François Gelis

# Initial state and thermalization

Factorization Thermalization

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François Gelis IPhT, Saclay

# Outline

1	Factorization and Universality
2	Thermalization and Isotropization
3	Summary

## In collaboration with :

K. Dusling	(NCSU)
T. Epelbaum	(IPhT)
R. Venugopalan	(BNL)

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### Stages of a collision

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## Stages of a collision

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This talk : evolution up to times  $\sim 1~\text{fm/c}$ 

- i. Partonic content of high energy nuclei
- ii. Gluon production in the collision
- iii. Evolution shortly after the collision, Thermalization

## Color Glass Condensate = effective theory of small x gluons

[McLerran, Venugopalan (1994), Jalilian-Marian, Kovner, Leonidov, Weigert (1997), Iancu, Leonidov, McLerran (2001)]

The fast partons (k<sup>+</sup> > Λ<sup>+</sup>) are frozen by time dilation
 ▷ described as static color sources on the light-cone :

$$\mathbf{J}^{\mu} = \delta^{\mu +} \boldsymbol{\rho}(\mathbf{x}^{-}, \mathbf{\vec{x}}_{\perp}) \qquad (0 < \mathbf{x}^{-} < 1/\Lambda^{+})$$

- The color sources  $\rho$  are random, and described by a probability distribution  $W_{\Lambda^+}[\rho]$
- Slow partons (k<sup>+</sup> < Λ<sup>+</sup>) may evolve during the collision
   ▷ treated as standard gauge fields

 $\rhd$  eikonal coupling to the current  $J^{\mu}$  :  $J_{\mu}A^{\mu}$ 

$$\mathcal{S} = \underbrace{-\frac{1}{4}\int F_{\mu\nu}F^{\mu\nu}}_{\mathcal{S}_{YM}} + \int \underbrace{J^{\mu}A_{\mu}}_{\text{fast partons}}$$

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# Factorization and Universality

[FG, Lappi, Venugopalan (2008)]

#### Leading Order

 All observables can be expressed in terms of classical solutions of Yang-Mills equations :

$$[\mathcal{D}_{\mu}, \mathcal{F}^{\mu\nu}] = J^{\nu}$$

- Boundary conditions for inclusive observables : retarded, with  ${\cal A} \to 0$  at  $x_0 = -\infty$ 

Inclusive spectra at LO

$$\frac{dN_1}{d^3\vec{p}}\Big|_{LO} \sim \int d^4x d^4y \ e^{ip \cdot (x-y)} \Box_x \Box_y \ \mathcal{A}(x) \ \mathcal{A}(y)$$
$$\frac{dN_n}{d^3\vec{p}_1 \cdots d^3\vec{p}_n}\Big|_{LO} = \frac{dN_1}{d^3\vec{p}_1}\Big|_{LO} \cdots \frac{dN_1}{d^3\vec{p}_n}\Big|_{LO}$$

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#### Next to Leading Order



- Valid for all inclusive observables, e.g. the energy-momentum tensor
- In the CGC, upper cutoff on the loop momentum : k<sup>±</sup> < Λ, to avoid double counting with the sources J<sub>1,2</sub>
   ▷ large logarithms of the cutoff

#### **Initial state logarithms**

#### **Central result**

$$\frac{1}{2} \iint_{\mathbf{u},\mathbf{v}} \left[ \mathbf{a}_{\mathbf{k}} \mathbb{T} \right]_{\mathbf{u}} \left[ \mathbf{a}_{\mathbf{k}}^{*} \mathbb{T} \right]_{\mathbf{v}} + \int_{\mathbf{u}} \left[ \boldsymbol{\alpha} \mathbb{T} \right]_{\mathbf{u}} = \\ = \log \left( \Lambda^{+} \right) \, \mathcal{H}_{1} + \log \left( \Lambda^{-} \right) \, \mathcal{H}_{2} + \text{terms w/o logs}$$

 $\mathfrak{H}_{1,2} = \mathsf{JIMWLK}$  Hamiltonians of the two nuclei

- No mixing between the logs of the two nuclei
- Since the LO⇔NLO relationship is the same for all inclusive observables, these logs have a universal structure

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#### **Factorization of the logarithms**

• The JIMWLK Hamiltonian  ${\mathcal H}$  is self-adjoint :

$$\int [\mathsf{D}\rho] W (\mathcal{H} \mathcal{O}) = \int [\mathsf{D}\rho] (\mathcal{H} W) \mathcal{O}$$

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nclusive observables at Leading Log accuracy  

$$\langle 0 \rangle_{_{\text{Leading Log}}} = \int \left[ D\rho_1 D\rho_2 \right] W_1 \left[ \rho_1 \right] W_2 \left[ \rho_2 \right] \underbrace{\mathcal{O}_{_{\text{LO}}}[\rho_1, \rho_2]}_{\text{fixed } \rho_{1,2}}$$

Logs absorbed into the scale evolution of W<sub>1,2</sub>

$$\Lambda \frac{\partial W}{\partial \Lambda} = \mathcal{H} W \qquad \text{(JIMWLK equation)}$$

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# Thermalization and Isotropization

[Dusling, Epelbaum, FG, Venugopalan (2010-12)] [Dusling, FG, Venugopalan (2011)] [Epelbaum, FG (2011)]

# Energy momentum tensor at LO

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## Energy momentum tensor at LO

# Factorization $\Omega^{-1}$ Thermalization $\mathsf{T}^{\mu\nu}$ for longitudinal $\vec{E}$ and $\vec{B}$ $T_{LO}^{\mu\nu}(\tau = 0^+) = diag(\varepsilon, \varepsilon, \varepsilon, -\varepsilon)$ ▷ far from ideal hydrodynamics

#### Weibel instabilities for small perturbations

[Mrowczynski (1988), Romatschke, Strickland (2003), Arnold, Lenaghan, Moore (2003), Rebhan, Romatschke, Strickland (2005), Arnold, Lenaghan, Moore, Yaffe (2005), Romatschke, Rebhan (2006), Bodeker, Rummukainen (2007),...]



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#### Weibel instabilities for small perturbations

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- Some of the field fluctuations  $\alpha_k$  diverge like  $exp\,\sqrt{\mu\tau}$  when  $\tau\to+\infty$
- Some components of  $T^{\mu\nu}$  have secular divergences when evaluated at fixed loop order
- When  $\alpha_{\bf k}\sim {\cal A}\sim g^{-1},$  the power counting breaks down and additional contributions must be resummed :

$$g e^{\sqrt{\mu \tau}} \sim 1$$
 at  $\tau_{max} \sim \mu^{-1} \log^2(g^{-1})$ 

$$1e-13 \underbrace{\textbf{E}}_{0} \underbrace{1}_{500} \underbrace{1}_{1000} \underbrace{1}_{500} \underbrace{2000}_{2500} \underbrace{2500}_{3000} \underbrace{3000}_{3500} \underbrace{3}_{9}^{2} \mu \tau$$

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## Improved power counting



## Improved power counting



#### Improved power counting



#### Leading terms at $\tau_{\text{max}}$

All disjoint loops to all orders

> exponentiation of the 1-loop result

# **Resummation of the leading secular terms**

$$\mathsf{T}^{\mu\nu}_{\text{resummed}} = \exp\left[\frac{1}{2} \int\limits_{\mathfrak{u}, \nu \in \Sigma} \underbrace{\int_{k} [\mathfrak{a}_{k} \mathbb{T}]_{\mathfrak{u}} [\mathfrak{a}_{k}^{*} \mathbb{T}]_{\nu}}_{\mathfrak{G}(\mathfrak{u}, \nu)} + \int\limits_{\mathfrak{u} \in \Sigma} [\mathfrak{a} \mathbb{T}]_{\mathfrak{u}}\right] \mathsf{T}^{\mu\nu}_{\text{LO}}[\mathcal{A}_{\text{init}}]$$

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#### **Resummation of the leading secular terms**

$$\begin{aligned} \mathbf{T}_{\text{resummed}}^{\mu\nu} &= \exp\left[\frac{1}{2} \int_{\mathbf{u},\nu\in\Sigma} \underbrace{\int_{\mathbf{k}} [\mathbf{a}_{\mathbf{k}}\mathbb{T}]_{\mathbf{u}} [\mathbf{a}_{\mathbf{k}}^{*}\mathbb{T}]_{\nu}}_{g(\mathbf{u},\nu)} + \int_{\mathbf{u}\in\Sigma} [\mathbf{\alpha}\mathbb{T}]_{\mathbf{u}} \right] \mathbf{T}_{\text{Lo}}^{\mu\nu}[\mathcal{A}_{\text{init}}] \\ &= \int [D\chi] \exp\left[-\frac{1}{2} \int_{\mathbf{u},\nu\in\Sigma} \chi(\mathbf{u}) \mathcal{G}^{-1}(\mathbf{u},\nu) \chi(\nu)\right] \mathbf{T}_{\text{Lo}}^{\mu\nu}[\mathcal{A}_{\text{init}} + \chi + \alpha] \end{aligned}$$

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- The evolution remains classical, but we must average over a Gaussian ensemble of initial conditions
- The shift  $\alpha$  can be absorbed into a redefinition of  $\mathcal{A}_{\text{init}}$
- The 2-point correlation  $\mathfrak{G}(u,\nu)$  of the fluctuations is known analytically

#### Analogous scalar toy model

 $\phi^4$  field theory coupled to a source

$$\mathcal{L} = \frac{1}{2} (\partial_{\alpha} \phi)^2 - \frac{g^2}{4!} \phi^4 + J \phi$$

Strong external source: 
$$J \propto \frac{Q^3}{q}$$

- In 3+1-dim, g is dimensionless, and the only scale in the problem is Q
- This theory has unstable modes (parametric resonance)
- Two setups have been studied :
  - Fixed volume system
  - Longitudinally expanding system

#### Pathologies in fixed order calculations





Oscillating pressure at LO : no equation of state

## Pathologies in fixed order calculations



- Oscillating pressure at LO : no equation of state
- Small NLO correction to the energy density (protected by energy conservation)
- Secular divergence in the NLO correction to the pressure

#### **Resummed energy momentum tensor**



- · No secular divergence in the resummed pressure
- The pressure relaxes to the equilibrium equation of state

## Time evolution of the occupation number



- Resonant peak at early times
- Turbulent Kolmogorov spectrum in the intermediate k-range?
- · Late times : classical equilibrium with a chemical potential
- $\mu \approx m$  + excess at k = 0 : Bose-Einstein condensation?

#### **Bose-Einstein condensation**



- Start with the same energy density, but an empty zero mode
- · Very quickly, the zero mode becomes highly occupied
- · Same distribution as before at late times

#### **Evolution of the condensate**



- Formation time almost independent of the coupling
- Condensate lifetime much longer than its formation time
- Smaller amplitude and faster decay at large coupling

#### Longitudinal expansion

• The EoM is singular when  $\tau \rightarrow 0$  : one must start at  $\tau_0 \neq 0$ 

$$\partial_{\tau}^2 \phi + \frac{1}{\tau} \partial_{\tau} \phi - \frac{1}{\tau^2} \partial_{\eta}^2 \phi - \nabla_{\perp}^2 \phi + \frac{g^2}{6} \phi^3 = 0$$

• The spectrum of fluctuations of the initial field also depends on  $\tau_0$ , precisely in such a way that the end result is independent of  $\tau_0$ 

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 $\tau = 0.1$  [40 × 40 × 320] 104  $10^{2}$  $10^{6}$  $10^{4}$  $10^{0}$  $10^{2}$  $10^{0}$ 10-2 50 100 200 k.

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- Note :  $\nu$  is the Fourier conjugate of the rapidity  $\eta$
- Initially, only the v = 0 modes are occupied

 $\tau = 10$  [40 × 40 × 320] 104  $10^{2}$  $10^{6}$  $10^{4}$  $10^{0}$  $10^{2}$  $10^{0}$  $10^{-2}$ 50 100 200 k.

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- Note :  $\nu$  is the Fourier conjugate of the rapidity  $\eta$
- Initially, only the v = 0 modes are occupied

 $\tau = 50$  [40 × 40 × 320]  $10^{4}$  $10^{2}$  $10^{6}$  $10^{4}$  $10^{0}$  $10^{2}$  $10^{0}$  $10^{-2}$ 100 200 k.

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- Note :  $\nu$  is the Fourier conjugate of the rapidity  $\eta$
- Initially, only the  $\nu=0$  modes are occupied
- Rapid expansion of the distribution in  $\nu$   $(\nu_{max} \sim \tau^{2/3}$  for a thermalized system)

 $\tau = 100$  [40 × 40 × 320]  $10^{4}$  $10^{2}$  $10^{6}$  $10^{4}$  $10^{0}$  $10^{2}$  $10^{0}$  $10^{-2}$ 100-150-200 k.

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- Note :  $\nu$  is the Fourier conjugate of the rapidity  $\eta$
- Initially, only the  $\nu=0$  modes are occupied
- Rapid expansion of the distribution in  $\nu$   $(\nu_{max} \sim \tau^{2/3}$  for a thermalized system)

 $\tau = 150$  [40 × 40 × 320] Factorization hermalization  $10^{4}$  $10^{2}$  $10^{6}$  $10^{4}$  $10^{0}$  $10^{2}$  $10^{0}$  $10^{-2}$ 100 200 k.

- Note :  $\nu$  is the Fourier conjugate of the rapidity  $\eta$
- Initially, only the  $\nu = 0$  modes are occupied
- Rapid expansion of the distribution in  $\nu$   $(\nu_{max} \sim \tau^{2/3}$  for a thermalized system)

 $\tau = 200$  [40 × 40 × 320]  $10^{4}$  $10^{2}$  $10^{6}$  $10^{4}$  $10^{0}$  $10^{2}$  $10^{0}$  $10^{-2}$ 100 200 k.

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#### Longitudinal expansion : equation of state



- After a short time, one has  $2P_{_{T}} + P_{_{L}} \approx \varepsilon$
- Change of behavior of the energy density:  $\tau^{-1} \rightarrow \tau^{-4/3}$ . Since one has  $\partial_{\tau} \varepsilon + (\varepsilon + P_{_L})/\tau = 0$ , this suggests that  $P_{_L}$  gets close to  $\varepsilon/3$

## Longitudinal expansion : isotropization



- At early times,  $P_{\rm L}$  drops much faster than  $P_{\rm T}$  (redshifting of the longitudinal momenta due to the expansion)
- Drastic change of behavior when the expansion rate becomes smaller than the growth rate of the unstability
- Eventually, isotropic pressure tensor :  $\mathrm{P_{L}} \approx \mathrm{P_{T}}$

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# Summary and Outlook

#### Summary

- Factorization of high energy logarithms in AA collisions
  - universal distributions, also applicable in pA or DIS
  - · controls the rapidity dependence of correlations
  - limited to inclusive observables

#### • Resummation of secular terms in the final state evolution

- stabilizes the NLO calculation
- · leads to the equilibrium equation of state
- isotropization even with longitudinal expansion
- full thermalization on longer time-scales
- · Bose-Einstein condensation if overoccupied initial state

(so far, all numerical studies done for a toy scalar model)

#### What's next? QCD

- approach straightforwardly generalizable to QCD
- seamless integration with the CGC description of AA collisions
- gauge invariant
- computationally expensive ( ~ [scalar case]  $\times 3(N_c^2 1)$ )

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# **Extra bits**

#### Lattice spacing dependence



- When  $\nu_{max}\sim \tau^{2/3}\gtrsim 2N,$  the longitudinal modes are artificially cut-off, and the longitudinal pressure decreases
- If N increases, the field acquires a mass  $m^2 \sim g^2 \log(N)$ Taking the limit  $N \to \infty$  requires a proper renormalization of the bare parameters of the theory

# **Dense-dilute collisions**

# receeee accedence 00000 g<sup>-4</sup> g-2 1 $g^2$ $g^4$ $\rho_1$ g<sup>-1</sup> AA g рA

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#### **Dense-dilute collisions**



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# **Exclusive processes**

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#### **Exclusive processes**

Example : differential probability to produce 1 particle at LO

$$\frac{\mathrm{d} P_1}{\mathrm{d}^3 \vec{\mathbf{p}}} \bigg|_{_{\mathrm{LO}}} = \mathbb{F}[\mathbf{0}] \times \int \mathrm{d}^4 x \mathrm{d}^4 y \, e^{\mathrm{i} \mathbf{p} \cdot (x-y)} \Box_x \Box_y \mathcal{A}_+(x) \mathcal{A}_-(y) \bigg|_{_{z=0}}$$

- The vacuum-vacuum graphs do not cancel in exclusive quantities : F[0] ≠ 1 (in fact, F[0] = exp(-c/g<sup>2</sup>) ≪ 1)
- A<sub>+</sub> and A<sub>-</sub> are classical solutions of the Yang-Mills equations, but with non-retarded boundary conditions



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# **Thermalization in Yang-Mills theory**

• Recent analytical work : Kurkela, Moore (2011)

- Going from scalars to gauge fields :
  - More fields per site (3 Lorentz components × 8 colors)
  - · More complicated spectrum of initial conditions



# **BEC and dilepton production**

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#### **BEC and dilepton production**

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 $\rhd$  excess of dileptons with  $k_\perp \ll M_{inv}$ 

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