

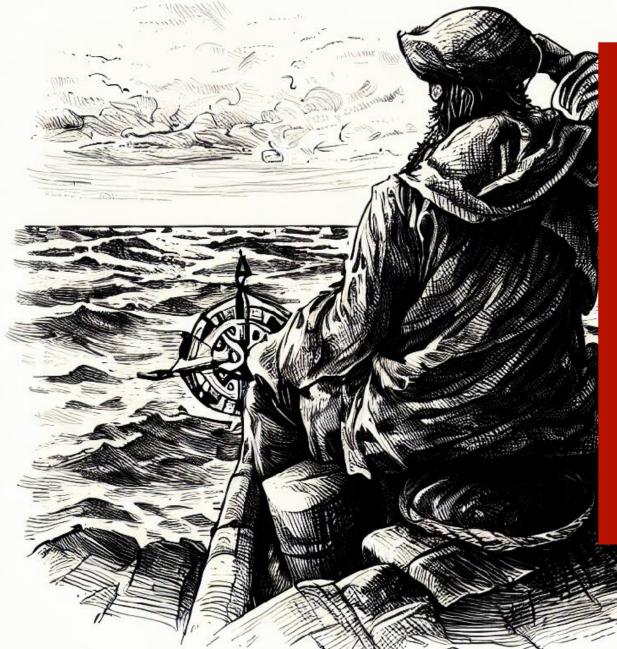
SEARCHING FOR A DARK SECTOR IN ATLAS AND CMS

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On behalf of the ATLAS and CMS Collaborations

ETH zürich

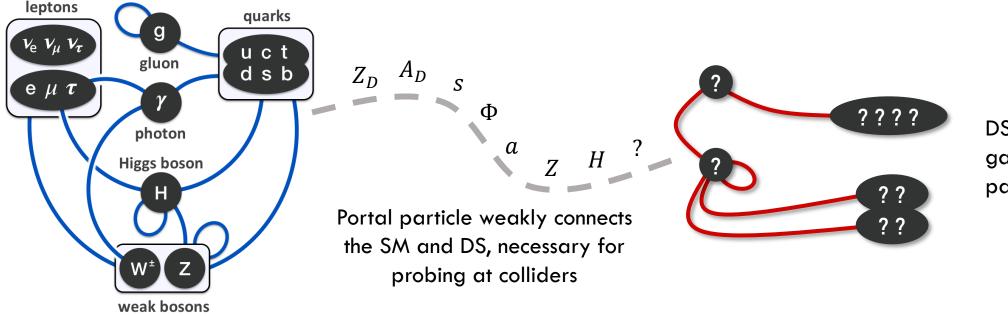
WIFAI 2024 Workshop Italiano sulla Fisica ad Alta intensità Bologna 12-15 Novembre 2024 Palazzo Hercolani, Aula Poeti Str. Maggiore, 45 - Bologna



WE KNOW IT'S OUT THERE, BUT NOT WHAT IT IS

A VERY BROAD CONCEPT

In a nutshell: any piece of Nature's lagrangian that is (mostly) decoupled from the Standard Model (SM) Motivated mainly by evidence for dark matter (DM), which we assumes makes up the dark sector (DS)



DS can have its own gauge structure and particle content

DISCLAIMER

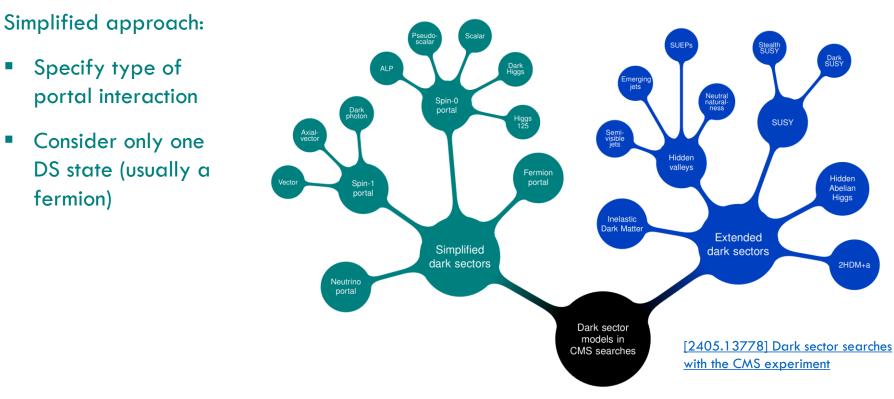
- Making justice to the dark matter programs of ATLAS and CMS in 20' is arguably impossible
- I will instead try to achieve two tings:
 - 1. Give you a picture of the landscape, leaving most details for you to discover
 - 2. Highlight some (subjectively picked) interesting new ideas and techniques in the field
- Helped greatly by the recently-released reviews by both collaborations [1, 2]



WE WILL NEED A COMPASS, OR TWO

WHERE DO YOU START?

How do you navigate a possible model space so big? Two complementary approaches



Extended approach:

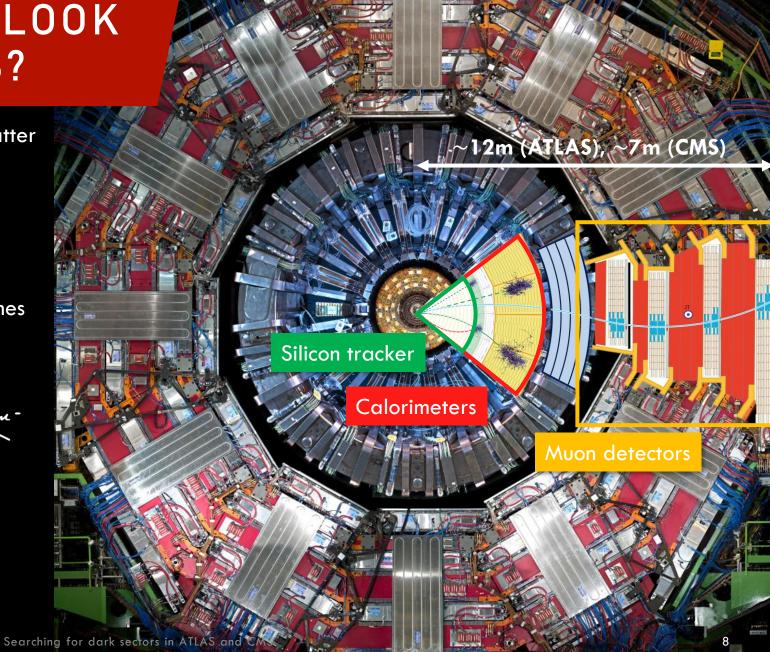
- More structured DS with rich dynamics
- Targeted searches with higher sensitivity, but less generality
- Exotic final states can evade searches targeting the "simplified" approach

AN EXPERIMENTALIST'S POINT OF VIEW

WHAT WOULD A DS LOOK LIKE IN ATLAS/CMS?

- By definition, DS states don't interact with matter
- Must detect decay products or particles produced in association

ne cT	Invisible	$ISR + p_T^{miss} \rightarrow (Mono)$ searches
Lifetime) _m +
		N M
	Displaced	Tracks not pointing to IP
	Prompt	Resonances

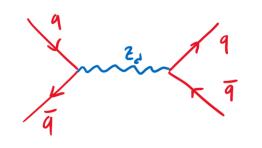


etecto

VISIBLE, INVISIBLE, OR...IN BETWEEN?

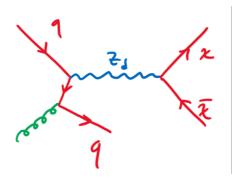
Consider as an example a simplified DS model with a vector portal Z_D and a dark fermion χ

• If $qq \rightarrow Z_D$, then $Z_D \rightarrow qq$



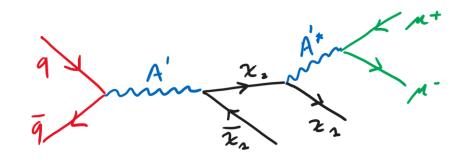
- Search for new dijet/dilepton resonances
- TLA/scouting* for low masses (more on that later)

• χ is detector-stable



- Exploit recoil against $\gamma/{
 m jet}/{
 m V}/{
 m H}$ and look for p_T^{miss}
- «Mono» or «DM+X» searches

What if the DS is more complicated?



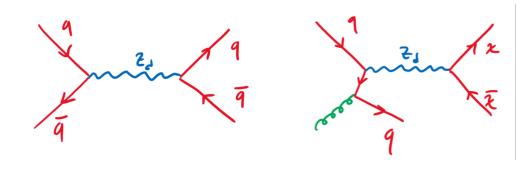
- Meta-stable states can result in displaced tracks
- Confining forces in the DS can lead to dark showers
- Such signatures would escape "traditional" searches



WE HAVE OUR MAP, LET'S DIVE IN

LET'S START GENERAL

 Little assumptions on the makeup of the DS → space for different interpretations



 Benchmark models common between ATLAS and CMS e.g., for a polar vector Z_d:

> $g_{\ell} = 0.01, g_q = 0.1, g_{\chi} = 1$ $g_{\ell} = 0, g_q = 0.25, g_{\chi} = 1$

Both ATLAS and CMS cover this space extensively

CMS

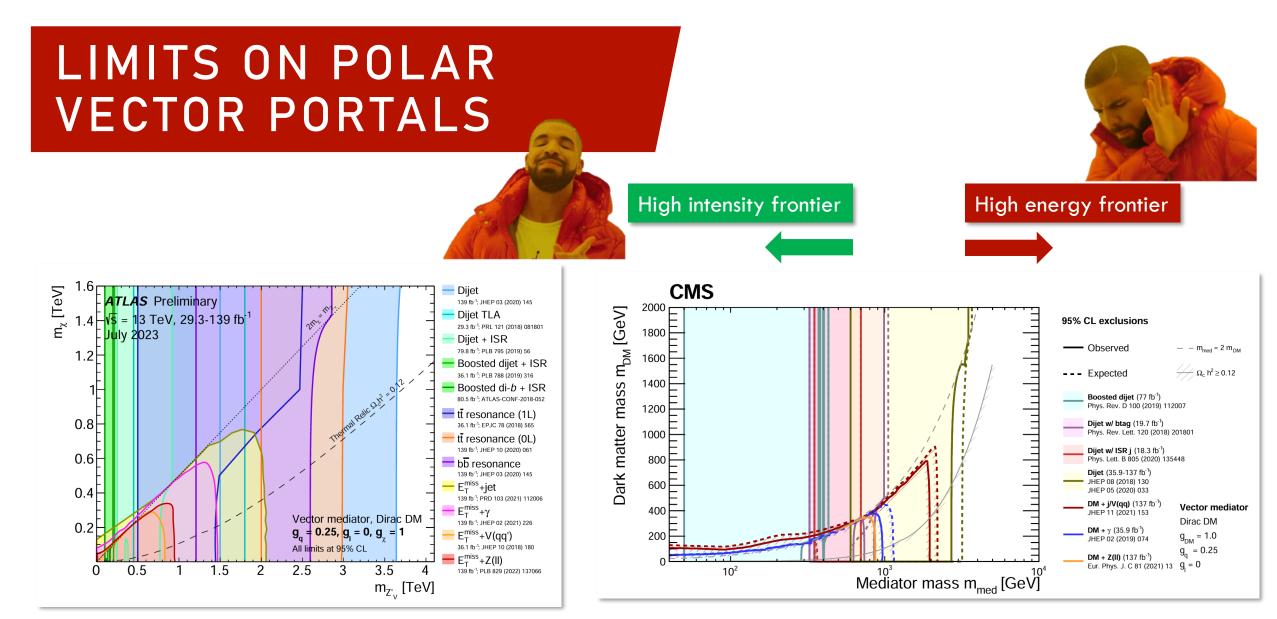
Boosted dijet search [3] B-tagged dijet search [4] Dijet+ISR with scouting* [5] High-mass dijet [6]

Mono-j/Mono-V(jj) [7] Mono-γ [8] Mono-V(ℓℓ) [9] Dilepton search [10]

ATLAS

(b-tagged) dijet search [10] Dijet search with TLA* [11] Dijet+ISR (resolved) [12] Dijet+ISR (boosted) [13] $t\bar{t}$ resonances [14, 15]

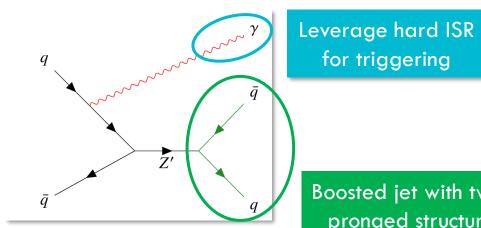
Mono-j [<u>16</u>] Mono-γ [<u>17</u>] Mono-V(jj/ℓℓ) [<u>18</u>, <u>19</u>] Dilepton search [<u>21</u>]



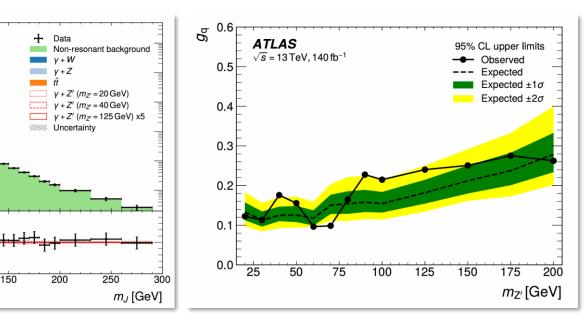
Other interpretations in the <u>backup</u>

- When the resonance mass becomes too low, the trigger system becomes a limit
- Overwhelming background rates (driven by dominant QCD cross section) → an online selection of potentially interesting events is required → trigger
- E.g., dijet analyses limited to resonance masses of ~1.5 TeV in CMS/ATLAS
- So how do we go lower?

- When the resonance mass becomes too low, the trigger system becomes a limit
- Overwhelming background rates (driven by dominant QCD cross section) \rightarrow an online selection of potentially interesting events is required \rightarrow trigger
- E.g., dijet analyses limited to resonance masses of ~ 1.5 TeV in CMS/ATLAS
- So how do we go lower?
- Let ISR help \rightarrow boosted topologies, e.g.:
- ATLAS DM+ISR photon search probes mediators down to 20 GeV



Boosted jet with twopronged structure



10.48550/arXiv.2408.00049

 $\sqrt{s} = 13 \,\text{TeV}$. 140 fb⁻¹

50

SR central tagged (post-fit)

ATLAS

Events / bin

 10^{7}

10⁶

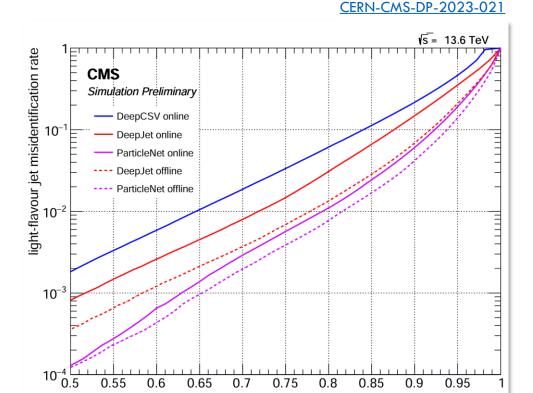
10⁵

10⁴

10³

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- E.g., dijet analyses limited to resonance masses of ~1.5 TeV in CMS/ATLAS
- So how do we go lower?
- Exploit jet substructure in the trigger
- Greatly suppresses QCD → higher rates

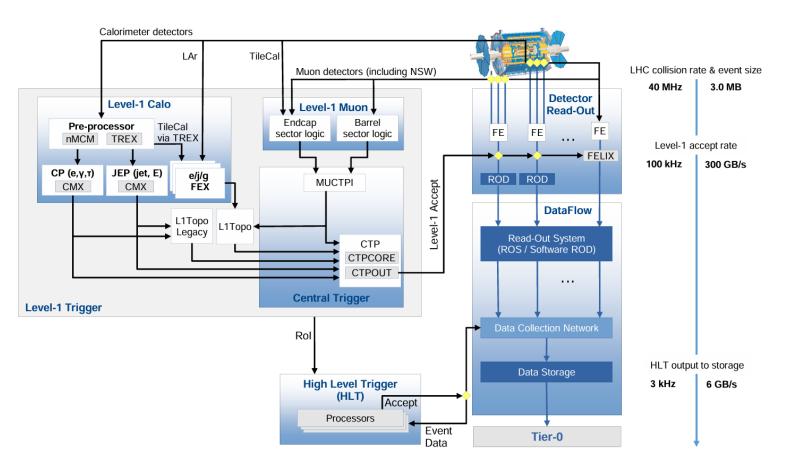


31.07.2024

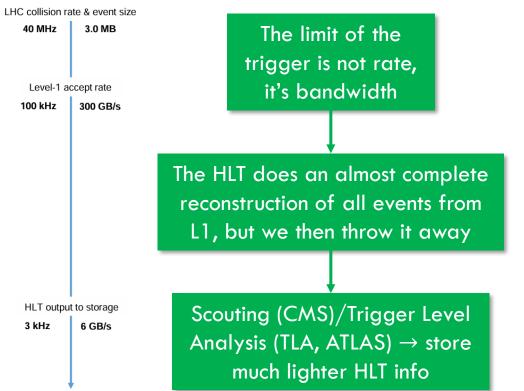
b jet identification efficiency

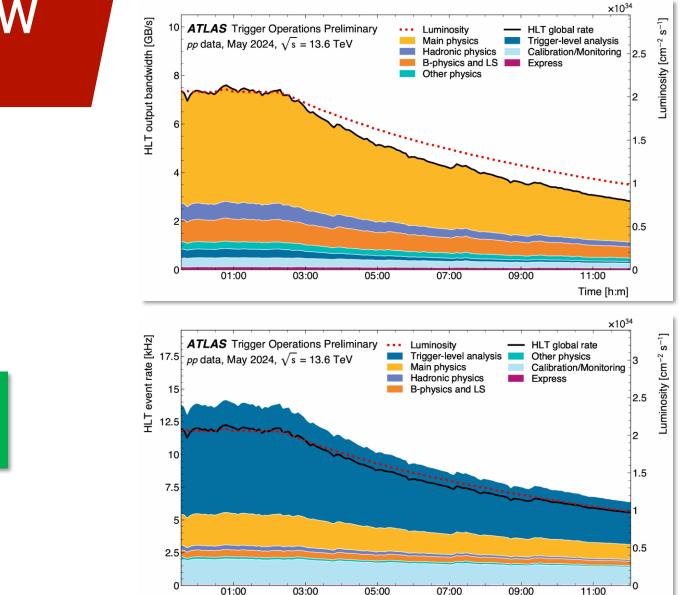
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- E.g., dijet analyses limited to resonance masses of ~1.5 TeV in CMS/ATLAS
- So how do we go lower?
- Look closer to what is actually limiting the trigger rate



 Look closer to what is actually limiting the trigger rate!





Time [h:m]

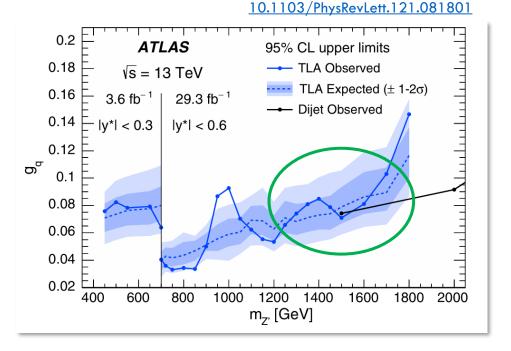
ATLAS trigger operations public results

LET'S SEE IT IN ACTION

To give but two examples of how this strategy can unlock a whole new level of potential sensitivity:

ATLAS TLA dijet search

- Record all events containing a L1 jet with $E_T > 100$ GeV (29.3 fb⁻¹)
- Subset of data with eve looser $E_T > 75$ GeV requirement (3.6 fb⁻¹)
- Store jet 4-momentum (+a few quality variables) \rightarrow 0.5% of size w.r.t. full event info
- Fully efficient for offline $p_T > 220$ GeV

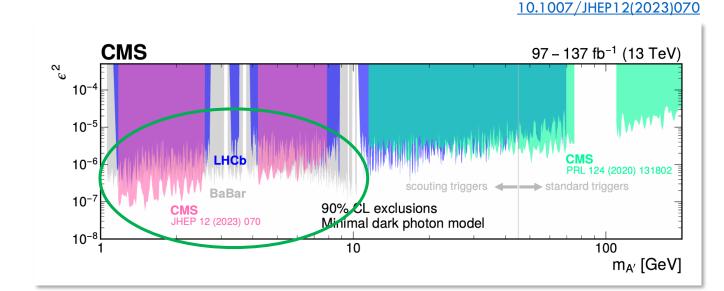


LET'S SEE IT IN ACTION

To give but two examples of how this strategy can unlock a whole new level of potential sensitivity:

CMS scouting dimuon search

- Record all L1 events with two muons having $p_T > 3 \text{ GeV}$
- Active in 2017 and 2018 for 97 fb^{-1} of lumi
- Store hits in tracker/muon system and calorimeter clusters
- Event size: ~ 1 MB (offline) $\rightarrow \sim 8$ kB (scouting)
- Event rate: 450 Hz (offline) \rightarrow 2 kHz (scouting)



Enabling some of the most stringent limits to date!

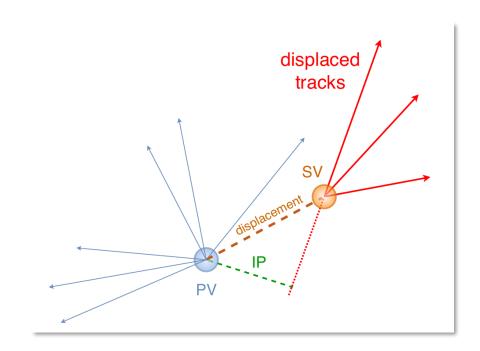
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COVERING OUR BLIND SPOTS

DISPLACED SIGNATURES

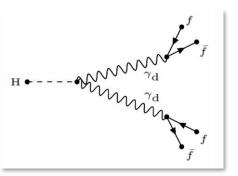
- The general approach to DS searches I discussed up to now is powerful, but it has failure modes
- Common reconstruction and analysis pipelines in ATLAS and CMS designed for prompt objects, but
- If DS states are meta-stable → displaced objects
- Require rethinking most of the criteria in reconstructing objects
- Powerful tool to extend reach of DS searches, but also new and challenging backgrounds to fight
- Will only give a few examples in the interest of time, much more in [1, 2]

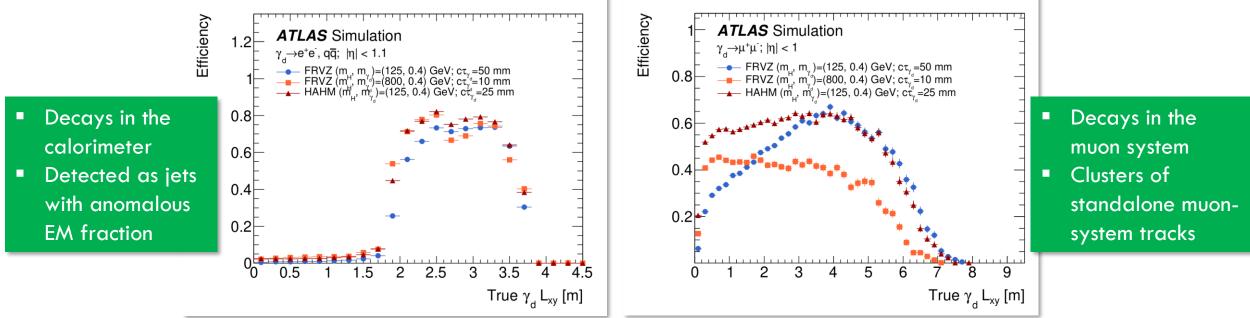


DARK-PHOTON JETS IN ATLAS

- Beautiful ATLAS searches considering different potential long-lived signatures
- Very different signature depending on dark photon lifetime

10.1007/JHEP06(2023)153

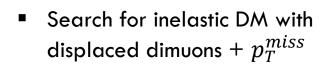


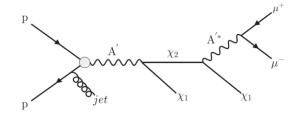


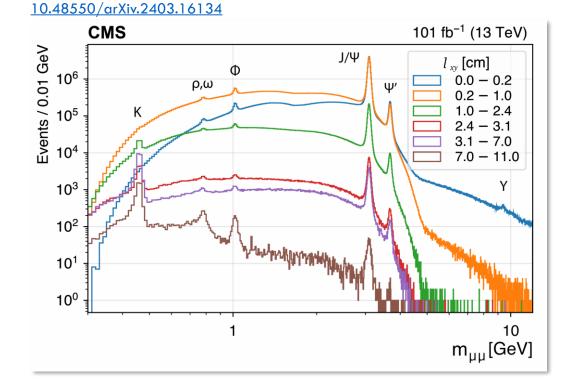
<u>10.1007/JHEP06(2023)153</u>

DISPLACED DIMUONS IN CMS

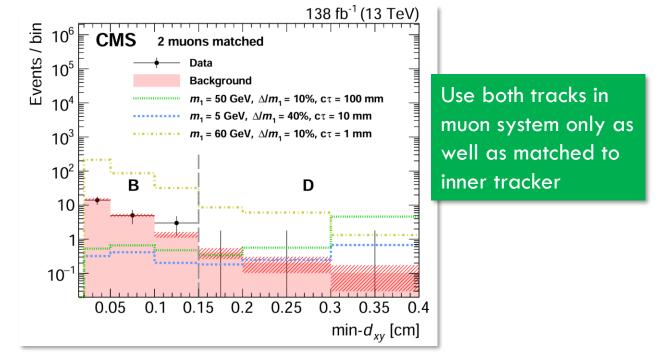
 Leveraging scouting data, displaced dimuons down to invariant masses of O(1) GeV







<u>10.1103/PhysRevLett.132.041802</u>

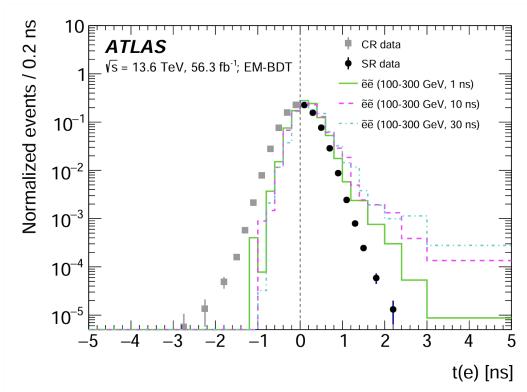


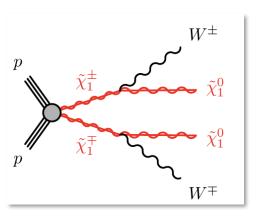
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DISPLACED LEPTONS IN ATLAS

- New for Run 3: exploit large radius tracking (LRT) directly in the trigger
- Exploit the capabilities of the ATLAS LAr EM calorimeter
- Two BDTs are built by using timing and pointing info from the LAr EM calo
- Discriminate displaced electrons reconstructed as either electrons or photons from prompt objects

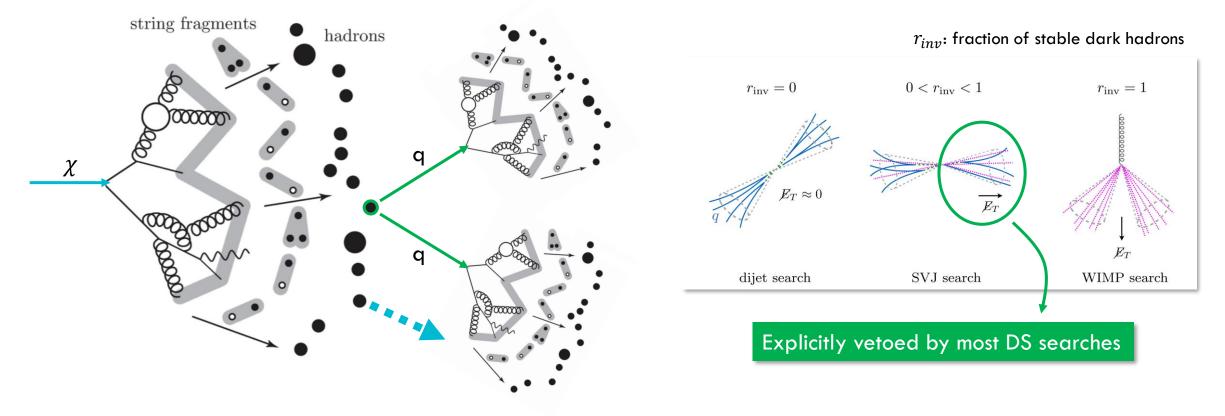
<u>10.48550/arXiv.2410.16835</u>





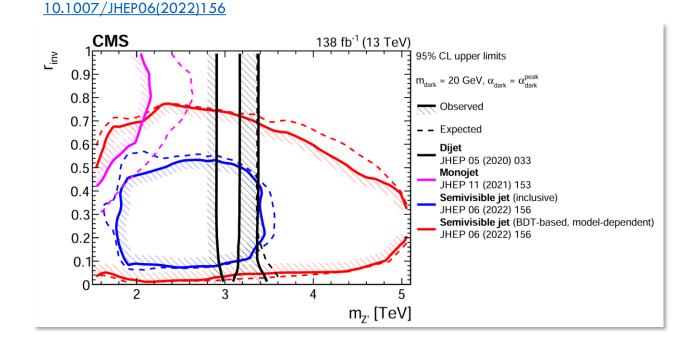
DARK QCD - SVJS

- If the DS is subject to a confining force, interesting signatures can arise
- Dark quarks shower and hadronized in DS \rightarrow "dark" iet \rightarrow stable dark hadrons decay back to SM \rightarrow semivisible jets

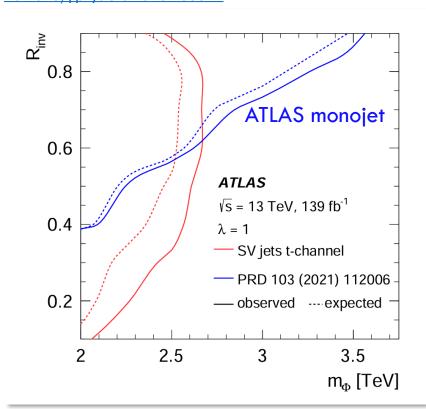


DARK QCD - SVJS

First search by CMS for vector mediator (s-channel), first search by ATLAS for bi-fundamental mediator (t-channel)

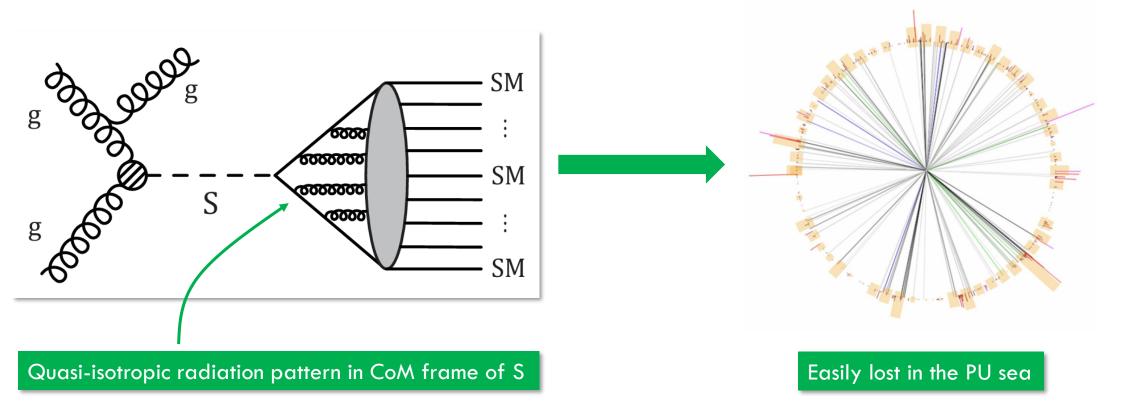


10.1016/j.physletb.2023.138324



DARK QCD - SUEPS

- Who says dark QCD should behave like SM QCD?
- In the large t' Hooft coupling regime you don't get jets anymore

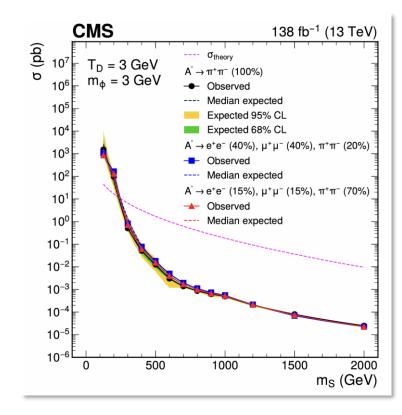


DARK QCD - SUEPS

First search for such signatures released by CMS

10¹⁰ CMS 138 fb⁻¹ (13 TeV) Events/bin H VR SR $S_{\text{boosted}}^{\text{SUEP}} \in (0.5, 1.0]$ S^{SUEP} ∈ (0.4, 0.5] S^{SUEP} ∈ (0.3, 0.4] É ∐G B Ć 1D F ĽΑ 10⁹ 10⁸ Observed $T_D = m_{\phi} = 3 \text{ GeV}$ 10^{-1} Post-Fit $A \rightarrow \pi^+ \pi^-$ (100%) 10⁶ (b. only) ---- m_s = 300 GeV 10⁵ Pre-Fit --- m_s = 800 GeV 10⁴ 10³ 10² 10¹ 10⁰ Pred./Obs. 100 150 200 50 100 200 50 50 150 100 150 200 n^{SUEP} constituent

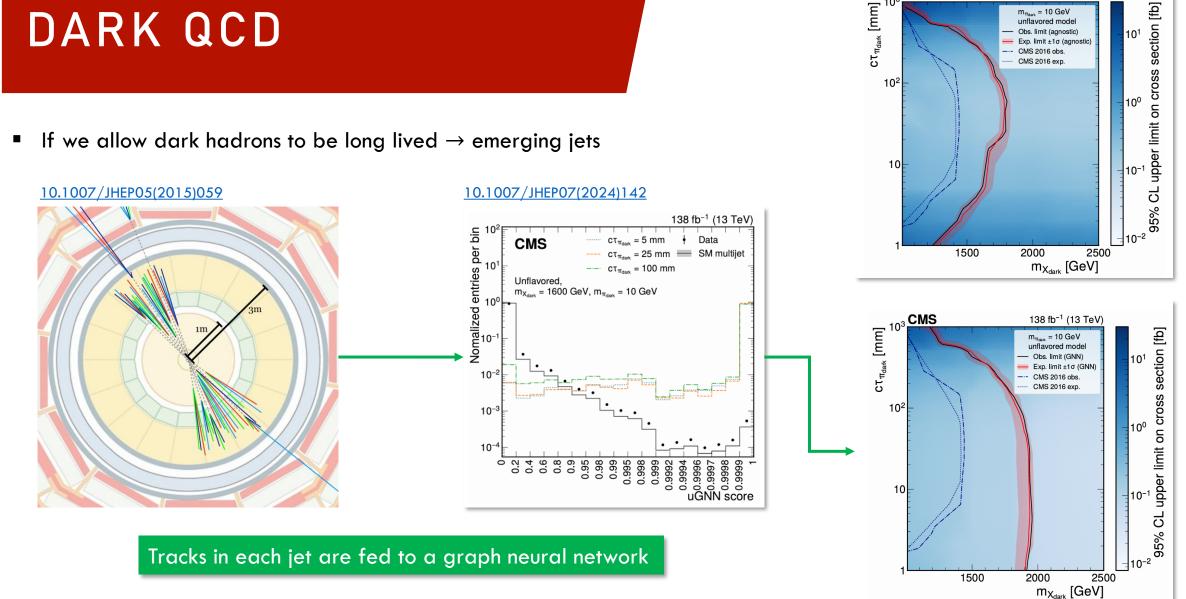
10.1103/PhysRevLett.133.191902



 Fully data-driven background estimation via extended ABCD, fit on n_{constituents}

DARK QCD

If we allow dark hadrons to be long lived \rightarrow emerging jets



10²

ст_{тdark}

138 fb⁻¹ (13 TeV)

 $m_{\pi_{dark}} = 10 \text{ GeV}$

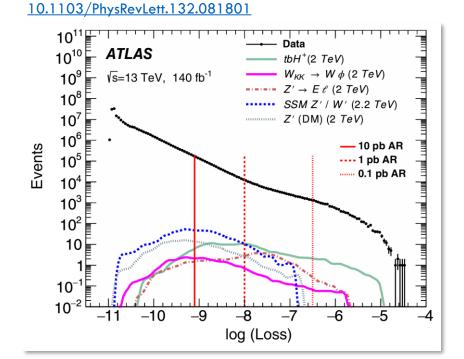
unflavored model Obs. limit (agnostic)

Exp. limit ±1 (agnostic) CMS 2016 obs. CMS 2016 exp.

10⁰

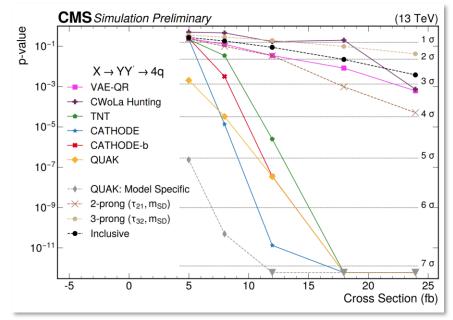
ANOMALY DETECTION

• Train a model to learn what the background looks like \rightarrow tag potentially anomalous data



Autoencoder trained on data represented as a <u>rapidity-mass matrix</u>

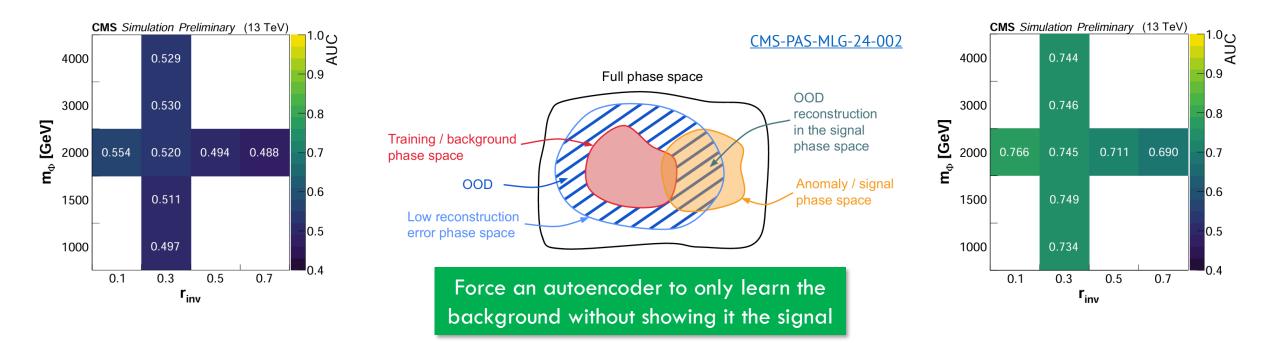




Various techniques, including variational autoencoders and normalizing-flow-based setups

ANOMALY DETECTION

- Train a model to learn what the background looks like \rightarrow tag potentially anomalous data
- Can also be powerful to reduce dependency of more targeted searches
- E.g.: Wasserstein Normalized Autoencoders for semivisble jet tagging





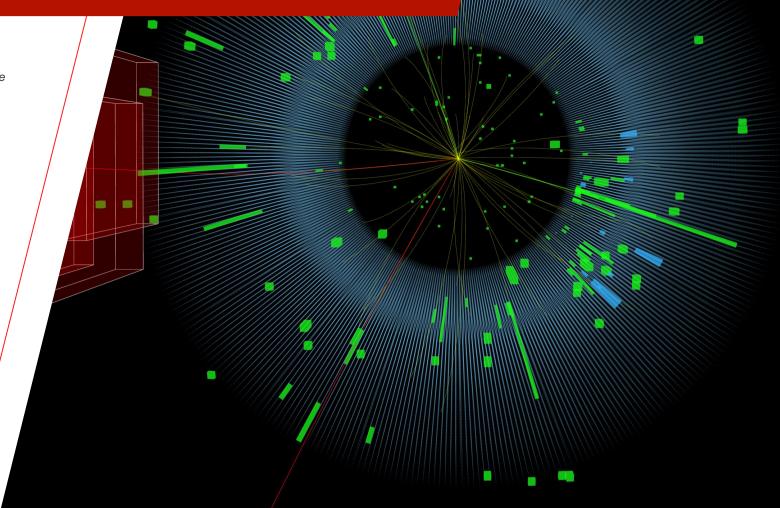
SUMMARY

- I hope I gave you an idea of the landscape of searches for dark matter in ATLAS and CMS
- From pushing "traditional" methods to ultimate sensitivity, to new ideas on how to tackle limits and challenges
- We have not found dark matter yet...but we are looking closer than ever

THANK YOU

« Ce qui est admirable, ce n'est pas que le champ des étoiles soit si vaste, c'est que l'homme l'ait mesuré. »

Jacques Anatole François Thibault





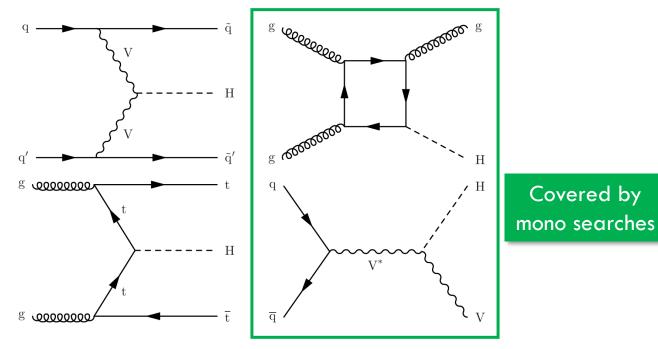
REFERENCES

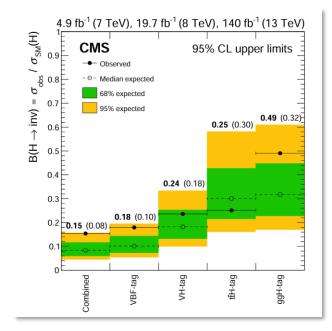
- [1]: <u>arXiv:2403.09292</u>
- [2]: <u>arXiv:2405.13778</u>
- [3]: <u>10.1103/PhysRevD.100.112007</u>
- [4]: <u>10.1103/PhysRevD.108.012009</u>
- [5]: <u>10.1016/j.physletb.2020.135448</u>
- [6]: <u>10.1007/JHEP05(2020)033</u>
- [7]: <u>10.1007/JHEP11(2021)153</u>
- [8]: <u>10.1007/JHEP02(2019)074</u>
- [9]: <u>10.1140/epic/s10052-020-08739-5</u>
- [10]: <u>10.1007/JHEP07(2021)208</u>
- [11]: <u>10.1007/JHEP03(2020)145</u>
- [12]: <u>10.1103/PhysRevLett.121.081801</u>
- [13]: <u>10.1016/j.physletb.2019.03.067</u>
- [14]: <u>10.1016/j.physletb.2018.09.062</u>
- [15]: <u>10.1140/epic/s10052-018-5995-6</u>

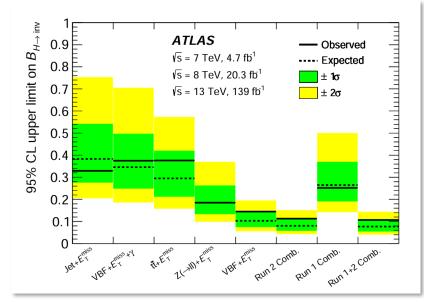
- [16]: <u>10.1007/JHEP10(2020)061</u>
- [17]: <u>10.1103/PhysRevD.103.112006</u>
- [18]: <u>10.1007/JHEP02(2021)226</u>
- [19]: <u>10.1007/JHEP10(2018)180</u>
- [20]: <u>10.1016/j.physletb.2022.137066</u>
- [21]: <u>10.1016/j.physletb.2019.07.016</u>

THE HIGGS BOSON AS A DARK SECTOR PROBE

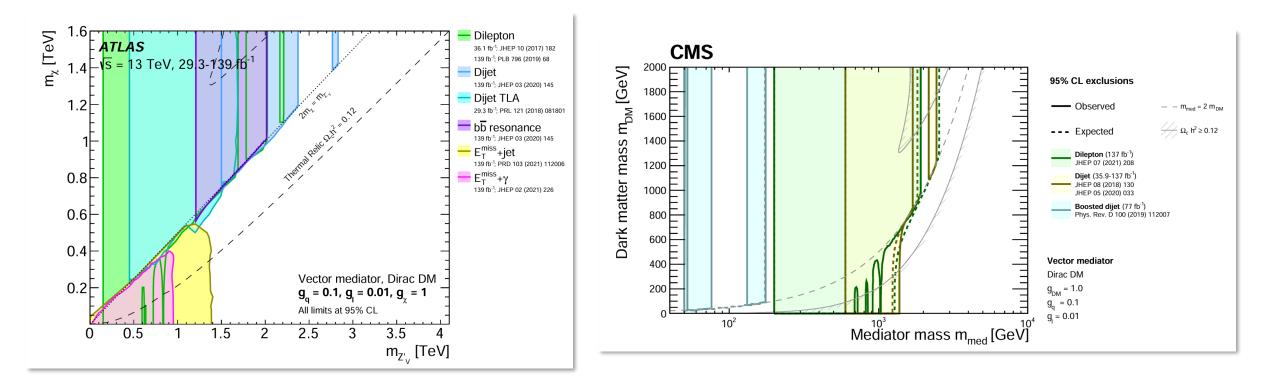
- We discovered the Higgs boson, why not start using it as a probe?
- In the SM, the Higgs boson can decay invisibly via $H \rightarrow ZZ \rightarrow 4\nu$
- Once $H \rightarrow ZZ$ and $Z \rightarrow \nu\nu$ branching fractions are known, searches for generic $H \rightarrow invisible$ events are powerful probes for DSs





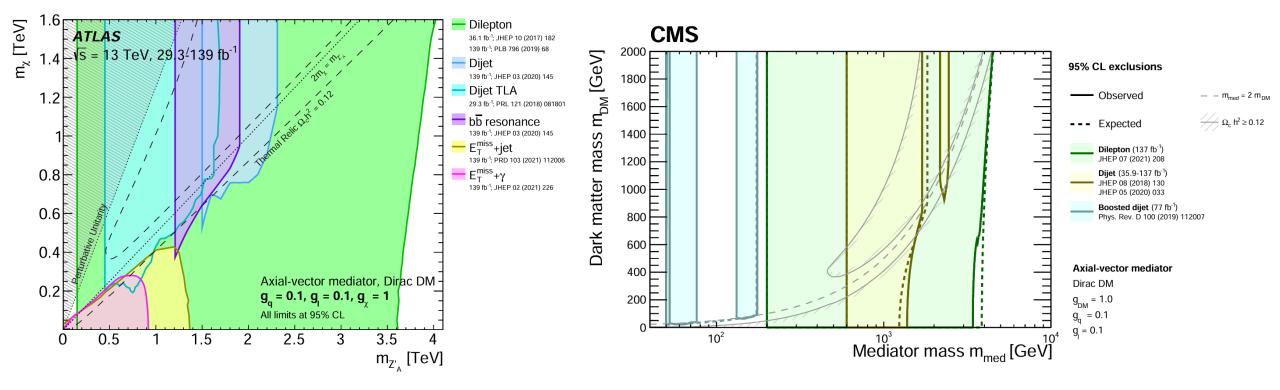


VECTOR PORTAL LIMITS



- Limits mostly dominated by dilepton searches if portal is allowed lepton couplings
- Lower masses can be reached by ISR+dijet searches or TLA/scouting (more on those later)

AXIAL VECTOR PORTAL LIMITS



LEPTOPHOBIC AXIAL VECTOR PORTAL LIMITS

