

Update on the coupling between the In-Flight RIB Facility EXOTIC and AGATA

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The EXOTIC facility

- In-flight production of **light weakly-bound RIBs**
- It employs **two-body inverse kinematics reactions** with heavy-ion beams delivered from the LNL-XTU tandem accelerator impinging on **gas targets** (p,d, ^3He)
- The **commissioning** of the facility EXOTIC was performed in **2004**

F. Farinon et al., NIM B 266, 4097 (2008)

- In **2012** some upgrades were operated on the facility
 - Installation of a cryogenic gas target
 - Installation of 2 γ -steerers ...

M. Mazzocco et al., NIM B 317, 223 (2013)

- During last year (**2023**), several other upgrades were operated **in order to couple the facility with AGATA**

The EXOTIC facility

Cryogenic Gas Target

1st set of
Quadrupole
triplets

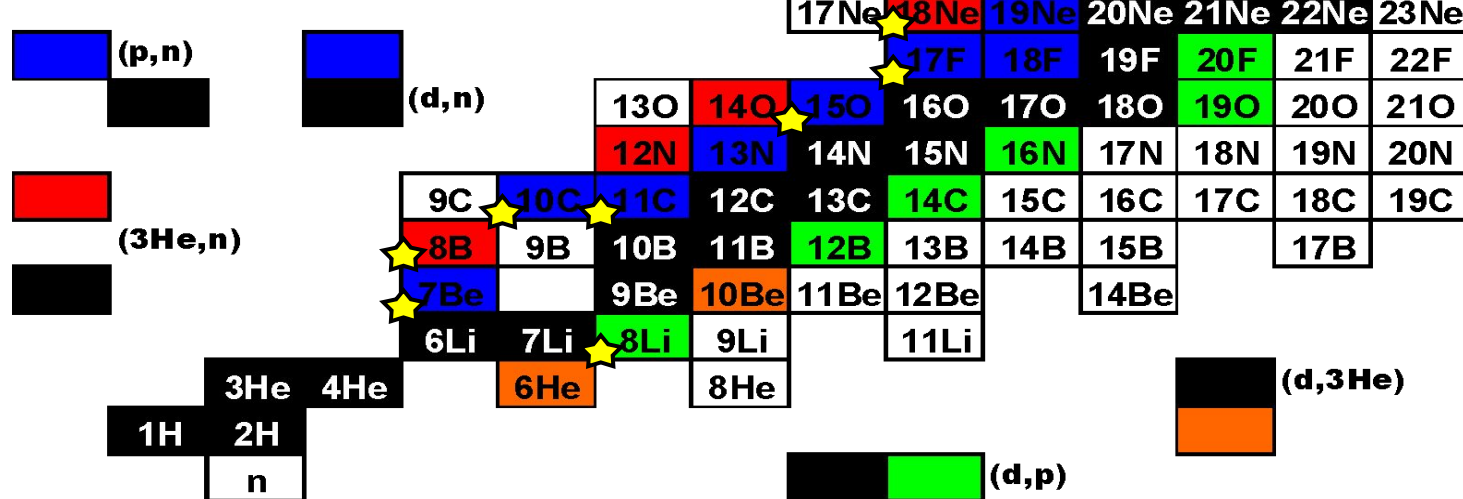
30° Dipole
bending magnet

S2 slits

2nd set of
Quadrupole
triplet

Wien Filter

Light RIBs



¹⁷F	E = 3–5 MeV/u	Purity: 93-96 %	Intensity: 10⁵ pps
⁸B	E = 3–5 MeV/u	Purity: 30-43 %	Intensity: ~ 10³ pps
⁷Be	E = 2.5–6 MeV/u	Purity: 99 %	Intensity: 10⁶ pps
¹⁵O	E = 1.3 MeV/u	Purity: 97-98 %	Intensity: 4*10⁴ pps
⁸Li	E = 2–2.5 MeV/u	Purity: 99 %	Intensity: 10⁵ pps
¹⁰C	E = 4 MeV/u	Purity: 99 %	Intensity: 5*10³ pps
¹¹C	E = 4 MeV/u	Purity: 99 %	Intensity: 2*10⁵ pps
¹⁸Ne	E = 1.3 MeV/u	Purity: 95 %	Intensity: 6*10³ pps (under dev.)

EXOTIC + AGATA

- **October 2023:** Installation of a new reaction chamber downstream the EXOTIC line



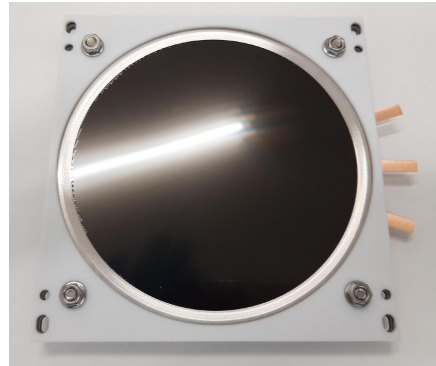
EXOTIC + AGATA

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EXOTIC + AGATA

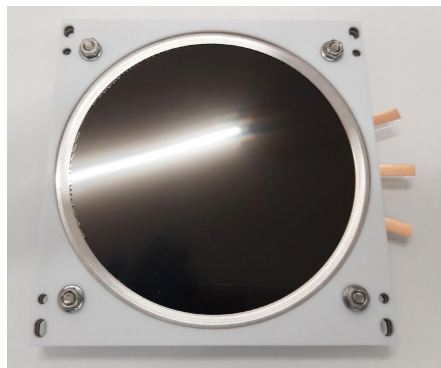
- **October 2023:** Installation of a new reaction chamber downstream the EXOTIC line
- Development of the 1st **Micro-Channel Plate (MCP)** detector



Coupling EXOTIC and AGATA

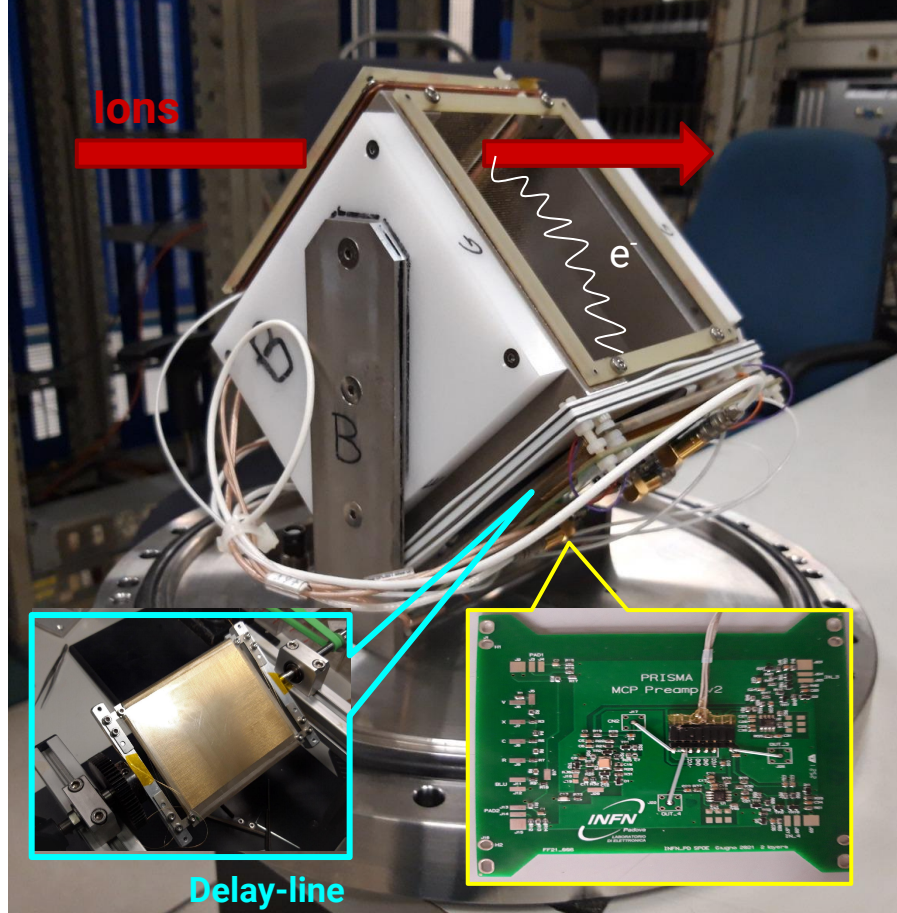
EXOTIC + AGATA

- **October 2023:** Installation of a new reaction chamber downstream the EXOTIC line
- Development of the 1st **Micro-Channel Plate** (MCP) detector
- **May-June 2024:** Installation of the second chamber and mounting of 2nd MCP

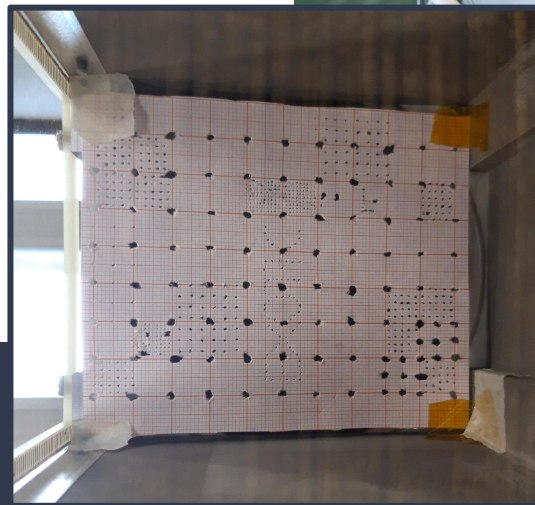
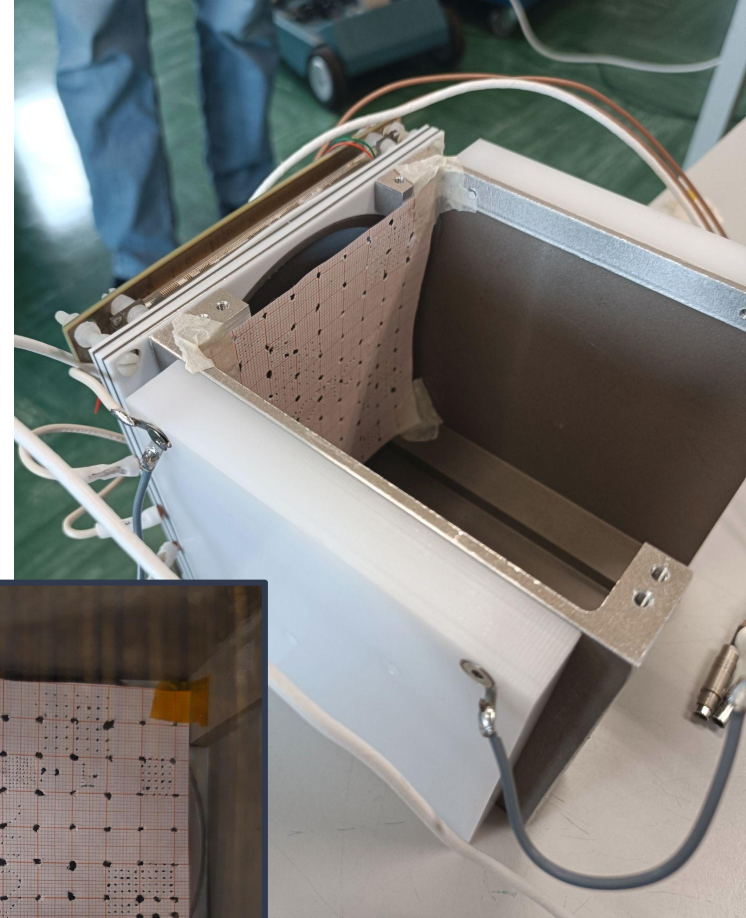


Micro-Channel Plate

- Aluminized Mylar foils
 - + LiF (to increase efficiency)
- Position-sensitive detector
 - Reconstruction of X and Y hit position
 - **Position-tracking of the RIB between 2 MCPs**
- Time signal of the hit
 - **Time-of-Flight (ToF) reconstruction between 2 MCPs**

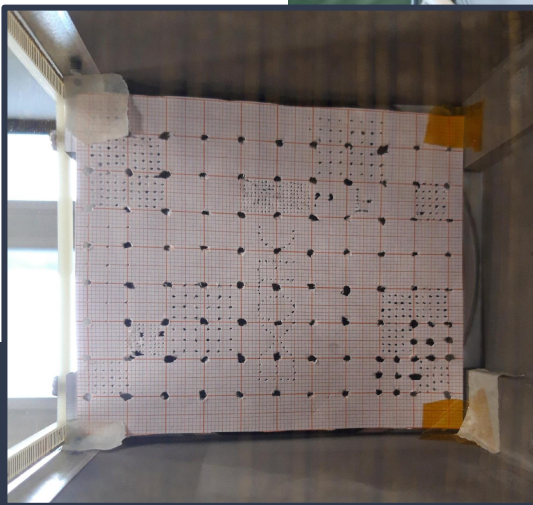
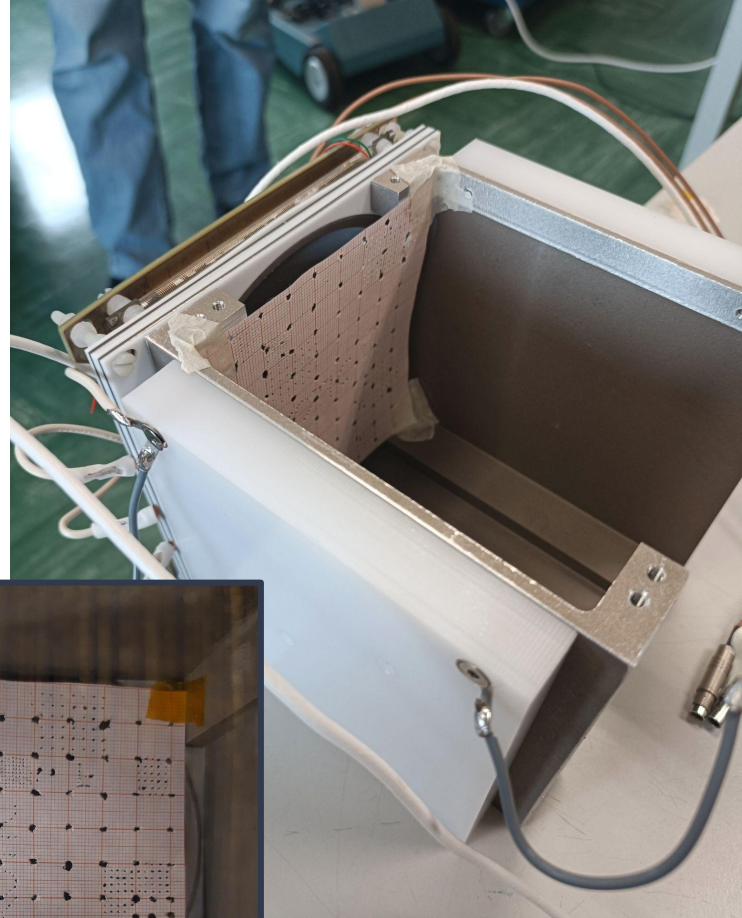
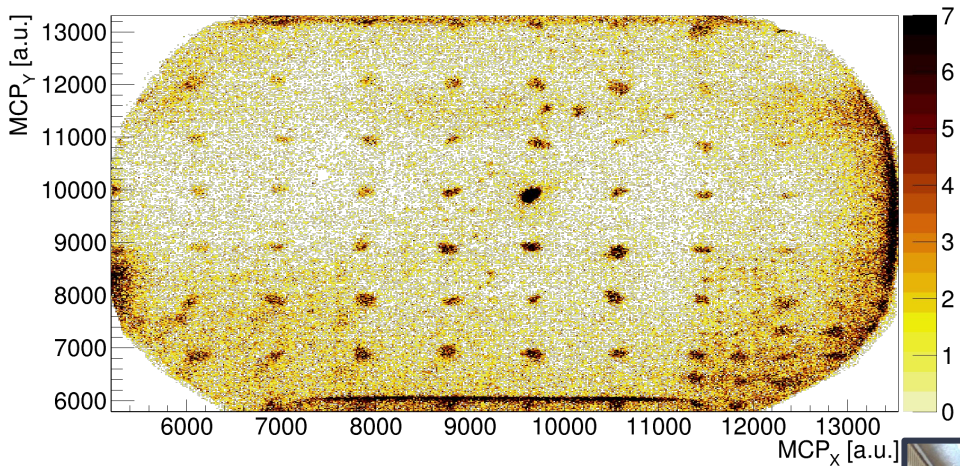


Position resolution



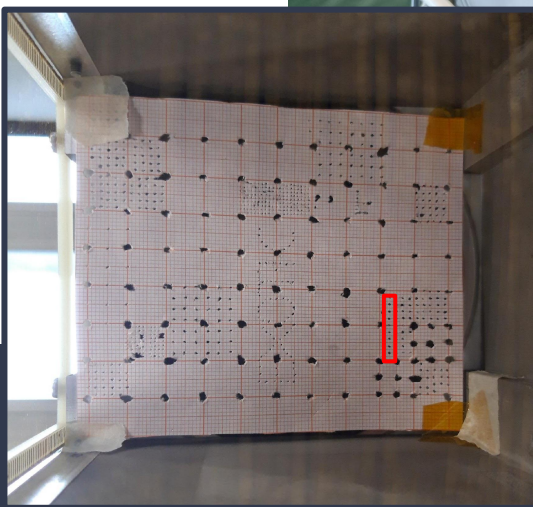
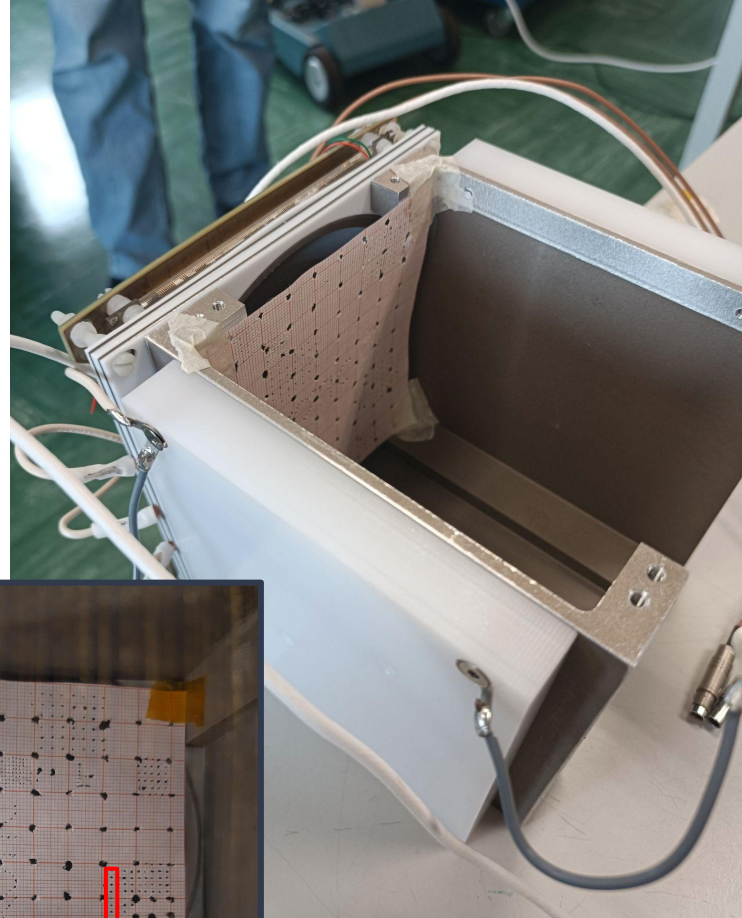
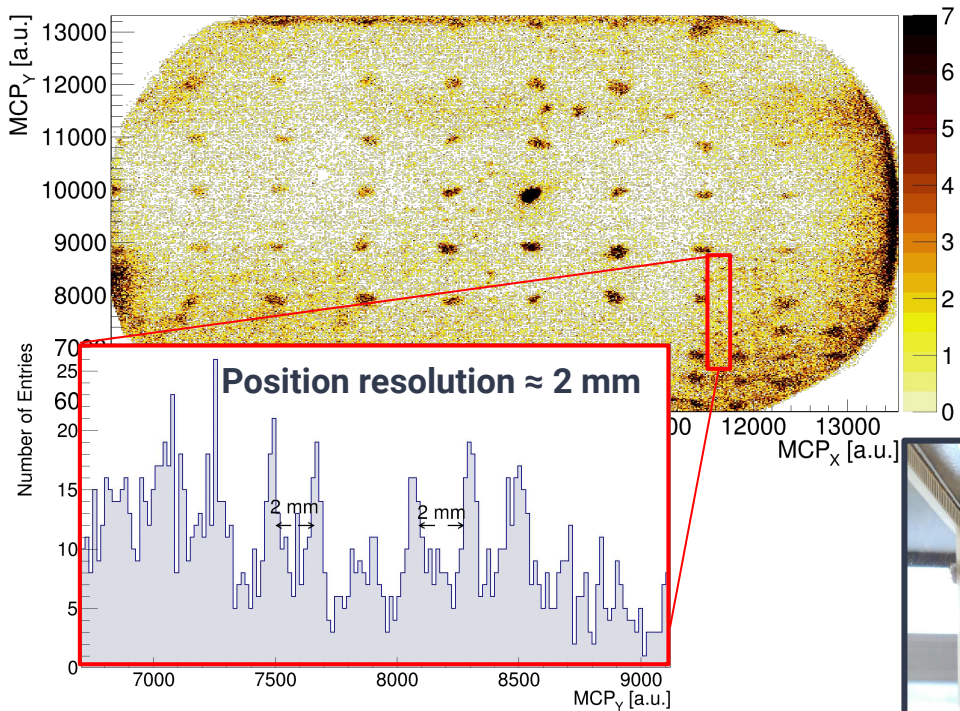
Characterization of the MCP detectors

Position resolution



Characterization of the MCP detectors

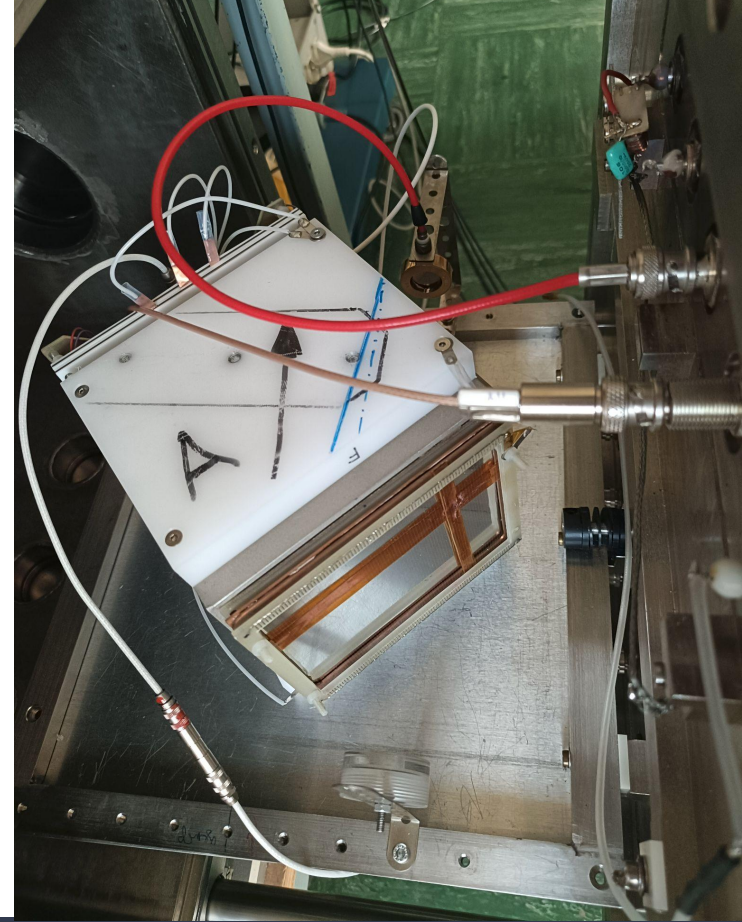
Position resolution



Characterization of the MCP detectors

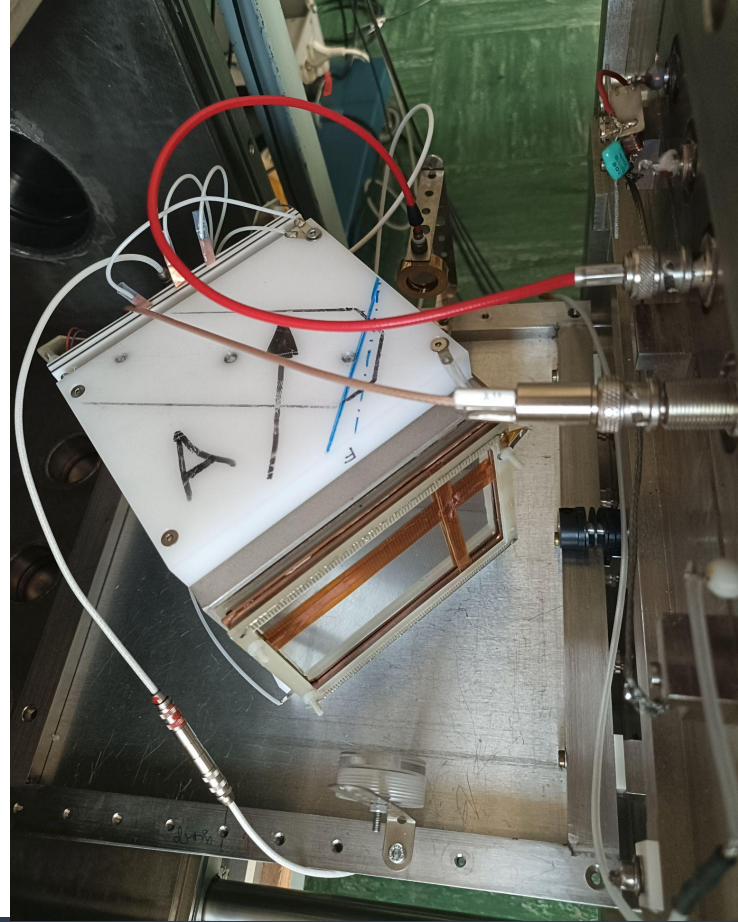
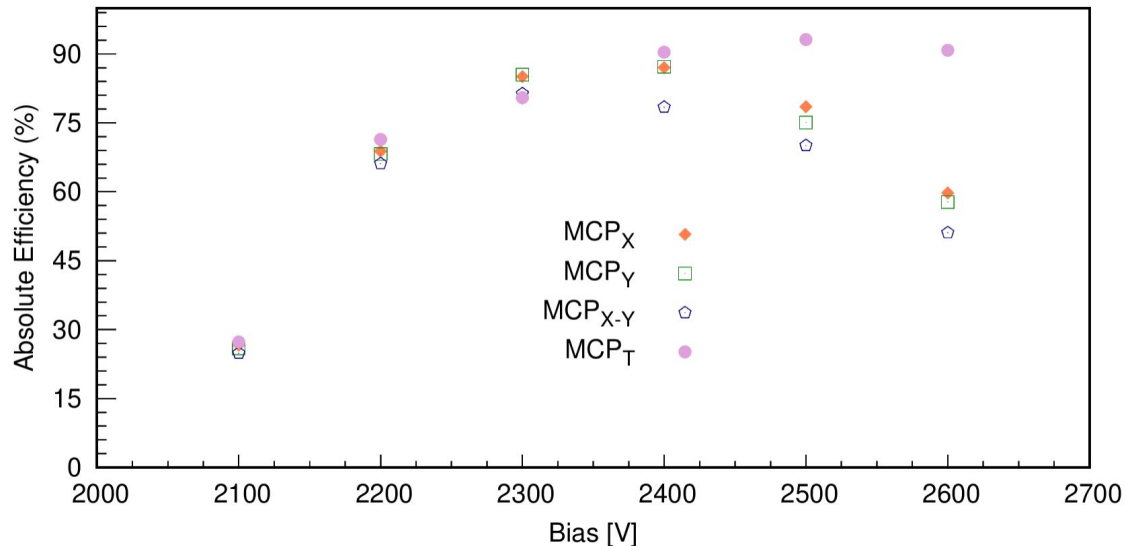
Efficiency and ToF resolution

- MCP + Surface barrier Silicon detector



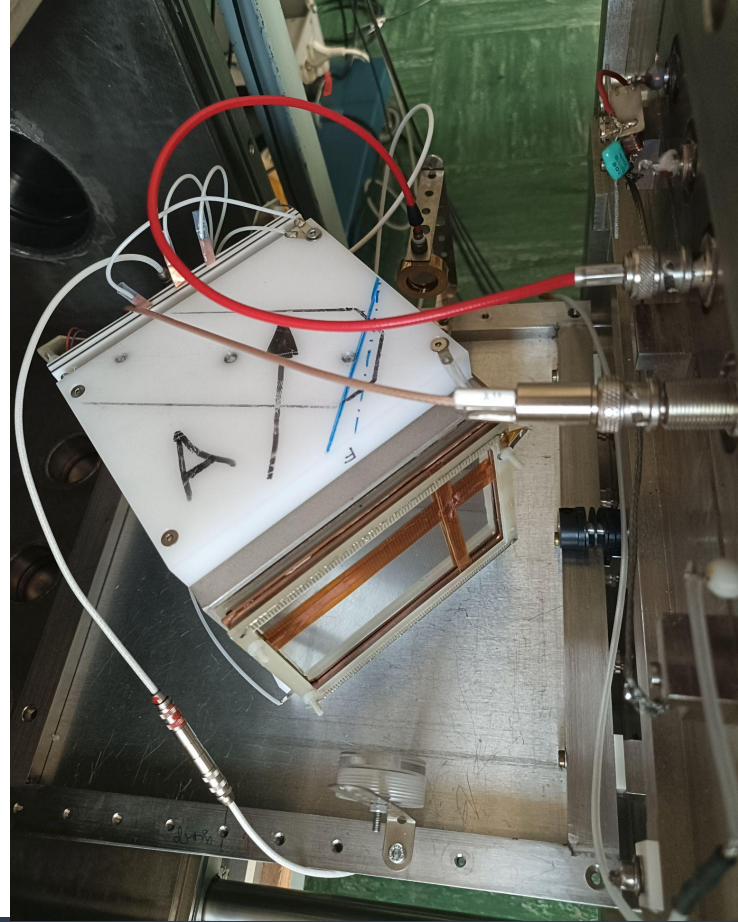
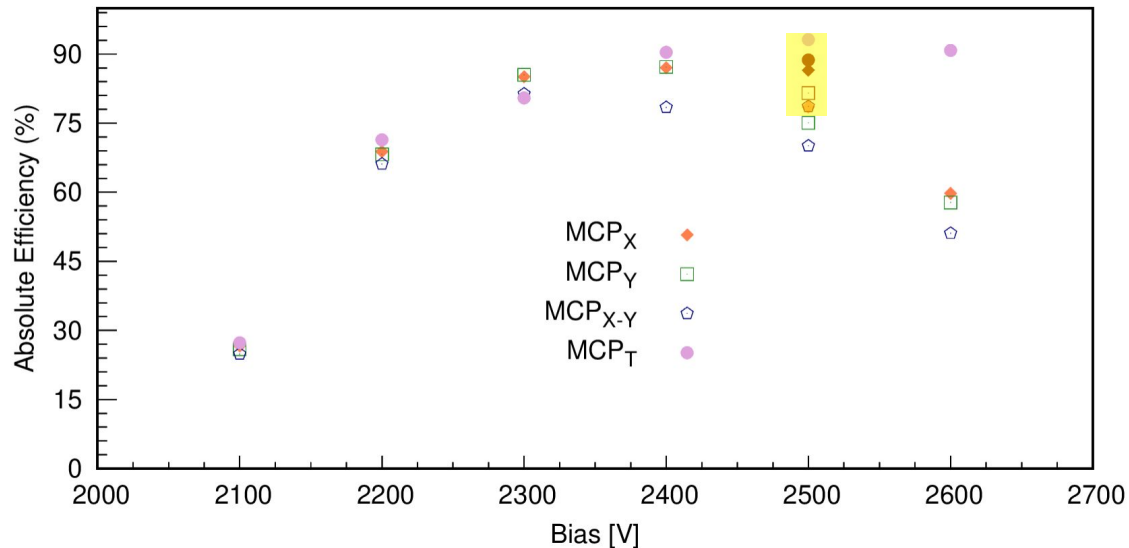
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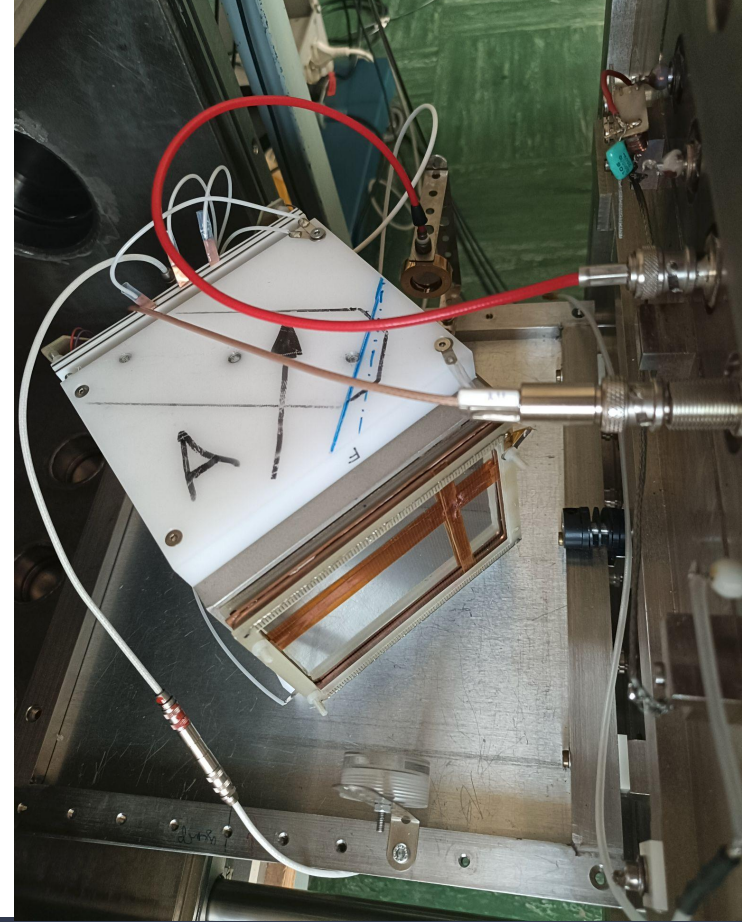
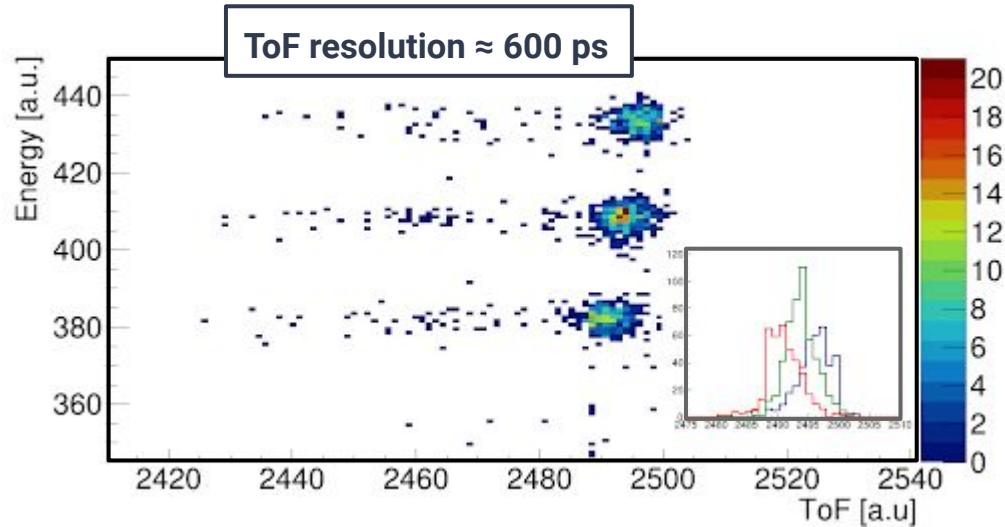
Efficiency and ToF resolution

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Efficiency and ToF resolution

- MCP + Surface barrier Silicon detector



Position tracking

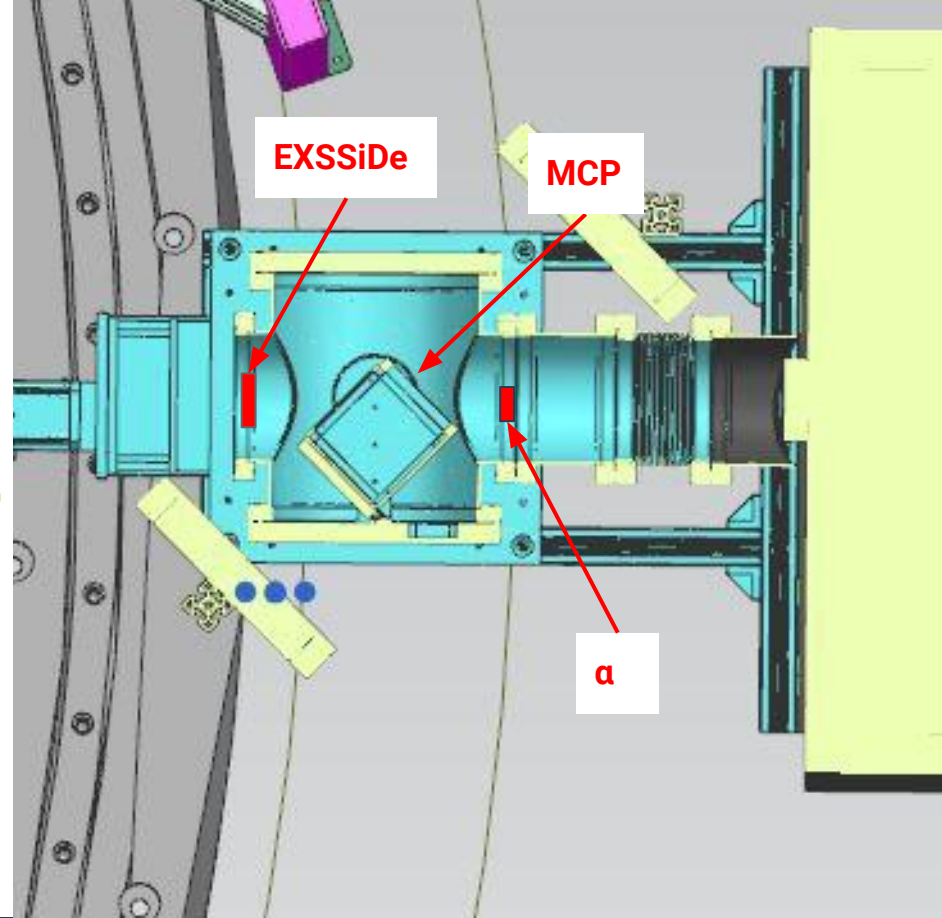
EXotic Silicon Strip Detector (EXSSiDe)

Area: 64 mm x 64 mm

Thickness: 1 mm

Strip: 32 vertical x 32 horizontal

Position Resolution: 2 mm x 2 mm



Position tracking

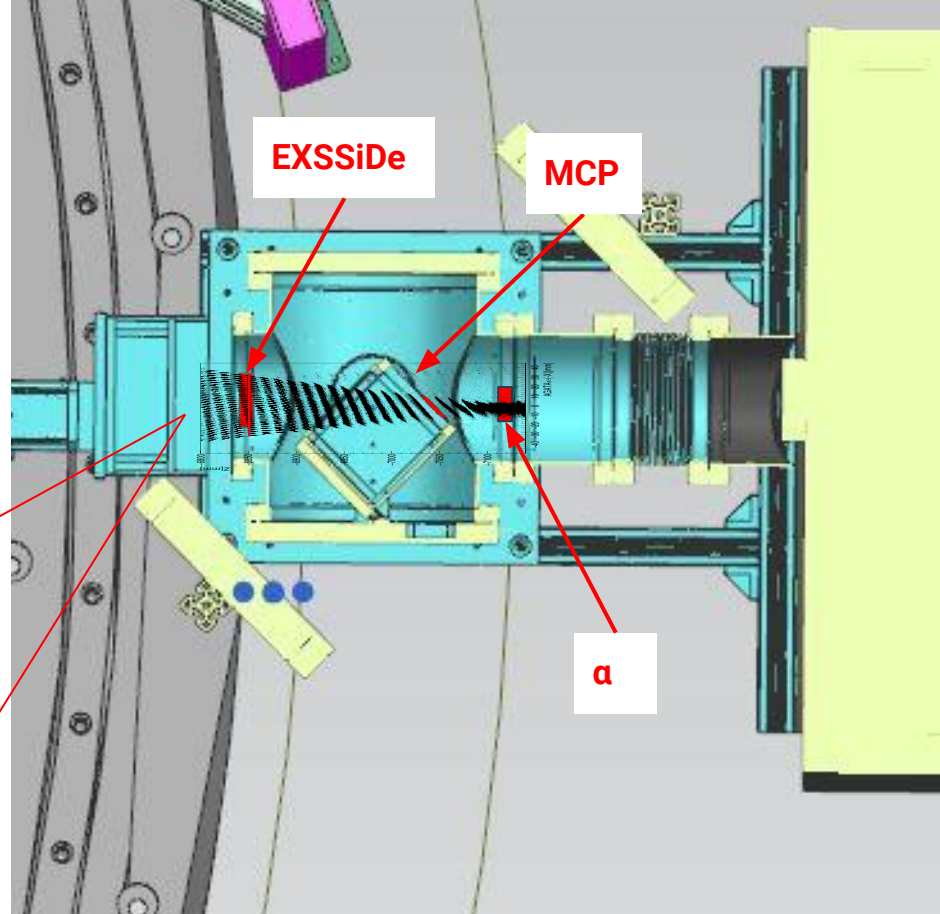
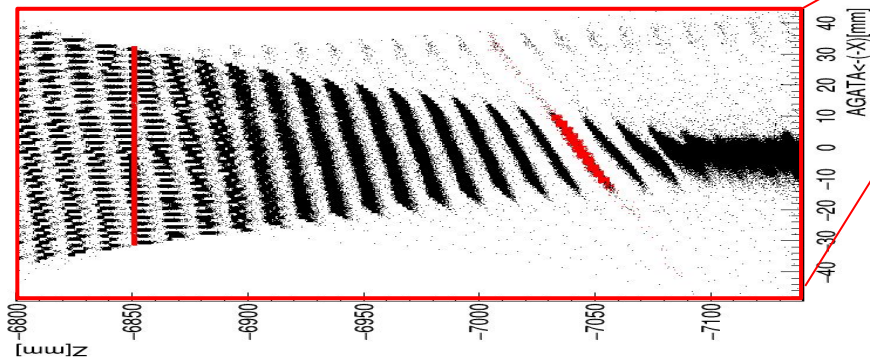
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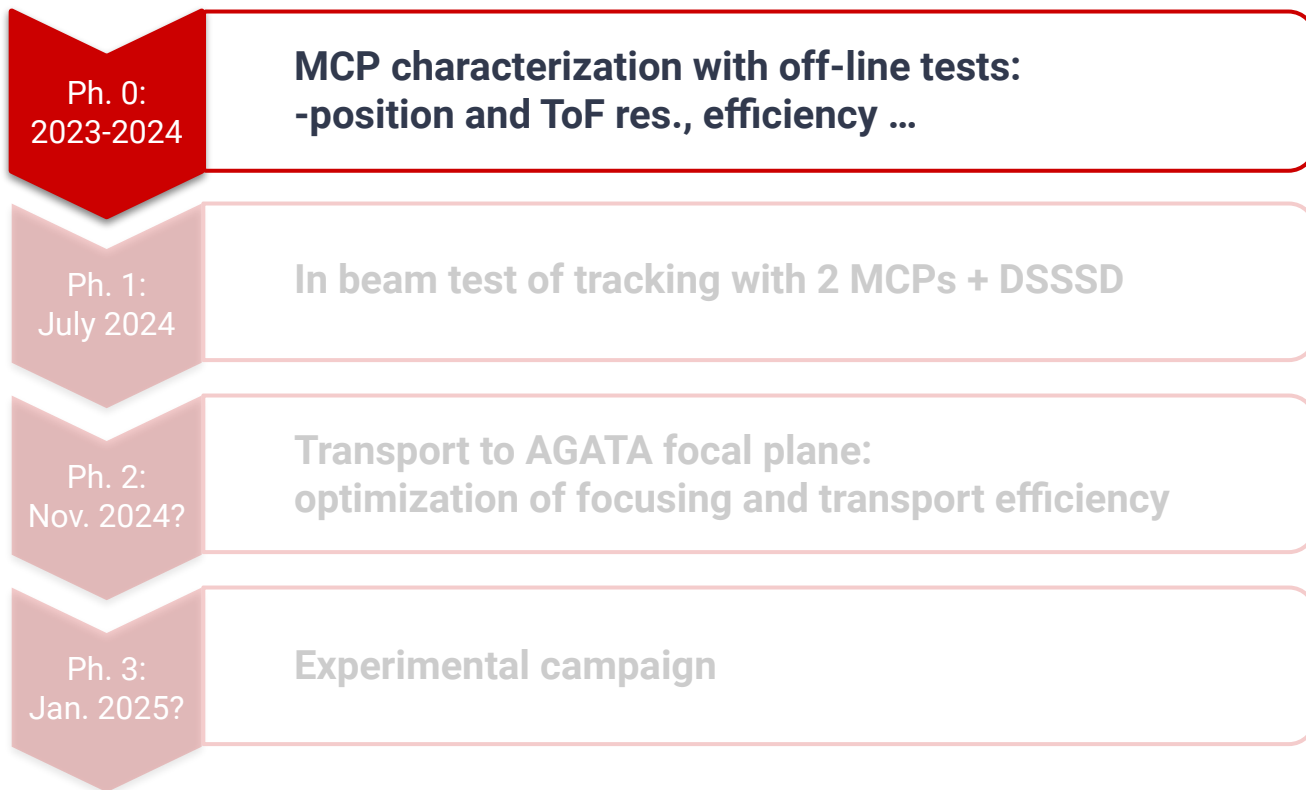
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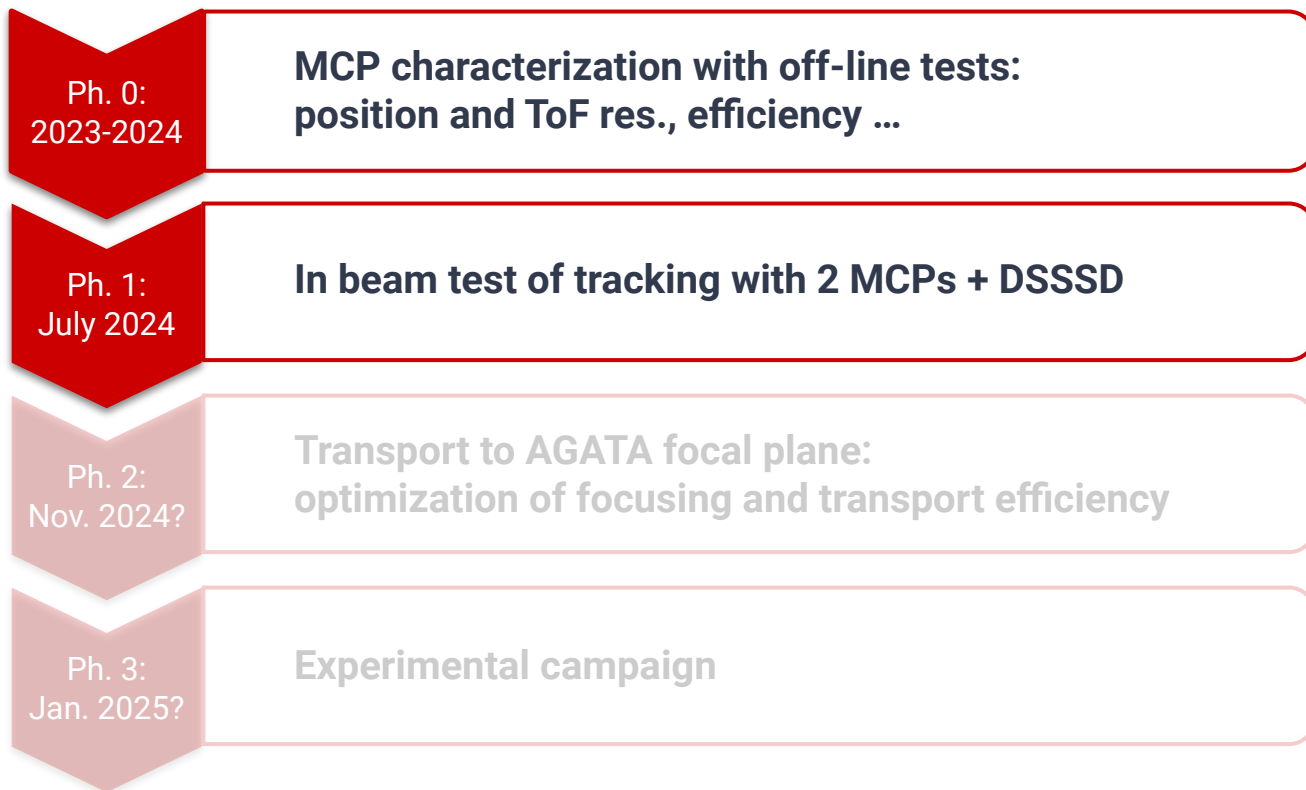
Position Resolution: 2 mm x 2 mm



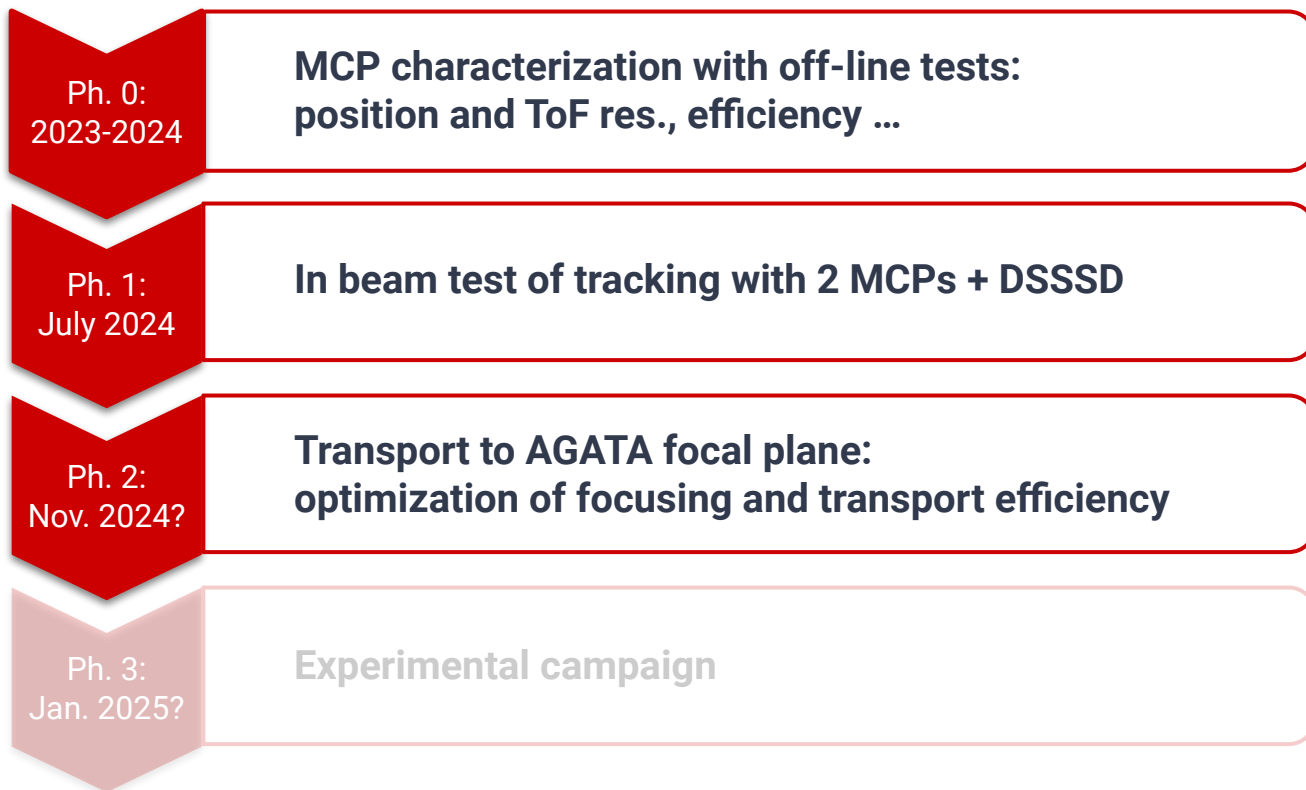
Timeline



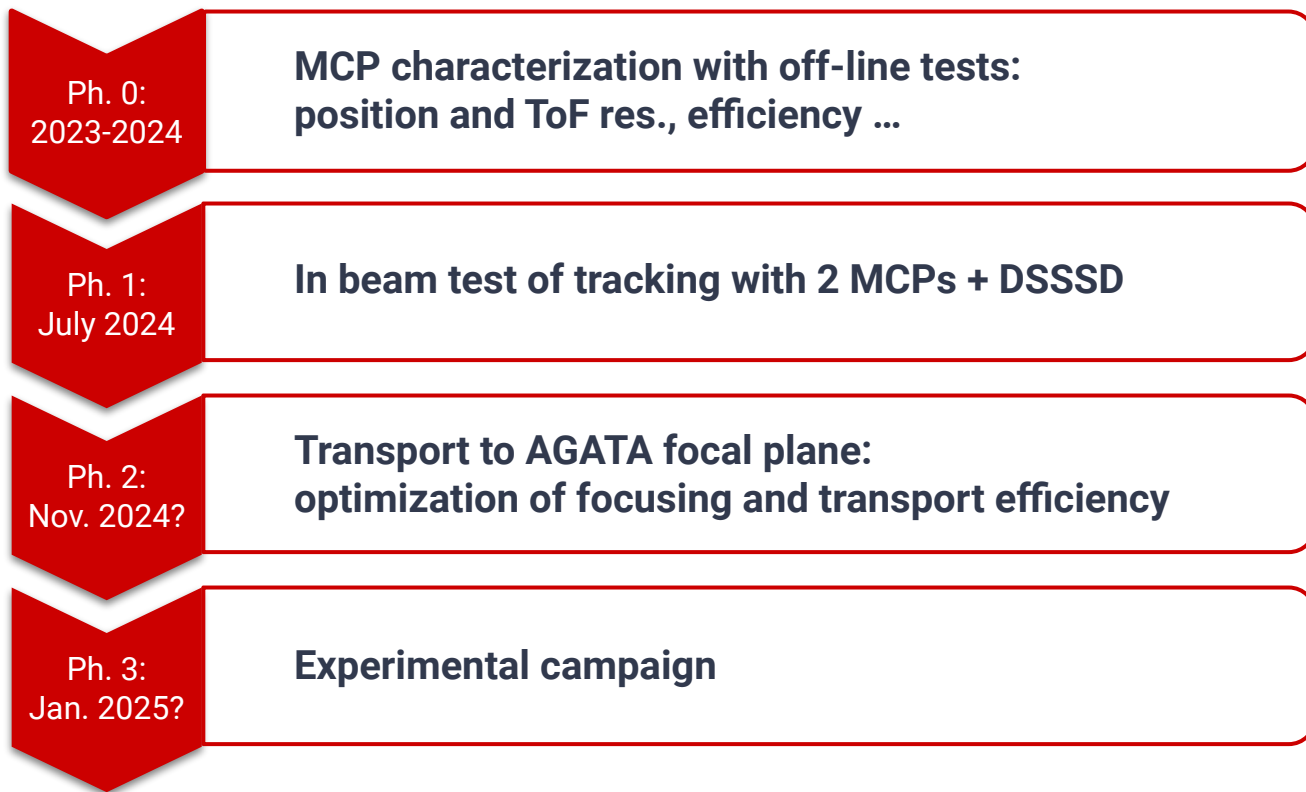
Timeline



Timeline



Timeline



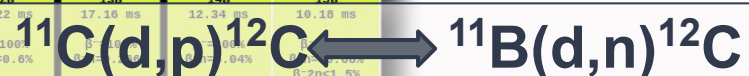
Isospin Symmetry Breaking for $A=11$ (A. Gottardo)

Are there **nuclear contributions** to breaking of the Isospin Symmetry?

Ab-initio calculations possible for $A=11$!

	110 2.31 meV 2p=100%	120 51 keV 2p=100%	130 0.58 ms $\epsilon+\beta+=100\%$ $\epsilon\rho=10.8\%$	140 70.62 s $\epsilon+\beta+=100\%$	150 122.266 s $\epsilon+\beta+=100\%$	160 STABLE 99.738-99.776%	170 STABLE 0.0367-0.04%	180 STABLE 0.187-0.222%
10N 2.5 meV p=100%	11N 780 keV p=100%	12N 10.996 ms $\epsilon+\beta+=100\%$ $\epsilon\alpha=1.93\%$	13N 9.967 min $\epsilon+\beta+=100\%$	14N STABLE 99.678-99.683%	15N STABLE 0.337-0.422%	16N 7.13 s $\beta^-=100\%$ $\beta^-\alpha=1.54e-3\%$	17N 4.171 s $\beta^-=100\%$ $\beta^-n=95.1\%$ $\beta^-a=2.5e-3\%$	
9C 126.5 ms $\epsilon+\beta+=100\%$ $\epsilon\rho=62\%$ $\epsilon\alpha=37.9\%$	10C 19.30 s $\epsilon+\beta+=100\%$	11C 20.34 min $\epsilon+\beta+=100\%$	12C STABLE 98.84-99.04%	13C STABLE 0.96-1.16%	14C 5686 y $\beta^-=100\%$	15C 2.451 s $\beta^-=100\%$	16C 750 ms $\beta^-=100\%$ $\beta^-n=99.28\%$	
8B 772.1 ms $\epsilon+\beta+=100\%$ $\epsilon\alpha=100\%$	9B 540 eV p=100%	10B STABLE 18.9-20.4%	11B STABLE 79.6-81.1%	12B 20.22 ms $\beta^-=100\%$ $\beta^-\alpha=0.6\%$	13B 17.16 ms $\beta^-=100\%$ $\beta^-n=0.04\%$	14B 12.34 ms $\beta^-=100\%$ $\beta^-n=0.002\%$ $\beta^-2n<1.5\%$	15B 10.18 ms $\beta^-=100\%$ $\beta^-n=0.002\%$ $\beta^-2n<1.5\%$	
7Be 53.3 d $\epsilon=100\%$	8Be 5.57 eV $\alpha=100\%$	9Be STABLE 100%	10Be 1.39e+6 y $\beta^-=100\%$	11Be 13.77 s $\beta^-=100\%$ $\beta^-a=3.3\%$	12Be 21.46 ms $\beta^-=100\%$ $\beta^-n=0.5\%$	13Be 450 keV n ?	14Be 4.65 ms $\beta^-=100\%$ $\beta^-n=86\%$ $\beta^-2n=5\%$	

Same reaction in Isospin Symmetry



Commissioning of the coupling between EXOTIC and AGATA (Part III)

EXOTIC

Spokespersons: S. Pigliapoco¹, D. Brugnara², M. Mazzocco¹, J. J. Valiente-Dobon²

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⁶ *Università di Napoli and INFN, Sezione di Napoli, Napoli, Italy*

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Thank you!