2<sup>nd</sup> SuperB Collaboration Meeting MDI Parallel session Dec. 14<sup>th</sup> 2011

# **Machine Background Report**

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### **Outline**

#### New BRN developments:

- Fwd-EMC: New geometry
- FDIRC: Cherenkov photons activated and instrumentation
- New Final Focus model
- Nov. 2011 Full-Sim production
  - Pairs
  - Touschek (LER/HER)

#### Machine background on the SuperB detector

- SVT
- DCH
- FDIRC
- EMC
- IFR

#### Summary and Outlook

# **New BRN Developments**

#### **Fwd-EMC**

- Request from Stefano Germani to test different options for Fwd-EMC device
  - Nominal configuration uses LYSO (Geometry\_CIPE\_V00-00-02)
  - New geometries being tested:
    - CSI: Csi with VPT readout (Geometry\_CIPE\_V00-00-02\_CSI)
    - > BGO: Bgo with PMT readout (Geometry\_CIPE\_V00-00-02\_BGO)
    - > PWO: Pwo with PMT readout (Geometry\_CIPE\_V00-00-02\_PWO)
- Nov. 2011 production:
  - Geometry\_CIPE\_V00-00-02\_PWO: Rad-Bhabha (~10k events)

#### **FDIRC**

#### Previously:

- Stand Alone G4 simulation (Doug Roberts)
- BRN: FDIRC geometrical model, no instrumentation
- Currently:
  - A lot of work to insert stand alone model in BRN (Andrea Di Simone and Doug Roberts)
  - Cherenkov photons in the bars can be activates/deactivated with an option on Bruno invovation (-O). No significant increase on computingtime/output-size

#### Nov. 2011 production:

Cherenkov photons activated for all samples produced

# **New Final Focus Model**

### New FF model: Cryostat and Magnets (I)



All magnetic elements are made of the same material (QD0\_mixture):

- Density: 7.57 gr/cm<sup>3</sup>
- Composition: Niobium (0.106), Titanium (0.119), Cooper (0.347) and Iron (0.428)

### New FF model: Cryostat and Magnets (II)

#### **BRN** implementation



#### New FF model: Cryostat and Magnets (III)



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#### **New FF model: Magnetic model**

- Previously:
  - detector solenoidal field turned off in final focus magnetic model
- This field is important for an accurate model of two-photon (pairs) backgrounds on SVT. Less important for Rad-Bhaha and Touschek

#### Implementation:

- Magnitude: 1.5 Tesla
- Direction: Z>0 (0.0,0.0,1.0)
- Volume: field different from zero only inside a cylinder of length 40cm and radius 40cm.

# Nov. 2011 Full-Sim Production

### **Pairs background**

- Use diag36 (fastsim) generator to generate pairs (2-photon) primaries
- Kinematic cuts:

X [mm]

- Study the minimum Pt(CM) cut at generator level to not to bias the pairs sample and increase efficiency
- Study the losses at the beam pipes from Pairs to set-up the Pt(CM) cut



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    - $\Rightarrow \sigma(Pt(CM) > 0.55 \text{ MeV/c}) = 4.47 \text{mb} (\sigma(total) = 7.3 \text{mb})$
- Use guinea pig generator to inject pairs primaries in BRN
  - N-int-bunch = Lumi  $\times \sigma(Pt(CM) > 0.55 \text{ MeV/c})/f_ = 19.5$  interactions
  - Each events has <N-int-bunch> interactions (4-primaries each)
  - N primaries per event ~ 78 (500 rad-bhabha) ⇒ much faster than Radbhabha

### **Touschek background: strategy**

#### Primaries for BRN: STAR code (Manuela Boscolo)

- Simulate both Touschek and the beam gas scattering along the beam line
- Transport the scattered particles along the lattice
- Detect the collisions of these particles with the beam pipes (scoring planes)

#### Typical output:



## **Touschek background: samples (I)**



## **Touschek background: samples (I)**



### **Touschek background: samples (II)**



# Nov. 2011 production summary

#### Rad-Bhabha (fullsim):

- Jobs: 1099 (25 exited), ~10k events
- Size: 1.4 TB
- Pairs (fullsim):
  - Jobs: 350 (22 exited), ~100k events
  - Size: 265 GB
- Touschek HER/LER:
  - Jobs: 1425 (65), ~180 (80k) primaries for LER (HER)
  - Size: 1.1TB

#### Rad-Bhabha (bg-frames):

- Jobs: 7324 (146 exited), ~900k events
- Size: 39.4G

# Machine Background on the SuperB Detector

# **SVT backgrounds**

- 2-photon background dominates
- Touschek-LER seems to have significant impact on outer layers
  - Results from usual macros
  - L0: +20-30% 2photons (see next slide), reduced RadBhabha
  - Touschek became relevant for outer layers (+50%)

LAYERS	May2011 [MHz/cm2] 2phot. Pixels	May2011 [MHz/cm2] 2photons	Dec 2011 [MHz/cm2] 2photons	Dec 2011 [MHz/cm2] Rad Bhabha	Dec 2011 [MHz/cm2] Tousc-HER	Dec 2011 [MHz/cm2] Tousc-LER
L0 phi	55.5	23.3	32.2	0.96	0.52	1.73
L0 z		29.9	40.6	1.6	1.45	4.37
L1 phi	2.0	1.5	1.7	0.12	0.18	0.74
L1 z		0.7	0.85	0.083	0.19	0.77
L2 phi	0.96	0.72	0.88	0.086	0.12	0.56
L2 z		0.35	0.45	0.064	0.14	0.61
L3 phi	0.25	0.194	0.44	0.084	0.055	0.31
L3 z		0.097	0.27	0.056	0.055	0.29
L4 phi	0.014	0.012	0.05	0.014	0.004	0.019
L4 z		0.0076	0.03	0.008	0.003	0.013
L5 phi	0.007	0.006	0.019	0.006	0.002	0.009
L5 z		0.0041	0.014	0.004	0.0016	0.007

# **DCH backgrounds**

D. Lindemann

R. Cenci

#### DCH occupancies for Rad-bhabha and Pairs



- Rad-bhabha and 2-photon background produce similar occupancies, both 1-2%. Total occupancy ~2-4%
- Touschek (both LER/HER) isn't a concern for DCH (around 1-2 order of magnitude smaller)

### **FDIRC background**

A. Pérez



# **FDIRC background**



# **EMC backgrounds**

#### S. Germani





- Rad-bhabha samples for London (Oct. 2011) and LNF (Dec. 2011) productions
- LNF sample has a somewhat lower/higher energy spectrum for Fwd/Barrel w.r.t London
- Main differences are solenoidal field and cryostat model

Brl Measured Energy [log(MeV)], MDI parallel session Dec 14th 2011

## **EMC backgrounds**

S. Germani



- There is no absolute dominating background
  - Energy and angular (Barrel/Fwd) dependence
  - Main background is Rad-bhabha, but
    - Pairs dominate at low energies (1<MeV) in central barrel</p>
    - > Touschek-LER at high energy (>10 MeV) in bwd-barrel

# **IFR backgrounds**





#### **Rad-bhabha**:

- Neutrons: rate is very high and dangerous for the Sipm
- $\gamma/e^{\pm}$ : rates are high but shouldn't be a problem

#### Other background sources (Pairs and Touschek) are small compared with Radbhabha

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### **Summary and Outlook**

#### Several Changes in the BRN code since last SuperB meeting

- FDIRC: Cherenkov photons and intrumentation
- Final Focus: Solenoidal field and Cryostat and super-conducting magnets
- Several background sources produced: Rad-bhabha, Pairs, Touschek

#### Backgrounds on the detector:

- SVT: Main background is Pairs
- DCH: Rad-bhabha and Pairs produce similar amounts of backgrounds, 1-2% occupancy each.
- FDICR: Main background is Rad-bhabha. Will increase shielding arounf photon-camera to reduce it.
- EMC: No absolute dominant background. Main is Rad-bhabha, Pairs (Touschek-LER) dominates for low (high) energy regions
- IFR: Rad-bhabha is main background. Neutron rate a little bit high



## Rad-bhabha bg-frames production (I)

- Current final focus (FF) model in FullSim is very complete, it covers from -16m to 16m
  - Rad-bhabha simulation takes ~10min per event
  - Impossible to produce the rad-bhabha bg-frames request of 1M events in a reasonable time

#### Approach to the problem:

- The reason of the long FF model is to have a realistic estimation of neutron rates on the subsystems (FDIRC, IFR, EMC)
- FastSim doesn't have a good simulation for neutrons
- Propose to build reduced version of the FF: ±8mts and ±5mts
- Run a small fullsim production with the reduced versions of the FF
  - Compare background rates on different subsystems for the different FF models: nominal (±16mts) and reduced ones (±8mts and ±5mts)

#### If rates are similar $\Rightarrow$ can use the reduced FF for the bg-frame production

### **Rad-bhabha bg-frames production (II)**



## Rad-bhabha bg-frames production (III)

- Summary of comparison of FF models:
  - Most of the subsystems see very similar rates for the different FF models
  - Only the IFR sees different rates. Can we leave with this? FastSim IFR experts yes
  - See link below for the reports full reports on this

- The reduced FF model (±5mts) is the only approach that the FullSim group can offer to generate the requested 1M Rad-bhabha events in a reasonable time
  - $\Rightarrow$  The reduced FF model of ±5mts have a factor of 10 lower execution time per event w.r.t. the nominal FF model (±16mts)
- Nov. 2011 production:
  - Use the ±5mts FF model (Geometry\_CIPE\_V00-00-02\_ShorterFF5mts)