

Simultaneous Searches for
WIMP Dark Matter and $0\nu\beta\beta$ decay
at the ton-scale
with a high-pressure xenon gas TPC

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For the NEXT collaboration

See also talk on NEXT by David Lorca Galindo

Xenon:
a popular choice for both
 $0\nu\beta\beta$ and WIMP search

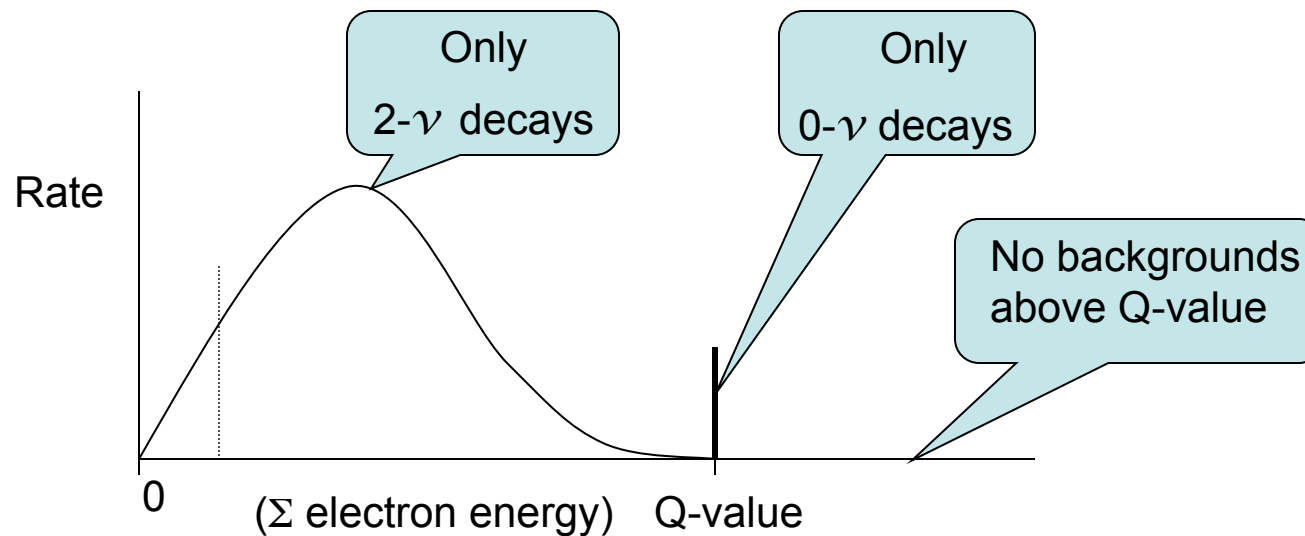
But not in the same detector!

Xenon: a popular choice for both $0\nu\beta\beta$ and WIMP search

- Both searches require very low backgrounds
- Simultaneous searches could save \$\$\$,...
- I present a case that both could be done, with little or no compromise, in a xenon gas electroluminescent TPC, perhaps at ton-scale

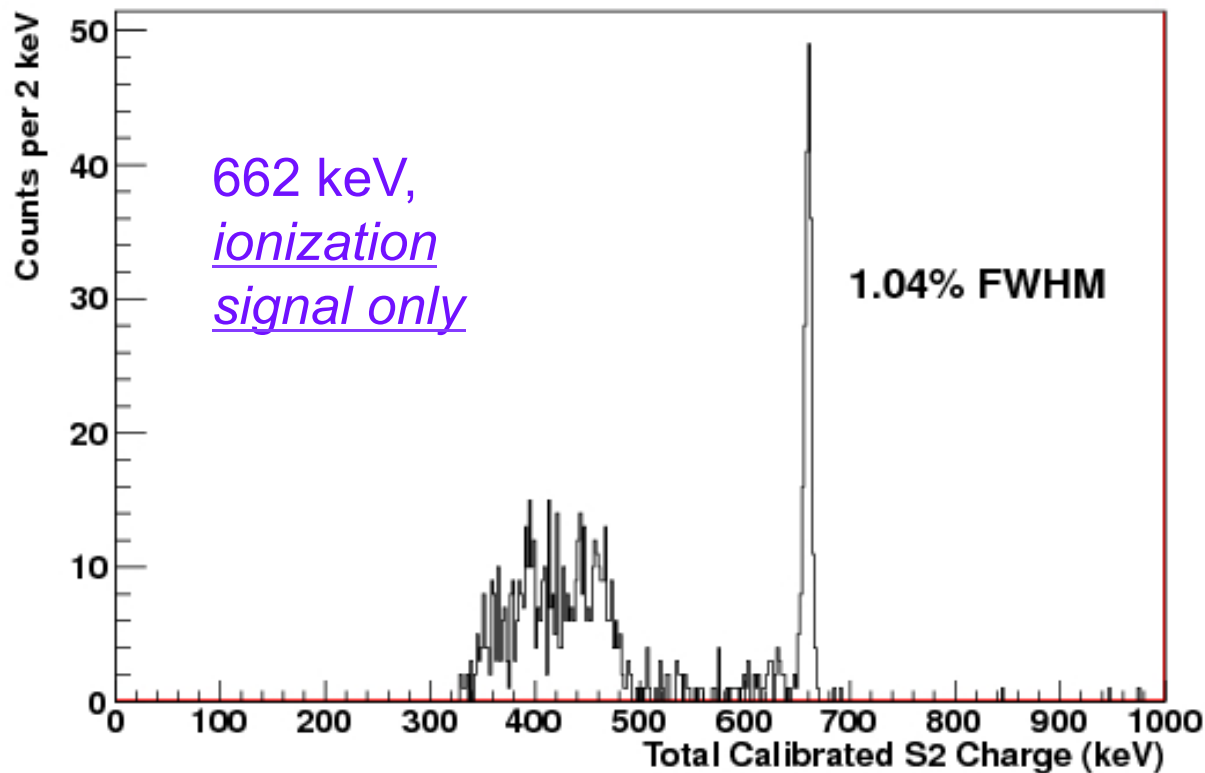
0- ν $\beta\beta$: Energy resolution is important!

Ideal case: 0- ν signal appears as a narrow peak



$\delta E/E < 1\%$ FWHM is needed for separation from 2- ν background and to avoid nearby γ -ray lines such as from ^{214}Bi

New: 1% FWHM energy resolution for ^{137}Cs 662 keV γ -rays in xenon!



Data from
LBNL-TAMU
HP Xe TPC

This result is
important for
both $0\nu\beta\beta$ &
WIMP searches

Resolution: ^{137}Cs γ -ray (662 keV)

$$\sigma_N = (F \cdot N)^{1/2}$$

F \equiv Fano factor: F = 0.15 (xenon gas)

$$N = Q/w = 662,000/25 \sim 26,500 \text{ primary electrons}$$

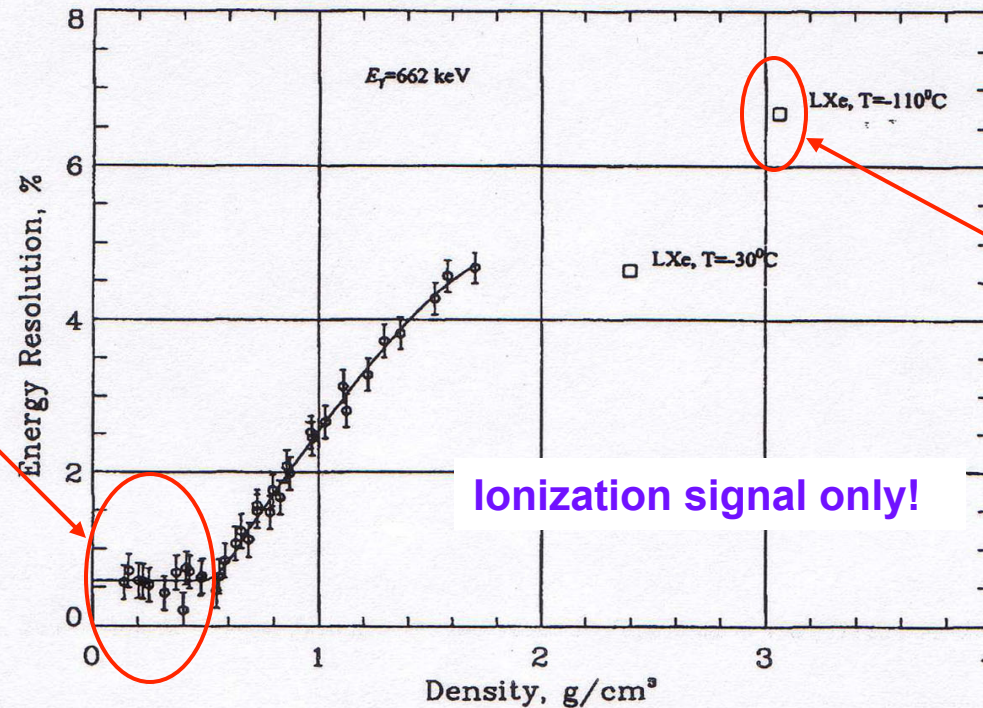
$$\sigma_N = 63, \text{ a } \underline{\text{very small number}} \text{ of electrons!}$$

$$\delta E/E = 2.35 \cdot \sigma_N/N = \underline{0.6\%} \text{ FWHM (intrinsic)}$$

Our 1.04% FWHM result is ~ 1.6 x intrinsic resolution
How does gas deliver such good performance?

Energy resolution in Xenon depends strongly on density

A. Bolotnikov, B. Ramsey / Nucl. Instr. and Meth. in Phys. Res. A 396 (1997) 360–370



Here, the fluctuations are normal

$$F = 0.15$$

Unfolded resolution:

$$\delta E/E \sim 0.6\%$$

FWHM

Very large fluctuations between light/charge!
 F ~ 20 !!
WIMPs:
S2/S1 suffers!

density dependencies of the intrinsic energy resolution (%FWHM) measured for 662 keV gamma-rays.

For $\rho < 0.55 \text{ g/cm}^3$, ionization energy resolution is “intrinsic”

Energy Partitioning in LXe

Anomalously large fluctuations in energy partition between **ionization** and **scintillation** appear in LXe, and generate a **Fano factor $F \sim 20$**

The large fluctuations in LXe are caused by ***delta-rays***, few in number, but with “Landau” fluctuations toward high local ionization density

A conduction band exists in LXe, **promoting** high recombination in regions of high ionization density – delta rays

The recombination process **amplifies** the non-Poisson statistics of the energy loss process of electrons in LXe, leading to large fluctuations

But not for xenon gas!

Energy resolution at $Q_{\beta\beta} = 2457$ keV

$$\delta E/E = 2.35 \cdot (F \cdot W/Q)^{1/2}$$

- $F \equiv$ Fano factor (HPXe) : $F = 0.15$
 - $w \equiv$ Average energy per ion pair: $w \sim 25$ eV
 - $Q \equiv$ Energy deposited from $^{136}\text{Xe} \rightarrow ^{136}\text{Ba}$:
- $N = Q/w \sim 100,000$ primary electrons
- $\sigma_N = (F \cdot N)^{1/2} \sim 124$ electrons rms!

$\delta E/E = 0.28\%$ FWHM intrinsic HPXe

Only about x3 worse than Ge diodes!

Energy resolution at $Q_{\beta\beta} = 2457$ keV

$\delta E/E = 0.28\%$ FWHM intrinsic HPXe!

**How can this performance be preserved
through the detection process?**

**Let “G” represent noise/fluctuations in EL gain
Uncorrelated fluctuations can add in quadrature**

Gain, noise & resolution

$$\sigma_n = ((F + G) \cdot N)^{1/2}$$

Require that $G \leq F = 0.15$

Only electroluminescence can provide this performance

EL: $G = J_{CP}/N_{UV} + (1 + \sigma_{PMT}^2)/N_{pe}$

N_{pe} = number of photo-electrons per primary electron

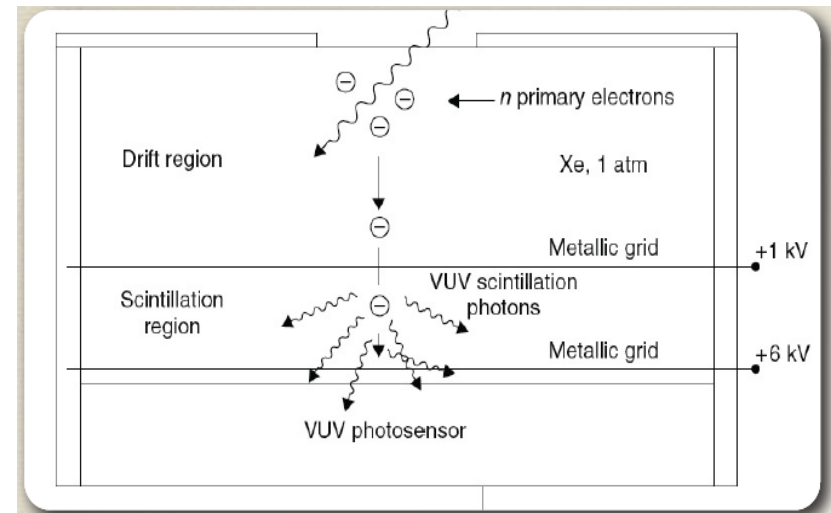
$$\sigma_{PMT}^2 \approx 2 \text{ (due to after-pulsing !!)}$$

$$G \approx 3/N_{pe}$$

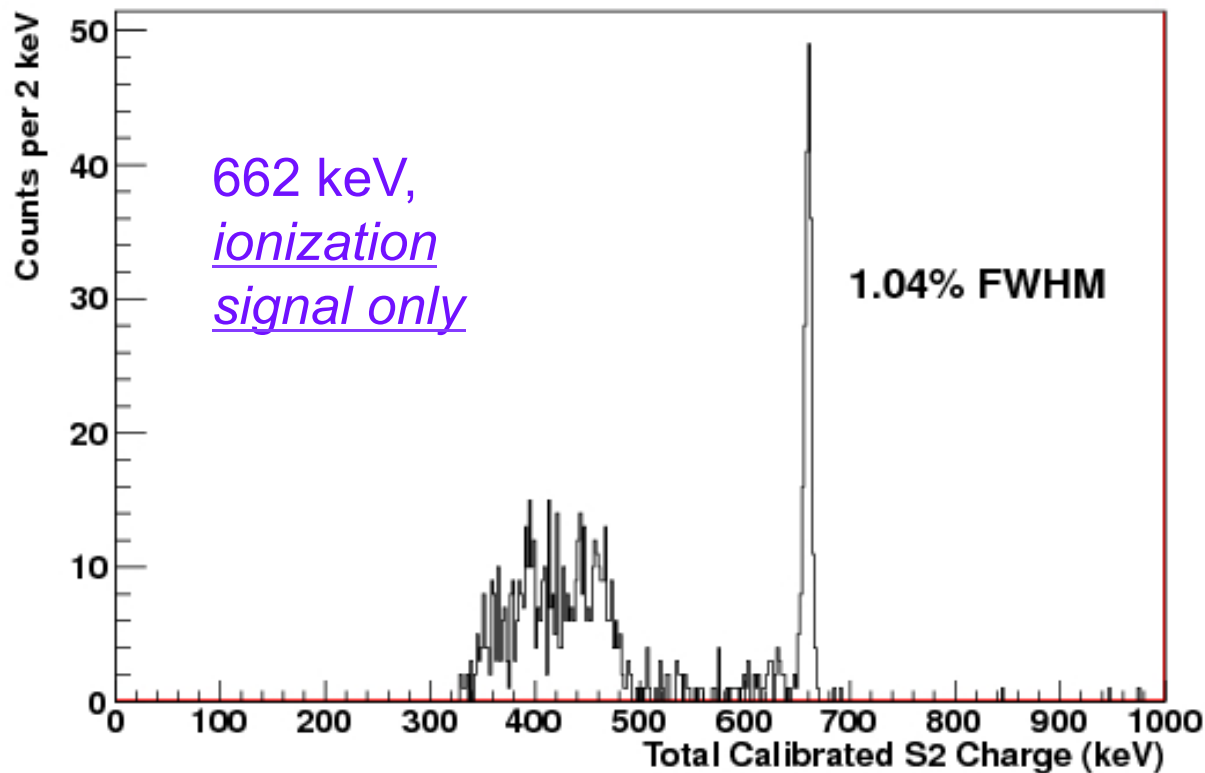
$$\Rightarrow N_{pe} > 20 \text{ per electron so that } G \leq F = 0.15$$

Electro-Luminescence (EL) is the key (aka: Gas Proportional Scintillation)

- Physics process generates ionization signal
- Electrons drift in low electric field region
- Electrons enter a high electric field region
- Electrons gain energy, excite xenon: 8.32 eV
- Xenon radiates VUV (≈ 175 nm, 7.5 eV)
- Electron starts over, gaining energy again
- Linear growth of signal with voltage
- Photon generation up to $>1000/e$, but no ionization
- **Sequential** gain; no exponential growth \Rightarrow fluctuations are very small
- $\delta N_{UV} = (J_{CP} \cdot N_{UV})^{1/2}$ (Poisson: $J_{CP} = 1$)
- Optimal EL conditions: $J_{CP} = 0.01$



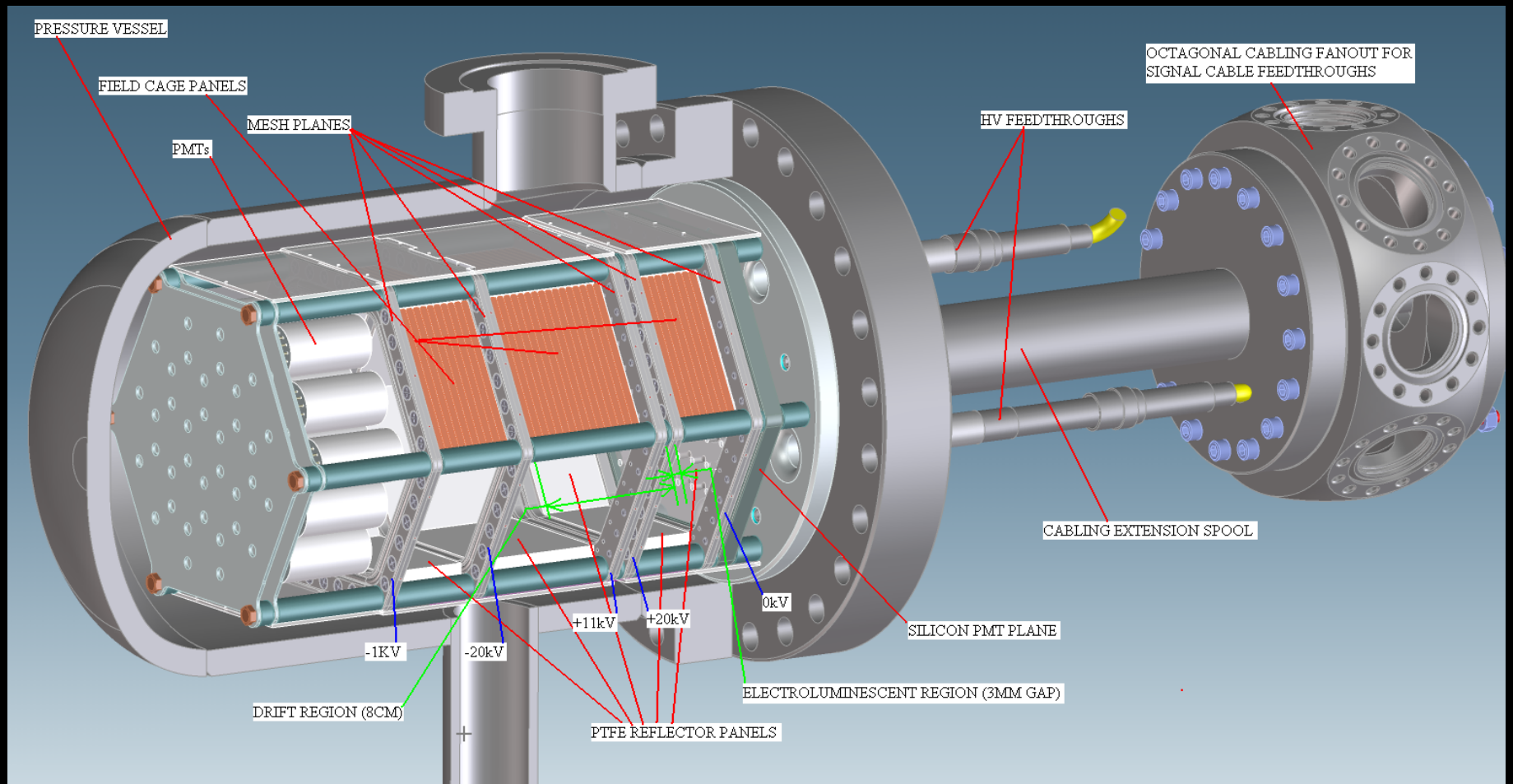
How was this 1% result obtained?



Data from
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HP Xe TPC

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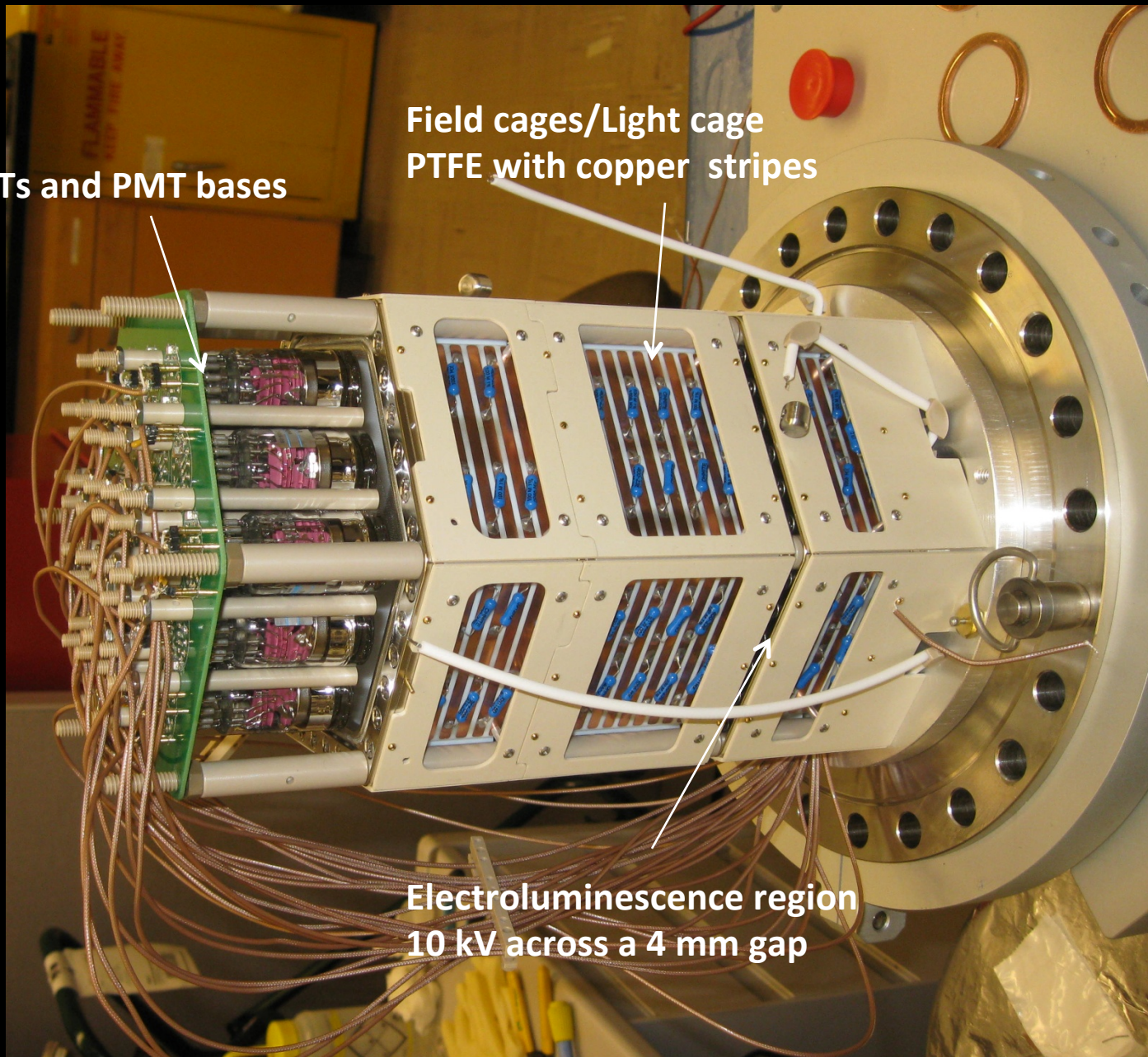
LBNL-TAMU TPC Prototype



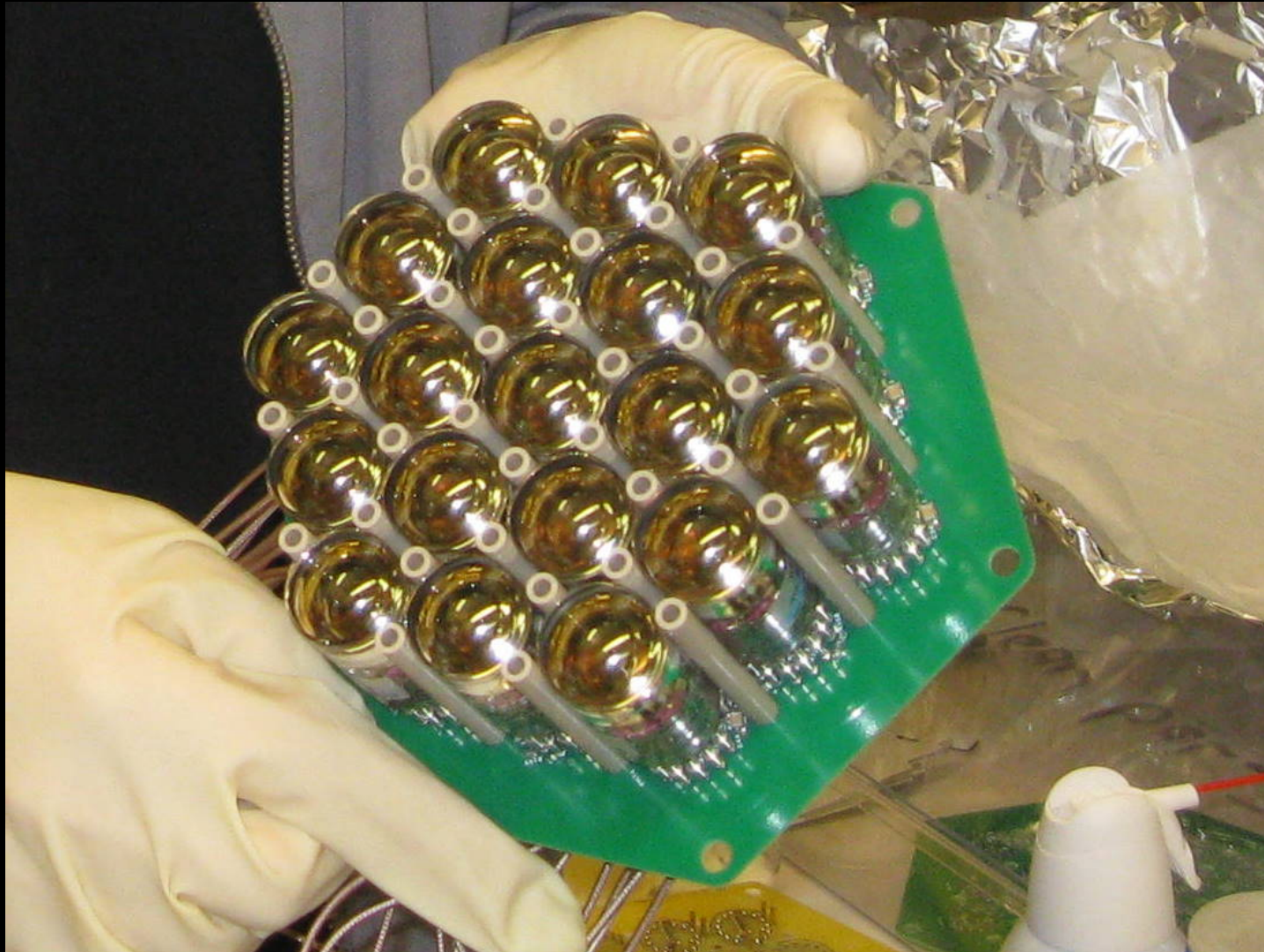
19 PMTs and PMT bases

Field cages/Light cage
PTFE with copper stripes

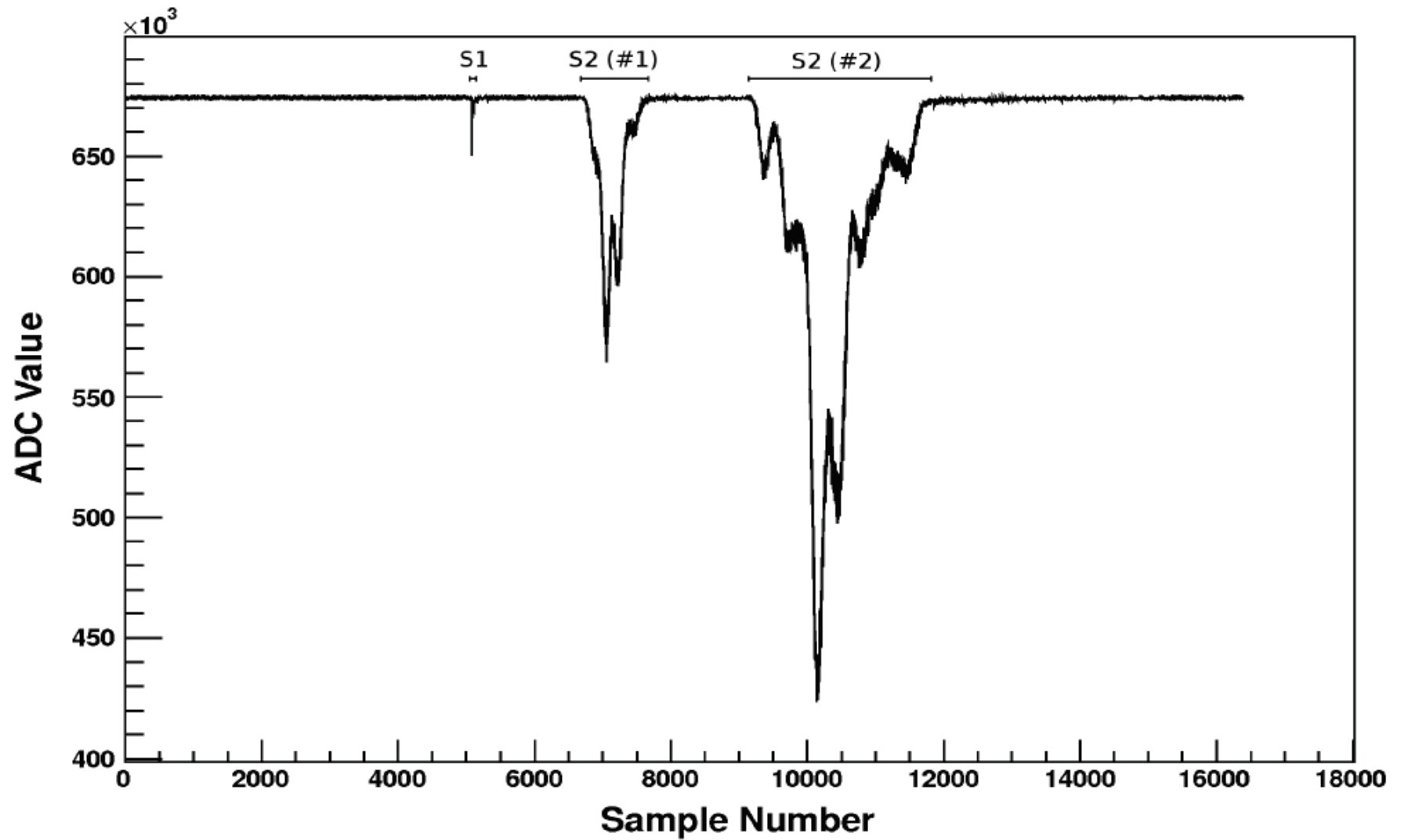
Electroluminescence region
10 kV across a 4 mm gap



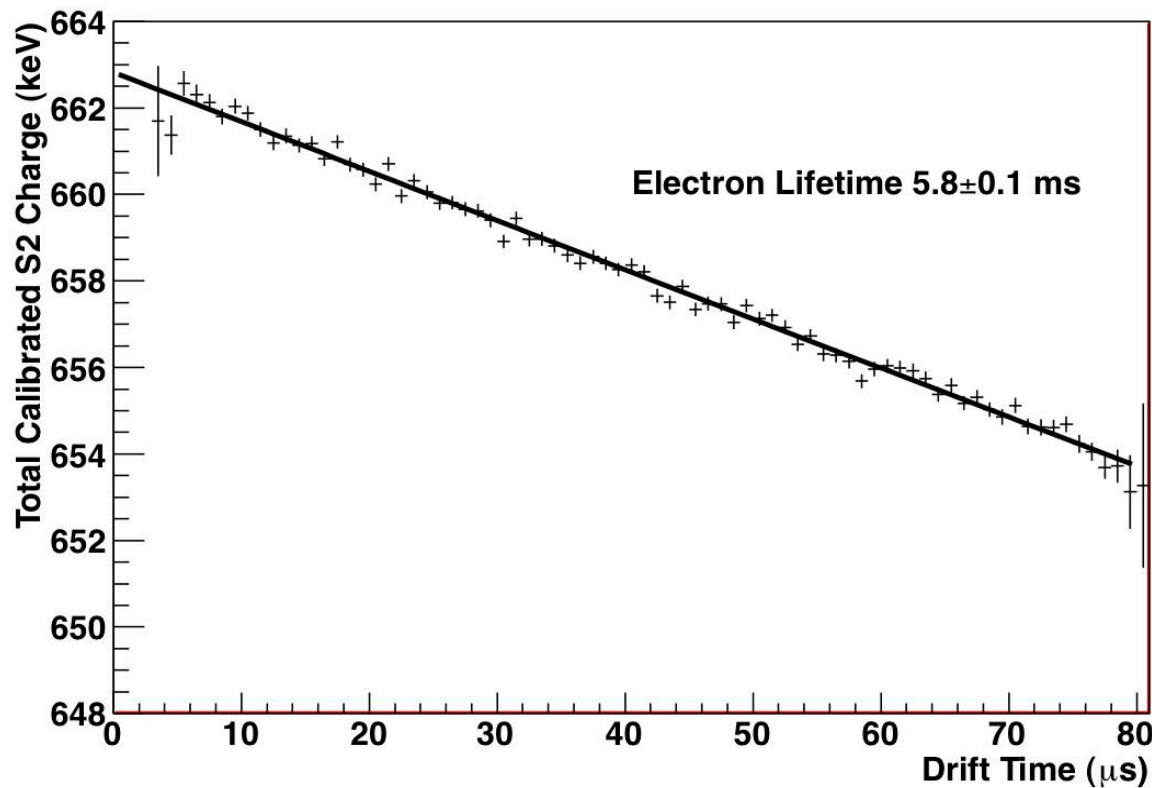
PMT Array: inside the pressure vessel Quartz window 2.54 cm diameter PMTs



Complex topologies are common!



Attenuation of electrons during drift is very small



← correction for attenuation is modest, but must be done with care for minimal error to energy

“*Neutrino Experiment Xenon TPC*”

NEXT

Is based on high-pressure xenon gas (HPXe) TPC ideas

Is optimized for $0\nu\beta\beta$ decay

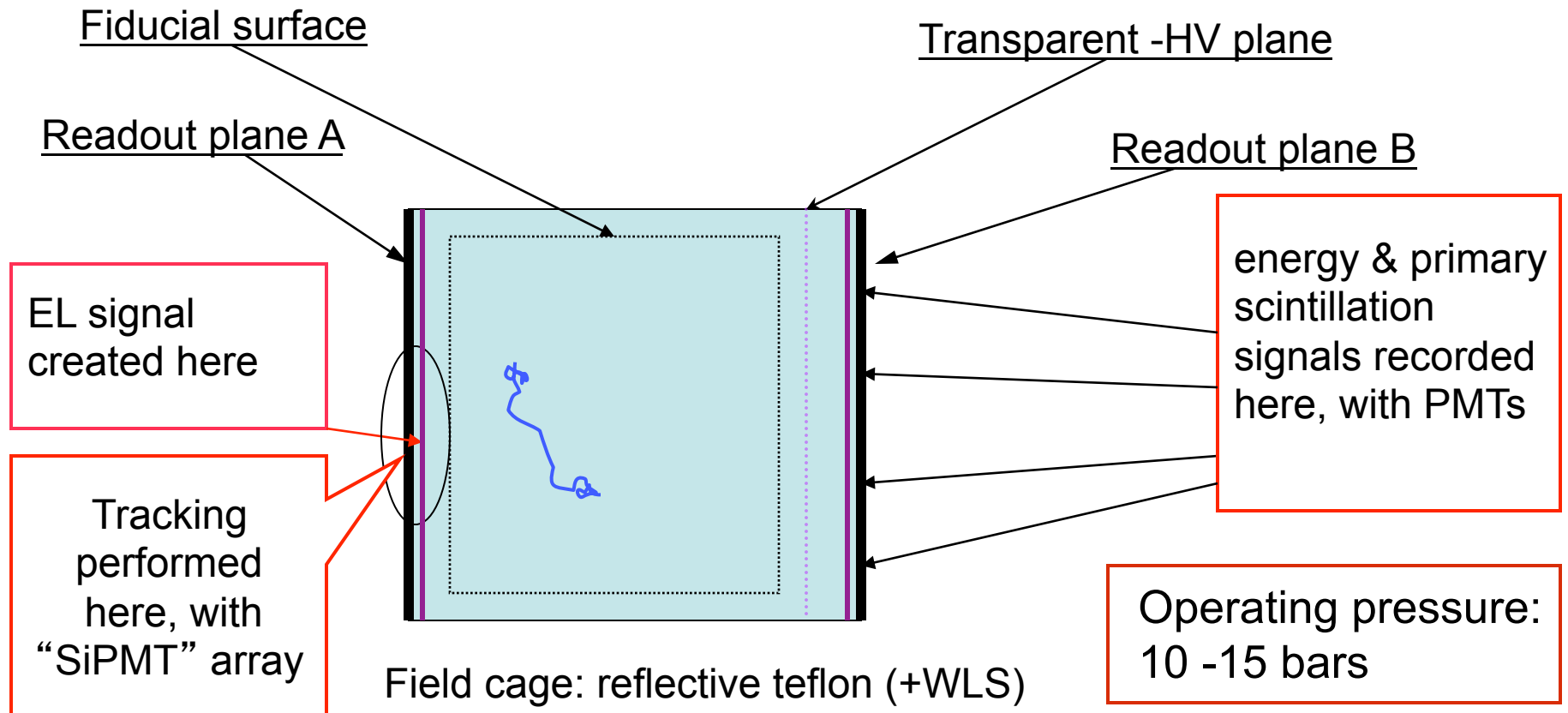
has 100 kg of enriched xenon (85% ^{136}Xe)

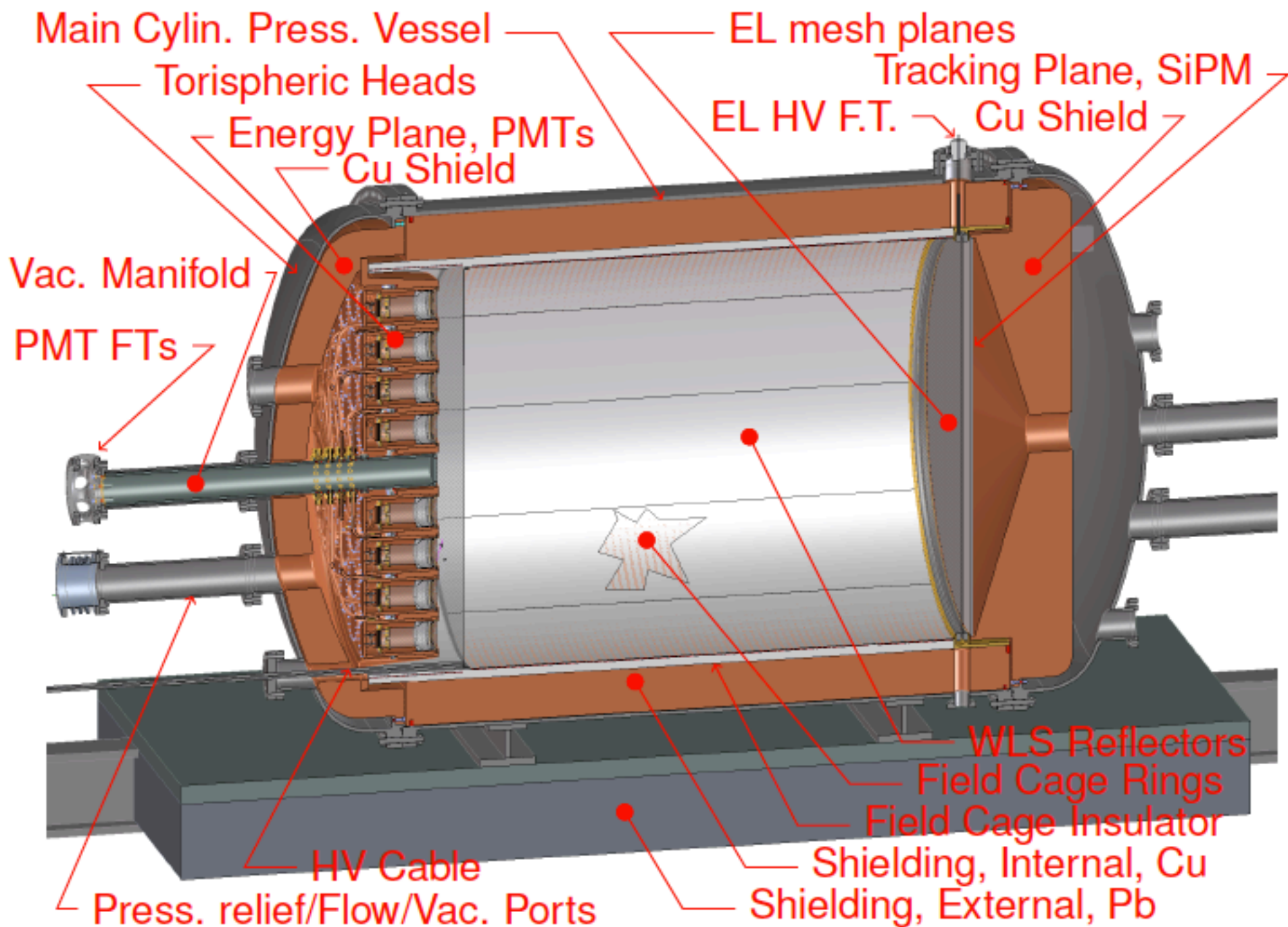
is located in Canfranc Underground Laboratory

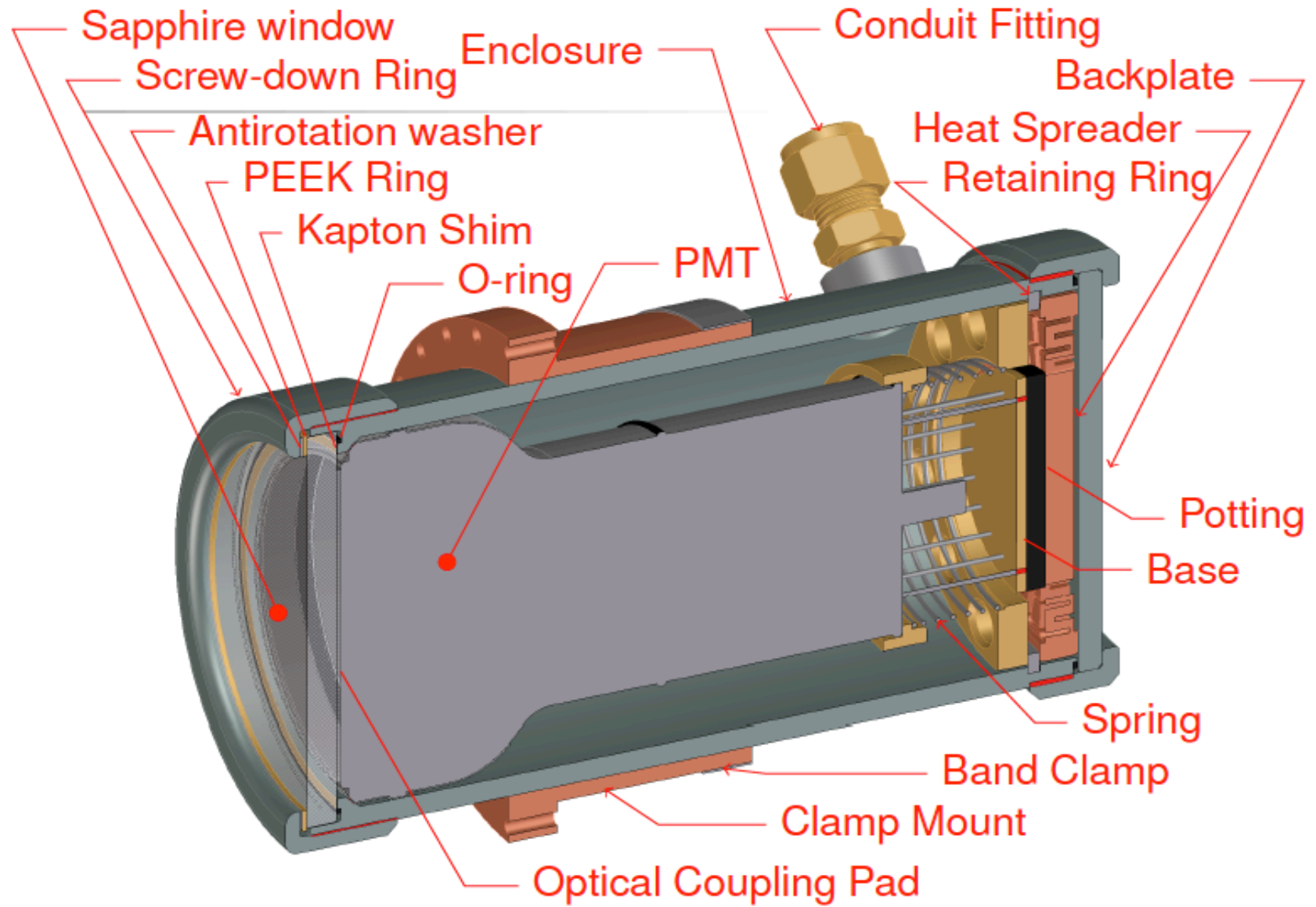
funding by Spanish Funding Agencies: € 6M+

NEXT Asymmetric TPC

“Separated function”





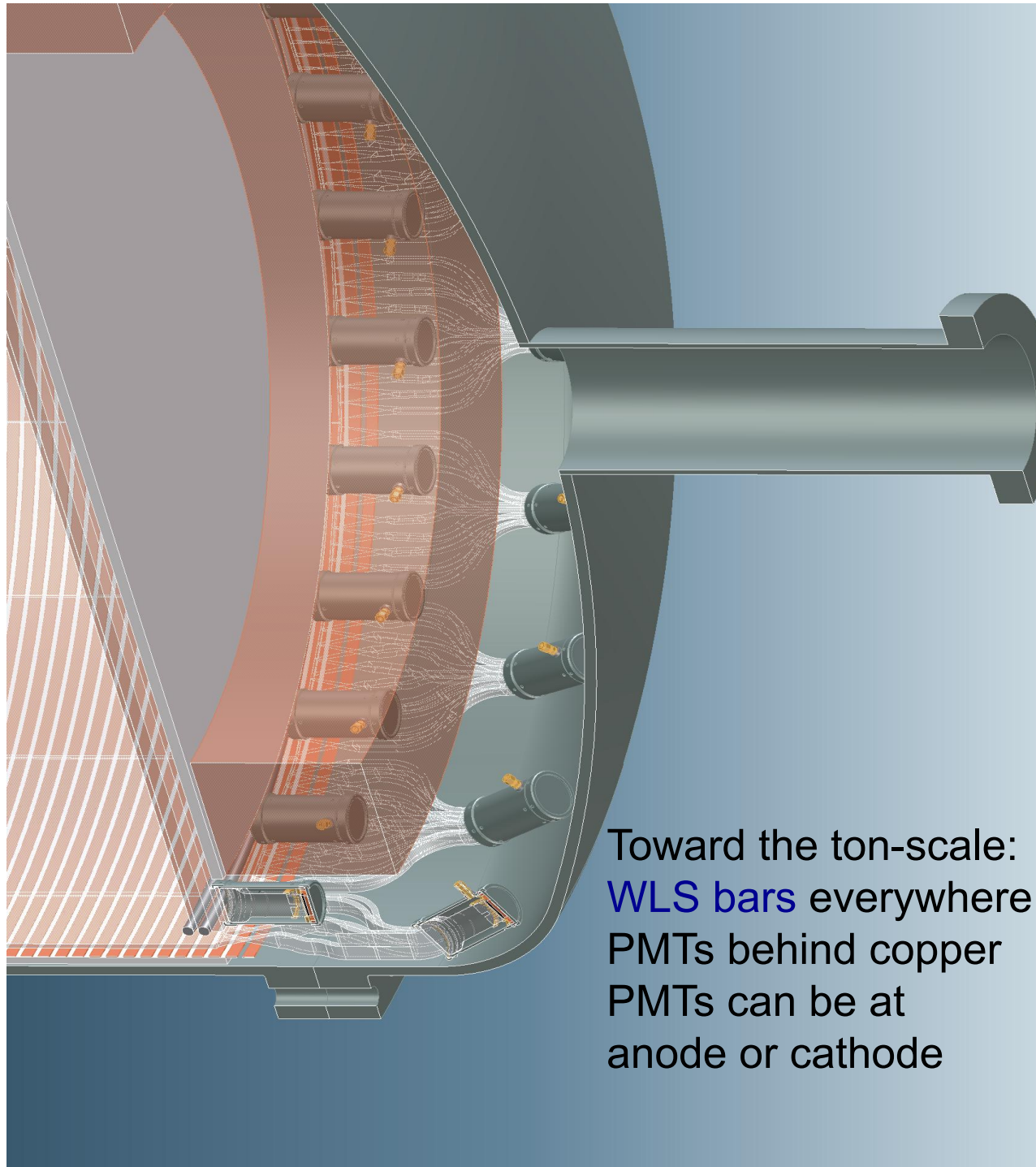


Direct Dark Matter Search?

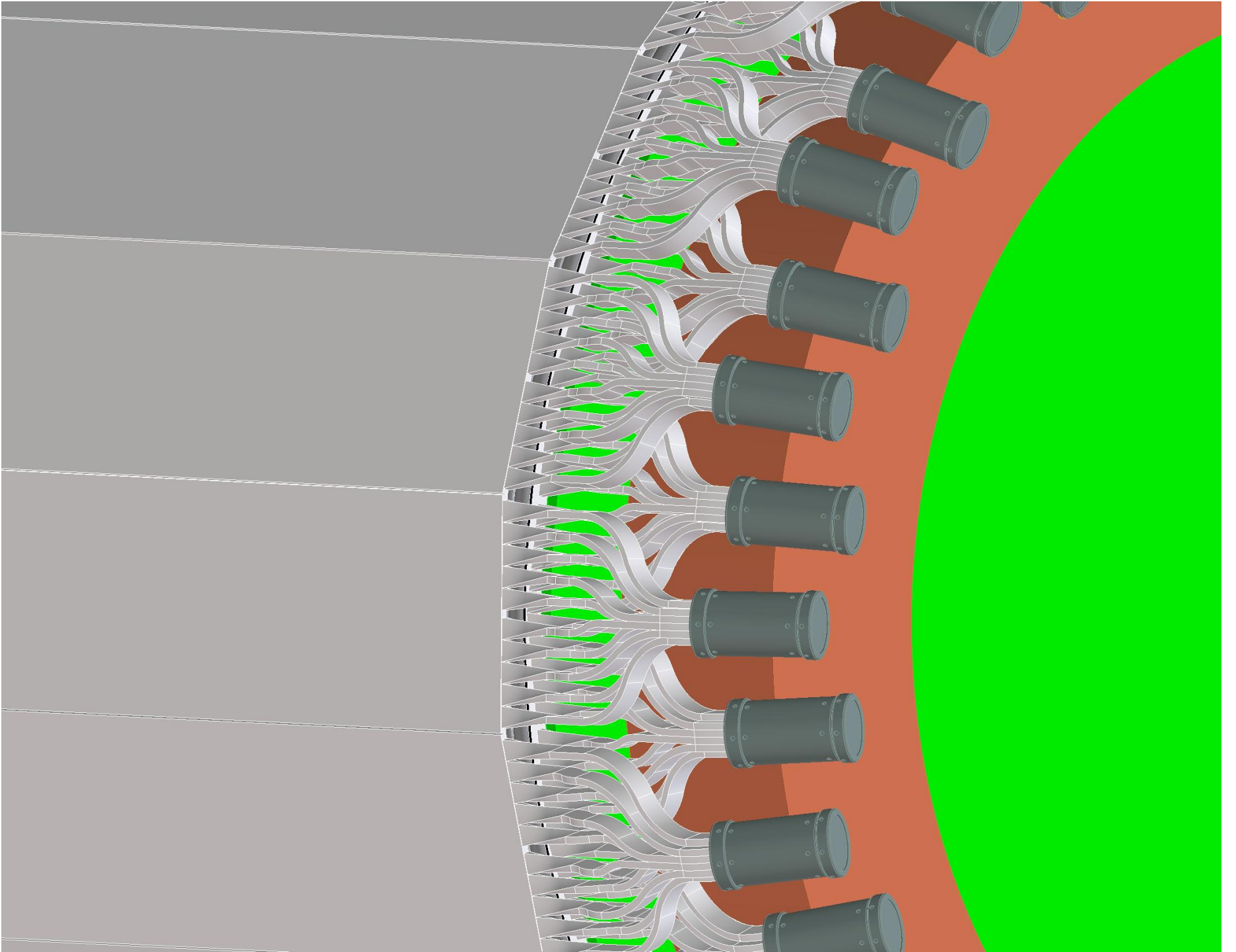
- NEXT-100 not optimized for Dark Matter search, but will serve as a spring board to understand potential
 - S1 detection efficiency determines recoil energy threshold
- Normal fluctuations in S2/S1 in HPXe should offer a huge benefit in electron/nuclear recoil discrimination.
 - F factor ratio of 11 enters exponentially in overlap of gaussian

Ton-scale concept: DM + $0\nu\text{-}\beta\beta$

- The number of PMTs should remain small, or smaller.
- The optical detection efficiency should be maximized to increase sensitivity to low-energy nuclear recoils.
- Approach: use wavelength shifting (WLS) plastic
 - cover the TPC interior completely
 - pipe light to remote PMTs, shielded by copper

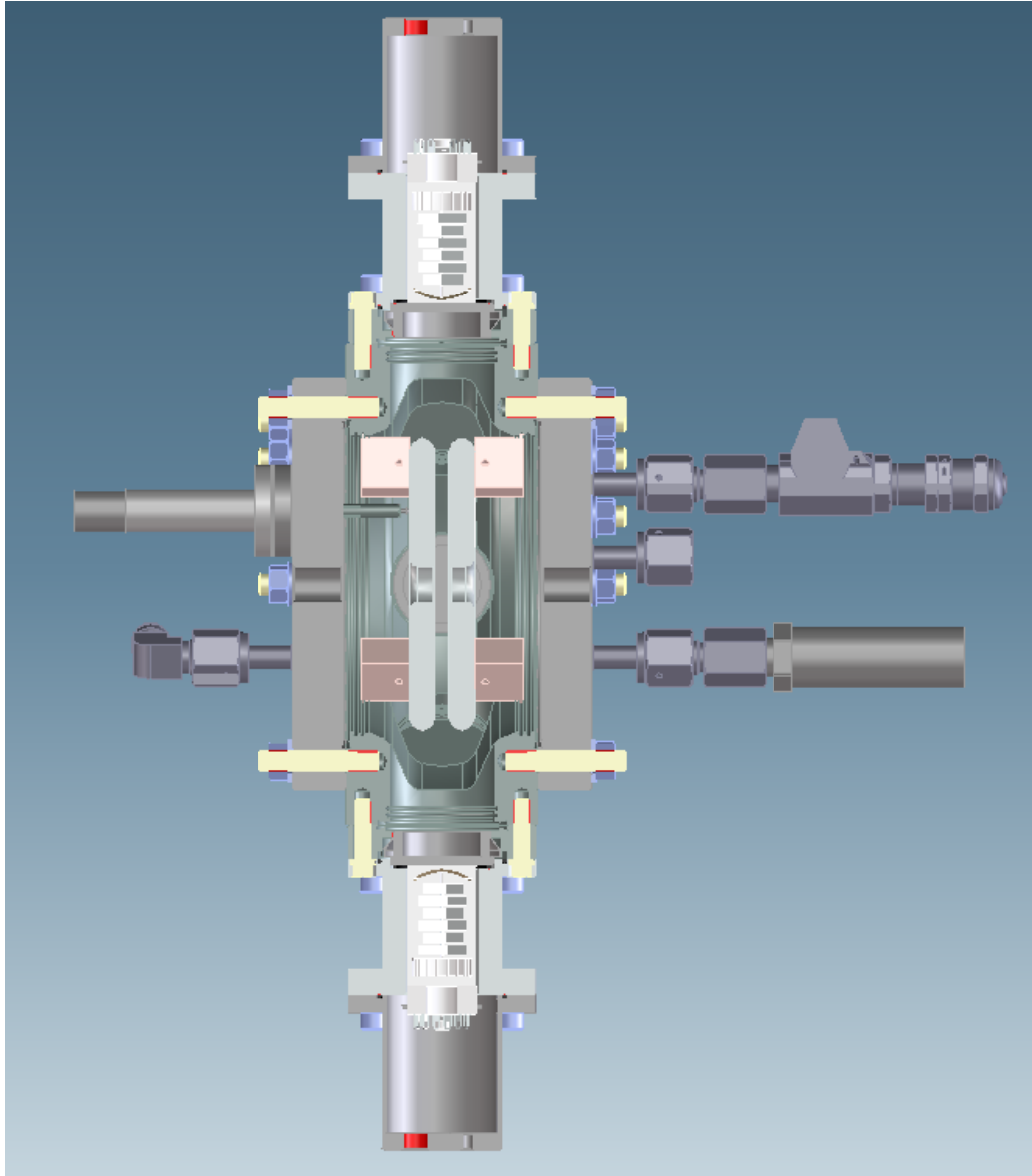


Toward the ton-scale:
WLS bars everywhere
PMTs behind copper
PMTs can be at
anode or cathode



Ton-scale concept...

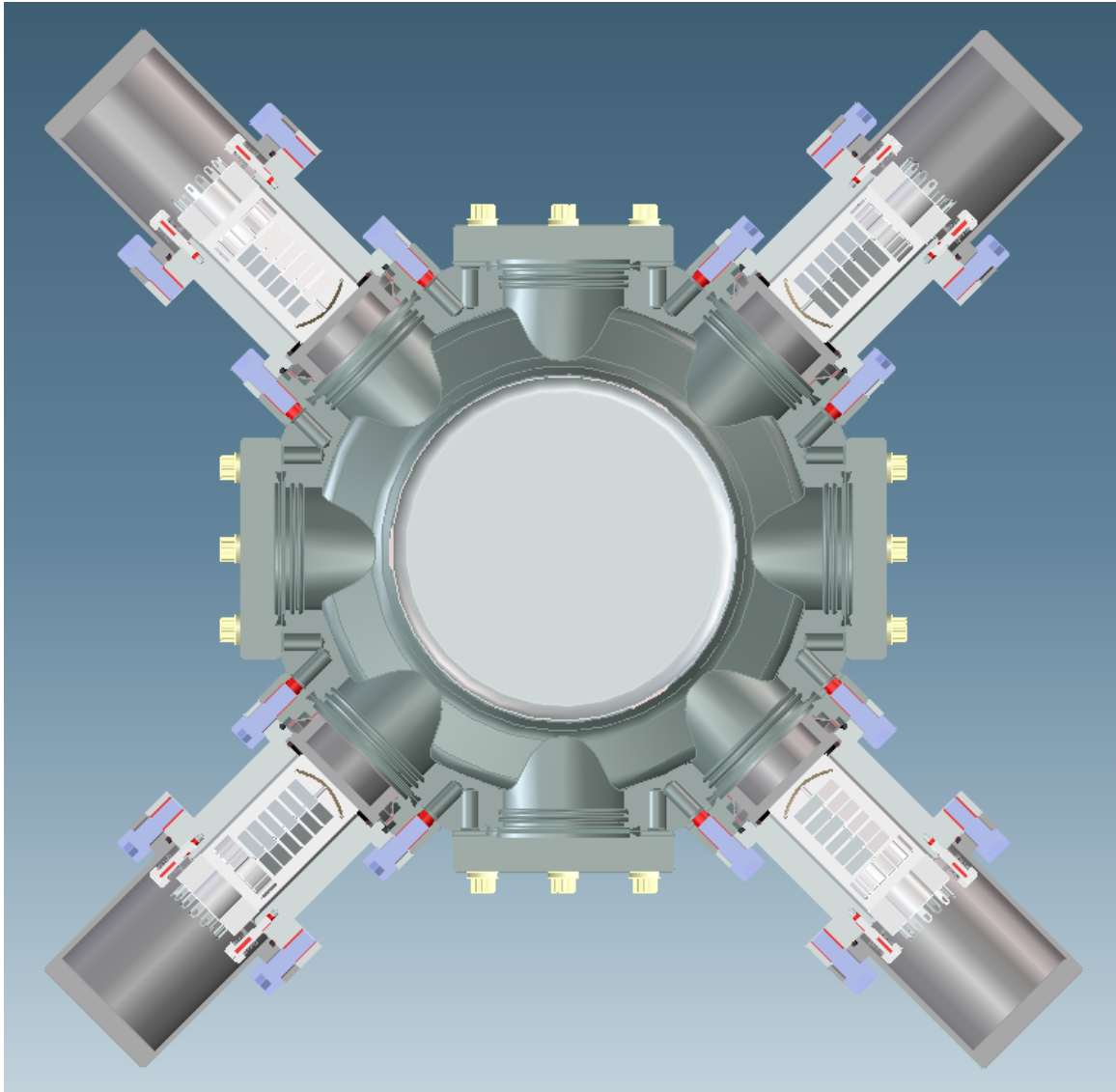
- Problem: xenon light is in VUV - 173 nm
 - WLS plastic response maximum is at 300 nm
 - WLS plastic response at 173 nm is ~zero !!
 - What to do?
- Use gaseous molecular wavelength shifters
 - Trimethylamine (TMA) and/or Triethylamine (TEA)
 - TMA and TEA fluoresce efficiently in bands 285 – 310 nm
 - TMA and TEA may also display Penning effect in xenon
 - Complex behavior expected with density and fraction!



Back to Basics

A parallel-plate ionization chamber with optical sensing, using 4 PMTs that look at the gap from the sides

We will measure both light and charge as functions of density, electric field, and fraction of TEA/TMA,



We can have eight PMTs if useful to do so.

We may modify the cathode and convert to TPC mode to study energy resolution and EL range

Perspective

- Gas phase offers superb energy resolution, event visualization and flexibility in operation
- Electroluminescent gain stage is the key element for near-intrinsic energy resolution for $0\nu\text{-}\beta\beta$ search
- Energy resolution may provide superb discrimination between electron and nuclear recoils through S2/S1
- Can switch easily from enriched to depleted xenon, with real benefits to both searches

Conclusion...

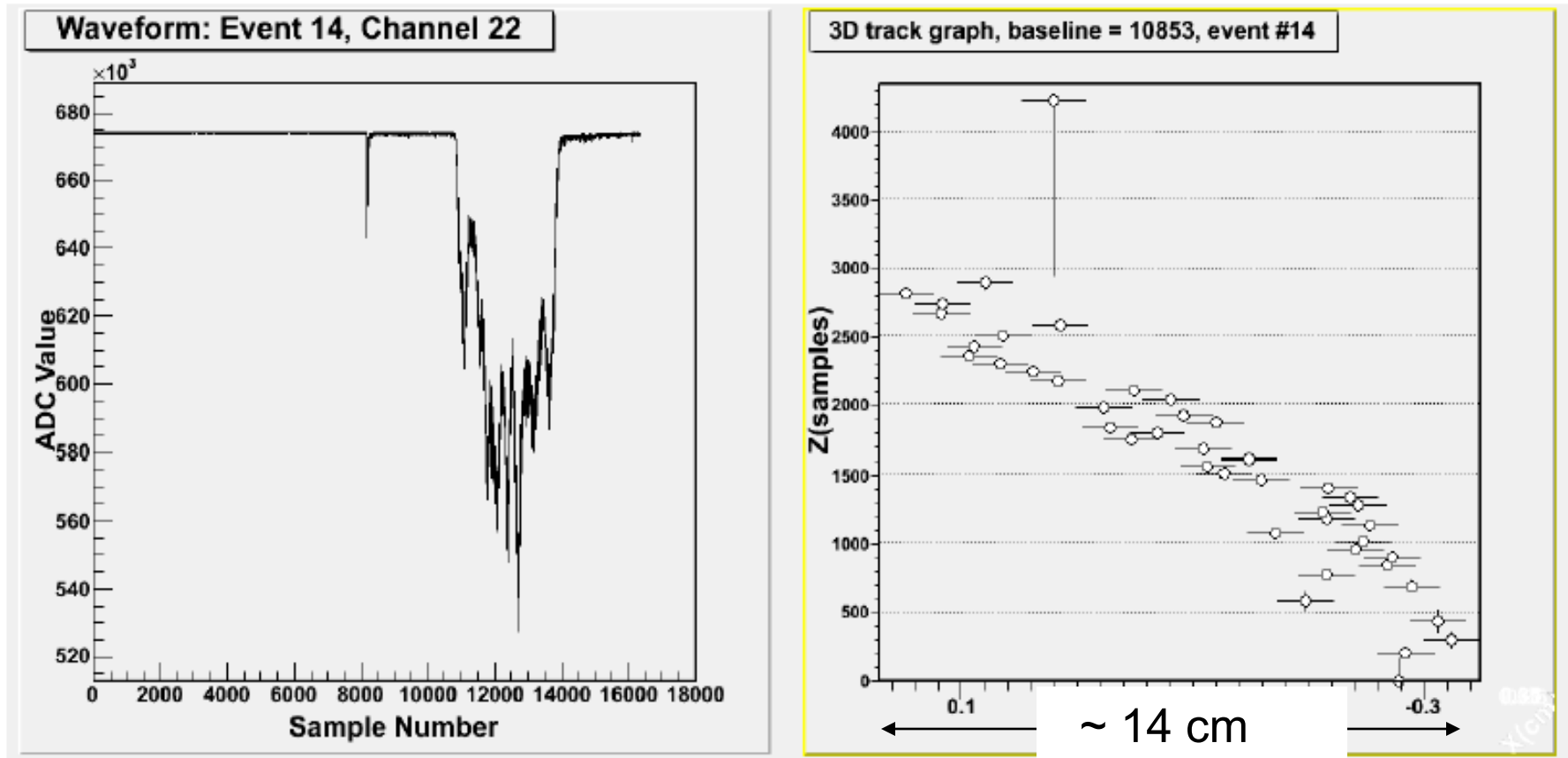
- Is this a true story or a fairy-tale?
 - We should know which in less than a year

Conclusion...

- Is this a true story or a fairy-tale?
 - We should know which in less than a year...
- Gas detectors continue to offer surprises!
 - “You can see a lot by looking” – Yogi Berra

Thank you

A Diagonal Muon Track - “reconstructed”; Signal depends on radius in chamber



Virtues of Electro-Luminescence in HPXe

- Linearity of gain versus pressure, HV
- Immunity to microphonics
- Tolerant of losses due to impurities
- Absence of positive ion space charge
- Absence of ageing, quenching of signal
- Isotropic signal dispersion in space
- Trigger, energy, and tracking functions are accomplished with **optical detectors**

WIMP search, with xenon

Xenon10 data

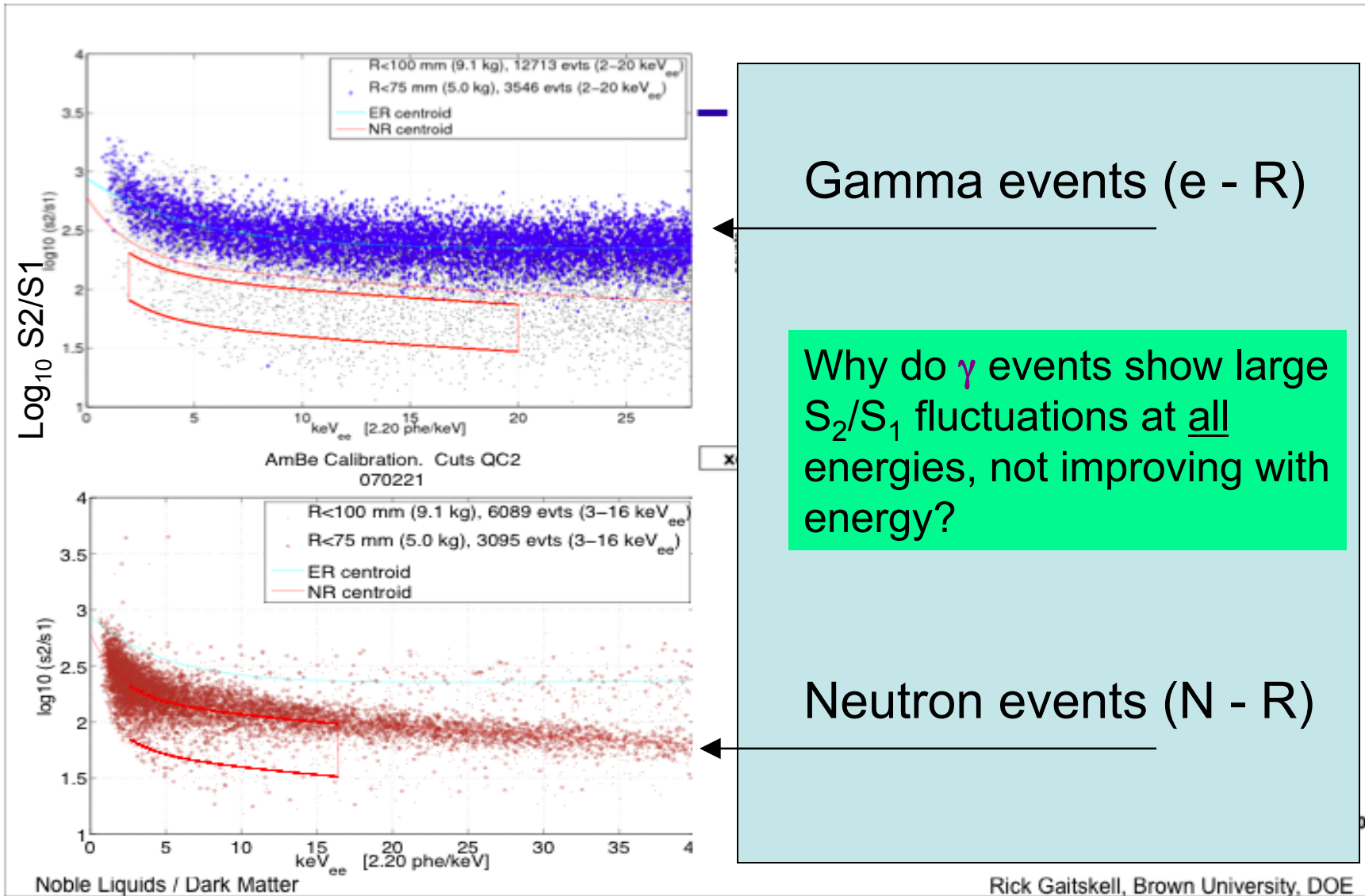


Photo-Luminescence of PMMA

Different WLS nature observed for two PMMA Samples

