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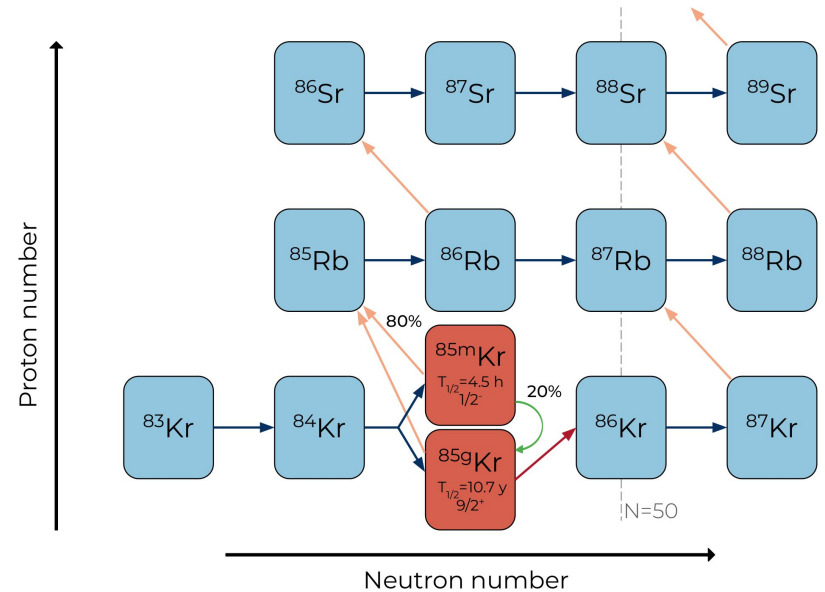
Study of the $^{85g}\text{Kr}(d,p\gamma)$ reaction for astrophysics at ANL

Sara Carollo
University and INFN Padova

Young GAMMA meeting - June 21ST 2024

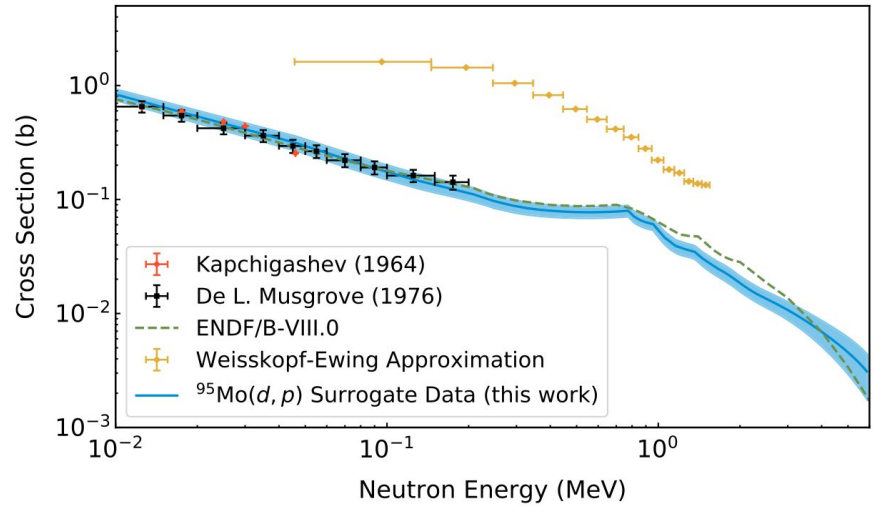
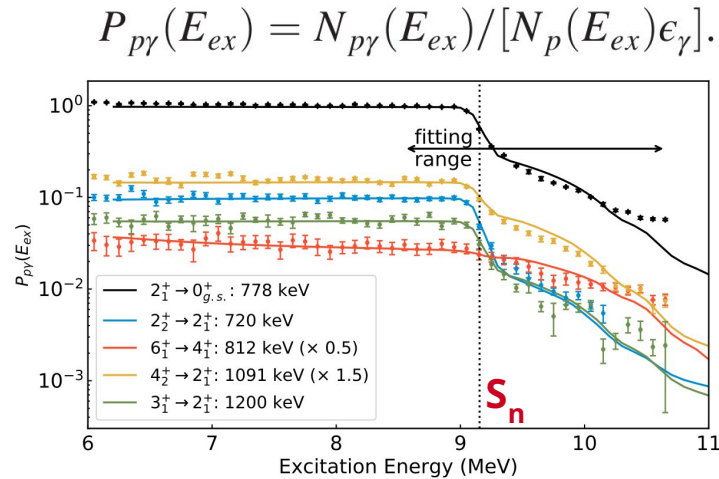
Motivation

- 50% of elements heavier than Fe are produced by the s-process: $\tau_{\beta} \lesssim \tau_n$
- Great uncertainty derives from the competition between n-capture and β -decay in some isotopes called **branching points**
- **^{85}Kr** is an important branching point of the s-process, that influences:
 - $^{86}\text{Kr}/^{82}\text{Kr}$ ratio in **presolar grains**
 - Abundances of heavy **Sr isotopes** that are produced also by r-process (lines in kilonova)



Surrogate reaction method: (n,γ) from $(d,p\gamma)$

^{85}Kr activity is too high to perform activation or ToF measurement \rightarrow Surrogate reaction method
 $(d,p\gamma)$ can be performed in inverse kinematics \rightarrow ^{85}Kr as beam \rightarrow $\geq 99\%$ purity!



J. E. Escher et al., Phys. Rev. Lett. 121, 052501 (2018)

A. Ratkiewicz et al. Phys. Rev. Lett. 122, 052502 (2019)

Experimental set-up

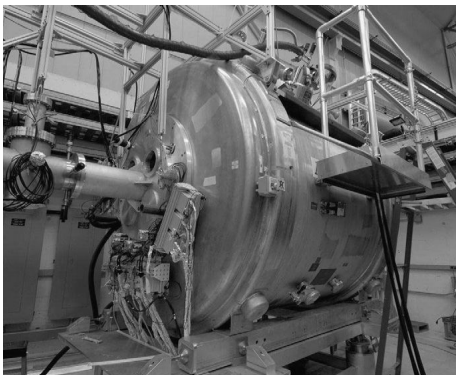
Reaction: $^{85}\text{Kr}(d, p\gamma)$

Beam: ^{85}Kr 10 MeV/u, 10^7 pps

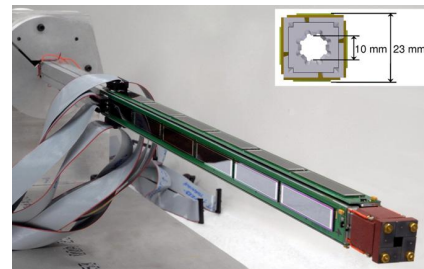
Targets: CD_2



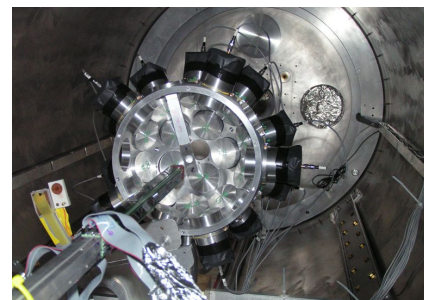
HELIOS: Solenoidal magnetic spectrometer
with $B=2.0$ T



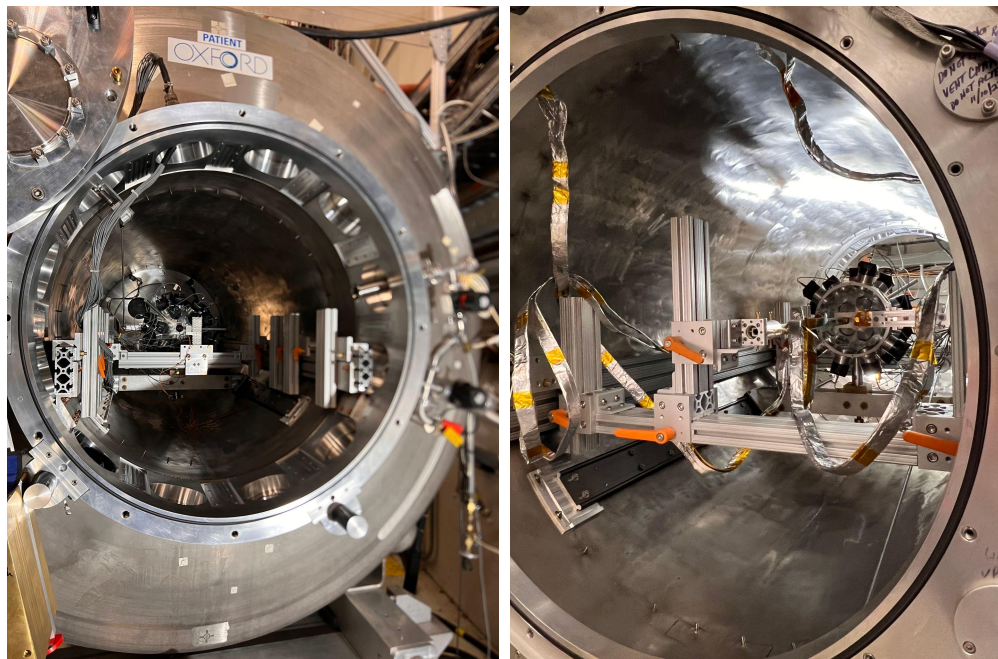
For **protons:** position sensitive Si array



For **γ -rays:** Apollo scintillator array, 5 LaBr + 15 CsI

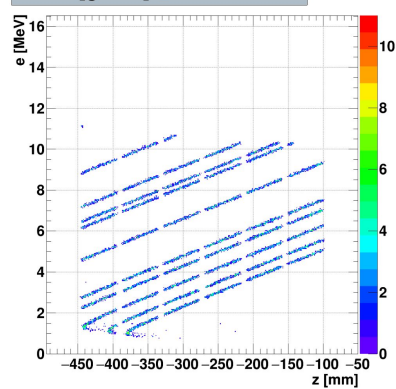


Experimental set-up

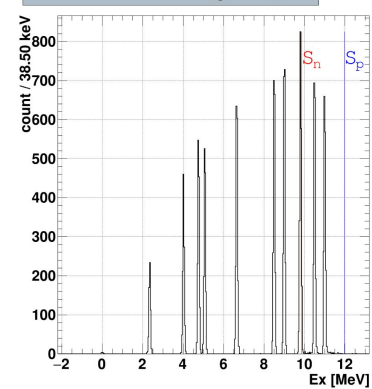


Target-array (1st) distance = 100 mm

e-z [gated] @ -100 mm



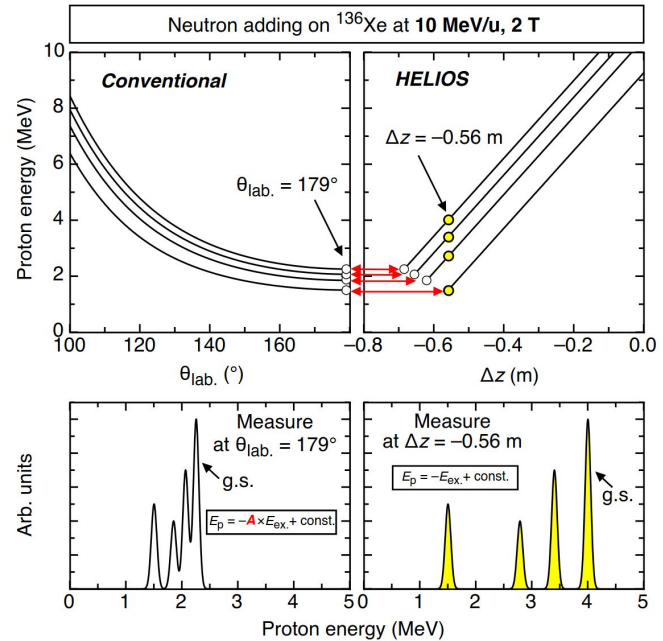
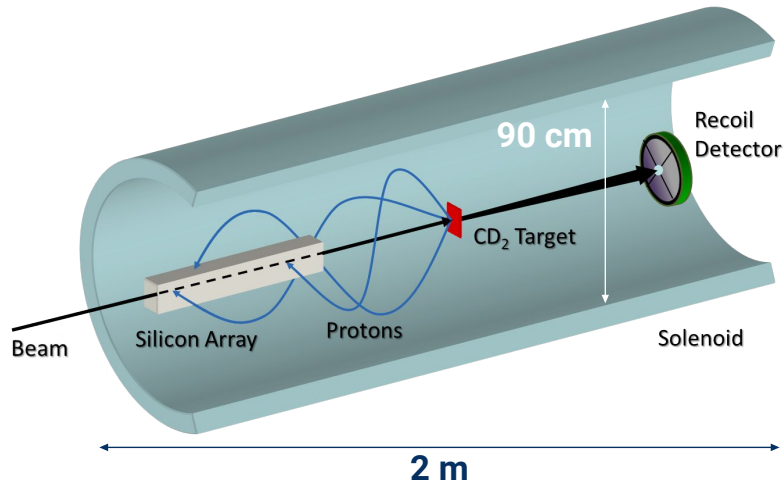
calculated Ex [gated]



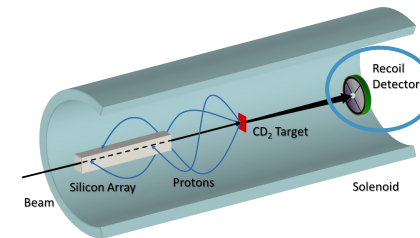
Q-value=7.63 MeV

HELIOS: Helical Orbit Spectrometer

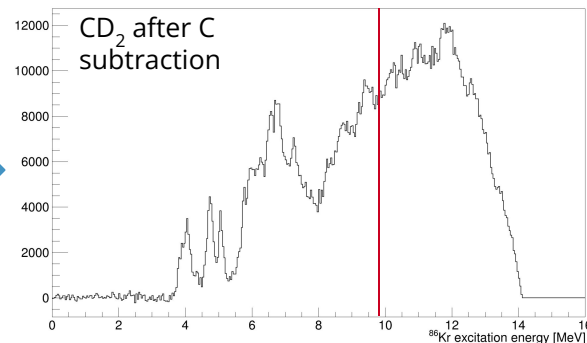
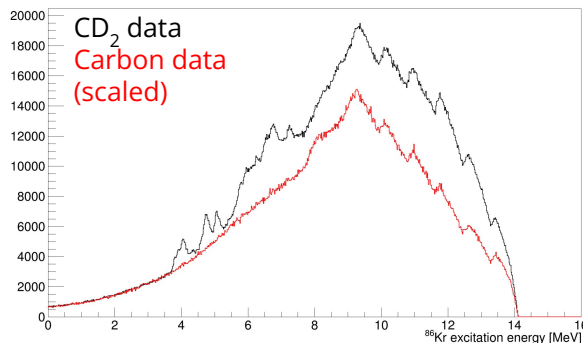
Solves the problem of **kinematic compression**!



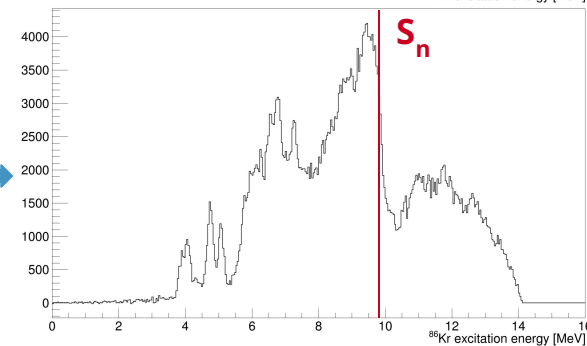
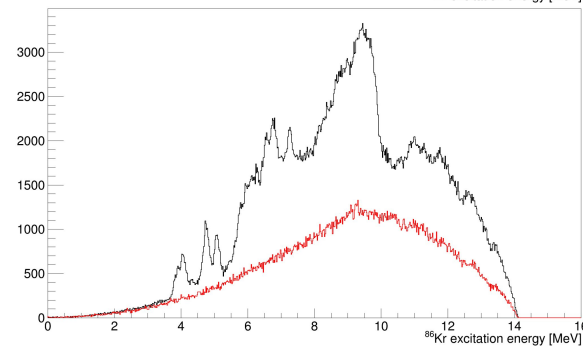
Analysis: C subtraction



Heavy recoils → can't use recoil detector → Need a run with C target to subtract
2 factors: for p only and for p- γ coincidences

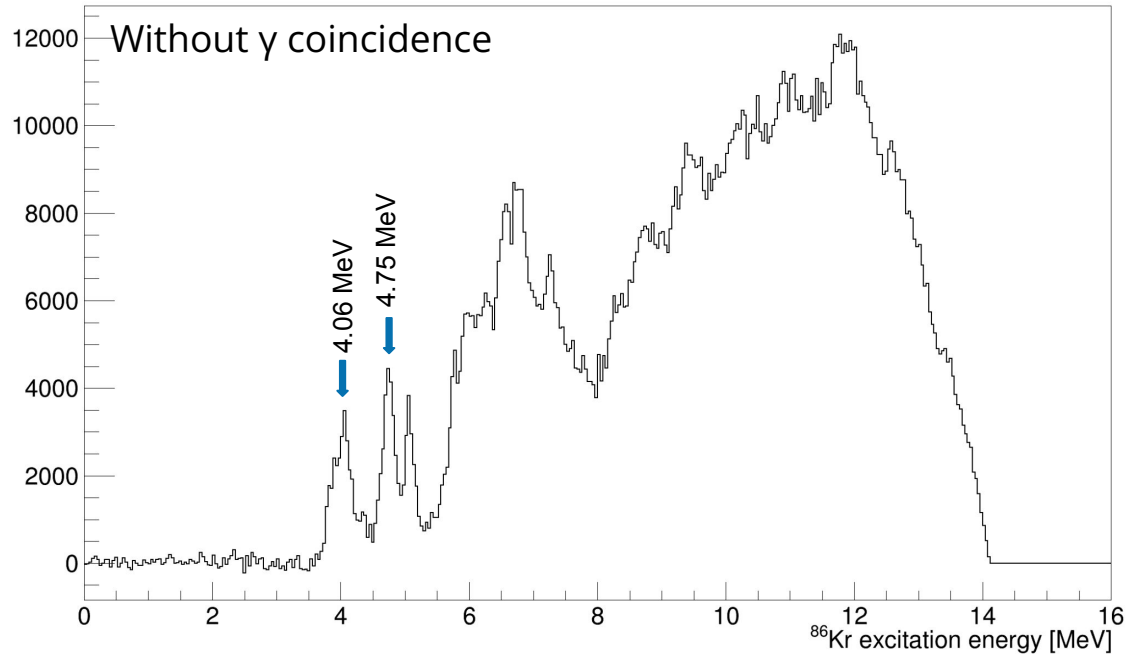


Only protons



p- γ coincidences

Analysis: ^{86}Kr excitation energy spectrum



Q-value FWHM ~ 150 keV

Only a few states already known

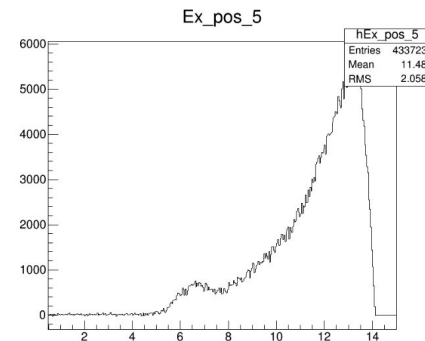
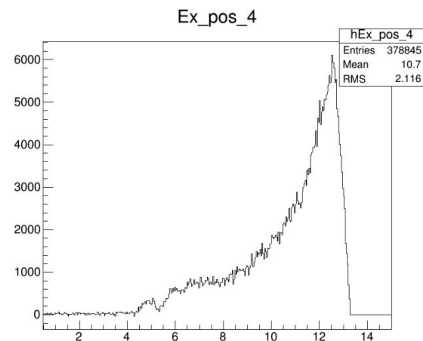
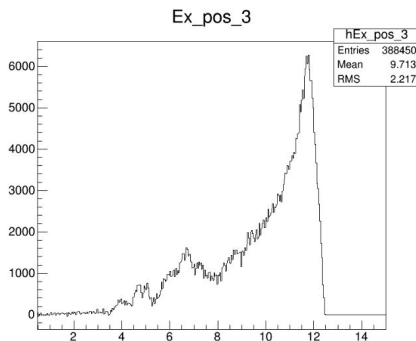
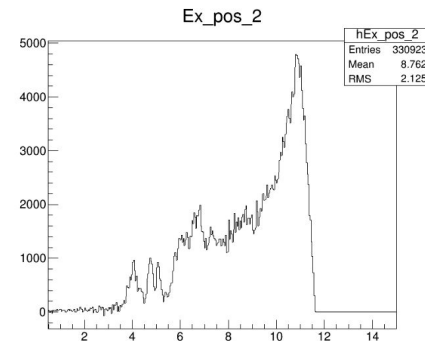
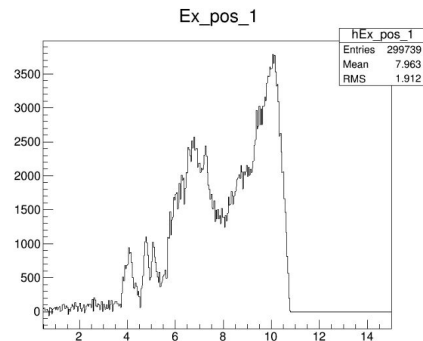
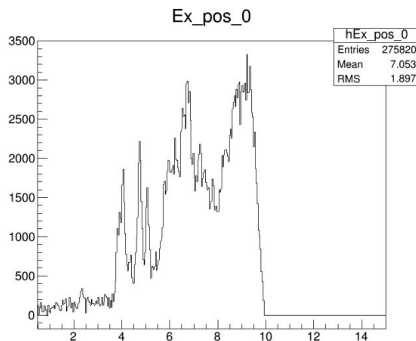
Analysis: angular distributions

Possible to get angular distributions:

Position in silicon array



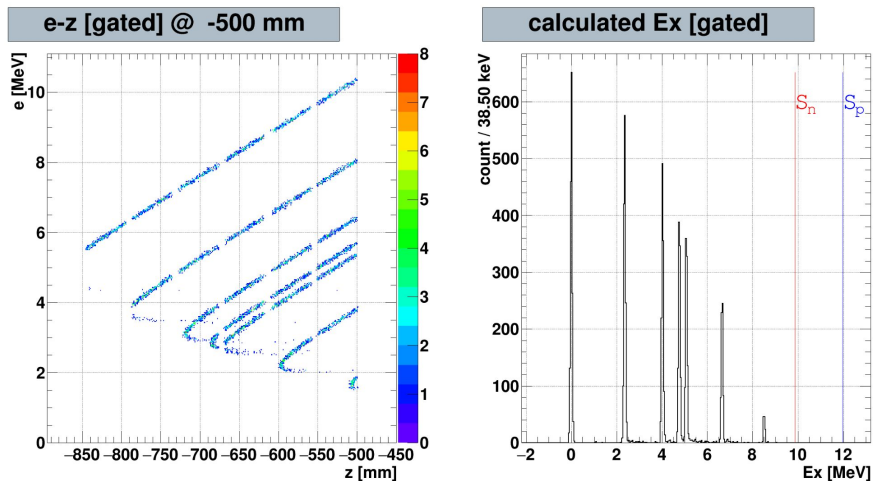
Angle of emission in CM
(and geometrical acceptance)



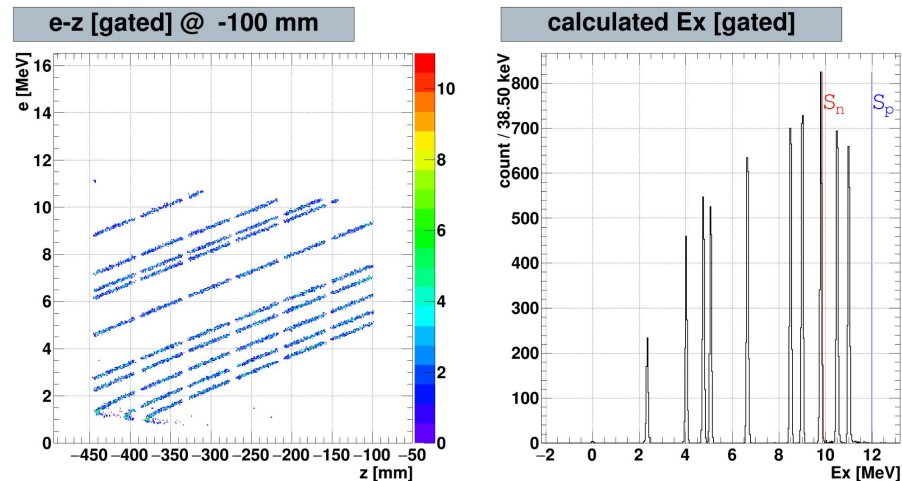
Without γ coincidence

Analysis: 2nd array position

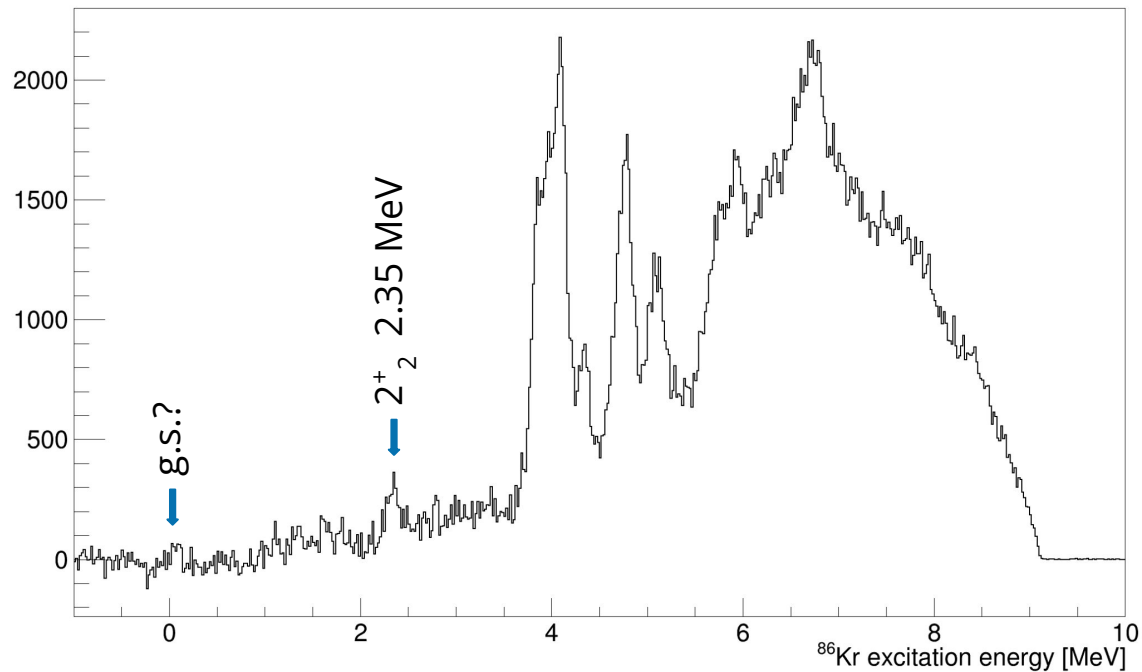
Target-array distance = 500 mm



Target-array distance = 100 mm

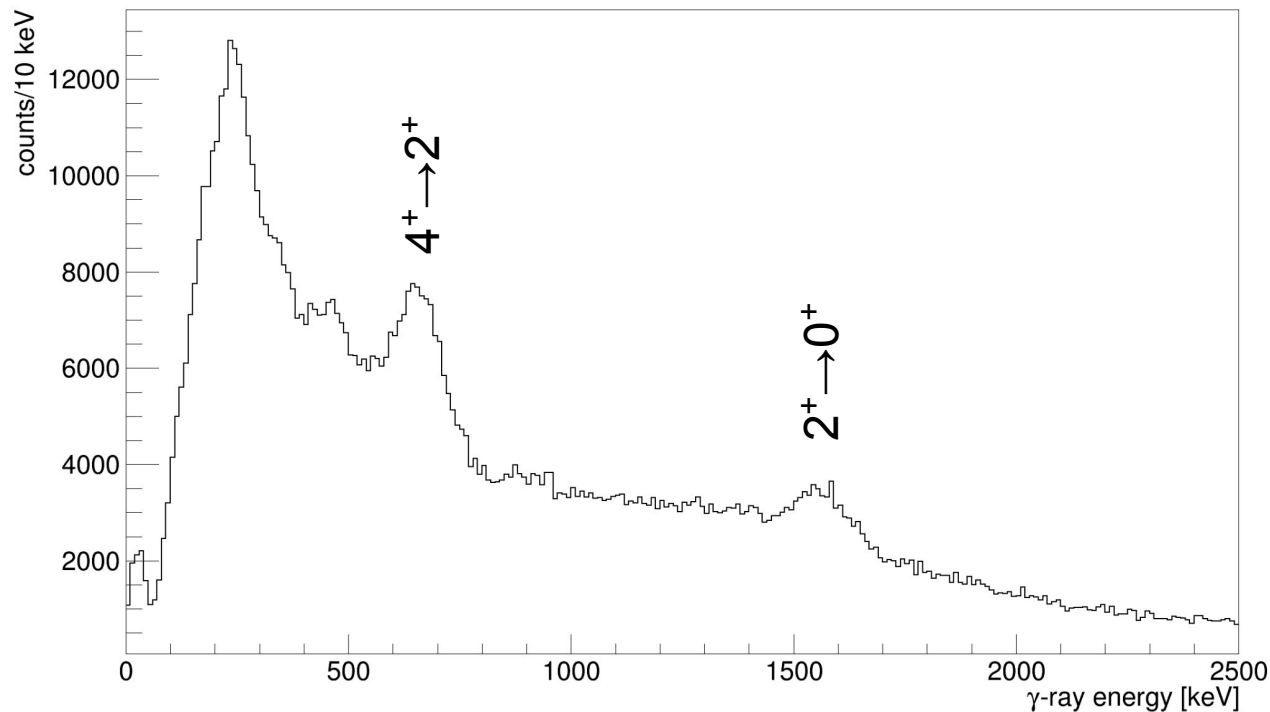


Analysis: 2nd array position



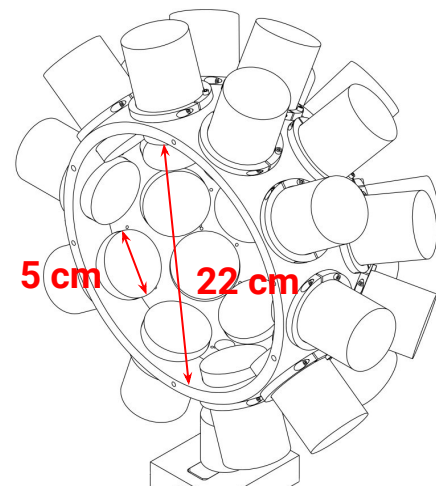
Without γ coincidence

Analysis: γ spectrum

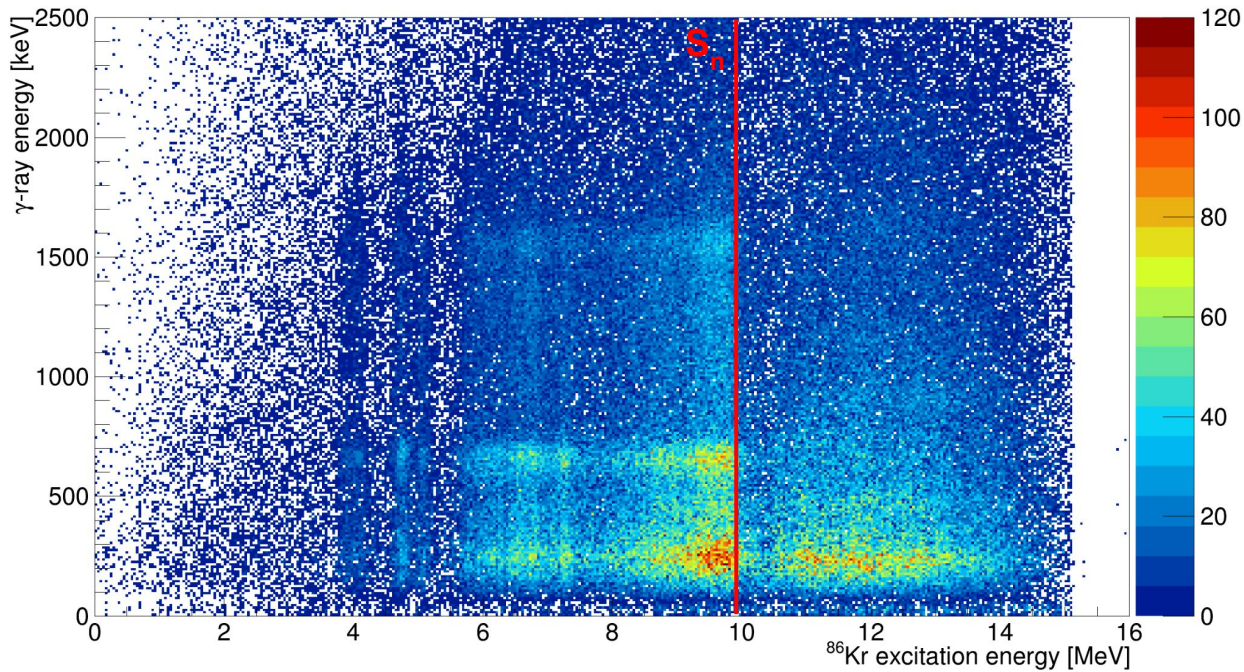


Full statistics

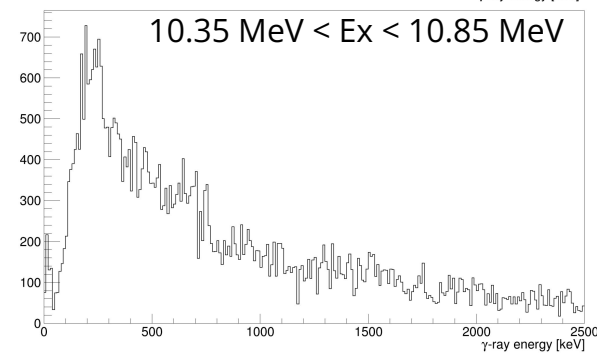
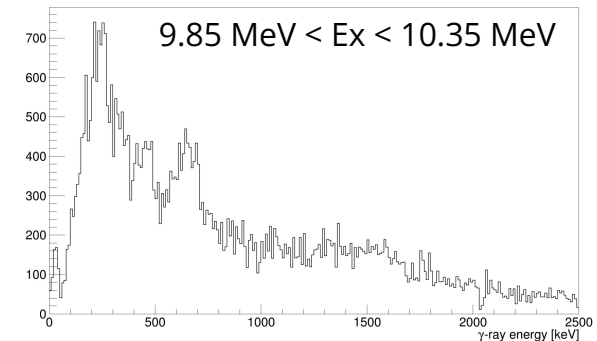
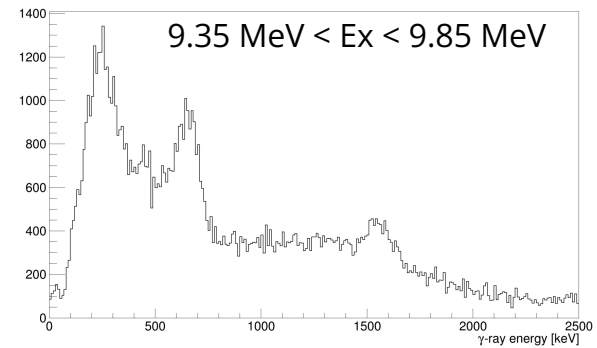
→ Large Doppler broadening



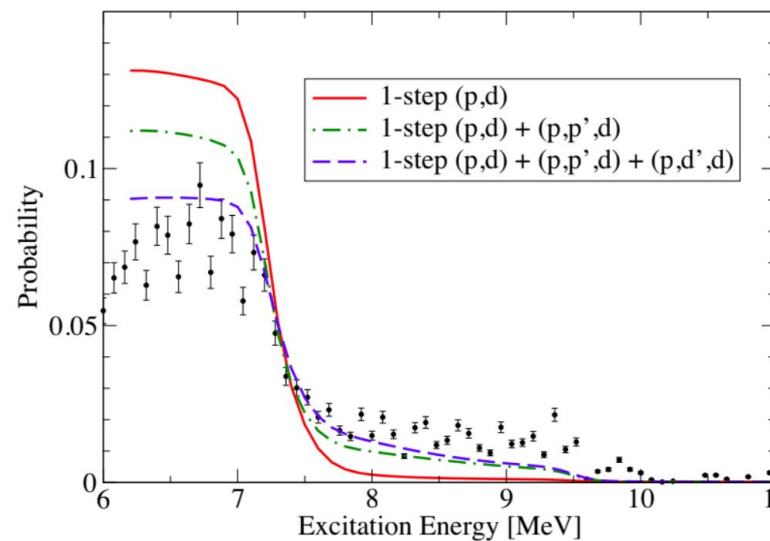
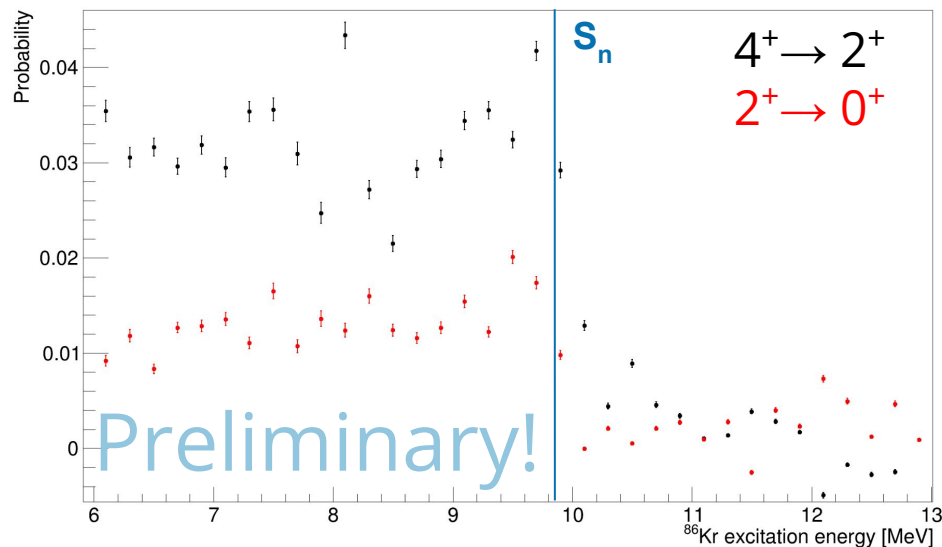
Analysis: γ vs excitation energy



$S_n = 9.85 \text{ MeV}$



Coincidence probability estimation



J. Escher et al. EPJ Web of Conferences 122, 12001 (2016)

Conclusion

- Coincidence between protons and γ s observed ✓
- C subtraction ✓
- First estimation of coincidence probability ✓

Next steps:

- 4^+ is an isomer ($T_{1/2}=3.1$ ns) \rightarrow need a simulation
- (n, γ) conversion

Thank you for your attention!

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