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CUBESATS AND DISTRIBUTED **ASTRONOMY: FROM THE** HERMES FLEET TO THE FLIGHT OF THE ALBATROS, SURFING THE WAVES OF QUANTUM SPACE-TIME

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ON BEHALF OF THE ALBATROS AND THE HERMES COLLABORATIONS

"QUAGRAP meeting" – Napoli – 18-19 April 2024

# SCIENTIFIC CHALLENGES FOR THE NEXT DECADES

# **Multi-Messenger Astronomy**

# **Testing Quantum Gravity theories**





# DEVELOPMENT OF MULTI-MESSENGER ASTRONOMY

**GW170817** 

- NS-NS merging
- Host galaxy NGC 4993
- ~ 40 Mpc
- 70 observatories





### MULTI-MESSENGER PARADOX



We need a high-energy All-sky Monitor with large area to allow Multi–Messenger Astronomy to develop from infancy to maturity!

### MONOLITHIC VS DISTRIBUTED HIGH ENERGY OBSERVATORIES



BeppoSAX

### AGILE



### Pros:

- Reliability
- long heritage



#### Pros:

- modularity
- limited cost
- quicker development
- Low risk

### HIGH ENERGY RAPID MODULAR ENSEMBLE OF SATELLITES - HERMES

# Modular X/gamma-ray ALL-SKY MONITOR

A swarm tens/hundreds of LEO nano/micro satellites equipped with:

- keV-Mev scintillators,
- sub µs time resolution
- large FoV



### IDEAL TARGETS: GAMMA-RAY BURSTS

- sudden and unpredictable bursts of hard-X / soft gamma rays with huge flux
- most of the flux detected from 10–20 keV up to 1–2 MeV,
- bimodal distribution of duration (0.1-1.0 s & 10.0-100.0 s)





# IDEAL TARGETS: GAMMA-RAY BURSTS



#### Crashing neutron stars can make gamma-ray burst jets







21.2 millie

Credit: NASA/AEI/ZIB/M. Koppitz and L. Rezzo

26.5 milliseco

# Long GRBs BH collaps of a massive star

# Short GRBs NS-NS binary system coalescence

# **GRB** – THE FIREBALL MODEL multiple collision of relativistic shells ( $\Gamma = [1 - (v_{iet}/c)^2]^{-1/2} \ge 100$ )

explains rapid variability •

ullet

synchrotron radiation and inverse Compton scattering



Data 40-700 keV (A=1136 cm2, courtesy of F. Frontera)





X-RAYS.

VISIBLE

LIGHT,

RADIO WAVES

### GRB LOCALISATION – TEMPORAL TRIANGULATION (IPN LEGACY)

Determination of source position through delays in Time of Arrival (ToA) of an impulsive (variable) signal over 3 (or more) spatially separate detectors

**GRB** front

position of the source in the sky:  $\alpha$ ,  $\delta$  (2 parameters, N<sub>PAR</sub> = 2)

Number of independent delays:  $N_{DEL} = N_{SATELLITES} \times (N_{SATELLITES} - 1) / 2$ 

Accuracy in determining  $\alpha$  and  $\delta$  with N<sub>SATELLITES</sub>:  $\sigma_{\alpha} \approx \sigma_{\delta} = c \sigma_{ToA} / < baseline > \times (N_{DEL} - N_{PAR} - 1)^{-1/2}$ 



### GRB LOCALISATION – TEMPORAL TRIANGULATION (IPN LEGACY)







#### QUANTUM GRAVITY MINIMAL LENGTH HYPOTHESIS, LIV AND DISPERSION REL<u>ATION FOR PHOTONS IN VA</u>CUO

Existence of a Minimal Length (String theories, etc.)

 $I_{MIN} \approx I_{PLANCK} = [Gh/(2\pi c^3)]^{1/2} = 1.6 \times 10^{-33} \text{ cm}$ 

implies:

i) Lorentz Invariance Violation (LIV): no further Lorentz contraction
ii) Space has the structure of a crystal lattice
iii) Existence of a dispersion law for photons *in vacuo*







Accumulation of delays in light propagation:  $\Delta t_{QG} = \xi \left( D_{TRAV} / c \right) \left[ \Delta E_{obs} / (\zeta E_{PI}) \right]^n$ The distance traveled by photons takes into account the cosmological



The distance traveled by photons takes into account the cosmological expansion:

### $D_{TRAV}(z) = (c/H_0) \int_0^z d\beta (1+\beta) / [\Omega_{\Lambda} + (1+\beta)^3 \Omega_M]^{1/2}$

**DISPERSION RELATION FOR PHOTONS IN VACUO AND TRAVEL TIME DELAYS** 

 $\Omega_{\Lambda}$ : ratio between the energy density due to the cosmological constant and the critical (closure) density of the Universe

 $\Omega_{M}$ : ratio between the energy density due to the matter and the critical (closure) density of the Universe

THE ENERGY AND REDSHIFT DELAY DEPENDENCE

 $\Delta t_{OBS} (E_{OBS}, z) = \Delta t_{INT} (E_{OBS}, z) + \Delta t_{QG} (E_{OBS}, z)$ 

#### Low z

Time lags caused by Quantum Gravity effects:

- $\propto |E_{phot}(Band II) E_{phot}(Band I)|$ 
  - $\propto D_{TRAV}(z_{GRB})$

Time lags caused by prompt emission mechanism:

- complex dependence from E<sub>phot</sub>(Band II) and E<sub>phot</sub>(Band I)
- independent of  $D_{TRAV}(z_{GRB})$

# QUANTUM GRAVITY: OBSERVATIONAL RESULTS



Fermi GBM & LAT detection of short ( $\Delta T < 1$  s) GRB 090510 z = 0.903(3), d = 1.8 × 10<sup>28</sup> cm ( $\Omega_{\Lambda}$  = 0.73,  $\Omega_{M}$  = 0.27, h = 0.71) (Abdo et al. 2009)

"Cleanest" constraints based on one photon detected at 31 GeV  $\Delta t_{31Gev} \le 859 \text{ ms} (+30 \text{ ms} \text{ because GRB started } 30 \text{ ms} \text{ before } 0)$  $\delta t/\delta E \le 30 \text{ ms/GeV} (35 \text{ Mev} - 31 \text{ GeV})$ 

LIV predictions: Relative Locality Models (Freidel, Smolin 2011):  $\xi = \frac{1}{2}$ ; n=1

Data of GRB 090510 imply:  $M_{QG} \ge 0.595 m_{PLANCK}$  ( $\Delta t_{31Gev} \le 859 + 30 ms; E_{ph} \ge 28 \text{ GeV}$ )

Caveats, assumptions:

i) photon at 31 GeV emitted after  $t_{\text{START GRB}} = -30 \text{ ms}$  (not before) ii) physical delays in emission process not considered

#### **VERY LARGE EFFECTIVE AREA**

Voyage 2050 - long term plan in the ESA science programme

# *GrailQuest*: hunting for Atoms of Space and Time hidden in the wrinkle of Space–Time

A swarm of nano/micro/small-satellites to probe the ultimate structure of Space-Time and to provide an all-sky monitor to study high-energy astrophysics phenomena

Contact Scientist: Luciano Burderi

#### **VERY LARGE EFFECTIVE AREA**

 $\sigma_{CCF} \approx 100 \ \mu s/ (N_{PHOT}/12000)^{-1/2} (GRB with ms variability)$ 10000 nano-satellites of A = 100 cm<sup>2</sup>

 $\Lambda CDM$  cosmology with  $\Omega_{\Lambda} = 0.6911$  and  $\Omega_{Matter} = 0.3089$ 

$rac{\mathbf{dN_E(E)}}{\mathbf{dA} \ \mathbf{dt}} = \mathbf{F}  imes \mathbf{F}$	$\int \left(\frac{\mathbf{E}}{\mathbf{E}_{\mathrm{B}}}\right)^{\alpha} \exp\{-(\alpha-\beta)\mathbf{E}/\mathbf{E}_{\mathrm{B}}\}$	$\}, \mathbf{E} \leq \mathbf{E}_{\mathrm{B}},$
	$\left( \left( \frac{\mathbf{E}}{\mathbf{E}_{\mathrm{B}}} \right)^{eta} \exp\{-(lpha - eta)\},  ight.$	$\mathbf{E} \geq \mathbf{E}_{\mathrm{B}}.$

Energy band	$\mathbf{E}_{\mathbf{AVE}}$	$\mathbf{N}$	$\mathbf{E_{CC}}(\mathbf{N})$	$\mathbf{N}$	$\mathbf{E_{CC}}(\mathbf{N})$	$\Delta$	${f \Delta T}_{ m LIV}~(\xi=1.0,~\zeta=$		
${f MeV}$	${f MeV}$	$egin{array}{l} (eta=-2.5) \ {f photons} \end{array}$	$\mu {f s}$	$egin{array}{l} (eta=-2.0) \ {f photons} \end{array}$	$\mu {f s}$	$\mu {f s}$	$\mu {f s}$	$\mu {f s}$	$\mu {f s}$
						$\mathbf{z} = 0.1$	$\mathbf{z} = 0.5$	z = 1.0	z = 3.0
0.005 - 0.025	0.0112	$3.80  imes \mathbf{10^6}$	0.38	$\mathbf{3.02  imes 10^6}$	0.43	0.04	0.25	0.51	1.42
0.025 - 0.050	0.0353	$1.40  imes \mathbf{10^6}$	0.62	$f 1.17 imes 10^6$	0.69	0.13	0.72	1.46	4.10
0.050 - 0.100	0.0707	$1.10 imes 10^6$	0.71	$9.98  imes \mathbf{10^5}$	0.74	0.27	1.43	2.93	8.21
0.100 - 0.300	0.1732	$\mathbf{8.98  imes 10^5}$	0.79	$f 1.00 imes f 10^6$	0.74	0.66	<b>3.51</b>	7.19	20.10
0.300 - 1.000	0.5477	$f 2.07 imes 10^5$	1.64	$\mathbf{3.82  imes 10^5}$	1.20	2.09	11.11	22.72	63.56
1.000 - 2.000	<b>1.4142</b>	$f 2.63 imes 10^4$	${\bf 4.56}$	$8.20  imes \mathbf{10^4}$	2.60	5.40	28.68	58.67	${\bf 164.12}$
2.000 - 5.000	3.1623	$f 1.07 imes 10^4$	7.19	$f 4.92 imes 10^4$	3.35	12.07	64.12	131.19	367.00
5.000 - 50.00	15.8114	${f 3.52 imes10^3}$	$\bf 12.54$	$2.95  imes \mathbf{10^4}$	4.33	<b>60.35</b>	320.62	656.00	1834.98

#### LARGE CONSTELLATIONS





Starlink Constellation 12,000 sats SpaceX (Elon Musk)

- 6000 @ 1200 km (by March 2024)
- 60 satellites launched on 16/05/2019
- LEO @~550 km
- optical inter-satellite links
- 100 ÷ 500 kg satellites (mass production) WHISHFUL THINKING:
  board a 100 cm<sup>2</sup> effective area GAGG crystal
  SDD photodetector (position sensitive + coded mask?) module on each satellite. 120 m<sup>2</sup> effective area All Sky Monitor!

#### THE ENERGY AND REDSHIFT DELAY DEPENDENCE

#### **GRB intrinsic spectral lag**

**Observer frame vs Rest frame** 

Lags come from GRB central engine emission mechanism at the rest frame:  $E_{rf} = E_{obs} (1+z)$ 

Moreover, cosmological dilation implies:  $\Delta t_{INT} (E_{OBS}, z) = \Delta t_{INT} (E_{rf}) (1+z)$ 

There is no direct evidence of a correlation between  $\Delta t_{INT}(E_{rf})$  and redshift z.

#### **Quantum Gravity delay**

**Observer frame vs Rest frame** 

$$\Delta t_{QG} (E_{OBS}, z) = \Delta t_{QG} (E_{rf}, z) = \frac{\xi}{H_0} \left(\frac{E_{rf}}{\zeta E_{pl}}\right)^n \times$$

 $\times \left(\frac{1+n}{2}\right) \left(\frac{1}{1+z}\right)^n \int_0^z \frac{(1+z')^n dz'}{\sqrt{\Omega_m (1+z')^3 + \Omega_\Lambda}}$ 

THE ENERGY AND REDSHIFT DELAY DEPENDENCE

Total observed spectral lag:

$$\Delta \mathbf{t}_{\text{total}} \left( \mathsf{E}_{rf}; \mathsf{z} \right) = \Delta t_{INT} \left( E_{rf} \right) \left( 1 + z \right) + \frac{\xi}{H_0} \left( \frac{E_{rf}}{\zeta E_{pl}} \right)^n \left( \frac{1 + n}{2} \right) \left( \frac{1}{1 + z} \right)^n \int_0^z \frac{(1 + z')^n dz'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + n}{2} \right) \left( \frac{1}{1 + z} \right)^n \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}} \right)^{n-1} \left( \frac{1 + z'}{\sqrt{\Omega_m (1 + z')^3 + \Omega_N z'}$$

Let's define the quantity:

$$au_{total}(E_{rf};\mathbf{z}) = rac{\Delta t_{total}(E_{rf};\mathbf{z})}{(1+\mathbf{z})} =$$

$$= \Delta t_{INT}(E_{rf}) + \frac{\xi}{H_0} \left(\frac{E_{rf}}{\zeta E_{pl}}\right)^n \left(\frac{1+n}{2}\right) \left(\frac{1}{1+z}\right)^{n+1} \int_0^z \frac{(1+z')^n dz'}{\sqrt{\Omega_m (1+z')^3 + \Omega_\Lambda}}$$

THE ENERGY AND REDSHIFT DELAY DEPENDENCE

We now define the function u(z):

$$u(z) = \left(\frac{1+n}{2}\right) \left(\frac{1}{1+z}\right)^{n+1} \int_0^z \frac{(1+z')^n dz'}{\sqrt{\Omega_m (1+z')^3 + \Omega_\Lambda}}$$

This formalism allows us to express the lags as follows:

$$\tau_{total}(E_{rf};z) = \Delta t_{INT}(E_{rf}) + \frac{\xi}{H_0} \left(\frac{E_{rf}}{\zeta E_{pl}}\right)^n u(z)$$

#### **OBSERVATIONAL HINTS ON GRB'S INTRINSIC DELAYS**



#### **THE POWER OF THE SAMPLE**

Let's consider a sample of N GRBs with known z and let's define a number M of common energy bands in the rest frame of the GRBs.

For each energy band  $E_{rf,i}$  we can fit the function

$$\tau_{total}(E_{rf,i}; \mathbf{z}) = \Delta t_{INT}(E_{rf,i}) + \frac{\xi}{H_0} \left(\frac{E_{rf,i}}{\zeta E_{pl}}\right)^n u(\mathbf{z})$$

Obtaining *M* values of the quantity

$$\eta(E_{rf,i}) = \frac{\xi}{H_0} \left(\frac{E_{rf,i}}{\zeta E_{pl}}\right)^2$$



#### **THE POWER OF THE SAMPLE**

Finally, we define the new variable

$$s(E_{rf}) = \frac{1}{H_0} \left(\frac{E_{rf}}{E_{pl}}\right)^n$$

Therefore, we can linearly fit the *M* sample  $\eta(E_{rf})$  with the function:

$$f(E_{rf}) = \left(\frac{1}{\zeta}\right)^n s(E_{rf}) = \Delta_{QG} s(E_{rf})$$

Interestingly, the uncertainty on  $\Delta_{QG}$  improves as



# THE ALBATROS MISSION

(Astonishingly Large Baseline Array Transient Reconnaissance Observatory from Space)



#### **Properties:**

- 3 satellites in heliocentric orbits
- 2x400 cm<sup>2</sup> detectors (~20 kg) per satellite pointing in opposite directions
- keV MeV energy band
- Sub-microsecond time resolution
- $4\pi$  steradians FoV (whole sky)
- 1 GRB/day detection rate
- 75% expected success in GRB redshift determination with ground-based facilities follow-up

# THE ALBATROS MISSION: CART-WHEEL ORBITS



3 satellites in "Cart-wheel" orbits (e.g., LISA orbits):

- 3 heliocentric orbits with a=1AU
- 3 slightly different small inclinations (i≈degrees) w.r.t. to eclipting plane
- Equatorial triangle of side  $D \approx 2.5 \ 10^6 \text{ km}$
- Contact to ground up to 23 hours per day
- Wet mass ~ 230 kg per satellite
- Dry mass ~ 165 kg per satellite

# THE ALBATROS MISSION: LOCALIZATION CAPABILITIES

Determination of source position through Delays in Time of Arrival (ToA) of an impulsive event (variable signal) over 3 (or more) spatially separate detectors

Transient source in the sky defined by time of the event, position in the sky:  $T_0$ ,  $\alpha$ ,  $\delta$  (3 parameters,  $N_{PAR} = 3$ )

Statistical accuracy in determining  $\alpha$  and  $\delta$  with N<sub>SAT</sub>:

 $\sigma_{\alpha} \approx \sigma_{\delta} = c \sigma_{ToA} / \langle baseline \rangle \times (N_{SAT} - N_{PAR})^{-1/2}$ 



 $\sigma_{\alpha} \approx \sigma_{\delta} \approx c \sigma_{ToA} / B \approx 24 \text{ arcsec} \times (B/2.5 \times 10^{6} \text{ km})^{-1} \times (\sigma_{ToA} / 1 \text{ ms})$ 

# HERMES PATHFINDER + SPIRIT IN A NUTSHELL

### In orbit demonstrator

- HERMES Pathfinder: six 3U cubesat equipped with advanced X-ray/gamma/ray wide field detector; funded by ASI & EC H2020
- SpIRIT: 6U cubesat managed by University of Melbourne and funded by ASA. Host 1 HERMES-PF
   X-ray/gamma-ray payload + S-band system.



# HERMES PATHFINDER: PAYLOAD





#### Silicon Drift Detector + GAGG Cristals

Scintillator Crystal size: 0.7×1.2×1.5 cm Crystal type: Photo detector: Energy range: Energy resolution: Effective area: FOV: Temporal resolution: Mass: Volume

60 GAGG crystals 120 SDD (1x0.5 cm) 3 keV ÷ ≥ 0.5 Mev ~ 10% at 30 keV  $\sim 56 \, {\rm cm}^2$ ~ 3 steradians (FWHM) ~ 0.5 µs ~1.5 kg < 10×10×12.5 cm



### HERMES PATHFINDER & SPIRIT PAYLOAD FAMILY PICTURE



FMI

**SpIRIT** 

PFM



# SPIRIT: STATUS UPDATE

- Launch! December 1st, 2023, OK
- Deployment OK, Detumbling OK
- Safe mode just after detumbling because of a failure of a magnetorque
- Communications OK, UHF beacon received by SatNogs
- Communications OK, about 15m per day of link with AU GS
- OBC OK



Communications between PDHU - PMS and MMS established. PDHU OK

# HERMES PATHFINDER: STATUS UPDATE

- •PFM fully integrated, environmental test to support qualification review started in Dec. 2023 @POLIMI labs. QR closed in Jan 2024
- •FM1, integration started end Nov. 2023, and full integration by Feb 2024.

•Acceptance review for six FMs second half of 2024

•Launch early 2025



## THE HERMES PROJECT: THE MOVIE



# **Thanks for the attention!**

Please join the HERMES Science Team: https://www.hermes-sp.eu/?page\_id=3643#ScienceTeam