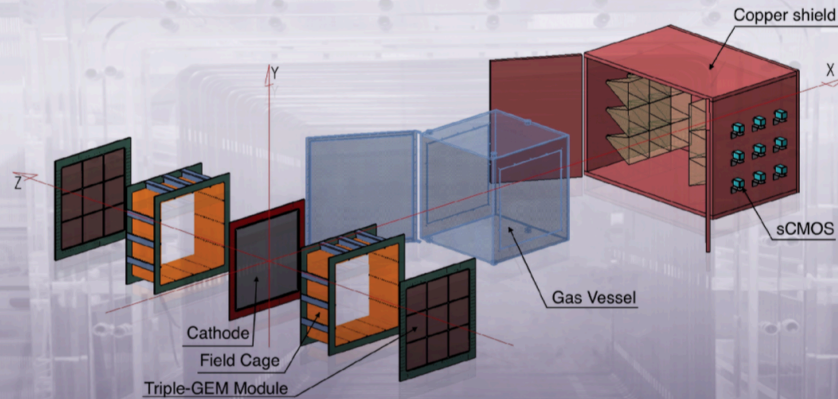




instruments



The CYGNO Experiment

Volume 6 · Issue 1 | March 2022

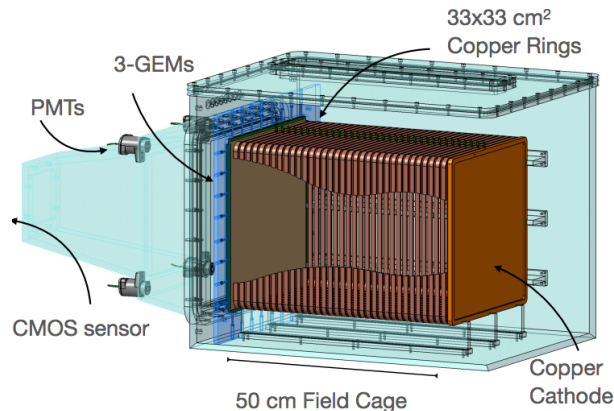
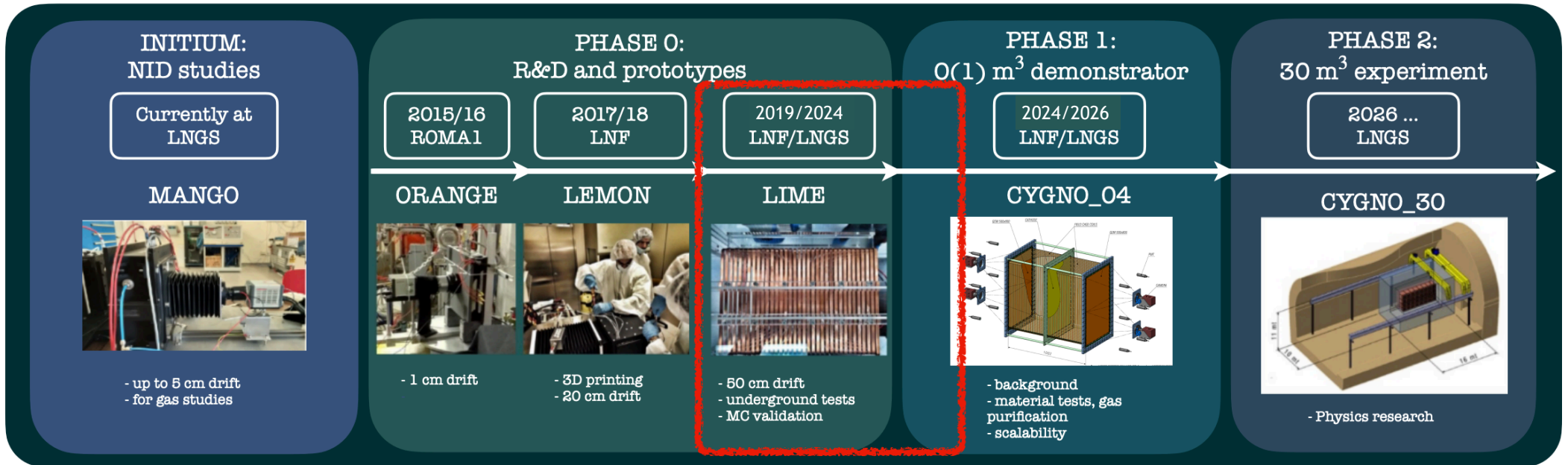


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JINST 15 (2020) 12, T12003
JINST 15 (2020) P08018
Measur.Sci.Tech. 32 (2021) 2, 025902

JINST 15 (2020) P10001
2019 JINST 14 P07011
NIM A 999 (2021) 165209



1 sCMOS + 4 PMT + 3 GEMs
33 x 33 cm² readout area
50 cm drift length
50 L active volume

LIME staged shielding underground campaign concludes today

Water shielding being dismantled to perform extended Run5 with 10 cm Cu only for underground neutron flux measurement (PRIN 2017)

	Shielding	Number of bkg pictures	Event rate	Period
Run1	none	4×10^5	35 Hz	Oct 2022
Run2	4 cm Cu	4.5×10^5	3.5 Hz	Jan-Mar 2023
Run3	10 cm Cu	2.7×10^6	1.3 Hz	May-Nov 2023
Run4	10 cm Cu + 40 cm H ₂ O	2.8×10^6	0.9 Hz	Dec 2023-Apr 2024

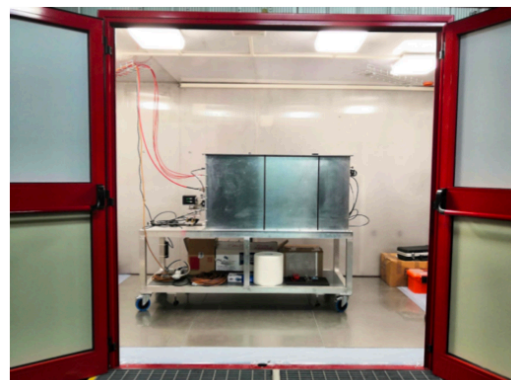
+ periodic calibrations with ⁵⁵Fe (as a function of distance from GEM from Run2)

+ AmBe measurement during Run3

+ ¹³³Ba, ¹⁵²Eu, ²⁴¹Am calibrations during Run3



Underground installation with full auxiliary systems



Run1



Run2 - Run3



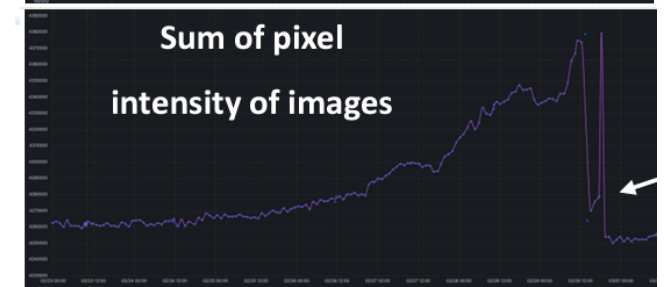
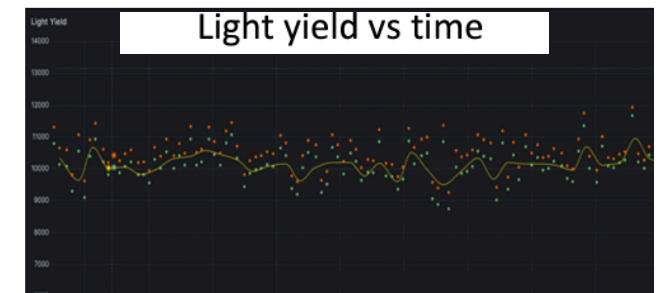
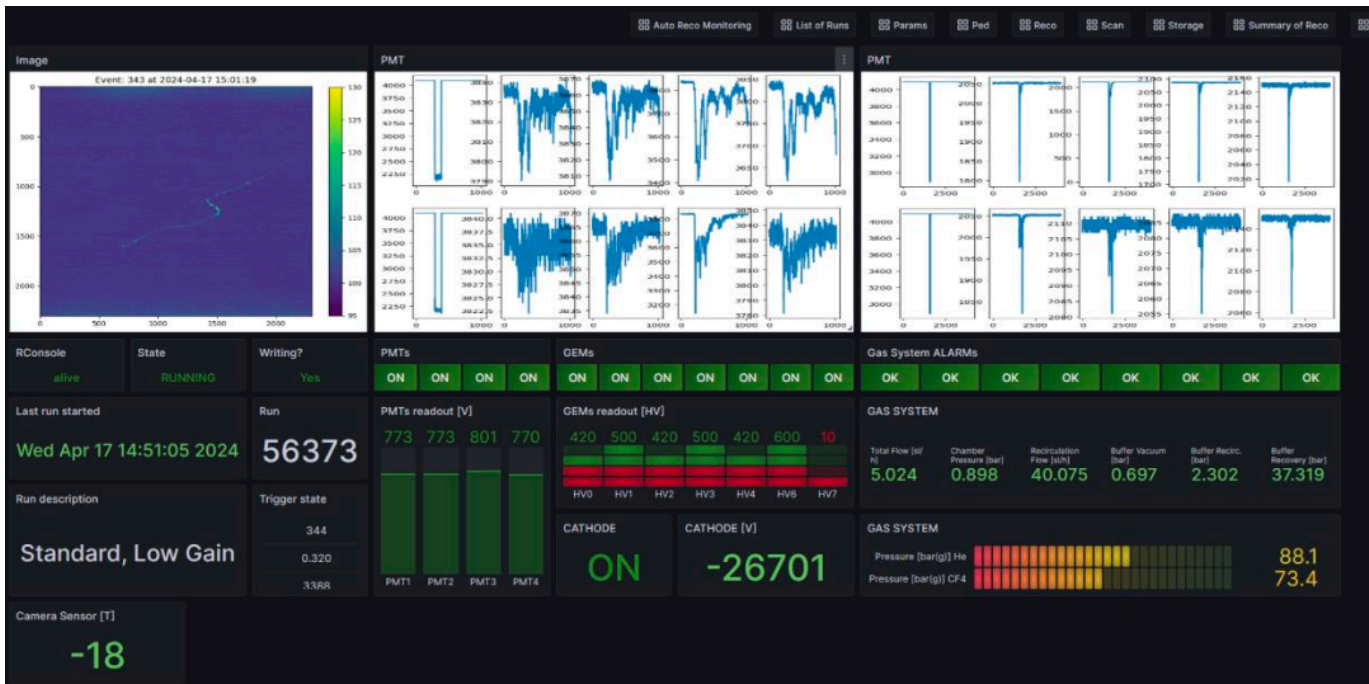
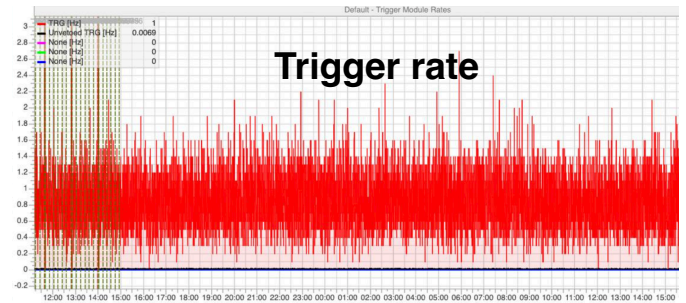
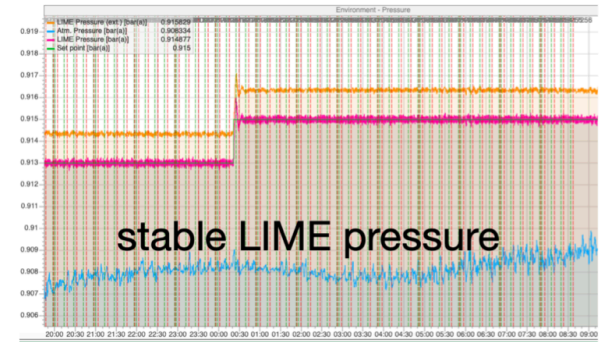
Run4



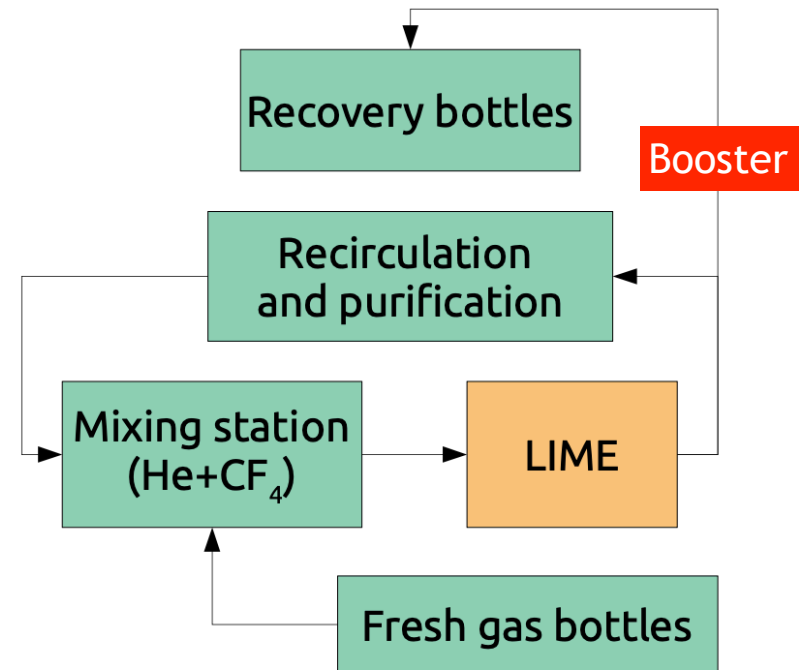
LIME underground operation with full auxiliary system configuration



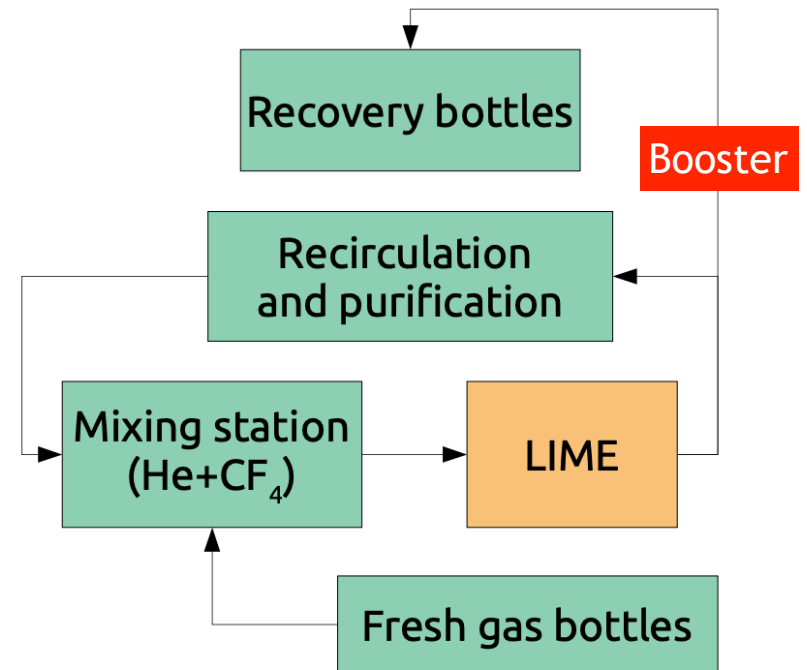
- Automated system developed to control and monitor remotely HV, gas system, environmental parameters, DAQ, trigger rate, ^{55}Fe calibrations, detector conditions and data taking
- Automatic data reconstruction implemented
- Complementary Grafana online monitor for fast interventions to critical issues
- Fully remote shifts 24/7 from Run4



- 📌 Operations started without recirculation to assess detector conditions independently from it (Run1 & Run2)
- 📌 Recirculation introduced at end of Run3
 - 📌 Leakages in the line induced LY instabilities due to humidity in the gas
 - 📌 Pump electronic board failures required periodic halts to data taking and interventions
 - 📌 Booster failures additionally prevented regular operations



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 - 📌 Leakages in the line induced LY instabilities due to humidity in the gas
 - 📌 Pump electronic board failures required periodic halts to data taking and interventions
 - 📌 Booster failures additionally prevented regular operations
- 📌 Interventions in September 2023 allowed to fix all pump and booster instabilities and main leaks down to:
 - 📌 0.5 cc/min leak in LIME
 - 📌 < 3 cc/min in the rest of the system
- 📌 Thanks to stability and improved gas system operations, optimisation of recirculation configuration performed in Oct-Nov 2023
- 📌 **Stable operation and high quality operations achieved in full gas system configuration from December 2023**





LIME operations in Run4



Uptime



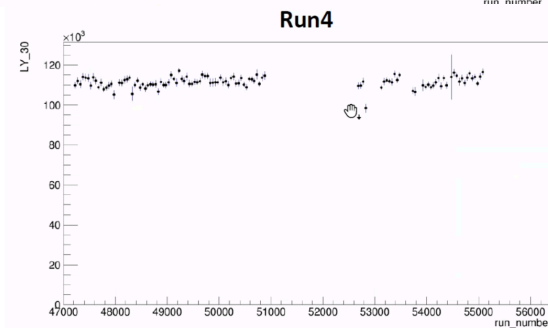
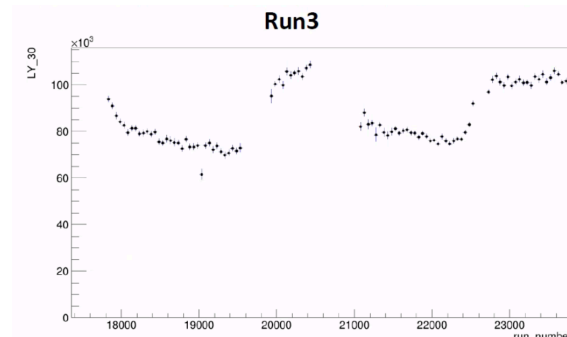
**Thanks to the interventions
on the gas system, an
uptime > 95% was achieved
for nearly 4 consecutive
months in Run4**

Uptime



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Light Yield



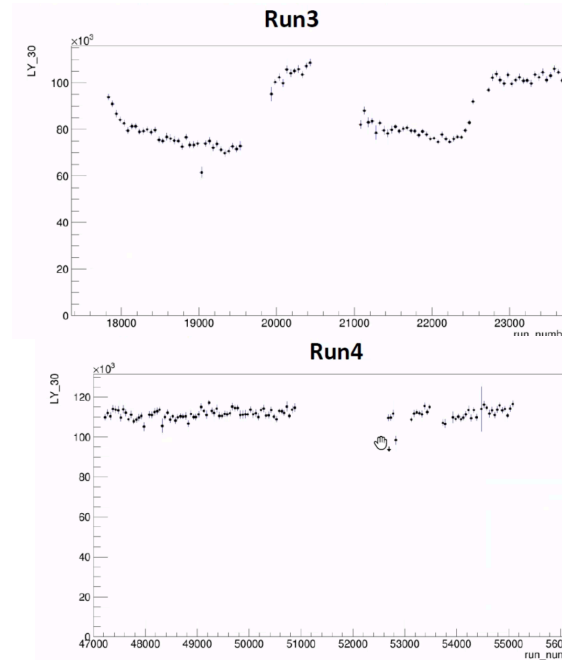
Thanks to the tests and optimisation of recirculation and filtering at the end of Run3, in Run4 we achieved a stable (within 5%) and high (2 counts/eV) LY

Uptime



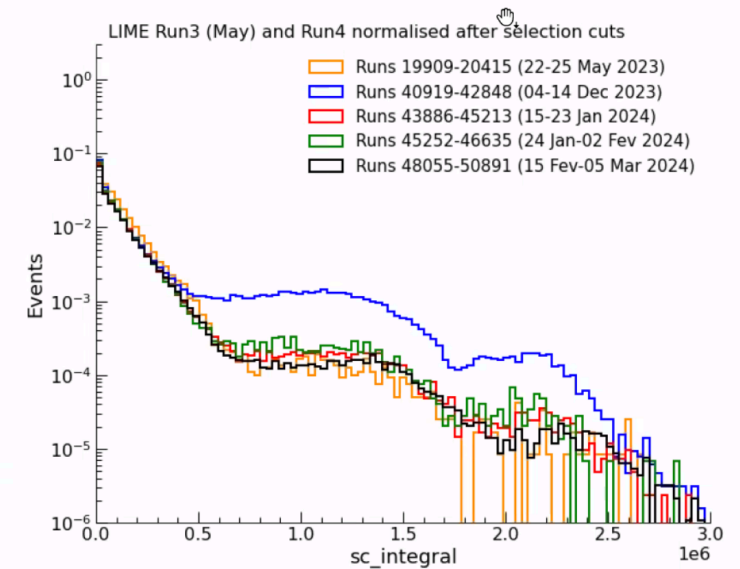
Thanks to the interventions on the gas system, an **uptime > 95%** was achieved for nearly 4 consecutive months in Run4

Light Yield



Thanks to the tests and optimisation of recirculation and filtering at the end of Run3, in Run4 we achieved a **stable (within 5%) and high (2 counts/eV) LY**

Alpha rate



Thanks to the tests and optimisation of recirculation and filtering at the end of Run3, in Run4 the alpha rate (likely related to Rn contamination in the gas, see later) was reduced to the same level of no recirculation

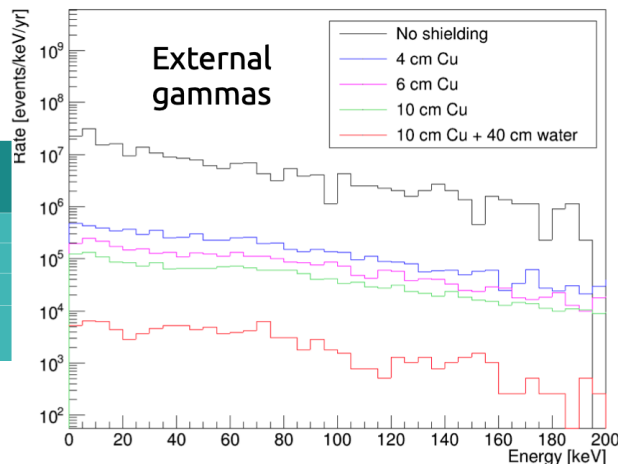


LIME underground background data analysis and comparison with MC prediction

Run4 is not discussed in this presentation since data analysis still ongoing

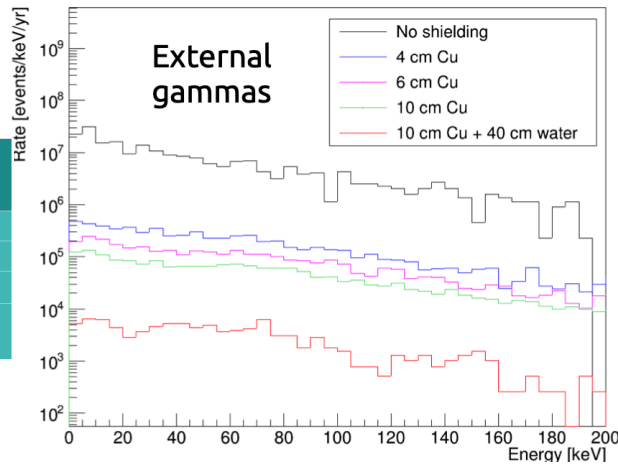
External backgrounds

Shielding	Gamma background [10 ⁶ ER yr ⁻¹]	Neutron background [NR yr ⁻¹]
Unshielded	(1140±30)	(1480±90)
4 cm Cu	(26.2±0.6)	(870±10)
6 cm Cu	(9.4±0.3)	(1000±30)
10 cm Cu + 40 cm H ₂ O	(0.5±0.2)	(2.0±0.2)

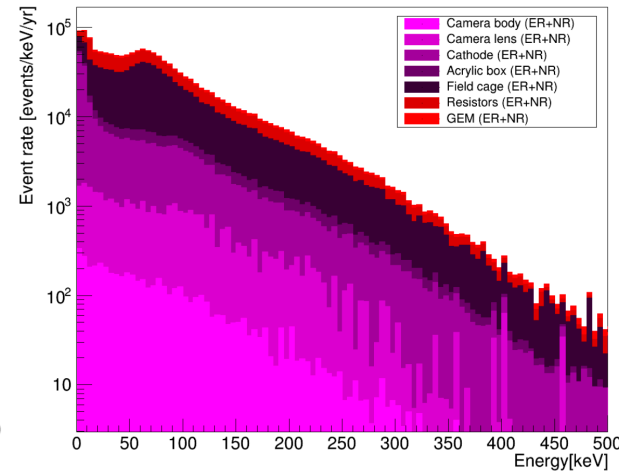


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Internal backgrounds

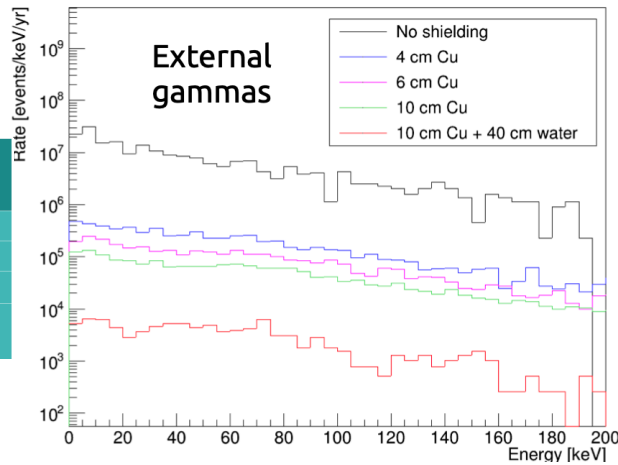


Source	Event rate [10 ⁶ yr ⁻¹]
Field cage	(3.57±0.01)
Resistors	(1.873±0.006)
Cathode	(1.095±0.001)
GEMs	(0.3891±0.0002)
Vessel	(0.268±0.001)
Camera lens	(0.151±0.004)
Camera body	(0.0242±0.0005)
TOTAL	(7.34±0.01)

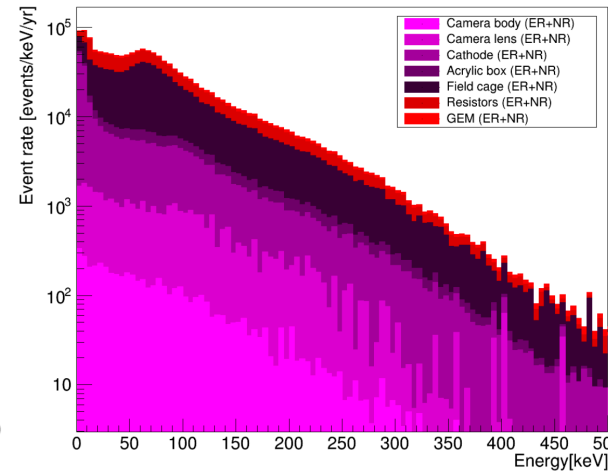
Please note LIME was NOT built with radioactive pure components

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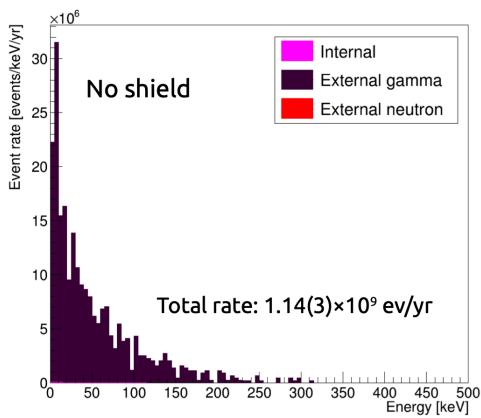
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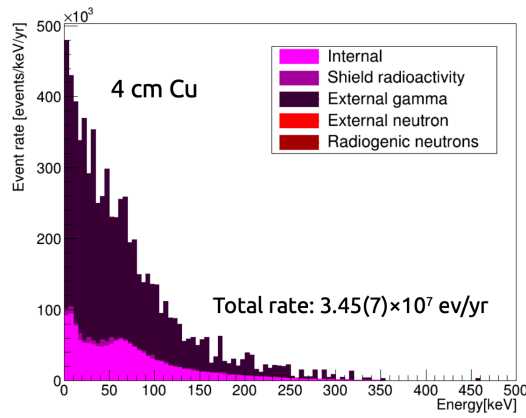
Please note LIME was NOT built with radioactive pure components

Total backgrounds



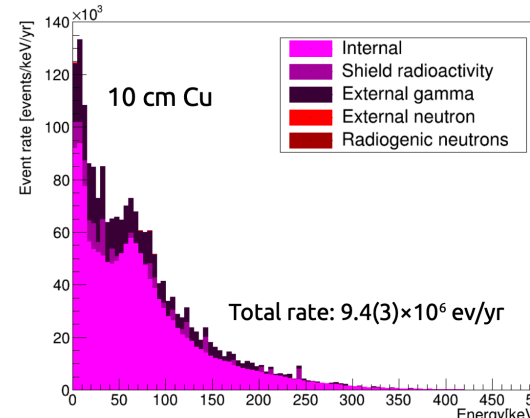
internal/external bkg

0.6%



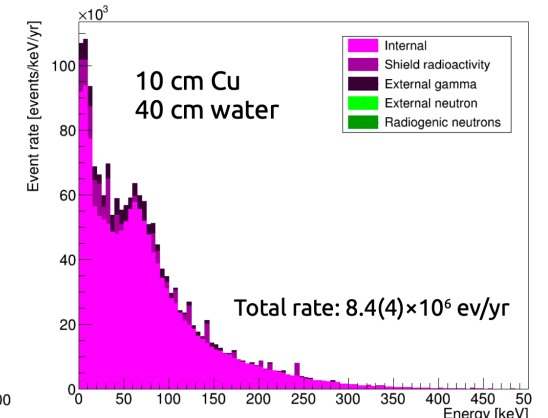
internal/external bkg

20%



internal/external bkg

78%

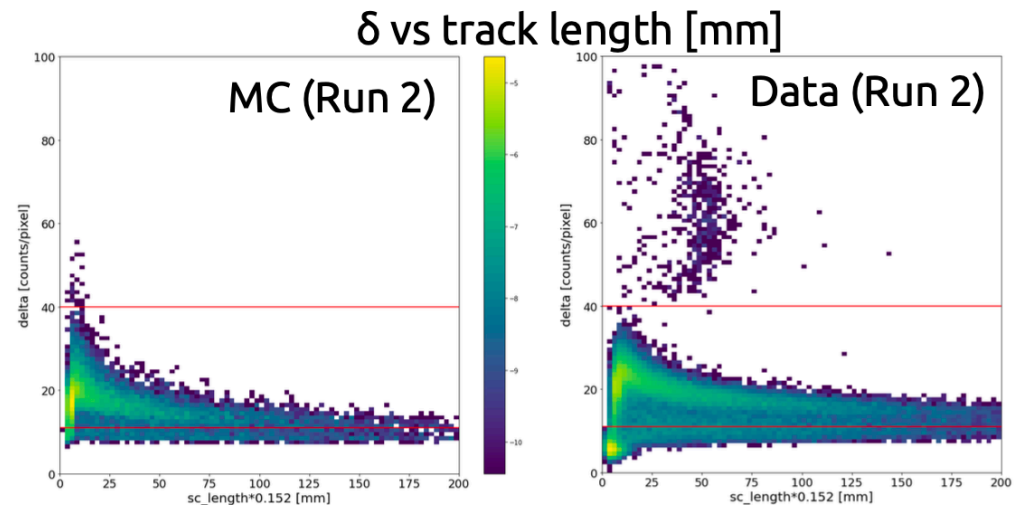
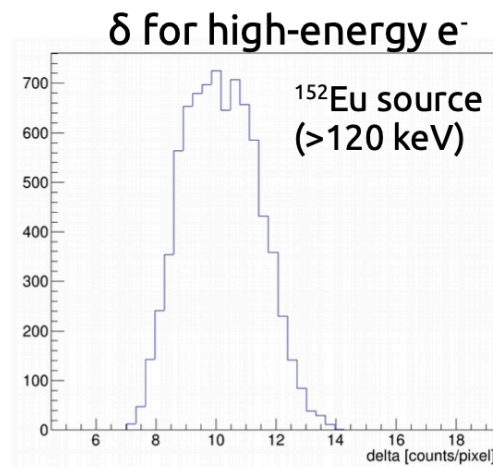
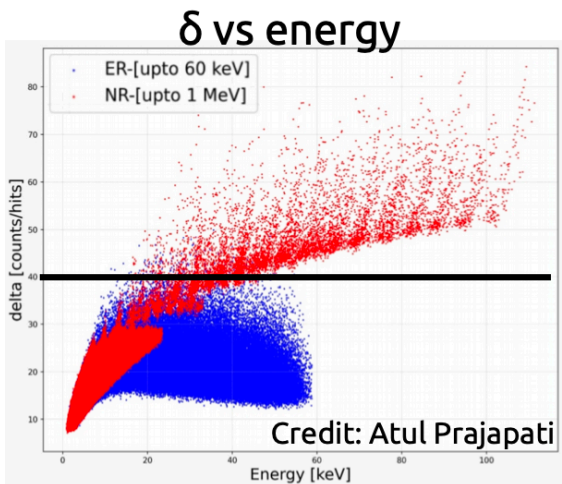


internal/external bkg

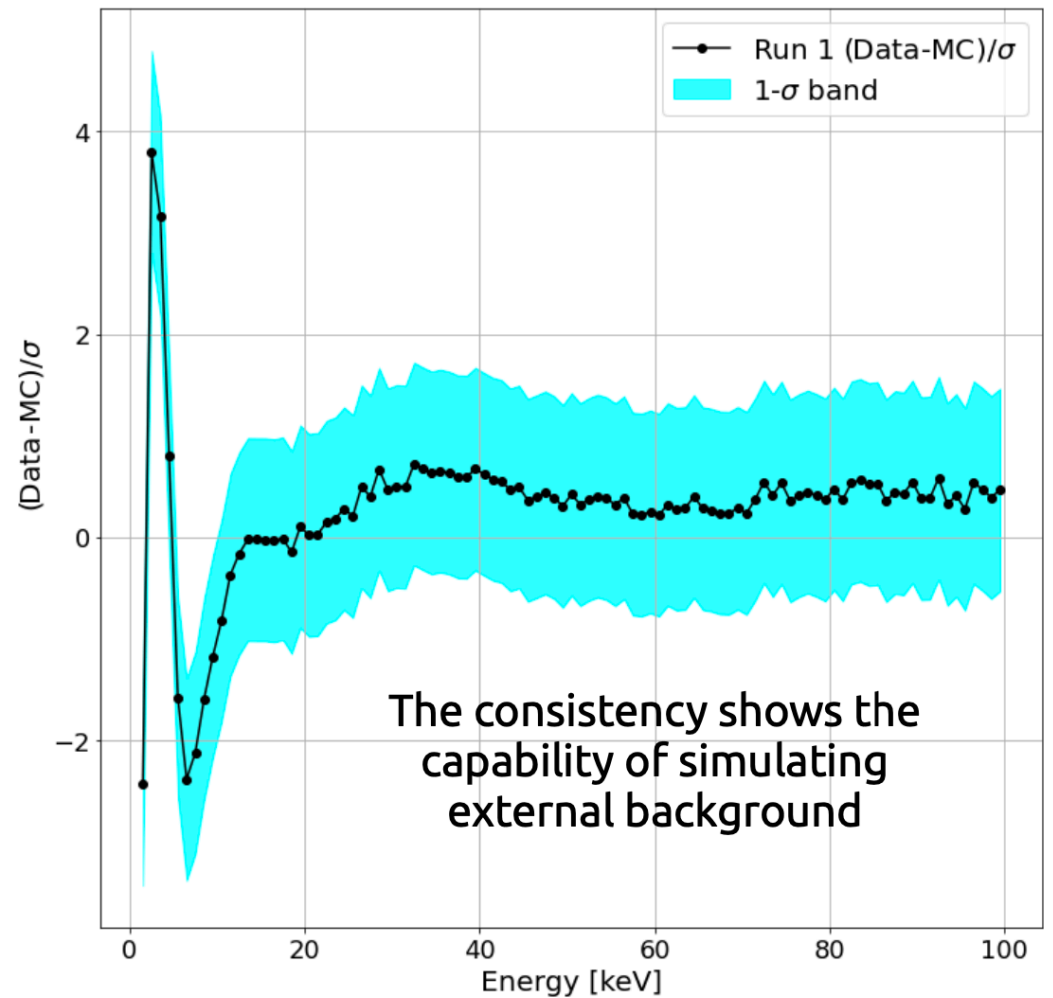
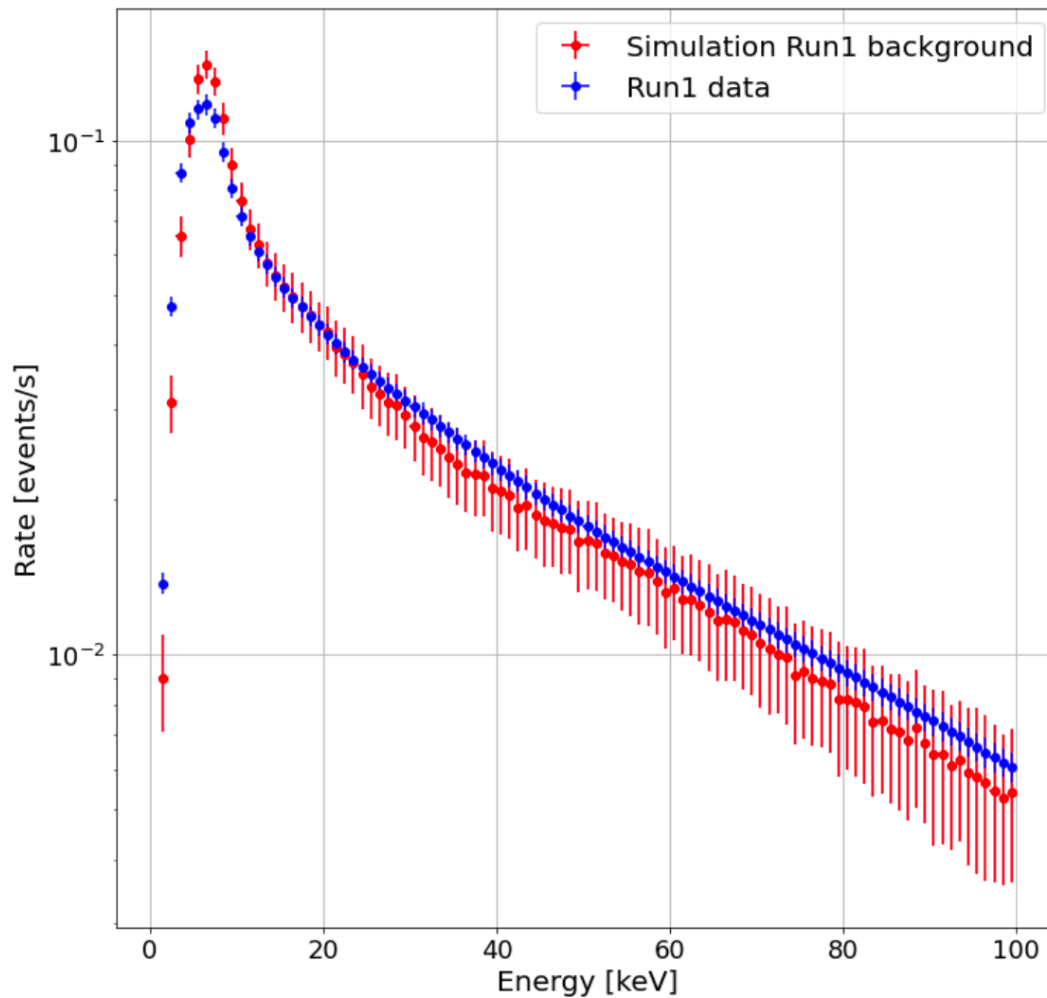
87%

Goal: maximise consistency between data and MC samples for proper data/MC comparison (rather than efficiency)

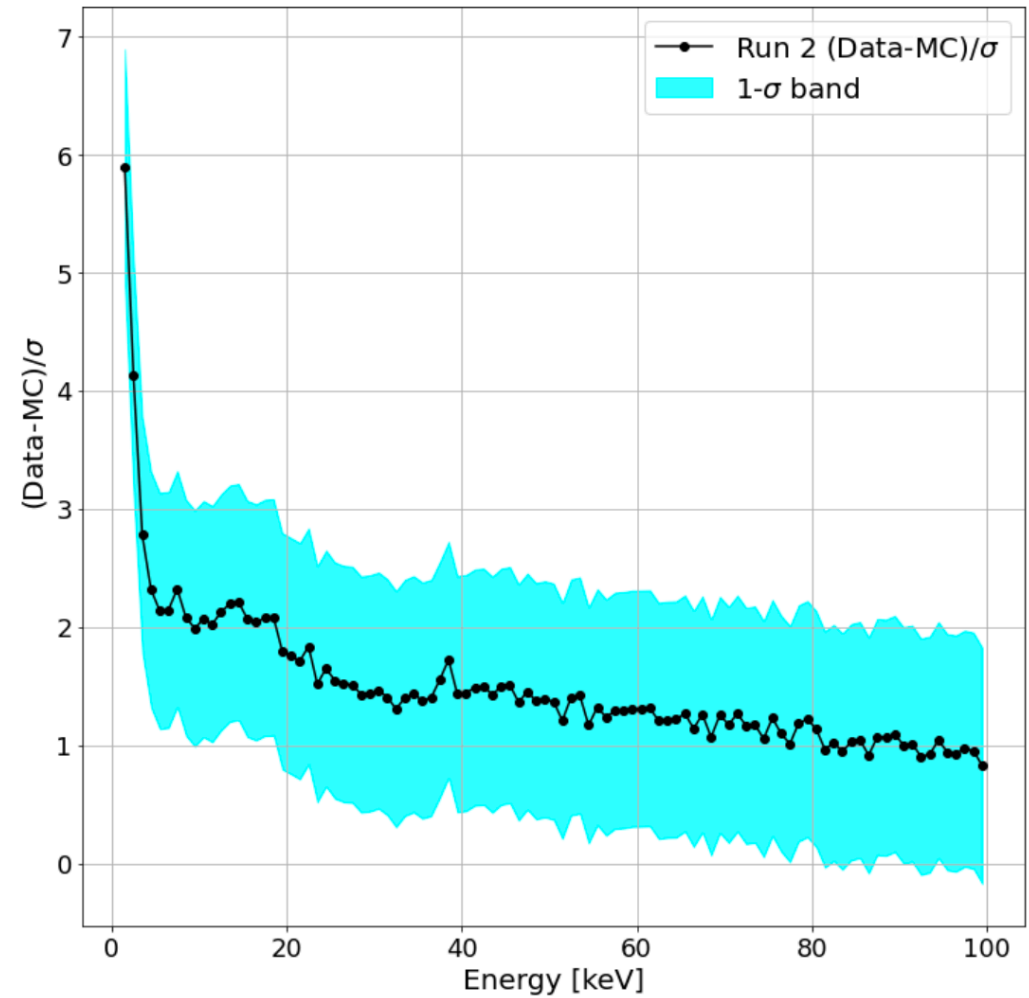
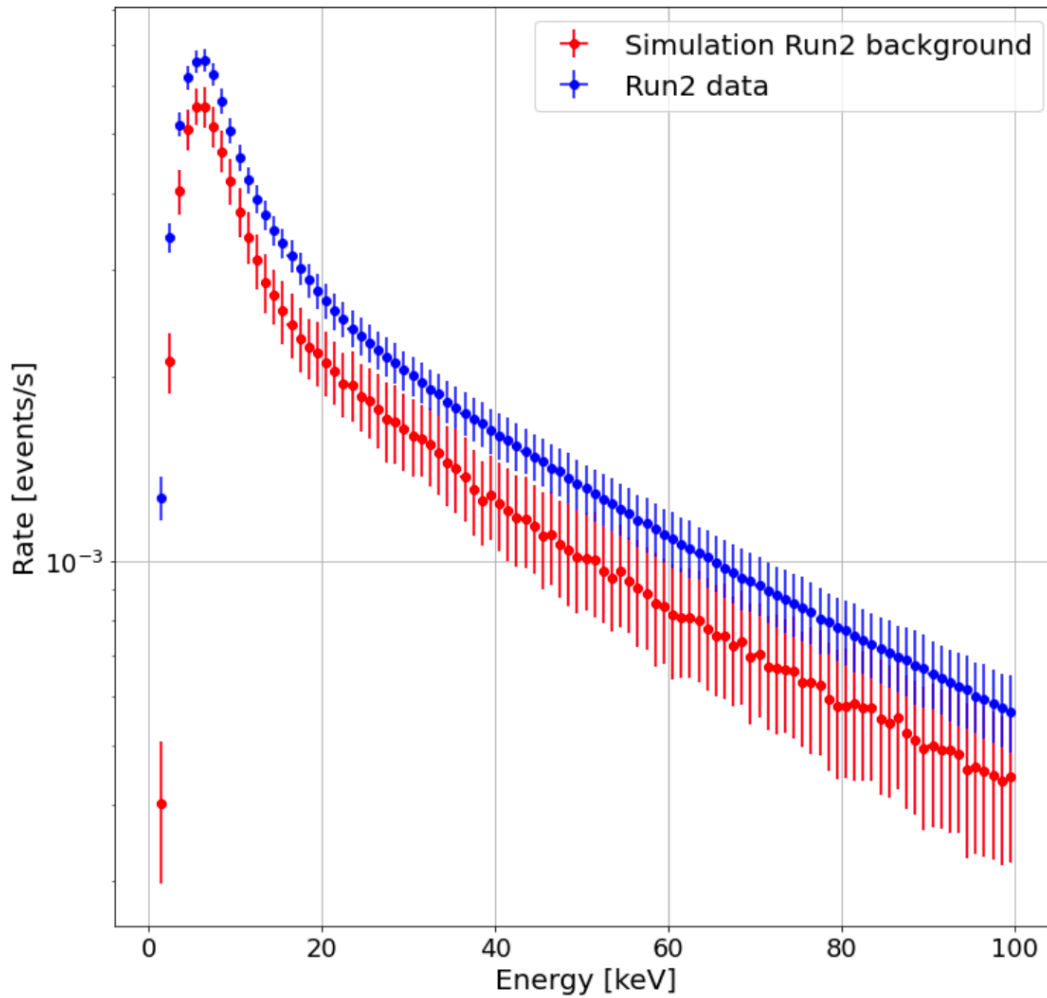
- 📍 **Fiducial cuts to exclude sCMOS sensor noisy borders**
 - 📍 A central square of 23 x 23 cm²
- 📍 **Quality cuts to remove fake clusters from residual sensor noise**
 - 📍 Remove uniform, compact and sharp-edged clusters from pedestal study
- 📍 **Track selection cuts based on $\delta = \# \text{ of counts} / \# \text{ of pixels}$ (i.e. uncalibrated dE/dx)**
 - 📍 To remove alphas (energy > 800 keV, not digitised, not of interest for DM searches)
 - 📍 To exclude MIP-like particles (not of interest for DM searches) whose reconstruction highly depends on LY and that might end up splitted in the reconstruction



Run1: dominated by external gamma background (internal 0.6%)

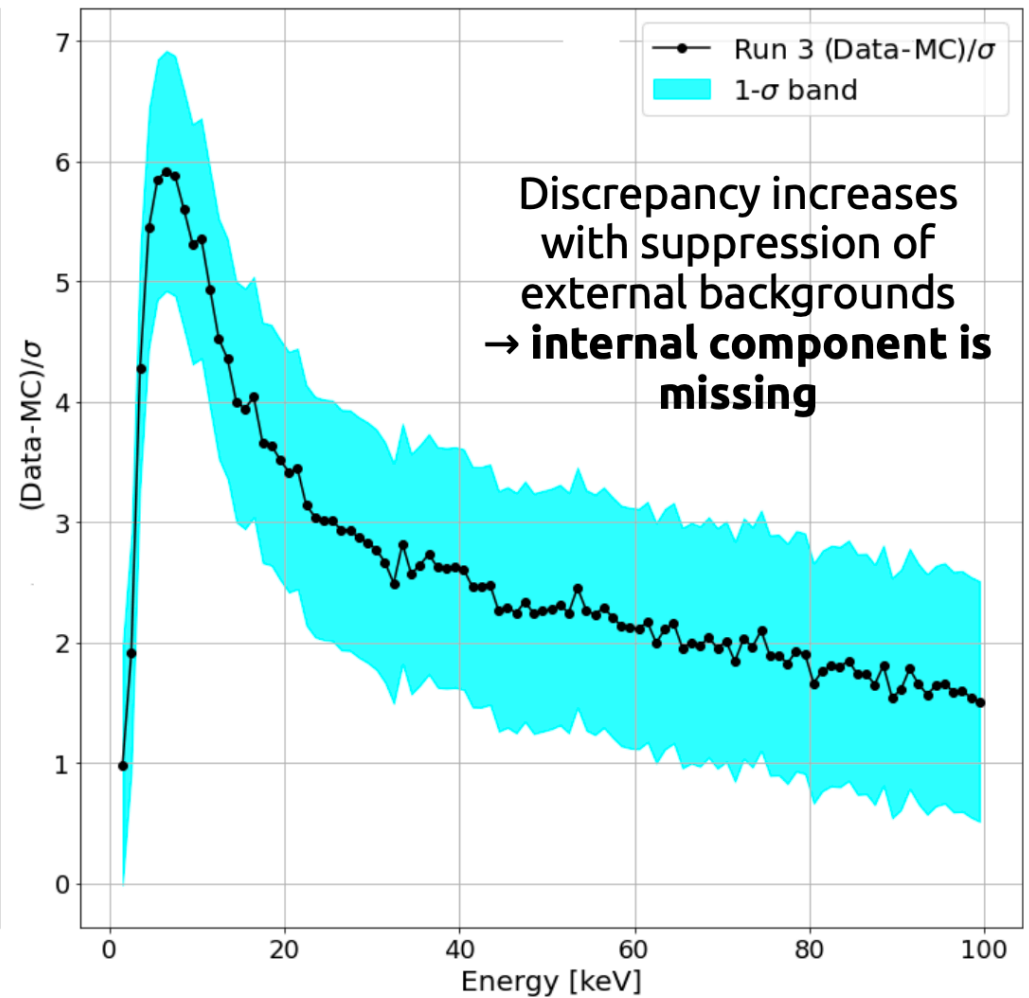
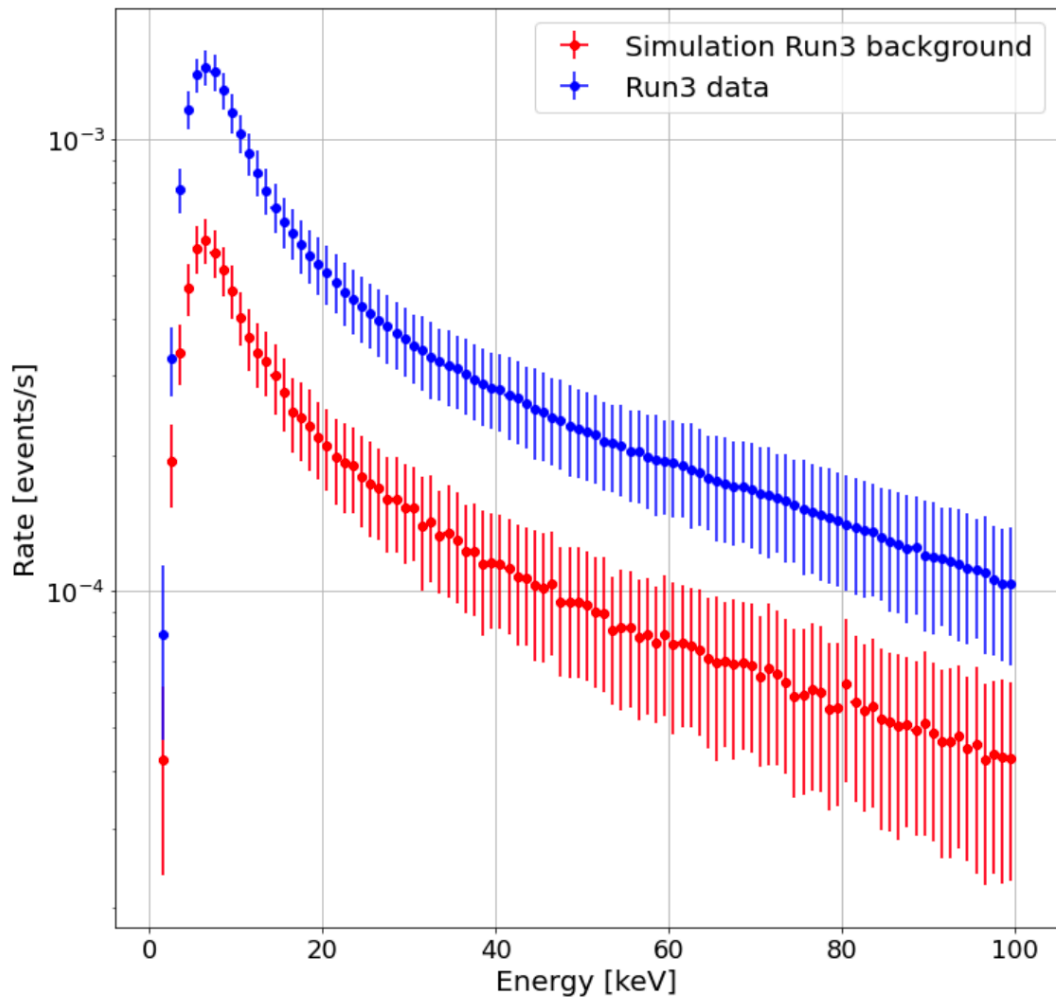


Run2: start probing internal gamma background (internal 20%)



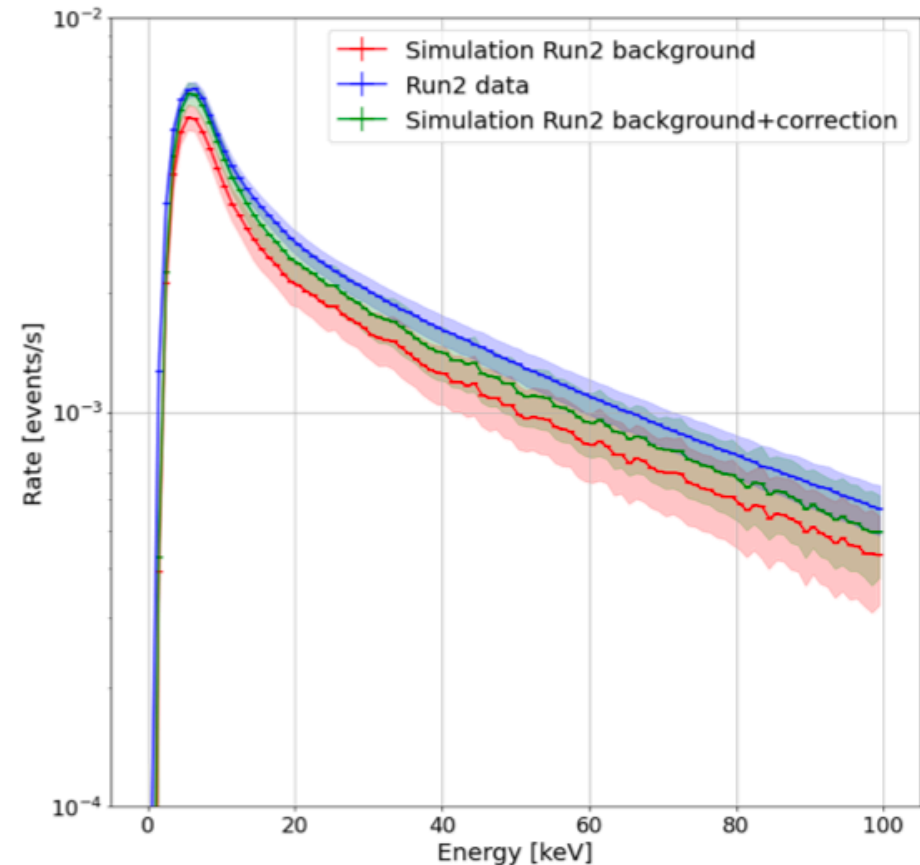
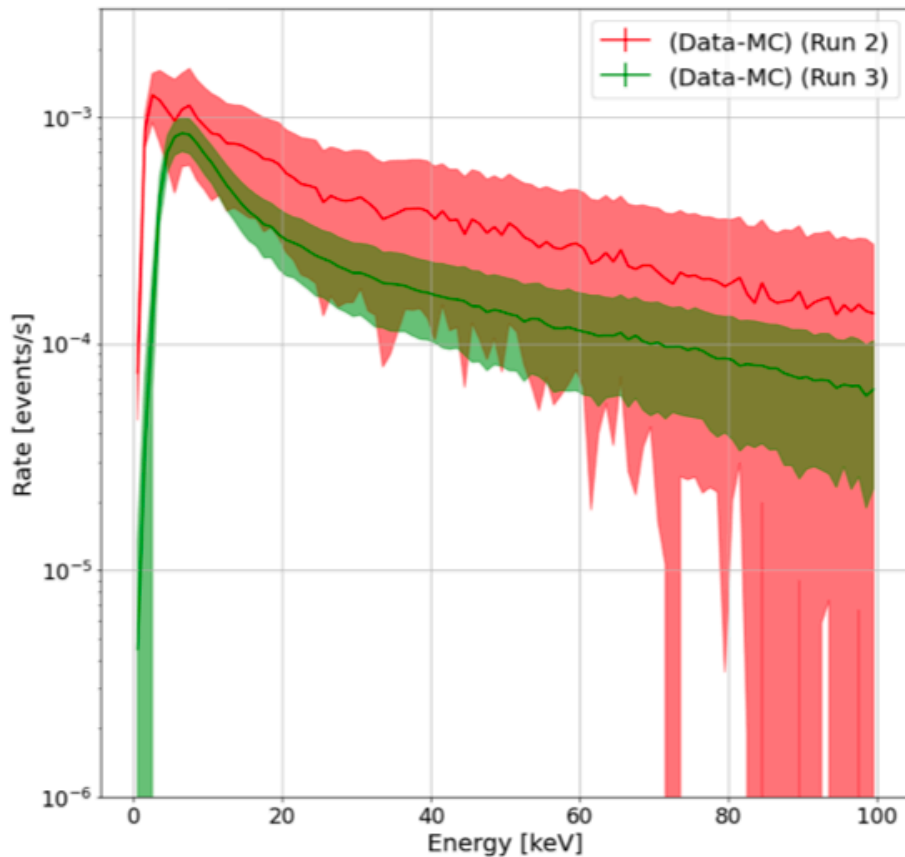
Data/MC difference $\pm 22\%$

Run3: dominated by internal gamma background (internal 86%)



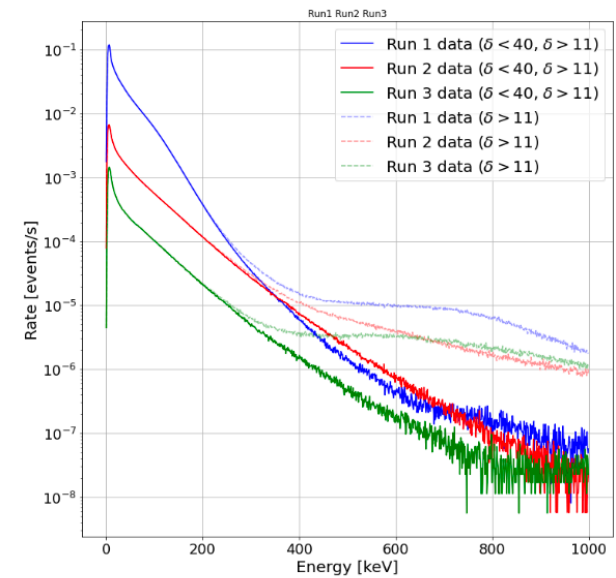
Data/MC difference $\pm 60\%$

The missing component

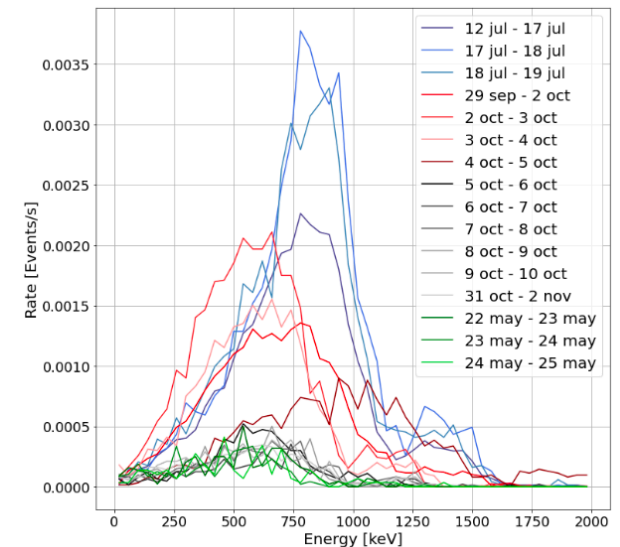
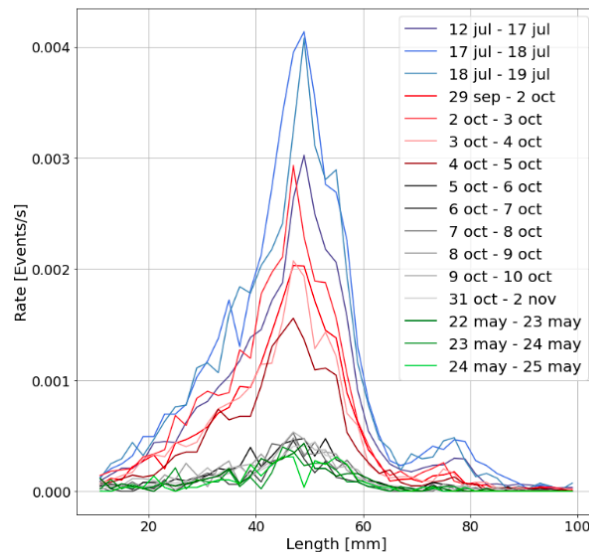
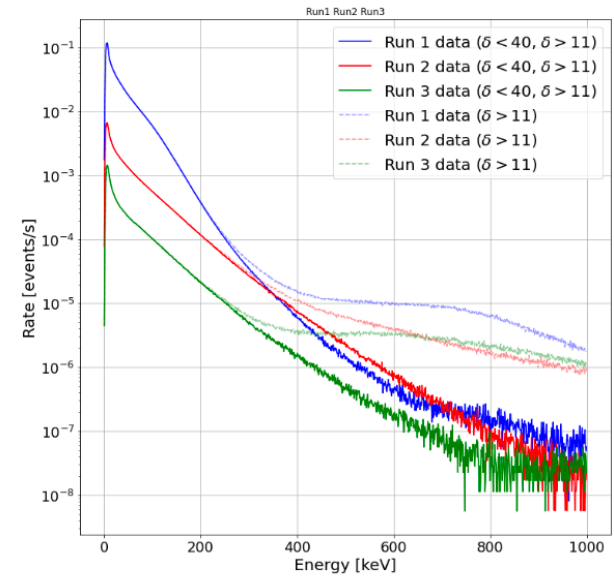


- Data/MC difference increasing with increasing shielding thickness/increasing suppression of external backgrounds
- Data/MC difference consistent between Run2 and Run3 even with an order of magnitude difference of overall rate
- **Indicating an internal component not considered in the MC simulation**
- Missing component of $O(10^{-2})$ events/s

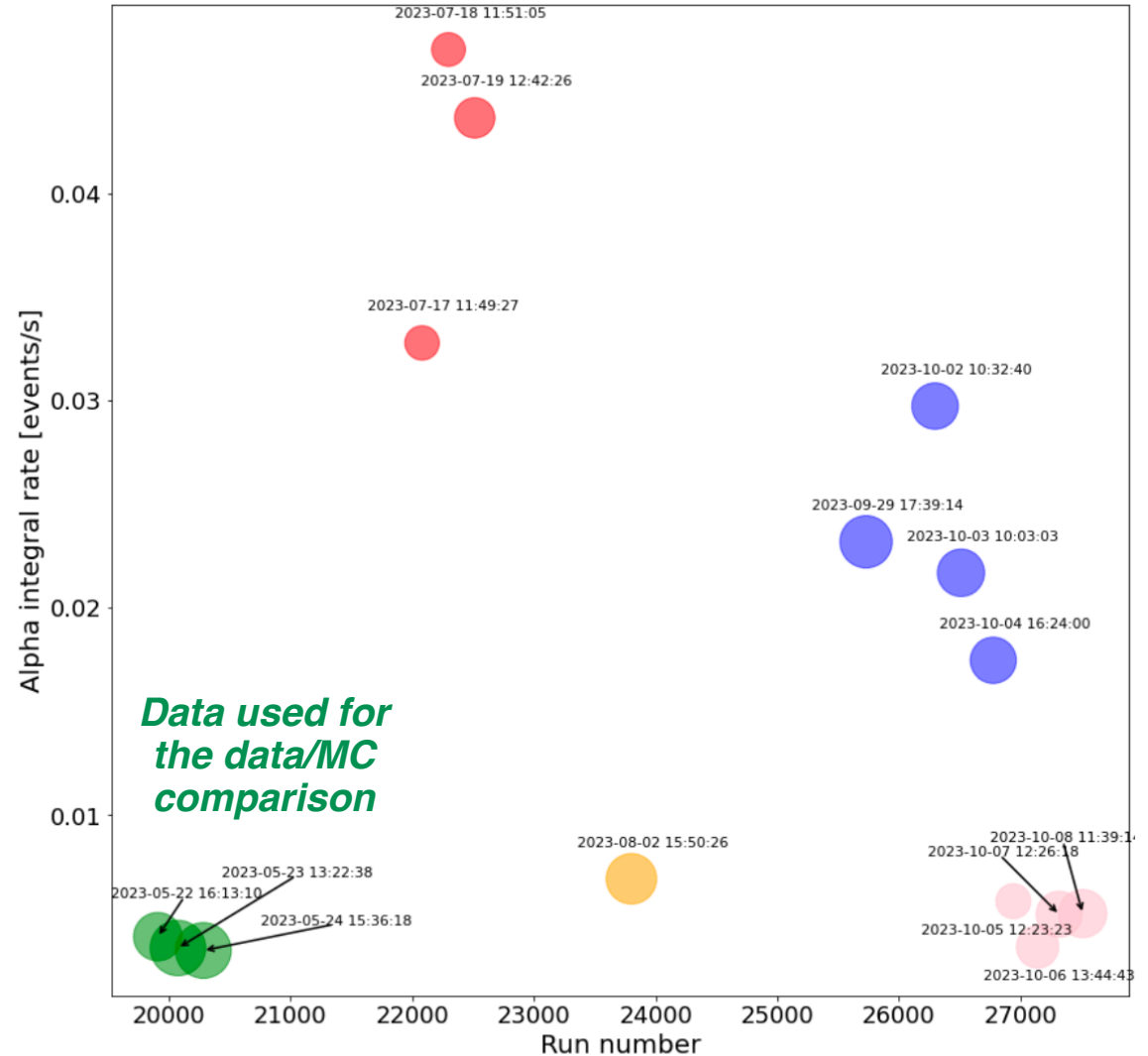
- **Excess of α events in all runs (long, dense tracks)**
 - Alphas from GEANT4 (not digitized) are not enough to explain the excess
- Due to varying gas conditions (charge gain saturation) energy measurement not feasible



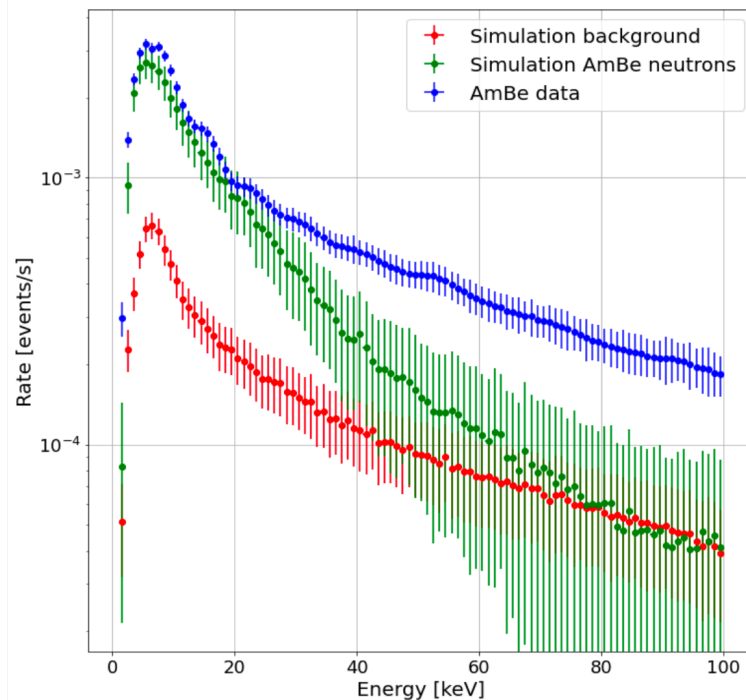
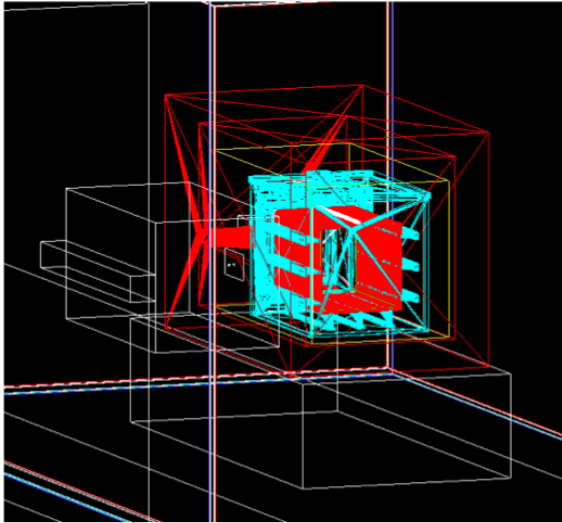
- **Excess of α events** in all runs (long, dense tracks)
 - Alphas from GEANT4 (not digitized) are not enough to explain the excess
- Due to varying gas conditions (charge gain saturation) energy measurement not feasible
- **Length** distribution indicates peaks around 5.9 MeV, 6.6 MeV, 8.1 MeV peaks (might be ^{222}Rn)
- Radioactive contamination might also induce beta and gamma events, populating the **low energy region**
- Further studies to identify the source (ongoing)



- Alpha event rate change over time, related to **gas system conditions**
 - May: **no recirculation**
 - July: **start of recirculation**
 - August (AmBe): **filter change**
 - **Gas system failure**
 - October: **filter change**
- Gas system interventions introduce **radioactive contaminants** (not expected, not simulated)
- Alpha contamination level $O(0.1-1)$ Bq/m³

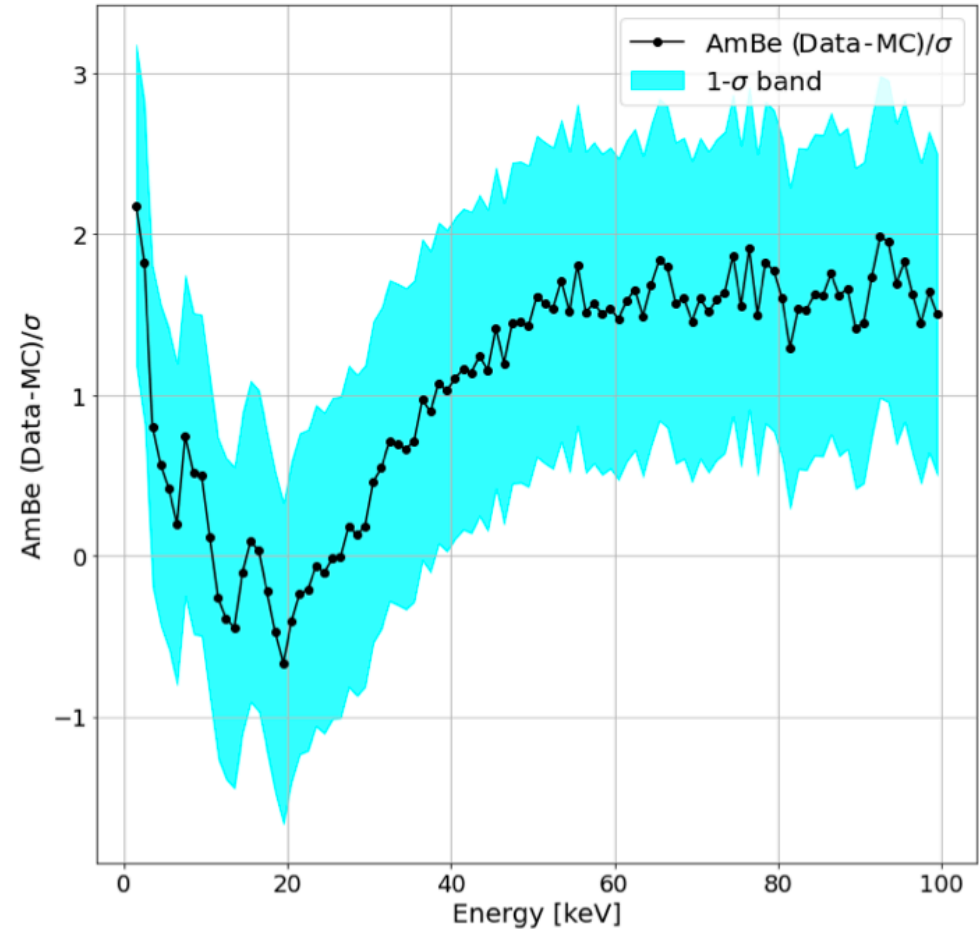
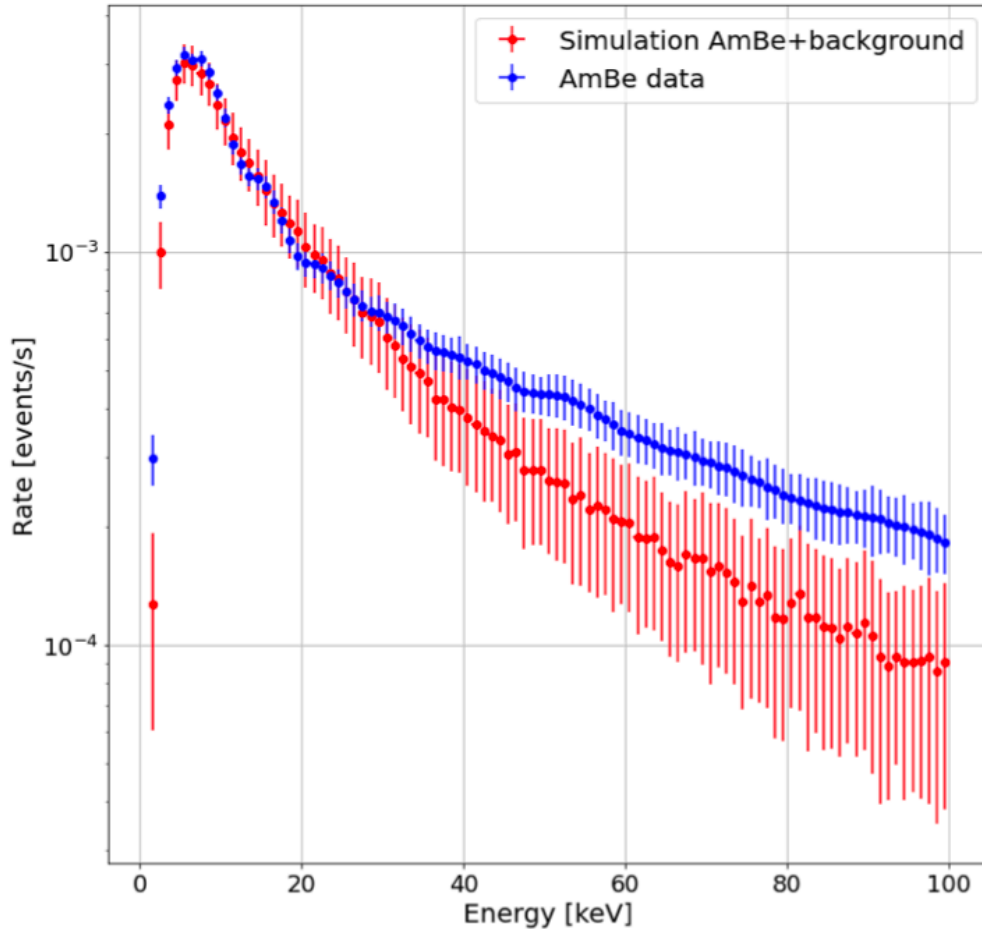


The size of the circle is proportional to the measured LY



- During AmBe ^{55}Fe calibrations only at the center of detector to minimise human intervention
- Same energy calibration, time normalisation and quality and selection cuts as background analysis
 - Except for $\Delta < 40$ to not remove NR
- AmBe induced events dominate backgrounds of 1 order of magnitude < 20 keV
- AmBe induced events comparable to backgrounds between 20 and 50 keV
- Backgrounds dominating > 50 keV

AmBe data/MC comparison



Missing component consistent between AmBe data and Run2/Run3, further validating our hypothesis of an internal common (likely gas-related) unforeseen radioactive contamination



Conclusions from LIME Run1, Run2, Run3 and AmBe data analysis and MC comparison



- Consistency of AmBe data and simulation below 20 keV where the AmBe contribution dominates over the backgrounds
 - **correct relative normalisation of data and MC and of the energy calibration**



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 - **we can properly simulate the response of our detector to the external gamma background**



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- Increase of the data/MC discrepancies with increasing shielding
 - **internal origin of the missing background components**



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- Increase of the data/MC discrepancies with increasing shielding
 - **internal origin of the missing background components**
- Consistency of the missing background component between Run2, Run3 and AmBe
 - **presence of an internal background not taken into account in the MC simulation**



Conclusions from LIME Run1, Run2, Run3 and AmBe data analysis and MC comparison



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 - correct relative normalisation of data and MC and of the energy calibration
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- Increase of the data/MC discrepancies with increasing shielding
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- Consistency of the missing background component between Run2, Run3 and AmBe
 - presence of an internal background not taken into account in the MC simulation
- Connection between the missing component with gas system operation, gas humidity and alpha particles spectra and rate
 - likely presence of Radon in the gas, at a level consistent with other underground gas detectors



Conclusions from LIME Run1, Run2, Run3 and AmBe data analysis and MC comparison



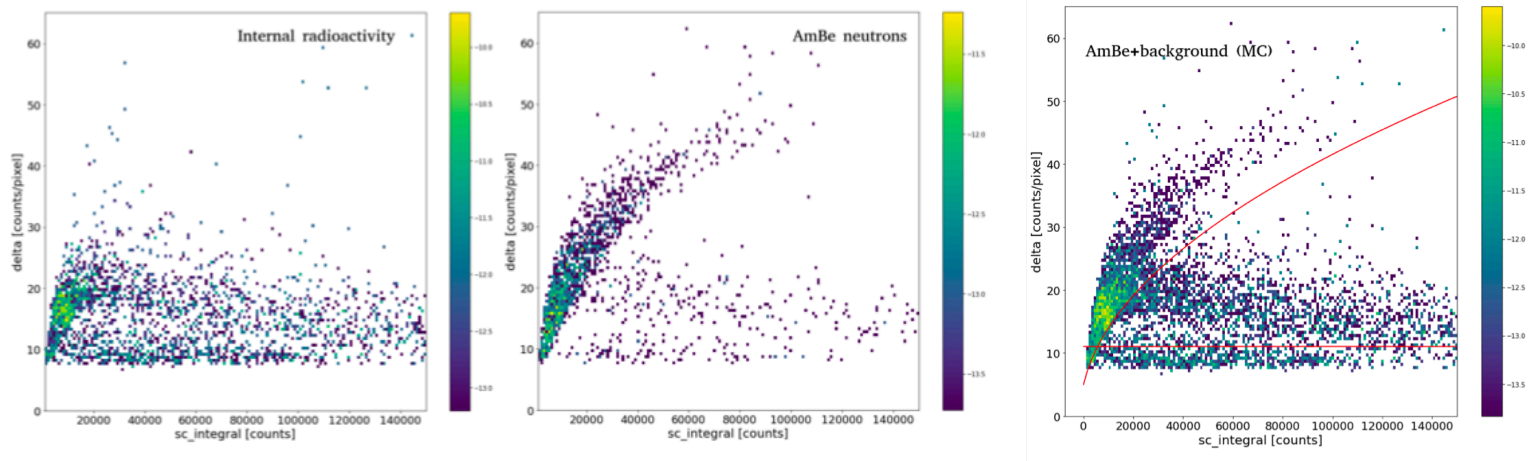
- Consistency of AmBe data and simulation below 20 keV where the AmBe contribution dominates over the backgrounds
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 - presence of an internal background not taken into account in the MC simulation
- Connection between the missing component with gas system operation, gas humidity and alpha particles spectra and rate
 - likely presence of Radon in the gas, at a level consistent with other underground gas detectors
- Low radioactivity Radon filter installed at the end of March 2024, Radon monitor to be installed soon



Preliminary estimate of ER/NR discrimination on LIME data

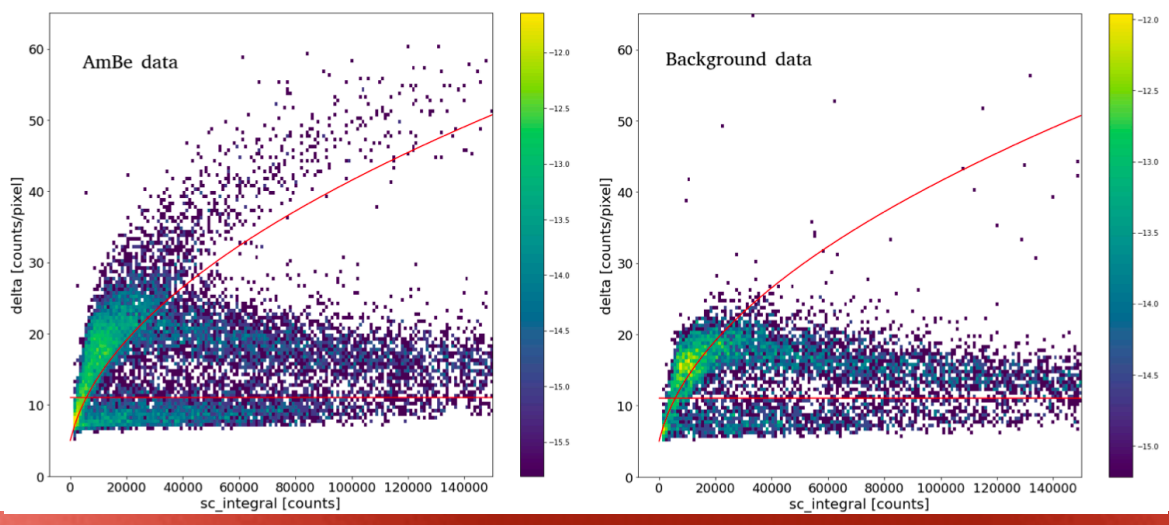
AmBe data: NR identification/ER rejection with classical approach

Same energy calibration, time normalisation and quality and selection cuts as background analysis, **except for $\delta < 40$ to not remove NR**



MC simulation

uncalibrated dE/dx versus uncalibrated E



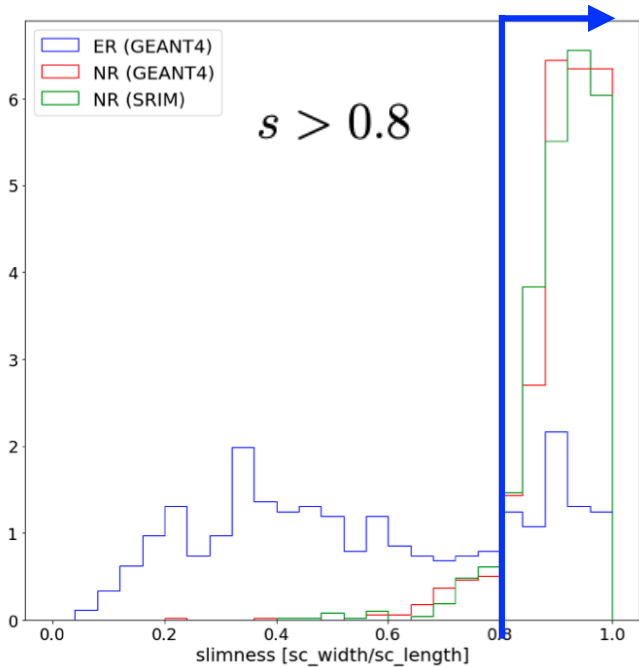
LIME AmBe/
background data

$$\delta > \sqrt{a + bI}$$

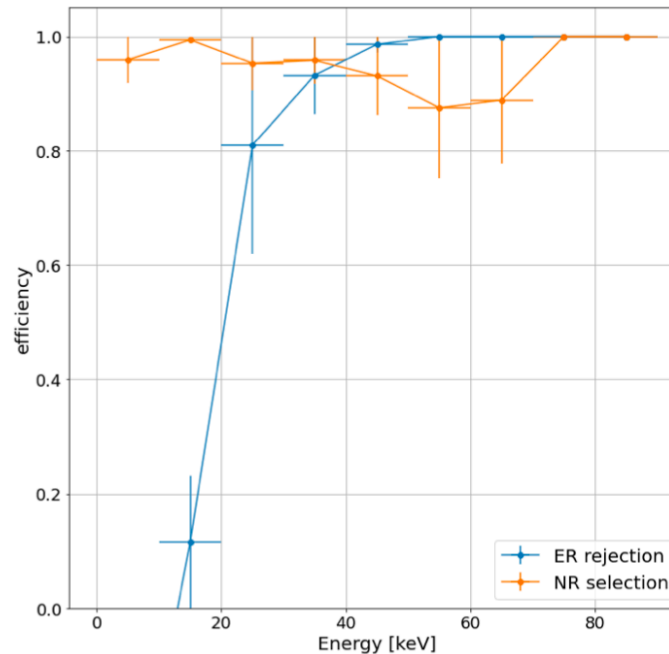
$$a = 25 \text{ and } b = 0.017$$

NR selection cut optimised on MC simulation

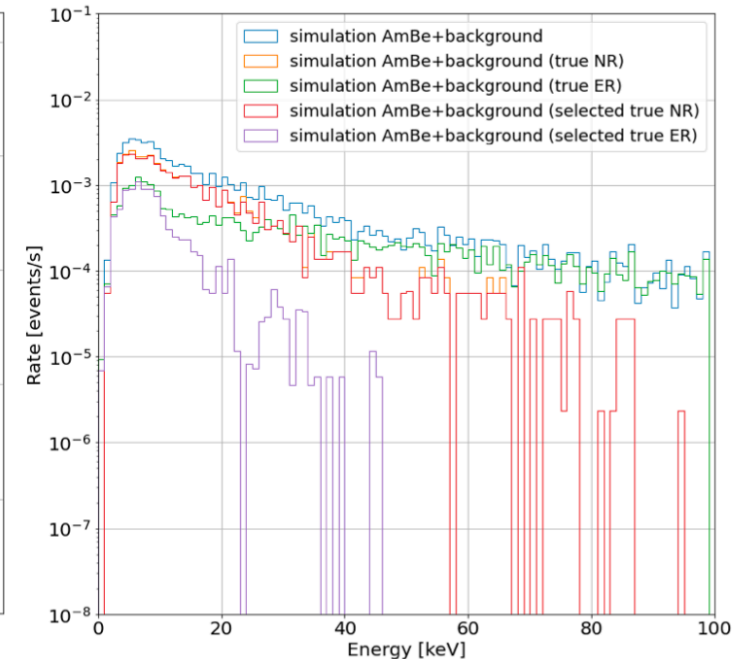
Ratio minor/major track axis



NR efficiency/ER rejection

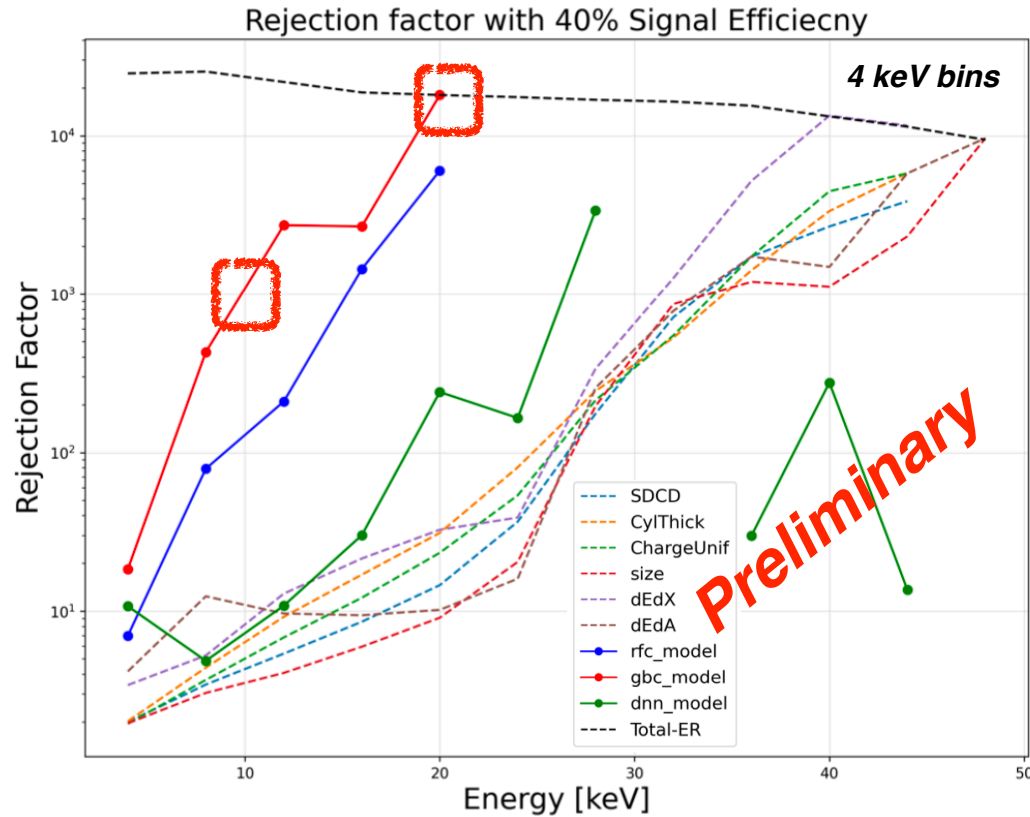


AmBe spectrum after selection



- Simple selection optimized on MC, cut on **track energy density** and **slimness** yields good **ER rejection** (>80% at 20 keV)
 - Preliminary demonstration of feasibility of neutron flux measurement (Run 5)
 - ML algorithm developments ongoing for ER/NR discrimination

Rejection factor on MC full simulation



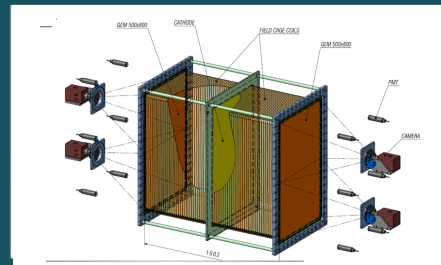
3 deep learning models developed and compare with classical analysis on track shape variable

indication of background rejection $> 10^4$ @ 20 keV

PHASE 1: O(1) m³ Demonstrator

2024/26
LNF/LNGS

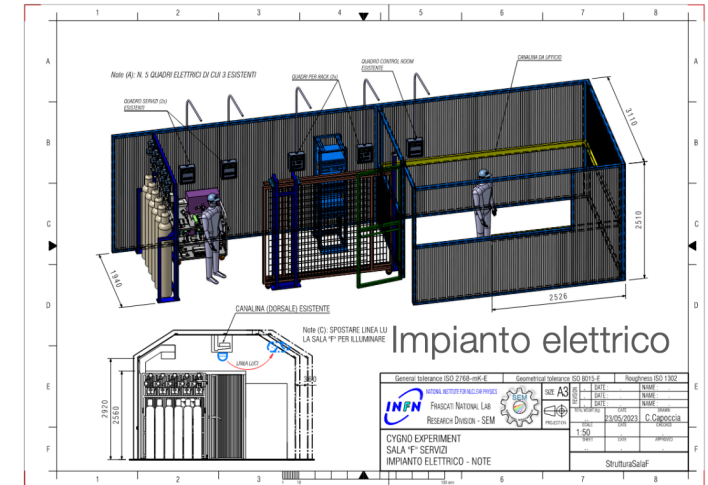
CYGNO_04



- background
- materials test, gas purification
- scalability

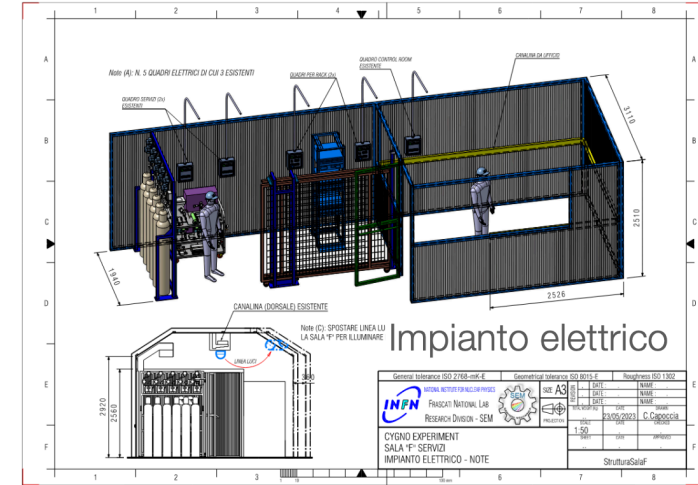
Towards CYGNO-04 realisation

- An (unforeseen necessary) formal agreement between GSSI and LNGS for the realisation of infrastructure works was signed by end of March 2024
- Technical design validation by the company to be completed by mid May 2024 will be immediately followed by formal awarding of the contract

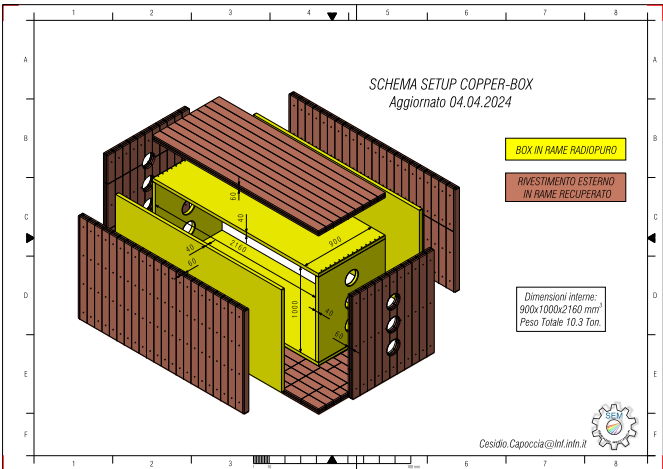


Towards CYGNO-04 realisation

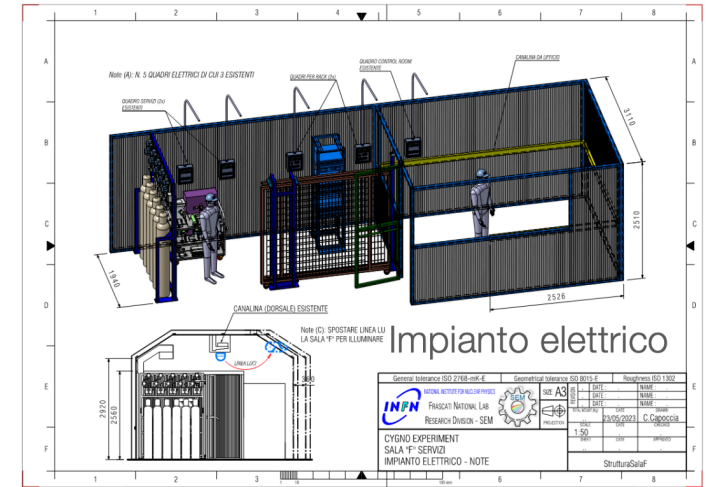
- An (unforeseen necessary) formal agreement between GSSI and LNGS for the realisation of infrastructure works was signed by end of March 2024
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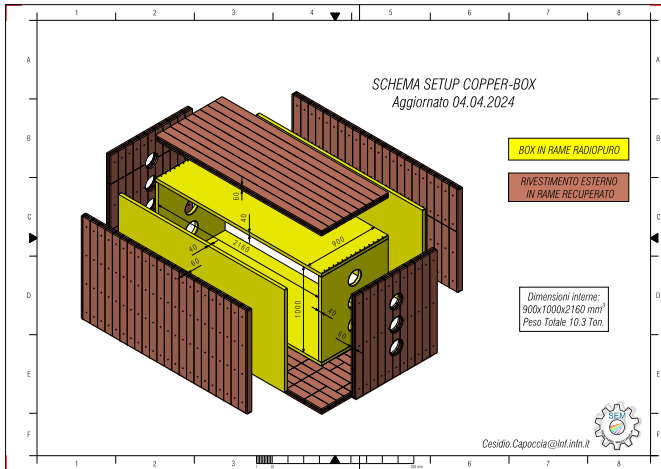
- Detector and shielding design under finalisation
- Main detectors materials and components identified and construction procedures under test and validation



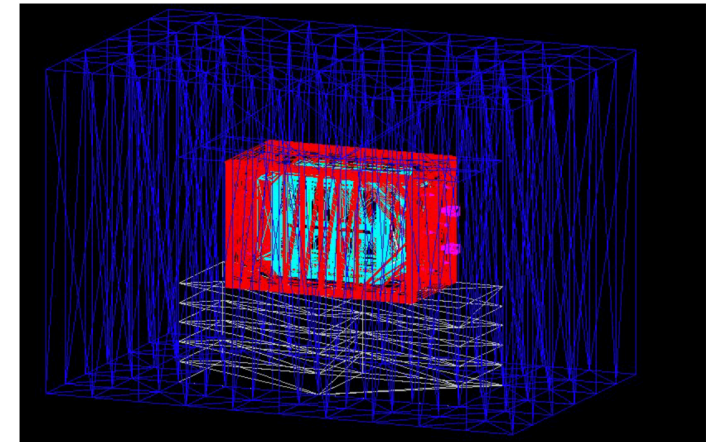
- An (unforeseen necessary) formal agreement between GSSI and LNGS for the realisation of infrastructure works was signed by end of March 2024
- Technical design validation by the company to be completed by mid May 2024 will be immediately followed by formal awarding of the contract



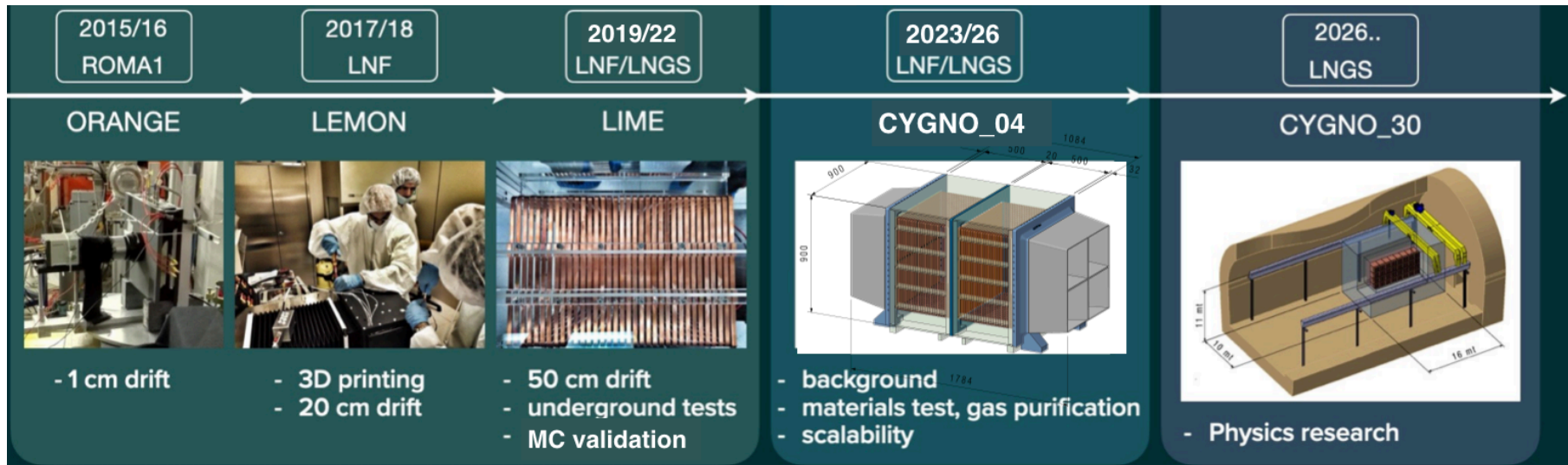
- Detector and shielding design under finalisation
- Main detectors materials and components identified and construction procedures under test and validation



- Nearly final design implemented in GEANT-4
- Preliminary evaluation of external gammas in full shielding configuration performed
- Full simulation of internal and external backgrounds ongoing

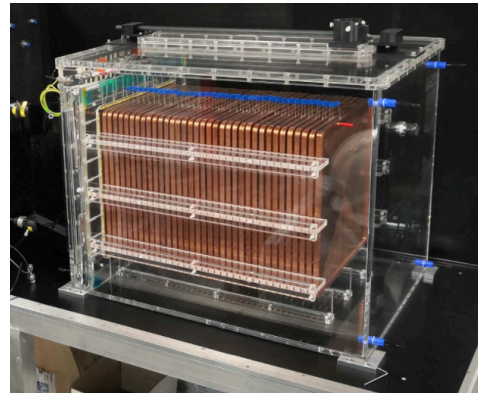
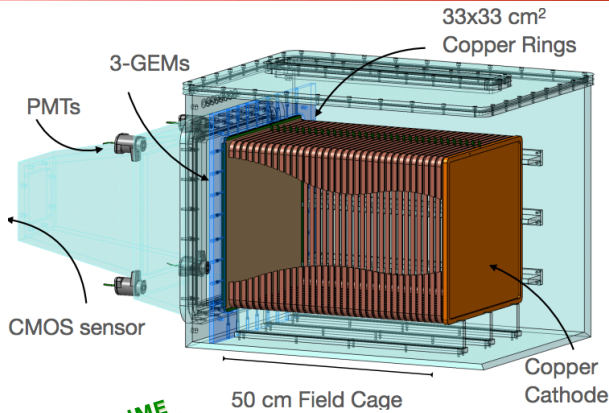


- **PHASE 0 has been successful in realising its goals**
 - LIME underground operation proceeding since > 1 year
 - Auxiliary systems improved and validated
 - Computing infrastructure realised and validated
 - External shielding effect on backgrounds validated
 - MC simulation validated and unforeen background contamination likely identified
 - **Stable and high quality detector operation achieved with full auxiliary systems configuration**
- **Development towards CYGNO-04 realisation advancing**
 - Infrastructures to be completed in a couple of months
 - Detector and shielding design under finalisation
 - Main detector materials and components identified and construction procedures under test
 - Full background simulation ongoing





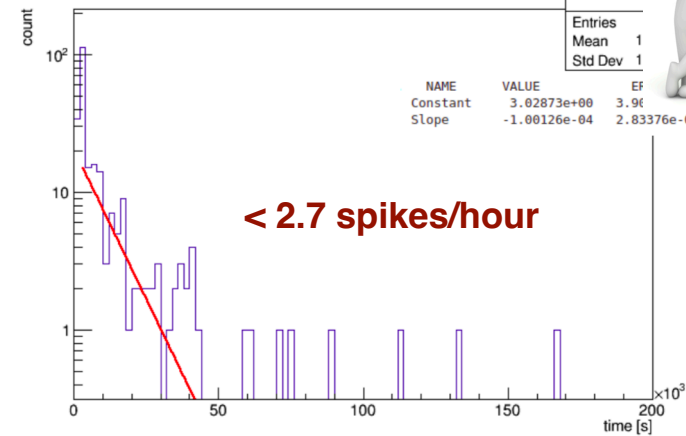
BACKUP



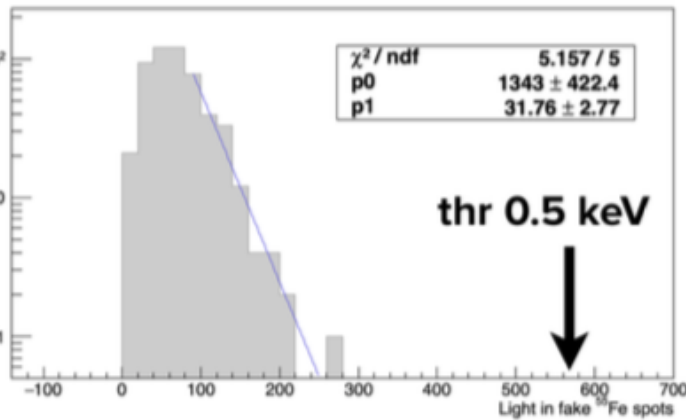
**He:CF₄
@ 1 atm**

**1 sCMOS + 4 PMT + 3 GEMs
33 x 33 cm² area
50 cm drift, 50 L active volume
*Base module of PHASE 1 design***

Stability



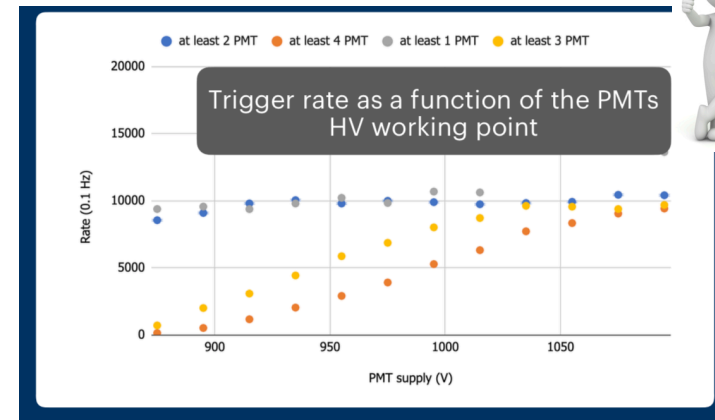
sCMOS fake clusters threshold



ORCA-Fusion

1 keV = 1200 photons

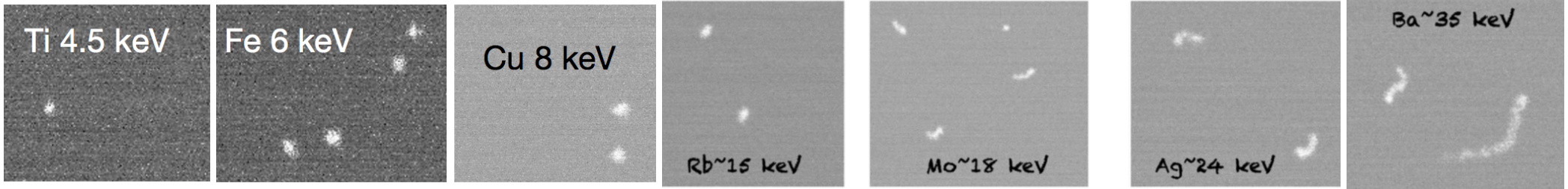
PMT trigger rate



Different Trigger logics were tested with ⁵⁵Fe;
All of them converge to the same rate of 1 kHz

LIME overground commisioning @ LNF

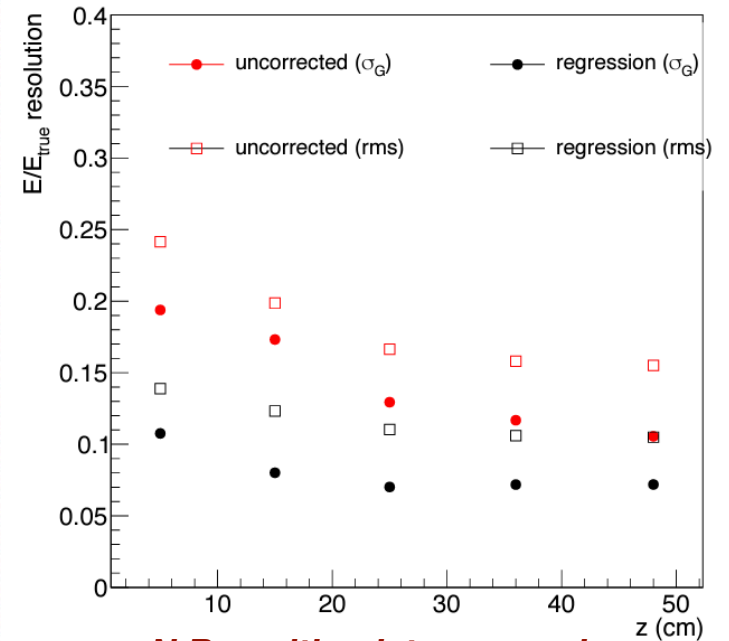
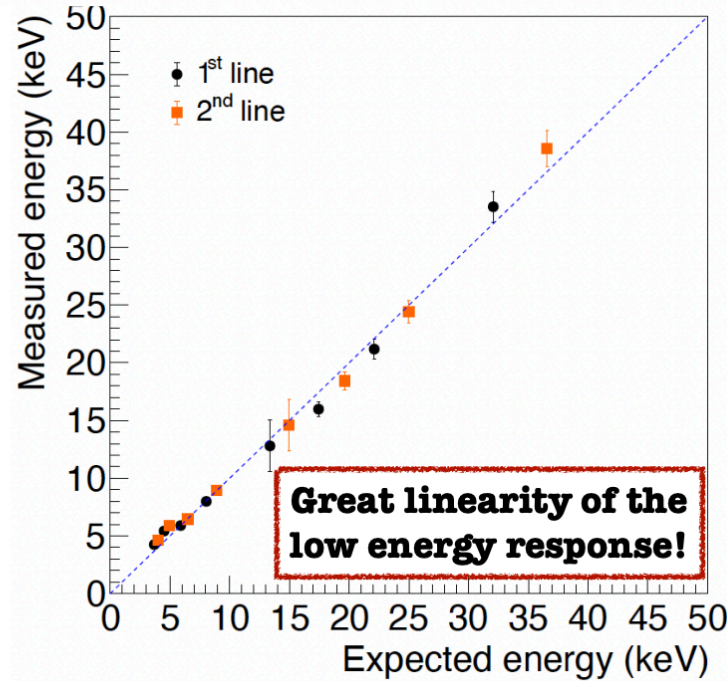
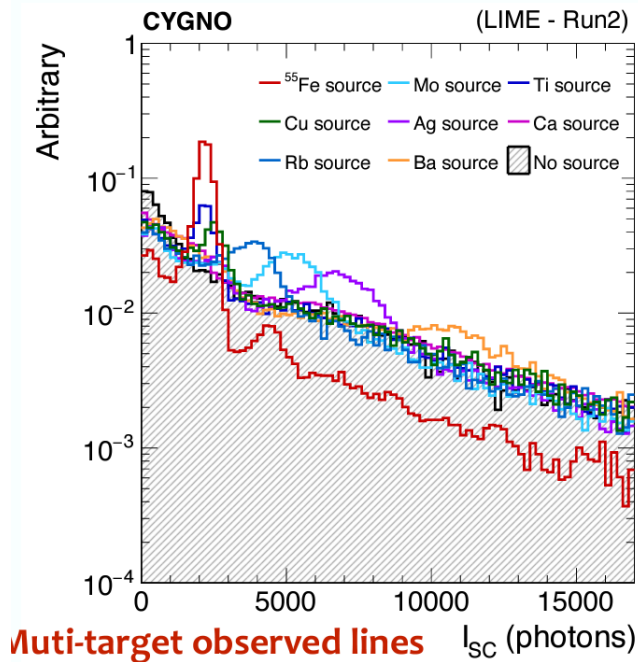
Electron recoils calibration



Multi-source + bkg spectrum

Energy response linearity

Energy resolution



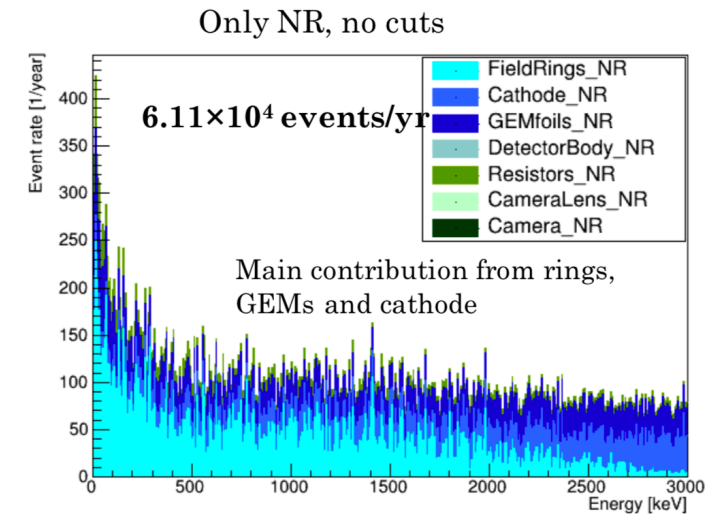
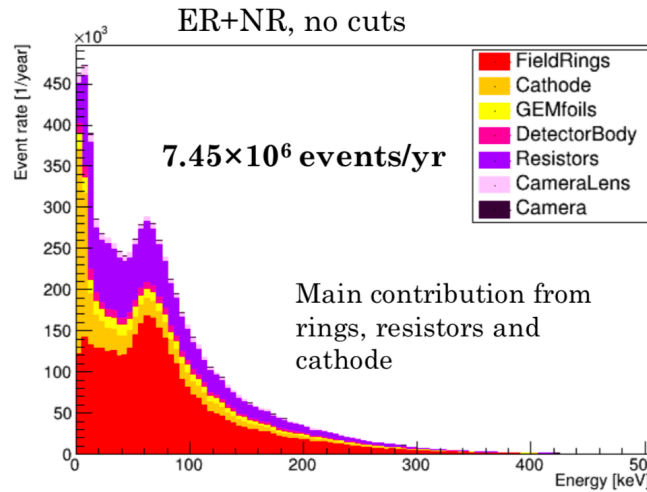
N.B. multivariate regression to correct for detector response disuniformities



PHASE 0 internal backgrounds simulation



- **Radioactivity** from natural radioactive decay chains ^{232}Th , ^{238}U , ^{235}U and other contaminants
- Activity of all main components of LIME were measured underground by M.Laubenstein:
 - Acrylic box
 - Field rings
 - Cathode
 - Resistors
 - GEMs
 - Camera (body + lens)



	Radionuclide	FieldRings	Cathode	Resistors	GEM	Acrylic	Camera body	Camera lens
^{238}U chain	^{234}Th	<2,10E-01	<2,10E-01	1,99E+01	1,63E-01	-	3,16E+00	4,22E+00
	^{234}mPa	<7,70E-02	<7,70E-02	2,19E+01	-	-	-	-
	^{226}Ra	<1,30E-03	<1,30E-03	2,16E+00	3,25E-02	<3,50E-03	8,13E-01	1,92E+00
	^{210}Pb	-	-	5,94E+02	-	-	-	-
^{232}Th chain	^{228}Ra	<1,10E-03	<1,10E-03	3,50E+00	<3,09E-02	<5,00E-03	9,49E-01	3,61E-01
	^{228}Th	<1,30E-03	<1,30E-03	3,36E+00	<1,56E-02	<4,50E-03	9,49E-01	3,65E-01
^{235}U chain	^{235}U	<1,60E-03	<1,60E-03	3,37E-01	<1,58E-02	-	1,81E-01	1,45E-01
	40K	<6,00E-03	<6,00E-03	<1,78E+00	<3,58E-01	<3,50E-02	8,59E-01	5,15E+01
Other	^{137}Cs	<4,70E-04	<4,70E-04	<7,35E-02	<8,13E-03	-	4,07E-02	<2,67E-02
	^{60}Co	<5,70E-04	<5,70E-04	<7,73E-03	<7,48E-03	-	<5,42E-03	<4,64E-02
	^{58}Co	9,00E-04	9,00E-04	<3,10E-03	-	-	-	-
	^{54}Mn	<4,30E-04	<4,30E-04	<3,27E-03	-	-	-	-
	^{138}La	-	-	-	-	-	-	2,44E+00

Thanks to
LNGS
Services

Simulation
performed
including both
measurements
and limits

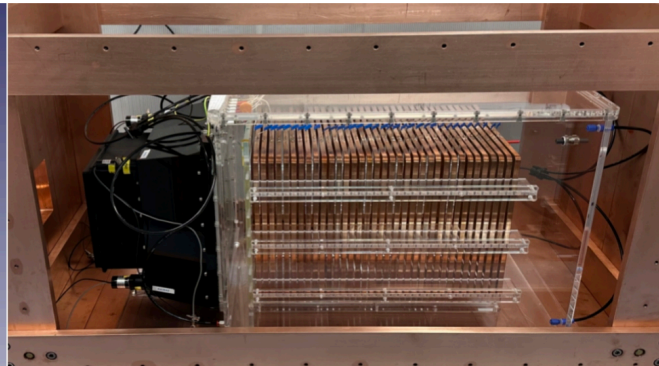
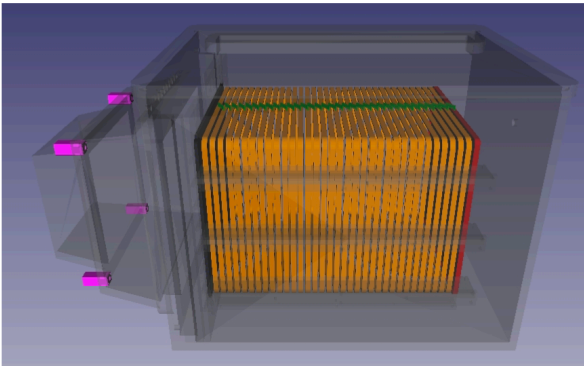
NOTE: internal background can be reduced of 96% (99% for NR) with fiducial cuts



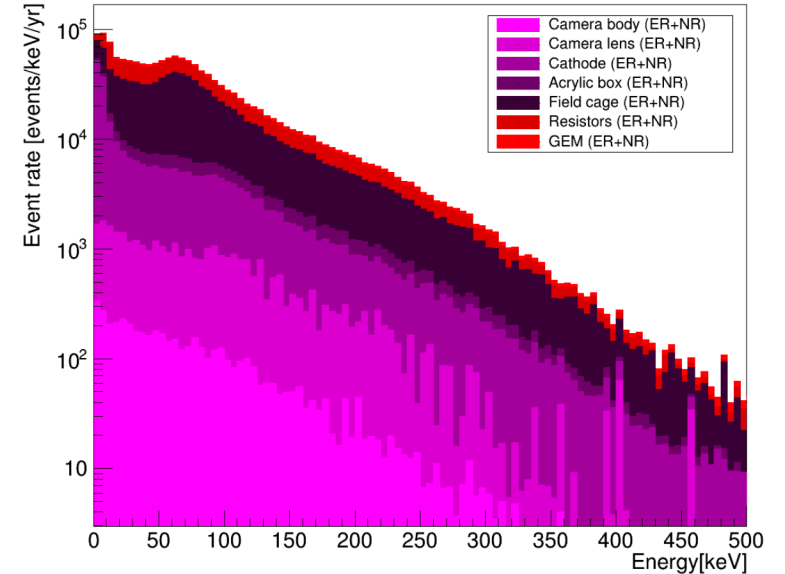
LIME backgrounds Geant4 simulation

- LIME components and shielding imported from technical CAD design
- Intrinsic radioactivity of main LIME materials as measured by LNGS Special Techniques Division
 - Cu field cage rings, field cage resistors, Cu cathode, GEMs, acrylic vessel, camera body, camera lens
 - camera body and lens results subdominant w.r.t. others
- Cu shielding radioactivity as measured by LNGS Special Techniques Division
 - From Opera dismissed Cu bars
 - Subdominant w.r.t. internal backgrounds
- Radiogenic and cosmogenic neutrons from shielding
 - Subdominant w.r.t. internal backgrounds

Many thanks to LNGS Special Techniques Division



Field cage Cu bars, resistors and Cu cathode dominate total background



Source	Event rate [10 ⁶ yr ⁻¹]
Field cage	(3.57±0.01)
Resistors	(1.873±0.006)
Cathode	(1.095±0.001)
GEMs	(0.3891±0.0002)
Vessel	(0.268±0.001)
Camera lens	(0.151±0.004)
Camera body	(0.0242±0.0005)
TOTAL	(7.34±0.01)

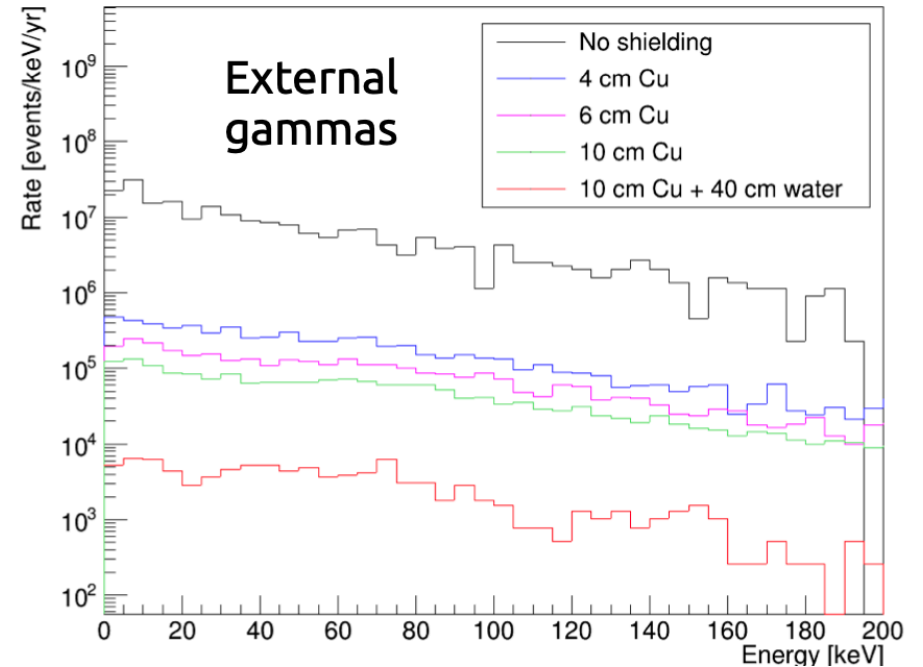
Mainly from ²³⁸U and ²³²Th chains

Please note LIME was NOT built with radioactive pure components

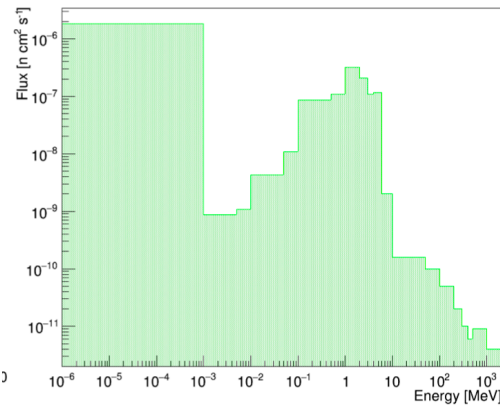
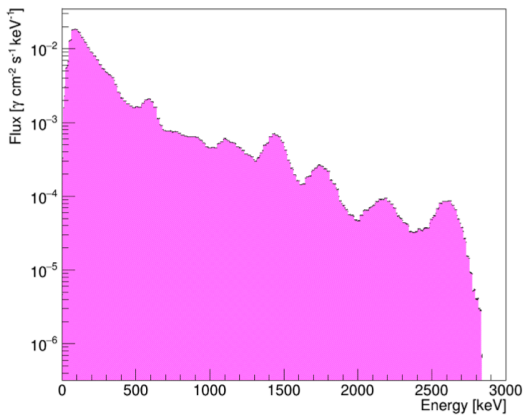


LIME backgrounds Geant4 simulation

- External gammas and neutron fluxes from previous measurements
- A staged shielding approach was devised to better:
 - measure and validate neutrons and gamma shielding capability to suppress external backgrounds
 - validate MC simulation of internal and external components
- After optimisation, four stages were chosen
 - No shield (Run1)
 - 4 cm Cu shield (Run2)
 - 10 cm Cu shield (Run3)
 - 10 cm Cu + 40 cm H₂O shield (Run4)



Input gamma (left) and neutron (right) fluxes for the simulation

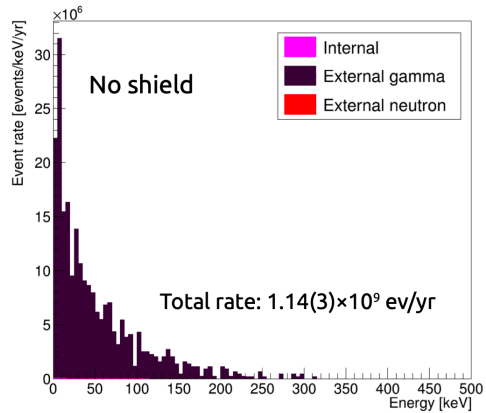


Phys.Rev. D73 053004 (2006)
 Il Nuovo Cimento A 101 (1989)
 Il Nuovo Cimento A 112.A (1999)

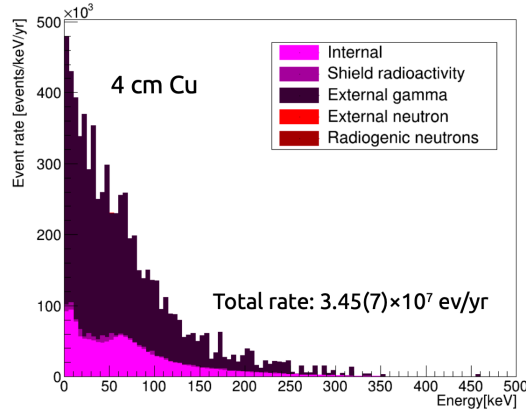
Shielding	Gamma background [10^6 ER yr^{-1}]	Neutron background [NR yr ⁻¹]
Unshielded	(1140±30)	(1480±90)
4 cm Cu	(26.2±0.6)	(870±10)
6 cm Cu	(9.4±0.3)	(1000±30)
10 cm Cu + 40 cm H ₂ O	(0.5±0.2)	(2.0±0.2)



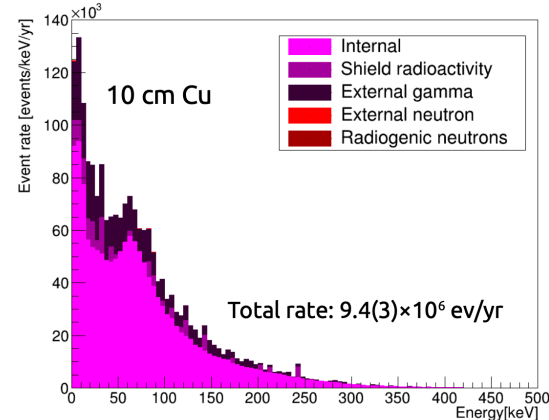
Total LIME expected backgrounds



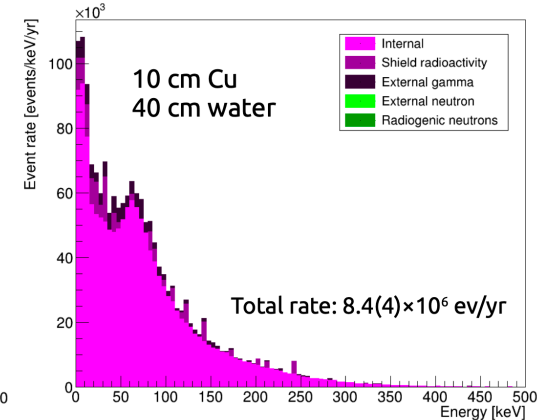
Completely dominated by external gammas
(internal is 0.6% of total)



Starting to probe internal background
(internal is 20% of total)



Internal background gives major contribution
(internal is 78% of total)



Largely dominated by internal background
(internal is 87% of total)

Phase	External		Internal		Total	
	ER [10^6 yr $^{-1}$]	NR [yr $^{-1}$]	ER [10^6 yr $^{-1}$]	NR [yr $^{-1}$]	ER [10^6 yr $^{-1}$]	NR [yr $^{-1}$]
Run 1	1140 ± 35	1480 ± 90	7.34 ± 0.01	79000 ± 470	1140 ± 35	80480 ± 480
Run 2	26.6 ± 0.6	870 ± 10	7.87 ± 0.3	79000 ± 470	34.5 ± 0.7	79870 ± 470
Run 3	1.49 ± 0.04	930 ± 25	7.88 ± 0.3	79000 ± 470	9.37 ± 0.3	79930 ± 470
Run 4	0.5 ± 0.2	2.0 ± 0.2	7.88 ± 0.3	79000 ± 470	8.38 ± 0.4	79000 ± 470

internal/external

0.6%

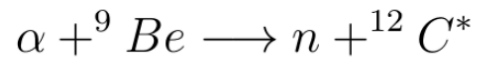
20%

78%

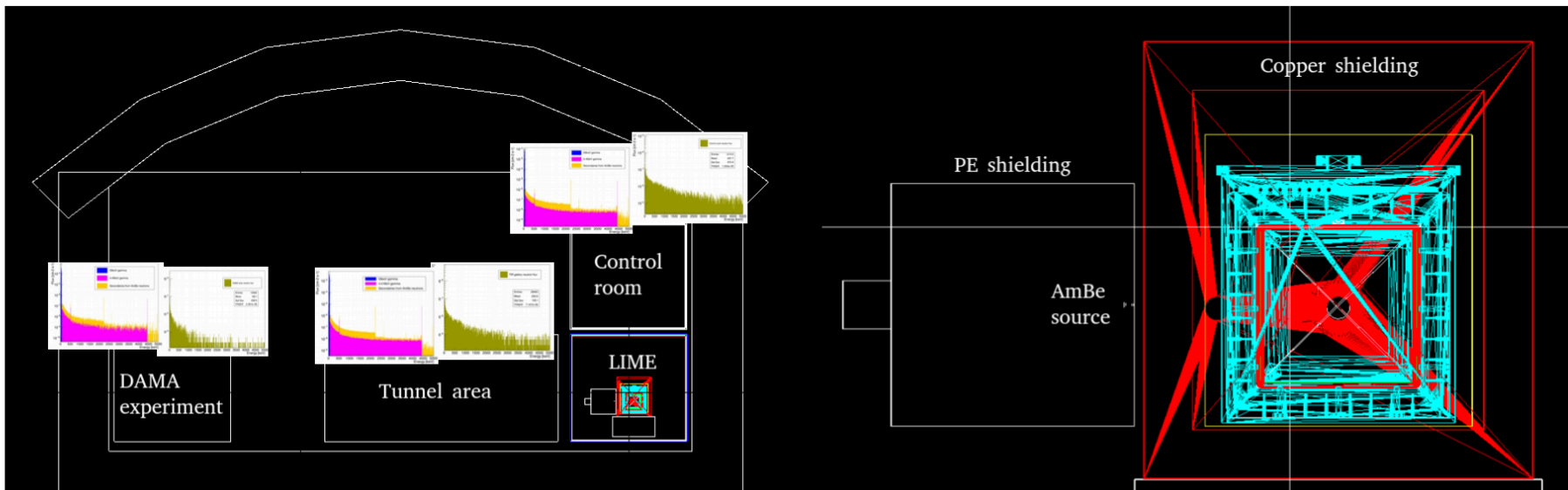
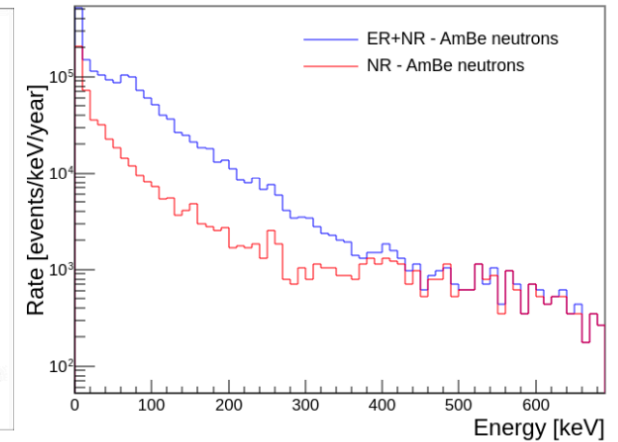
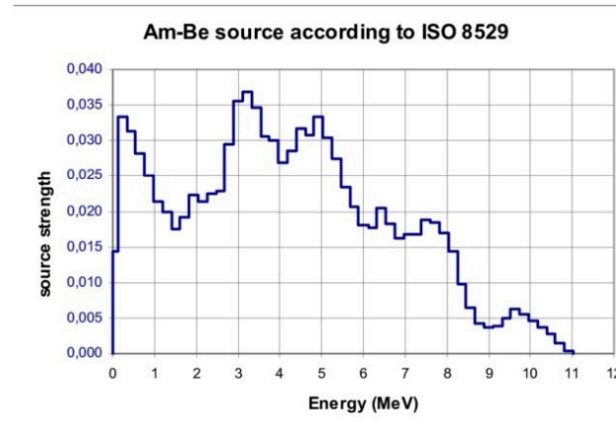
87%

AmBe setup & simulation

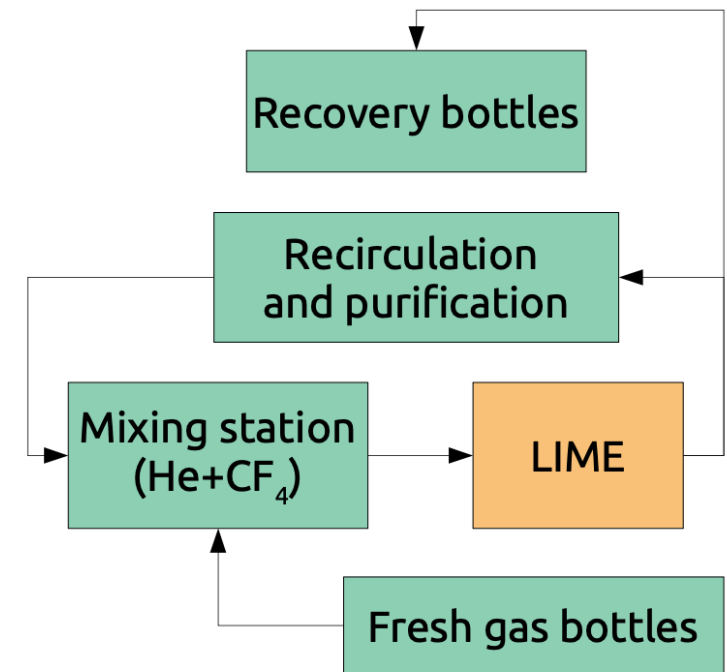
- AmBe neutrons*, 4.43 MeV gamma, 59 keV gamma isotropically emitted



- Neutron-induced events dominate over gamma, expected event rate (0.472 ± 0.004) ev/s (ER+NR), (0.146 ± 0.002) NR/s



- **GEM voltage** at 420 V (440 V) on each GEM in Run 1 (2, 3, 4)
- **Drift field** 800 V/cm
- **Gas system:**
 - Fresh gas flux: between 1 and 20 L/h (to study LY)
 - Recirculation system activated for part of Run 3 (to reduce gas consumption)
- **DAQ:**
 - 300 ms camera exposure
 - Periodic pedestal runs
 - Acquisition trigger on coincidence of 2 out of 4 PMTs



- MC energy spectra normalized to equivalent time of simulation
- Data normalized to total live time of runs, corrected for dead time

$$\tau = T_{cam} + N_{wf} \cdot t_{wf} \quad \text{Dead time: time for readout of camera and PMT waveforms}$$

$$r_{true} = \frac{k + m}{T_{run}} \quad \text{True event rate: measured rate + rate of events missed during dead time}$$

$$D = 1 + \frac{m}{k} = 1 + \frac{T_{cam} + R_{PMT} \cdot T_{window} \cdot t_{wf}}{T_{window}} \quad \text{Correction mainly depends on total trigger rate}$$

Run 1

($R_{PMT} \sim 30$ Hz)

$D = 1.42 \pm 0.06$

Run 2

($R_{PMT} \sim 3.5$ Hz)

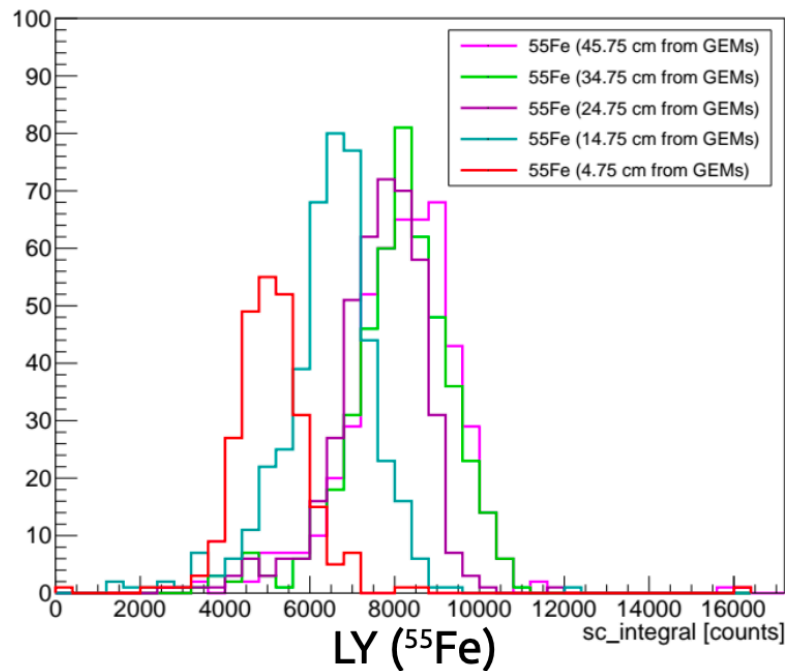
$D = 1.104 \pm 0.009$

Run 3

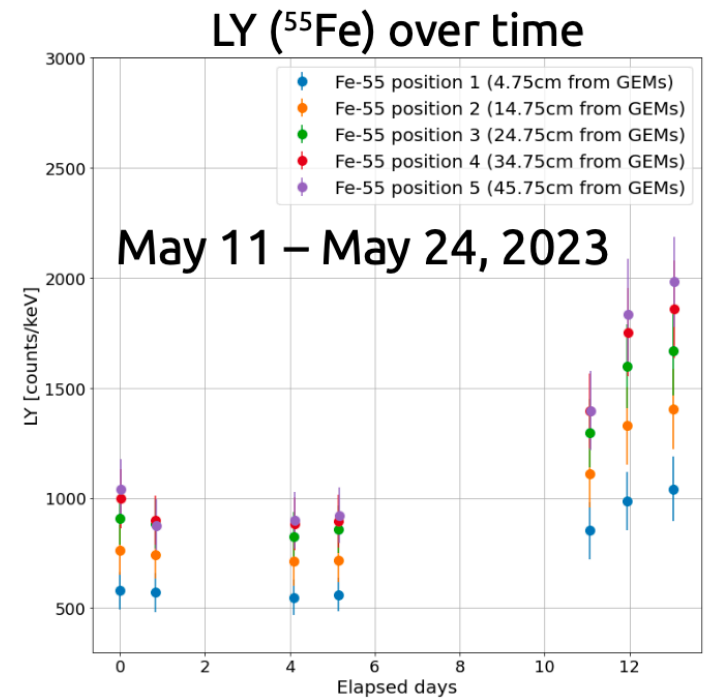
($R_{PMT} \sim 1.6$ Hz)

$D = 1.081 \pm 0.007$

- 5.9 keV X-ray ⁵⁵Fe source on top face of LIME, pointing towards gas
- **Daily dataset** taken in 5 positions (only Run 2,3)



Standard candle for **energy calibration**



LY monitoring over time



LIME background datasets considered



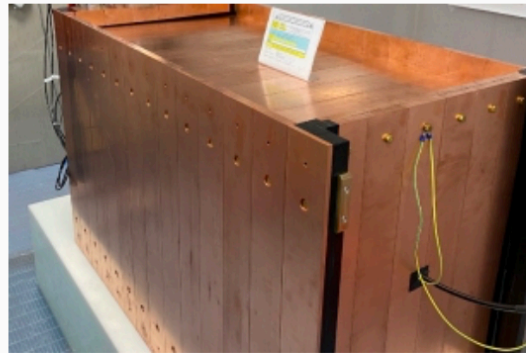
In order to maximise consistency among runs, Run3 data analysed are before the introduction of recirculation, i.e. fresh gas flux

A subset of runs were analysed for MC comparison (good LY, stability, gas system operation)



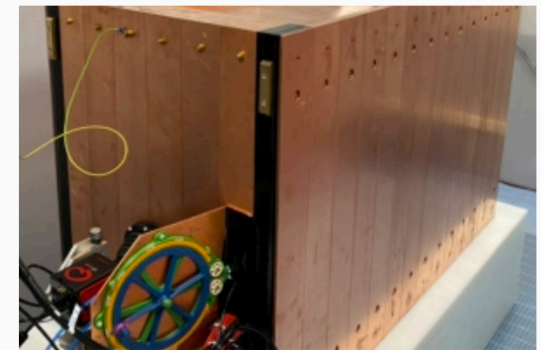
Run 1 (no shield)
Oct 8 - Dec 6 2022

285665 images
~49 hr, Dec 2–6 2022
10 L/h flux



Run 2 (4 cm Cu)
Feb 15 - Mar 9 2023

297992 images
~53 hr, Mar 6–9 2023
20 L/h flux



Run 3 (10 cm Cu)
May 5 - Nov 7 2023

171579 images
~53 hr, May 22–25 2023
20 L/h flux

Run4 is not discussed in this presentation since still on-going to these days



LIME background MC dataset



- From GEANT4 output, only main contributors were digitized and reconstructed: **external gammas, intrinsic radioactivity** of field cage rings, resistors, cathode, GEMs, acrylic box

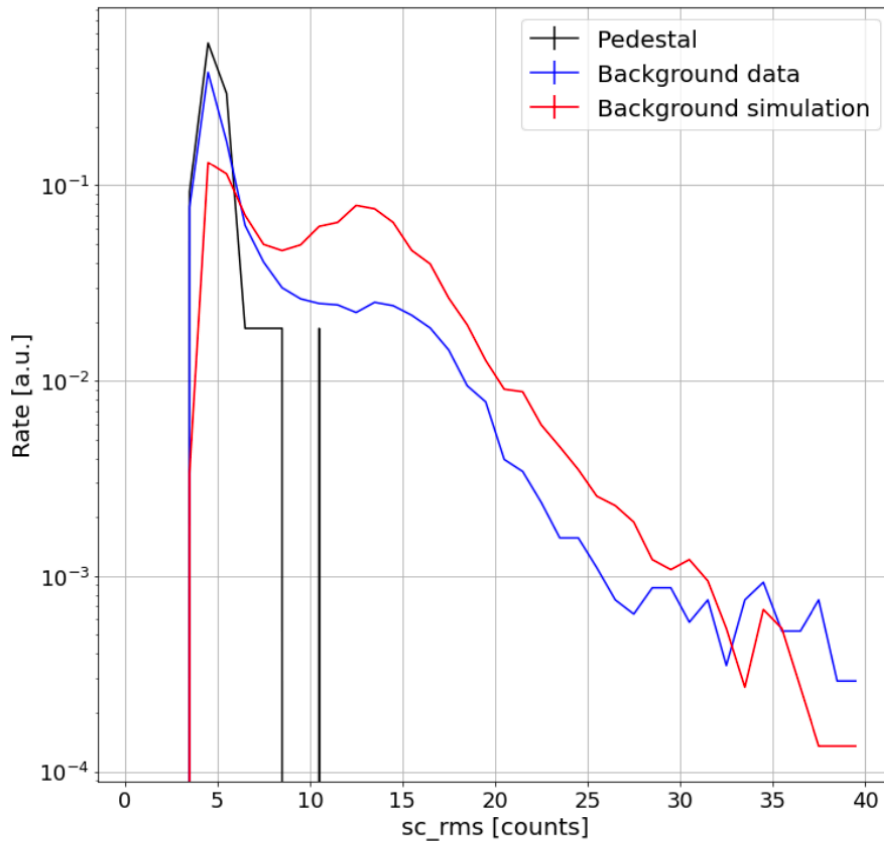
Digitised with standard CYGNO detector response simulation (see backup for details) with:

- Sensor noise added from real pedestal images *from underground data*
- Only events with energy deposits <800 keV digitized
- Diffusion parameters

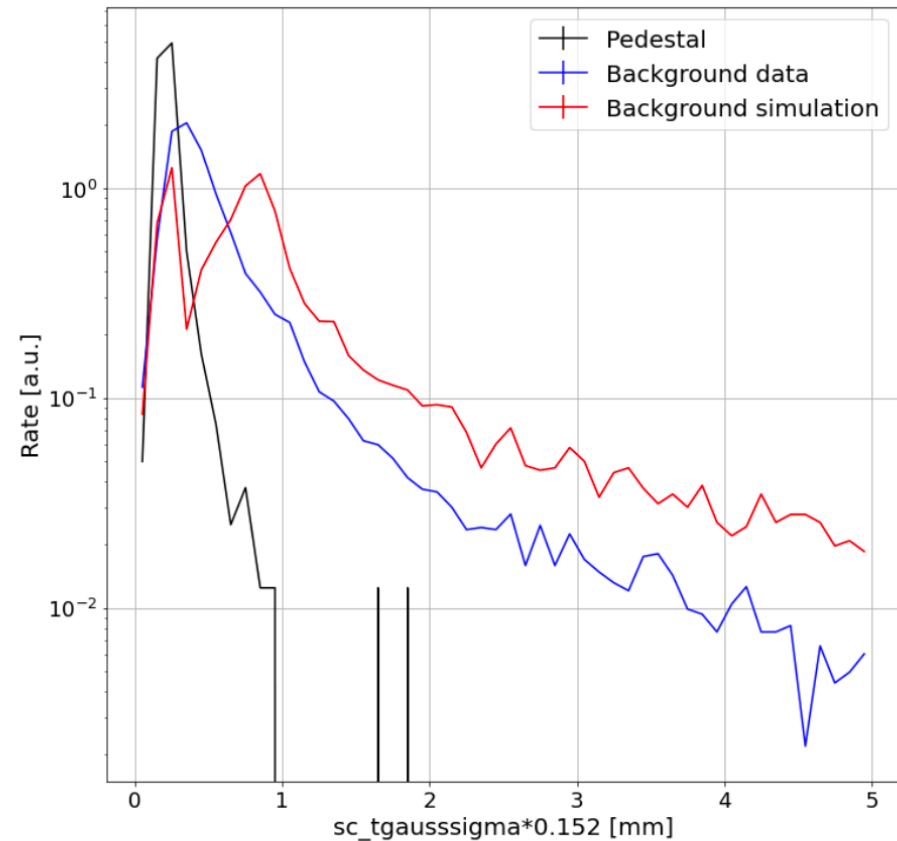
$$\sigma_{0T} = 350 \mu m, \quad \sigma_T = 115 \mu m / \sqrt{cm}, \quad \sigma_{0L} = 260 \mu m, \quad \sigma_L = 100 \mu m / \sqrt{cm}$$

- Images reconstructed with same reconstruction parameters as corresponding data

Corresponding phase	Shielding	V _{GEM} [V]	Background source	Equivalent time [s]
Run 1	None	420	Gammas	3129
			Radioactivity	432000
Run 2	4 cm Cu	440	Gammas	57813
			Radioactivity	432000
Run 3	10 cm Cu	440	Gammas	172700
			Radioactivity	432000

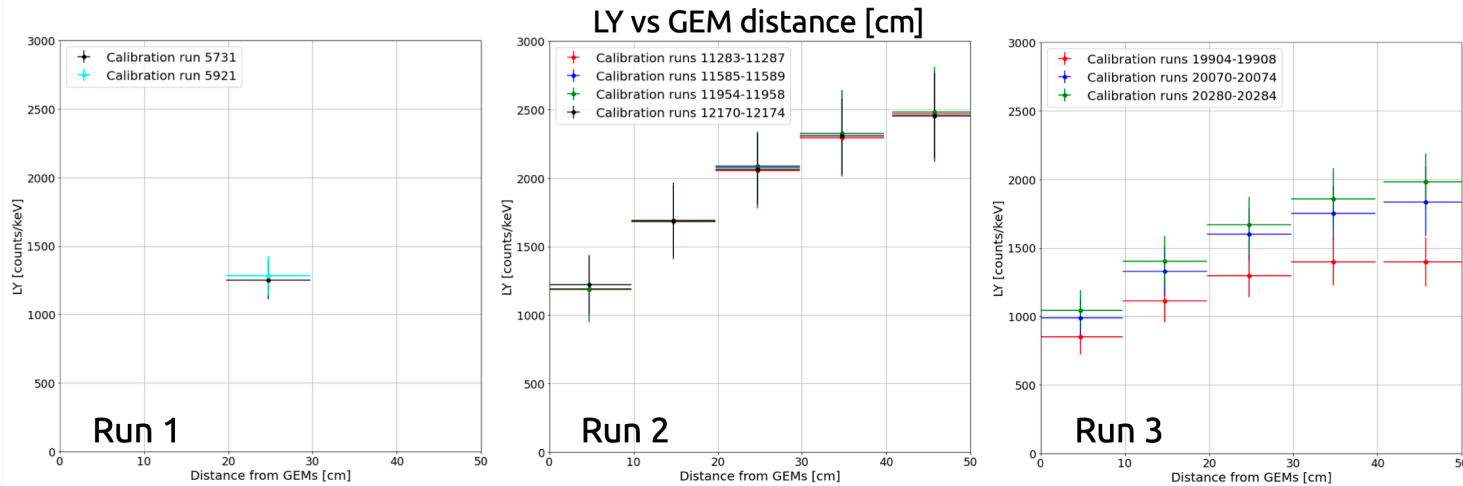


Standard deviation of the counts of all the pixels in a cluster. The lower it is, the more uniform the pixel intensity (and therefore the dE/dx) of the track is



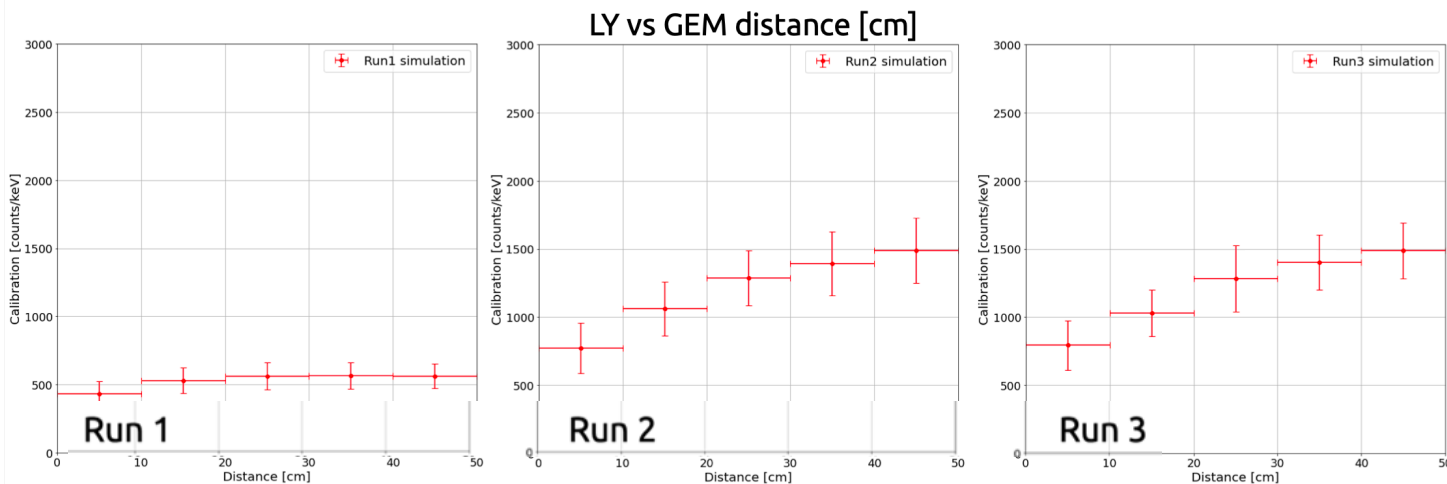
Sigma of the Gaussian transverse profile. It is a measurement of the spread of the energy deposit wrt the original path of the event in the gas

Run1, Run2 and Run3 data energy calibration



- LY depends on distance from GEM (Z)
- Event Z position evaluation still preliminary and not precise enough yet (about 10 cm resolution) to correct data
- Random uniform Z extraction, random Gaussian LY extraction, bootstrap sampling

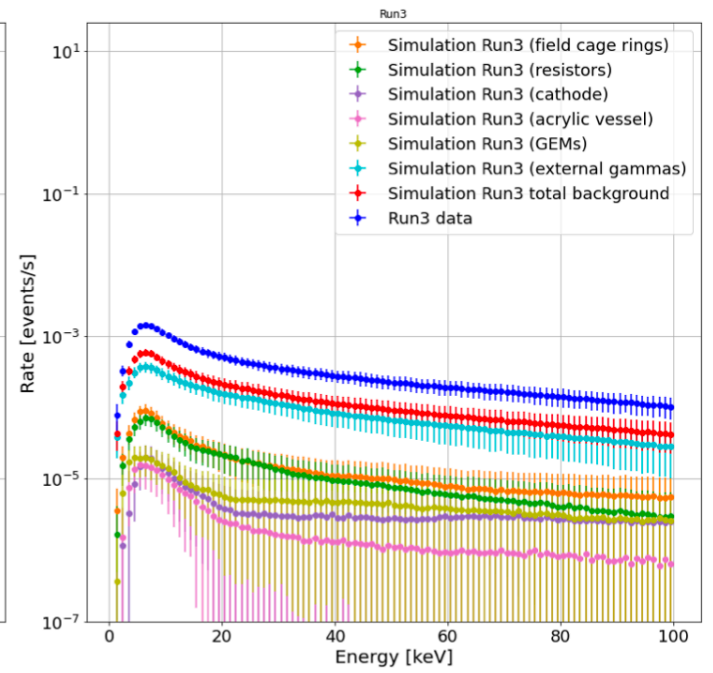
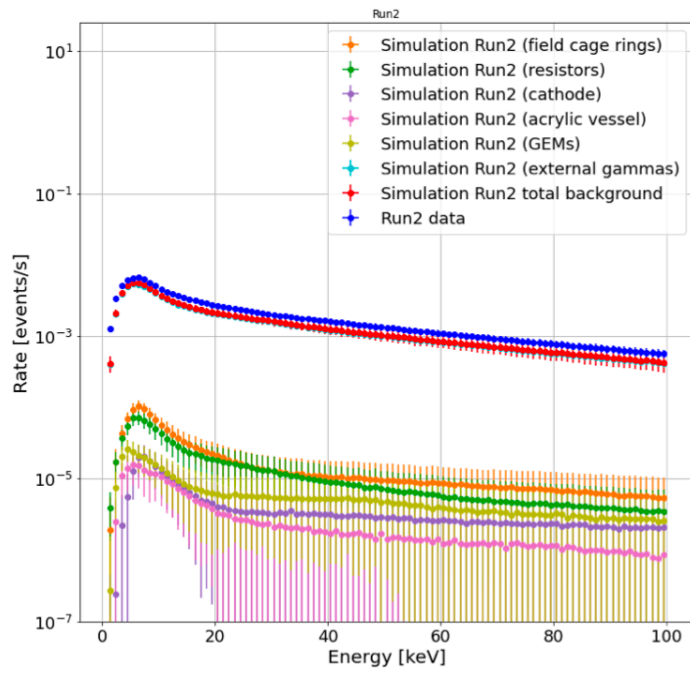
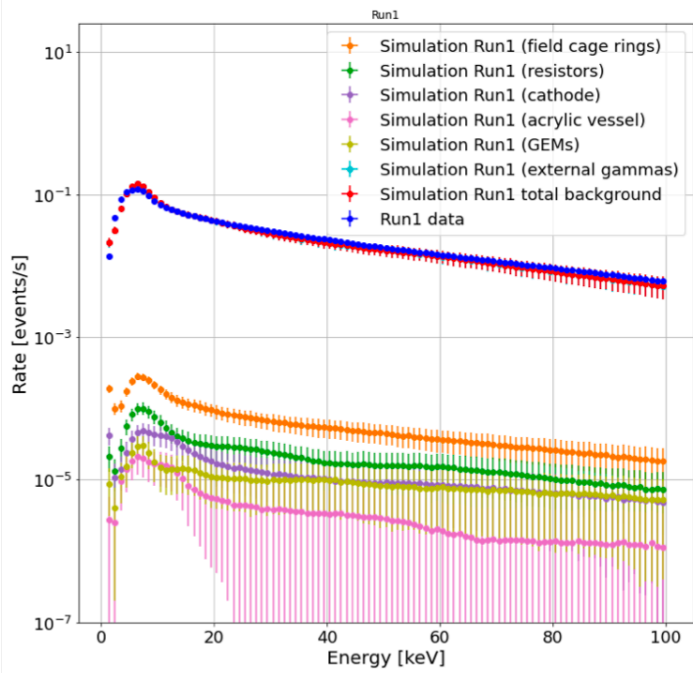
Run1, Run2 and Run3 MC energy calibration



- LY from MC sample with energy between 2 and 10 keV
- Same method used for data, except for LY variation over time
- Lower LY observed in MC (optimised on overground data), strongly dependent on specific data conditions

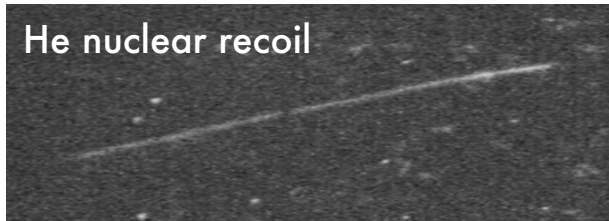
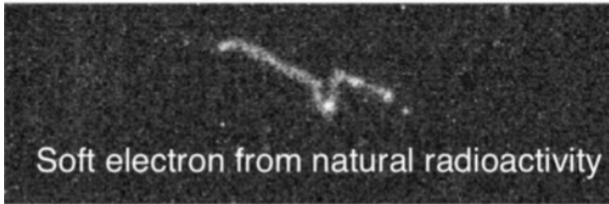


Data/MC comparison by background component

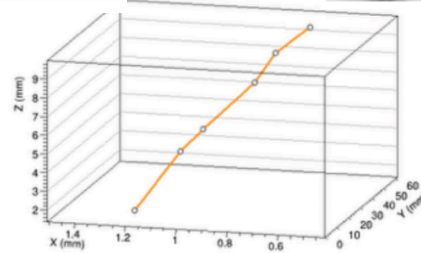
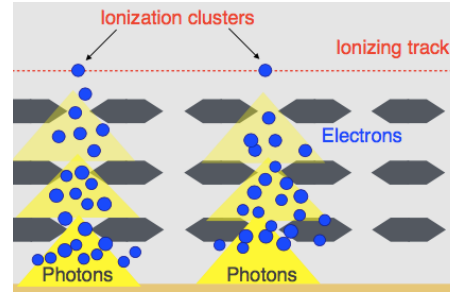


sCMOS:

high granularity
X-Y + energy measurements



JINST 13 (2018) no.05, P05001

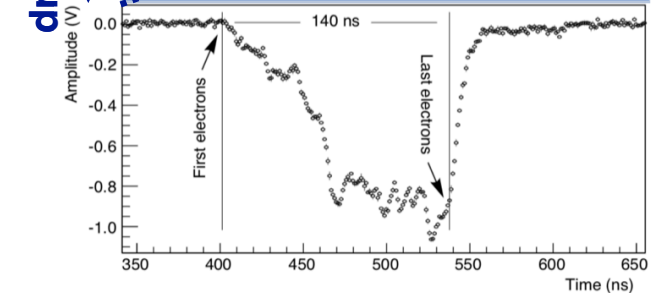
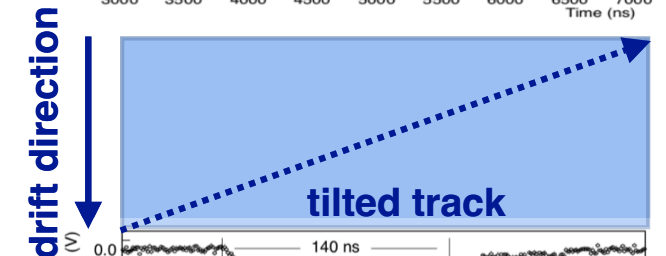
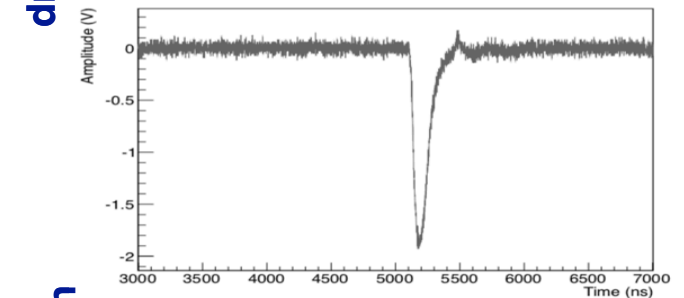
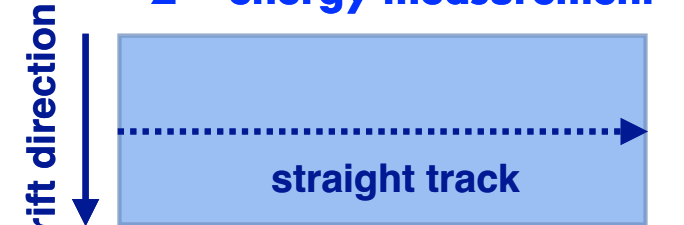


+ SF₆ for negative ion drift



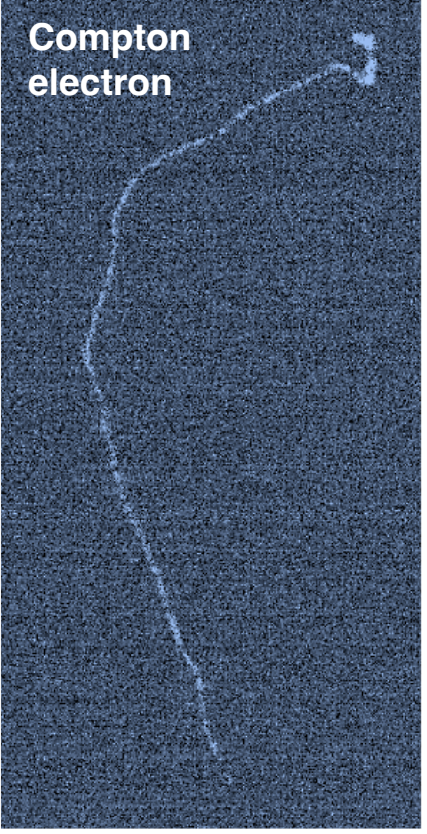
PMT:

integrated
Z + energy measurement

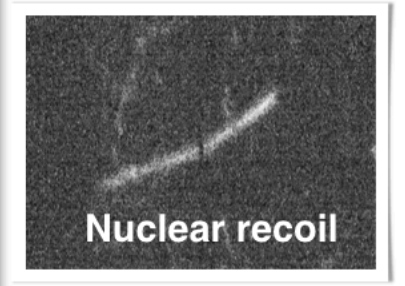
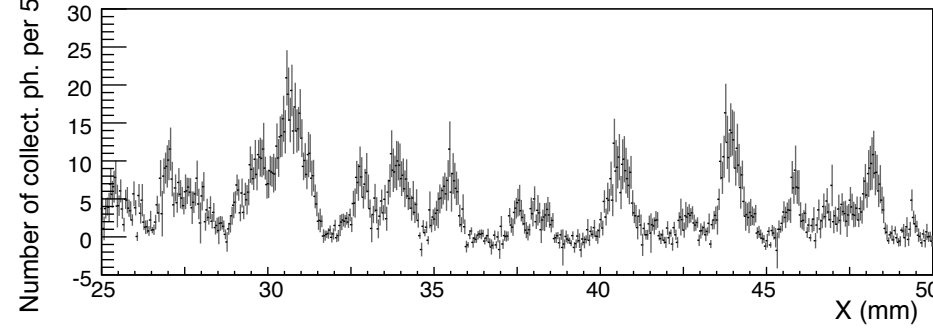
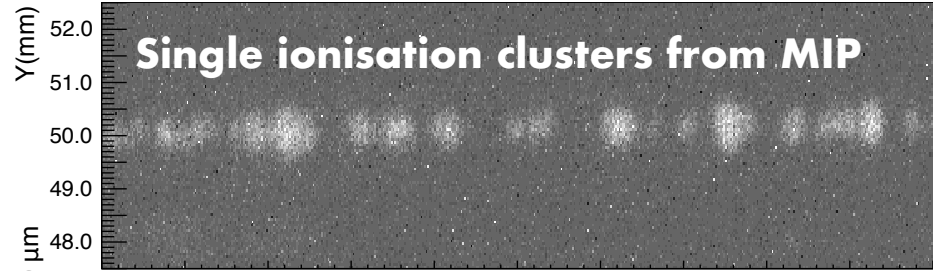
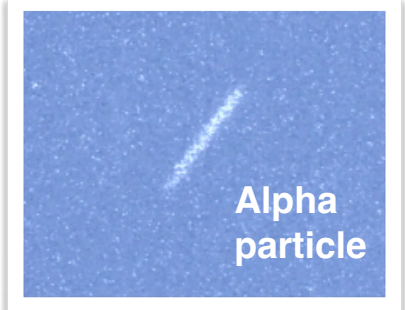
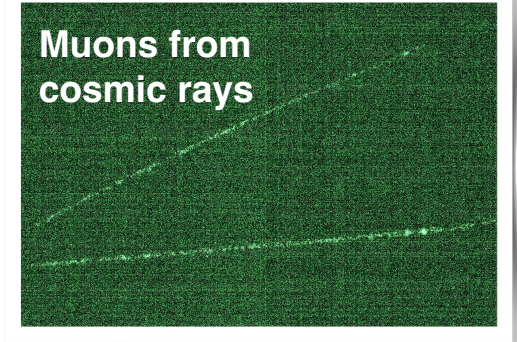
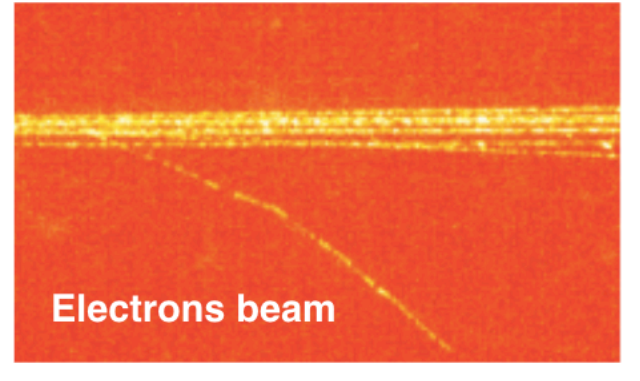
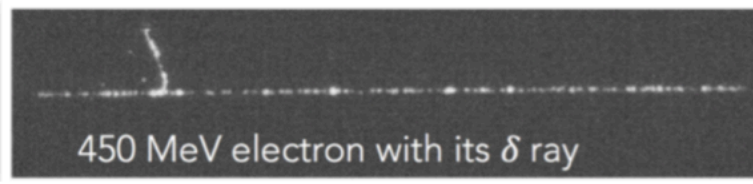
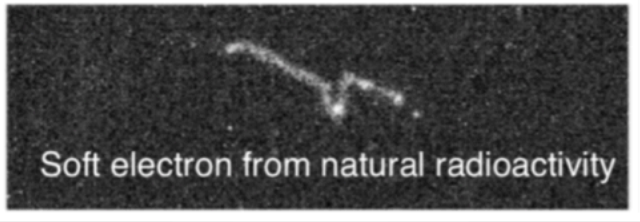


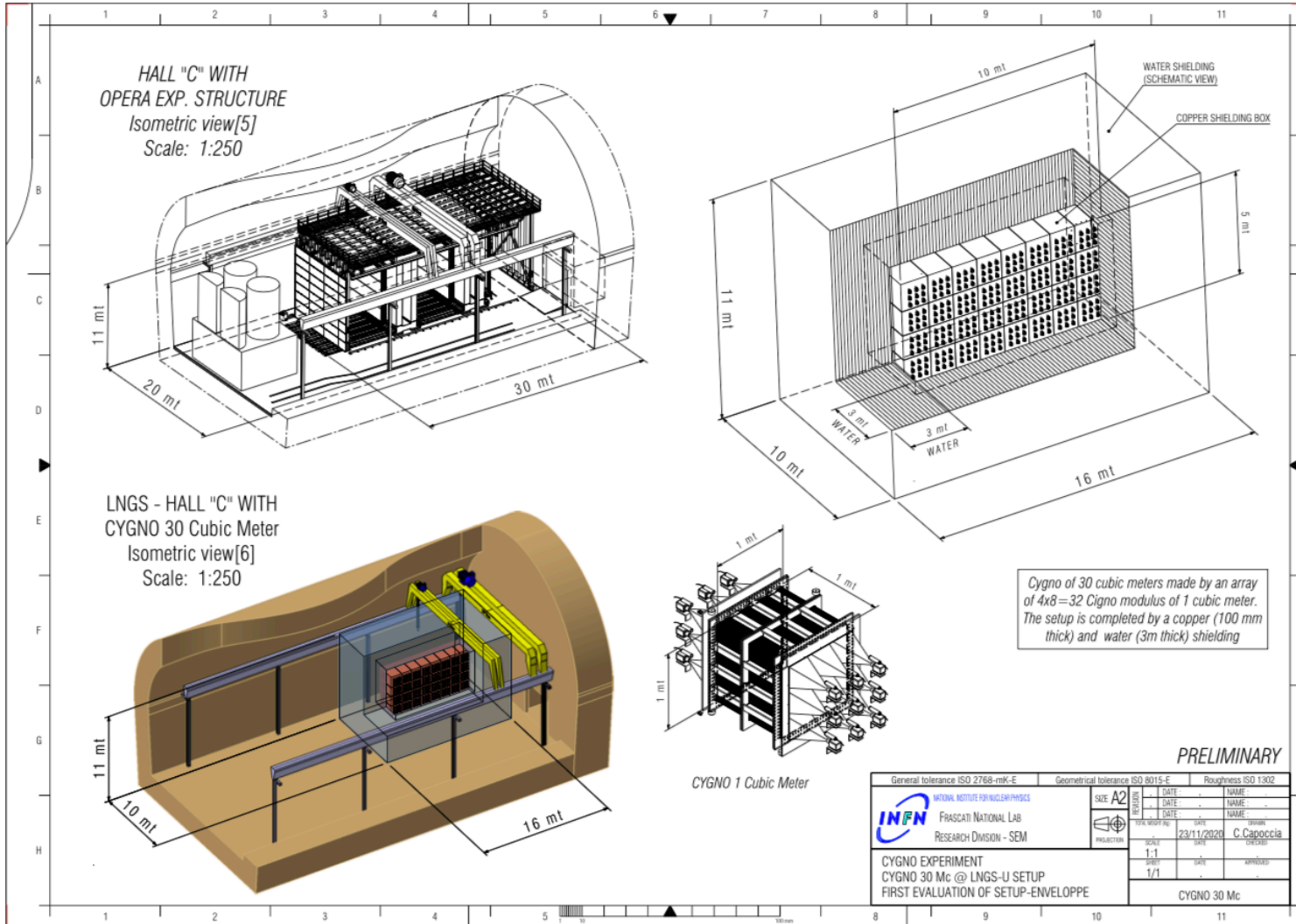
- **1/3 noise w.r.t. CCDs**
- **Market pulled**
- **Single photon sensitivity**
- **Decoupled from target**
- **Large areas with proper optics**

<https://web.infn.it/cygnus/>



He:CF₄ @ 1 atm





PHASE 2:
30 m³ Experiment

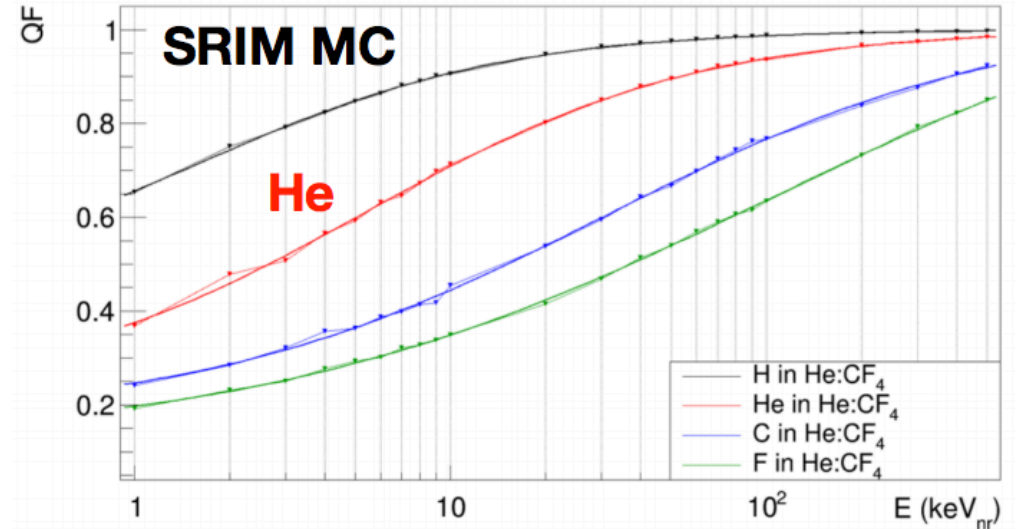
2026..
LNGS

CYGNO_30

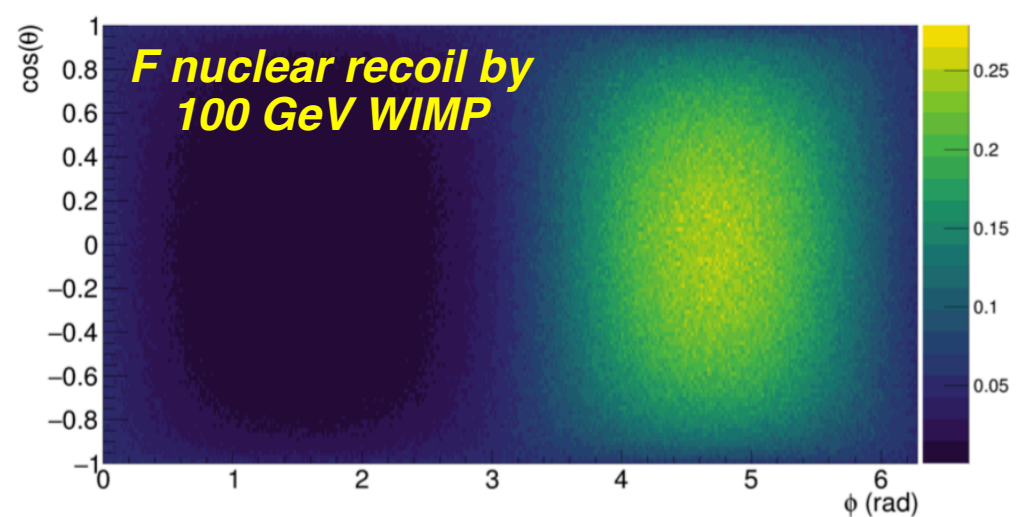
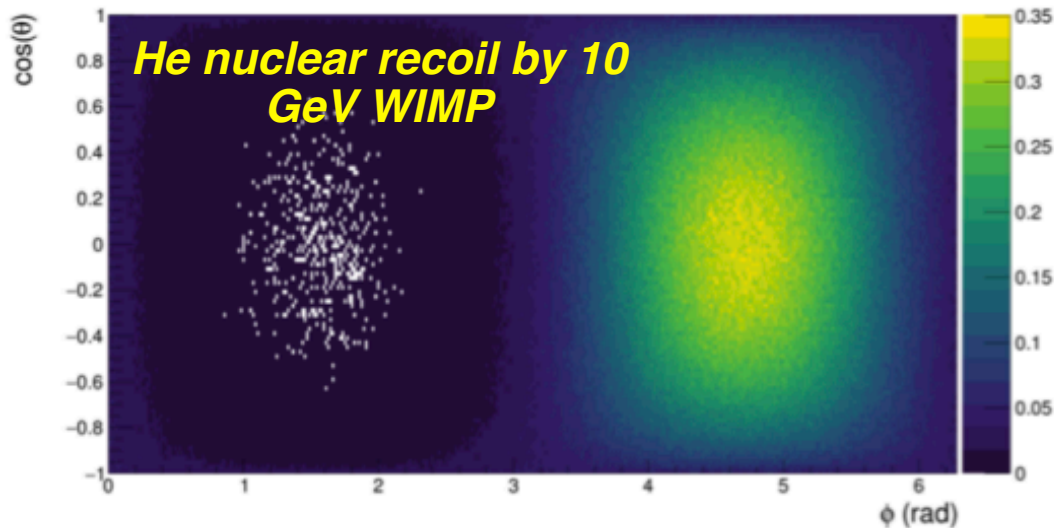
- Physics research

- ▶ Use 1 keV_{ee} threshold
- ▶ Evaluate QF with SRIM
- ▶ Introducing **angular distribution** as discriminating
- ▶ Full head/tail recognition
- ▶ Using a 30 deg resolution

Quenching Factor



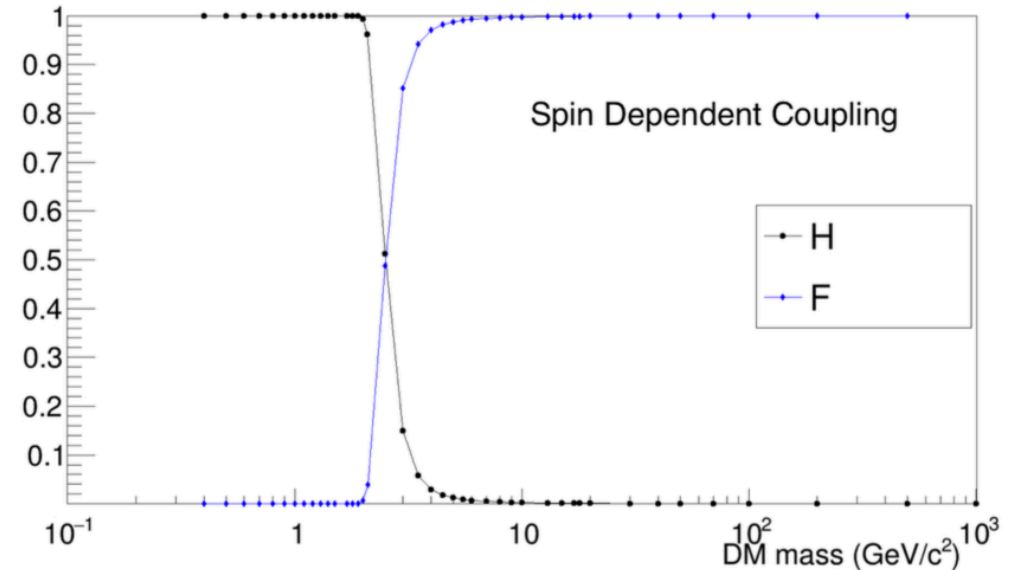
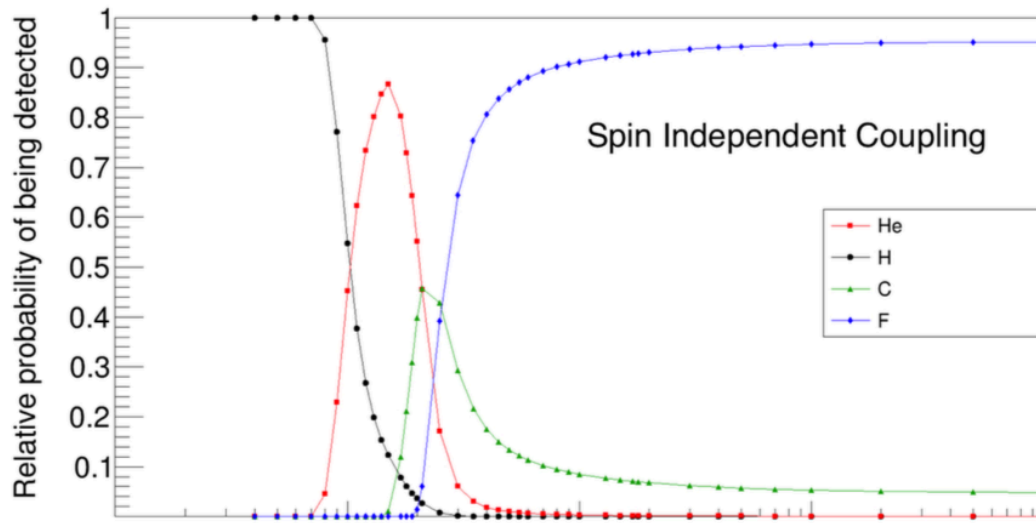
Examples of expected measured angular distribution in Galactic coordinates



Since CYGNO is a multi-target DM experiment, both the kinematics of the expected DM-nucleus interaction and the expected rate calculation influence the probability of each element to be detected differently as a function of the DM mass

The region of the DM velocity distribution accessible to detection is limited at lower values by the energy threshold and at higher values by the local escape velocity (here taken as 544 km/s)

	Minimum detectable DM mass for 0.5 keV _{ee} energy threshold	Minimum detectable DM mass for 1 keV _{ee} energy threshold
H	300 MeV/c ²	500 MeV/c ²
He	700 MeV/c ²	1 GeV/c ²
C	1.4 GeV/c ²	1.9 GeV/c ²
F	1.9 GeV/c ²	2.5 GeV/c ²



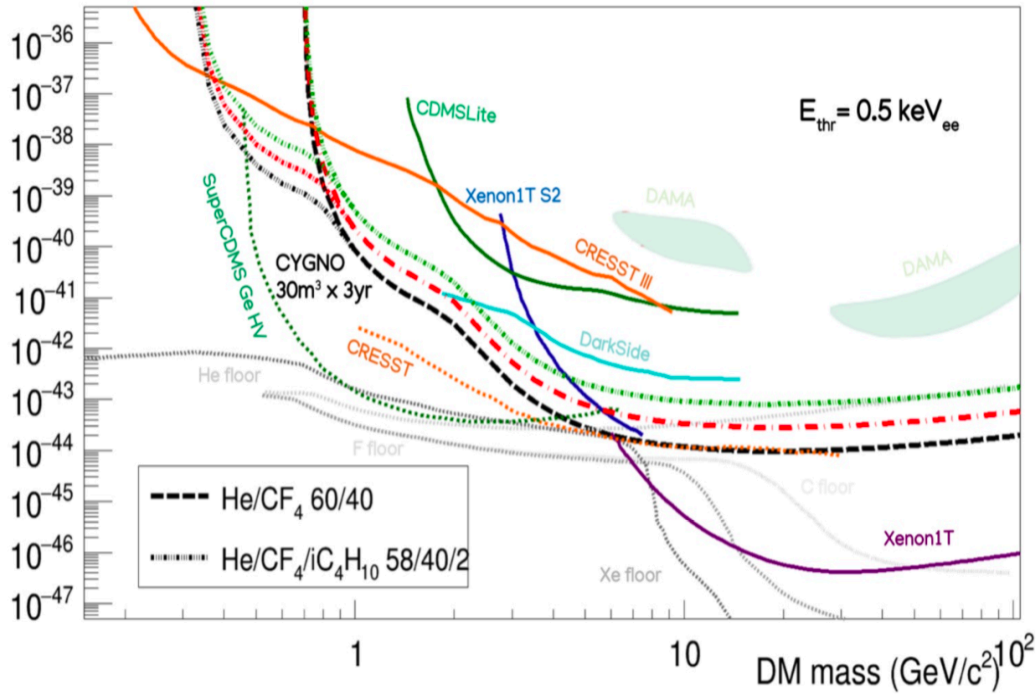
Target nuclei relative probability of being detected for 1 keV_{ee} energy threshold



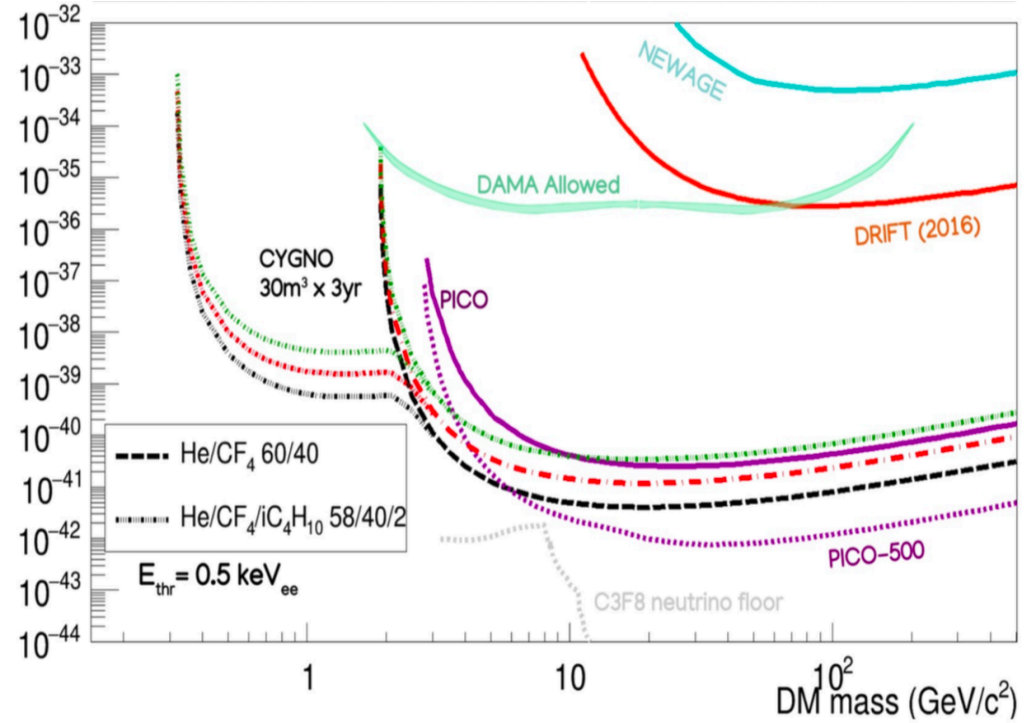
CYGNO 30 m³ preliminary sensitivity studies



Spin Independent

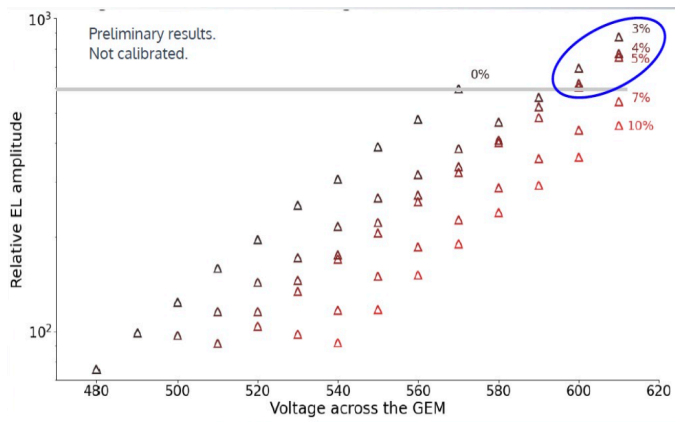


Spin Dependent

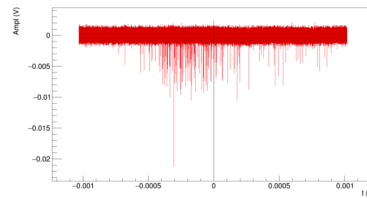
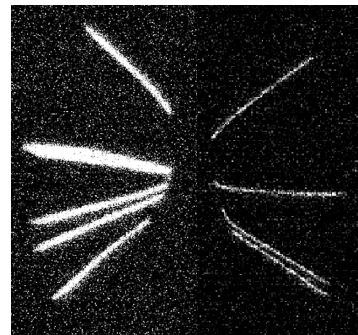


- $n_{\text{BKG}} = 10^2$
- - - $n_{\text{BKG}} = 10^3$
- ⋯ $n_{\text{BKG}} = 10^4$

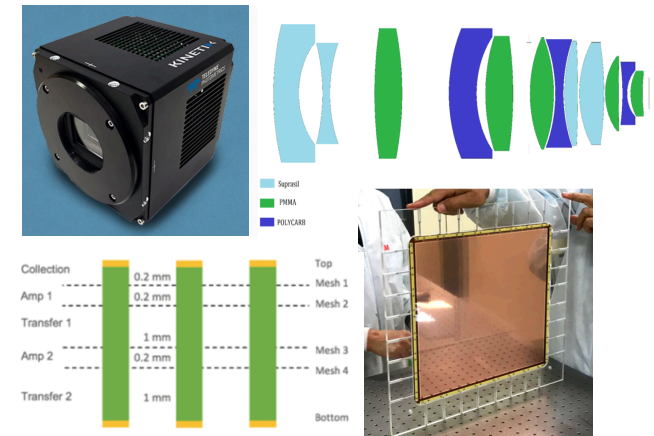
Improve sensitivity at low < 1 GeV WIMP masses by means of Hydrogen target



Improve tracking by means of Negative Ion Drift operation



Minimise internal radioactivity and optimise optical system & amplification



He:CF₄:SF₆
59:39.4:1.6



- First ever demonstration of NID operation at atmospheric pressure with optical readout of both sCMOS and PMT
- 5 MeV alpha particles and possibly Ba133 observed
- Opens a completely new window of possibility of optimisation of the gas mixtures
- Systematics studies ongoing

- Develop custom sCMOS sensor with photon sensitivity & radioactivity budget optimised for CYGNO
- Realisation of custom lens with large aperture & low radioactivity
- Optimisation of amplification structures in terms of gain and radioactivity budget

- R&D with iC₄H₁₀ and CH₄ demonstrated good light yield achievable
- Future studies on Fluorine-based molecule with H (CHF₃, CH₂F₂)
- R&D work on eco-friendly gas mixture as substitute to CF₄ (doi: 10.1109/NSS/MIC42101.2019.9059721)