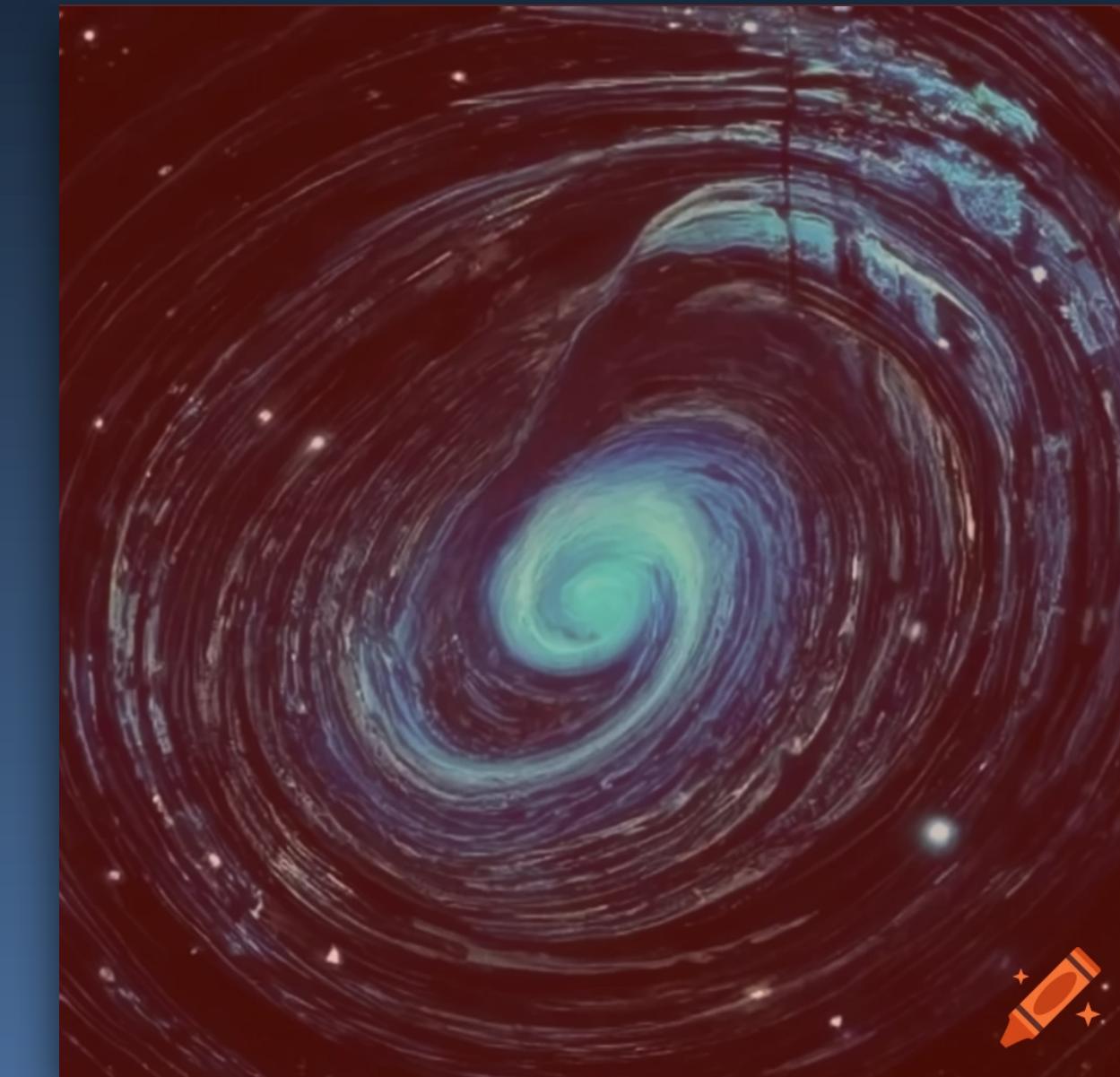
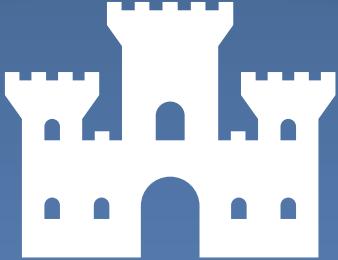


Testing Gravity through the Distortion of Time



Sveva Castello 



UNIVERSITÉ
DE GENÈVE

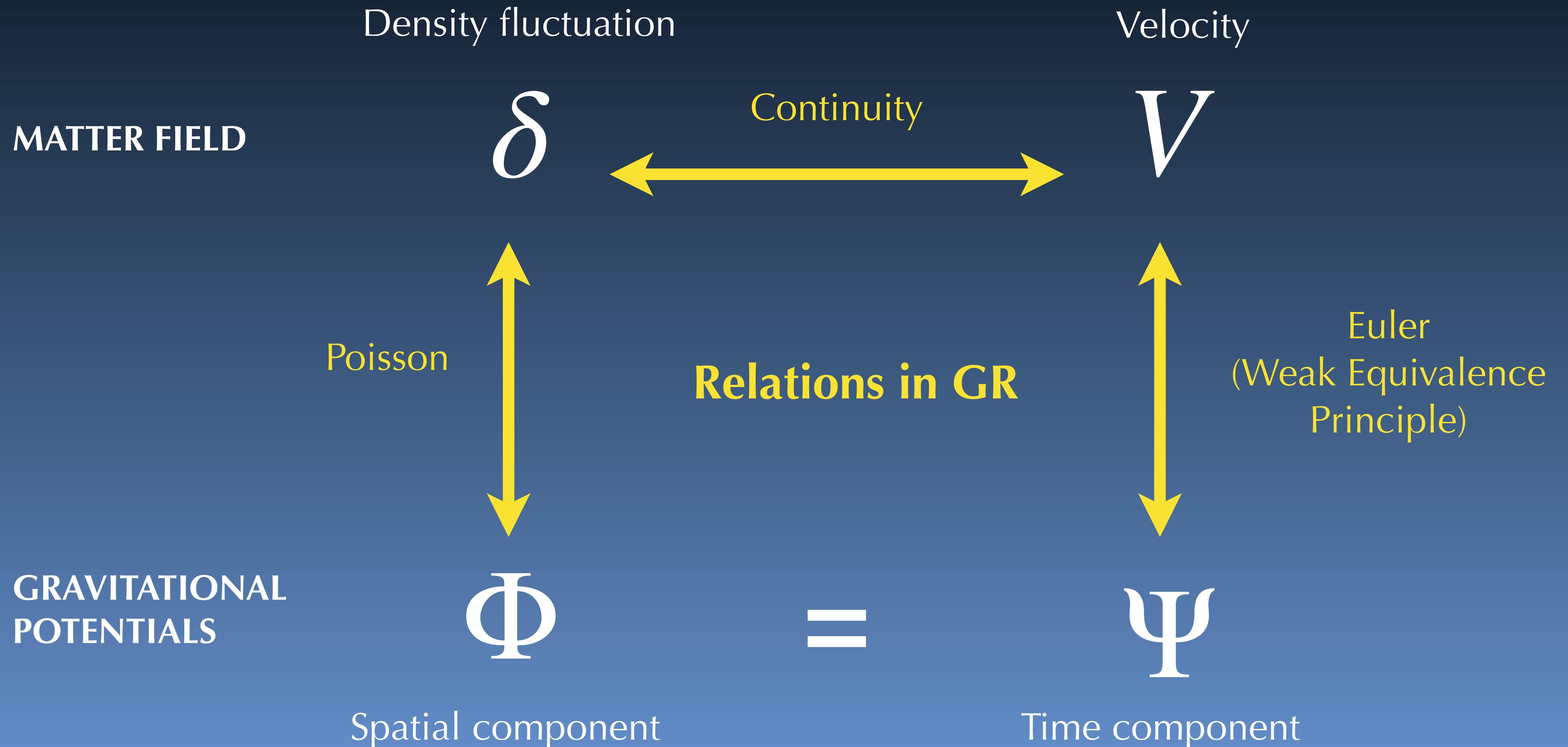
FACULTÉ DES SCIENCES

CASTLE 2024
Tagliolo Monferrato, September 19th, 2024

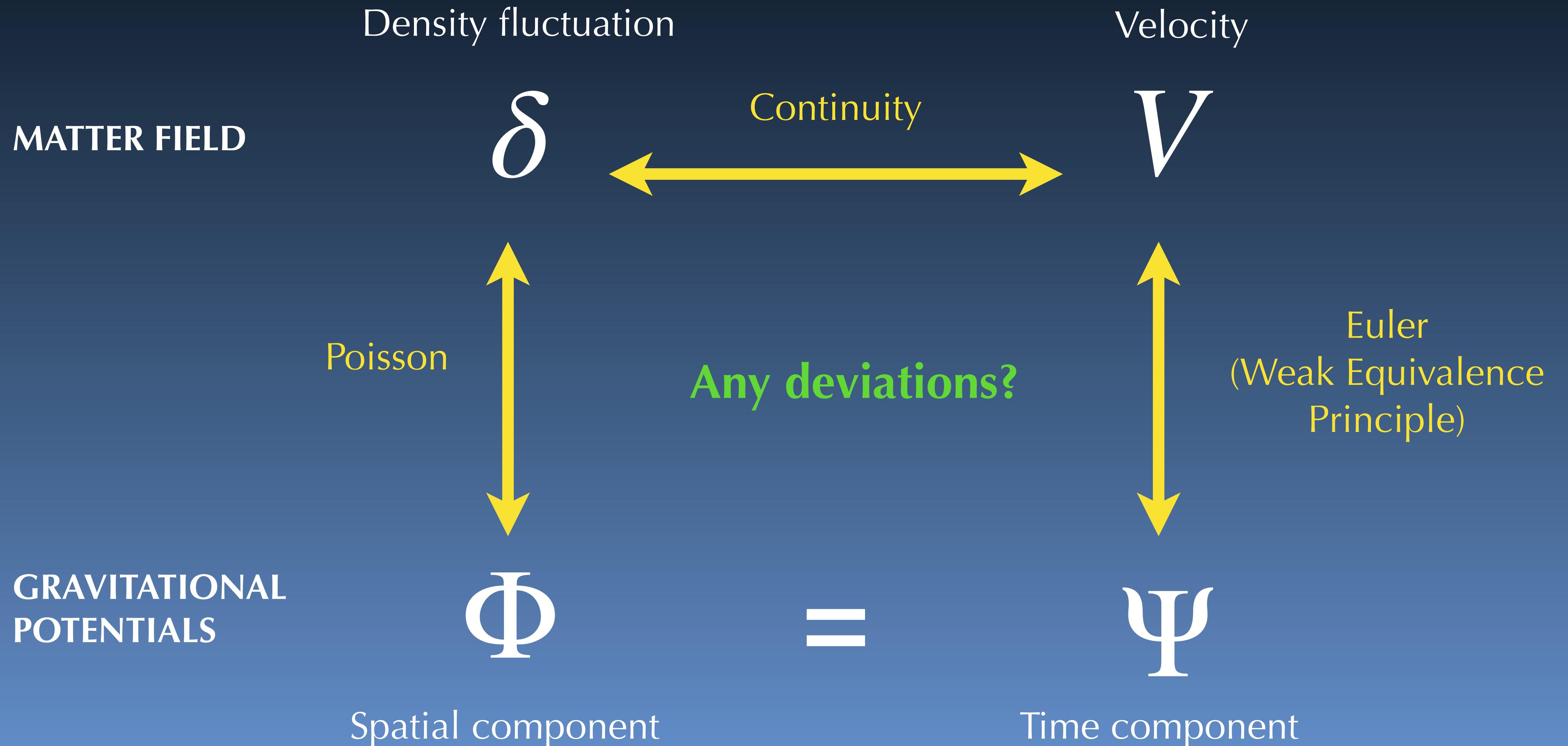


erc
European Research Council
Established by the European Commission

Describing the Universe with four fields

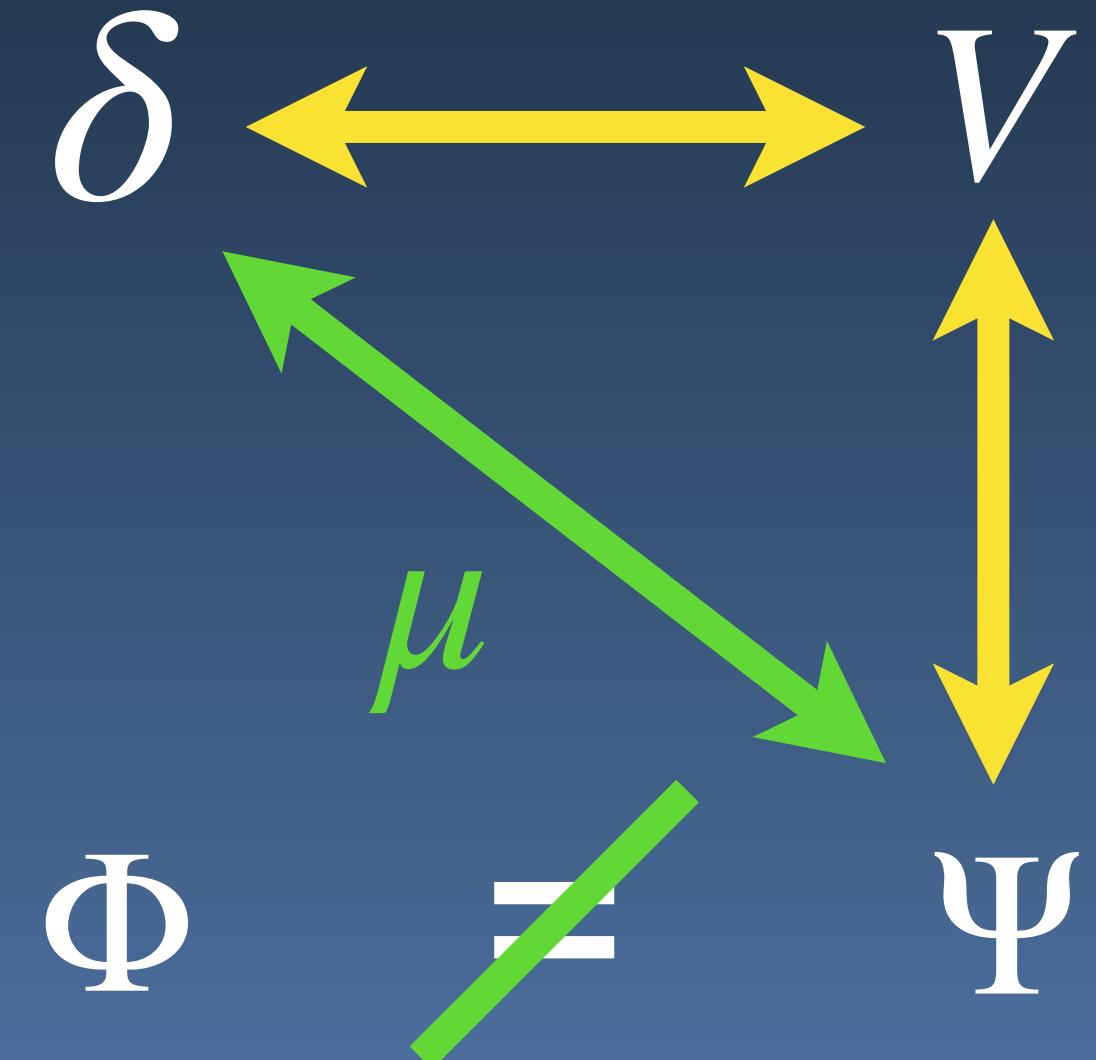


Describing the Universe with four fields

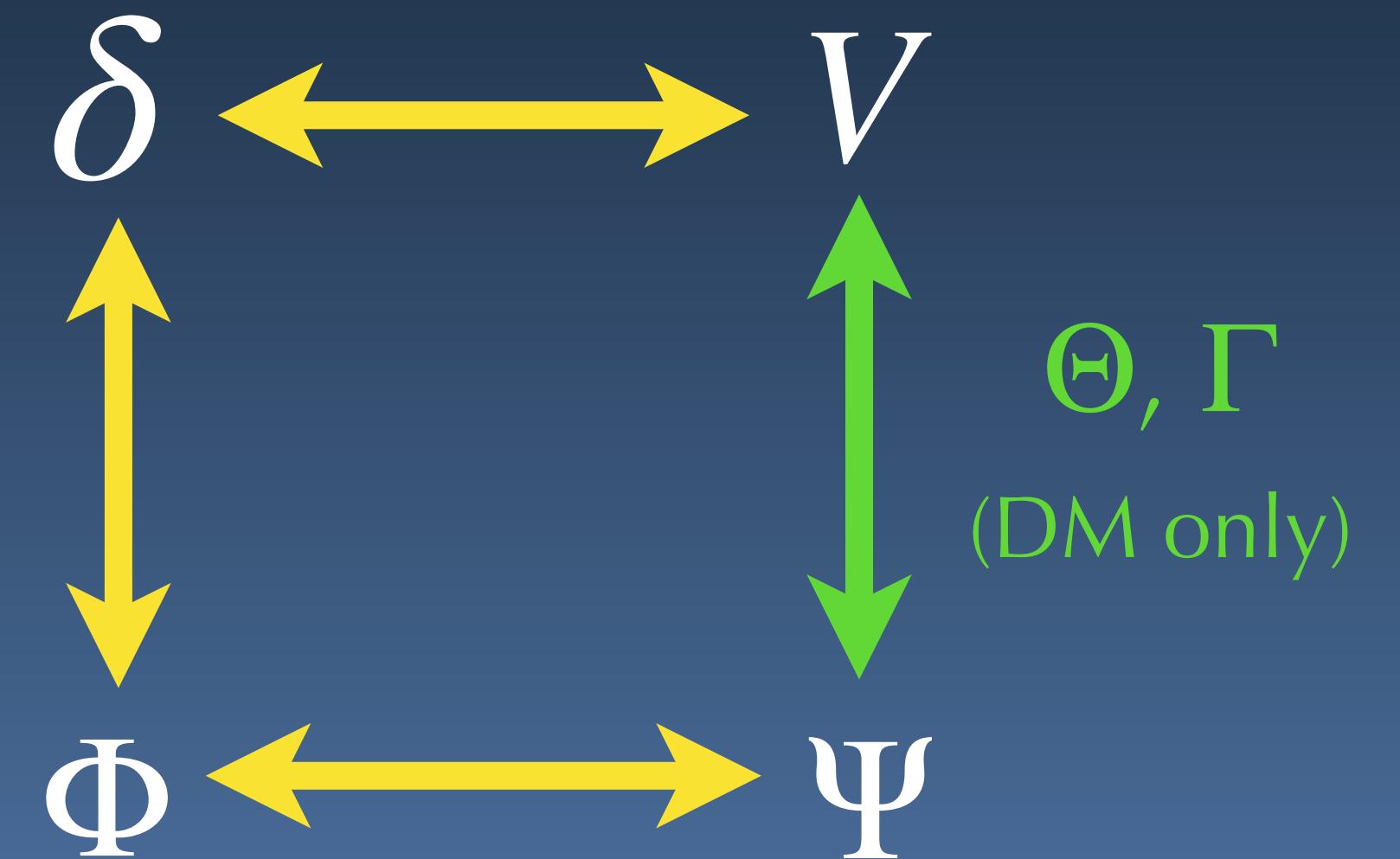


Two scenarios

Gravity modifications



Dark sector interactions



Can we distinguish between the two?

Galaxy clustering

Fluctuations in galaxy number counts

$$\Delta(z, \mathbf{n}) = b \delta_{\text{DM}} - \frac{1}{\mathcal{H}} \partial_r (\mathbf{V} \cdot \mathbf{n})$$

DM density
x galaxy bias

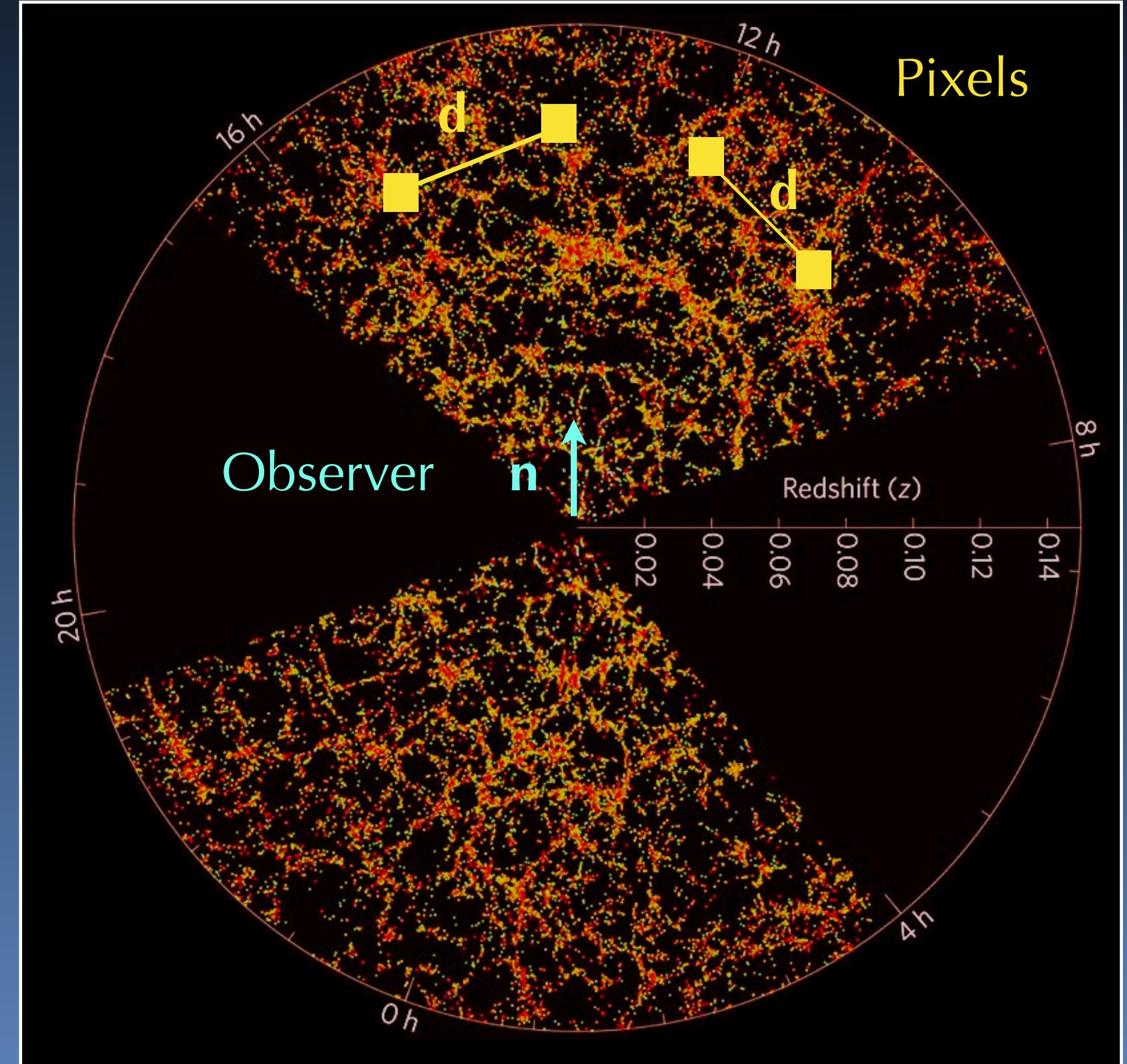
Redshift-space
distortions (RSD)

Two-point correlation function

$$\xi \equiv \langle \Delta(z, \mathbf{n}) \Delta(z', \mathbf{n}') \rangle$$



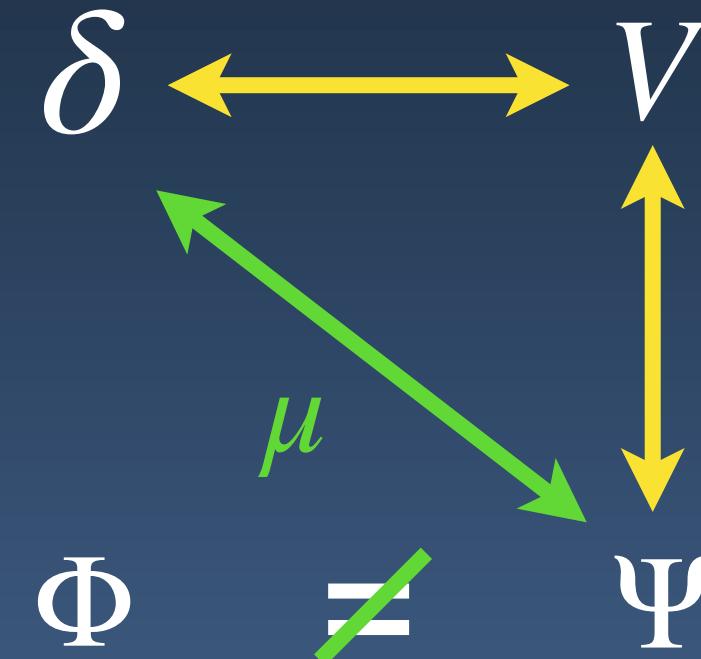
Extracted from observations and
compared with theoretical predictions



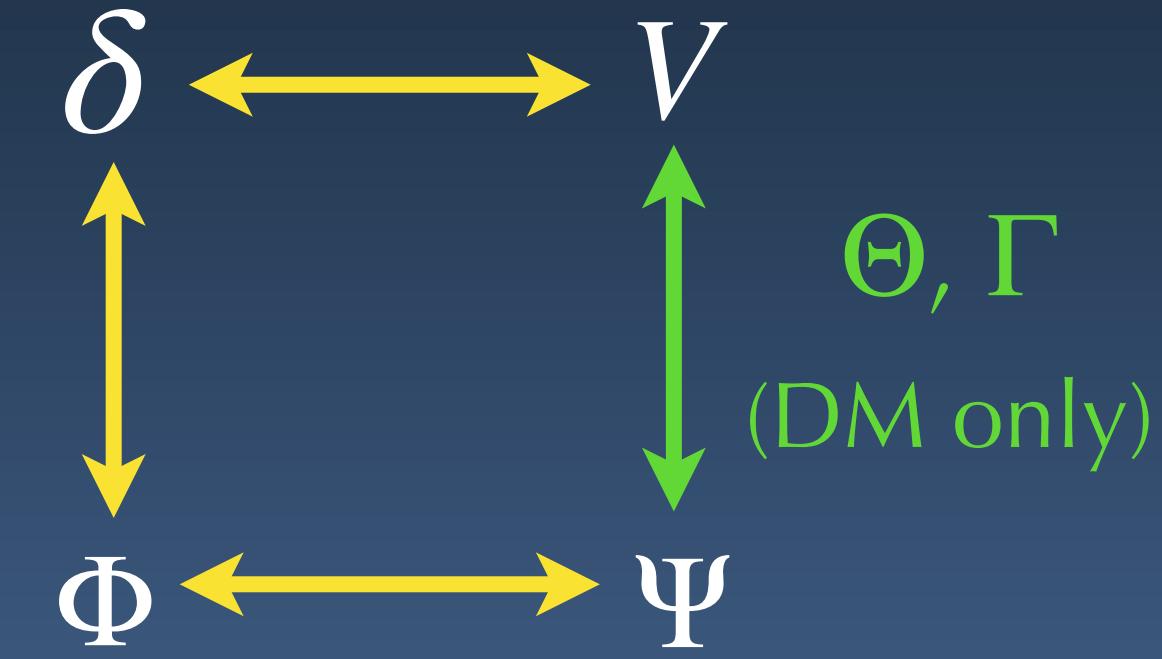
Credits: M.Blanton, SDSS

Impact of the modifications

Gravity modifications



Dark sector interactions



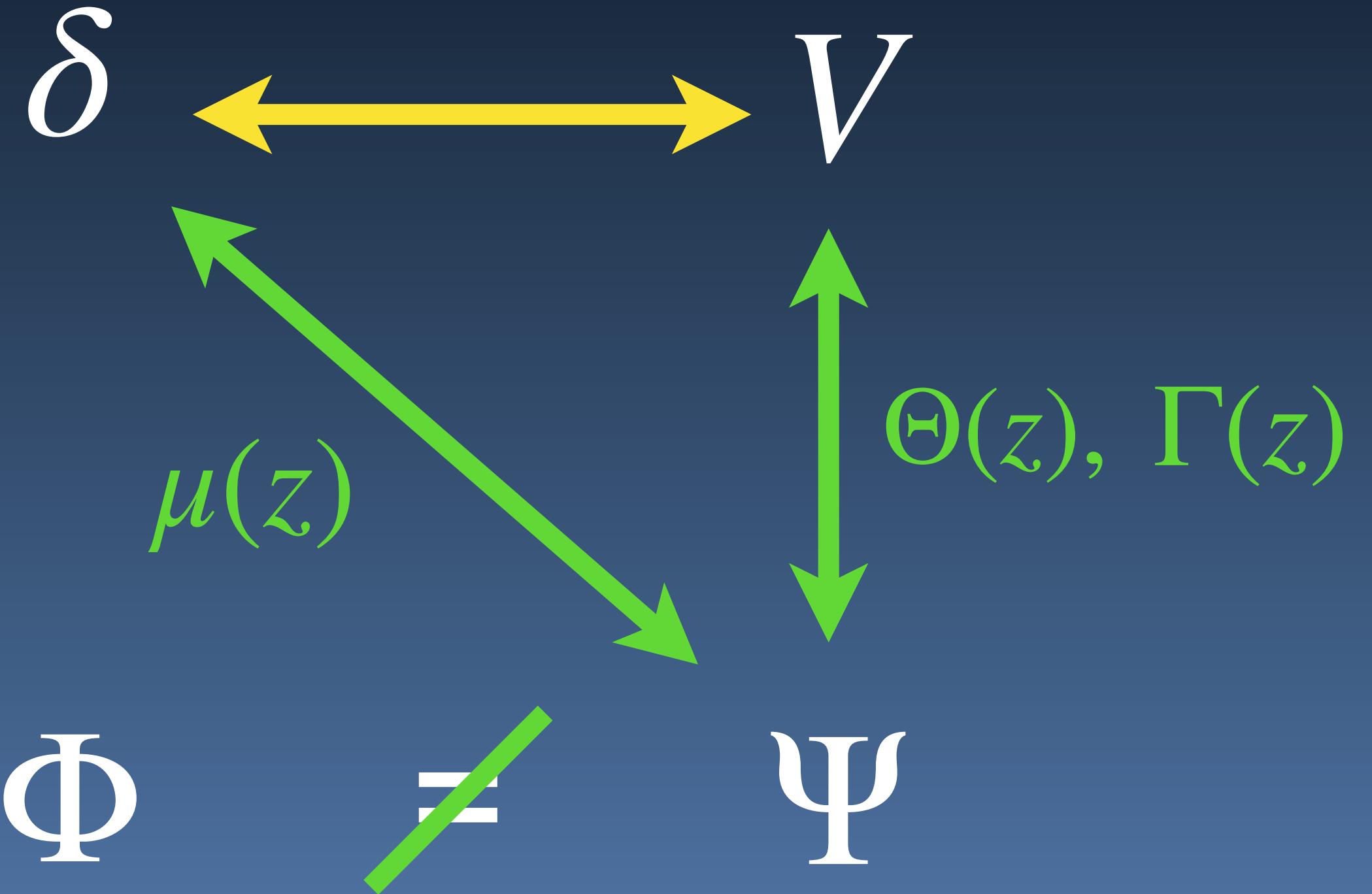
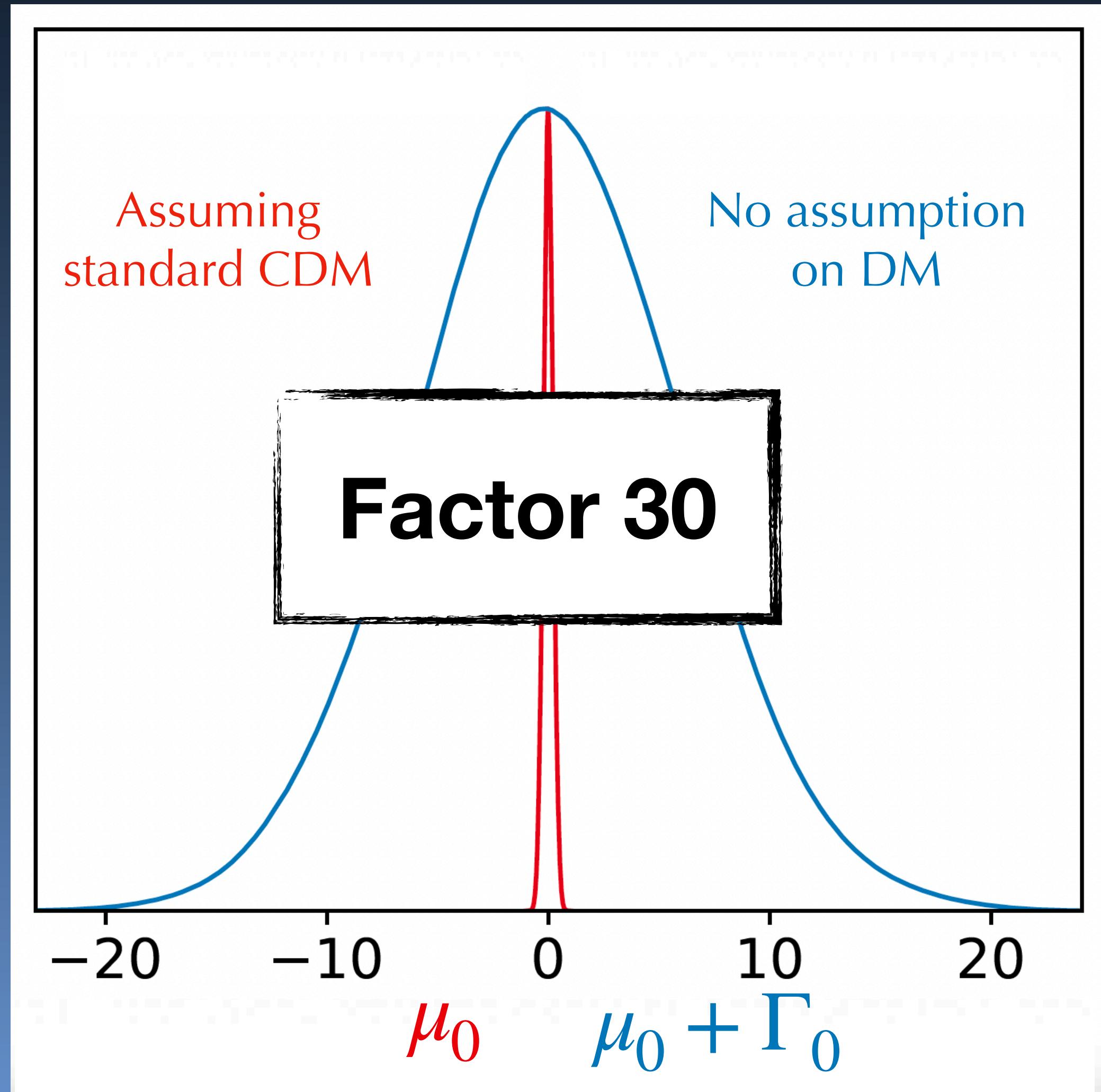
$$\delta'' + \left(1 + \frac{\mathcal{H}}{\mathcal{H}'}\right) \delta' - \frac{3}{2} \frac{\Omega_{m,0}}{a} \left(\frac{\mathcal{H}_0}{\mathcal{H}}\right)^2 \mu \delta = 0$$

$$\delta'' + \left(1 + \frac{\mathcal{H}}{\mathcal{H}'} + \Theta\right) \delta' - \frac{3}{2} \frac{\Omega_{m,0}}{a} \left(\frac{\mathcal{H}_0}{\mathcal{H}}\right)^2 (\Gamma + 1) \delta = 0$$

DEGENERATE EFFECTS

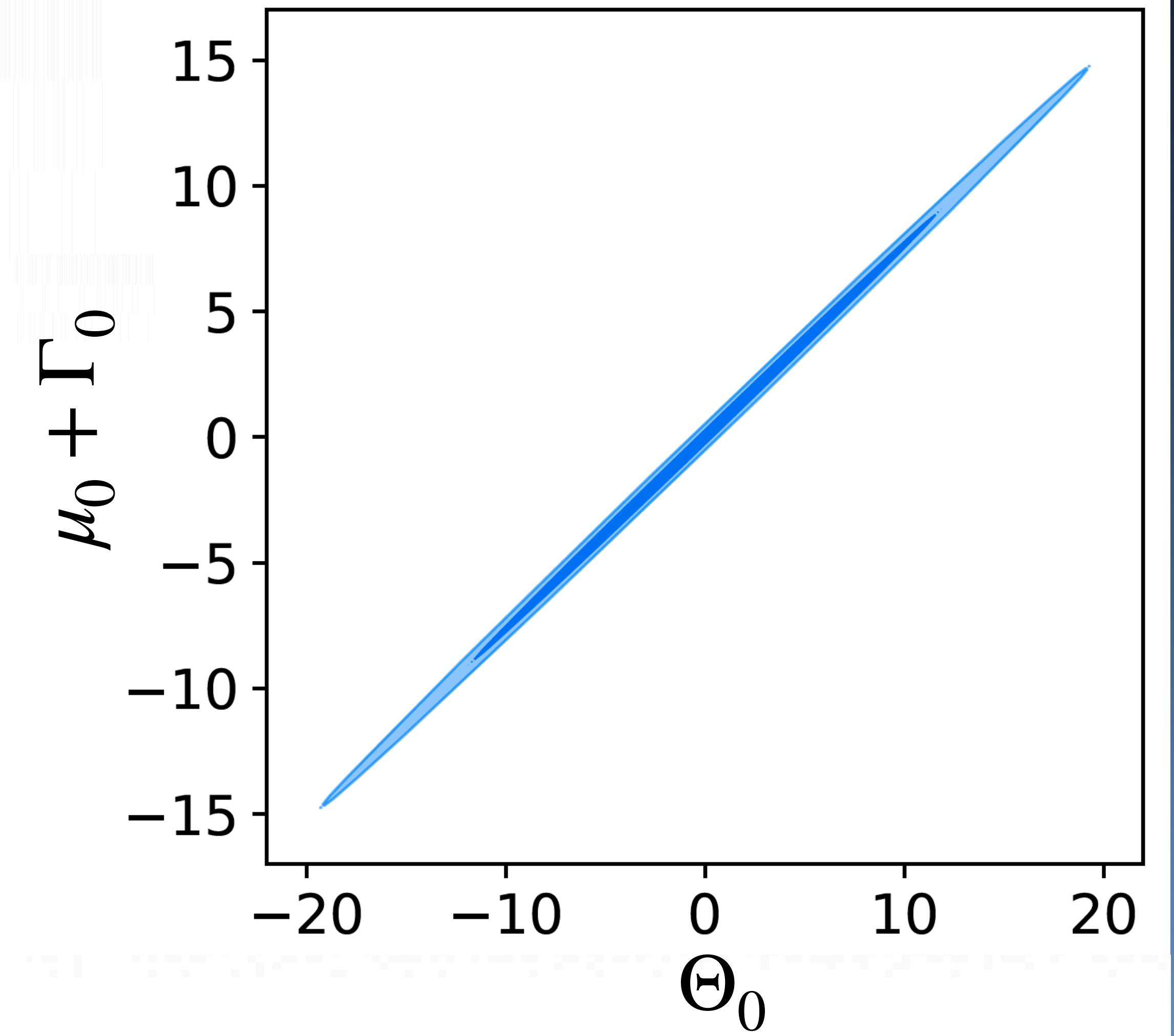
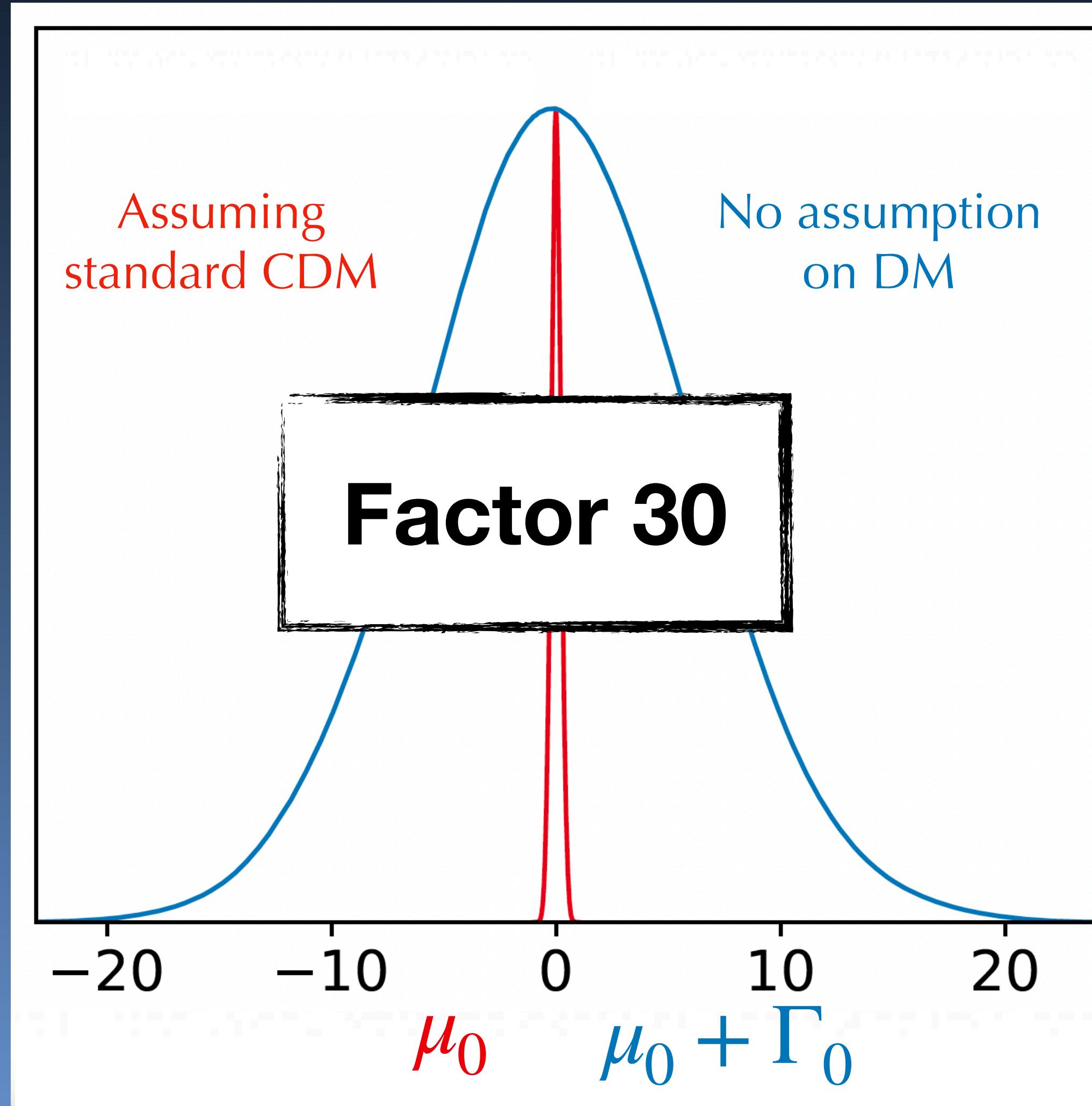
Precision with SDSS data

SC, Grimm and Bonvin (2022)



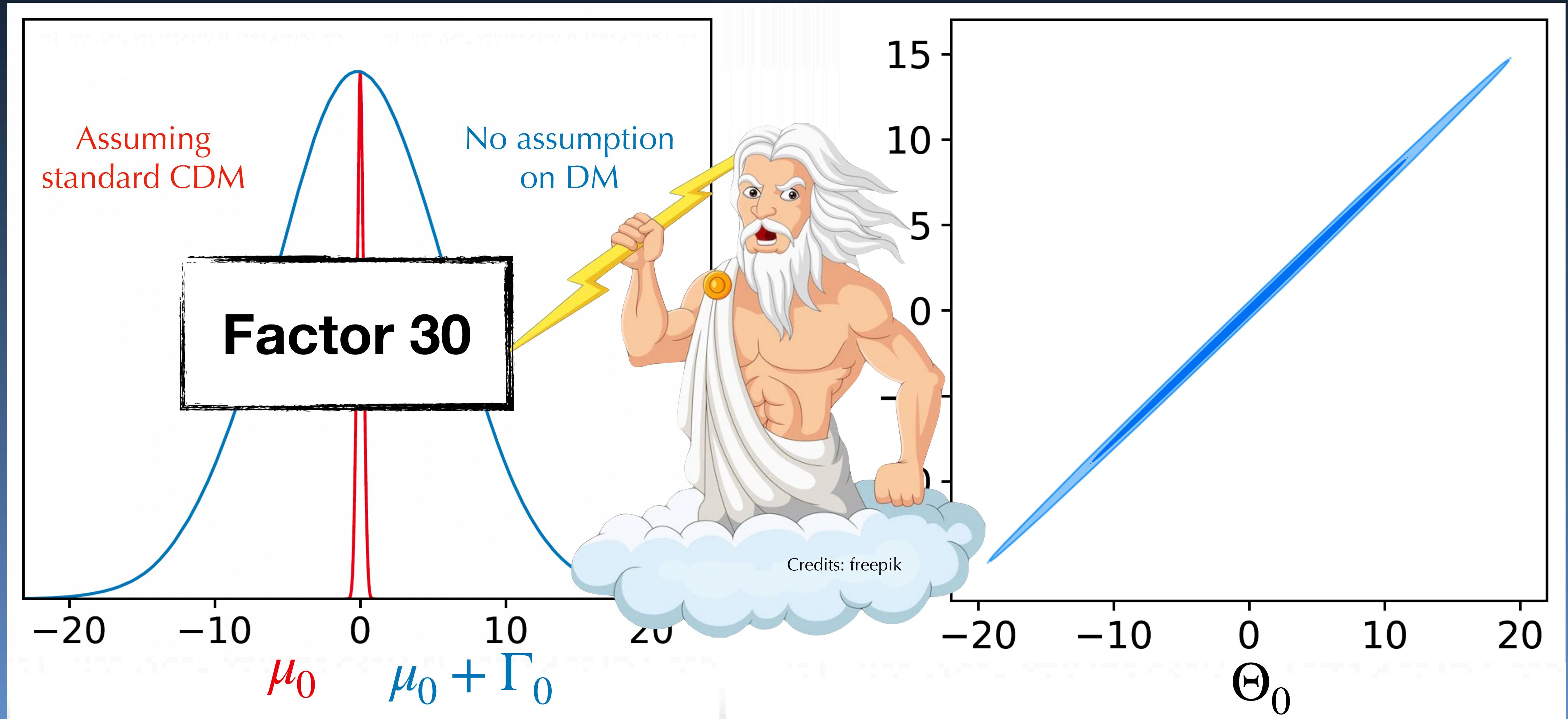
Precision with SDSS data

SC, Grimm and Bonvin (2022)



Precision with SDSS data

SC, Grimm and Bonvin (2022)



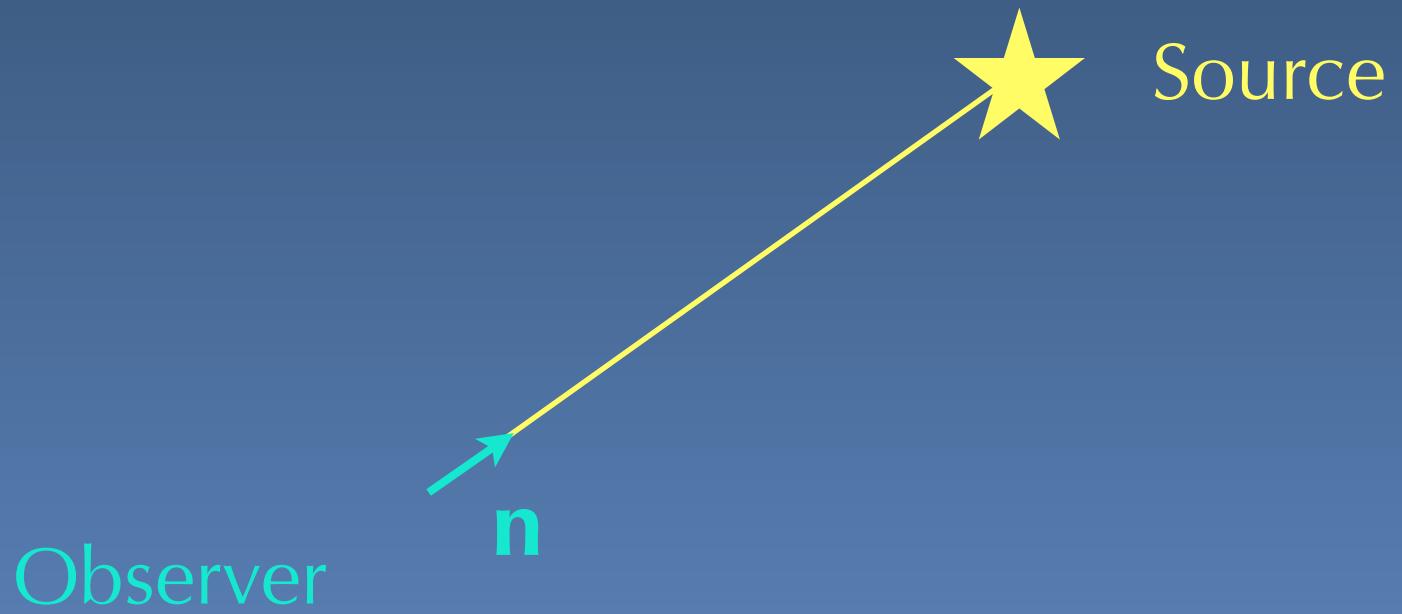
What we really observe

Yoo et al. (2010)
Bonvin and Durrer (2011)
Challinor and Lewis (2011)

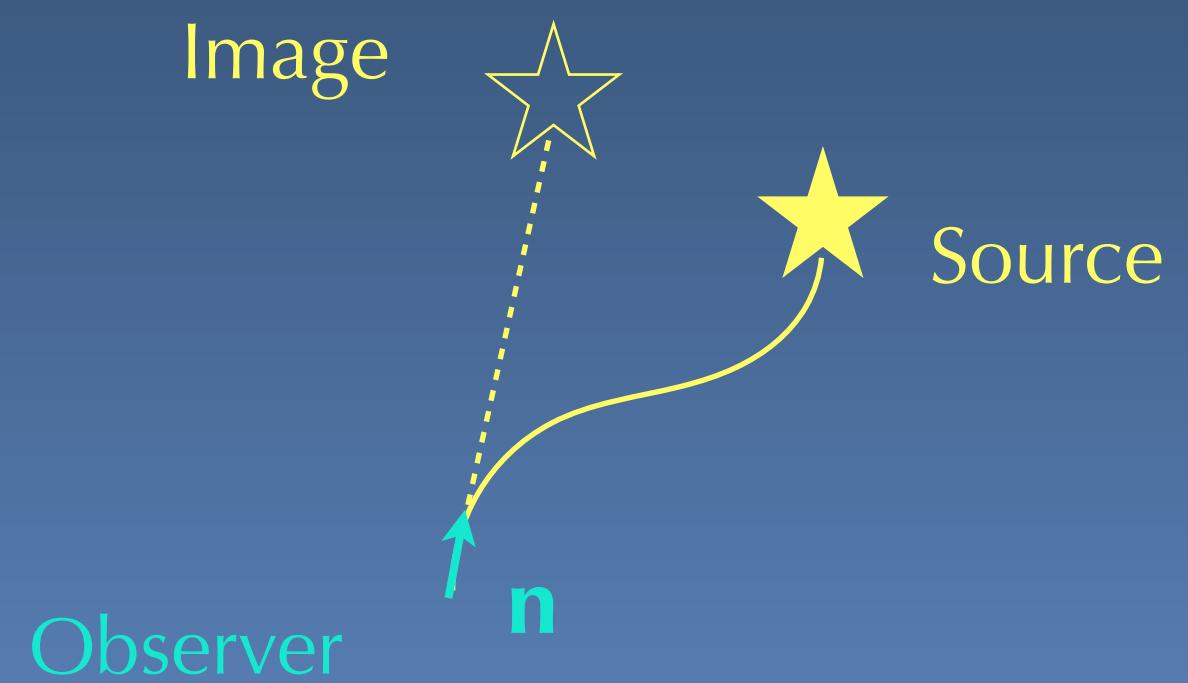
$$\Delta(\mathbf{n}, z) = b \delta_{\text{DM}} - \frac{1}{\mathcal{H}} \partial_r (\mathbf{V} \cdot \mathbf{n})$$

+ relativistic corrections

Homogeneous Universe



Inhomogeneous Universe



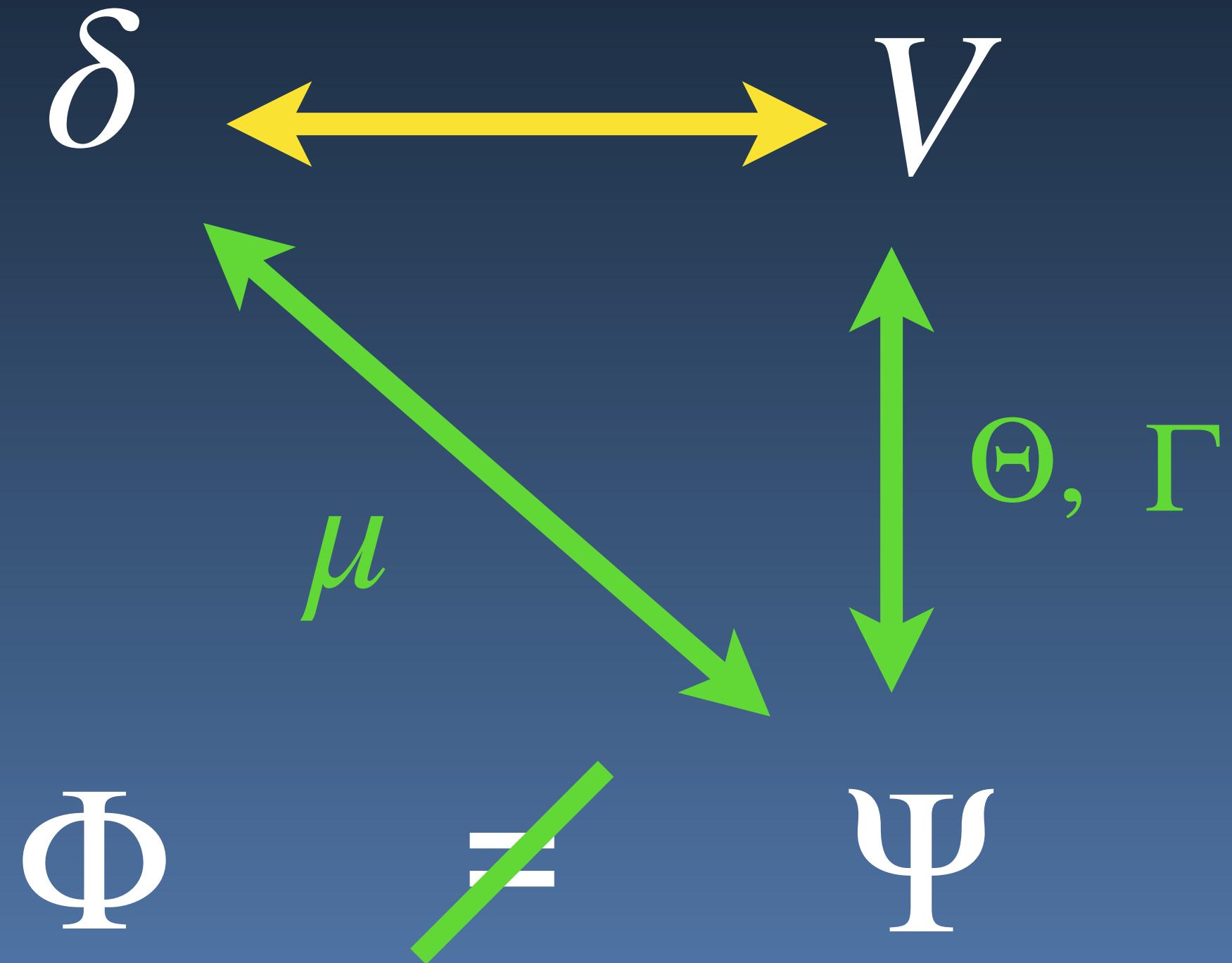
Calculate effects within
linear perturbation theory

Deus ex machina: gravitational redshift

SC, Grimm and Bonvin (2022)

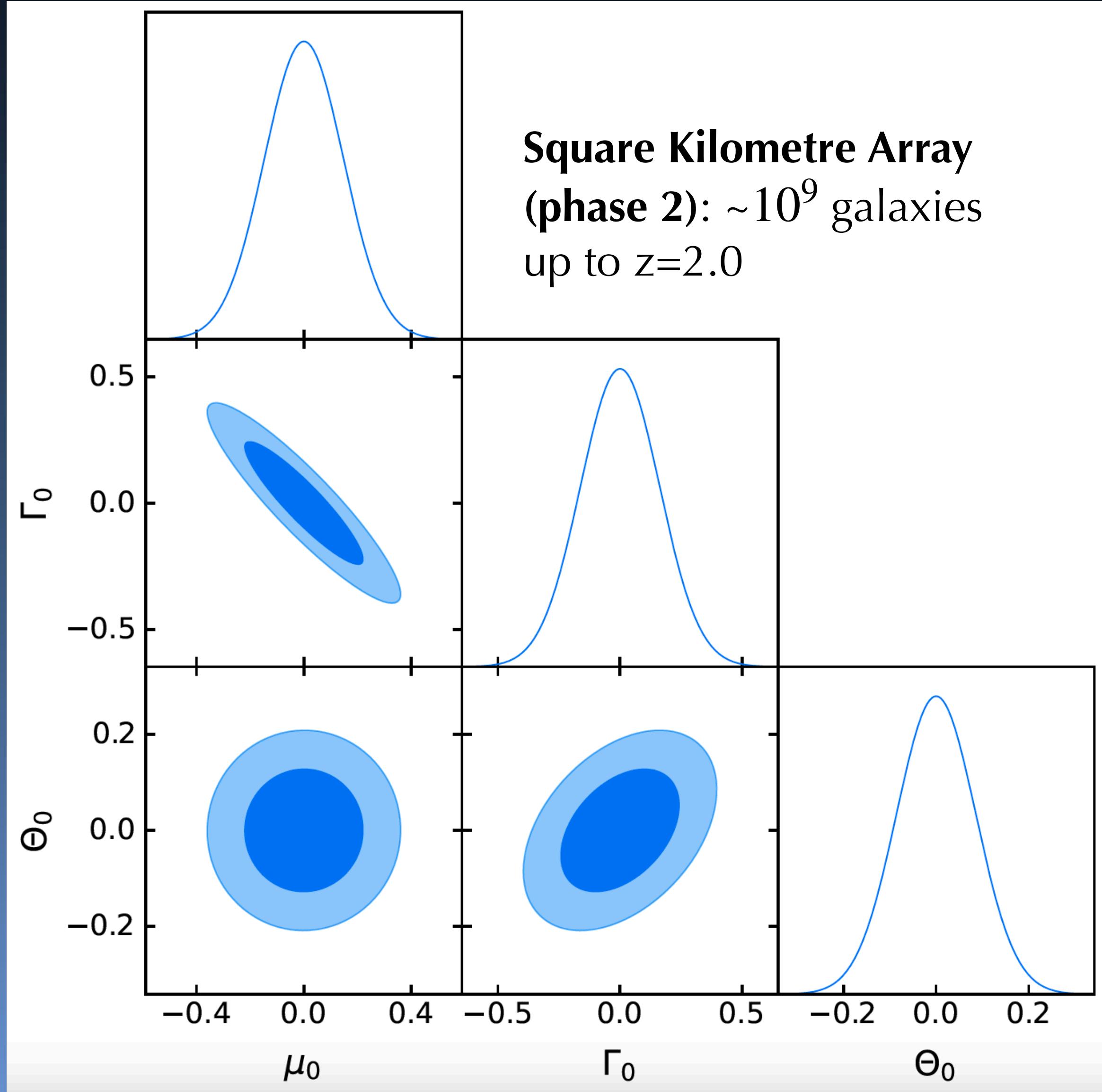


$$\Delta_{\text{gr}} = \frac{1}{\mathcal{H}} \partial_r \Psi$$

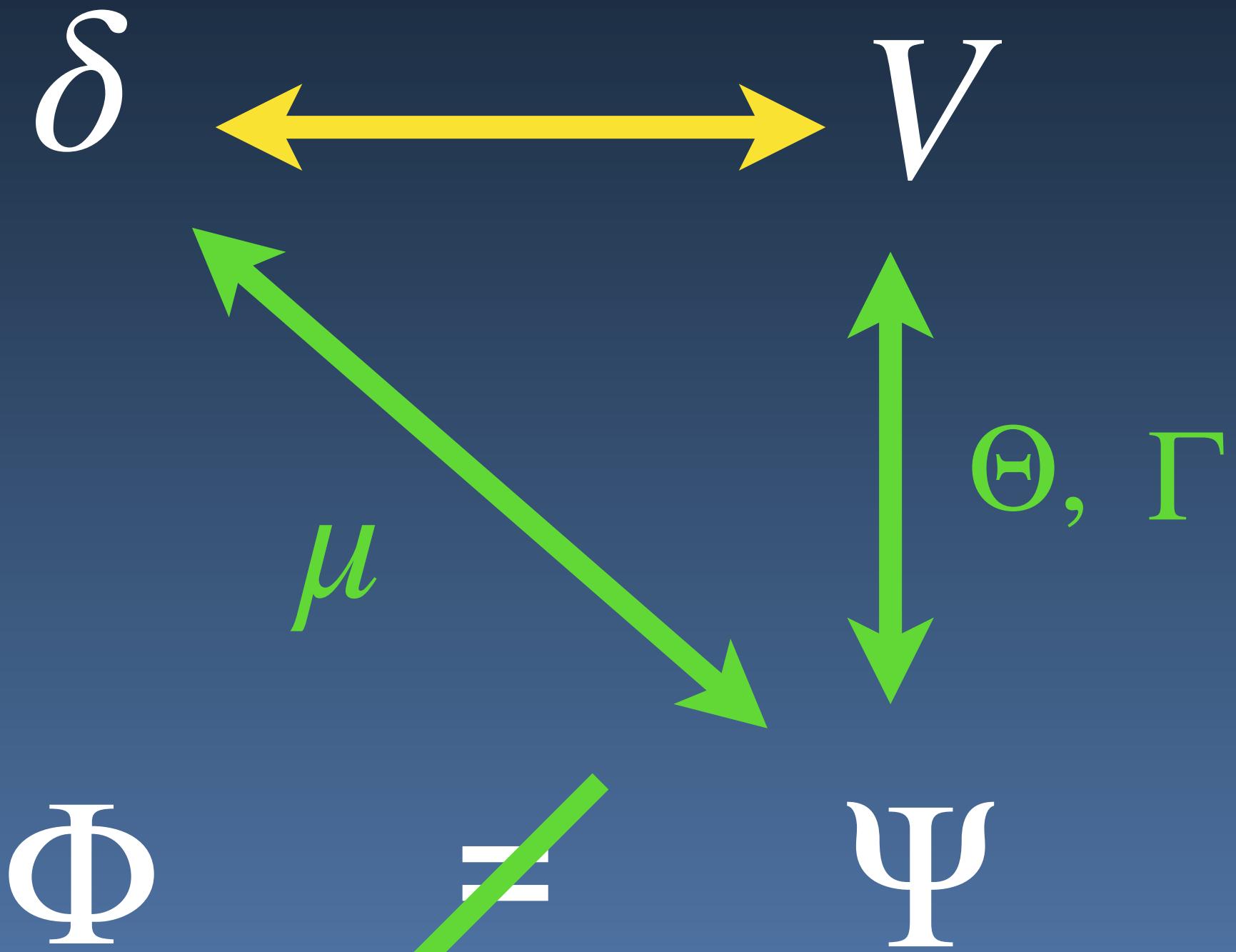


- Much smaller than RSD
- Observable by upcoming surveys
- Extracted by correlating two populations

Deus ex machina: gravitational redshift



SC, Grimm and Bonvin (2022)



Effective theory of interacting dark energy

SC, Mancarella et al. (2024)

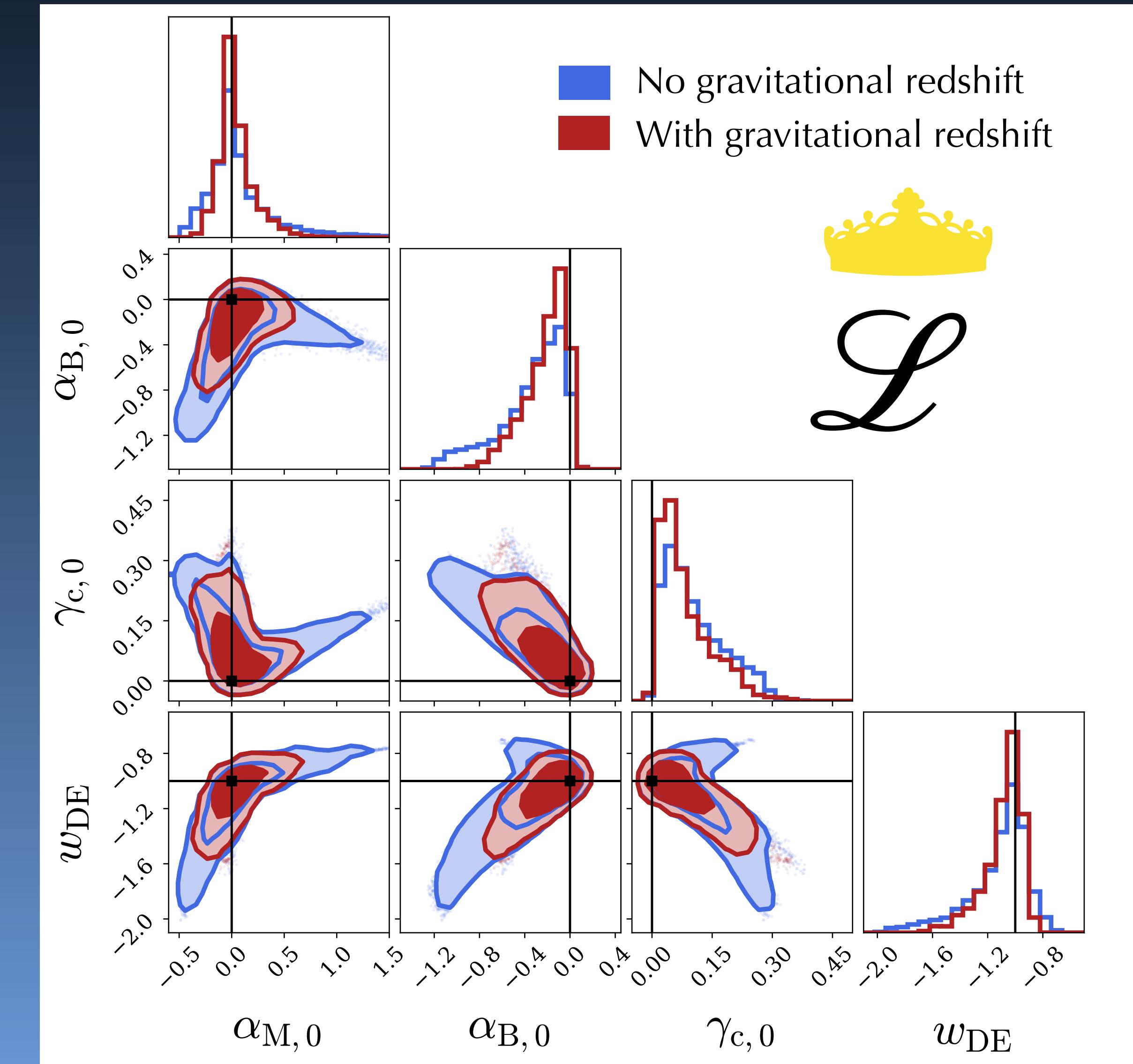
EF-TIGRE package



Gravity modifications
 α_M, α_B

WEP breaking
 γ_c

Equation of state of DE
 w_{DE}

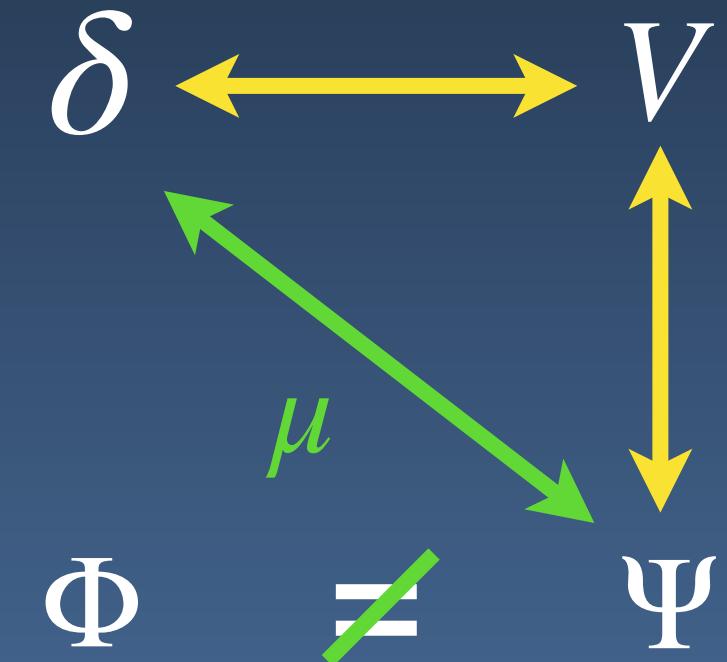


Two example models

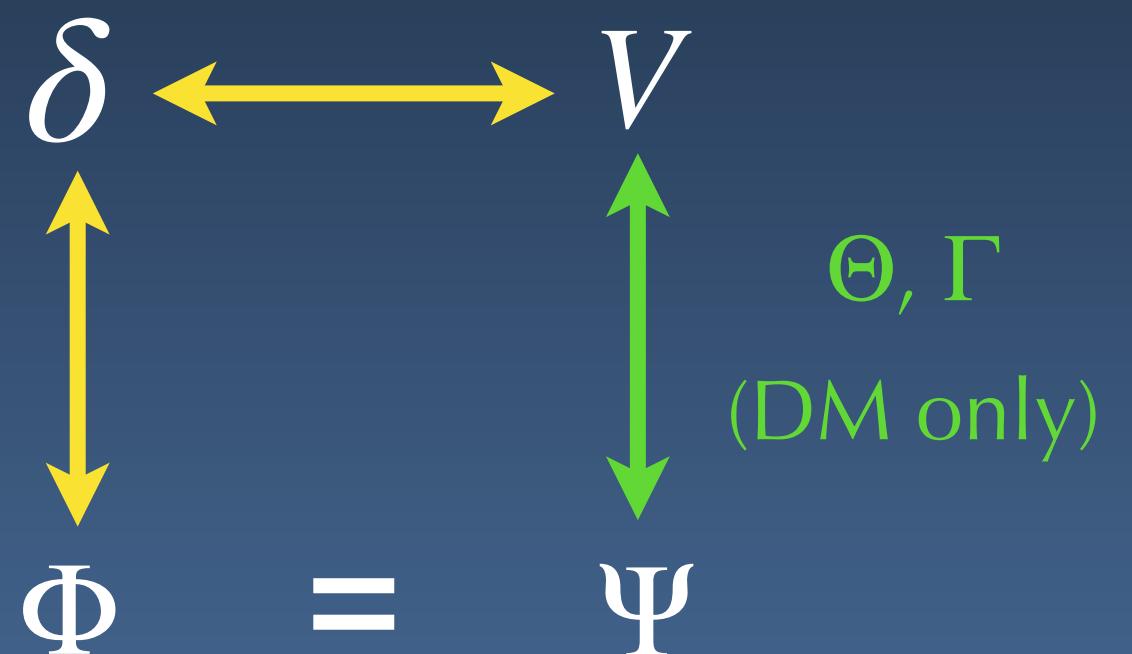
Bonvin & Pogosian (2022)

SC, Wang, Dam, Bonvin, Pogosian (2024)

Gravity modifications affecting
all constituents



Breaking of the WEP
for DM only



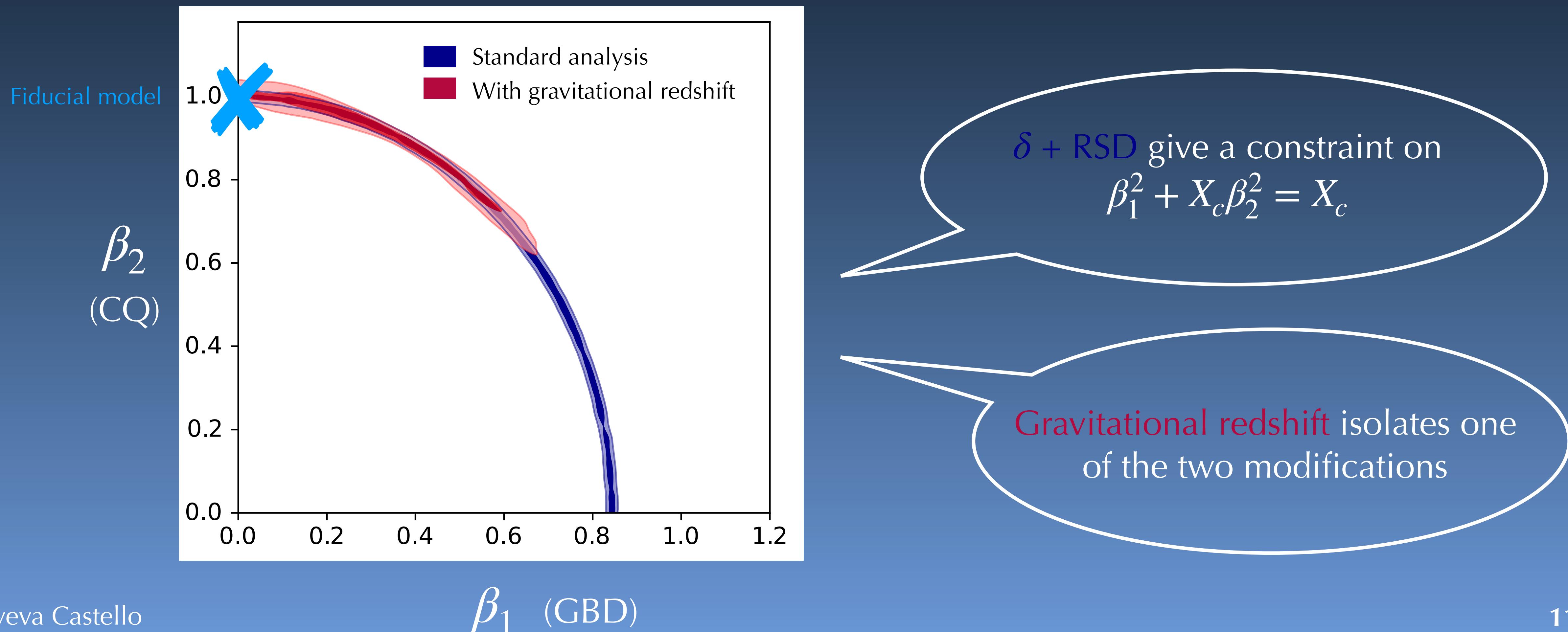
→ Generalised Brans-Dicke
Universal coupling β_1

→ Coupled quintessence
DM-only coupling β_2

Forecasts for SKA2

SC, Wang, Dam, Bonvin, Pogosian (2024)

- Generate mock data with one type of modification (e.g. $\beta_1 = 0, \beta_2 = 1$)
- Fit with both models (galaxy clustering + CMB + weak lensing)

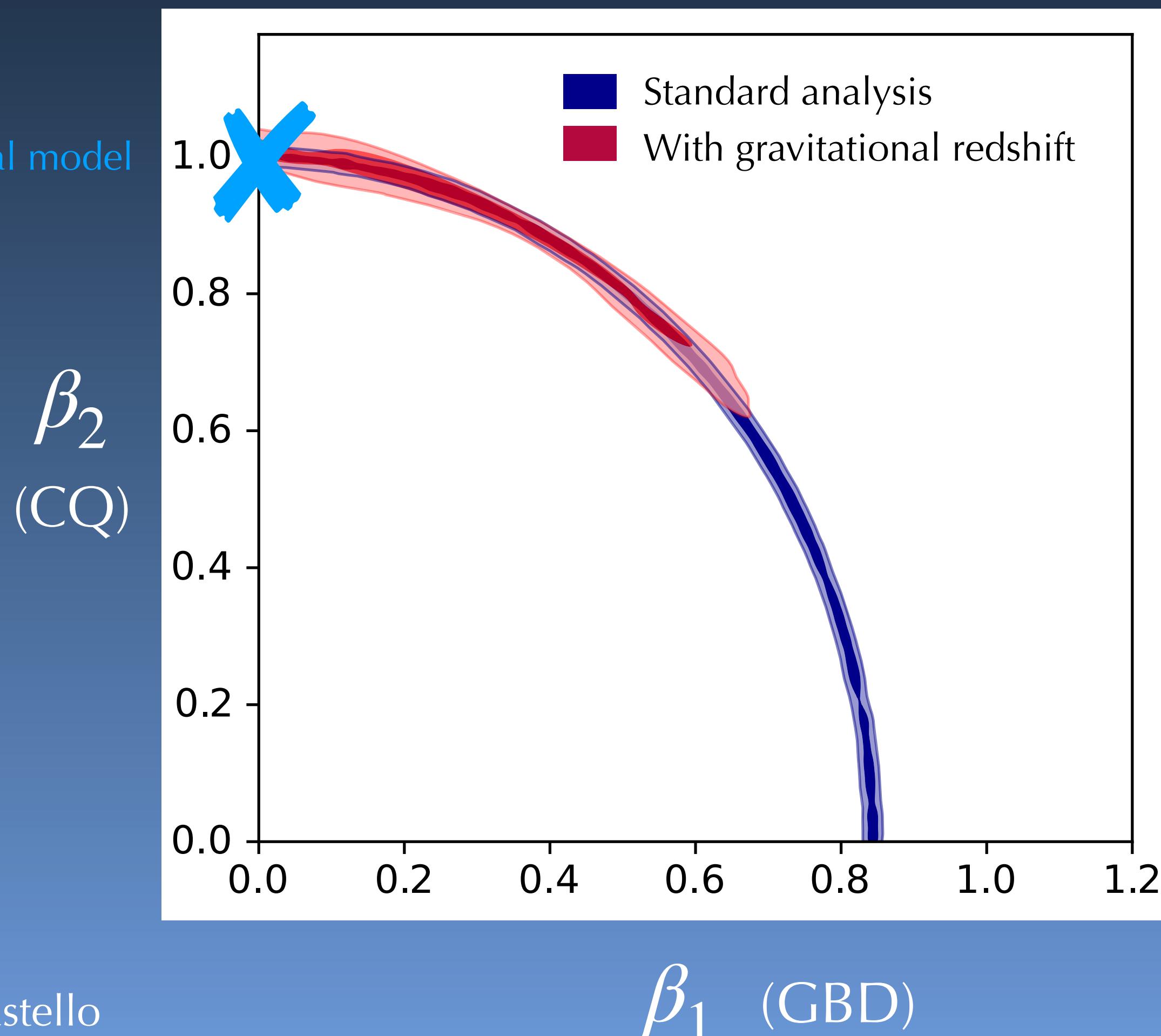


Forecasts for SKA2

SC, Wang, Dam, Bonvin, Pogosian (2024)

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Fiducial model

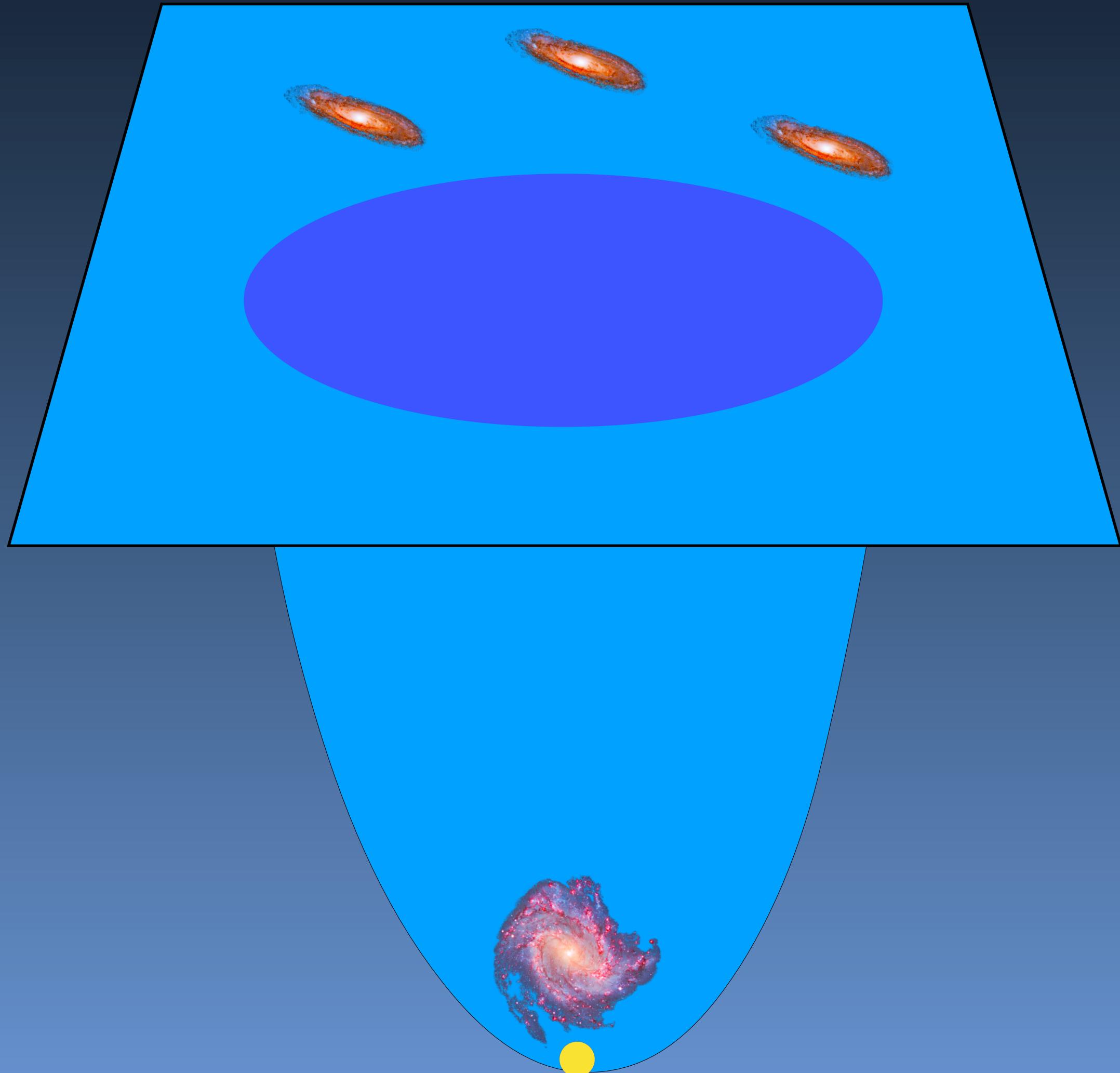


What is the threshold for gravitational redshift to help?

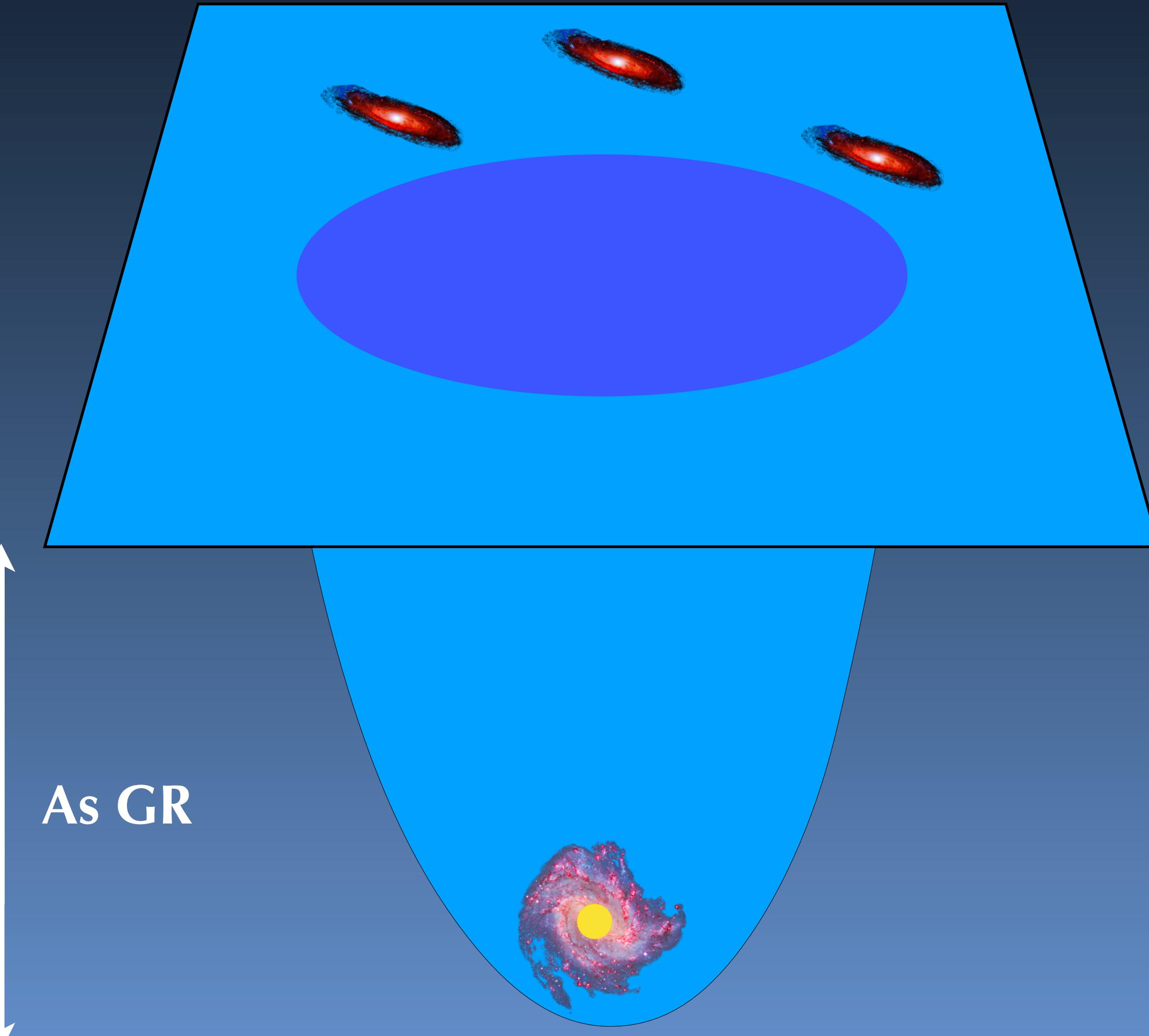
$$\beta_2 = 0.7 \text{ for } m = 0.1 \text{ Mpc}^{-1}$$
$$\beta_2 = 0.4 \text{ for } m = 0.01 \text{ Mpc}^{-1}$$

Some physical intuition

Modified gravity



Dark interactions



Some ongoing work...

Small scales: Gravitational redshift from galaxy clusters

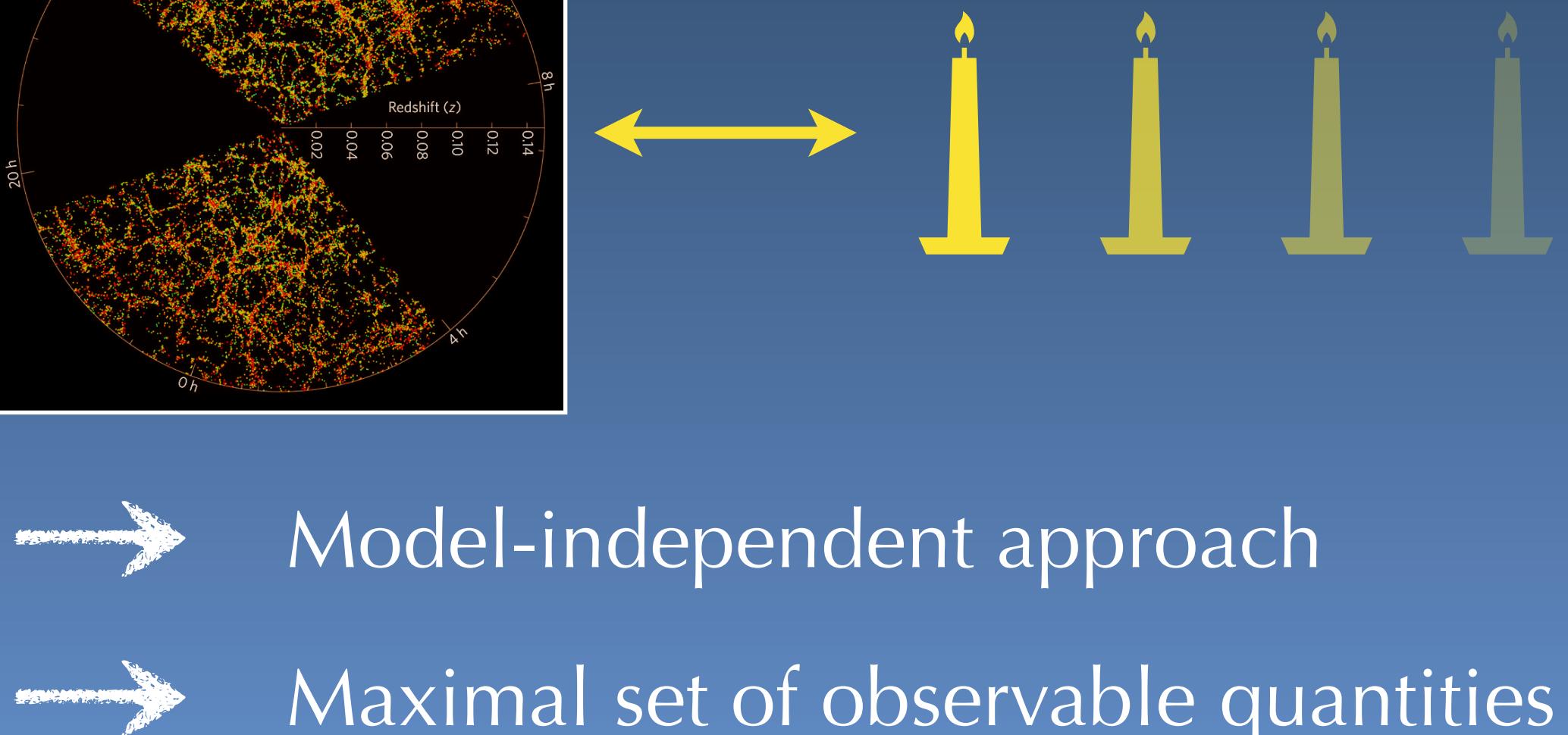
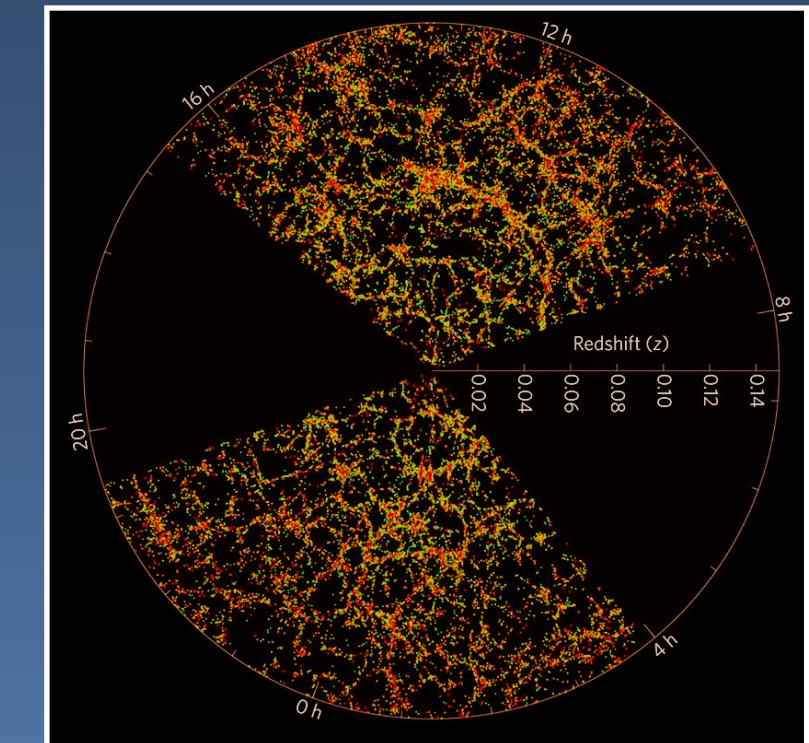
With C. Bonvin, Ø. Christiansen, E. Di Dio, D. Mota, H. Winther



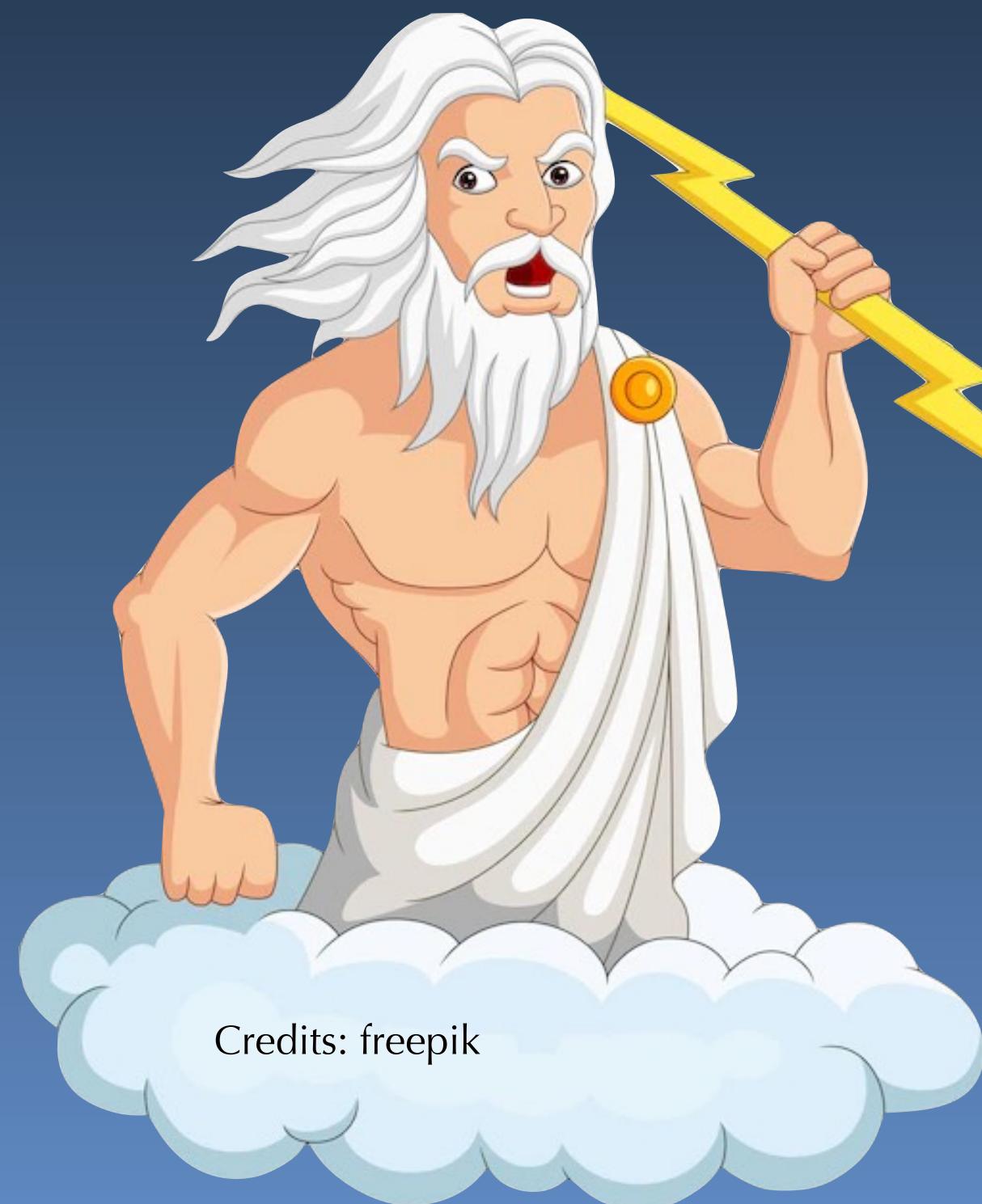
- Fully relativistic modelling of the signal
- Test of the equivalence principle

Cross-correlation: luminosity distance fluctuations

With L. Amendola, C. Bonvin, Z. Zheng



Take-home message



Credits: freepik

Gravitational redshift can break the degeneracy between modified gravity and a dark fifth force!



Happy to chat live or at sveva.castello@unige.ch :)