

SAŠO GROZDANOV



THERMAL SPECTRA IN QFTS AND THEIR RECONSTRUCTIONS

OUTLINE

hydrodynamics, kinetic theory and thermal spectra

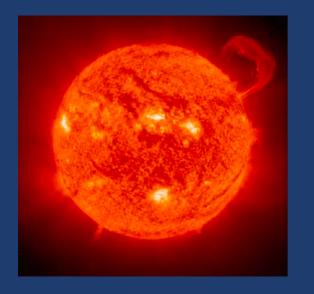
holography and reconstruction of spectra

pole-skipping and the reconstruction

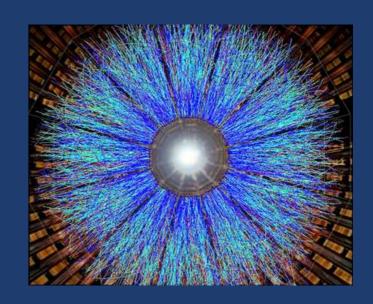
summary and future directions

HYDRODYNAMICS, KINETIC THEORY AND THERMAL SPECTRA

THERMAL FIELD THEORY







$$Z[\beta = 1/T] = \int \mathcal{D}\Phi \, e^{-\beta H} e^{\frac{i}{\hbar} \int d^d x \, \mathcal{L}(\Phi, \lambda)}$$

expansions of observables:

$$\mathcal{O} = \mathcal{O}_{\lambda,0} + \lambda \, \mathcal{O}_{\lambda,1} + \dots$$

$$\mathcal{O} = \mathcal{O}_{p,0} + p \, \mathcal{O}_{p,1} + \dots$$

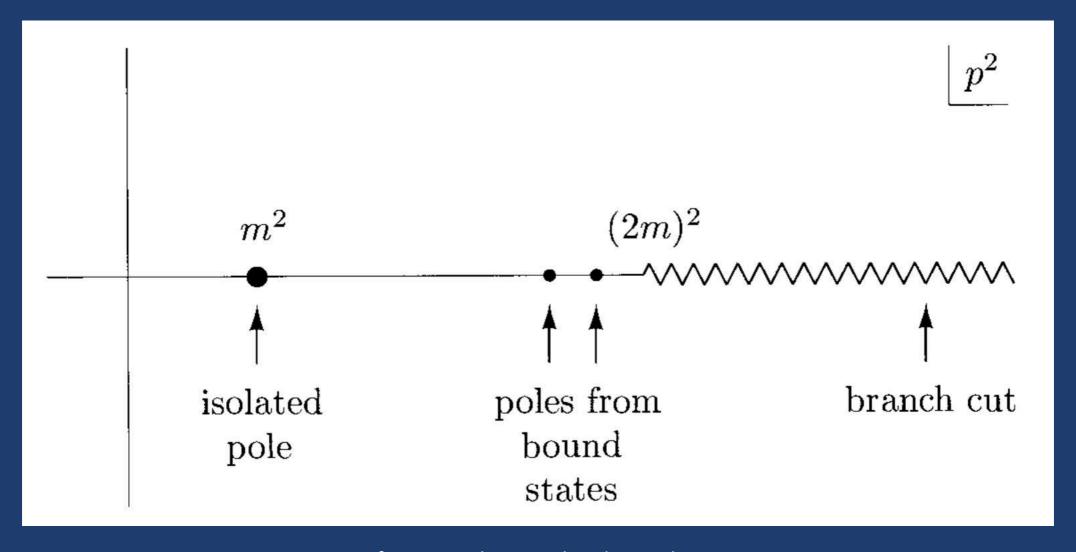
$$\mathcal{O} = \mathcal{O}_{N,0} + \frac{1}{N} \, \mathcal{O}_{N,1} + \dots$$

$$\mathcal{O} = \mathcal{O}_{d,0} + \frac{1}{d} \, \mathcal{O}_{d,1} + \dots$$

 $|\mathcal{O}| = \mathcal{O}_{T,0} + T \overline{\mathcal{O}_{T,1} + \dots}$

SPECTRUM OF A SIMPLE T=0 CORRELATOR

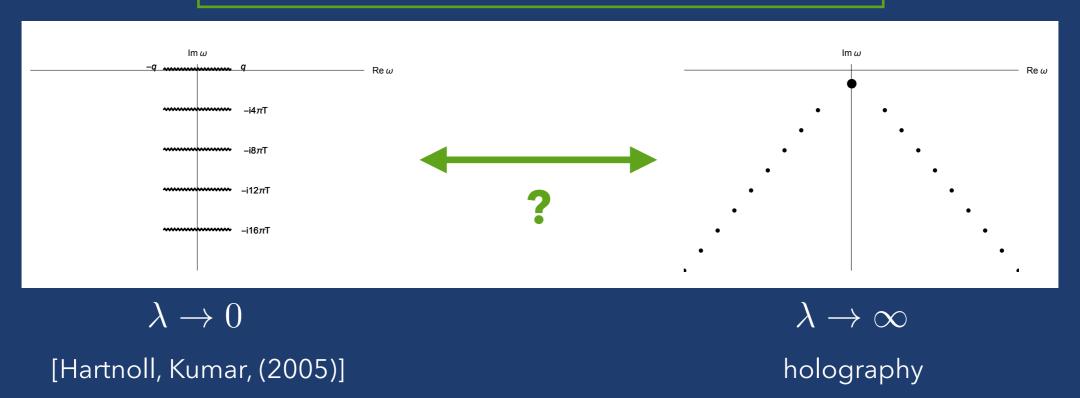
$$\langle \phi(p)\phi(-p)\rangle = \frac{Z(p^2)}{p^2 - m^2 + \Sigma(p^2)}$$



[from Peskin and Schroeder]

ANALYTIC STRUCTURE OF THERMAL CORRELATORS

$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)}$$



what is the structure of thermal correlators and what is the minimal information necessary to determine them completely?

- kinetic theory: solving the collision integral
- perturbative QFT: calculation of Feynman diagrams
- holography: solving differential equations

HYDRODYNAMIC PART OF THE THERMAL SPECTRUM

- low-energy limit of QFTs a Schwinger-Keldysh effective field theory [SG, Polonyi (2013); Crossley, Glorioso, Liu (2015); Haehl, Loganayagam, Rangamani (2015); ...]
- conservation laws (equations of motion) of globally conserved operators

$$\nabla_{\mu}T^{\mu\nu} = 0 \qquad \nabla_{\mu}J^{\mu} = 0$$

$$\dots \nabla_{\mu} J^{\mu\nu} = 0$$

higher-form currents in MHD [SG, Hofman, Iqbal, PRD (2017)]

tensor structures (symmetries, gradient expansions) and transport coefficients (QFT)

$$T^{\mu\nu} = \sum_{n=0}^{\infty} \left[\sum_{i=1}^{N} \lambda_{i}^{(n)} \mathcal{T}_{(n)}^{\mu\nu} \right] \qquad \frac{\nabla_{\mu} T^{\mu\nu} = 0}{u^{\mu} \sim T \sim e^{-i\omega t + iqz}}$$

$$\nabla_{\mu} T^{\mu\nu} = 0$$

$$u^{\mu} \sim T \sim e^{-i\omega t + iqz}$$

$$\omega(q) = \sum_{n=1}^{\infty} \alpha_n q^n$$

$$\partial u^{\mu} \sim \partial T \ll 1$$

$$\omega/T \sim q/T \ll 1$$

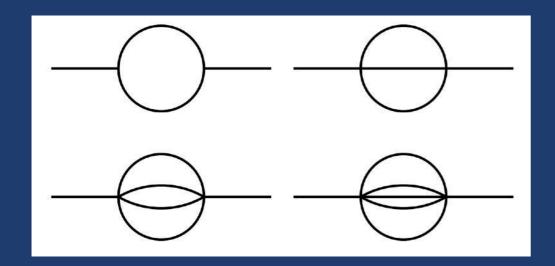
dispersion relations:

shear diffusion sound $\omega = -iDq^2 \qquad \omega = \pm v_s q - i\Gamma q^2$

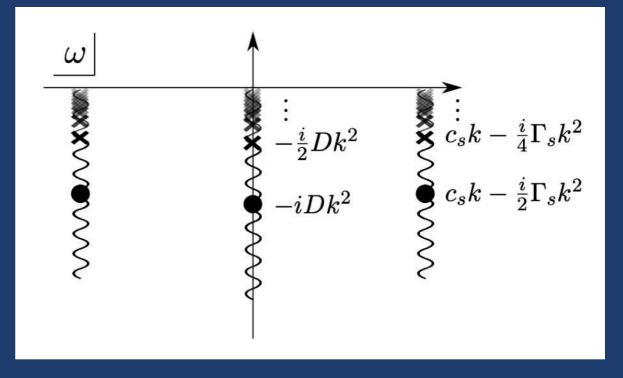
equilibrium temperature
$$q=\sqrt{{f q}^2}$$

HYDRODYNAMIC PART OF THE THERMAL SPECTRUM

- Schwinger-Keldysh effective field theory of diffusion to all loops: long time tails... [Chen-Lin, Delacretaz, Hartnoll (2019); Delacretaz (2020); SG, Lemut, Soloviev, to appear]
- tree-level result (classical hydrodynamics) has one diffusive pole at $\omega = -iDq^2$
- one-loop result has one (renormalised) diffusive pole at $\omega=-i(D+\delta D)q^2$ and one branch point at $\omega=-iDq^2/2$
- *n*-loop result has a branch point at $\omega = -iDq^2/(n+1)$



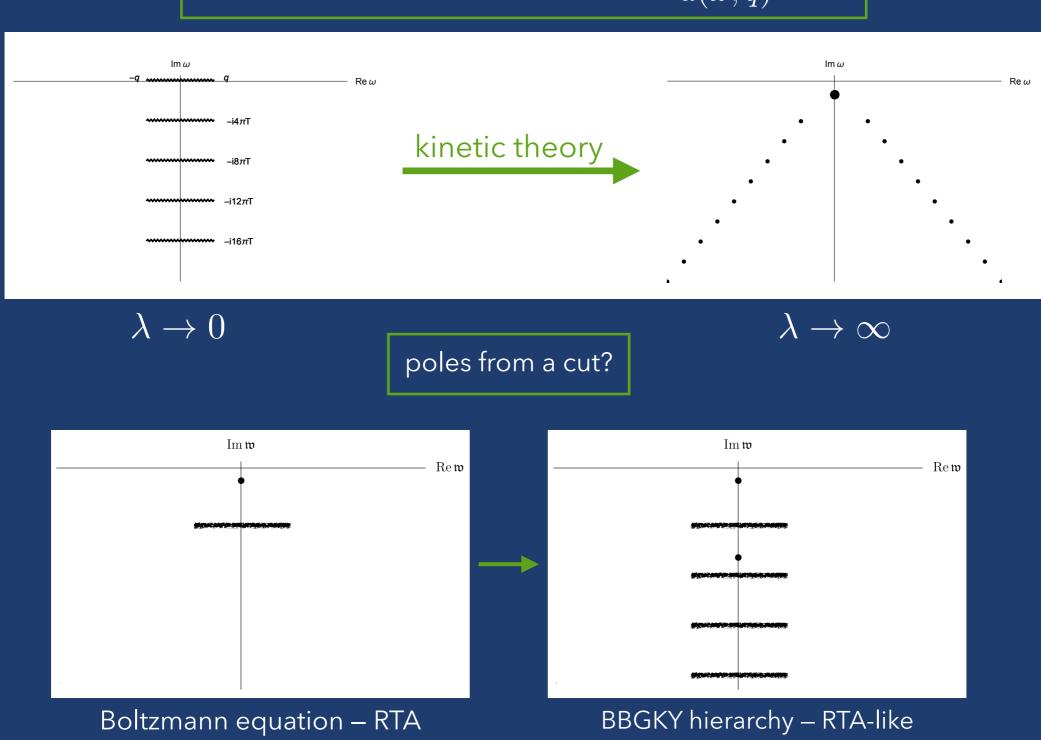
dominant one-, two-, three- and fourloop diagrams responsible for the cut



[plot from Delacretaz (2020)]

KINETIC THEORY AND THERMAL SPECTRUM

$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)}$$



truncations [SG, Soloviev, to appear]

[Romatschke, (2016)]

KINETIC THEORY AND THERMAL SPECTRUM

• two-point functions of $T^{\mu\nu}$ and J^{μ} from RTA kinetic theory at finite T, μ, Γ [Bajec, SG, Soloviev, to appear]

$$[p^{\mu}\partial_{\mu} + F^{\mu}\nabla_{p^{\mu}}] = C[f] = \frac{p^{\mu}u_{\mu}}{\tau_{R}}(f - f^{\text{eq}})$$

$$\nabla_{\mu} T^{\mu 0} = 0, \ \nabla_{\mu} T^{\mu i} = -\Gamma T^{0i}$$
$$\nabla_{\mu} J^{\mu} = 0$$

- analytic expression for all correlators that exhibit hydrodynamics (diffusion and sound), quasihydrodynamics, a branch cut running between $\omega=-i/\tau_R\pm q$, thermoelectric effect
- a 'periodic system' of correlators:

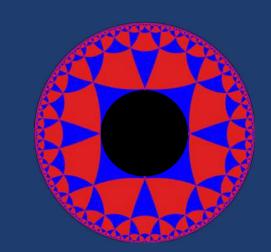
	$\langle JJ \rangle$		$\langle T$	$T\rangle$	$\langle TJ \rangle$	
	even	odd	even	odd	even	odd
$T \neq 0$	Z W		• • • • • • • • • • • • • • • • • • • •	*****	_	_
$T,\Gamma_{\parallel} \neq 0$	ZWY S	· · · · · · · · · · · · · · · · · · ·	******	******	_	-
$T,\Gamma_{\perp} eq 0$	Zwis S	·····	• • •	······	_	_
$T, \mu \neq 0$	ZWY.	······	• • • • • • • • • • • • • • • • • • • •	······	• • • • • • • • • • • • • • • • • • • •	******
$T,\mu,\Gamma_\parallel eq 0$	Z WY	·····	******	·····	******	***************************************
$T,\mu,\Gamma_{\perp} eq 0$	Z	·····	• • •	·····	*****	www.

	$\langle JJ \rangle$		$\langle TT \rangle$			$\langle TJ \rangle$	
	spin 0	spin 1	spin 0	spin 1	spin 2	spin 0	spin 1
$T \neq 0$		****	• • •		******	_	_
$T,\Gamma_{\parallel}\neq 0$		~~~	******	******	·····	_	_
$T, \Gamma_{\perp} \neq 0$		*****	• • •		*****	_	_
$T, \mu \neq 0$	*****	****	• • •	· · · · · · · · · · · · · · · · · · ·	*****	• • •	*****
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$T,\mu,\Gamma_{\perp} \neq 0$	******	*******	•	******	*******	• • • • • • • • • • • • • • • • • • • •	

HOLOGRAPHIC SPECTRA

HOLOGRAPHY AND HYDRODYNAMICS

- duality: theory A = theory B
- a result of string theory (quantum gravity) [Maldacena (1997)]



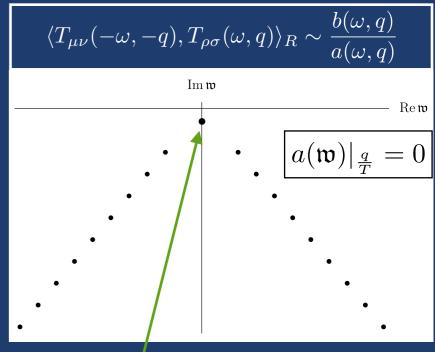
strongly coupled quantum theory

(extremely hard)

weakly coupled gravity

(much easier)

- perturbations of black holes (quasinormal modes) give spectra of QFT operators for $\mathfrak{w}\equiv rac{\omega}{2\pi T}\in\mathbb{C}$
- large-N QFT calculations -> ODEs
- invaluable explicit (toy) models: the $\mathcal{N}=4$ supersymmetric Yang-Mills theory [SG, Kovtun, Starinets, Tadić, JHEP (2019)]



sound:

shear diffusion:

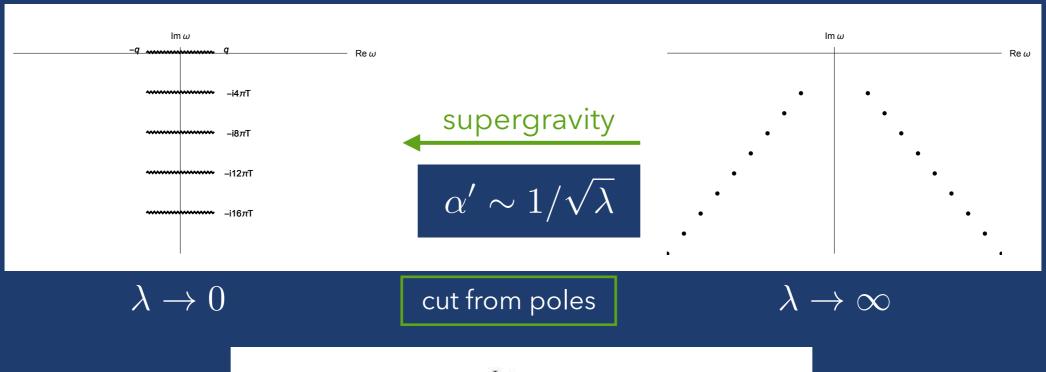
$$\omega = \pm \frac{1}{\sqrt{3}} q - \frac{i}{6\pi T} q^2 \pm \frac{3 - 2\ln 2}{24\sqrt{3}\pi^2 T^2} q^3 - \frac{i(\pi^2 - 24 + 24\ln 2 - 12\ln^2 2)}{864\pi^3 T^3} q^4 \pm \cdots$$

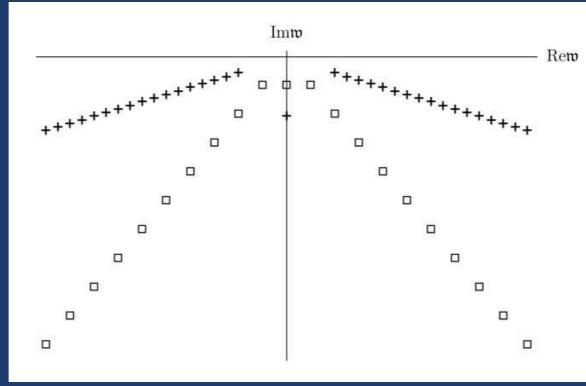
$$\omega = -\frac{i}{4\pi T} q^2 - \frac{i(1 - \ln 2)}{32\pi^3 T^3} q^4 - \frac{i(24\ln^2 2 - \pi^2)}{96(2\pi T)^5} q^6$$

$$-\frac{i\left[2\pi^2(\ln 32 - 1) - 21\zeta(3) - 24\ln 2(1 + \ln 2(\ln 32 - 3))\right]}{384(2\pi T)^7} q^8 + \cdots$$

HOLOGRAPHY AND THERMAL SPECTRUM

$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)} = \frac{B(\omega, q)}{\prod_{i=0}^{\infty} (\omega - \omega_i(q))}$$

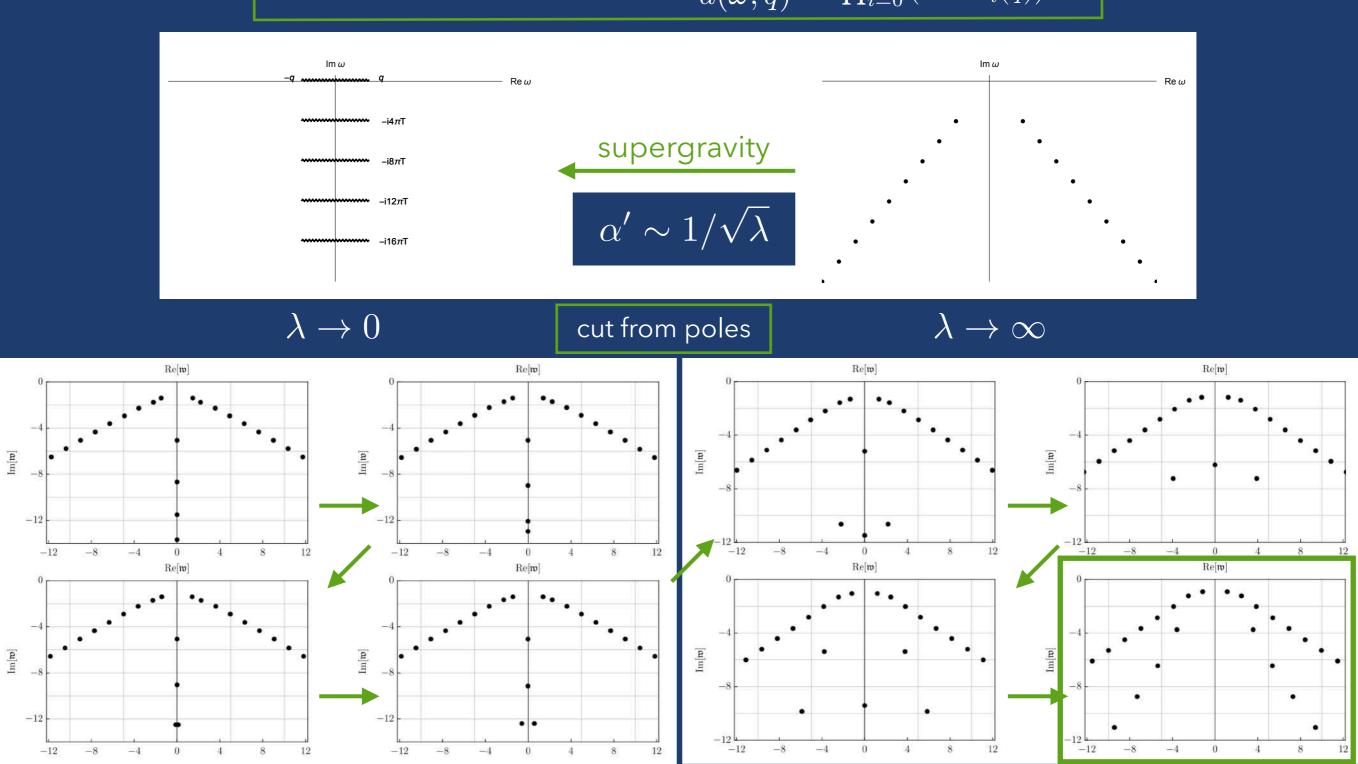




[SG, Starinets ..., several papers]

HOLOGRAPHY AND THERMAL SPECTRUM

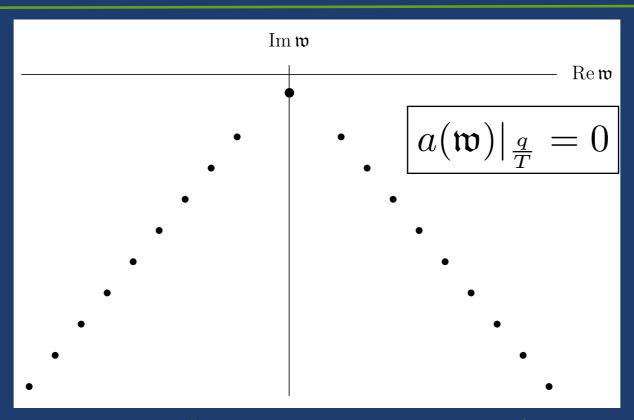
$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)} = \frac{B(\omega, q)}{\prod_{i=0}^{\infty} (\omega - \omega_i(q))}$$



[SG, Starinets ..., several papers]

HOLOGRAPHY AND THERMAL SPECTRUM

$$\langle T_{\mu\nu}(-\omega, -q), T_{\rho\sigma}(\omega, q) \rangle_R \sim \frac{b(\omega, q)}{a(\omega, q)} = \frac{B(\omega, q)}{\prod_{i=0}^{\infty} (\omega - \omega_i(q))}$$



meromorphic momentum space correlator

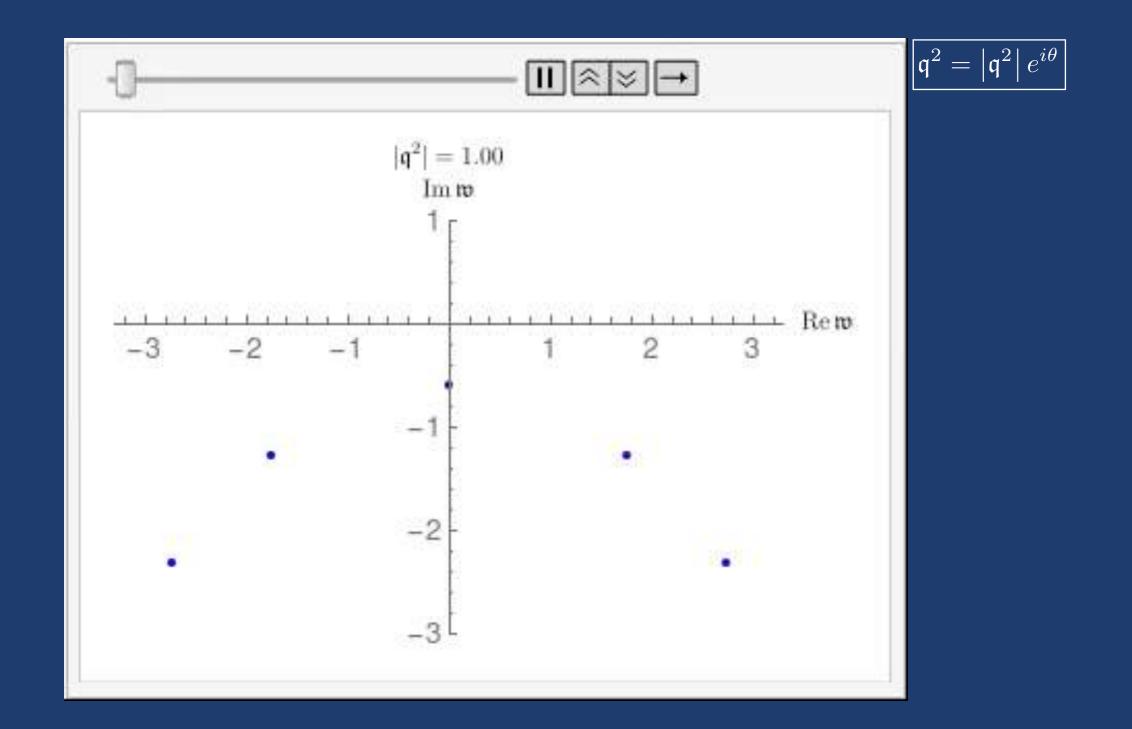
quantum field theory

spectra of linear non-Hermitian operators

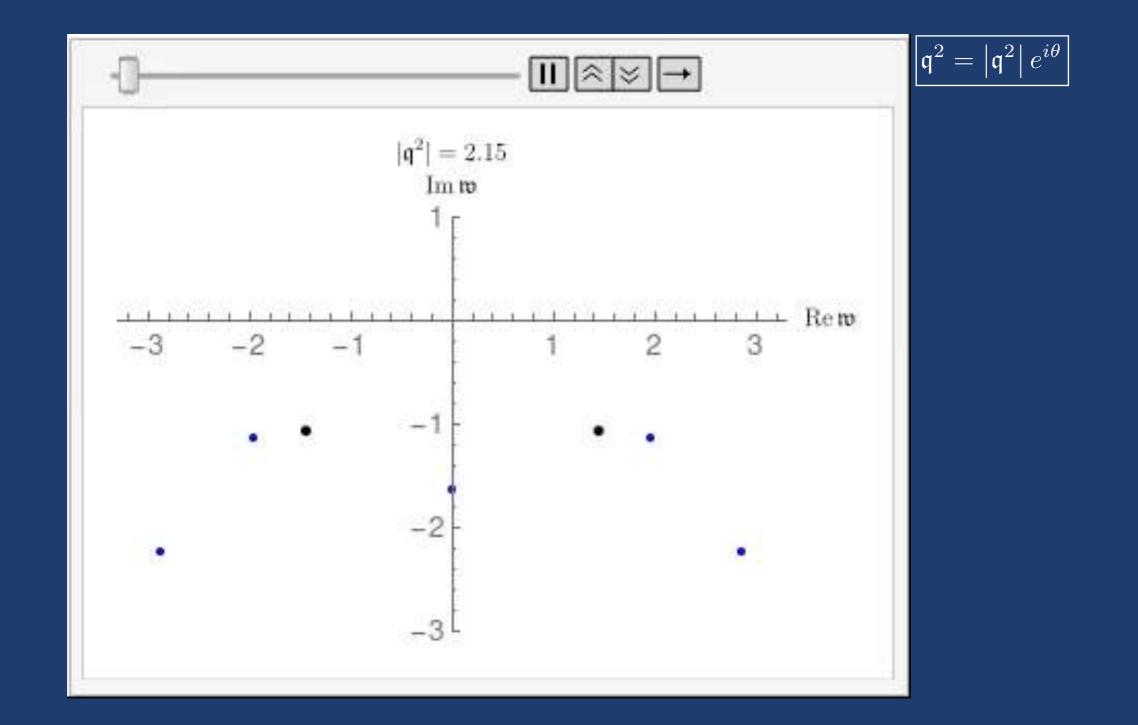
quasinormal mode spectrum of black holes

zeros of (algebraic) equations

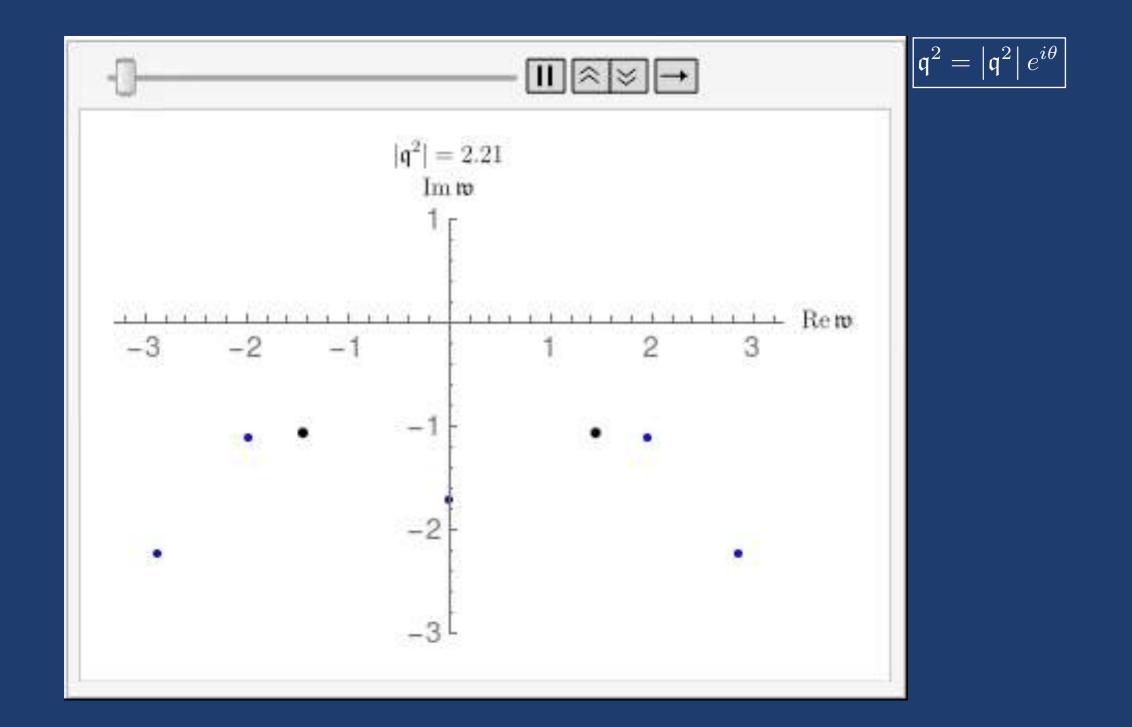
radius of convergence of $\mathfrak{w}(\mathfrak{q}) = \sum_{i=1}^{\infty} c_n \mathfrak{q}^n$, $|\mathfrak{q}| < \mathfrak{q}_*$, is set by the lowest momentum at which the hydro pole collides (level-crossing): $\mathfrak{q}_* = \min[|\mathfrak{q}_{\text{collision}}|]$



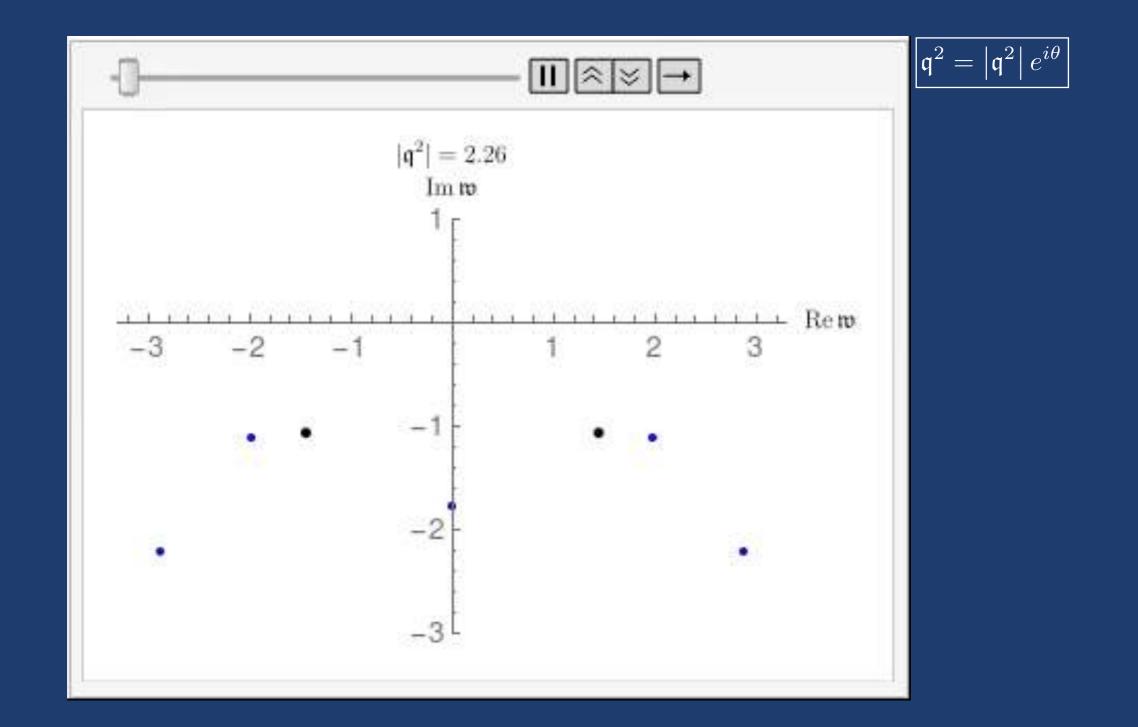
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• hydrodynamic series are convergent Puiseux series (shear p=1, sound p=2) [SG, Kovtun, Starinets, Tadić, PRL (2019); ...; see also Withers; JHEP (2018); Heller, et.al. (2020, ...)]

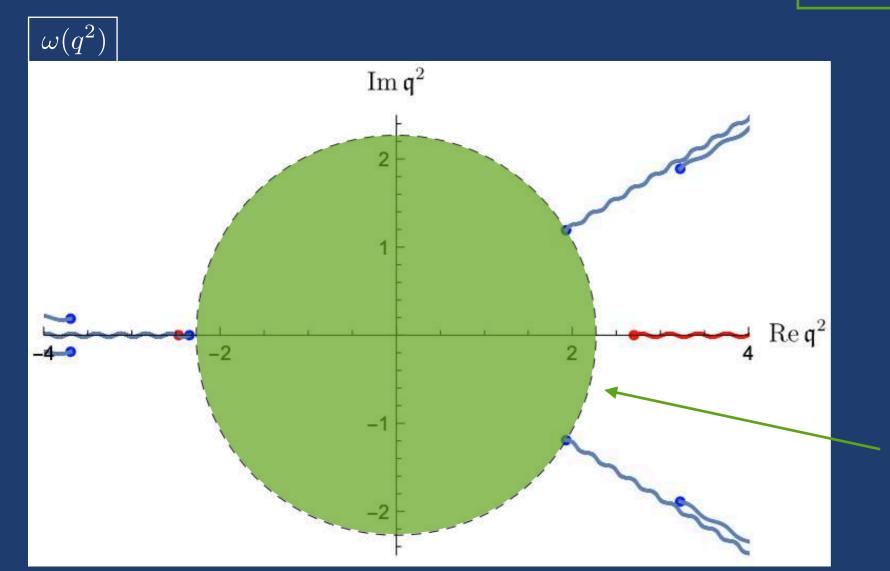
$$\mathfrak{w}_{\mathrm{shear}} = -i\sum_{n=1}^{\infty} c_n \left(\mathfrak{q}^2\right)^n = -i\mathfrak{D}\mathfrak{q}^2 + \dots$$

$$\mathfrak{w}_{\text{sound}} = -i \sum_{n=1}^{\infty} a_n e^{\pm \frac{i\pi n}{2}} \left(\mathfrak{q}^2\right)^{n/2} = \pm v_s \mathfrak{q} - \frac{i}{2} \mathfrak{G} \mathfrak{q}^2 + \dots$$

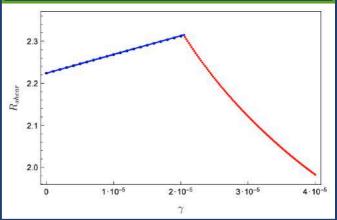
 dispersion relations are holomorphic in a disk

$$R_{\text{shear}}(\lambda) = 2.22 \left(1 + 674.15 \,\lambda^{-3/2} + \cdots \right)$$

 $R_{\text{sound}}(\lambda) = 2 \left(1 + 481.68 \,\lambda^{-3/2} + \cdots \right)$



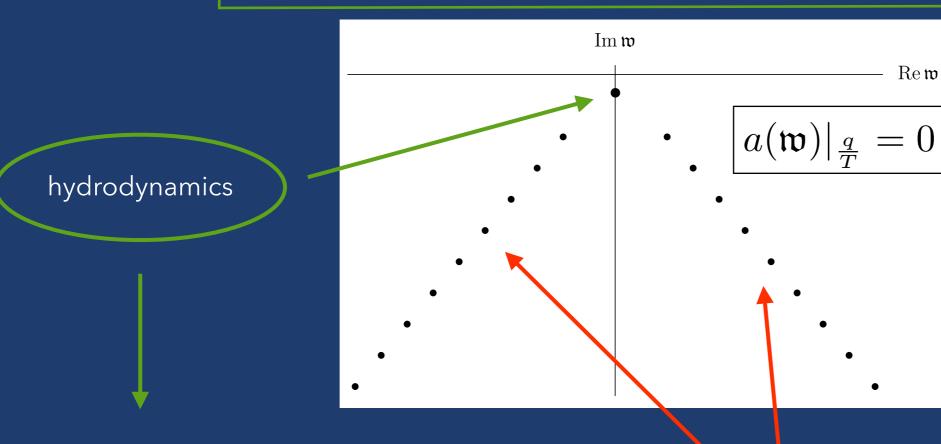
N=4 SYM radius convergence [SG, Starinets, Tadić, JHEP (2021)]

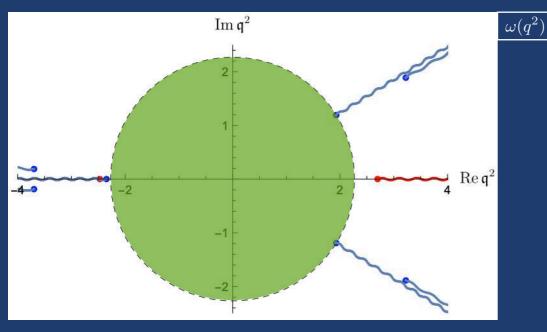


holomorphic disk

RECONSTRUCTION OF SPECTRA







gapped modes with $\omega(q=0) \neq 0$

[SG, Lemut, JHEP (2023)]

[see also Withers, JHEP (2019)]

PUISEUX AND DARBOUX THEOREMS

Puiseux theorem

Around a critical point of order p, we expect p branches of solutions

$$f(x_* = 0, y_* = 0) = 0, \ \partial_y f(0, 0) = 0, \ \dots, \ \partial_y^p f(0, 0) \neq 0$$

$$y = Y_j(x) = \sum_{k \ge k_0}^{\infty} a_k x^{k/m_j}, \quad j = 1, \dots, p$$

If some $\,m_j>1$, we necessarily have a family of $\,m_j$ solutions

$$y = Y_l(x) = \sum_{k \ge k_0}^{\infty} a_k \left(e^{\frac{2\pi i l}{m_j}} \right)^k x^{k/m_j}, \quad l = 0, 1, \dots, m_j - 1$$

recall: sound

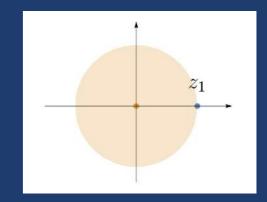
$$\mathfrak{w}_{\text{sound}} = -i \sum_{n=1}^{\infty} a_n e^{\pm \frac{i\pi n}{2}} \left(\mathfrak{q}^2\right)^{n/2} = \pm v_s \mathfrak{q} - \frac{i}{2} \mathfrak{G} \mathfrak{q}^2 + \dots$$

PUISEUX AND DARBOUX THEOREMS

Darboux theorem

Consider a power series

$$f(z) = \sum_{n=0}^{\infty} a_n z^n$$



that converges up to a critical point of order $\nu = -1/p$, which can be computed

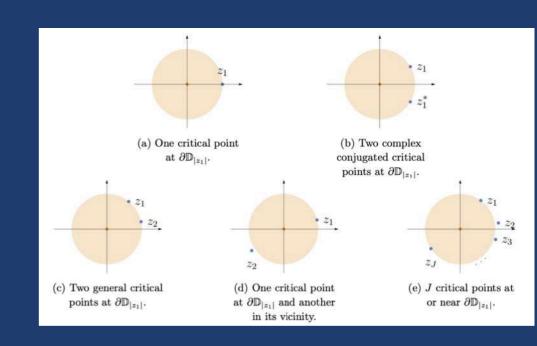
$$f(z) \sim (z - z_1)^{-\nu} [-1/2] r(z) + q(z)$$

$$\nu = \lim_{n \to \infty} \left[z_1(n+1) \frac{a_{n+1}}{a_n} - n \right]$$

as well as all coefficients in the expansion and subleading (non-singular) terms

$$r(z) = \sum_{m=0}^{\infty} r_m (z - z_1)^m$$

$$r_m = \lim_{n \to \infty} \left[\frac{(-1)^{m-\nu} n! z_1^{n-m+\nu} a_n}{(\nu - m)_n} - \sum_{k=0}^{m-1} \frac{(-1)^{m-k} (\nu - k)_n r_k}{(\nu - m)_n z_1^{m-k}} \right]$$



 Problem: need the location of the critical point or the exponent... this is resolved by Hunter and Guerrieri (1980), which we generalise and apply to finite number of coefficients

RECONSTRUCTION OF 'ALL' UV MODES

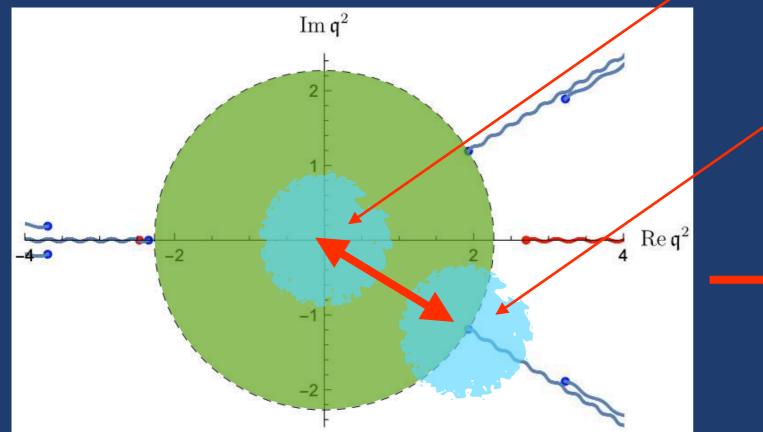
claim: systematic reconstruction of *all* modes connected via *level-crossing* is possible by exploration (analytic continuations) of the Riemann surface connecting physical modes

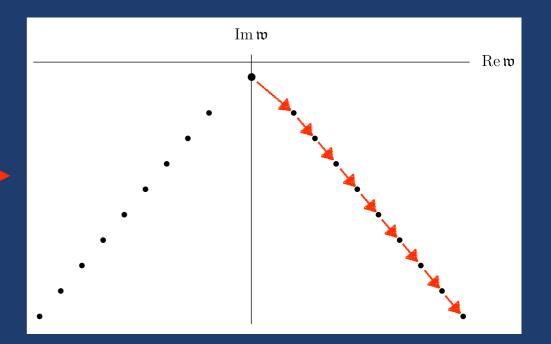
- statement should hold for spectra that are 'sufficiently complicated' Heun function
- momentum space analogue of resurgence in position space everything is convergent

$$\omega_0(z) = \sum_{n=0}^{\infty} a_n z^n$$

$$\omega_0(z) = -i \sum_{n=0}^{\infty} e^{\frac{i\pi n}{2}} b_n (z - z_1)^{n/2}$$

$$\omega_1(z) = -i \sum_{n=0}^{\infty} e^{-\frac{i\pi n}{2}} b_n (z - z_1)^{n/2}$$





all UV modes from one IR mode

EXAMPLE: MOMENTUM DIFFUSION OF M2 BRANES

- start from 300 hydrodynamic coefficients $\omega_0(z) = \sum_{n=0}^{\infty} a_n z^n$
- use algorithm with 2 c.c. critical points, 'recover' 12 coefficients and compute the gap with analytic continuation on the same sheet (Padé approximant, ...)

$$\mathfrak{w}_1(z) = \sum_{n=0}^{(N_1=12)-1} b_n (z-z_1)^{n/2}$$



$$\mathfrak{w}_1^{\text{calc}}(0) = 1.23506 - 1.76338i$$

$$\mathfrak{w}(0) = 1.23455 - 1.77586i$$

(re)compute the first 300 coefficients, use algorithm with 2 general critical points,
 'recover' 12 coefficients and compute the gap

$$\mathfrak{w}_2(z) = \sum_{n=0}^{(N_2=12)-1} c_n (z - z_2)^{n/2}$$



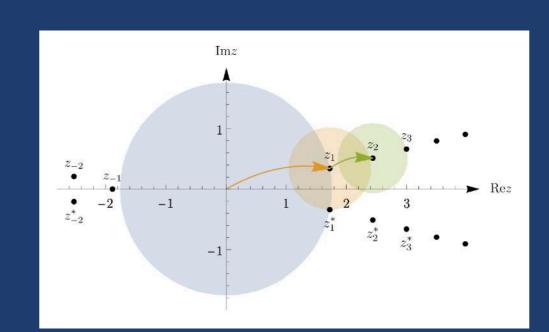
$$\mathfrak{w}_2^{\text{calc}}(0) = 2.16275 - 3.25341i$$

$$\mathfrak{w}_2(0) = 2.12981 - 3.28100i$$

... exploration continues ...



 conceptually useful and instructive, practically not quite (yet)...



POLE-SKIPPING AND THE RECONSTRUCTION

POLE-SKIPPING

- precise analytic relation between hydrodynamics and quantum chaos
 [SG, Schalm, Scopelliti, PRL (2017); Blake, Lee, Liu, JHEP (2018); Blake, Davison,
 SG, Liu, JHEP (2018); SG, JHEP (2019)]
- resumed all-order hydrodynamic series (e.g. sound) $\omega(q) = \sum_{n=1}^{\infty} \alpha_n \left(T, \mu_i, \langle \mathcal{O}_j \rangle, \lambda \right) q^n$

passes through a 'chaos point' at where the associated 2-pt function is "0/0":

$$\omega(q = i\lambda_L/v_B) = i\lambda_L = 2\pi Ti$$

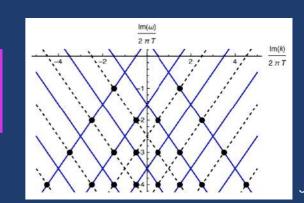
$$G_R = \frac{0}{0} = N(\delta\omega/\delta q)$$

relation to quantum chaos as measured by the out-of-time-ordered correlation functions

$$C(t,\mathbf{x}) = \left\langle [W(t,\mathbf{x}),V(0,\mathbf{0})]^{\dagger}[W(t,\mathbf{x}),V(0,\mathbf{0})] \right\rangle_{T} \sim \epsilon \, e^{\lambda_{L}(t-|\mathbf{x}|/v_{B})}$$
 butterfly velocity 'quantum' Lyapunov exponent

- triviality of Einstein's equations at the horizon
- infinite number of such '0/0' points:
 [SG, Kovtun, Starinets, Tadić, JHEP (2019);
 Blake, Davison, Vegh, JHEP (2019)]

$$\omega_n(q_n) = -2\pi Tin$$

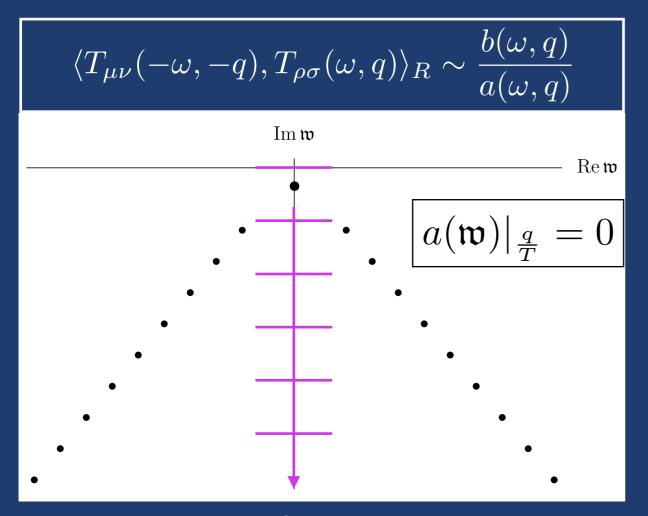


[from Blake, Davison, Vegh, JHEP (2019)]

DIFFUSION AND SPECIAL POLE-SKIPPING POINTS

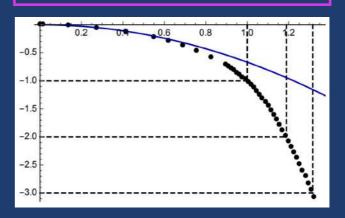
consider diffusion in a neutral 3d CFT dual to AdS₄-Schwarzschild

$$\mathfrak{w}_{\mathrm{shear}} = -i\sum_{n=1}^{\infty} c_n \left(\mathfrak{q}^2\right)^n = -i\mathfrak{D}\mathfrak{q}^2 + \dots$$



for increasing real q





[plot from Blake, Davison, Vegh, JHEP (2019)]

analytic result known for 4d bulk [SG, PRL (2021)]

$$q_n = \frac{4\pi T}{\sqrt{3}} n^{1/4}, \ n = 0, 1, 2, \dots$$

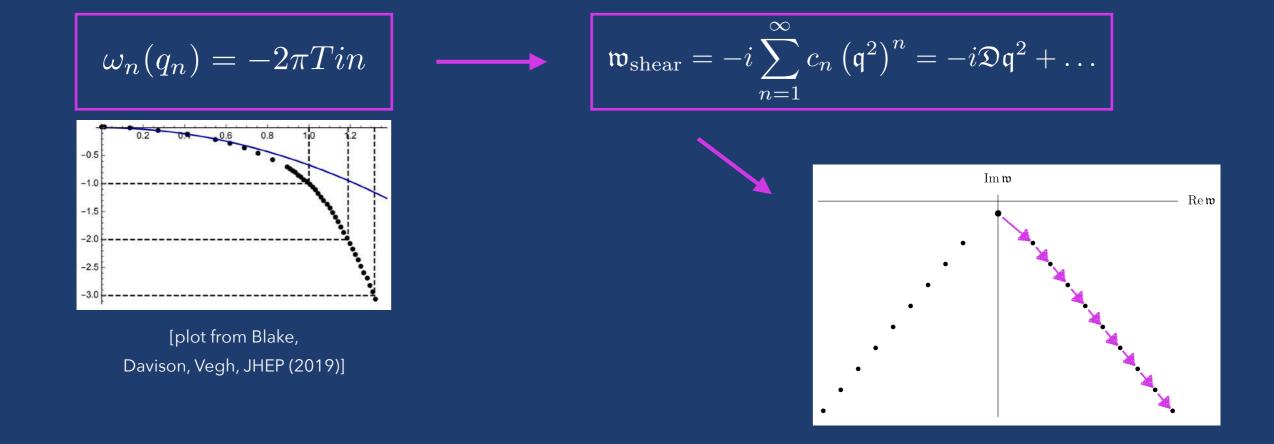
algebraically special points

4d odd-even duality: Darboux transform [Chandrasekhar (1983); SG, Vrbica (2023)]

how much information is required to reconstruct a QFT spectrum?
 so far in the talk: the knowledge of one dispersion relation

claim: in holographic theories of the type discussed here (N=4 SYM, M2, M5, ...), the entire spectrum can be computed from only a discrete set of pole-skipping points

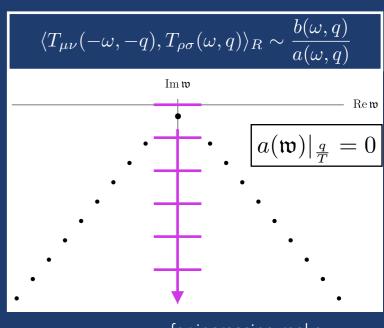
[SG, Lemut, Pedraza, PRD (2023)]



interpolation problem:

$$\omega_n(q_n) = -2\pi Tin$$

$$\mathbb{w}_{\text{shear}} = -i\sum_{n=1}^{\infty} c_n \left(\mathfrak{q}^2\right)^n = -i\mathfrak{D}\mathfrak{q}^2 + \dots$$



- for increasing real q
- unique solutions to interpolation problems are extremely hard in the absence of special known properties of the function (e.g. Weierstrass, Hadamard, Nevanlinna-Pick, ...)
- trick: 'analytic continuation' to d spacetime dimensions and expansion around infinite d
- general relativity in large *d* drastically simplifies [review by Emparan, Herzog, 2003.11394]

$$V \sim 1/r^d$$

- recall: large-d limit of quantum mechanics allows a solution of the Helium problem
- convergence of such series depends on the details

interpolation: $\omega_n(q_n) = -2\pi Tin$ $\mathfrak{w}_{\mathrm{shear}} = -i\sum_{\mathbf{1}} c_n \left(\mathfrak{q}^2\right)^n = -i\mathfrak{D}\mathfrak{q}^2 + \dots$

'analytic continuation' to d spacetime dimensions and expansion around infinite d

$$\omega_0(q) = -i\left(\frac{q}{\sqrt{d}}\right)^2 - i\sum_{m=2}^{\infty} \frac{1}{d^m} \sum_{j=2}^m c_{m,n} \left(\frac{q}{\sqrt{d}}\right)^{2j} \qquad \frac{q_n}{\sqrt{d}} = \sqrt{\frac{nd}{2}} \left(1 + \sum_{m=1}^{\infty} \frac{b_{n,m}}{d^m}\right)$$

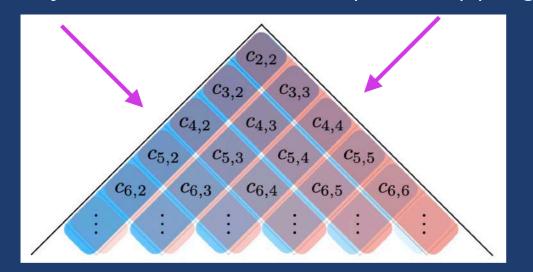
$$b_{n,1}=-\sum_{m=2}^{\infty}\frac{n^{m-1}c_{m,m}}{2^m}$$

$$b_{n,2}=-\frac{b_{n,1}^2}{2}-\sum_{m=2}^{\infty}\frac{n^{m-1}\left(c_{m+1,m}+2mb_{n,1}c_{m,m}\right)}{2^m}$$
 second analytic continuation
$$n\in\mathbb{Z}\to x\in\mathbb{R}$$

$$n \in \mathbb{Z} \to x \in \mathbb{R}$$

hydrodynamics

pole-skipping



$$c_{m,m} = -\frac{2^m}{(m-1)!} \partial_x^{m-1} b_1(0)$$

$$c_{m+1,m} = -\frac{2^m \partial_x^{m-1} b_2(0)}{(m-1)!} + \sum_{j=2}^{m-1} \left(j - \frac{1}{4}\right) c_{j,j} c_{m-j+1,m-j+1}$$

generating functions

$$\omega_0(q) = -i\left(\frac{q}{\sqrt{d}}\right)^2 - i\sum_{m=2}^{\infty} \frac{1}{d^m} \sum_{j=2}^m c_{m,n} \left(\frac{q}{\sqrt{d}}\right)^{2j} \qquad \frac{q_n}{\sqrt{d}} = \sqrt{\frac{nd}{2}} \left(1 + \sum_{m=1}^{\infty} \frac{b_{n,m}}{d^m}\right)$$

first level:
$$b_{n,1} = -\frac{1}{2}H_n = -\frac{1}{2}\sum_{k=1}^{n}\frac{1}{k}$$
 $H_n \to H(x) = \sum_{k=1}^{\infty}\frac{x}{k(x+k)}$

$$c_{m,2} = c_{m,m} = (-1)^m 2^{m-1} \zeta(m)$$

$$\omega_0(q) = -i\bar{q}^2 - i\frac{\bar{q}^2}{d} H_{2\bar{q}^2/d} + \dots$$

partial (convergent) resummation

second level benefits from numerical approach (Padé interpolation):

$$b_{n,2} = a_n + b_n H_n + c_n H_n^{(2)} + d_n H_n H_n^{(2)} + e_n H_n^2 + \sum_{k=1}^{n-3} \left(j_{n,k} + k_{n,k} H_n + l_{n,k} H_k + m_{n,k} H_k^{(2)} \right)$$

$$+ \frac{n}{8(n+1)} \left[\sum_{k=1}^{n-3} \frac{1}{k+1} \left(\frac{n+2}{n} \right)^k \right] \left[\sum_{k=1}^{n-3} \frac{1}{k+1} \left(\frac{n}{n+2} \right)^k \right] - \frac{3n}{8(n+1)} \sum_{k=2}^{n-3} \frac{1}{k+1} \left(\frac{n}{n+2} \right)^k \sum_{j=0}^{k-2} \frac{1}{j+1} \left(\frac{n+2}{n} \right)^j$$

$$+ \frac{1}{2(n+1)} \sum_{k=2}^{n-3} H_k \left(\frac{n}{n+2} \right)^k \sum_{j=0}^{k-2} \frac{1}{j+1} \left(\frac{n+2}{n} \right)^j + \frac{n+2}{4n(n+1)} \sum_{k=2}^{n-3} \frac{1}{k-1} \left(\frac{n}{n+2} \right)^{k-1} \sum_{j=k}^{n-3} \frac{1}{j+1} \left(\frac{n+2}{n} \right)^j$$

$$+ \frac{n+2}{4n(n+1)} \sum_{k=2}^{n-3} \frac{1}{k-1} \left(\frac{n}{n+2} \right)^{k-1} \sum_{j=k}^{n-3} H_j \left(\frac{n+2}{n} \right)^j + \frac{3n^2}{8(n+1)(n+2)} \sum_{k=1}^{n-3} \frac{1}{k} \sum_{j=k+1}^{n-3} \frac{1}{j} \left(\frac{n+2}{n} \right)^{n-j}, \quad (B17)$$

$$\begin{split} a_n &= \frac{1-2n}{4\left(n^2-1\right)} \left(\frac{n}{n+2}\right)^{n-3} + \frac{12-n(2(n-3)(n-1)n+11)}{8n^2\left(n^2-4\right)^2\left(n^2-1\right)} \left(\frac{n+2}{n}\right)^n \\ &- \frac{n^9-17n^8+77n^7-103n^6-44n^5+4n^4+378n^3-544n^2+392n-96}{8(n-2)^2(n-1)^3n^2(n+1)(n+2)}, \\ b_n &= \frac{n^7-10n^6+21n^5-32n^4-28n^3+72n^2-96n+96}{16(n-2)^2(n-1)^2n(n+1)(n+2)} \\ &+ \frac{(5n+6)}{8(n+1)} \left(\frac{n}{n-2}\right)^{n-2} - \frac{3(n-2)}{16(n+1)} \left(\frac{n+2}{n}\right)^{n-2}, \\ c_n &= -\frac{n\left(6n^2-16n+9\right)}{8(n+1)\left(n^2-3n+2\right)}, \\ d_n &= -\frac{n}{4n+4}, \\ e_n &= \frac{5n^3-8n^2-10n+10}{8\left(n^3-2n^2-n+2\right)}, \end{split}$$

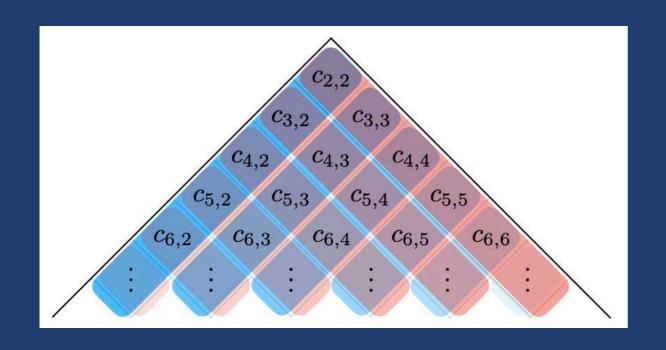
```
\begin{split} j_{n,k} &= \left(\frac{(k(k(5k+24)+23)+8)n^2 - 2(3k+1)(k+1)^2n + 4k(k+1)^2}{8k(k+1)^2(n-2)(n-1)n(n+1)(k-n+1)} \right. \\ &+ \frac{-k(k+2)n^5 + (k(k(3k+10)+11)+3)n^4 + (-k(k+2)(7k+12)-9)n^3}{8k(k+1)^2(n-2)(n-1)n(n+1)(k-n+1)} \left. \left(\frac{n+2}{n}\right)^k \right. \\ &+ \frac{k(n(n(n+3)+6)-4) + n(6-(n-3)n(n+1)) - 4}{8k(k+1)(n^2-1)(k-n+1)} \left(\frac{n}{n+2}\right)^k \\ &+ \frac{2k^2n^2 - 5k^2n - 2kn^3 + 8kn^2 - 9kn + n^3 - 3n^2 + 2n}{8k^2(k+1)(n-2)(n-1)(n+1)}, \\ k_{n,k} &= \frac{2k^2(k+1)^2 + (k(k+4)+1)n^2 - 2k(k+1)^2n - n^3}{8k^2(k+1)n(n+1)(k-n+1)} \left(\frac{n}{n}\right)^k \\ &+ \frac{-2k(n+1)(n+6) + n(n(2n+7)-18) - 24}{8k(k+1)(n+2)(k-k-n-2)(-k+n-1)} \left(\frac{n}{n}\right)^k, \\ l_{n,k} &= \frac{n(4k(n+2) - 3n^2 + 4n + 8)}{8(k+1)(n+1)(n+2)^2} \left(\frac{n+2}{n}\right)^{n-k} - \frac{(n-2k)}{8k(k+1)(n+1)} \left(\frac{n+2}{n}\right)^k \\ &+ \frac{4(k(k+2) - 2)(k(k+1)1 + 6)n^2 - 8k(k+2)(k-1)(k-1)n + 16k(k+2)(k+1)^2}{8k^2(k+1)^2(k+2)n(n^3 - 5n^2 + 4)} \\ &+ \frac{4(k(k+2) - k(k+1) + 4)n^2 + 2k(k+2)(k+1)(k-1)n + 16k(k+2)(k+1)^2}{8k^2(k+1)^2(k+2)n(n^3 - 5n^2 + 4)} \\ &+ \frac{4(k(3-5k-2)n^9 + (k(10 - 3k(k(3k+7) - 2)) - 4)n^2}{8k^2(k+1)^2(k+2)n(n^4 - 5n^2 + 4)}, \\ k_{n,k} &= \frac{4(k^3 - 5k - 2)n^9 + (k(10 - 3k(k(3k+7) - 2)) - 4)n^2}{4k(k+2)(n+1)}, \end{split}
```

results:

$$c_{2,2} = 2\zeta(2),$$

 $c_{3,2} = -4\zeta(3), \quad c_{3,3} = -4\zeta(3),$
 $c_{4,2} = 8\zeta(4), \quad c_{4,3} = 7 \times 8\zeta(4), \quad c_{4,4} = 8\zeta(4)$

- symmetries
- 'polylogs'[see e.g. Aminov et.al.2307.10141]



$$b_{n,1} \longrightarrow c_{m,2} = c_{m,m} = (-1)^m 2^{m-1} \zeta(m)$$

$$b_{n,2} \longrightarrow$$

$$c_{3,2} \approx -1.000 \times 4\zeta(3),$$

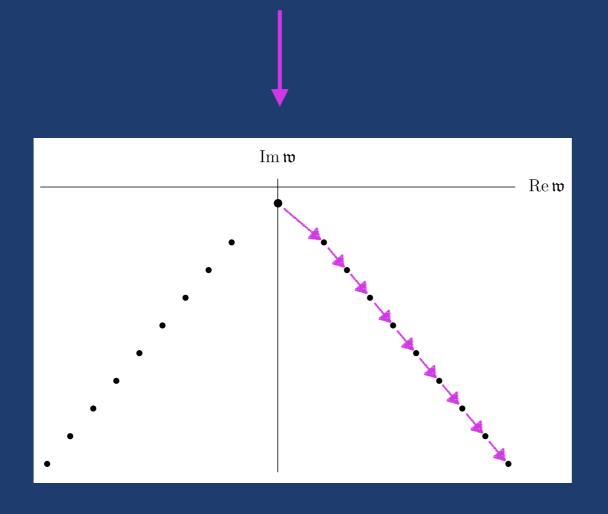
 $c_{4,3} \approx 7.001 \times 8\zeta(4),$
 $c_{5,4} \approx -15.548 \times 16\zeta(5),$
 $c_{6,5} \approx 27.546 \times 32\zeta(6)$

$$b_{n,3} \longrightarrow$$

$$c_{4,2} \approx 1.000 \times 8\zeta(4),$$

 $c_{5,3} \approx -15.502 \times 16\zeta(5)$

the rest of the spectrum follows from a reconstruction discussed before



- complete reconstruction of the spectrum using only algebraic near-horizon manipulations (local instead of global (ODE/PDE) analysis)
- 'full' correlator follows from poles product formula by Dodelson, et.al., 2304.12339

SUMMARY AND FUTURE DIRECTIONS

SUMMARY AND FUTURE DIRECTIONS

- complex analytic structures of transport are a powerful tool for exploring physics
- new results about the nature of thermal QFTs
- in momentum space we can deal with convergent series, but, 'morally', this is equivalent to resurgence in position space
- useful not only in QFTs, also for QNMs and general linearised operator spectra
- improve practical aspects of reconstructions given a limited number of known coefficient
- can these techniques be used in realistic QFTs (Euler-Heisenberg, chiral Lagrangian)?
- claim: reconstruction from IR to UV is possible from a discrete set of pole-skipping points
- large-N QFT calculation \longrightarrow differential equations \longrightarrow algebraic manipulations

THANK YOU!