New ideas about DS at the intensity frontier

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Evidence for DM





100 Mpc



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Evidence for DM



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- Stable on cosmological timescales

Possible interactions?

SM-DM interactions



Optimist

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Pessimist

SM-DM interactions





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Mass range from freeze out



Standard WIMPS

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 $m_{\chi} \sim \text{TeV}$

Mass range from freeze out



Standard WIMPS

> Light **WIMPS**

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 $\sigma v \sim \alpha_{\rm EW}^2 / m_{\chi}^2$

 $m_{\gamma} \sim \text{TeV}$



Prime target for Intensity Experiment!

Experimental landscape



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Experimental landscape



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3 Experimental Initiatives 3.1 BDX 3.2 Belle II 3.3 CODEX-b 3.4 CODEX- β 3.5 DarkQuest 3.6 DarkMESA 3.7 FASER 3.8 FASER2 3.9 FASERv 3.10 FASERv2 3.11 FerMINI 3.12 FLArE 3.13 FNAL- μ 3.14 FORMOSA 3.15 HPS 3.16 JPOS 3.17 LDMX 3.18 LHCb 3.19 milliQan 3.20 NA64 3.21 NA64 $_{\mu}$ 3.22 PIONEER 3.23 PIP2-BD 3.24 POKER 3.25 REDTOP 3.26 SHADOWS 3.27 SND@LHC

2206.04220

And many more!

Great potential for repurposing/parasiting runs!

Beam dump & ν experiments



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Renormalizable portal models

 $\epsilon F'_{\mu
u}B^{\mu
u}$

кSH[†]H yLHN

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Renormalizable portal models

The dark photon $F_{\mu\nu} = \partial_{\mu}A'_{\nu} - \partial_{\nu}A'_{\mu}$

(target) p_t

кSH[†]H yLHN



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Renormalizable portal models

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кSH[†]H yLHN



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Dark photon is coupled to DM particle χ : $A'_{\mu}(\bar{\chi}\gamma^{\mu}\chi)$

> Typically mass relation fixed to $m'_A \sim 3m_{\gamma}$

A' typically produced on-shell, then decays into SM or χ

Dark photon portal results

Invisibly decaying Dark Photon



Mass shell of MET or decay products is peaked at m'_A

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Dark photon portal results



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Non-renormalizable portals

Non-renormalizable

 $g_{a\gamma}aF_{\mu\nu}\tilde{F}^{\mu\nu}$



 $\frac{c_J}{\Lambda_{\rm UV}^2} J_{\mu,\rm DS} J_{\rm SM}^{\mu}$

 $\frac{c_N}{\Lambda\Delta - 5/2} LH \mathcal{N}_{\rm DS}$

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Non-renormalizable portals

Non-renormalizable

 $g_{a\gamma} a F_{\mu\nu} \tilde{F}^{\mu\nu}$

 $\mathcal{O}_{\rm DS} |0\rangle \sim |{\rm DS}\rangle \sim \psi\psi...$

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D2

 c_N

 $\Delta - 5/2$

Darme+2001.01490 Cheng+2110.10691 Bertuzzo+2011.04735

 $\Lambda_{\rm UV}$ heavy mediator scale: cannot be produced on-shell!

> **DS** operator generically excites multiparticle state! (Controlled by Δ)



 ψ lightest DS particle. **Not necessarily DM!**

Non-renormalizable portals

Non-renormalizable

 $g_{a\gamma}aF_{\mu
u}\tilde{F}^{\mu
u}$

 $\mathcal{O}_{\rm DS} | 0 \rangle \sim | {\rm DS} \rangle \sim \psi \psi \dots$

 ψ lightest DS particle. **Not necessarily DM!**

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DS

 c_N

 $\Delta - 5/2$

Darme+2001.01490 Cheng+2110.10691 Bertuzzo+2011.04735

 $\Lambda_{\rm UV}$ heavy mediator scale: cannot be produced on-shell!

Can be **weakly coupled**: LDSP ψ directly excited by DS operator. Like SM ν in neutral current!

 $\frac{g_{\rm EW}^2}{M_Z^2}\bar{p}p\ \bar{\nu}\nu$

DS operator generically excites multiparticle state! (Controlled by Δ)



Can be **strongly coupled**: LDSP ψ comes from DS confinement

Cosmology depends on these details! We will not talk about it!

WIFAI2024

Common in BSM scenarios such as Hidden Valley, Neutral Naturalness, mirror worlds...





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Missing energy



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DS

Contino+2012.08537

$$J^{\mu}_{\rm DS}/\Lambda^2_{\rm UV}$$

 $p_{\rm DS}^2 \sim \sqrt{s}$

Missing energy



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DS

Scattering/disintegration



i) ψ stable: reaches detector and disintegrates against SM particles

> SM recoils is different from scattering against light, single particle ν DIS

Unique opportunity for ν detectors!

DUNE, Faser ν , SHiP,...

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Scattering/disintegration



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i) ψ stable: reaches detector and disintegrates against SM particles

> SM recoils is different from scattering against light, single particle ν DIS

Unique opportunity for ν detectors!

DUNE, Faser ν , SHiP,...

ii) ν disintegrates into ψ : **INCLUSIVE** and model independent!



Borrello MC Redigolo (WIP)

ν disintegration against protons



Need kinematic cuts to improve bkg rejection!



Displaced vertex



Depends on ψ lifetime $\langle 0 | \mathcal{O}_{\rm DS} | \psi \rangle$

$$\langle 0 | \mathcal{O}_{\rm DS} | \psi \rangle \sim \Lambda_{\rm IR}^{\Delta - 1} / 4\pi$$

$$S \sim \frac{N_{\text{POT}}}{\sigma_{pN}} \times \sigma_{\text{DS}} \times P_{\text{dec}} \times \epsilon_{\text{geo}}$$

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Conclusions

- SM-DS portal interactions well motivated possibility.
- Typical scales: $\Lambda_{IR} \sim MeV \div GeV$
- Non-renormalizable portals: not unusual in BSM scenarios. Can be seen as exciting multi-particle dark states
- Missing energy, disintegration signatures at future experiments have distinct kinematical signature controlled by Δ .
- Displaced decays, prompt more similar to standard case
- Probing up to $\Lambda_{\rm UV} \sim {\rm TeV}$.

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Thanks for the attention!

Backup

Cosmology



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Redi, Tesi 2107.14801 Hong, Kurup, Perelstein 1910.10160

Freeze-in is a possibility

(although for very high cutoffs in model indep. scenario)

$$f_L H f_R \to CFT$$

Specifics

Experiment	$N_{\rm POT}$ (total)	$E_{\rm beam}~({ m GeV})$	<i>l</i> (m)	<i>d</i> (m)	Off-axis angle, θ_{det} (rad)	$ heta_{ m acc} \ ({ m rad})$	
	10						
CHARM [78, 85, 86]	2.4×10^{18}	400	480	35	0.01	0.003	
NO ν A-ND [47, 84]	$3 imes 10^{20}$	120	990	14.3	0.015	0.002	
MicroBooNE (KDAR) [81]	$1.93 imes 10^{20}$	120	100	10.4	_	0.013	
ICARUS-NuMI [11, 16]	3×10^{21}	120	803	19.6	0.097	0.005	
DUNE-MPD[14, 87]	1.47×10^{22}	120	579	5	0	0.004	
SHiP [65, 77]	2×10^{20}	400	64	50	0	0.078	

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Production cross-section



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Dark Photon-DM constraints



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Dark photon produced on-shell

Dark photon decays into $\chi, \bar{\chi}$

 χ scatters against detectors

Scalar portal



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