



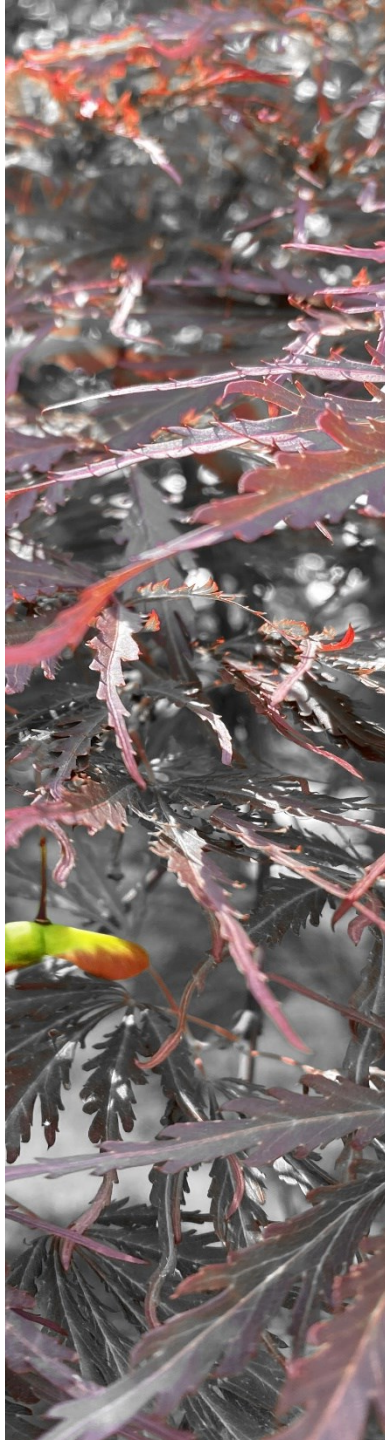
Cristian **Pira**
on behalf of LNL SRF group

S  **MARA**
SRF R&D @LNL

The section header features the word 'Samarium' where the 'O' is replaced by a stylized green and yellow atom symbol. Below it, 'SRF R&D @LNL' is written in a blue, sans-serif font.

Meeting CSN5 LNL

13/03/2024





Superconducting **A**lternative **M**aterials
for **A**ccelerating cavities and
haloscope **R**esonators for **A**xions

Material	Tc	Hc2
Nb	9.2 K	0.4 T
Nb₃Sn	18.3 K	30 T

**WORKS IN VERY HIGH
MAGNETIC FIELD**
Axion Haloscopes

**REDUCES
CRYOGENIC POWER
BY A FACTOR 3**
Accelerating Cavities

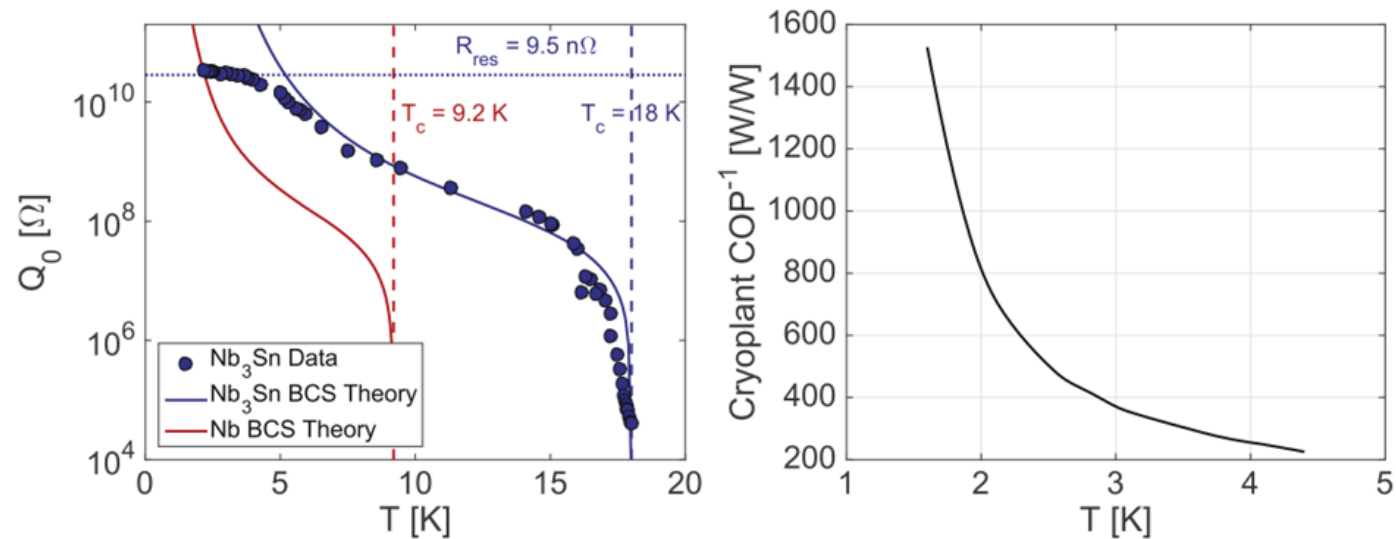
Accelerating Cavities

Energy saving is mandatory for the **next generation accelerators**...

...**cryogenics** is one of the **larger energy cost** in modern SRF accelerators



Move from bulk Nb @2K to **Nb₃Sn @4.5 K**
reduces cryogenic power by a factor of 3



Quantum Sensing

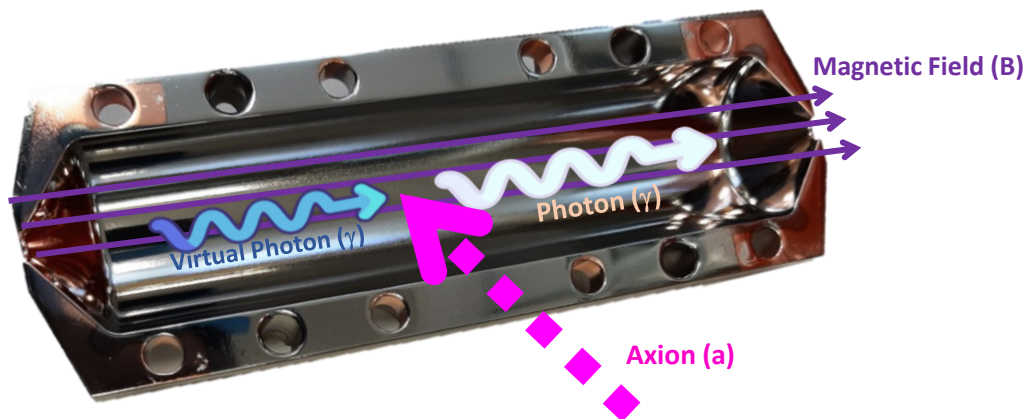
SRF Axion Haloscopes

- ▶ The axion haloscope was designed to scope microwave photon signals from the axions in our galactic halo
- ▶ The signal power can be enhanced by the resonance effect that occurs when the axion mass matches the natural frequency of the detection system

$$P_{a\gamma\rightarrow\gamma} = g_{\alpha\gamma\gamma}^2 \frac{\rho_a}{m_a} B^2 \omega_0 V C \frac{Q_a Q_c}{Q_a + Q_c}$$

Conversion Power: $P_{a\gamma\rightarrow\gamma}$
 Coupling Constant: $g_{\alpha\gamma\gamma}^2$
 Dark Matter Axion Density: ρ_a
 Axion Mass: m_a
 Magnetic Field: B^2
 Resonant Frequency: ω_0
 Volume: V
 Cavity Quality Factor: C
 Axion Quality Factor (10^6): Q_a
 Cavity Quality Factor: Q_c

D. Kim et al. JCAP03(2020)066

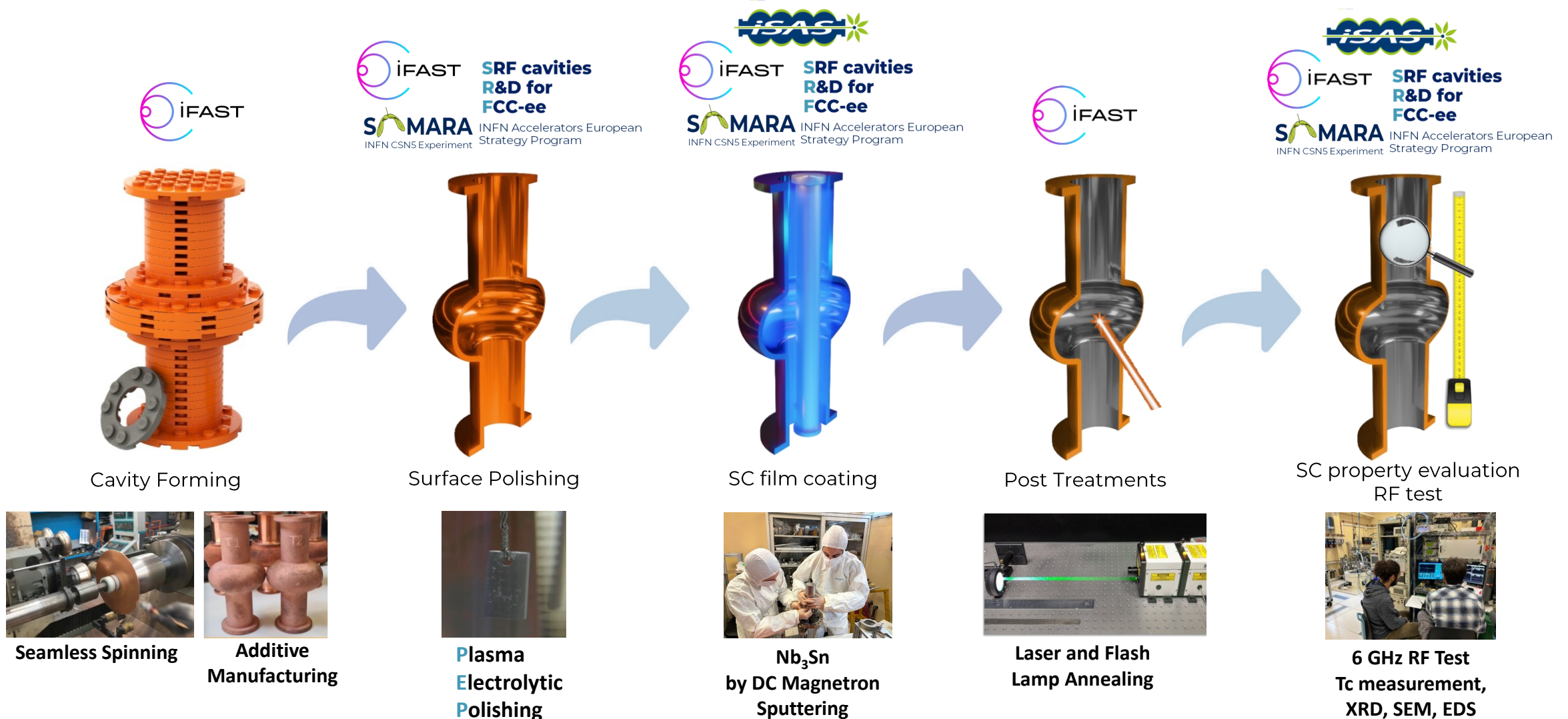


Q factor of Cu haloscope limited
by anomalous skin effect

to improve Q

SRF in High Magnetic Field
(Mixed State)

R&D in all the production chain



LNL Personnel involved in SRF R&D activities



Permanent Staff

Name	Profile	%
Cristian Pira	Tecnologo	45
Giorgio Keppel	Primo Tecnologo	15
Fabrizio Stivanello	Coll. Tecnico E.R.	55



Non Permanent Staff

Name	Profile	%
Davide Ford	Coll. Tecnico E.R.	100
Dorothea Fonnesu	Assegnista	100



Non Permanent Staff

Name	Profile	%
Roberta Caforio	Borsista	20
Giovanni Marconato	Borsista	100
Alessandro Salmaso	Borsista	80



Non Permanent Staff

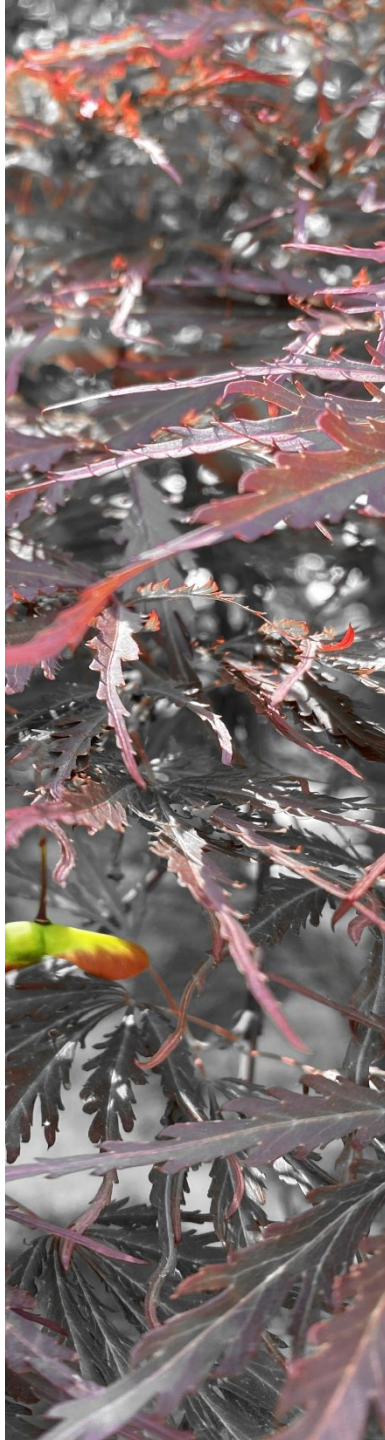
Name	Profile	%
Eduard Chyhyrynets	Tecnologo	100
Thomas Bortolami	Coll. Tecnico	100

2 Laureandi: Anita Fetaj e Matteo Lazzari





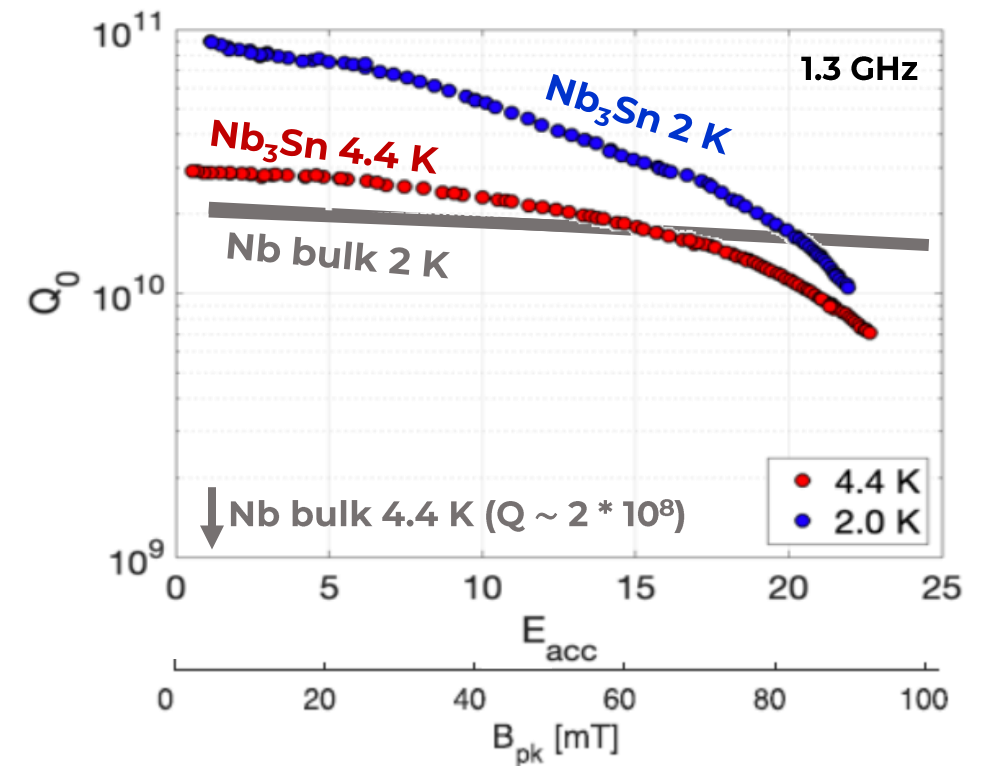
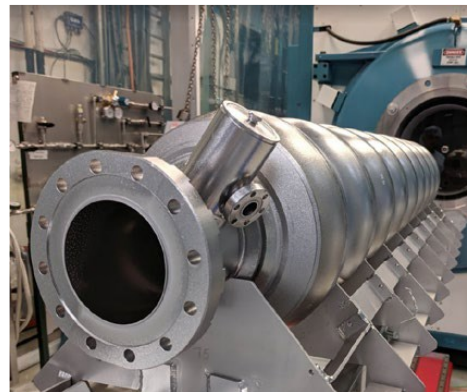
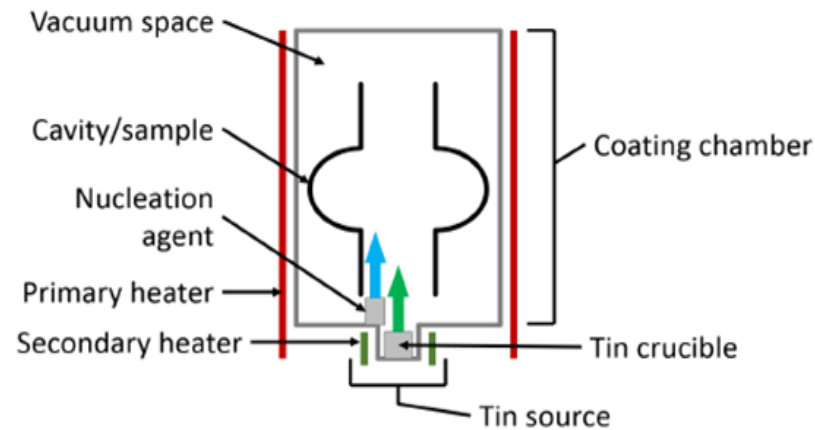
S  **MARA**
BACKUP SLIDES



Nb₃Sn state of the art

Vapor Tin Diffusion

Cornell, Fermilab, JLab, KEK



S. Posen, SRF 2019 proceedings (elaborated)

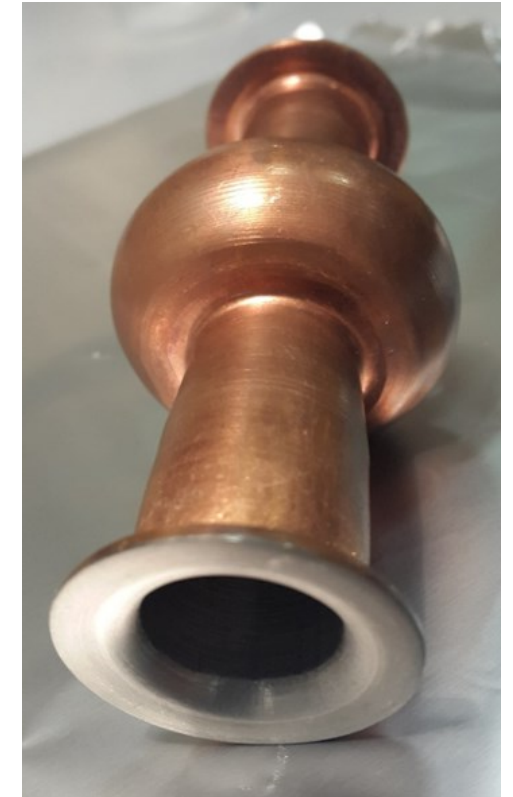
Technology limitation:

- ▶ Reproducibility
- ▶ Substrate cost

A different approach: Nb₃Sn on Cu

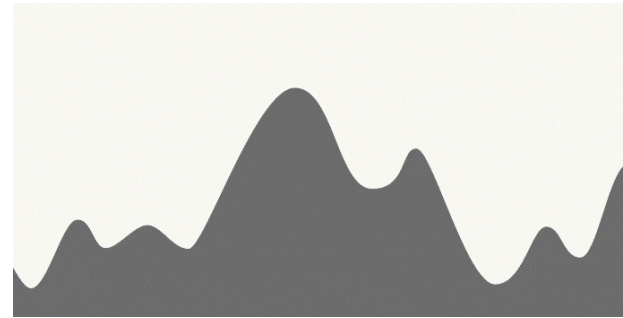
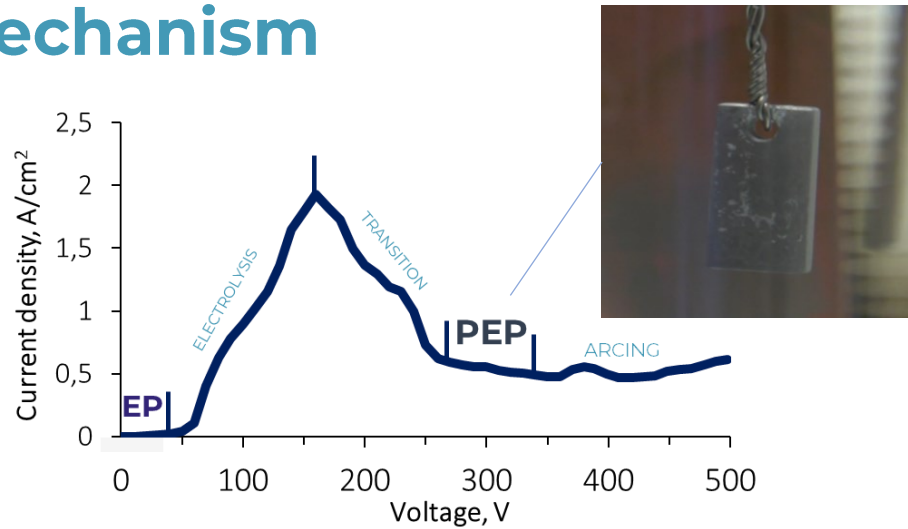
Cu substrate as several advantages:

- ▶ **Cheaper** than Nb
- ▶ Higher **thermal conductivity**
- ▶ Higher **mechanical stability**
- ▶ **PVD technology** (Nb on Cu) already used for LEP, LHC, HIE-ISOLDE @ CERN
ALPI @ INFN LNL
- ▶ **Interlayer** can be added to engineering the surface



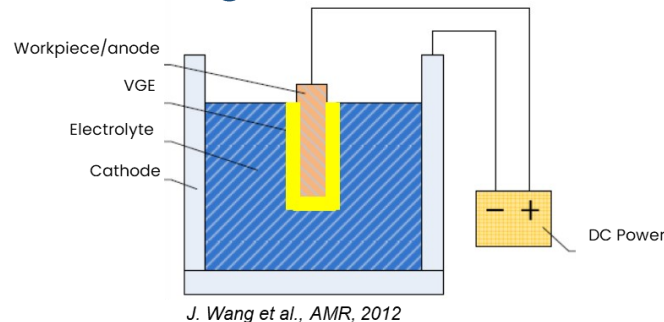
Plasma Electrolytic Polishing PEP

Mechanism



Advantages

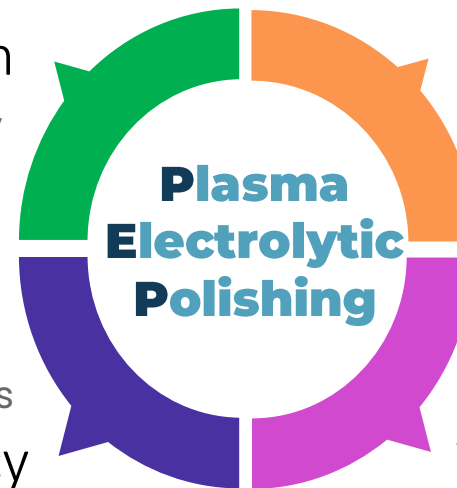
Same EP set-up
Different regime



Green
Diluted water solutions,
environmentally friendly

Equal thickness removal yield
lowest roughness among
competitors

Efficiency



Fast
The fastest
non-destructive
polishing

Less sensitive to the
cathode shape!
AM compatible

Versatility



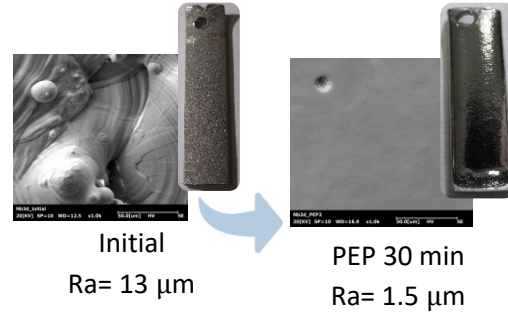
Surface Polishing



Plasma Electrolytic Polishing PEP Results

1x Nb 3x Cu
Solution Patentees by INFN

Additive Manufacturing

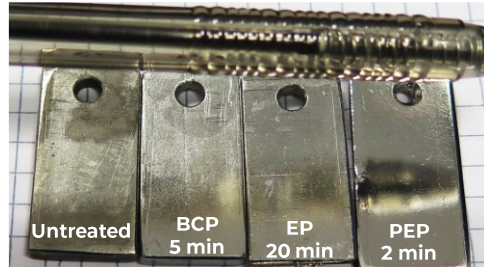


QPR Samples Helmholtz Zentrum Berlin

Nb QPR polishing optimization on-going

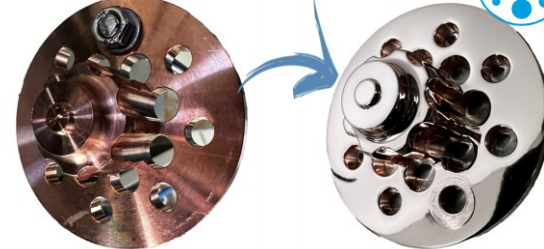
Full Cu QPR ready for coating

Nb planar samples



6.5 μm removed

Cu Photocathodes



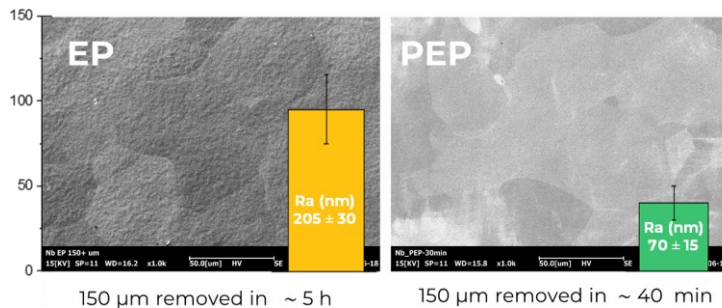
Ra ~ 8 nm!!!

6 GHz Cu cavity



No internal cathode!

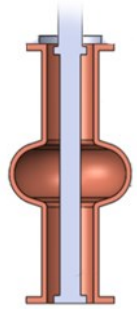
70 μm removed in 10 minutes
30 A (100 $\text{cm}^2 \rightarrow 1.3 \text{ GHz} \sim 300 \text{ A}$)



Surface Polishing



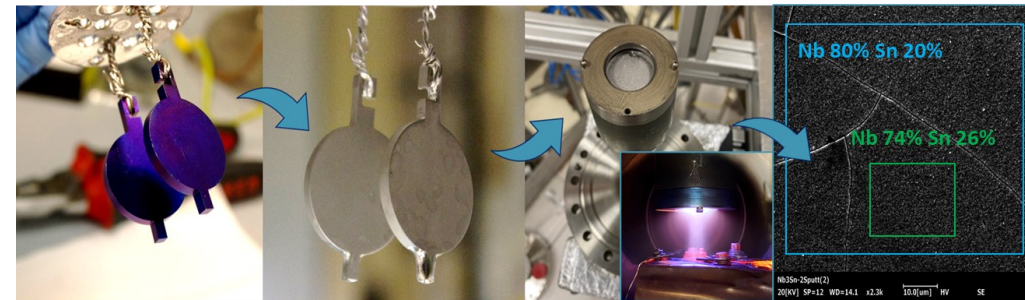
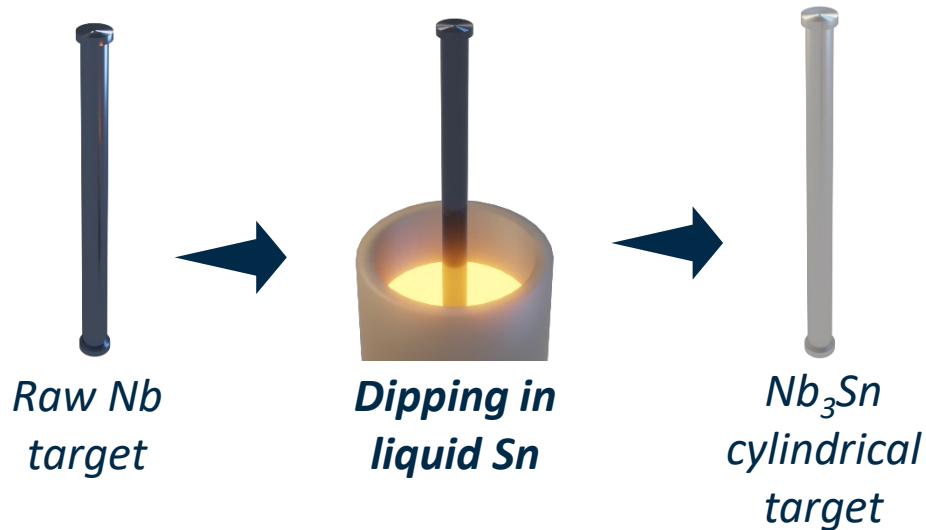
Nb₃Sn coatings: target production



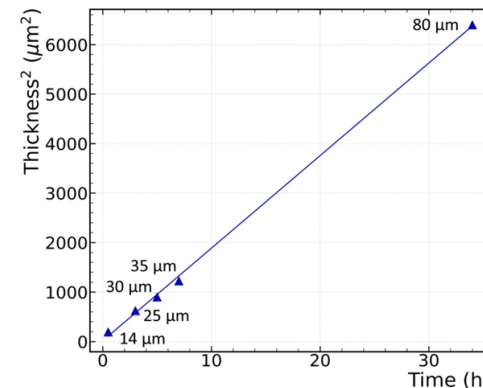
Single target configuration **easiest to scale** onto elliptical geometry

Nb₃Sn cylindrical target are not commercially available

LNL Strategy for Nb₃Sn cylindrical target production for 6 GHz cavities



Proof of concept



Nb₃Sn **thickness** related to **dipping time**

Possible **tin content modulation**



SC coatings



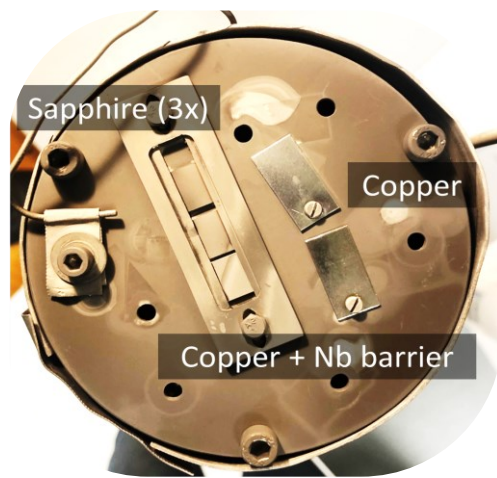
Nb₃Sn coatings

Sputtering parameter optimization

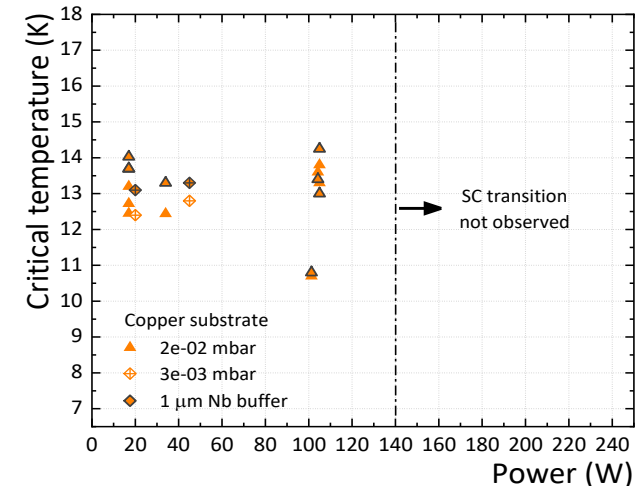
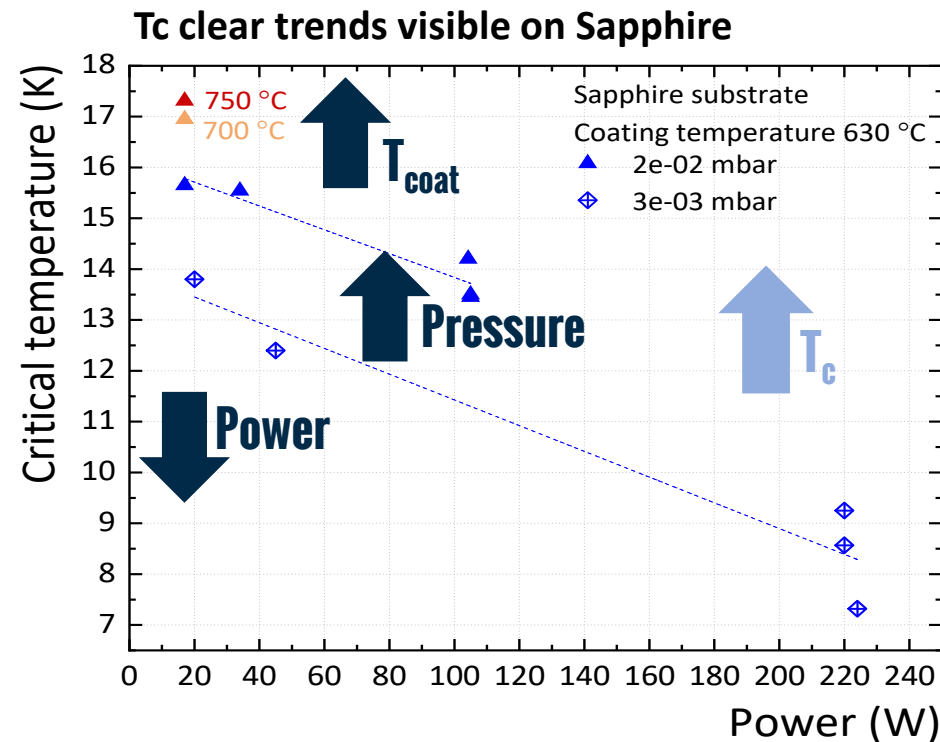
Nb₃Sn deposited via DCMS from 4" planar stoichiometric target in Ar atmosphere



Single target configuration **easiest to scale** onto elliptical geometry



In the same run different substrates are coated in the same conditions



Nb₃Sn T_c sticks at 14 K on Cu



Interface effect

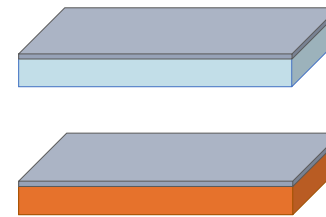
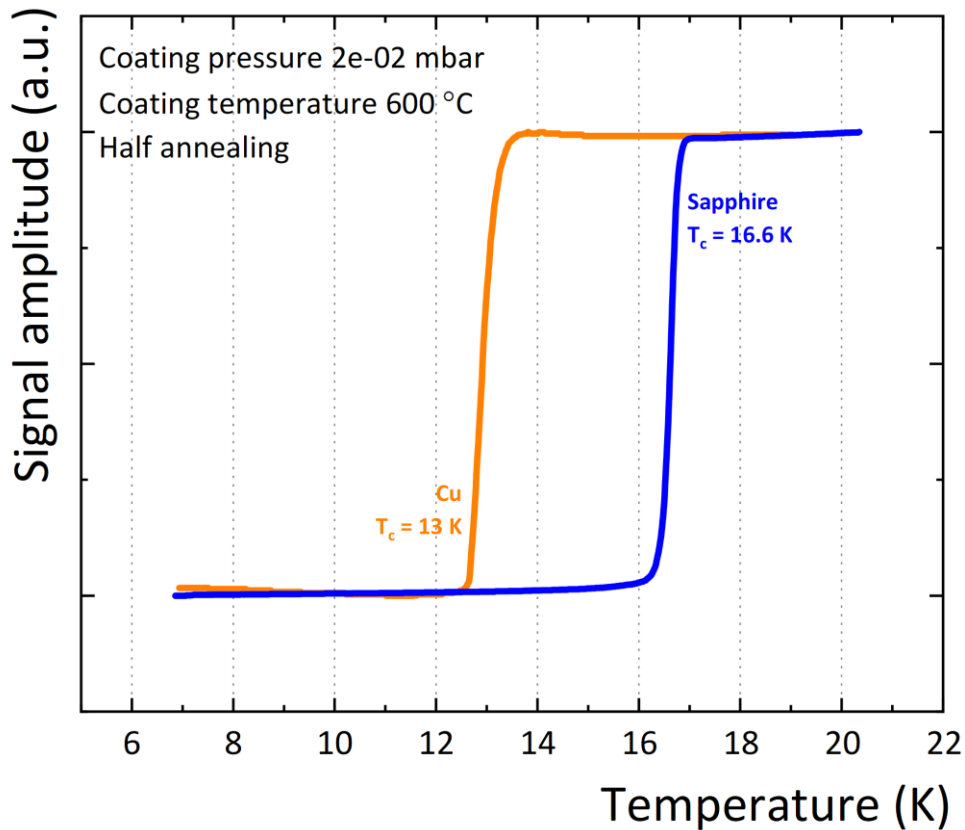
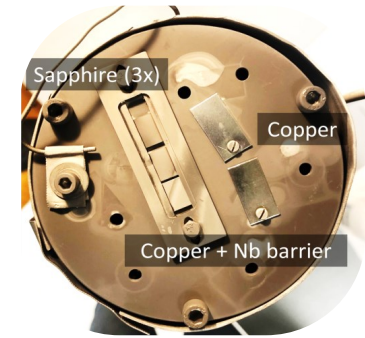


SC coatings



Nb₃Sn coatings

Sputtering parameter optimization



Sapphire + 1 μ Nb₃Sn

Cu + 1 μ Nb₃Sn

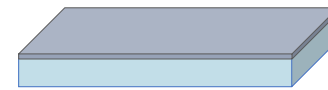
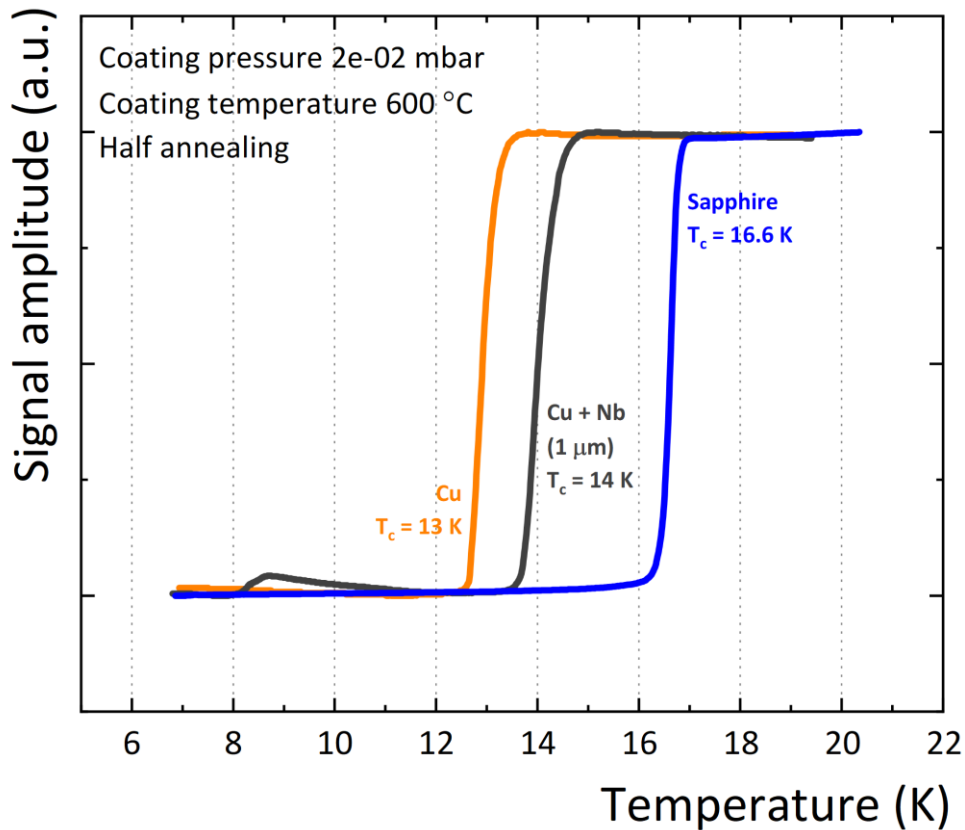
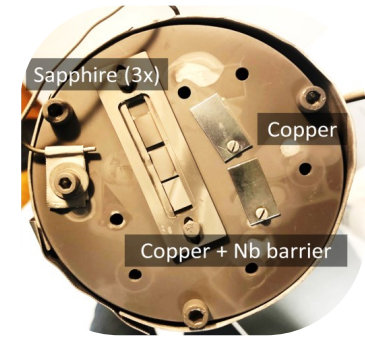


SC coatings



Nb₃Sn coatings

Sputtering parameter optimization



Sapphire + 1 μ Nb₃Sn



Cu + 1 μ Nb₃Sn



Cu + 1 μ Nb + 1 μ Nb₃Sn

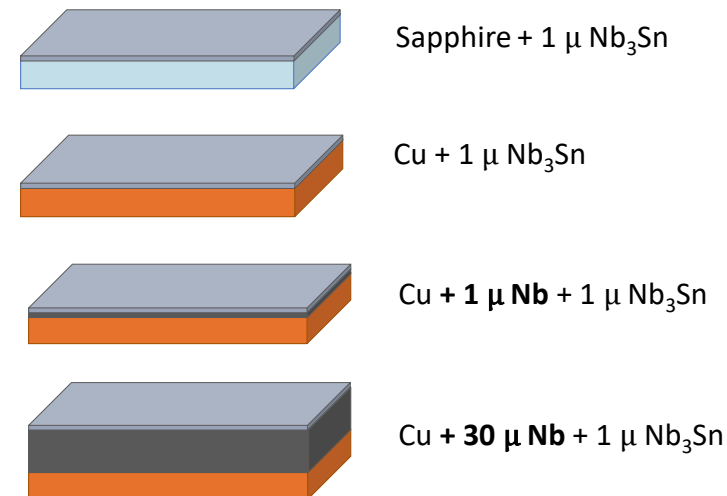
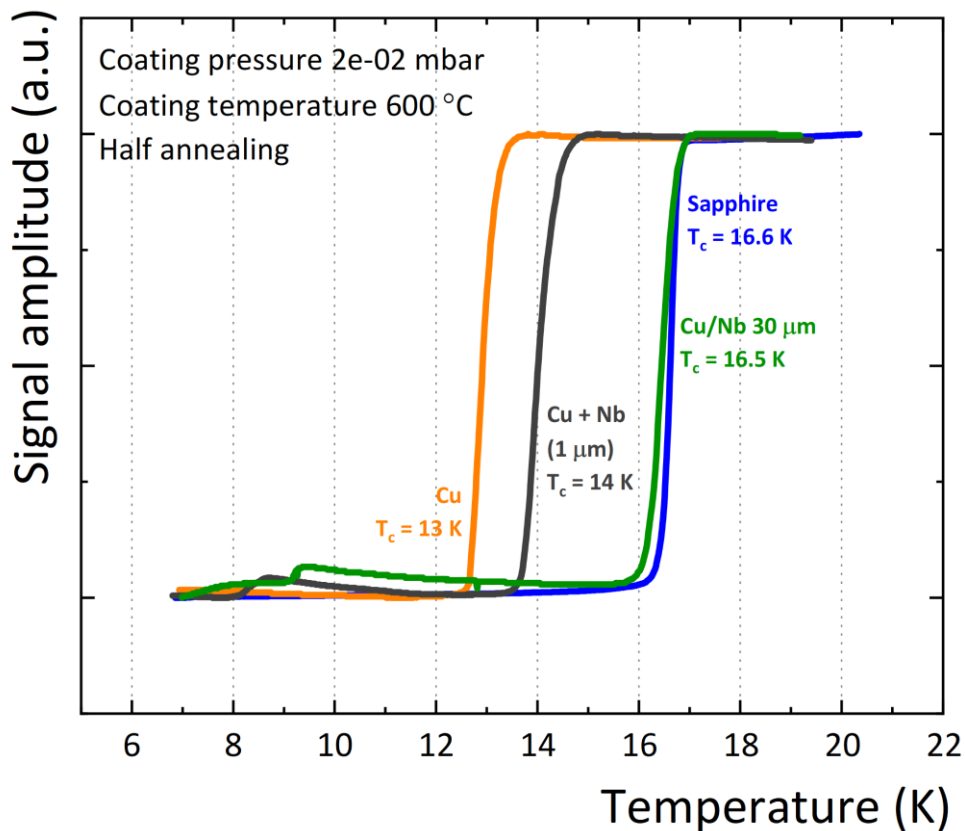
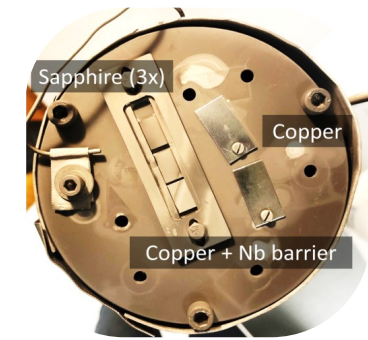


SC coatings



Nb₃Sn coatings

Sputtering parameter optimization



A very thick Nb layer enhance T_c close to Nb₃Sn nominal values

**$T_c = 17.3 \pm 0.25$ K
on Cu+50 um BL @ $T_{dep}=600$ °C**

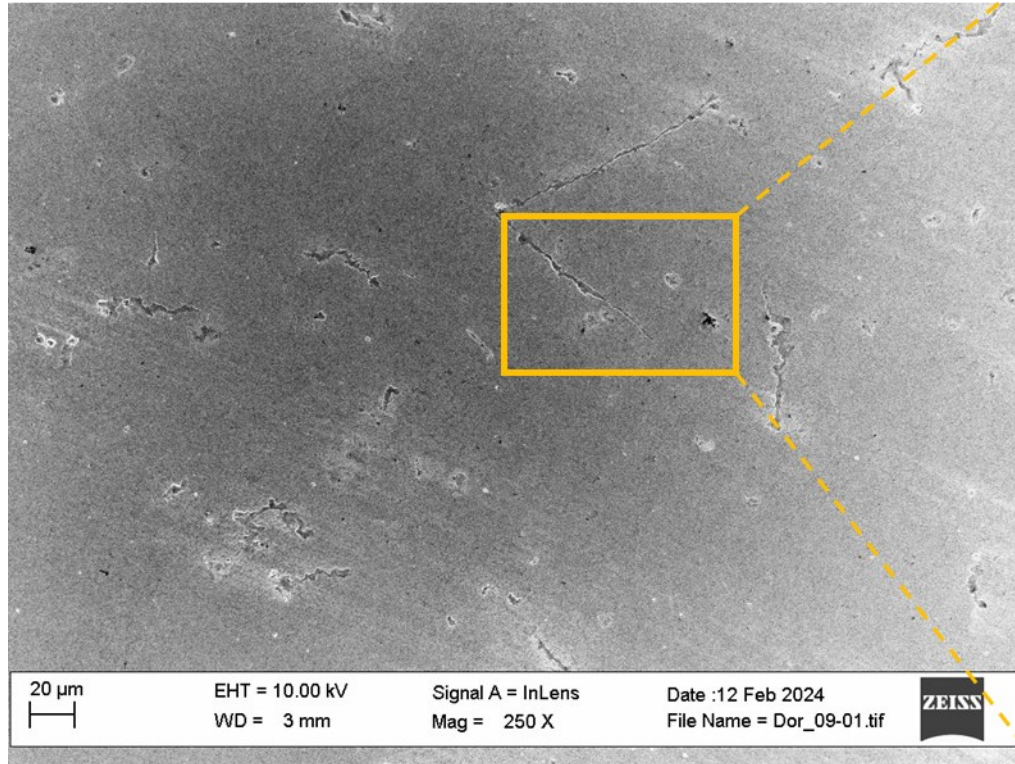
Barrier or accommodation effect?



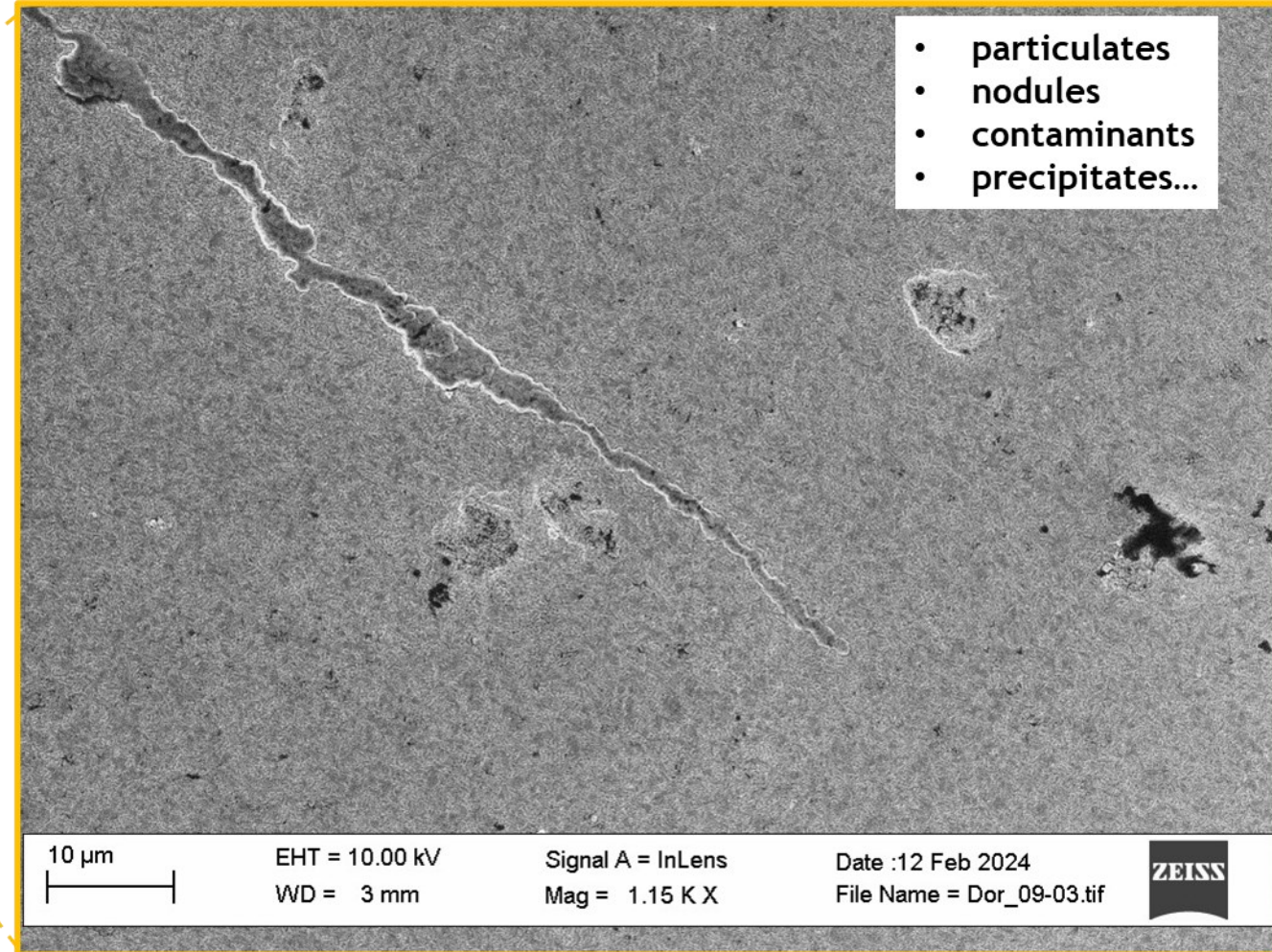
SC coatings



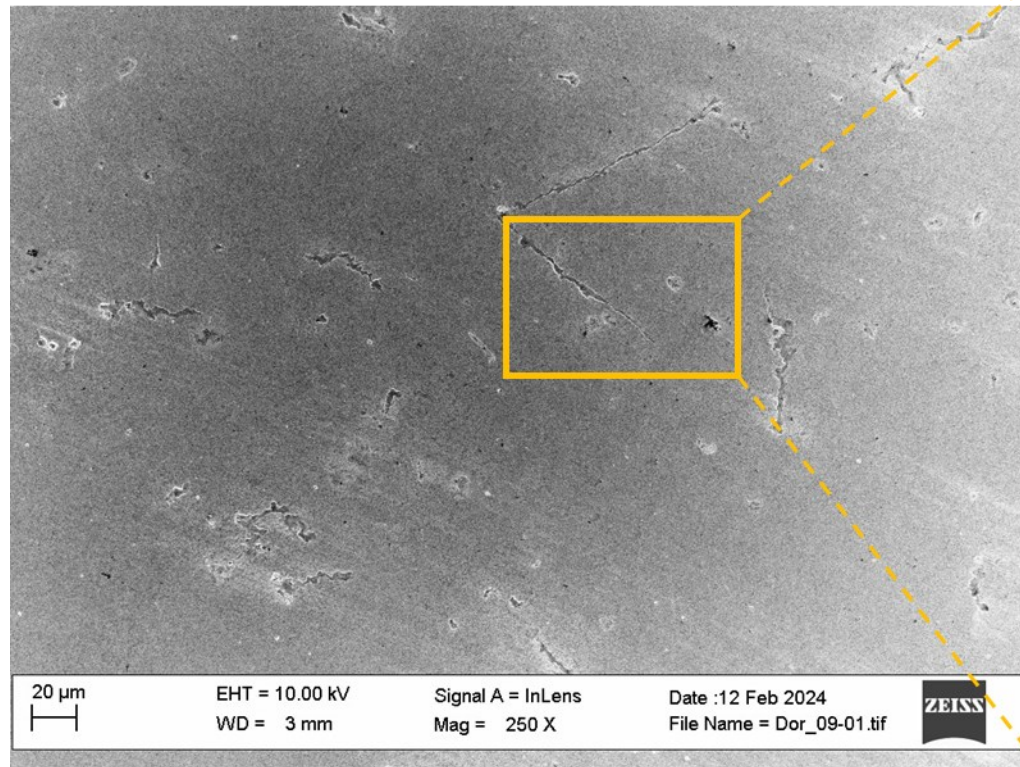
Nb₃Sn coatings



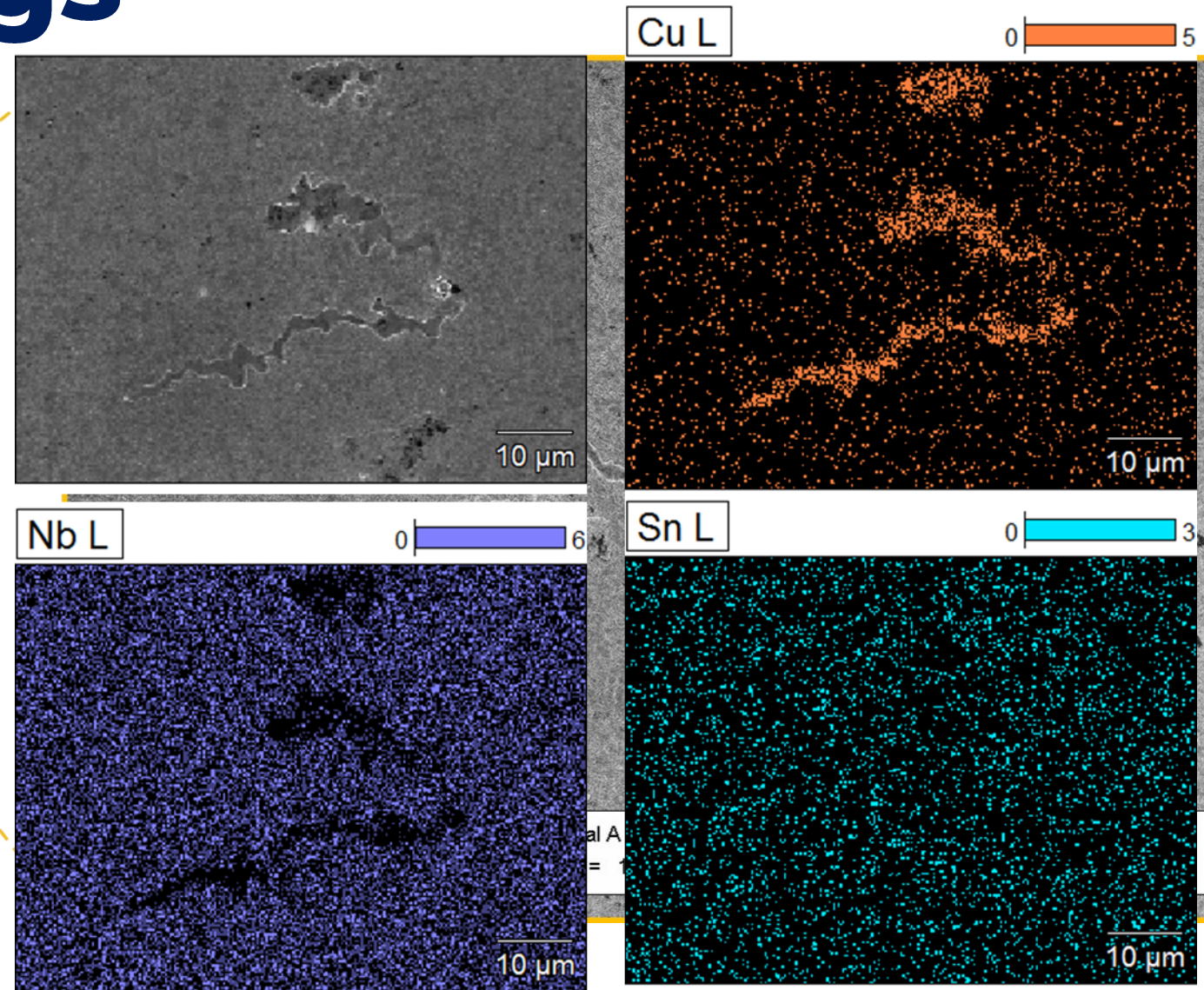
Nb₃Sn on 1 μm Nb buffer Layer



Nb₃Sn coatings



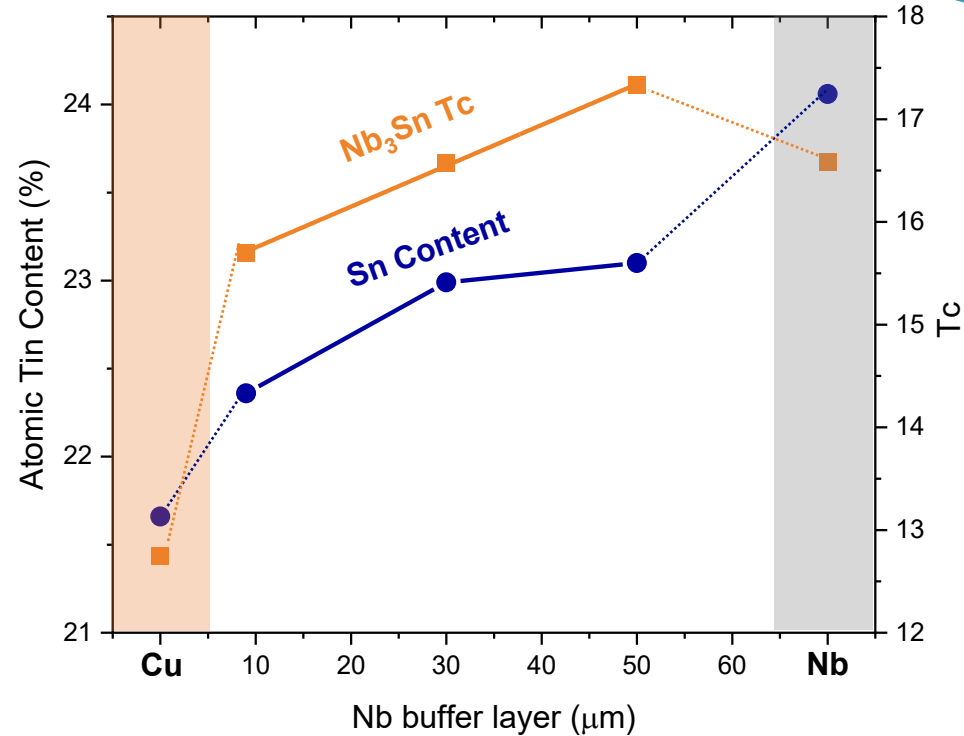
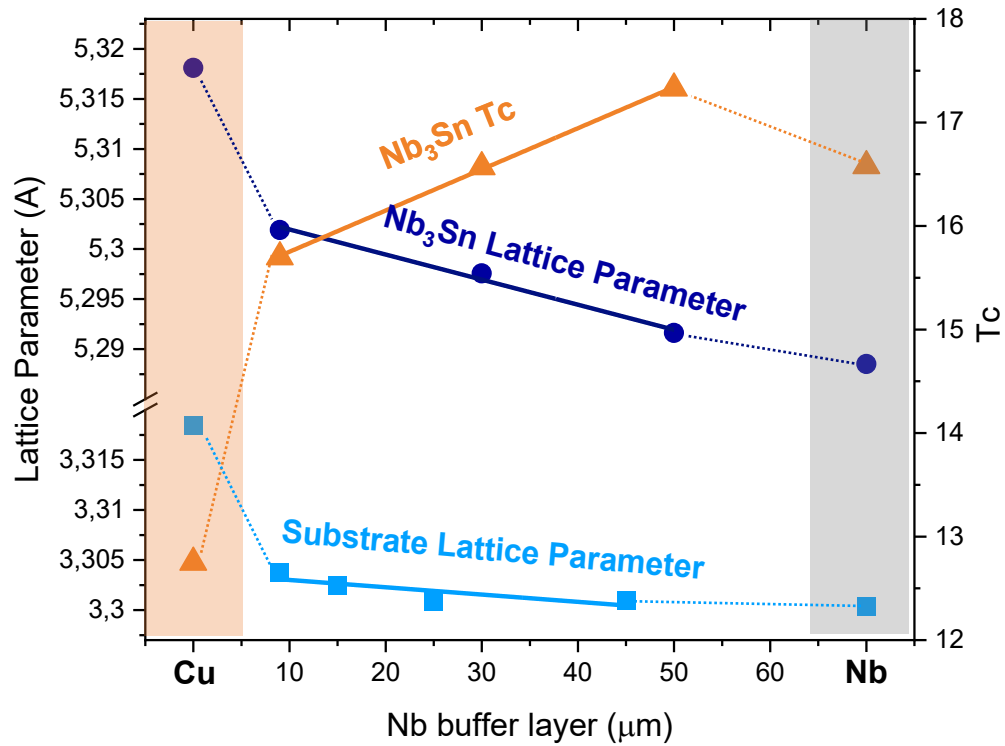
Nb₃Sn on 1 μm Nb buffer Layer



Nb₃Sn coatings

Sputtering parameter optimization

PRELIMINARY RESULTS



The role of the thick Nb layer is to accommodate the Nb₃Sn lattice parameter



Nb₃Sn coatings

Path to final prototype

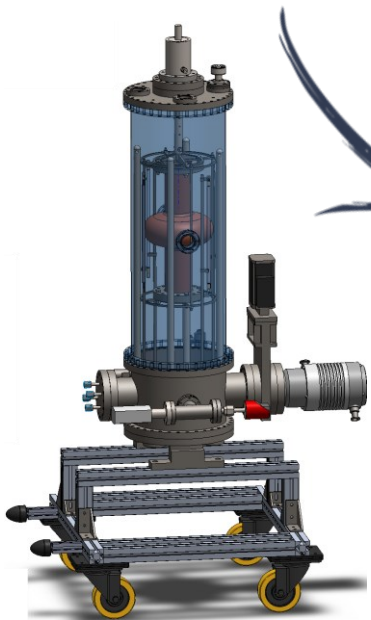
Nb₃Sn on bulk Nb to validate coating performances

Develop Nb thick coating on 1.3 GHz Elliptical Cavities

Nb₃Sn on Cu with thick Nb coating on 1.3 GHz Elliptical Cavities



RF test on QPR ongoing @ HZB



1.3 GHz coating system

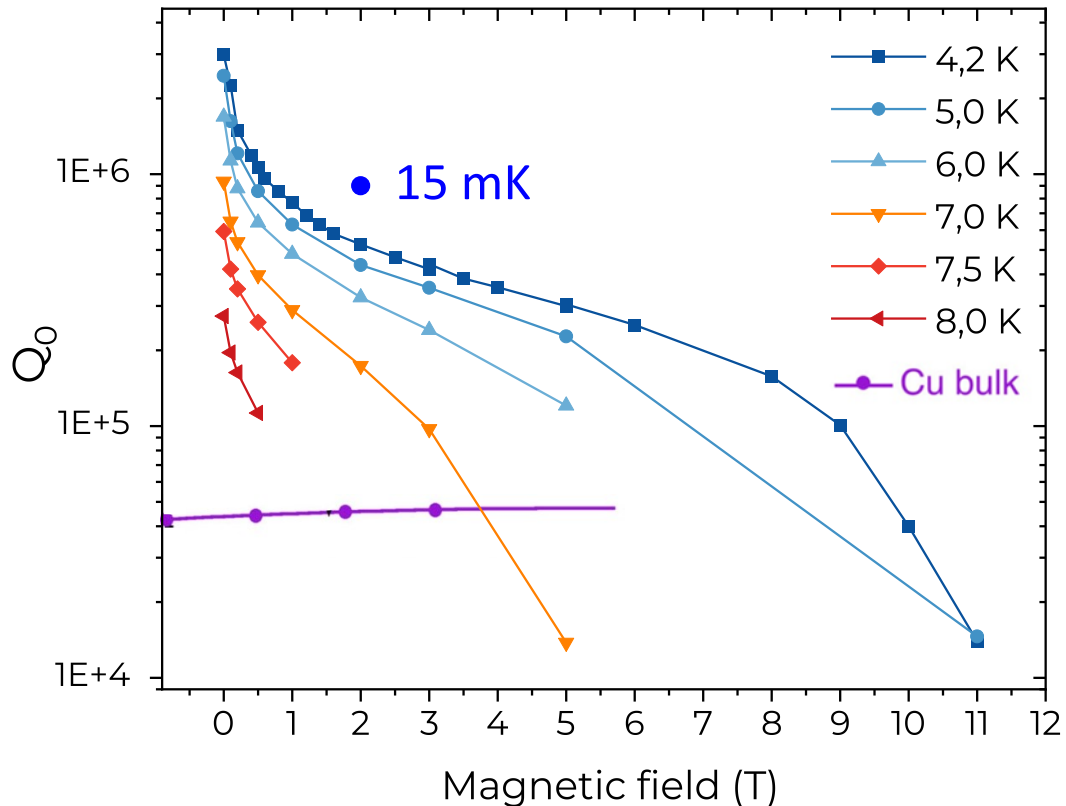


SC coatings



Quantum Sensing

NbTi Hybrid Haloscope



- ▶ $Q_0 \approx Q_0$ simulation \rightarrow limited by Cu cones
- ▶ $Q_0 \text{ NbTi} > 20 * Q_0 \text{ Cu}$ @ 15 mK, 2T
- ▶ $Q_0 \text{ NbTi} > 10 * Q_0 \text{ Cu}$ @ 4K, 5T
- ▶ $Q_0 \text{ NbTi} > Q_0 \text{ Cu}$ up to 9 T @ 4K

Next Step: Nb_3Sn

