

Center for Axion and Precision Physics Research



Advancements on CAPP's Axion Research

Woohyun Chung Center for Axion and Precision Physics Research (CAPP) Institute for Basic Science (IBS)

19th Patras Workshop on Axions, WIMPs and WISPs



IBS-CAPP in 10 years



- Last 10 years, CAPP has grown to be one of the best axion search facilities in the world
 - CAPP-MAX: the most sensitive axion experiment (when NT is right...)
- Whole axion community has grown much more than before
- However, there are "WIDE" range of frequencies waiting to be explored. If you want to do
 it in your lifetime, you need serious R&Ds (Improvements/Innovations)
 - Powerful, big bore Magnet
 - Quantum noise-limited JPA / Single Photon Detector
 - HTS superconducting cavities
 - Higher frequency w/o sacrificing volume
- CAPP is phasing out at the end of 2024
 - We may continue in other form, which we don't know yet



September 19th 2024

19th PATRAS Workshop at Patras, Greece Woohyun Chung







CAPP Experimental Hall (LVP)





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Status of CULTASK at CAPP



| Refrigerators | | | | | | | S | SC Magi | Experiments | | | | |
|-------------------|----------|--------|--------------------|------|---------|------------|--------|--------------------|------------------------|---------|-----------|------------|----------------------|
| Vendor | Model | BaseT | Power | Туре | Install | B field | Bore | Mat. | Vendor | Install | Exp. | Status | Freq (GHz) |
| Leiden | DRS1000 | 10 mK | <u>1.3mW@120mK</u> | wet | 2020 | 12T | 32cm | Nb ₃ Sn | Oxford | 2020 | CAPP-MAX | Running | 1.0~1.5 |
| BlueFors (BF3) | LD400 | 10 mK | 580uW@100mK | dry | 2016 | 12T | 10cm | Nb₃Sn | AMI | 2021 | CAPP-HF | Running | 4.0~10.0 |
| BlueFors (BF4) | LD400 | 10 mK | 580uW@100mK | dry | 2016 | | | | | | CAPP-QNA | Testing | JPA test |
| BlueFors (BF5) | LD400 | 10 mK | 580uW@100mK | dry | 2017 | 8T | 12cm | NbTi | AMI | 2016 | CAPP-PACE | Running | 2.0~3.0 (5.6-5.9) |
| BlueFors (BF6) | LD400 | 10 mK | 580uW@100mK | dry | 2017 | 8T | 16.5cm | NbTi | AMI | 2017 | CAPP-8TB | Running | 1.5~2.0 (5.8-6.0) |
| Janis | He3 | 300 mK | 25uW@300mK | wet | 2017 | 9T | 12cm | NbTi | Cryo- Magnetic s | 2017 | CAPP-HF | Installing | 3.5~4.5 (~10.0) |
| Oxford | Kelvinox | 30 mK | 400uW@120mK | wet | 2017 | 18T | 7cm | GdBCO | SuNAM | 2017 | CAPP-18T | upgrading | 4.5~6.0 |







How much can you improve?



- Maximize Signal (B²VQ)
 - 12T 32cm bore LTS magnet by Oxford
 - Improve Q-factor of cavity SC cavity
- > x100 faster scan (than 8T-10cm) x10-100 faster scan than Cu

~ x50 faster scan compared to HEMTs

- Minimize Noise (T_{system} = T_{physical} (T_{amp})
 Quantum noise-limited amplifier JPA

 - Optimize cryo-RF receiver chain (SQL)
 - Single photon detector in the future (esp. for high freq.)
- Higher frequencies without sacrificing volume ullet
 - Pizza Cavity
 - Dielectric rings (TM $_{030}$ and TM $_{050}$)
 - Photonic cell cavity \bullet
 - Meta material + Superfluid LHe tuning ullet





Quantum Noise-Limited Amp (S. Uchaikin)

- Collaboration with Nakamura's group (U. of Tokyo, RIKEN)
 - > 200 JPA chips delivered (1.0 6.0) GHz (4 batches) + Packaged at CAPP
 - 1.0 GHz
 - DAQ running ~ 1.0 GHz data in CAPP-MAX (DFSZ)
 - 2.3 GHz
 - Excellent Noise Temperature (close to SQL) ~120 mK
 - Completed taking axion dark matter physics data around 2.3 GHz
 - 6.0 GHz
 - Wider bandwidth: ~ MHz (covers ~ 300 MHz of frequencies)
 - Taken 5.5-6.0 GHz physics data using 6 & 8 cell Pizza Cavity (two DRs)
 - Shows excellent NT: ~ 250 mK
 - 6 JPAs bundled to give wider coverage
- Wideband (1~15 GHz) TJWPA (INFN) and TWJPA (Lancaster)
- Single photon detector for higher frequency in progress
 - Collaboration with Aalto U., INFN and U. of Hamberg





- Quantum-limited noise
- Noise Squeezing
- T_N ≤ 167 mK @5.6 GHz
- SQL at 5.6GHz 260 mK

Split-band (Dulcimer) Amplifier (up to 6 JPAs)



Uchaikin S.V. et al. Front. Phys., v.12-2024 https://doi.org/10.3389/fphy.2024.1437680



HTS Superconducting Cavity (OJ Kwon)

KAIST 1571

- > SC Resonant Cavity in Axion Haloscope Exp.
 - \checkmark Could improve axion to photon conversion power: enhancing scanning speed (> x10)
 - \checkmark Requirement: should sustain high enough Q-factor in high magnetic field (up to 43 T?)
- Choice of SC at CAPP: well-known ReBCO
 - \checkmark Reasonably low surface resistance
 - ✓ Very high Hc2 (~100 T) and high depinning frequency (~100 GHz)
 - ✓ "biaxially textured"
 - ✓ Technically challenging to grow on 3D surface
 - ✓ But, many "grain aligned" high quality tapes are commercially available.
- How can we attach those tapes on the inner surface of resonant cavity?
 3D surface with planar objects: polygon cavity based on TM010 mode (called "melon cut")
 Excellent results (Q>500,000) even up to 8 T (further improvement in progress)
 Physics data run w/ 0.5M Q-factor is complete: will publish soon
- Much higher Q-factor (>10,000,000) achieved
 Bigger HTS superconducting cavity?







HTS Superconducting Cavity (tapes)



Overlaying test results over existing data Extend to lower temp. (mK) Test cavity (w/ rutile in the middle) Comparing Fujikura w/superpower @150 mK



STE OF C



HTS Superconducting Cavity

KAIST 1571 19

Eliminating edge defects Reaches Q ~ 3.7M, first time Q > Q_{axion}





13 million Q-factor HTS Cavity







- \succ HTS cavity Q-factor can be more than 10 times larger than axion's (~ 10⁶)
- > The scan rate could be enhanced two orders of magnitude compared to using copper cavities



HTS Microwave Cavity Design Simulation





> HTS tape surface quality matters

We learned...

- Gaps between tapes has to be minimized
- Mechanical considerations
- Tape alignment



Axion Experiments at CAPP



| | CAPP- PACE | CAPP- 8TB | CAPP- HF | CAPP- PACE -JPA | CAPP- PACE -JPA- 6cell | CAPP- 8TB -JPA- 8cell | CAPP- PACE -JPA-SC | CAPP- MAX-1.0 | CAPP- AQN-SC | CAPP- HeT-SC | CAPP- 12T-HF- 3cell | CAPP- MAX-1.2 |
|--------------------------|---|--------------------------------------|---------------------------------------|-----------------------|---------------------------------|---------------------------------|-----------------------------|---|---|--|--|--|
| Year | 2018 | 2019 | 2019 | 2020 | 2021 | 2021 | 2021 | 2021 | 2023 | 2023 | 2023 | 2024 |
| Magnet [T] | 8 | 8 | 9 | 8 | 8 | 8 | 8 | 12 | 8 | 8 | 12 | 12 |
| m _a [GHz] | ~2.5 | ~1.6 | ~4.0 | ~2.3 | ~5.6 | ~5.8 | ~2.3 | 1.0 ~ 1.2 | ~2.3 | ~5.4 | ~5.3 | 1.2 ~ 1.5 |
| Δm_a [MHz] | 250 | 200 | 250 | 30 | 80 | >100 | 30 | ~20 | - | > 50 | ~30 | ~ 150 |
| Sensitivity | 10*KSVZ +KSVZ | 4*KSVZ | 10*KSV Z | 2*KSVZ | 3*KSVZ | KSVZ | KSVZ | DFSZ | DFSZ | KSVZ | KSVZ | DFSZ |
| T _{phy} [K] | < 0.05 | < 0.05 | ~2 | ~0.05 | ~0.05 | ~0.03 | ~0.04 | ~30 mK | 60 mK | 30 mK | 30 mK | ~30 mK |
| T _{sys} [K, mK] | ~1 K | ~1 K | ~2 K | ~200 mK | <300 mK | <300 mK | <200 mK | <300 mK | ~200 mK | ~400 mK | ~400 mK | <300 mK |
| Comments | R&D machine: First physics run (coldest axion data) | First result published by CAPP | First multi- cell cavity result | First run with JPA | First run with JPA+6-cell | First run with JPA+8-cell | First run with JPA+SC | CAPP's main axion detector with JPA | Axion Quark Nugget + SC cavity (Q~1.6M) | First run with He tuning + SC cavity (Q~10M) | 3-cell with 12T mag + JPA SC cavity (future) | CAPP's main axion detector with JPA |
| Publication | Published in PRL | Published in PRL | Published in PRL | Published in PRL | | Will publish | Will publish | Published in PRL | Will publish | | Published in PRL | Published in PRX |

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Grahal-Capp



GrAHal-CAPP ► Focus on 1-3 µeV axion mass (200-600 MHz)



- Challenging task of fabricating 538 liter cavity
- SC cavity for > 35 T environment

P. Pugnat's presentation on Tue.

Toward the most sensitive Haloscope worldwide ► Focus first on 1-3 μeV axion mass (200-600 MHz)





GrAHal-CAPP : Phase 1 @ 4K - 50 K cryo-stage operational

- @ t₀+18 months - 4 K cryo-stage operational
- @ t_0 +24 months → 1st run
- GrAHal-CAPP : Phase 2 @ 50 mK - Operational @ t_0 + 42 months
- $\rightarrow 2^{nd}$ run reaching DFSZ, in 2-year integration time







CAPP's main flagship experiment
 Powerful 12 Tesla Nb3Sn superconducting magnet with big 32 cm bore
 Achieved DFSZ sensitiviity around 1 GHz
 Total T_{sys} reaches 100 ~ 200 mK
 Physical temp. of cavity (30 mK)



➢ First scan: ~20 MHz (4.51~4.59 µeV) Published in PRL
 ➢ Second + Third scan: ~58 + 150 MHz scan ~@1 GHz: complete, published in PRX







Fourth scan: ~300 MHz scan (1.200 - 1.500 GHz) next

- Target physical temperature of cavity T_{phy} < 30 mK
- Bundled 6 JPAs : T_{sys} < 200 mK \rightarrow > 3 MHz / day
- Adding SC cavity:
 - Speed up more: 36 liter 100% SC cavity is implemented
 - Q-factor > 1M will speed up the search ~ 10 times
- Engineering Run in progress







Fourth scan: ~300 MHz (1.2 - 1.5 GHz) Preparing 6 JPAs : T_{sys} < 200 mK → > 3 MHz / day



> 5th scan: ~250 MHz (1.50 - 1.75 GHz)











From O. Kwon's presentation









From O. Kwon's presentation





Axion Experiments at CAPP







CAPP-MAX + JPA + SC cavity + HF

- 12T LTS magnet working as expected
- Physics data (w/ 6 JPAs)
- $\Delta f \sim 300 \text{ MHz}$ for DFSZ sensitivity
- Ultra-light 100% Superconducting Cavity (36 liters)
- High frequency R&D in progress for next freq. range





Axion Experiments at CAPP In PRX (Aug. 2024) KAIST



TUTE OF SC

September 19th 2024









1mm



Axion Experiments at CAPP

Beyond 2024



SUTE OF SE



September 19th 2024



CAPP's Winning Strategy



- Objective: pinpointing axion mass (discovery) in mass range 1-10 (near future) and 10-100 GHz (requires more R&D)
- > Scanning Speed is the key to cover the possible axion mass range, asap
- Required Technology: cryogenics + big & powerful magnet, quantum noise limited amplifiers and high Q-factor superconducting cavity
- CAPP is equipped with everything required now (plus multiple fridges) and ready to race!

CAPP's advantage:

- Powerful Flagship exp. w/ 12T-32cm SC magnet + 4 dry dilution refrigerators
- > Achieved the lowest system noise temperature (< 200 mK) among existing experiments
- Physics data w/ HTS superconducting cavities (> 10 times boost) (new gen. SC cavity shows >10M Q-factor even under 8T field)







CAPP has successfully established multiple haloscope axion dark matter experiments in Korea, the best axion research facility in the world

Through R&Ds we are ready to speed up our axion search and presently, we are collecting data with DFSZ sensitiivity for axion dark matter search in our flagship experiment, CAPP-MAX

>HTS SC cavity in all our Haloscope

Major improvement is expected with CAPP-MAX for next 5 years QA + SC + HF (1-10 GHz)

>Stay tuned!





Thank You For Your Attention!



Discovered by WC in Bogota, Colombia (Sep. 14th, 2018)



Global Trend in Axion Research



Axion vs WIMP (INSPIRE HEP)



- Axion solves both strong CP problem and a dark matter puzzle
- Two most popular dark matter candidates: Axion and WIMP, but axion is gaining momentum compared to WIMP
- Presently, there are >30 axion experiments worldwide with front-runners including



Axion Detection (experimentalist's view)



- Haloscope: the most sensitive method known today, but
- Still very wide range of mass to scan...





- Needs innovations/breakthroughs to speed up the search!
- Taking advantage of the advancement of Superconductor technology (quantum noise-limited amp + superconducting tapes)







- What we achieved
 - > 7 refrigerators (6 magnets) operational at CAPP
 - > 4 (better than KSVZ sensitivity) experiments running...
- R&Ds needed to search 1~10 GHz in 5 years
 - > Optimizing cryo-RF receiver with quantum amplifiers (JPA) at or near SQL
 - ✓ Ultra-low physical temperature (< 25 mK)</p>
 - $\checkmark~$ Reduce the total noise ~ 120 mK @ 1 GHz
 - ✓ >5 MHz /day scanning speed
 - > Developing Superconducting Cavity under High Magnetic Field
 - $\checkmark~$ HTS cavity could enhance the scan speed $\,$ > 10 times
 - Developing High-Frequency Cavities
 - \checkmark Avoid sacrificing volume of the cavity
 - Multi-cavity, dielectric cavity and photonic gap crystal
 - ✓ Superfluid He tuning





Multi-cell (8) cavity + JPA (6 GHz) (S. Lee)





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Soohyung Lee (IBS-CAPP)



Axion Search w/ SC Cavity (D. Ahn)





- Cavity Tuning
 - Tunability (2270 2295 MHz)
 - Form factor ($C \approx 0.6$)
- Total System Noise ($T_{sys} = T_{eff} + T_{add}$)
 - Josephson Parametric Amplifier + HEMT
 - Effective cavity noise temperature ($T_{eff} \approx 60 \text{mK}$)
 - Added noise by the receiver chain ($T_{add} \approx 115$ mK) \bullet





AQN Experiment at CAPP (Jinsu Kim)



CAPP's first Axion Quark Nugget experiment w/ haloscope
 SC taped cavity (Q > 1.6 M) is implemented

1e6 For Axion Quark Nugget Search Session 02, Thu, Dr. Jinsu Kim Quality Factor TM010 TM011 Magnetic Field (T)

Demonstrate the ability of analyzing high-Q data

Develop AQN analysis (can be done in parallel using haloaxion data)





Superfluid Helium Tuning w/ SC cavity (H. Byun)



- Applicable to superconductor cavity and metamaterial cavity
- No significant change in field distribution, form factor and Q-factor

DAQ run with JPA + SC cavity (Q~13M) + superHe is in progress



Concept of Superfluid
Helium (
$$\epsilon_r \approx 1.057$$
) tuning

•
$$f_{TM010} = \frac{1}{2\pi\sqrt{\mu\epsilon}} \frac{2.405}{R}$$

•
$$\frac{f_{empty} - f_{LHe}}{f_{LHe}} = \sqrt{\epsilon_{LHe}} - 1$$
$$\approx 0.028, ~3\% frequence$$



SNITT

High-Frequency Cavity Design (J. Jeong)



High-Frequency Cavity Designs at IBS-CAPP Multiple-cell Cavity [J. Jeong *et al.*, Phys. Lett. B **777**, 412] Wheel Tuning Mechanism [J. Kim *et al.*, J. Phys. G: Nucl. Part. Phys. **47**, 035203] Tunable Photonic Crystal [S. Bae *et al.*, Phys. Rev. D **107**, 015012]

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