



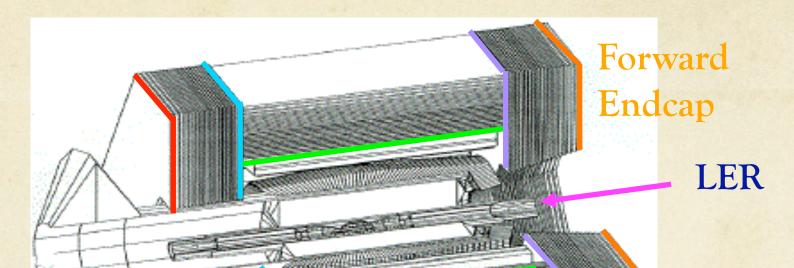
'IFR Background Report

Valentina Santoro INFN Ferrara

II-SuperB Collaboration Meeting 14 Dec 2011

Hot regions





Backward

Endcap

HER

Barrel: innermost layers, mostly neutrons

FWD encaps (hottest region): inner layer and outer layers (BEAM halo),

neutrons, electron and photons

BWD encaps: inner layer and small radii

Barrel

Outline

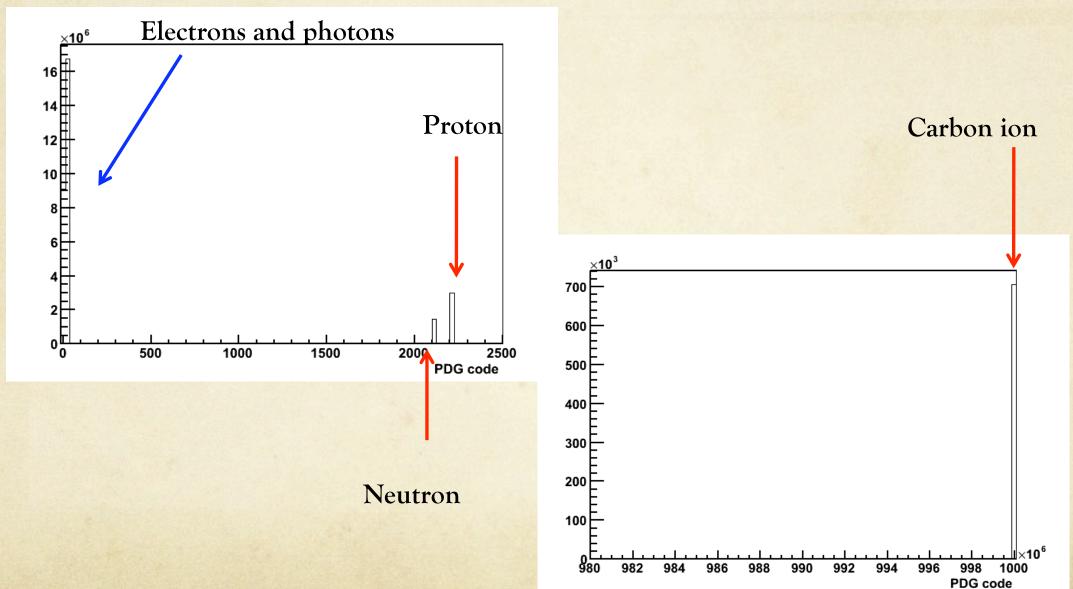


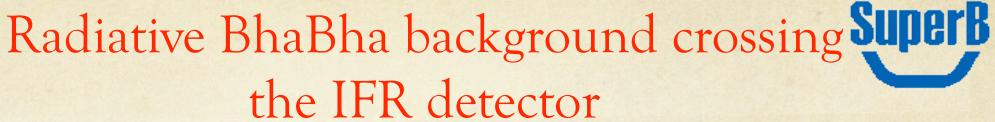
- Radiative BhaBha Background Studies (neutrons, photons and electron)
- ✓ Touschek background (neutrons, photons and electron)
- ✓ Pair background (neutrons, photons and electron)
- ✓ Background Studies on our FEEs

Beam Composition

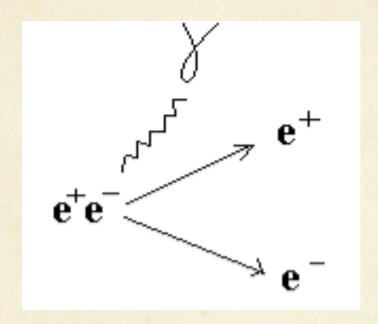
For BhaBha, Touschek and Pair events the particle crossing the IFR are photons, electron, protons, neutrons and heavy nuclei











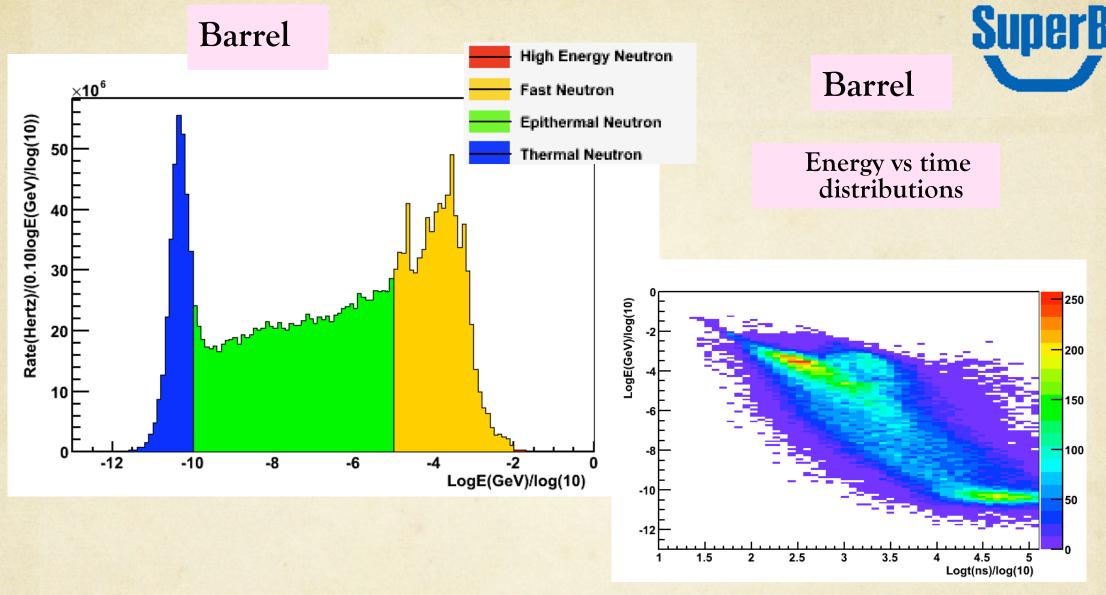
Neutrons



Why do we have to worry about neutrons

- Neutrons damage silicon devices → Neutrons damage
 SiPM
- The silicon damage function has a strong dependance on the energy spectrum therefore we scaled all the doses in this presentation to 1MeV equivalent accordingly to ASTM E 722 93.

Neutron Energy Distributions for Radiativa BhaBha wents



The Energy distribution for FWD and BWD Endcap are similar

Neutron Rates

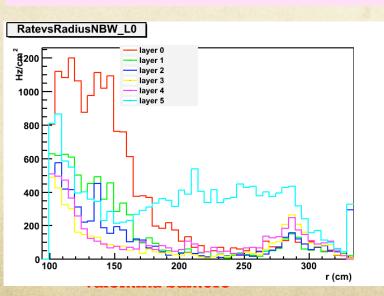
Rate vs Z-coordinate for Barrel

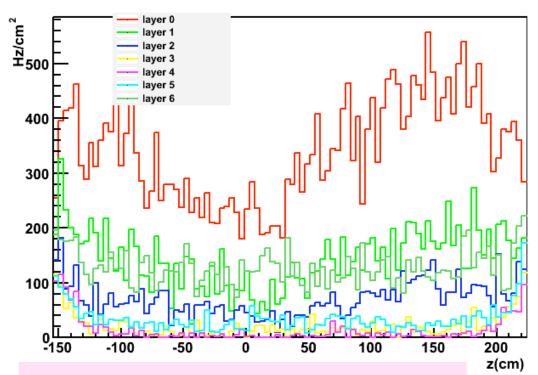
Cunny

Rate of $450\text{Hz/cm}^2 \rightarrow \text{about}$ $3x10^9 \text{ neutrons/cm}^2 \text{ for a year}$

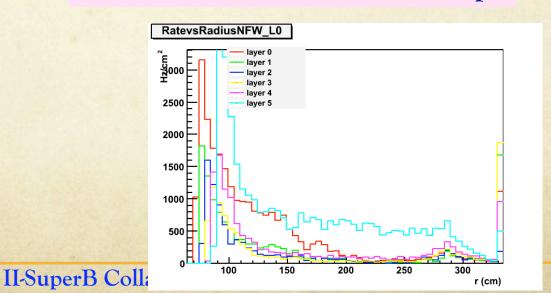
All the rate are normalized to 1MeV energy

Rate vs radius for BWD Endcap





Rate vs radius for FWD Endcap



Photons



Why do we have to worry about photons

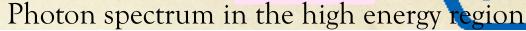
High Energy Photons convert in e⁺e⁻ that produce signal in the detector

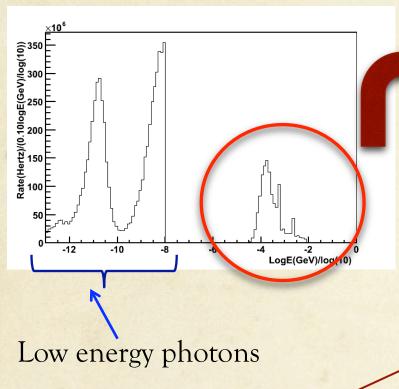
Photon Energy Distributions for Radiativa BhaBha events



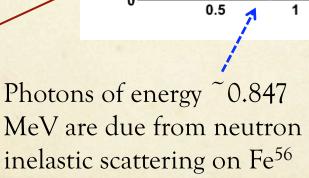
Barrel

anarmy ragion





Photons of energy ~0.512 MeV are from annihilation radiation



6000

5000

4000

3000

2000

1000

Photons of energy ~2.223 MeV are from neutron capture on Hydrogen

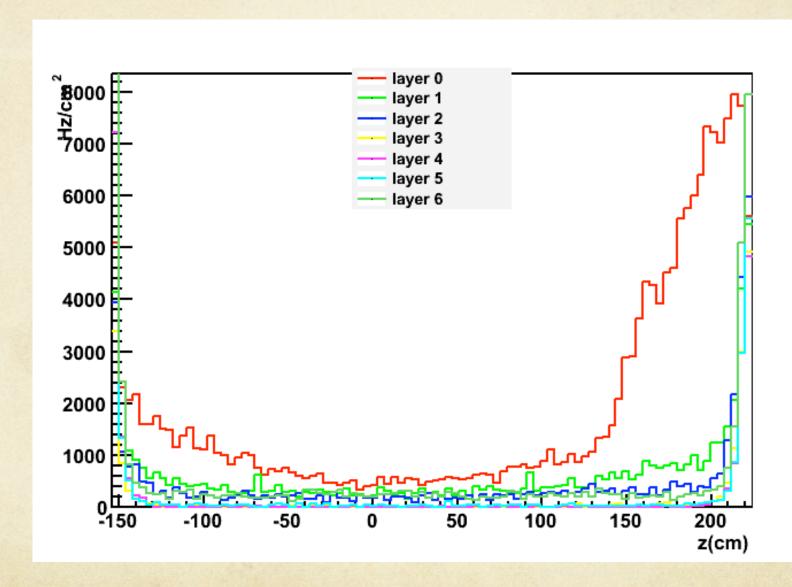
E(MeV)

The Energy distribution for FWD and BWD Endcap are similar

γ rates for Different layers

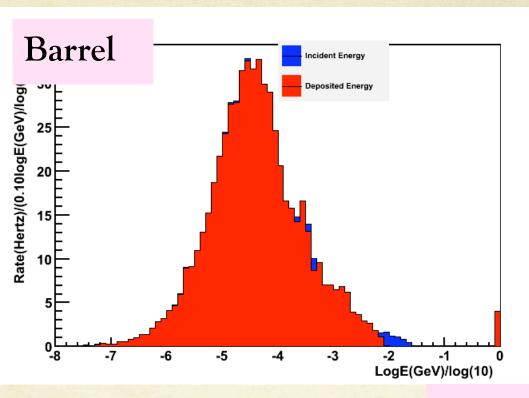
Rate vs Z-coordinate for Barrel





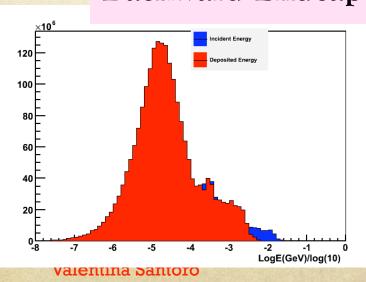
This rate is due only to photon with E > 10keV

Electron Energy Distributions for Radiativa BhaBha wents

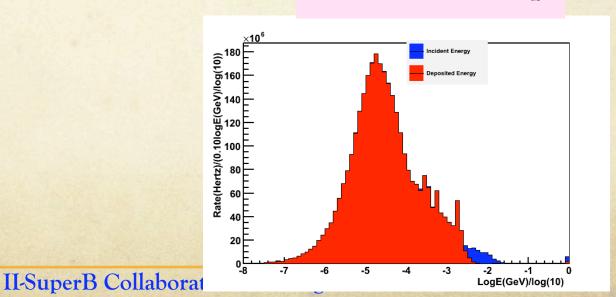




Backward Endcap



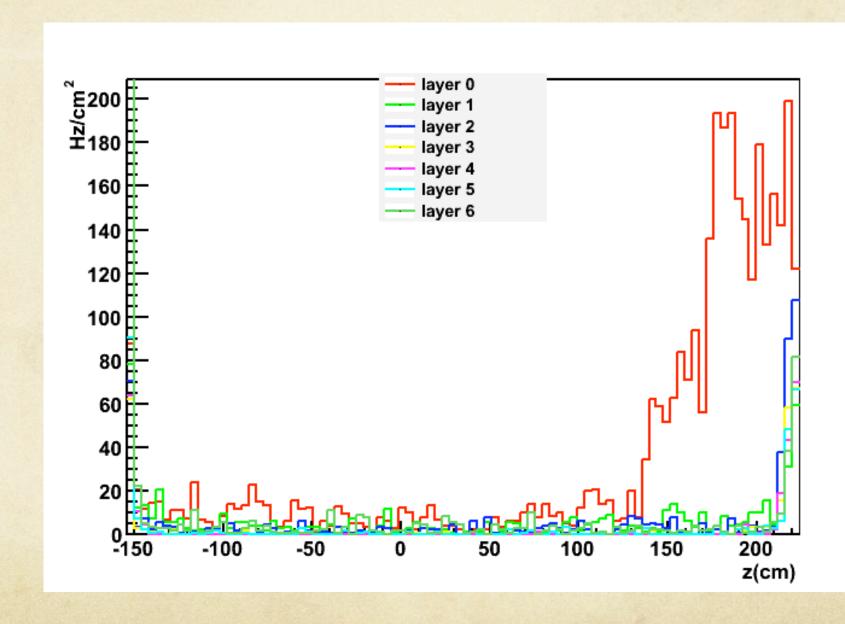
Forward Endcap



Electron rates for Different layers

Rate vs Z-coordinate for Barrel





Summary on BhaBha studies

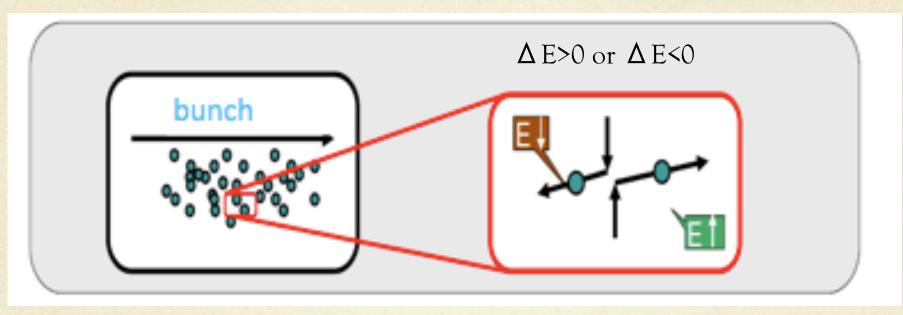


- ✓ The neutrons rates are very high and dangerous for our Sipm
- ✓ The photons and electrons rates are high but they should not be a problem

Touschek events studies

SuperB

Touschek scattering results from a Coulomb collision of two relativistic electrons in a particle beam, producing an instantaneous change in particle energy

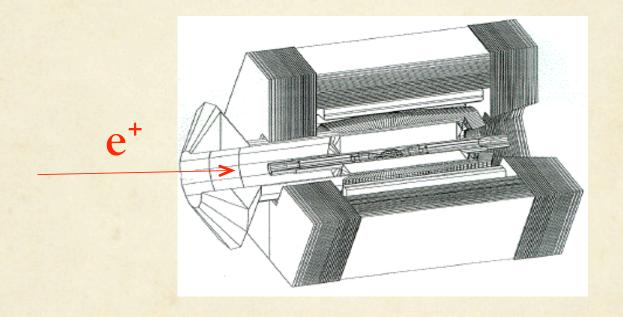


Why we don't like Touschek events:

Scattered e⁺/e⁻ goes off trajectory -> lost at beam pipe wall near IP ->creates shower ->reach detector



Touschek events HER

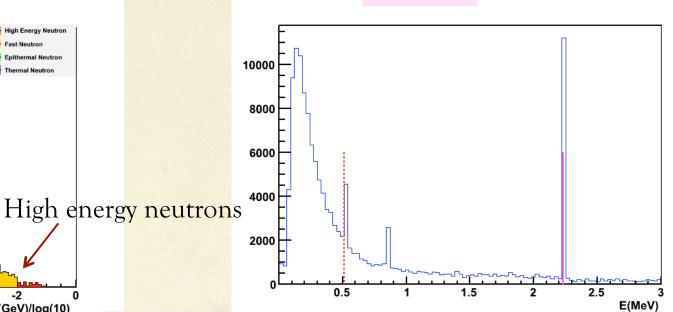


Energy distribution for Touschek events



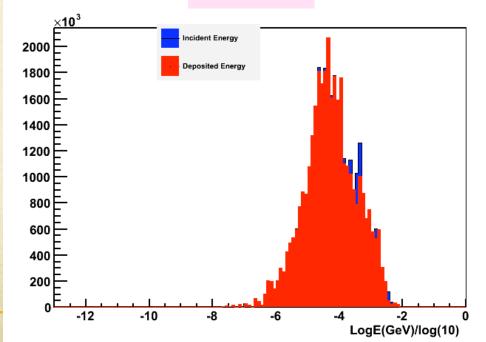
-2

LogE(GeV)/log(10)



Photons





-6

Rate(Hertz)/(0.10logE(GeV)/log(10))

-12

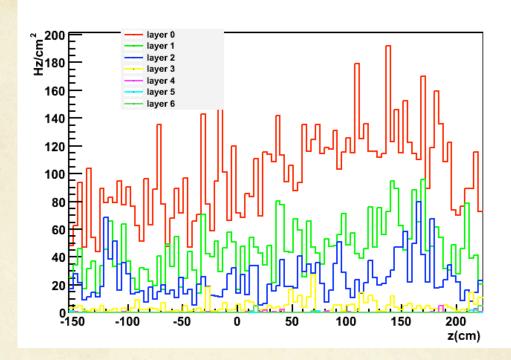
-10

Neutron Rates for Touschek HER events

Rate vs Z-coordinate for Barrel

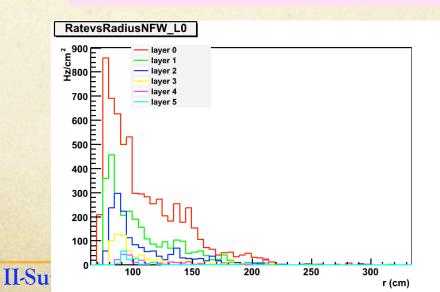


Rates are small but not negligible



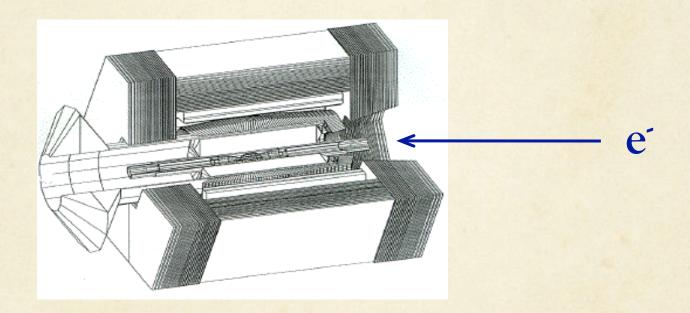
Rate vs radius for BWD Endcap

Rate vs radius for FWD Endcap



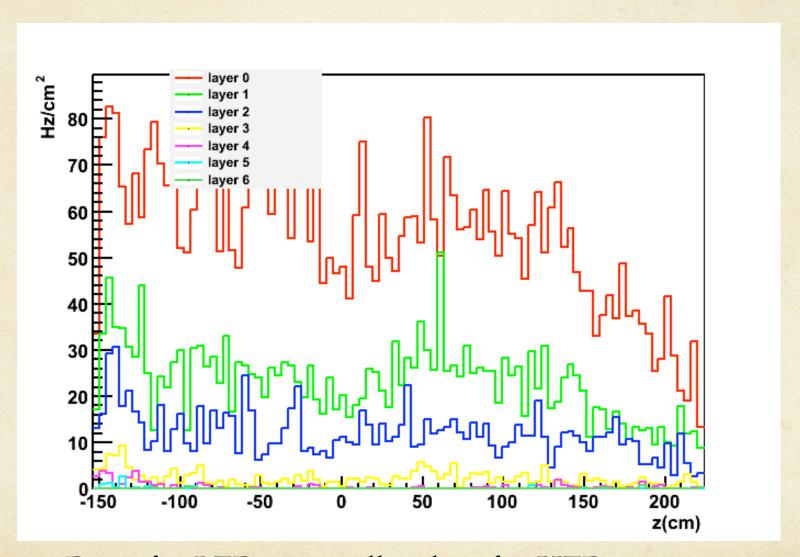


Touschek events LER



SuperB

Rate vs Z-coordinate



Rates for LER are smaller than for HER

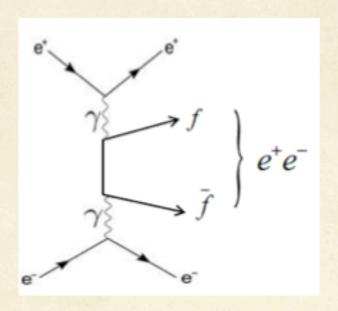
Summary on Touschek studies



- ✓ Touschek background studied for the HER and LER
- Results for the HER and LER show that the rate are small compared to the BhaBha one.

Pair Production

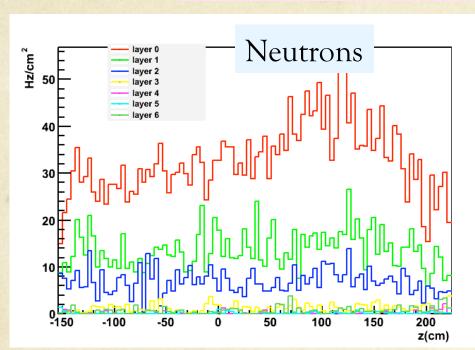


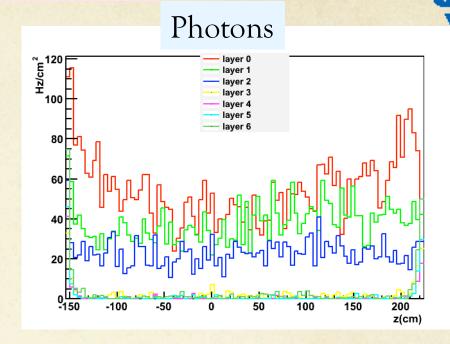


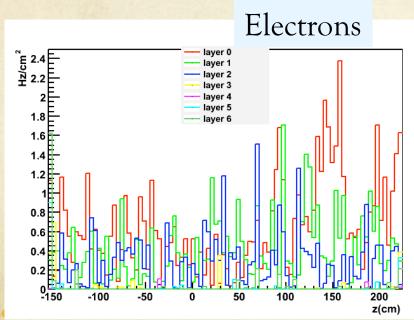
$$e^+e^-\rightarrow e^+e^-\gamma\gamma\rightarrow e^+e^-e^+e^-$$

Neutron, Photon and Electrons Rates for Pair events

Rate vs Z-coordinate for Barrel





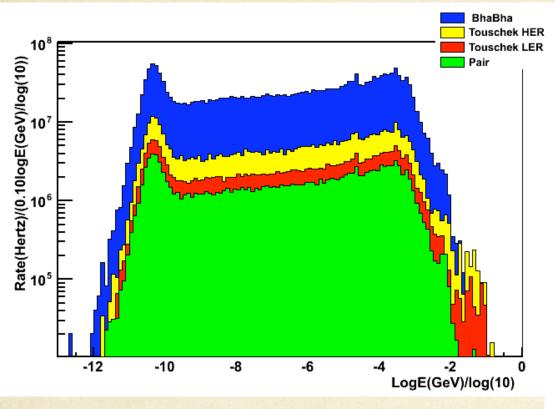


Pair background (neutrons, photons and electrons) is small compared to the BhaBha one and to the Touschek events

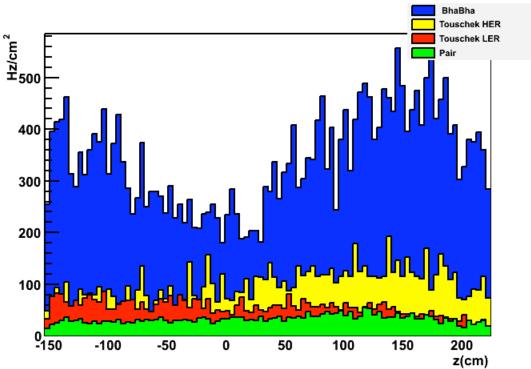
Neutron Rates from different background Sources

Energy distribution: Barrel





Rate for Barrel LO

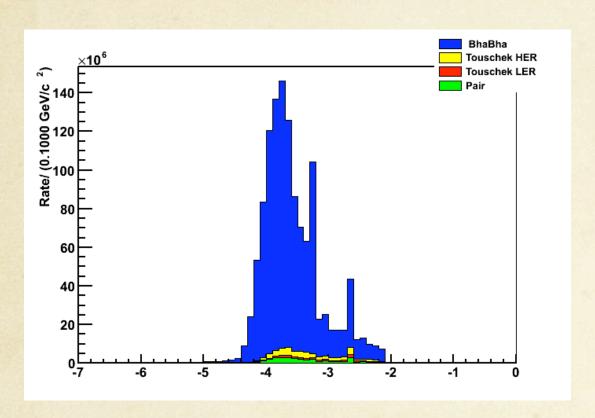


Valentina Santoro II-Super

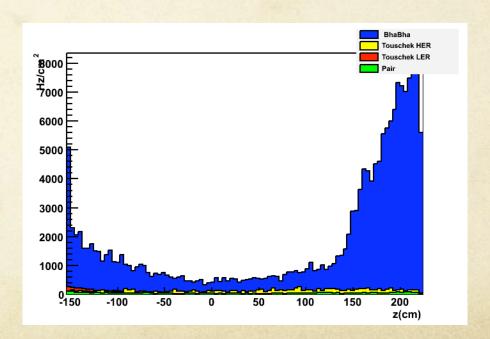
Gamma Rates from different background Sources

Energy distribution: Barrel





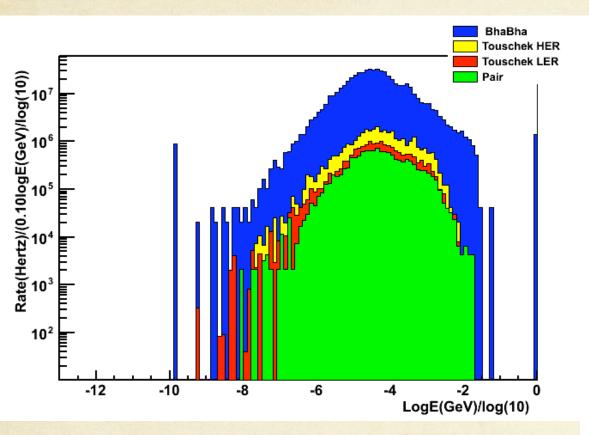
Rate for Barrel LO



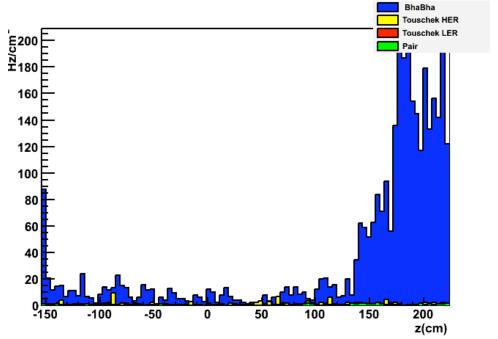
Electron Rates from different background Sources

Energy distribution: Barrel





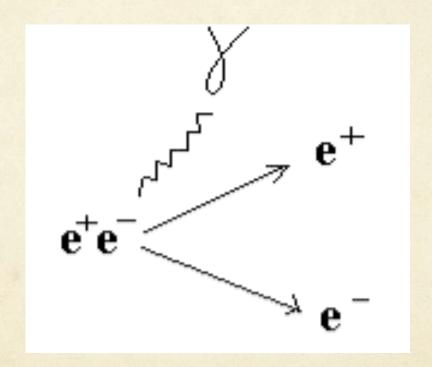
Rate for Barrel LO



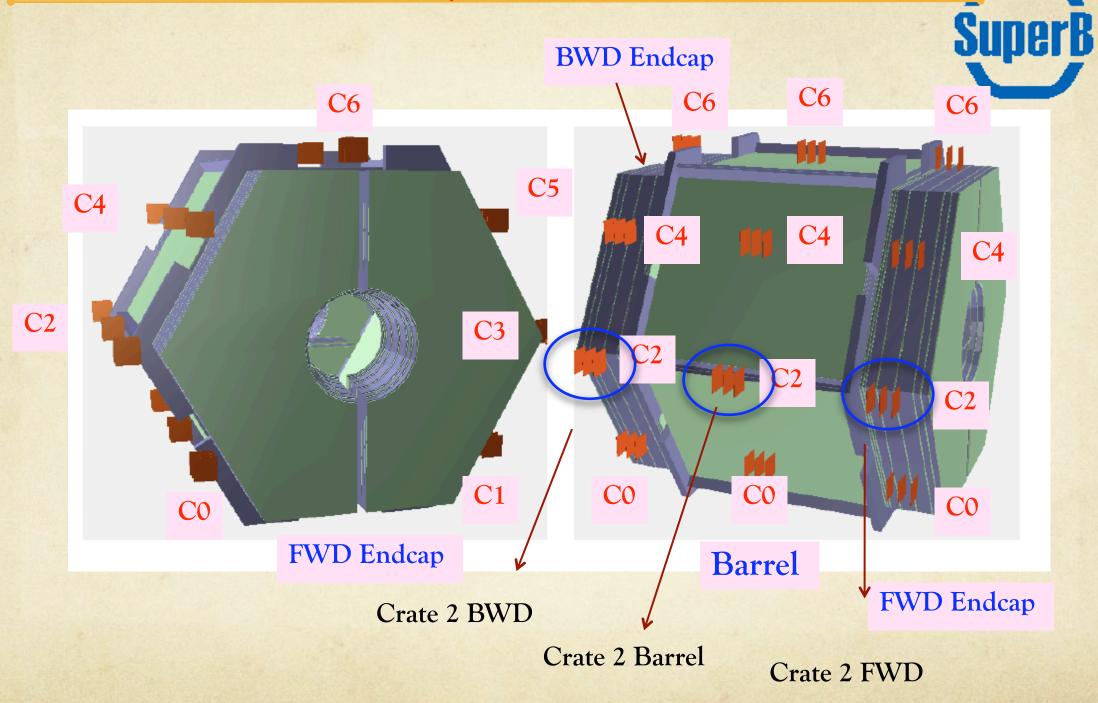
Valentina Santoro II-SuperB C



Radiative BhaBha background crossing the IFR FEE boards

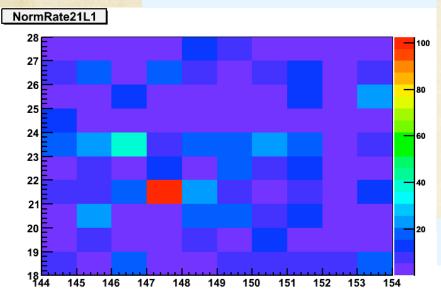


Present layout of the IFR crates

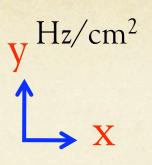


Neutron Rates for FEEs Electronics

2D view of one FEE

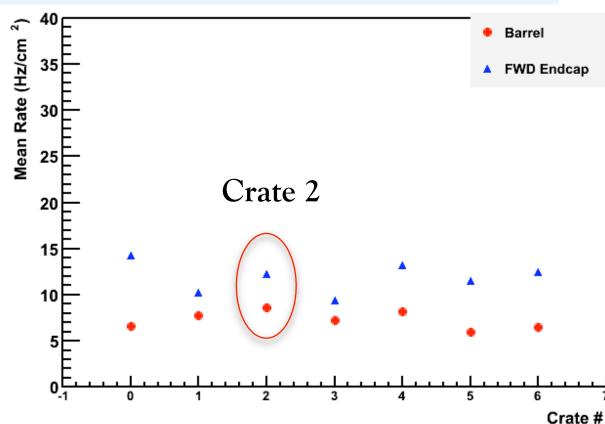


Crates located in the FWD have systematically higher rates compared to that one in the Barrel

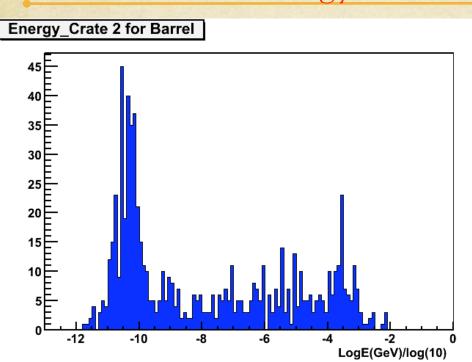




Mean Rate for each FEE in different Crates

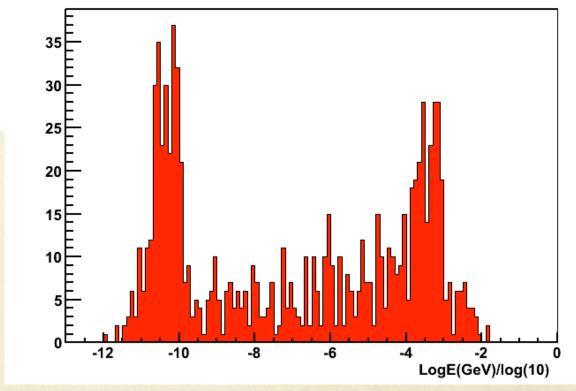


Neutron Energy Distribution for FEEs Electronics





Energy_Crate 2 for FWD



The Energy distribution

change for Electronics crate in

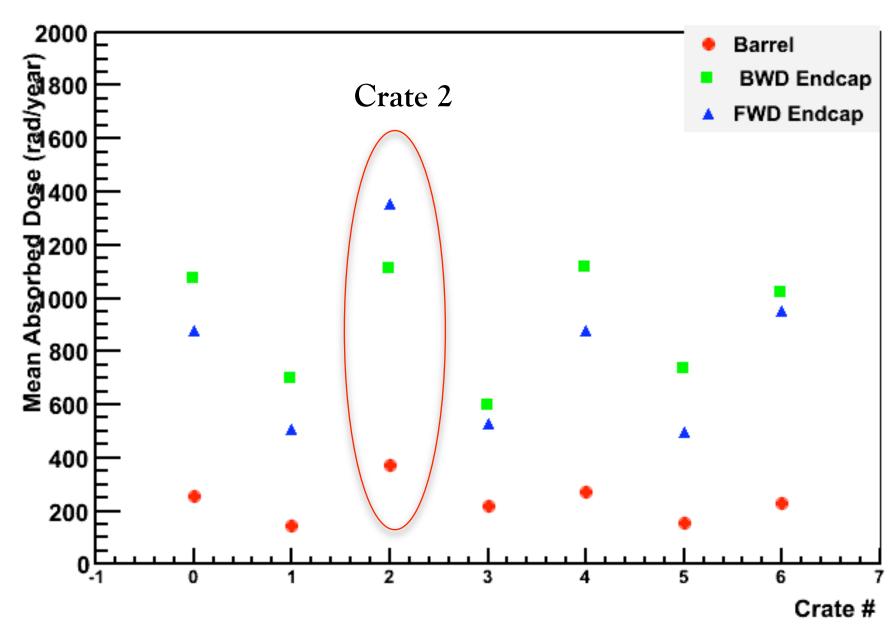
the Barrel respect to the one in

the FWD and BWD endcap



Absorbed Dose for each FEE Crates





Summary



- ✓ Radiative BhaBha, Touschek and Pair backgrounds have been studied in details.
- ✓ The effect of these backgrounds have been also studied on our FEEs
- ✓ IFR TDR background on writing



BACK-UP SLIDES

Electrons



Why do we have to worry about electrons

Electrons are charged particle that produce signals