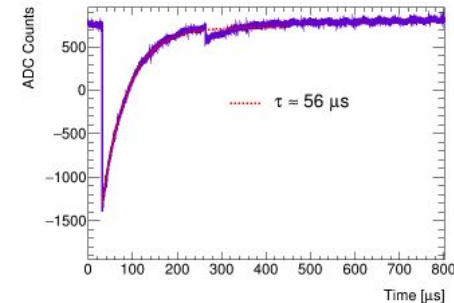
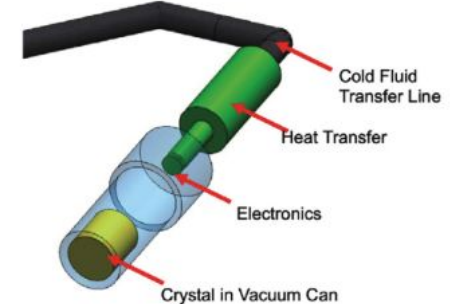
3d → 2p 62% of stops emit 66 keV	2p → 1s 62% of stops emit 347 keV	24.2% of stops DIO	37.8% of stops OMC	Emit 844 keV, 1809 keV X-rays or nothing
2p → 1s 18% of stops emit 347 keV	7% of stops DIO	11% of stops OMC	Emit 844 keV, 1809 keV X-rays or nothing	
stopped in: 1s 20% of stops	7.8% of stops DIO	12.2% of stops OMC	Emit 844 keV, 1809 keV X-rays or nothing	

Lower rates (<100kHz)
High resolution (~1keV for 1809keV)

LaBr3

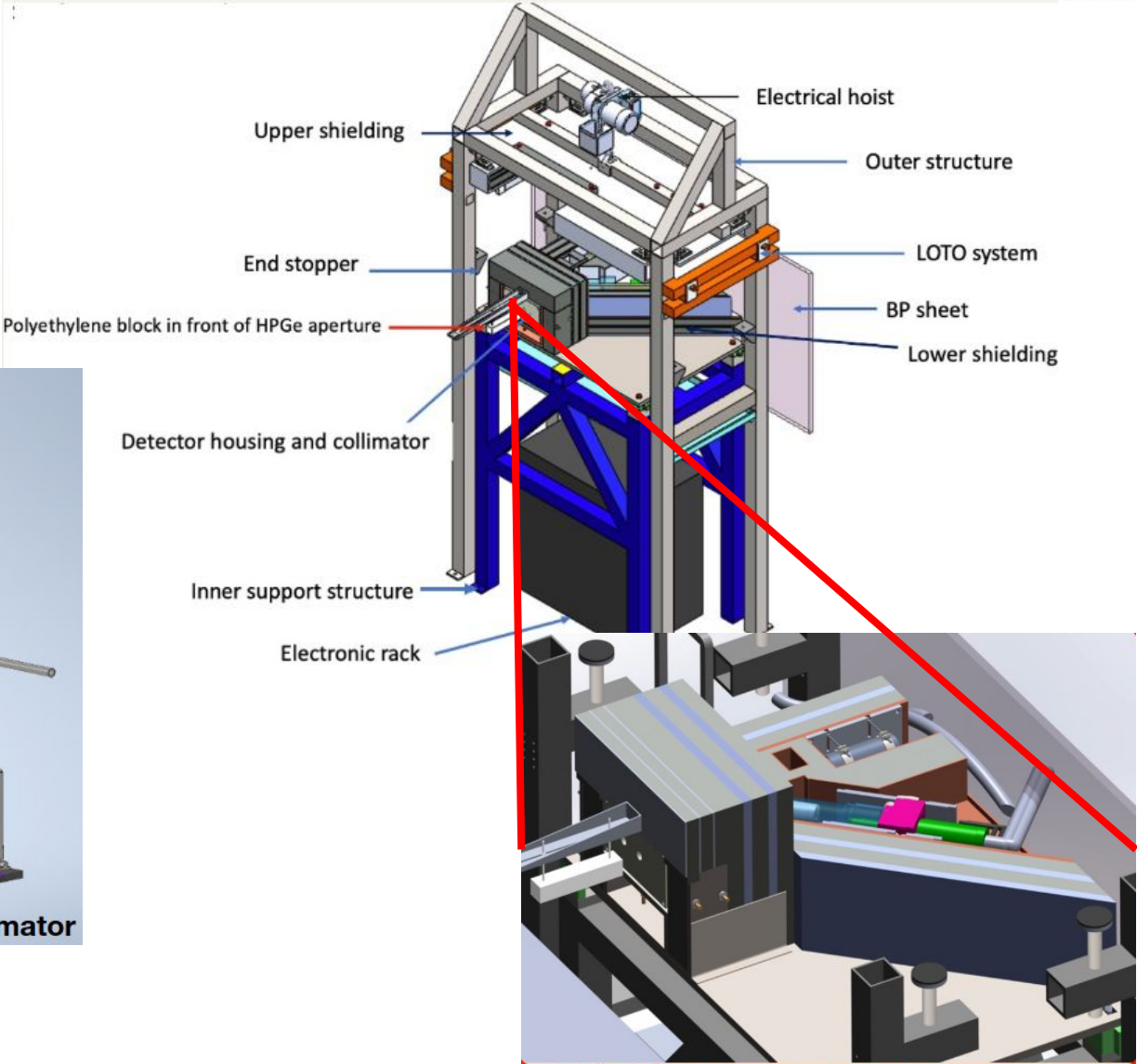
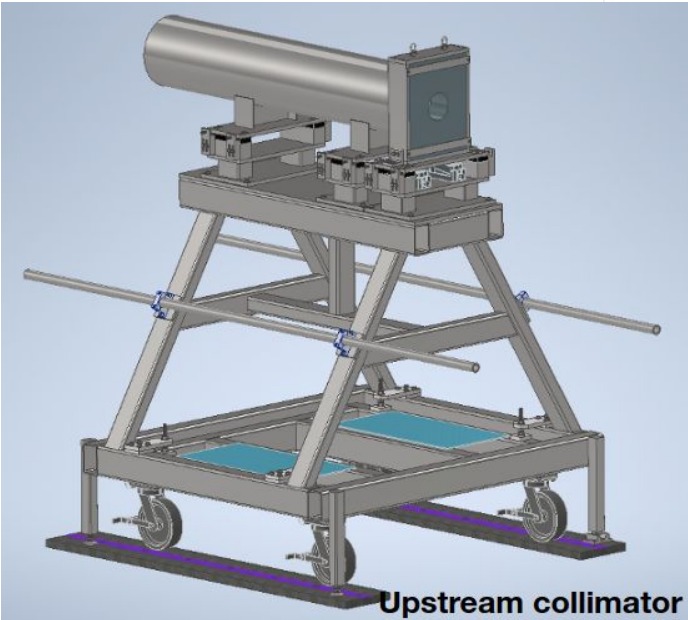
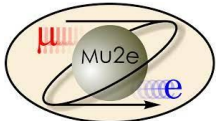


HPGe



High rates (>100kHz)
Low resolution (~12keV for 1809keV)

STM Infrastructure



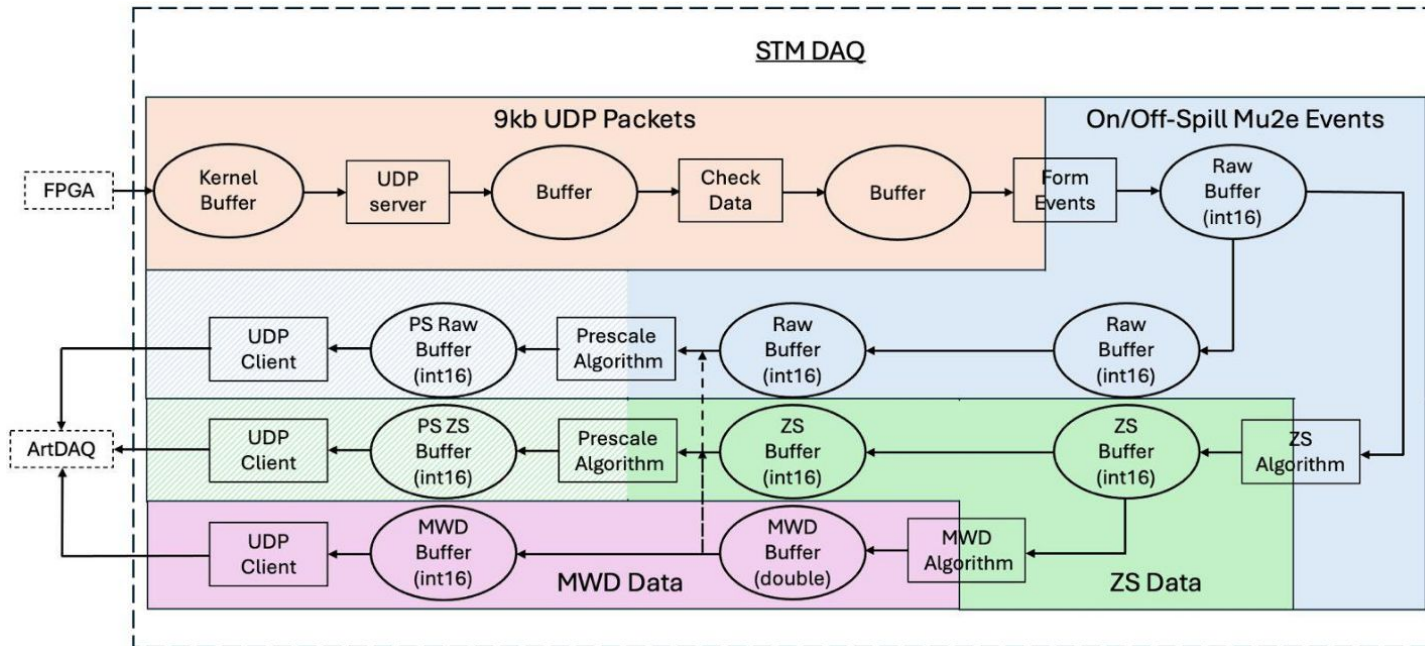
STM Data Acquisition (DAQ)

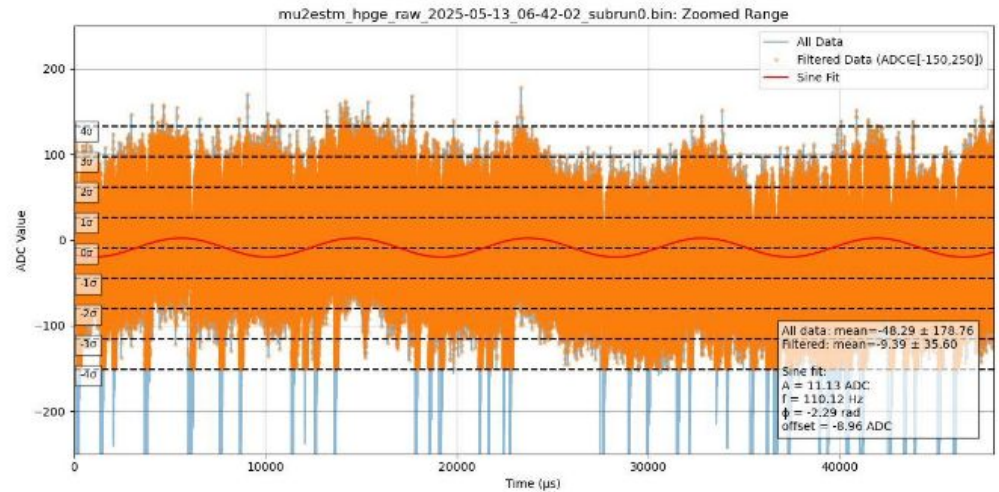
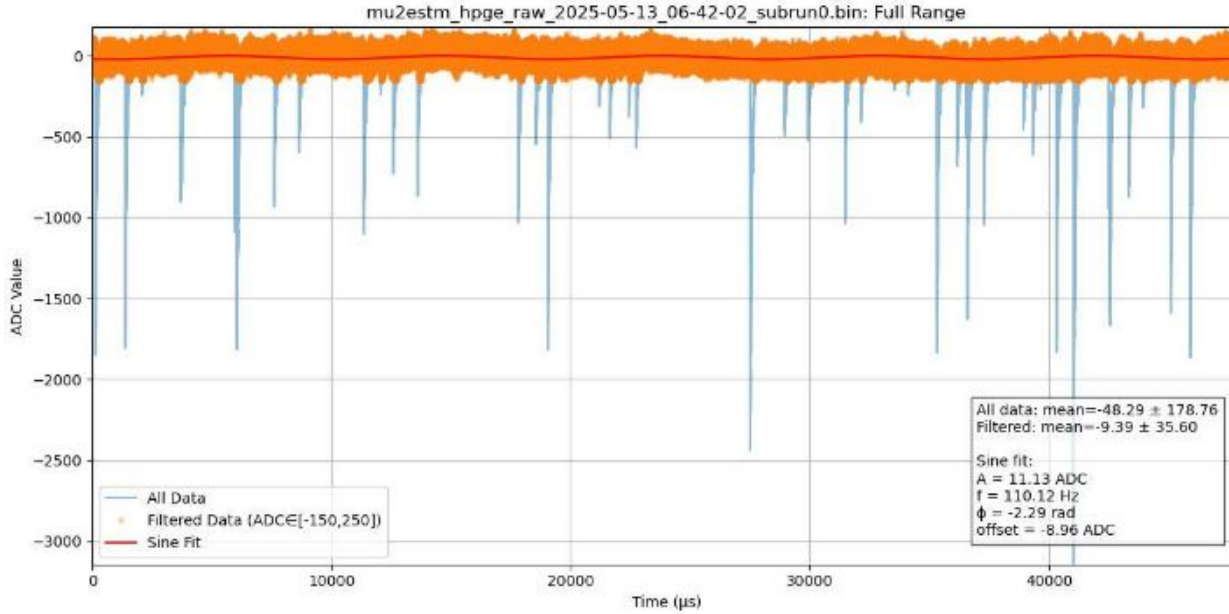


Detectors

FPGA

Server





Outlook: Status & Plans for first cosmic ray run



- STM Firmware is in a great state and ready to be implemented in coming weeks
- The STM DAQ is in a great state
- We can now run for O(10hrs)
- Online zero suppression is complete for LaBr3 and HPGe
- Timing studies performed in real time
- Have online DQM in standalone and OTSDAQ packages

TO DO:

- We need to push the data rates! where do we fall over?
- Ensure we understand cross sub-system timing
- Source testing of the detectors. What is our current achievable resolution for known energy peaks