

In Memoriam, a Colleague and Beloved Friend

Sydney Benjamin Galès

November 1, 1943 – November 29, 2024



Corrine and Sydney Galès, Jaffa, Tel Aviv, April 27, 2019

Nuclear Astrophysics With TPCs and Gamma-Beams *

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<http://Astro.uconn.edu>

**Sheffield
Hallam
University**



1. Oxygen Formation in Stellar Helium Burning/ the $^{12}\text{C}(\alpha,\gamma)$ Reaction
2. Status of World Best Data (**Stuttgart's Heroic Effort**)
3. UConn Measurement, Optical Readout TPC (O-TPC @ HI γ S)
4. Present/ Future: The Warsaw electronic readout (eTPC @ HI γ S)

* Supported in part by the USDOE grant No. DE-FG02-94ER40870.

VII Workshop..., Bormio, February 2-9, 2025

Nuclear Astrophysics in the Era of Windows on the Universe

Multi-Messenger Astrophysics (WoU-MMA)

SN1987A: First MMA, **Type II Supernova**

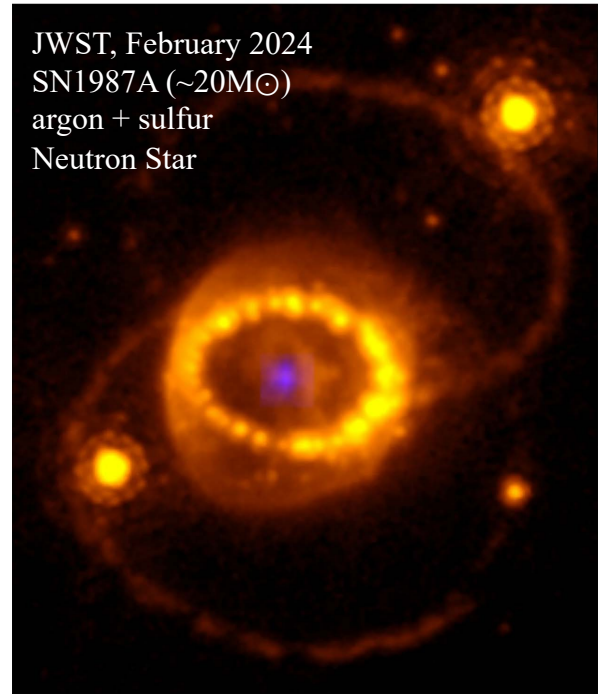
Observed Neutrinos & 4 HR Later Light Curve (EM)/ MMA object

Progenitor: Sanduleak -69 202 (Sk -69 202) **Blue Supergiant** $\sim 20M_{\odot}$

SN1987A (JWST 2024): Neutron Star, Not Black Hole

Type II SN: Neutron Star or Black Hole, Determined by C/O

JWST, February 2024
SN1987A ($\sim 20M_{\odot}$)
argon + sulfur
Neutron Star



Helium Burning: $3\alpha \rightarrow {}^{12}\text{C}$ ($\sim 11\%$) **“Hoyle State”**

${}^{12}\text{C}(\alpha, \gamma){}^{16}\text{O}$ @300 keV ???

${}^{12}\text{C}(\alpha, \gamma) \rightarrow \text{C/O} = ?$

Two partial waves:

p-wave $S_{E1}(300)$

d-wave $S_{E2}(300)$

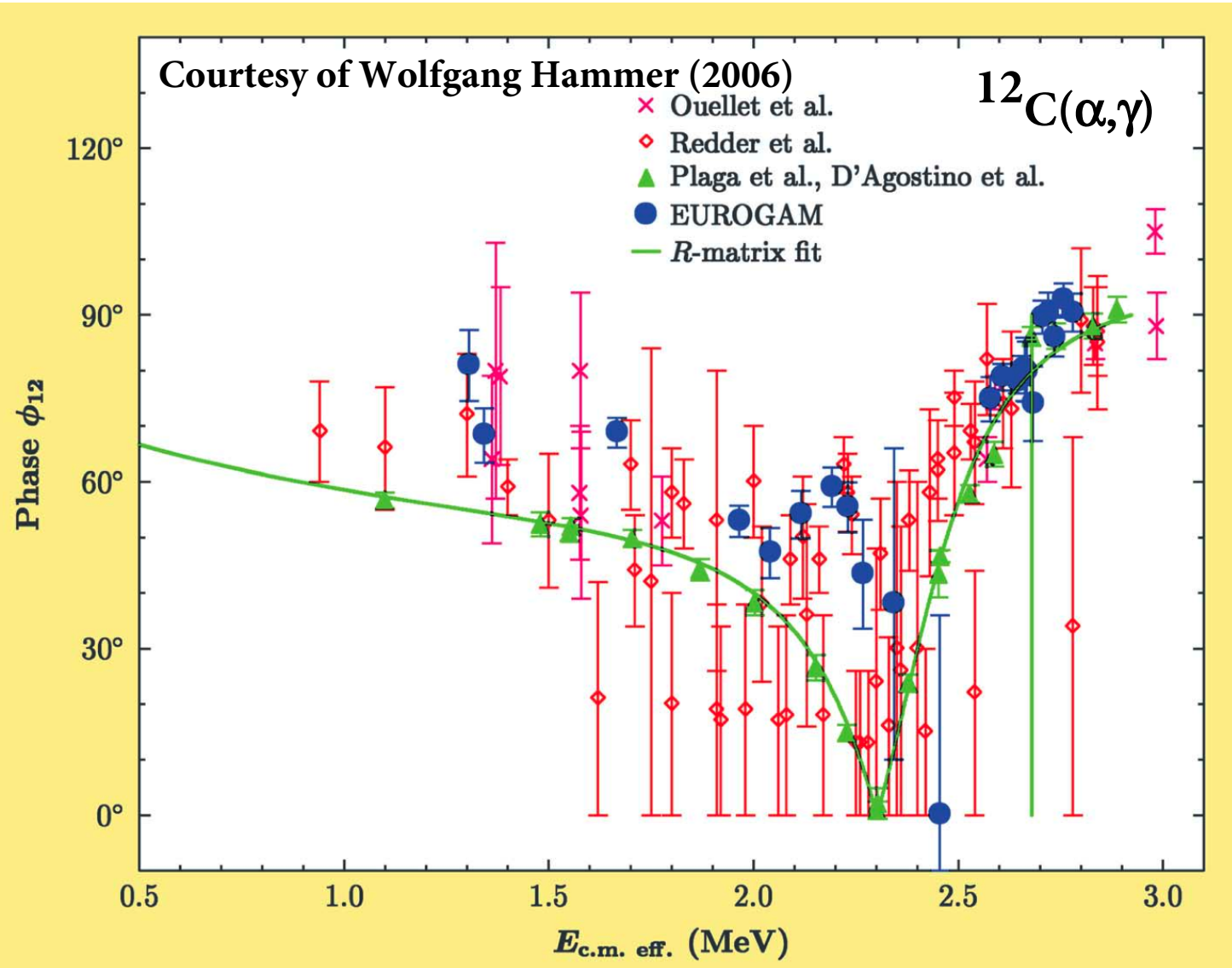
$E1$ - $E2$ Mixing Phase Angle (ϕ_{12})



W.A. Fowler: Rev. Mod. Phys. 56, 149 (1984)
“The ${}^{12}\text{C}(\alpha, \gamma)$ reaction is of paramount importance”

$$\phi_{12} = \delta_2 - \delta_1 + \arctan(\eta/2)$$

F.C. Barker and T. Kajino, Aust. J. Phys. 44, 369 (1991), R-Matrix Theory.



E1-E2 Mixing Phase Angle (ϕ_{12})

M. Gai, Phys. Rev. C 88, 062801(R) (2013).

C. R. Brune, Phys. Rev. C 64, 055803 (2001).

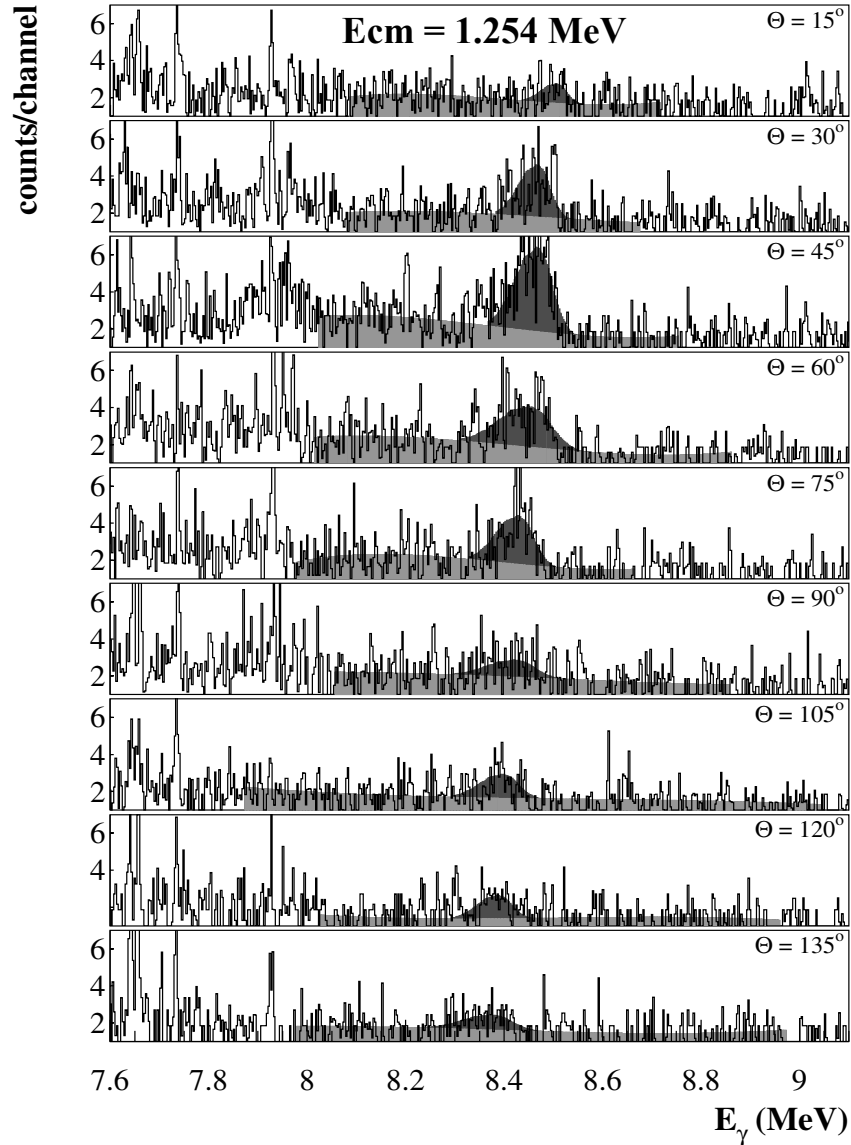
L.D. Knutson, Phys. Rev. C 59, 2152 (1999).

K.M. Watson, Phys. Rev. 95, 228 (1954).

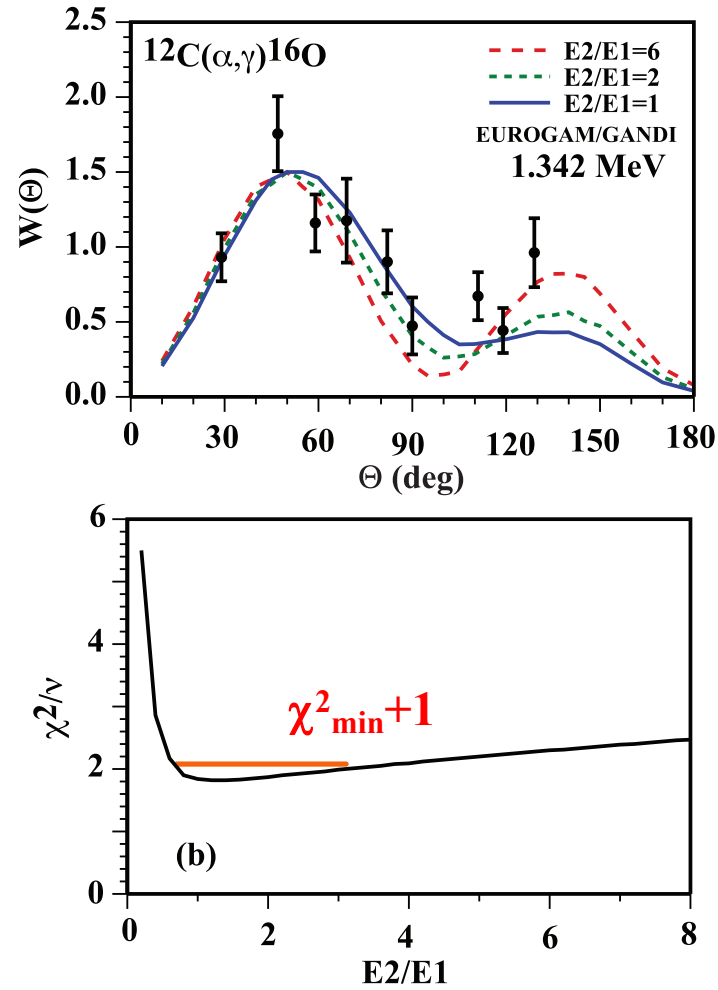
Required by Unitarity

**The (heroic) Stuttgart Effort: 1) 450 μ A 2) 700 HRs 3) Four x 100% HPGe
4) EUROGAM 5) 0.01% ^{13}C [x100 Reduced $^{13}\text{C}(\alpha,n)$]**

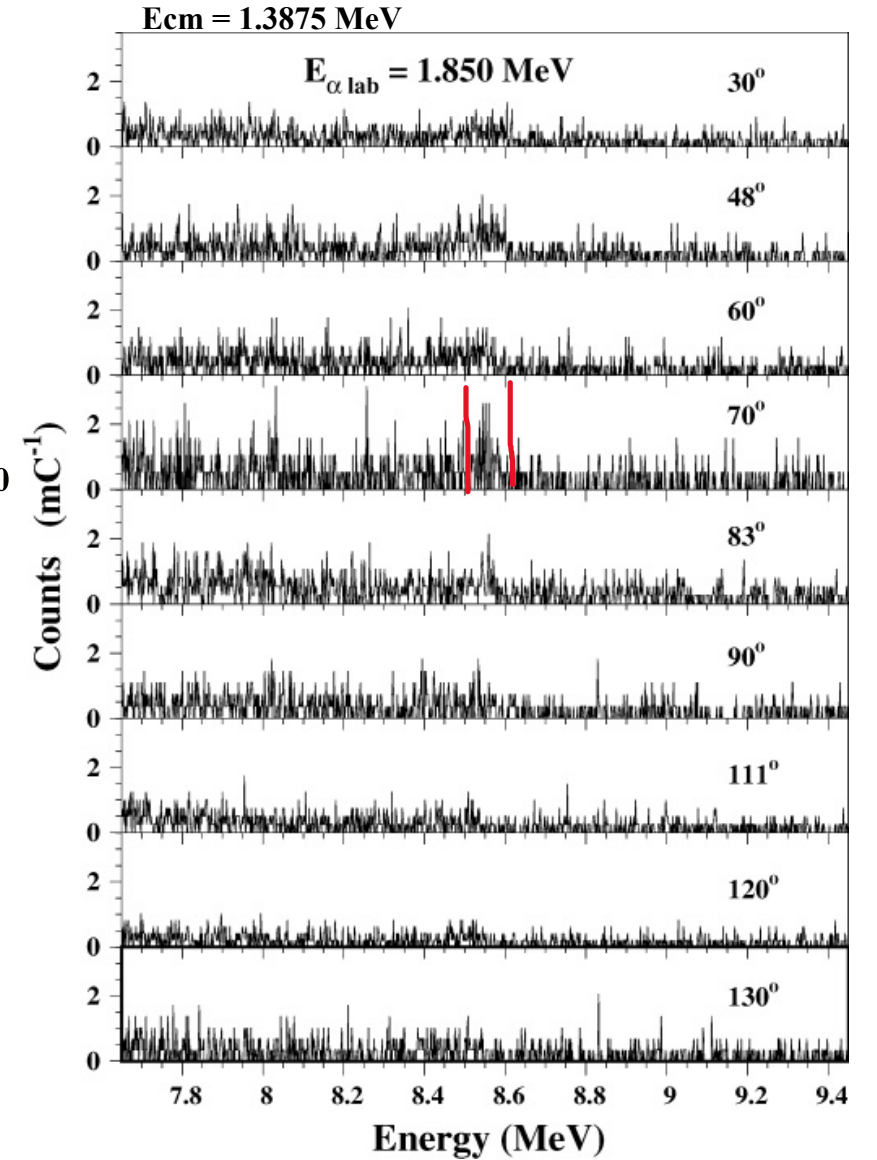
R. Kunz *et al.*, PRL, 86, 3244 (2002)



M. Gai, PRC, 88, 062801(R) (2013)



M. Assuncao *et al.*, PRC 73, 055801 (2006)



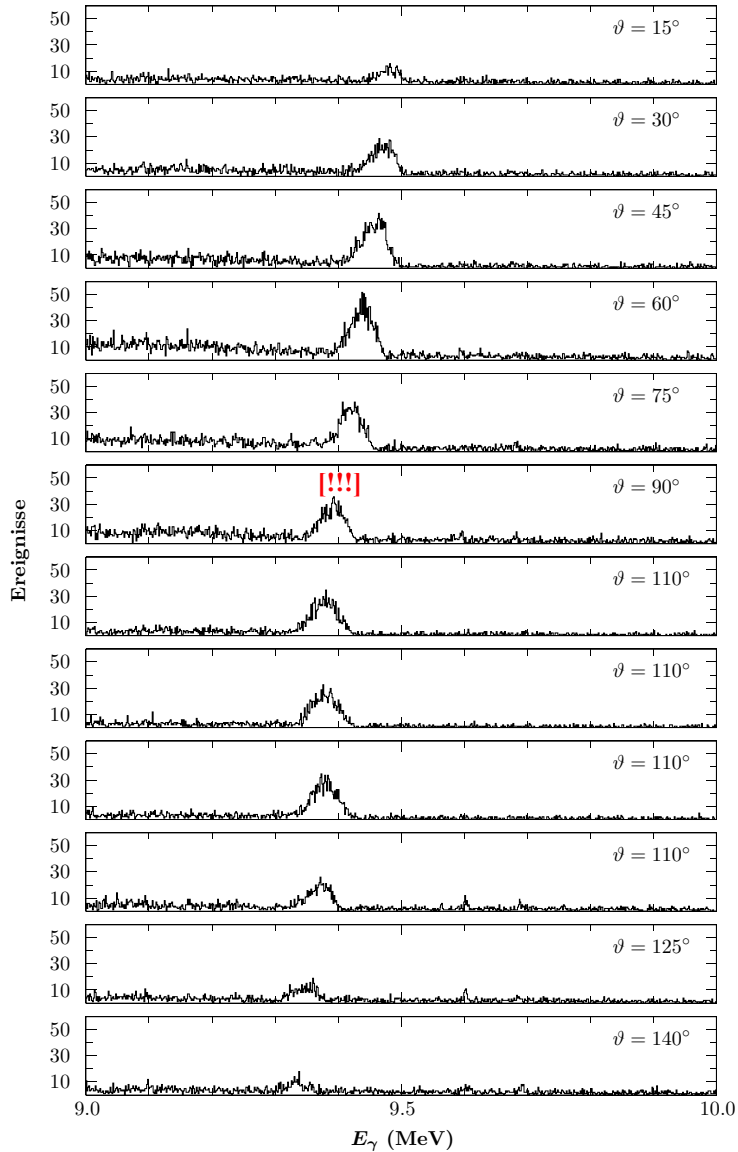


Abbildung C.35: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 2.209 \text{ MeV}$. $E_I = 2.945 \text{ MeV}$

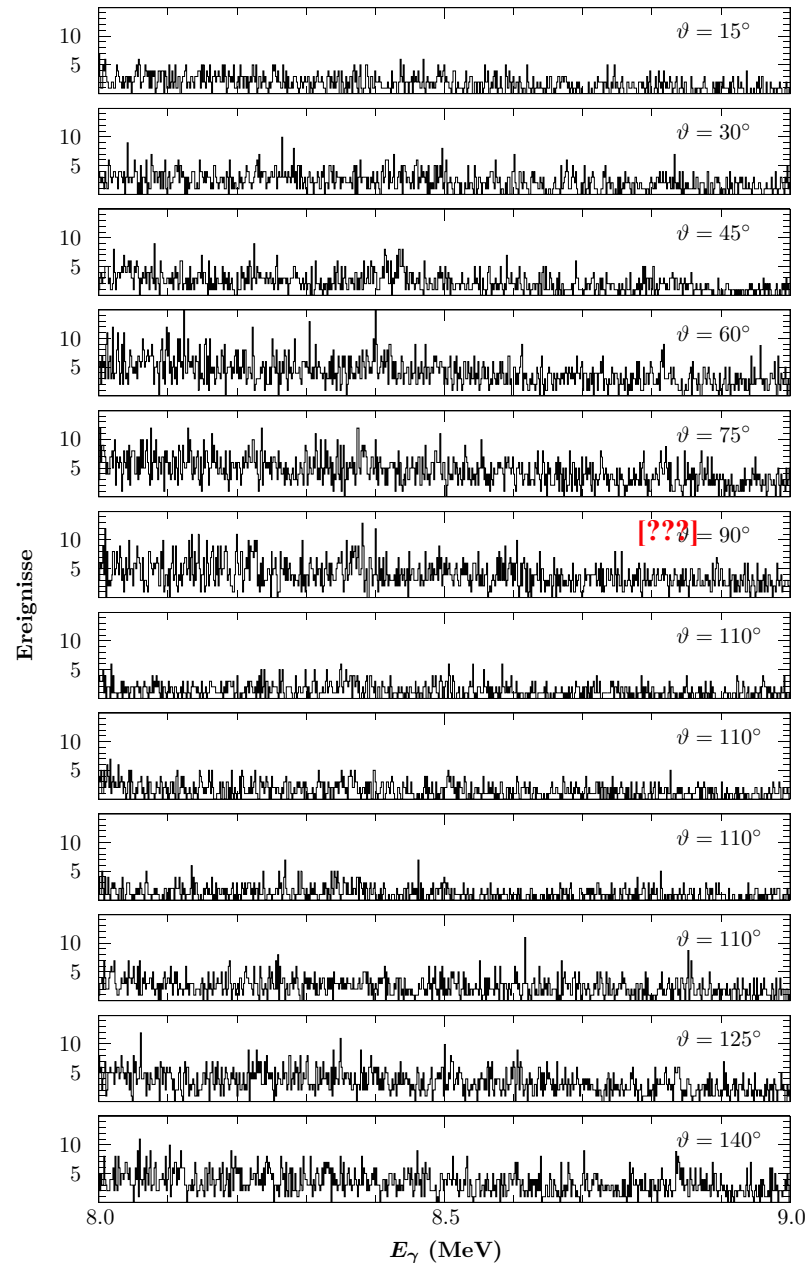


Abbildung C.34: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 1.696 \text{ MeV}$. $E_I = 2.261 \text{ MeV}$

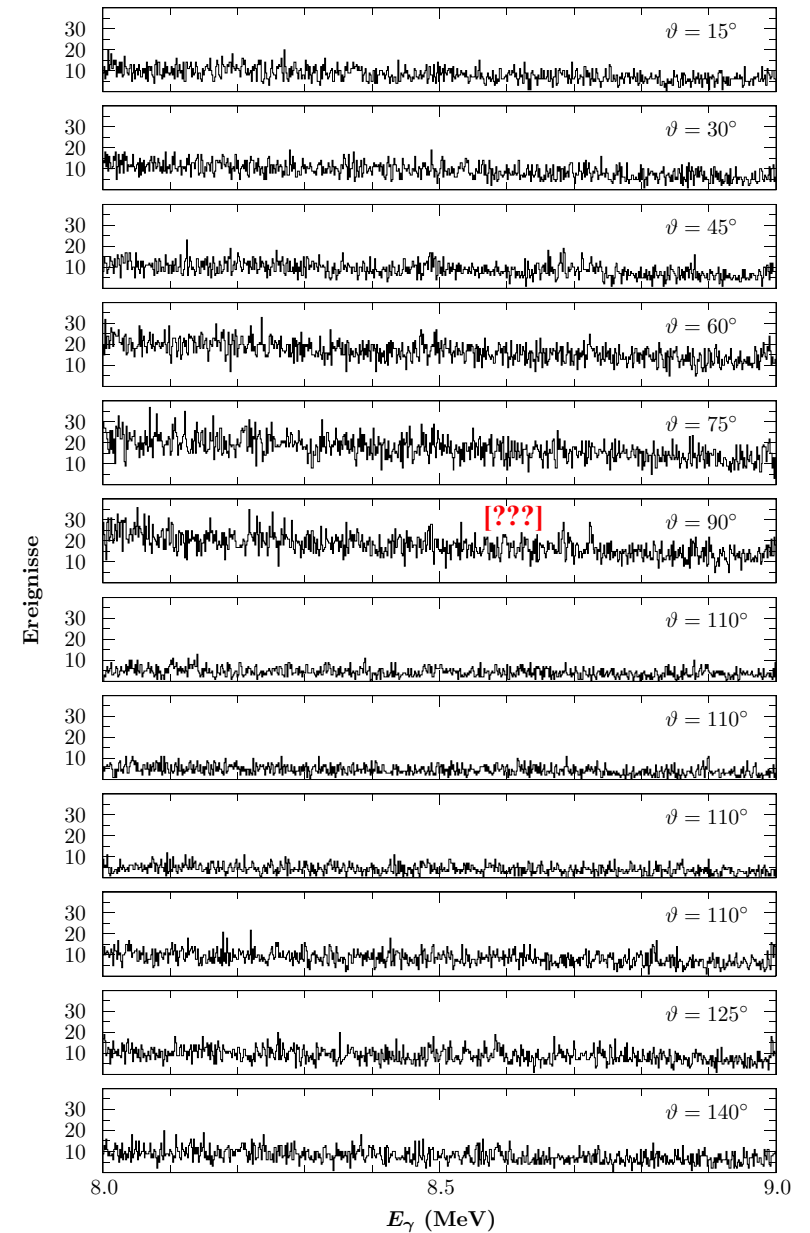


Abbildung C.33: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 1.452 \text{ MeV}$. $E_I = 1.936 \text{ MeV}$

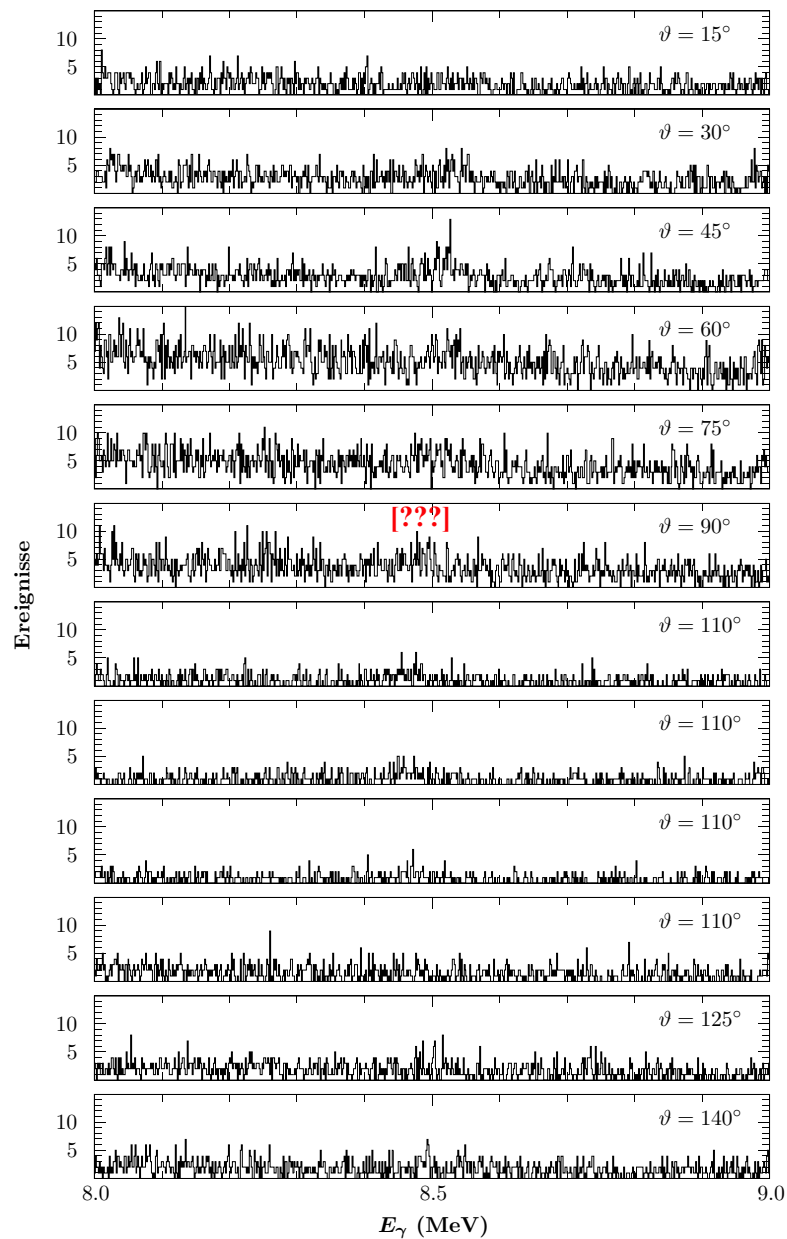


Abbildung C.32: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 1.308$ MeV. $E_L = 1.744$ MeV

C.2. DREHTISCH-EXPERIMENT

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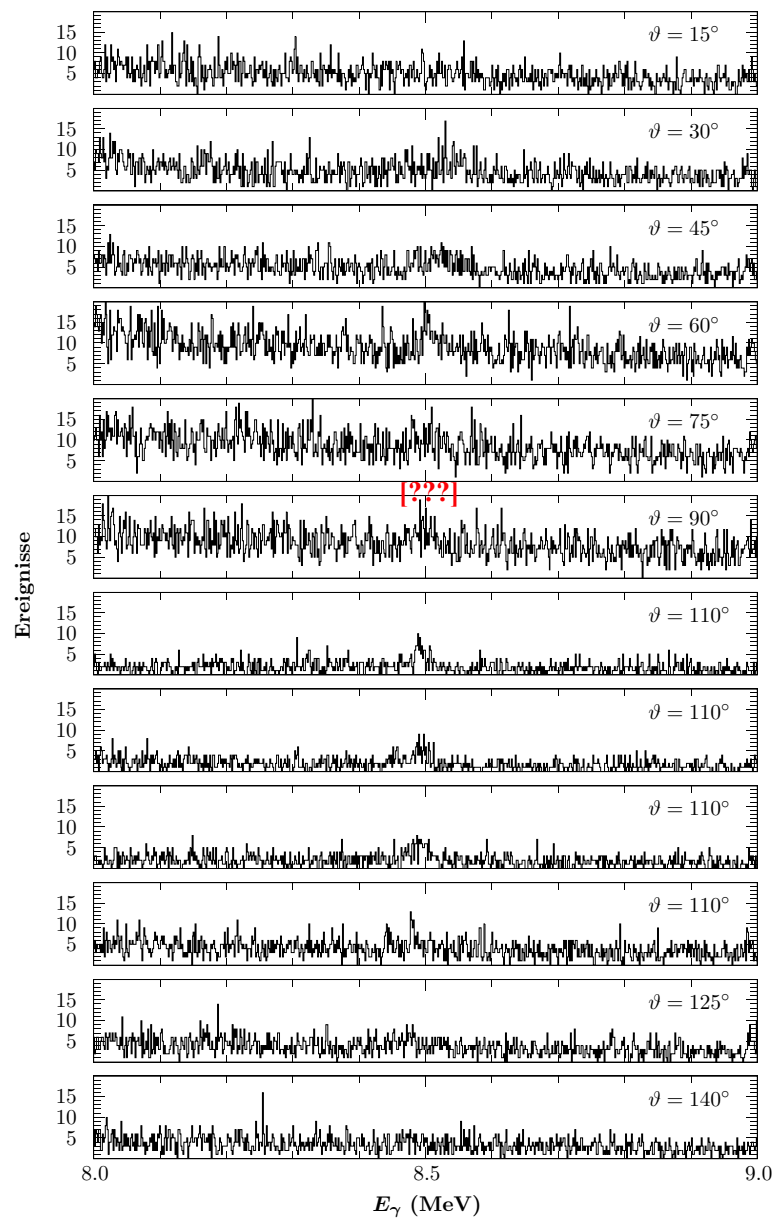


Abbildung C.31: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 1.305$ MeV. $E_L = 1.740$ MeV

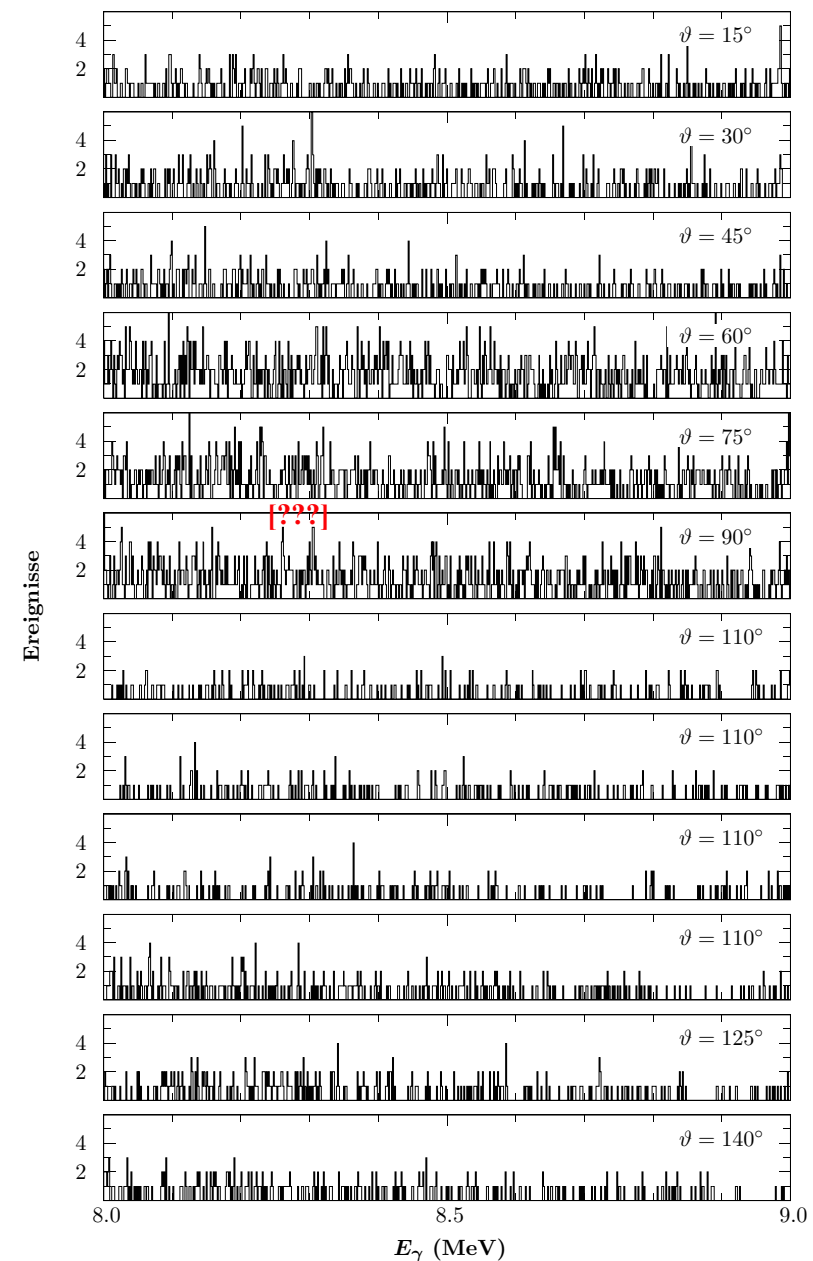


Abbildung C.30: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 1.103$ MeV. $E_L = 1.470$ MeV

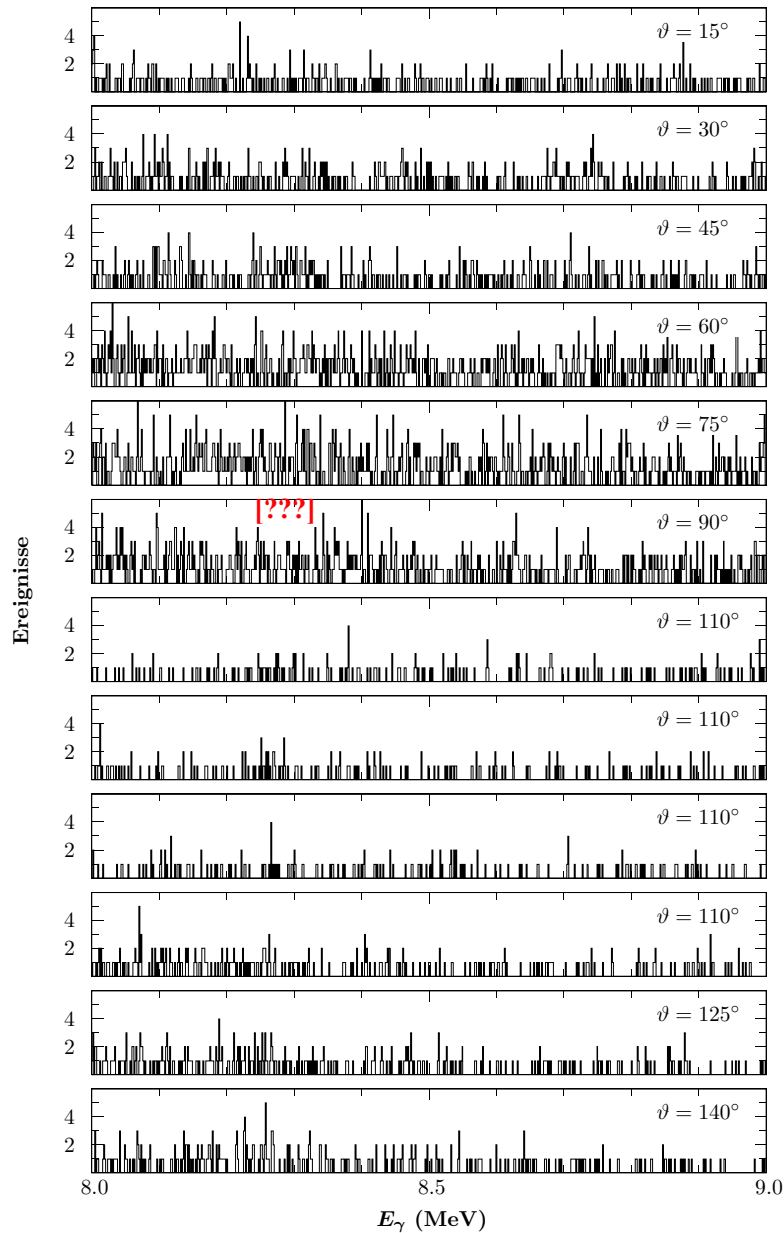


Abbildung C.29: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 1.102 \text{ MeV}$. $E_1 = 1.469 \text{ MeV}$

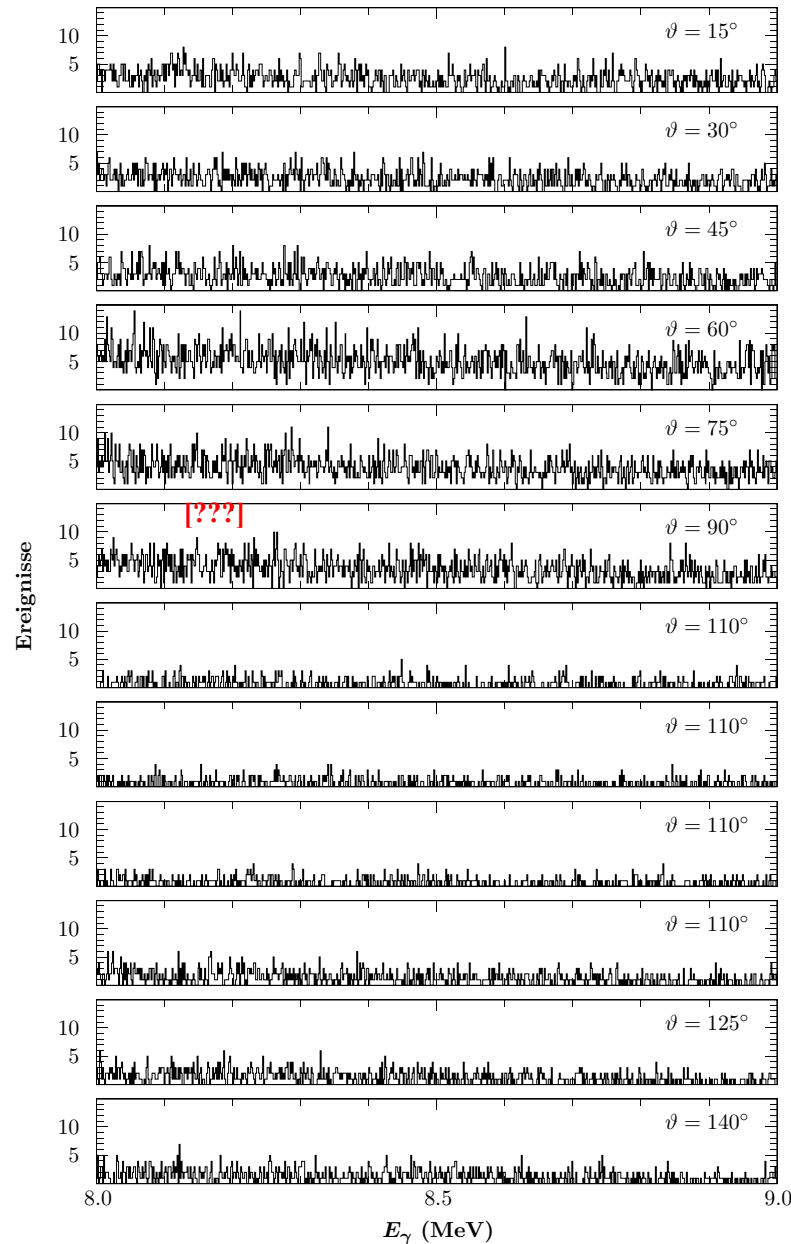


Abbildung C.28: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 1.099 \text{ MeV}$. $E_1 = 1.465 \text{ MeV}$

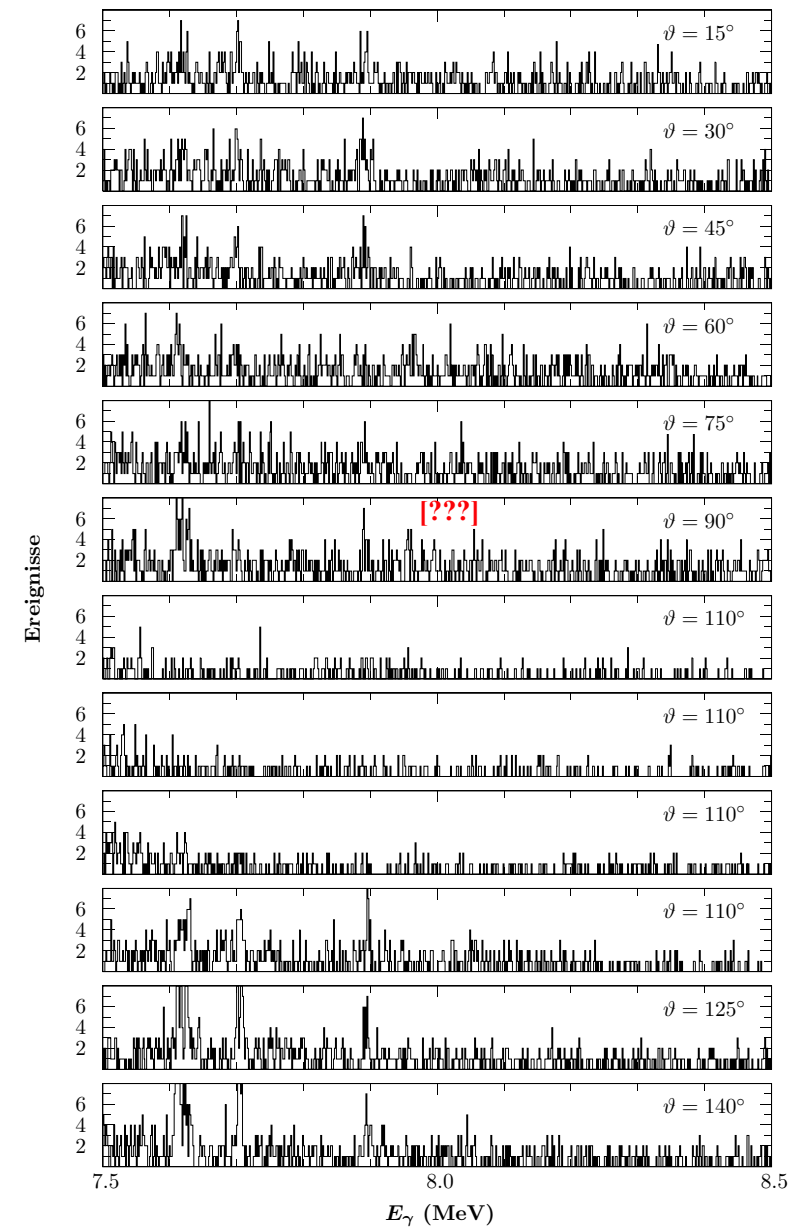
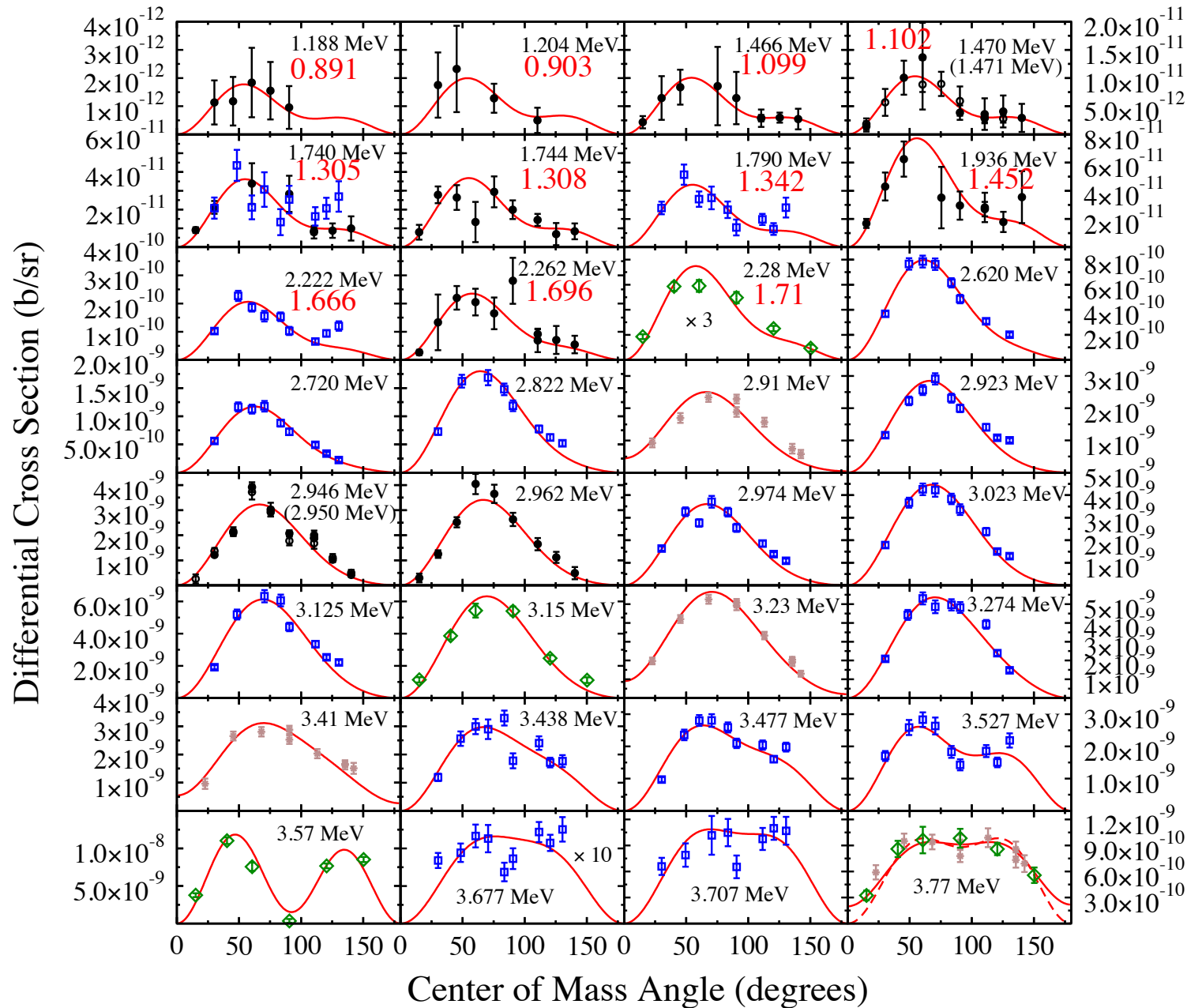


Abbildung C.27: Im Rahmen des Drehtisch-Experiments gemessene γ -Roh-Spektren bei $E_{c.m.} = 0.903 \text{ MeV}$. $E_1 = 1.204 \text{ MeV}$



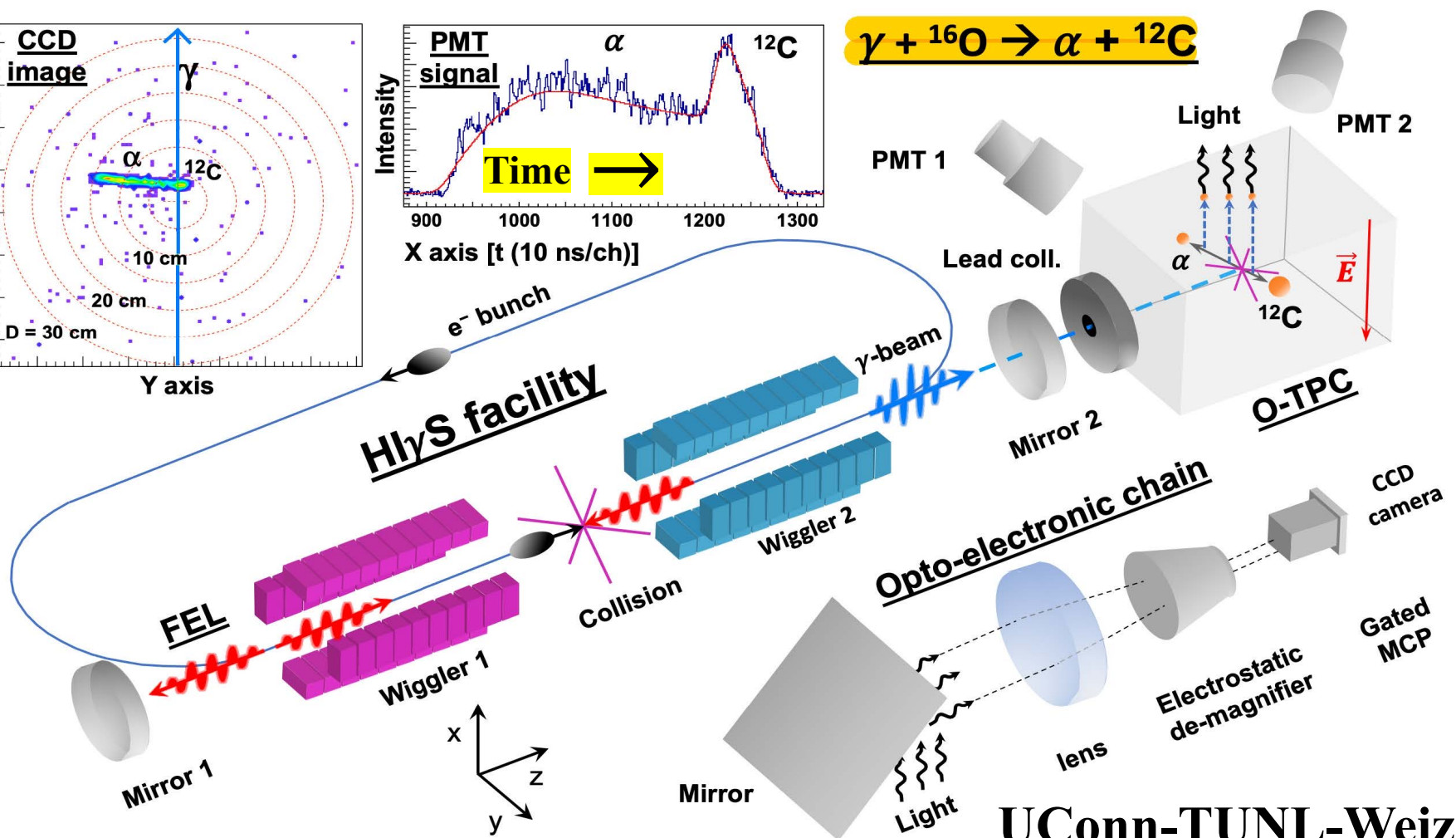
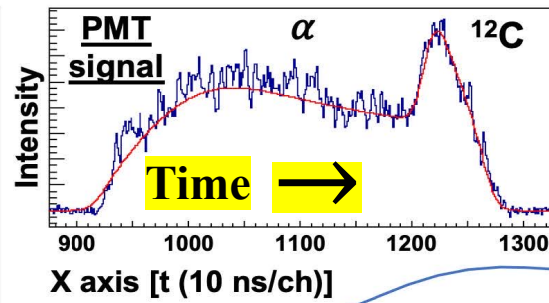
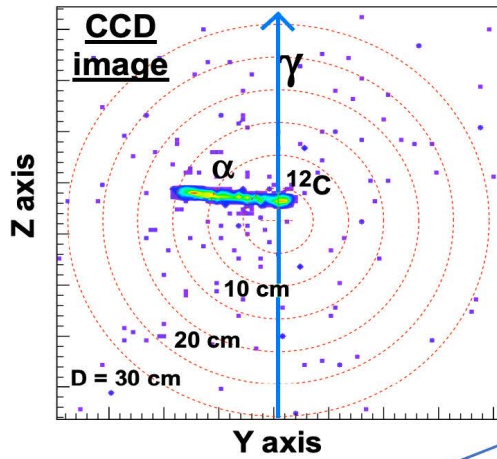
**deBoer *et al.* rely on the ANC
The S-factor is derived from
Alpha-transfer, e.g. (${}^7\text{Li},t$)
Not from capture gamma-ray
Indirect Method ala 1980's**

**51 years after Dyer & Barnes
We still do not have the needed
Capture Gamma-Ray Data**

**This is the status of our field
(NOT a personal criticism)**

Detailed Balance: $^{12}\text{C} + \alpha \rightarrow ^{16}\text{O} + \gamma$ (in Stars)
 $^{16}\text{O} + \gamma \rightarrow ^{12}\text{C} + \alpha$ (at HI γ S)

(O-TPC: CO₂)



Active Target TPC
M. Gai et al,
JINST 5, 12004 (2010)

UConn-TUNL-Weizmann-PTB (2012)

R. Smith, M. Gai, D.K. Schweitzer, S.R. Stern and M.W. Ahmed,
Nature Communications, 12, 5920 (2021).
<https://www.nature.com/articles/s41467-021-26179-x>

SHU-UConn-TUNL (2021)

Detailed Balance:

(Inverse, Time Reversed Reaction)

$$\sigma[{}^{12}\text{C}(\alpha, \gamma){}^{16}\text{O}] = \frac{2 k_{\gamma}^2}{k_{\alpha}^2} \sigma[{}^{16}\text{O}(\gamma, \alpha){}^{12}\text{C}] *$$

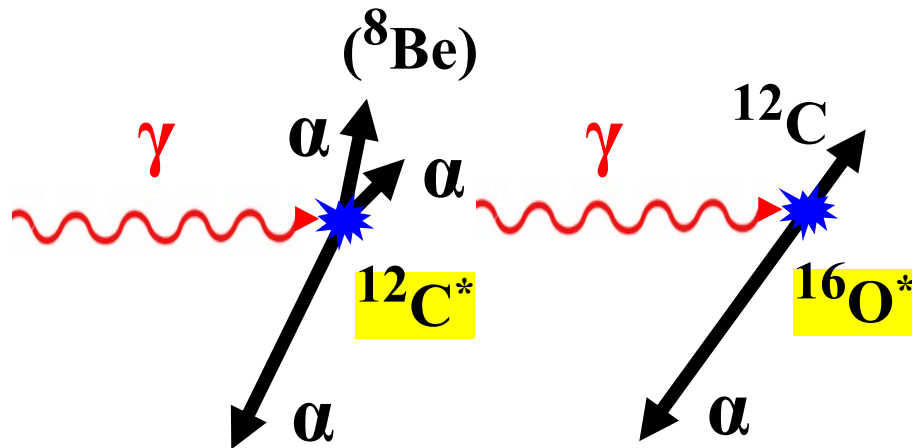
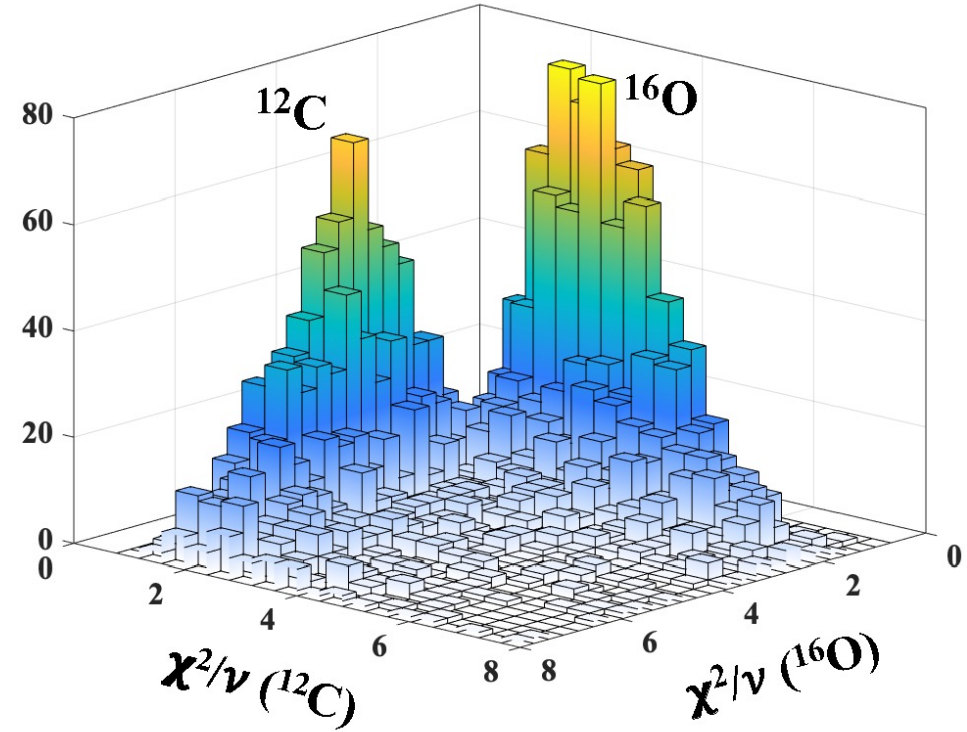
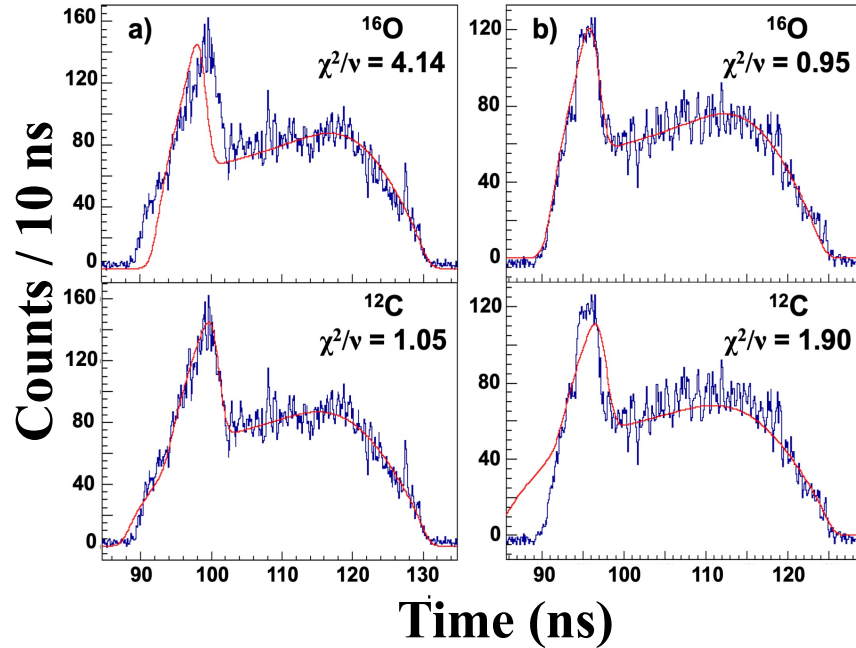
$$\sigma[{}^{16}\text{O}(\gamma, \alpha){}^{12}\text{C}] \approx \sim 50 \times \sigma[{}^{12}\text{C}(\alpha, \gamma){}^{16}\text{O}]$$

* For Real Photons $2S+1 = 2$ (not 3)

Not a “Surrogate Reaction”

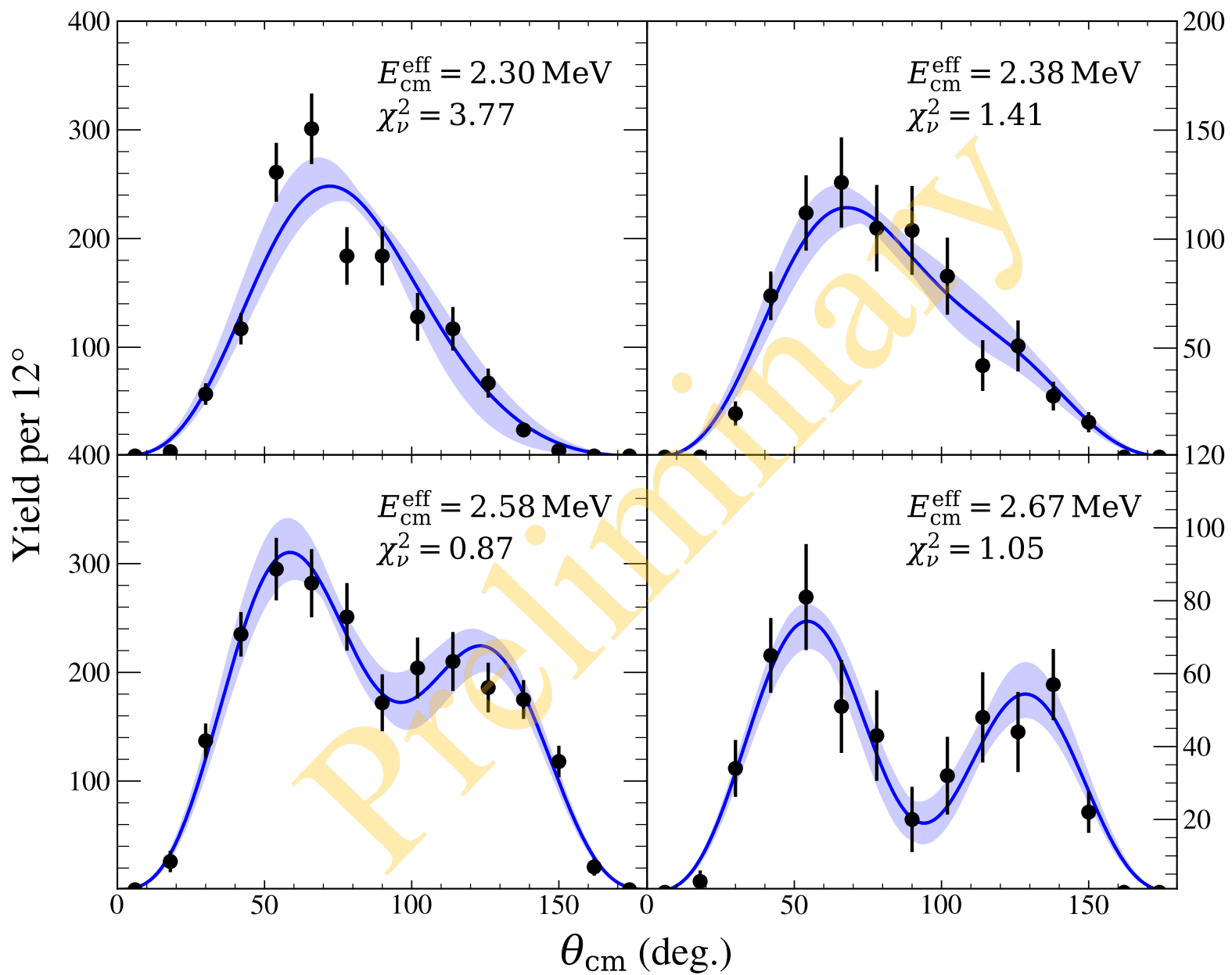
Not an Indirect Measurement

Line Shape Analysis (CO₂ Gas)



Machine Learning

$$Q(^{16}\text{O}^*) - Q(^{12}\text{C}^*) = 112 \text{ keV}$$

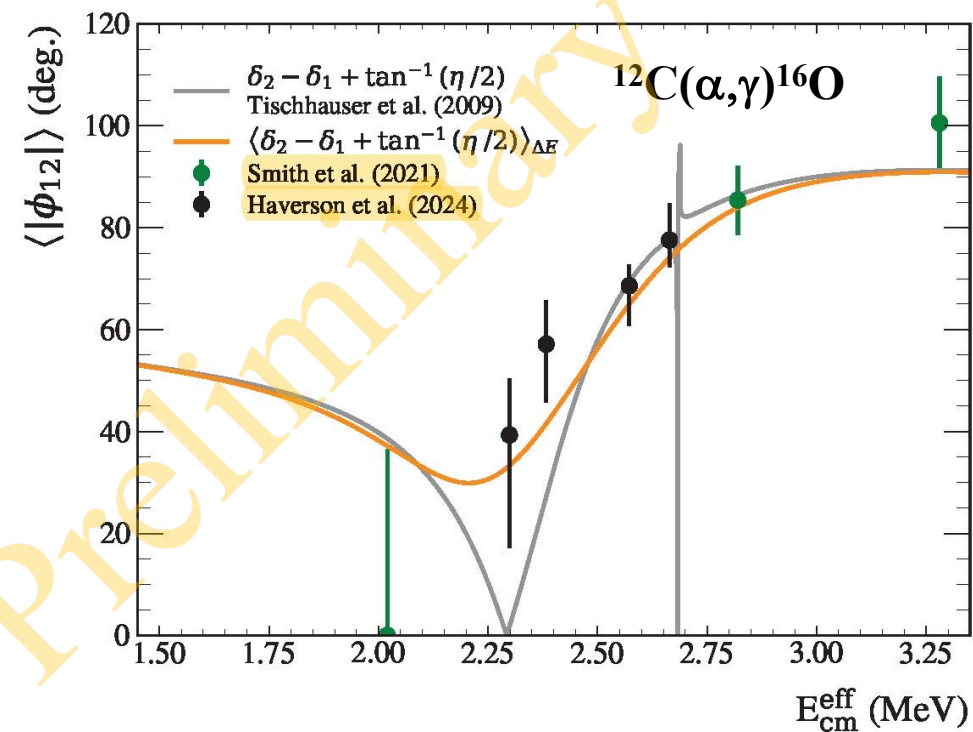
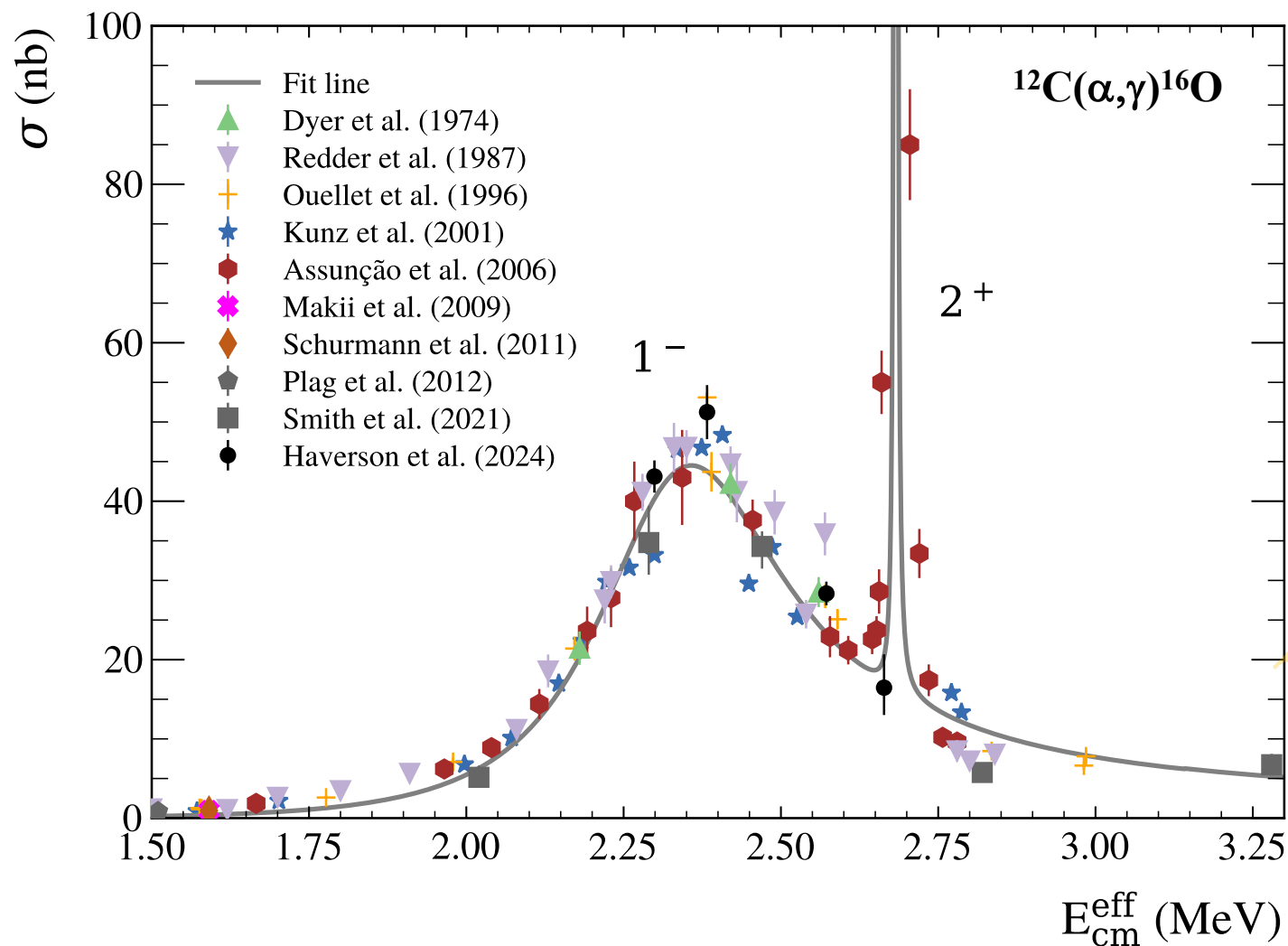


**O-TPC Data N_2O gas
Angular distributions
measured at 17 angles**

**Kristian C.Z. Haverson
@ SHU, UConn-SHU (2024)**

O-TPC (Nature + N₂O) Data Benchmarked against World Data

First Agreement of data on ϕ_{12} with Quantum Mechanics



UConn O-TPC @ HI γ S (2012)



Warsaw eTPC @ HI γ S (2022)



Looking forward to ELI-NP, Bucharest/ M. Gai Fulbright US Scholar, 2025

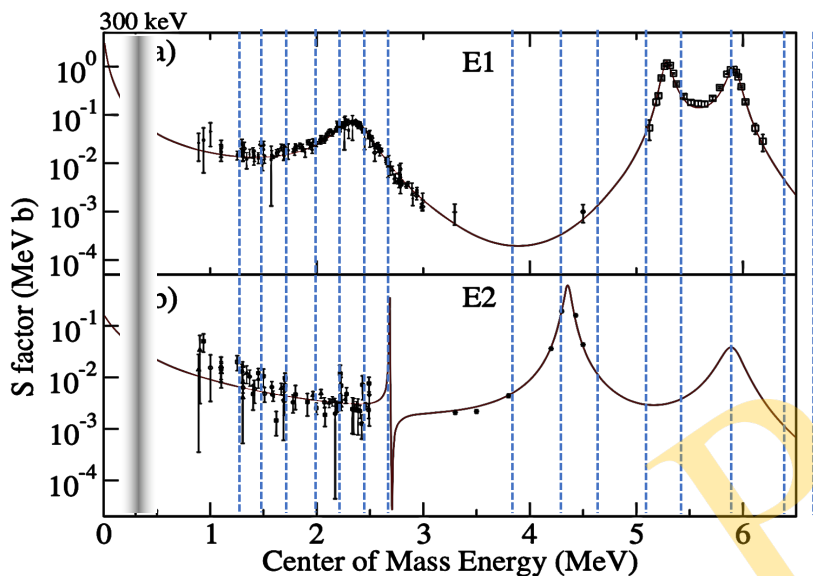
Sneak peek, 2022 Measurement with the Warsaw eTPC:

Mateusz Fila, Ph.D. Thesis, Warsaw 2023

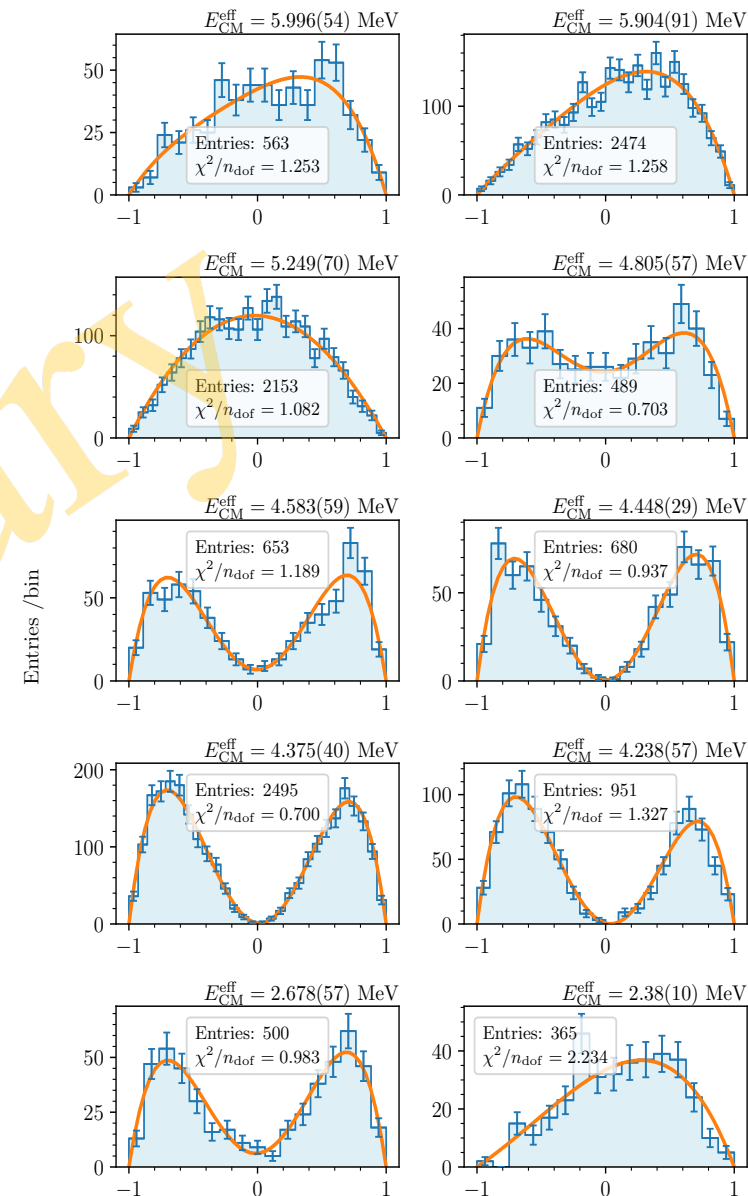
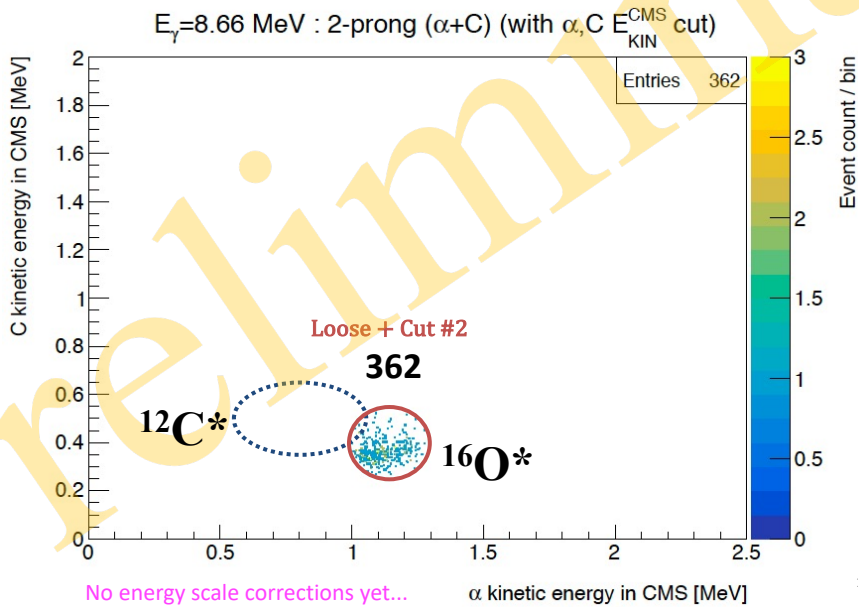
<https://repozytorium.uw.edu.pl/bitstreams/5049fd46-45eb-4b32-9edb-fa121cba7812/download>

Complete Angular Distributions, 20 Angular Bins down to ~ 1.3 MeV

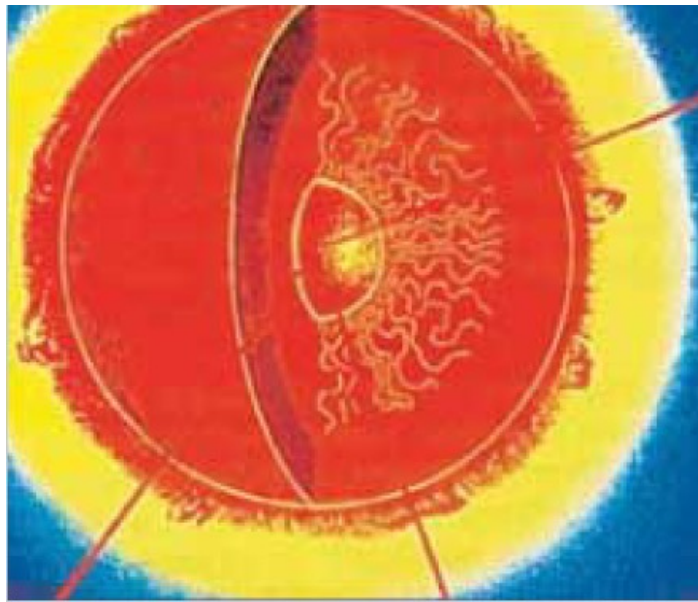
We measured at the shown 15 Energies
(Curve E1&E2 from deBoer *et al.*)



Exploratory Measurement:
Mikołaj Ćwiok, $E_\gamma = 8.66$ MeV
 $E_{cm} = 1.5$ MeV



α track $\cos \theta_{BEAM}$



University of Connecticut
Laboratory for Nuclear Science
at Avery Point
(AKA Laboratory for Astrophysics)

Conclusions

TPC data of unprecedented quality:

1. **Low background, if any**
 2. **Measurement in one detector (response, simple Monte Carlo)**
 3. **Complete angular distribution ($0^\circ - 180^\circ$)**
(Measured at 17-20 bin-angles)
 4. **First Physics Result, First Agreement with Unitarity**
 5. **New Criteria for Judging Data (Agreement with QM)**
 6. **Further data measured at HIγS, Warsaw TPC, 2022**
- (Please do not publish or analyze data that disagree with QM)**