

Updates on CMB polarization cosmology

CSN2 Meeting - Catania
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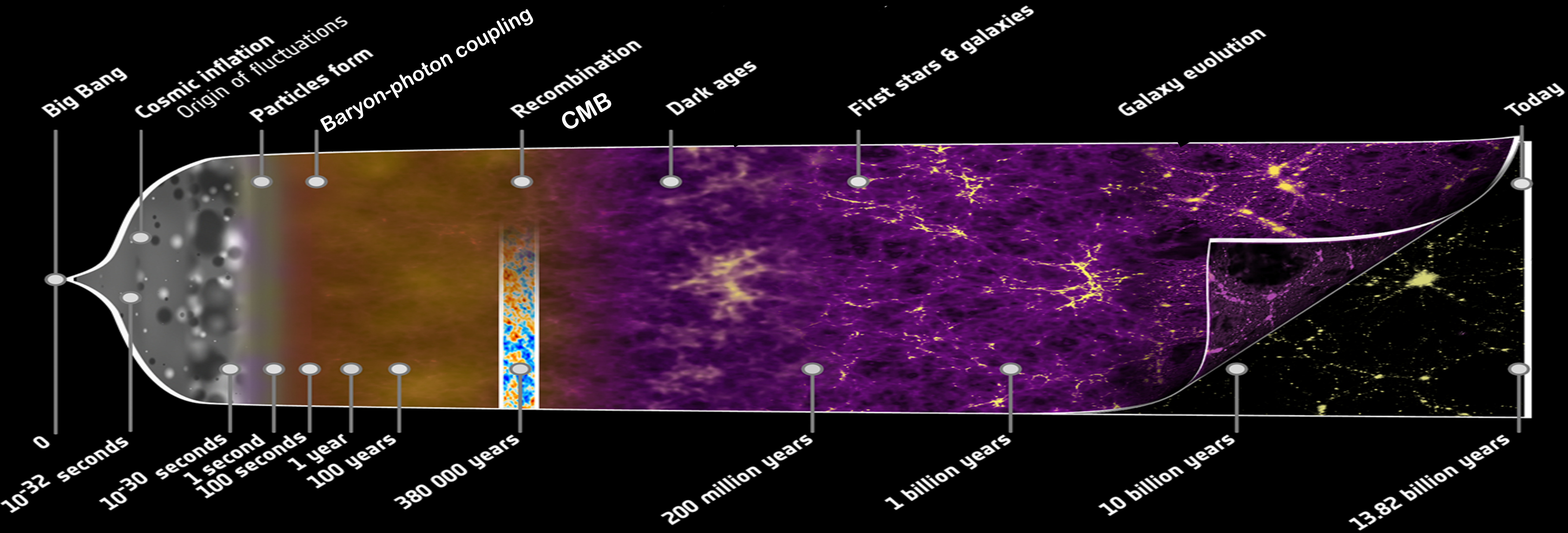


Università
di Catania



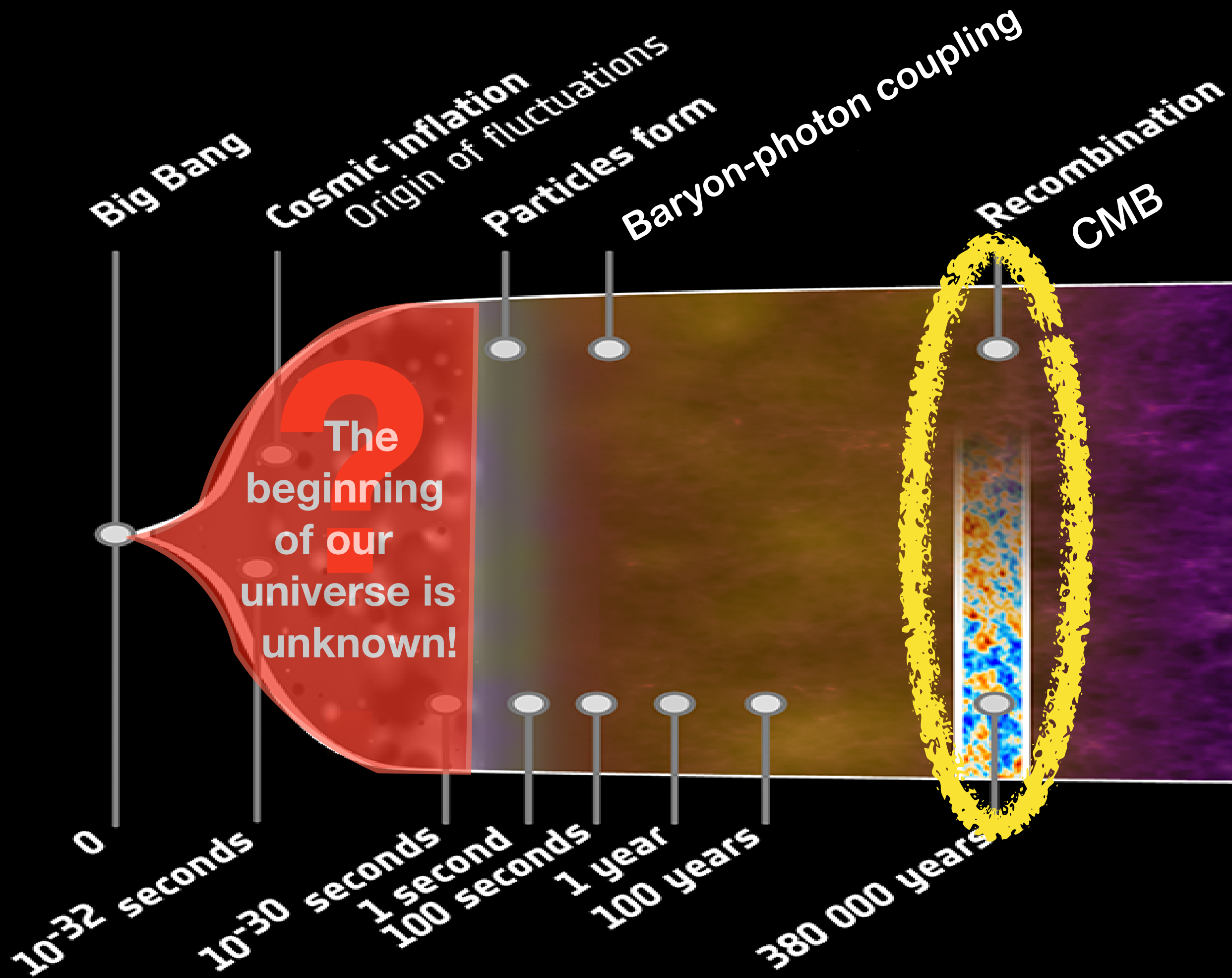
Istituto Nazionale di Fisica Nucleare

The Standard model of Cosmology: Λ CDM



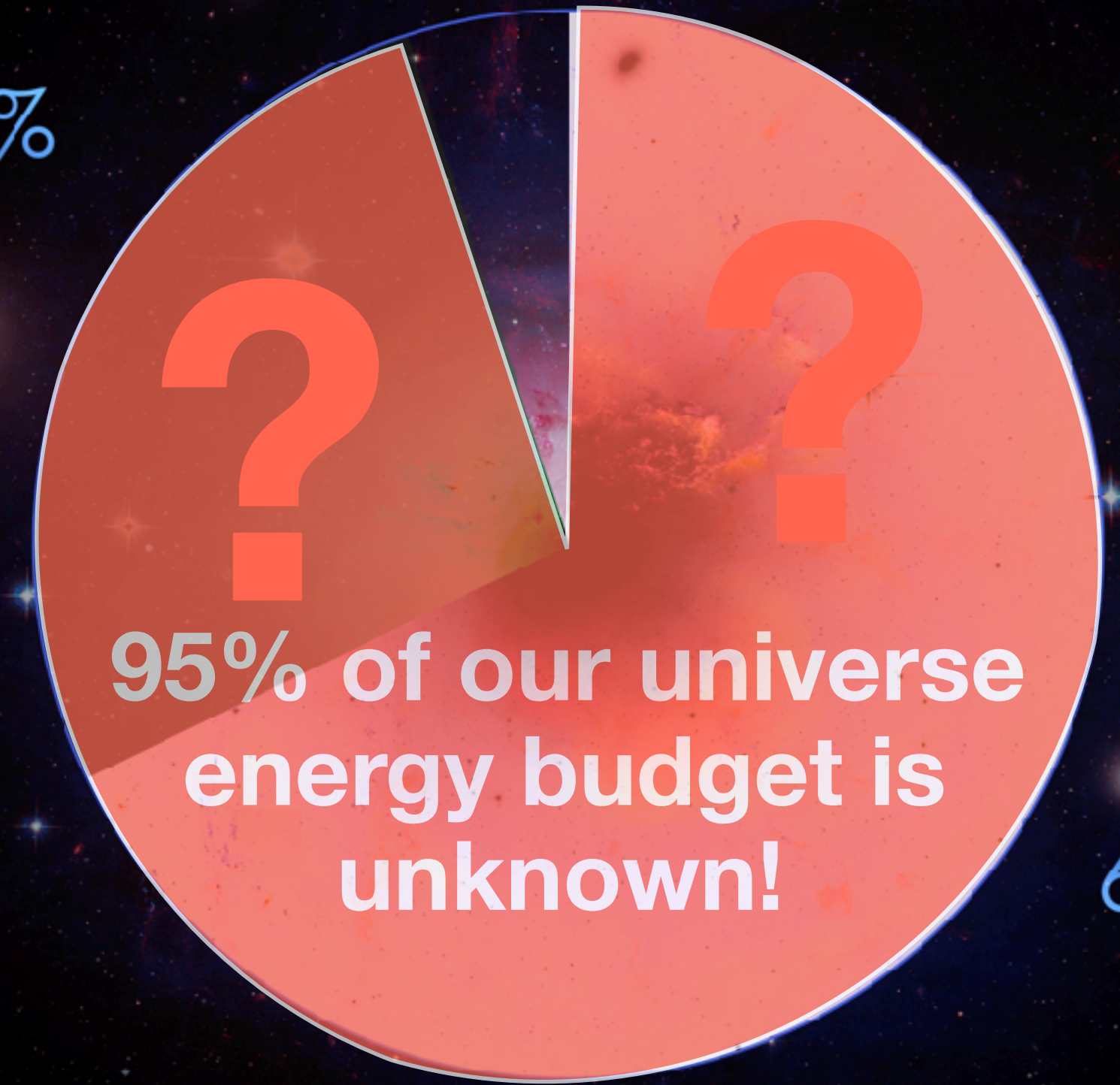
Credits: ESA & Planck Collaboration

The Standard model of Cosmology: Λ CDM



Dark matter
27%

Visible matter
5%



68%
Dark energy

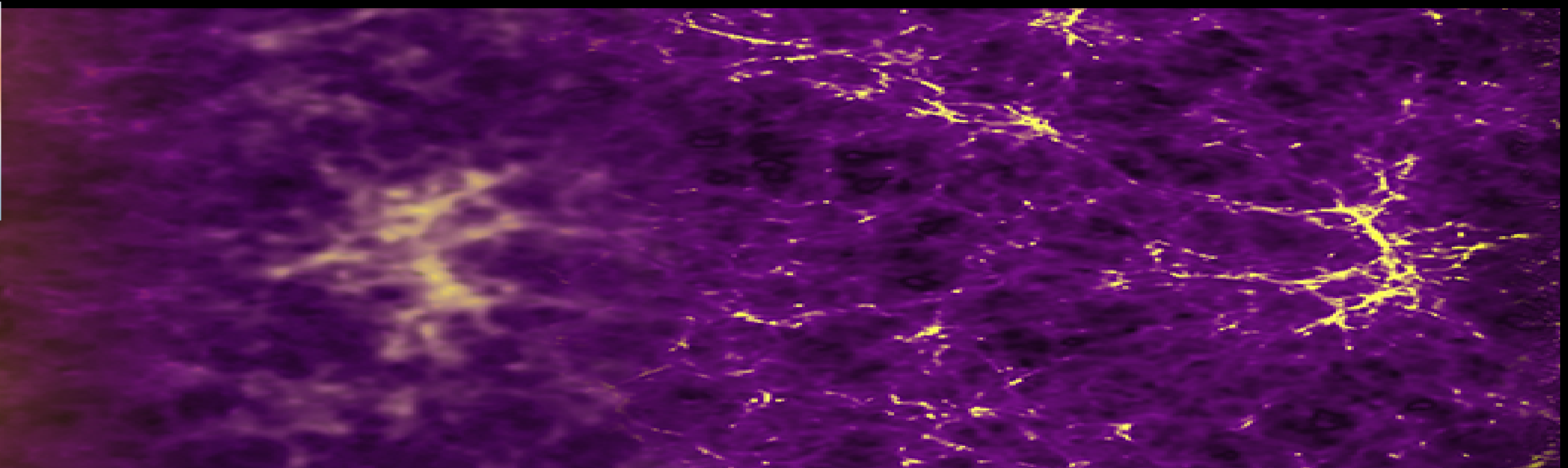
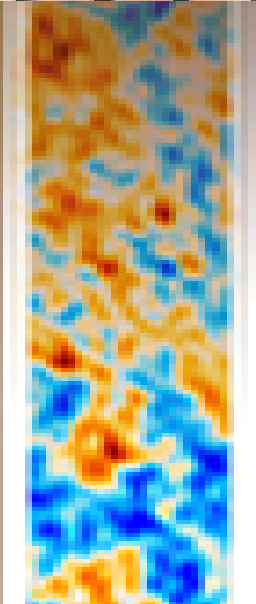
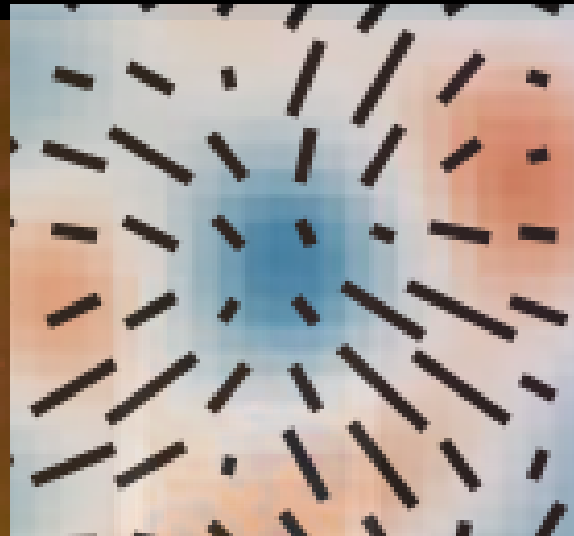
Polarization Cosmic Microwave Background

CMB in a nutshell:

- Black-body emission at $T \sim 2.7$ K (Penzias & Wilson 1965)
- Emission peaks at 100 GHz (3mm)
- Anisotropies $\delta T \sim 100 \mu K$ (COBE 1992)

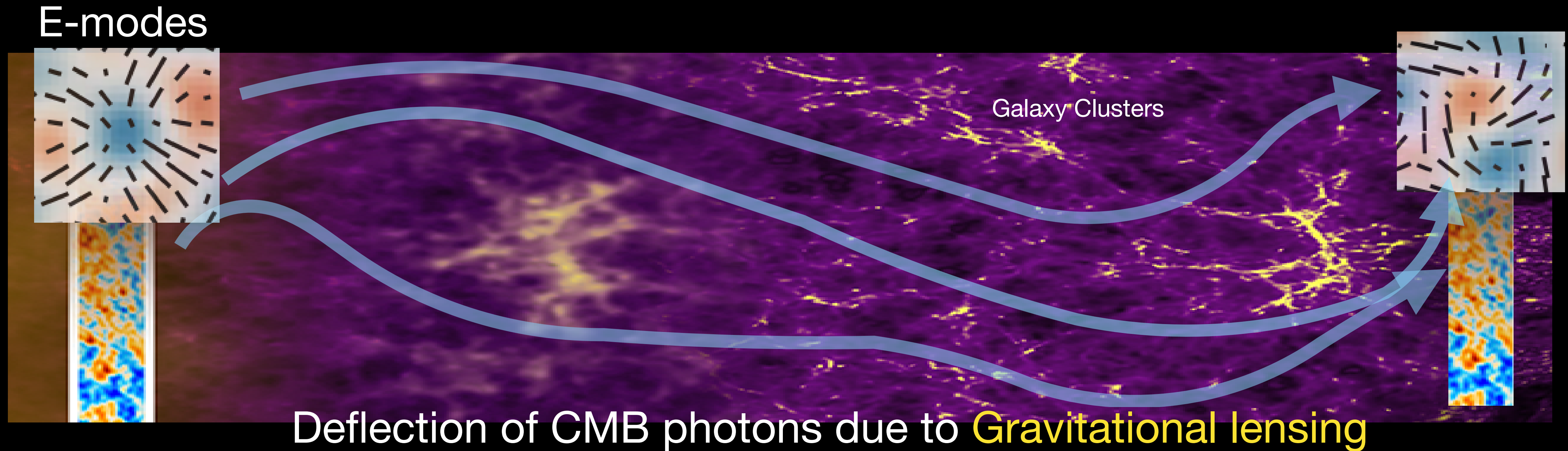
CMB Polarization
firstly detected by
DASI (Kovac et al. 2002)

E-modes



Polarization Cosmic Microwave Background

Large scale structures (e.g. galaxy clusters) gravitationally distort the CMB anisotropies, \Rightarrow **lensing B-modes**



Lensing B-modes firstly detected by Polarbear Collaboration 2014

What are Dark Matter and Dark Energy made of?

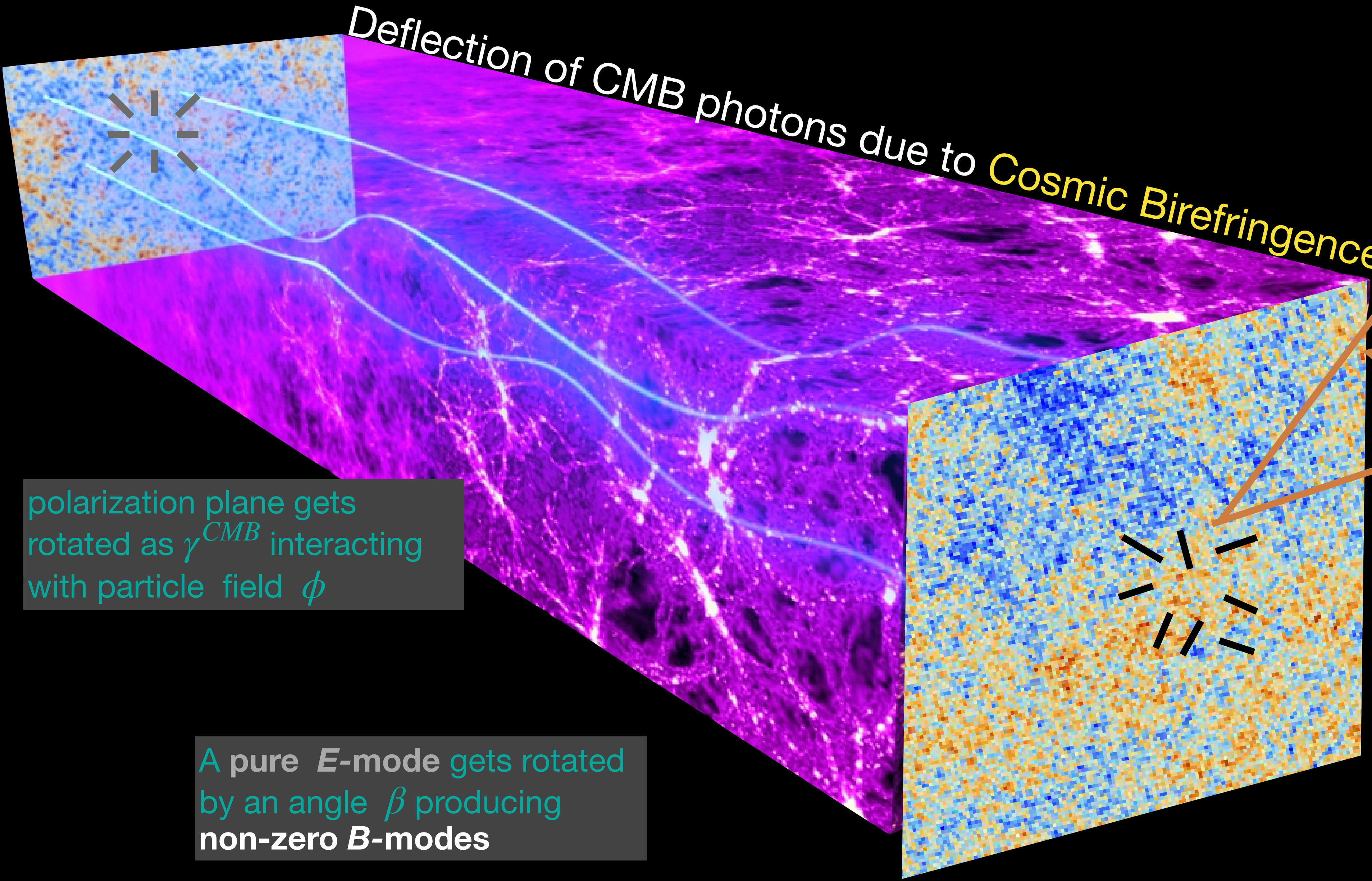
Deflection of CMB photons due to Cosmic Birefringence

$$\beta \neq 0$$

polarization plane gets rotated as γ^{CMB} interacting with particle field ϕ

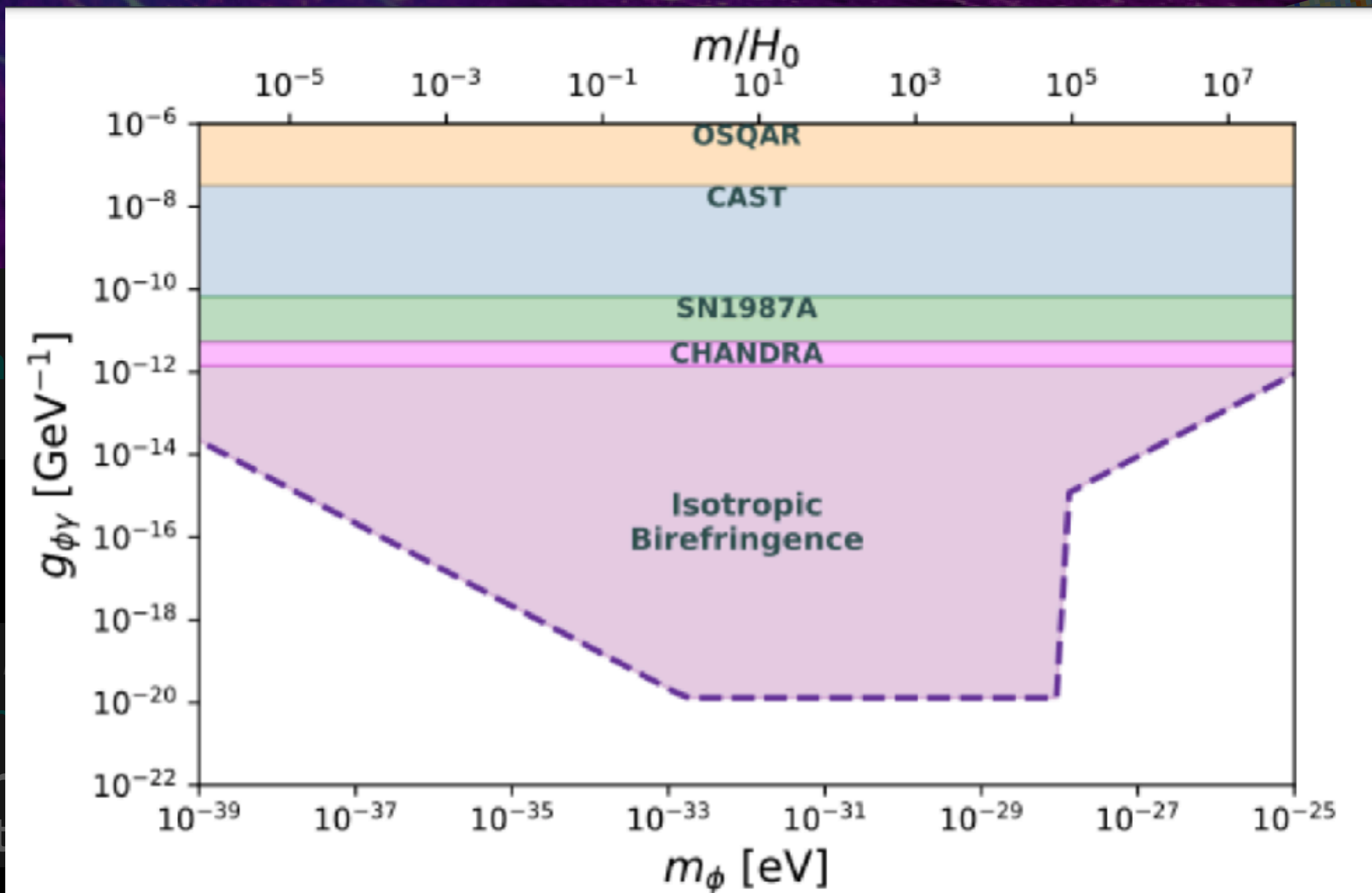
A pure *E*-mode gets rotated by an angle β producing non-zero *B*-modes

Eskilt&Komatsu (2022) claimed a 3.6σ of $\beta = 0.34^\circ$, using *Planck/WMAP* CMB polarization data



What are Dark Matter and Dark Energy made of?

If confirmed, we could explain 95% of the energy-matter budget in the universe.

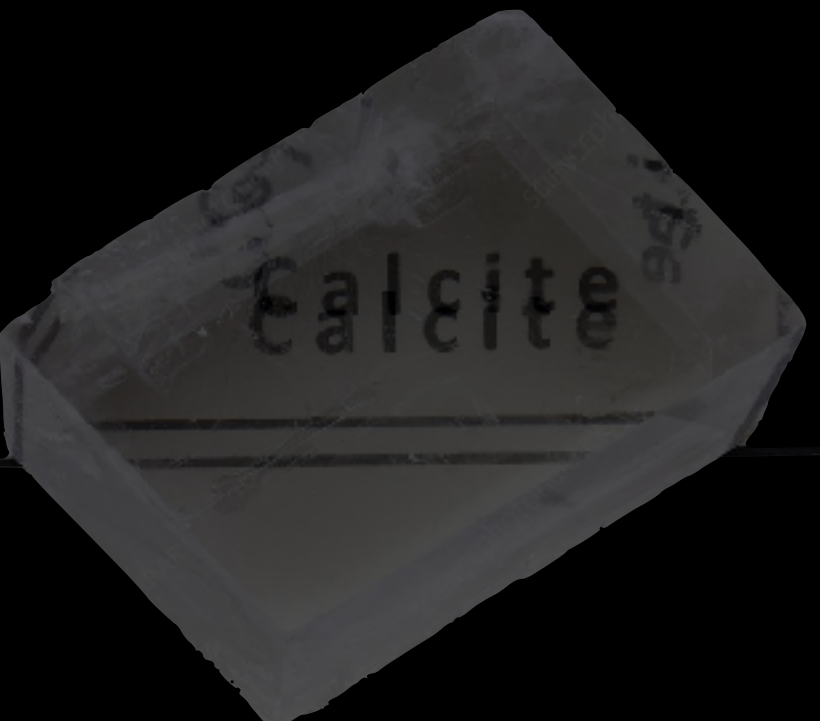


$\beta \neq 0$

Eskilt&Komatsu (2022) claimed a 3.6σ of $\beta = 0.34^\circ$, using *Planck*/*WMAP* CMB polarization data

$\implies m_\phi \sim 10^{-23} \text{ eV}$

polarization plane gets rotated as γ^{CMB} interact with a pseudo-scalar field ϕ



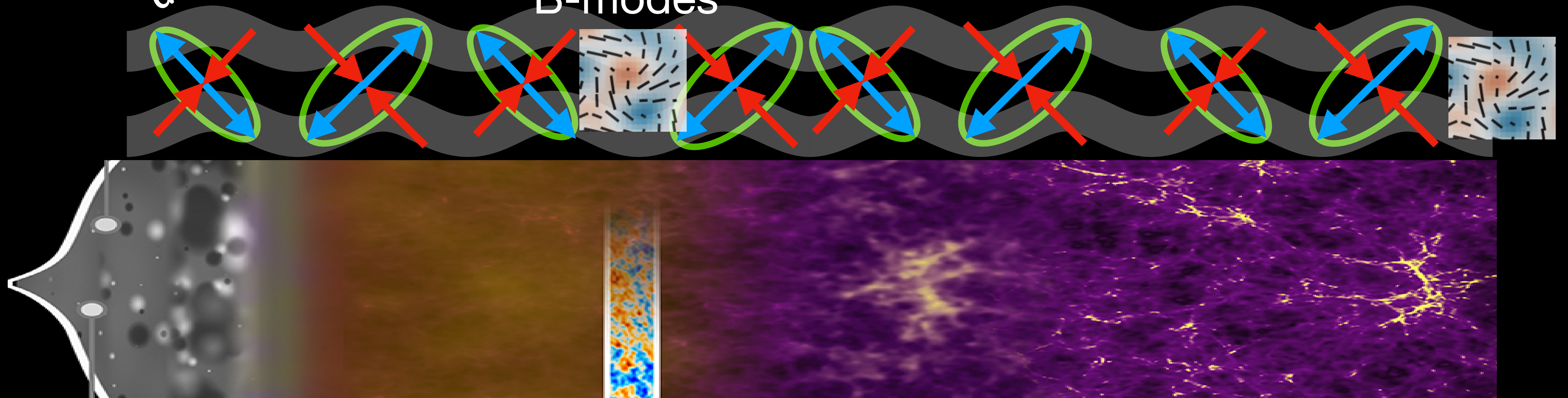
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Deflection of CMB photons due to Cosmic Birefringence

Did Inflation happened?

Gravitational waves emitted during inflation

Primordial B-modes



Did Inflation happened?

Gravitational waves emitted during inflation

Primordial B-modes

If detected, we characterize the universe $\sim 10^{-32}$ s after the Big Bang, at the energy of $\sim 10^{16}$ GeV!

⚠️ To date **Upper-limits** (Bicep/Keck Collab. 2021)

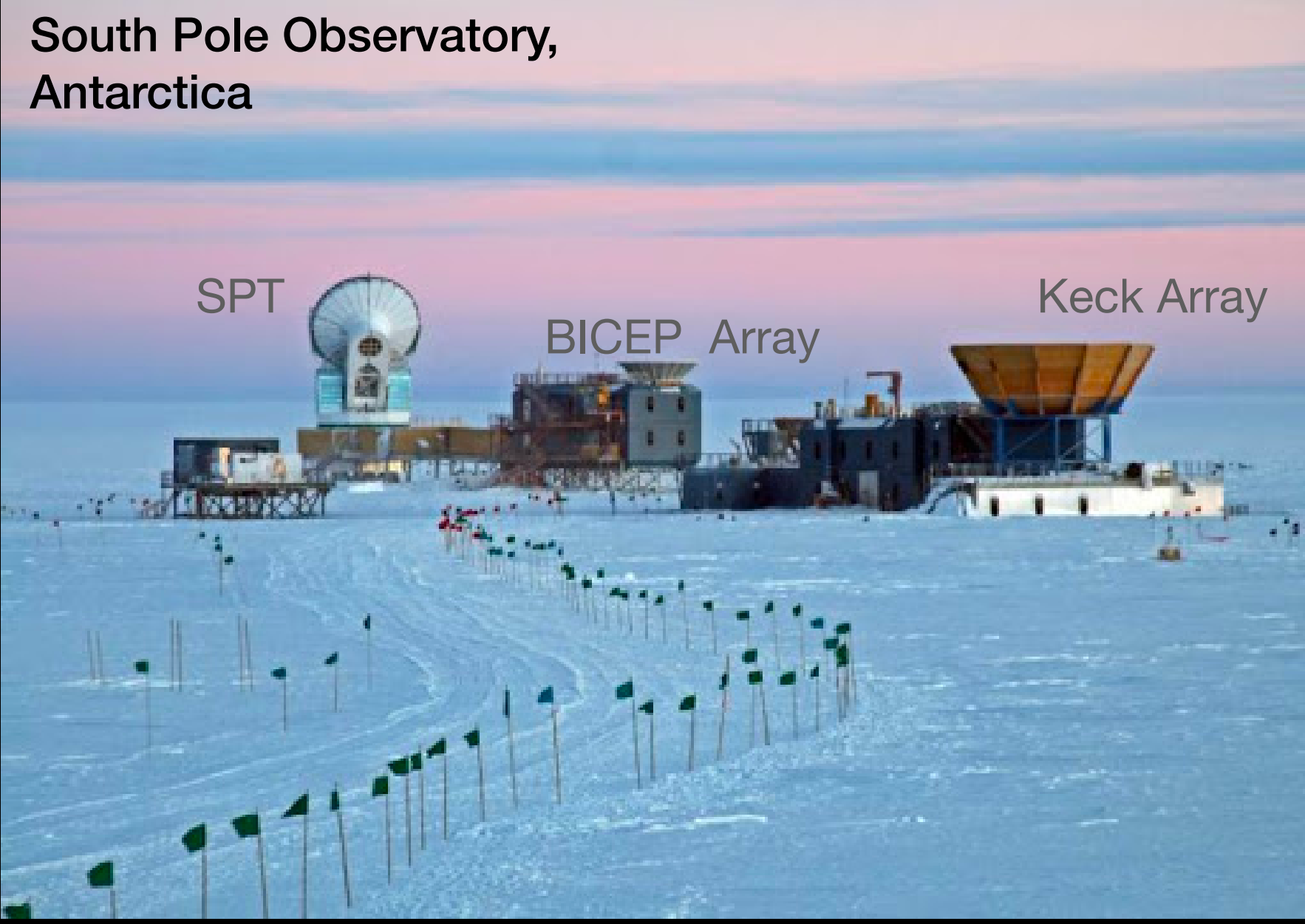
CMB forthcoming experiments

Ground-based

Simons Observatory,
Atacama Desert, Chile

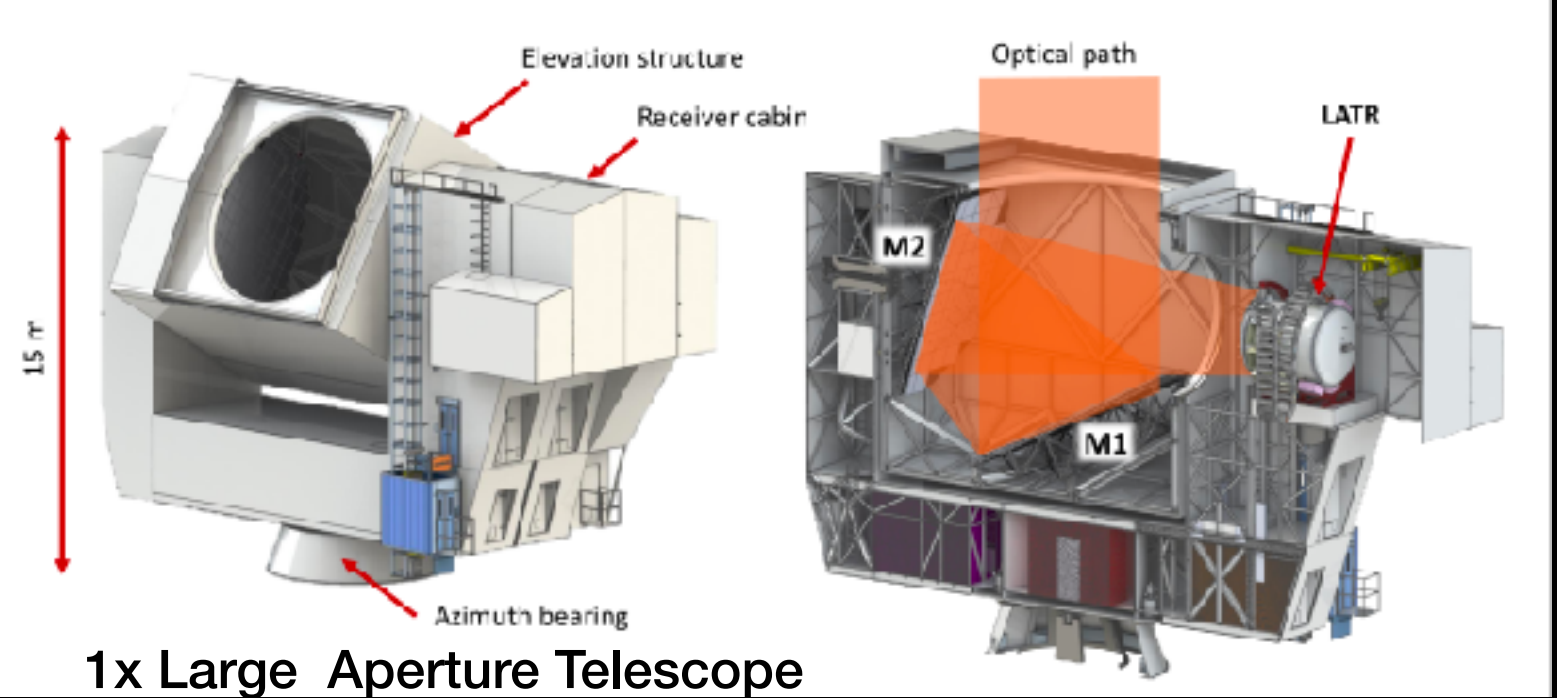
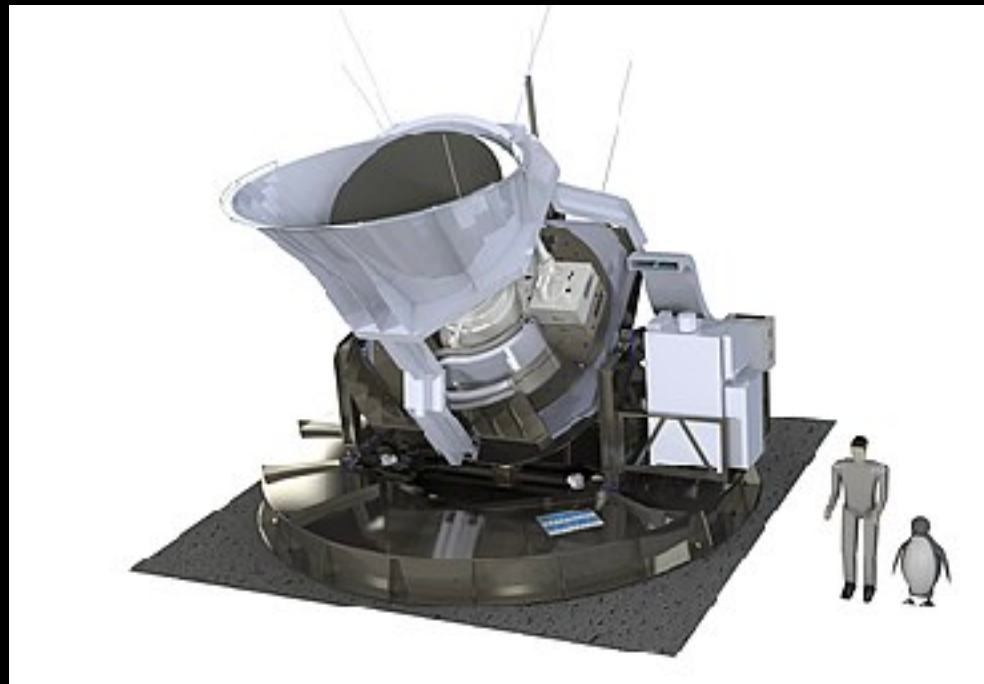


South Pole Observatory,
Antarctica



CMB-S4

$r \sim 0.003$ at $>5\sigma$ C.L.



- $>100,000$ detectors observing at:
- multiple resolutions (both degree and arcminute)
- Multiple frequency bands 20-200 GHz

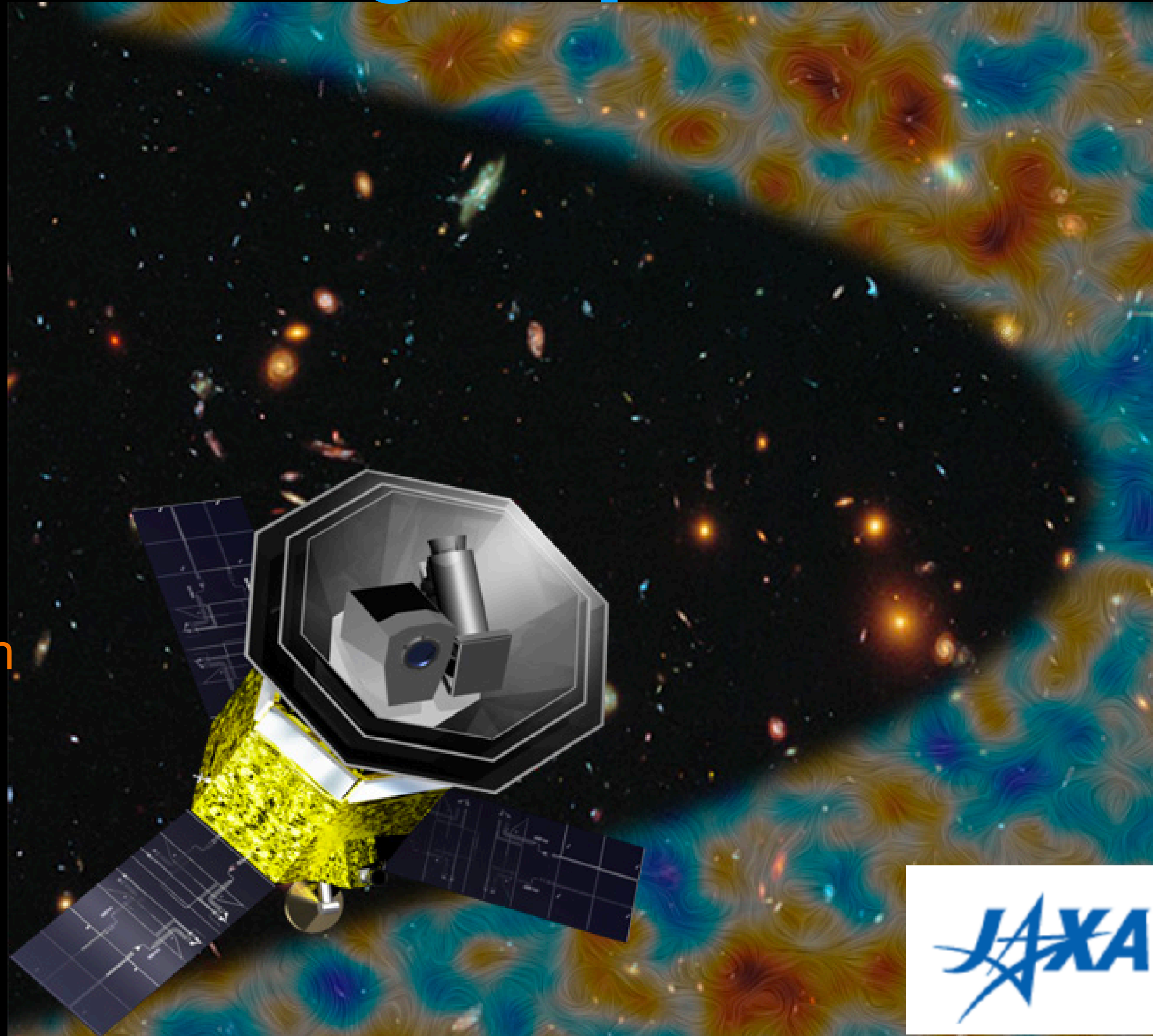
CMB forthcoming experiments

Space-mission

LiteBIRD

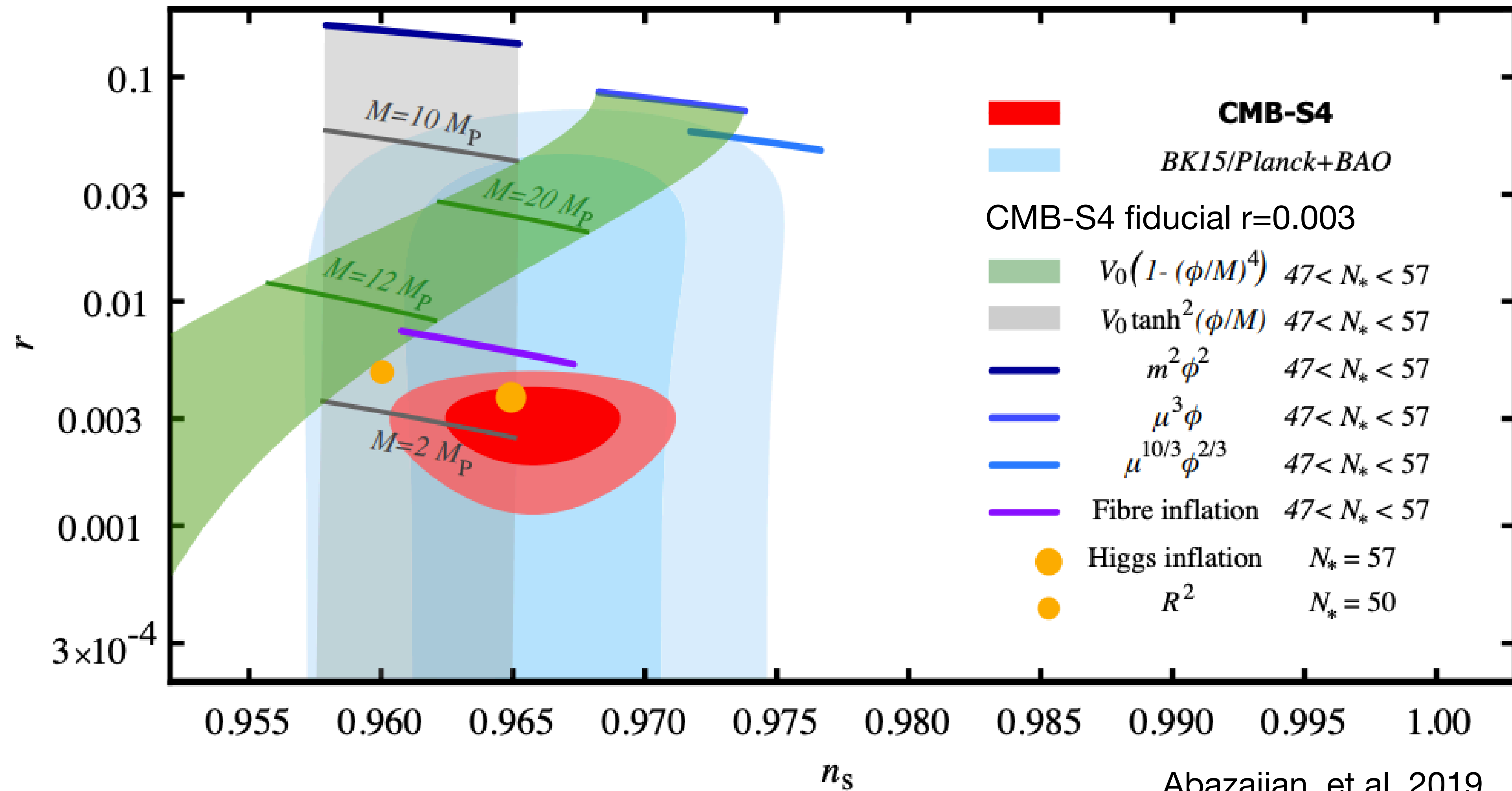
$r \sim 0.001$ at $>5\sigma$ C.L.

- 22 frequency bands (40-400 GHz)
- 10-70 arcmin resolution
- ~4500 detectors

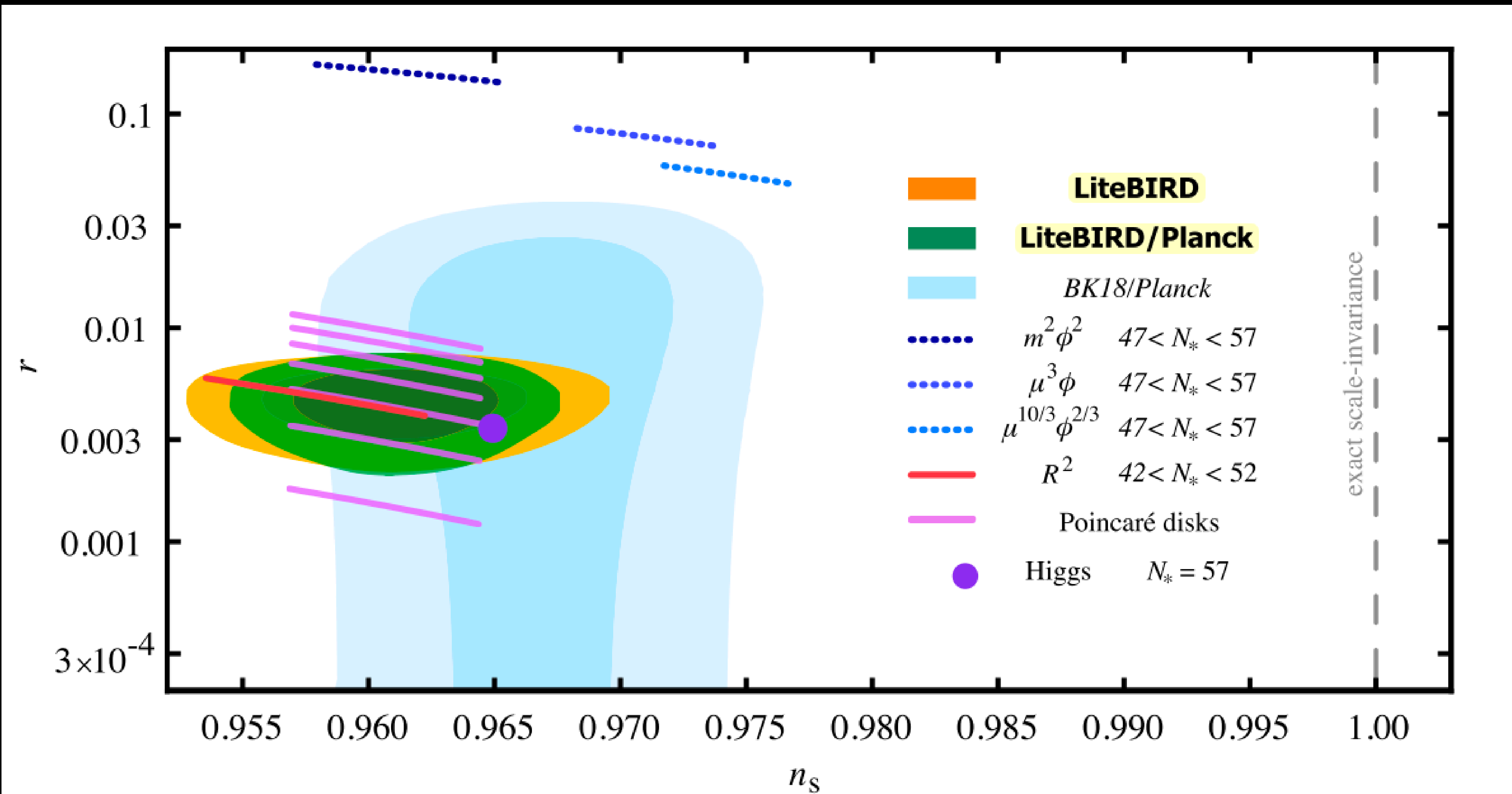


Japan Aerospace Exploration Agency

CMB-Stage 4 forecasts on r



LiteBIRD forecasts on r



Summary

- **Coming soon:** new data will bring to 5σ the claim of Cosmic birefringence
- In the coming decade Litebird and CMB-S4 will achieve first light to set stringent constraints on tensorial perturbations, $r \sim 0.001$ at 5σ C.L., shedding light to the universe state at 10^{16} GeV
- Multimessenger Astrophysics (microwave): Tidal Disruption Events, Stellar Flares, Variable AGNs

Thank you